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## A TREATISE

## S U R V E Y I N G

COMPRISING THE THEOR Y AND THE PRACTICE

BY<br>WILLIAM M. GILLESPIE, LL. D. formerly professor of civil engineering in union college

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## PREFACE.

Gillespie's "Land-Surveying" was first printed in 1851, for use in Professor Gillespie's classes in Union College. It was published in 1855, and very soon became the standard authority on land-surveying.

In the preface to the first edition Professor Gillespie says:
"Land-surveying is perhaps the oldest of the mathematical arts. Indeed, geometry itself, as its name-'land-measuring' -implies, is said to have arisen from the efforts of the Egyptian sages to recover and to fix the landmarks annually swept away by the inundations of the Nile. The art is also one of the most important at the present day, as determining the title to land, the foundation of the whole wealth of the world. It is, besides, one of the most useful as a study, from its striking exemplifications of the practical bearings of abstract mathematics. But, strangely enough, surveying has never yet been reduced to a systematic and symmetric whole. To effect this, by basing the art on a few simple principles and tracing them out into their complicated ramifications and varied applications (which extend from the measurement of 'a mowing-lot' to that of the heavens), has been the earnest endeavor of the present writer.
"The work, in its inception, grew out of the author's own needs. Teaching surveying, as preliminary to a course of civil engineering, he found none of the books in use (though very excellent in many respects) suited to his purpose. He was, therefore, compelled to teach the subject by a combination of
faniliar lectures on its principles and exemplifications of its practice. His notes continually swelling in bulk, gradually became systematized in nearly their present form. His system has thus been fully tested, and the present volume is the result.
"A double object has been kept in view in its preparation: viz., to produce a very plain introduction to the subject, easy to be mastered by the young scholar or the practical man of little previous acquirement, the only prerequisites being arithmetic and a little geometry; and at the same time to make the instruction of such a character as to lay a foundation broad enough and deep enough for the most complete superstructure which the professional student may subsequently wish to raise upon it."

In the preface to the "Land-Surveying," Professor Gillespie announced that another volume, on "Leveling and Higher Surveying," was to follow. This work was, at the time of his death, in 1868, unfinished.

The same method was pursued in its preparation as for the "Land-Surveying." Parts of it had been printed for class use, and a large part of the book had been given in the form of lectures to the Professor's classes. It was completed by the editor of this volume, and published in $18 \% 0$.

The two volumes, "Land-Surveying" and "Leveling and Higher Surreying," are now revised and united in this rolume.

The best authorities have been consulted, in order to render the work as reliable as possible.

The writer is under obligations to many friends for assistance in the work of revision, and especially to E. P. Dicker, C. E., for a large part of "Mining-Surveying;" and to Professor T. W. Wright, C. E., for the formula and table in gradienter measurement, and other valuable assistance.

Cady Stalet.

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## geveral division of The subject.

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## PART I.

## LAND-SURVEYING.

## CHAPTER I.

General privciples and fundamental operations.

## DEFINITIONS AND METHODS.

1. Surveying is the art of making such measurements as will determine the relative positions of any points on the surface of the earth ; so that a $M a p$ of any portion of that surface may be drawn, and its Content calculated.
2. The position of a point is said to be determined, when it is known how far that point is from one or more given points, and in what direction therefrom ; or how far it is in front of them or behind them, and how far to their right or to their left, etc. ; so that the place of the first point, if lost, could be again found by repeating these measurements in the contrary direction.

The "points" which are to be determined in Surveying are not the mathematical points treated of in Geometry, but the corners of fences, boundary stones, trees, and the like, which are mere points in comparison with the extensive surfaces and areas which they are the means of determining. In strictness, their centers should be regarded as the points alluded to.

A straight Line is "determined," that is, has its length and its position known and fixed, when the points at its extremities are determined ; and a plane Surface has its form and dimensions determined when the lines which bound it are determined. Consequently, the determination of the relative positions of points is all that is necessary for the principal objects of Surveying ; which
are to make a map of any surface, such as a field, farm, State, etc., and to calculate its content in square feet, acres, or square miles. The former is an application of Drafting, the latter of Mensuration.

The position of a point may be determined by a variety of methods. Those most frequently employed in Surveying are the following-all the points being supposed to be in the same plane :
3. First Method. By measuring the distances from the required point to two given points.

Thus, in Fig. 1, the point $S$ is "determined," if it is known to Fig. 1. be one inch from $A$, and half an inch from
 B ; for its place, if lost, could be found by describing two arcs of circles, from $A$ and $B$ as centers, and with the given distances as radii. The required point would be at the intersection of these arcs.

In applying this principle in surveying, $S$ may represent any station, such as a corner of a field, an angle of a fence, a tree, a house, etc. If, then, one corner of a field be 100 feet from a second corner, and 50 feet from a third, the place of the first corner is known and determined with reference to the other two.

There will be two points fulfilling this condition, one on each side of the given line, but it will always be known which of them is the one desired.

In Geography, this principle is employed to indicate the position of a town ; as when we say that Buffalo is distant (in a straight line) 295 miles from New York, and 390 from Cincinnati, and thus convey to a stranger acquainted with only the last tro places a correct idea of the position of the first.

In Analytical Geometry, the lines AS and BS are known as "Focal Co-ordinates," the general name "co-ordinates" being applied to the lines or angles which determine the position of a point.
4. Second Method. By measuring the perpendicular distance from the required point to a given line, and the distance thence along the line to a given point.

Thus, in Fig. 2, if the perpendicular distance SC be half an
inch, and CA be one inch, the point S is "determined"; for its place could be again found by measuring one inch from A to C, and half an inch from C, at right angles to Fic. 2. A C, which would fix the point $S$.

The public lands of the United States are laid out by this method, as will be explained in. Chapter VII.

In Geography, this principle is employed under the name of Latitude and Longitude.

Thus, Philadelphia is one degree and fifty-two minutes of longitude east of Washington, and one degree and three minutes of latitude north of it.

In Analytical Geometry, the lines A C and CS are known as "Rectangular Co-ordinates." The point is there regarded as determined by the intersection of two lines, drawn parallel to two fixed lines, or "Axes," and at a given distance from them. These Axes, in the present figure, would be the line A C, and another line, perpendicular to it and passing through A , as the origin.
5. Third Method. By measuring the angle between a given

Fig. 3.
 of it to the required point; and also the length of this latter line.

Thus, in Fig. 3, if we know the angle BAS to be a third of a right angle, and AS to be one inch, the point $S$ is determined; for its place could be found by drawing from $A$, a line making the given angle with AB , and measuring on it the given distance.

In applying this principle in surveying, S , as before, may represent any station, and the line AB may be a fence, or any other real or imaginary line.

In "Compass Surveying," it is a north-and-south line, the direction of which is given by the magnetic needle of the compass.

In Geography, this principle is employed to determine the relative positions of places, by " bearings and distances"; as when we say that San Francisco is 1, 750 miles nearly due west from St.

Louis ; the word "west" indicating the direction, or angle which the line joining the two places makes with a north-and-south line, and the number of miles giving the length of that line.

In Analytical Geometry, the line AS, and the angle BAS, are called "Polar Co-ordinates."
6. Fourth Method. By measuring the angles made with a given line by two other lines starting from given points upon it,

Fig. 4.
 and passing through the required point.

Thus, in Fig. 4, the point $S$ is determined by being in the intersection of the two lines $A S$ and $B S$, which make respectively angles of a half and of a third of a right angle with the line $\mathrm{A} B$, which is one inch long ; for the place of the point could be found, if lost, by drawing from A and B lines making with A B the known angles.

In Geography, we might thus fix the position of St. Louis, by saying it lay nearly due north from New Orleans, and due west from Washington.

In Analytical Geometry, these two angles would be called "Angular Co-ordinates."
7. In Fig. 5 are shown together all the measurements necessary for determining the same point S , bs each of the four preceding methods. In the First Method, we measure the distances AS and BS ; in the Second Method, the distances AC and CS, the latter at right angles to the former ; in the Third Method, the distance A S, and the angle S A B ; and, in
 the Fourth Method, the angles S A B and S B A. In all these methods the point is really determined by the intersection of two lines, either straight lines or ares of circles. Thus, in the First Method, it is determined by the intersection of two circles; in the Second, by the intersection of two straight lines; in the Third, by the intersection of a straight line and a circle ; and, in the Fourth, by the intersection of tro straight lines.
8. Fifth Method. By measuring the angles made with each other by three lines of sight passing from the required point to three points whose positions are known.

Thus, in Fig. 6, the point S is determined by the angles ASB and BSC, made by the three lines SA, SB , and S C.

Geographically, the position of Chicago would be determined by three straight lines passing from it to Washington, Cincinnati, and Mobile, and making known angles with each other ; that of the first and second lines being about one third,

Fig. 6.
 and that of the second and third lines, about one half of a right angle.

From the three lines employed, this may be named the Method of Trilinear Co-ordinates.
9. The position of a point is sometimes determined by the intersection of two lines, which are themselves determined by their

Fig. 7.
 extremities being given. Thus, in Fig. '7, the point S is determined by its being situated in the intersection of $A B$ and $C D$. This method is sometimes employed to fix the position of a station on a railroad line, etc., when it occurs in a place where a stake can not be driven, such as in a pond, and in a few other cases, but is not used frequently enough to require that it should be called a sixth principle of Surveying.
10. These five methods of determining the positions of points produce five corresponding systems of Surveying, which may be named as follows:
I. Diagonal Surveying.
II. Perpendicular Surveying
III. Polar Surveying.
IV. Triangular Surveying.
V. Trilinear Surveying.

The above division of Surveying has been made in harmony with the principles involved and the methods employed.

The subject is, however, sometimes divided with reference to the instruments employed; as the chain, either alone or with cross-staff ; the compass; the transit or theodolite ; the sextant; the plane-table, etc.
11. Surveying may also be divided according to its objects.

In Land Surveying, the content, in acres, etc., of the tract surveyed, is usually the principal object of the survey. A map, showing the shape of the property, may also be required. Certain signs on it may indicate the different kinds of culture, etc. This land may also be required to be divided up in certain proportions; and the lines of division may also be required to be set out on the ground. One or all of these objects may be demanded in Land Surveying.

In Topographical Surveying, the measurement and graphical representation of the inequalities of the ground, or its "relief," i. e., its hills and hollows, as determined by the art of "Levelling," is the leading object.

In Maritime or Hydrographical Surveying, the positions of rocks, shoals, and channels are the chief subjects of examination.

In Mining Surreying, the directions and dimensions of the subterranean passages of mines are to be determined.
12. Surveying may also be divided according to the extent of the district surveyed into Plane and Geodesic. Geodesy takes into account the curvature of the earth, and employs Spherical Trigonometry. Plane Surveying disregards this currature, as a needless refinement except in very extensive surreys, such as those of a State, and considers the surface of the earth as plane, which may safely be done in survers of moderate extent.
13. In all the methods of Land Surreying, there are three stages of operation :

1. Measuring certain lines and angles, and recording them;
2. Drawing them on paper to some suitable scale;
3. Calculating the content of the surface surveyed.

## MAKING THE MEASUREMENTS.

14. The Measurements which are required in Surveying may be of lines or of angles, or of both, according to the Method employed. Each will be successively considered.

## Measuring Straight Lines.

15. The lines, or distances, which are to be measured, may be either actual or visual.

Actual lines are such as really exist on the surface of the land to be surveyed, either bounding it, or crossing it; such as fences, ditches, roads, streams, etc.

Visual lines are imaginary lines of sight, either temporarily measured on the ground, such as those joining opposite corners of a field ; or simply indicated by stakes at their extremities or otherwise. If long, they are "ranged out" by methods to be given.

Lines are usually measured with chains, tapes, or rods, divided into yards, feet, links, or some other unit of measurement.

## 16. Gunter's Chain.

 This is the measure most commonly used in Land Surveying. It is 66 feet, or 4 rods long.* Eighty such chains make one mile.It is composed of one hundred pieces of iron or steel wire, or
 links, each bent at the end into a ring, and connected with the

[^1]ring at the end of the next piece by another ring. Sometimes two or three rings are placed between the links. The chain is then less liable to twist and get entangled or "kinked." Two or more swivels are also inserted in the chain, so that it may turn around without twisting. Every tenth link is marked by a piece of brass, having one, two, three, or four points, corresponding to the number of tens which it marks, counting from the nearest end of the chain.* The middle or fiftieth link is marked by a round piece of brass.

The hundredth part of a chain is called a link. $\dagger$ The great advantage of this is that, since links are decimal parts of a chain, they may be so written down, 5 chains and 43 links being $5 \cdot 43$ chains, and all the calculations respecting chains and links can then be performed by the common rules of decimal arithmetic. Each link is $7 \cdot 92$ inches long, being $=66 \times 12 \div 100$.

The following table will be found convenient:

| Chains. | Feet. | Chains. | Feet. |
| :---: | :---: | :---: | :---: |
| 0.01 | $0 \cdot 66$ | $1 \cdot 0$ | 65. |
| $0 \cdot 02$ | $1 \cdot 32$ | $2 \cdot$ | 132. |
| 0.03 | $1 \cdot 98$ | 3. | 198. |
| $0 \cdot 04$ | $2 \cdot 64$ | 4. | 264 - |
| $0 \cdot 05$ | $3 \cdot 30$ | 5. | $330^{\circ}$ |
| $0 \cdot 06$ | $3 \cdot 96$ | 6. | 396 |
| $0 \cdot 07$ | $4 \cdot 62$ | 7. | 462 . |
| $0 \cdot 08$ | $5 \cdot 28$ | S. | 528. |
| 0.09 | $5 \cdot 94$ | 9. | 594. |
| $0 \cdot 10$ | 660 | $10^{\circ}$ | $660^{\circ}$ |
| $0 \cdot 20$ | 13.20 | 20. | 1320. |
| $0 \cdot 30$ | 19.80 | $30^{\circ}$ | $1980^{\circ}$ |
| $0 \cdot 40$ | $26 \cdot 40$ | $40^{\circ}$ | $2640^{\circ}$ |
| $0 \cdot 50$ | $33 \cdot 00$ | 50. | $3300^{\circ}$ |
| $0 \cdot 60$ | $39 \cdot 60$ | 60. | $3960{ }^{\circ}$ |
| $0 \cdot 70$ | 4620 | 70. | $4620^{\circ}$ |
| $0 \cdot 80$ | $52 \cdot 80$ | $80^{\circ}$ | 5280 |
| 0.90 | $59 \cdot 40$ | $90^{\circ}$ | 5940 - |
| 1.00 | 66.00 | $100^{\circ}$ | $6600^{\circ}$ |


| Feet. | Links. | Feet. | Litils. |
| :---: | :---: | :---: | :---: |
| $0 \cdot 10$ | $0 \cdot 15$ | 10. | $15 \cdot 2$ |
| 0.20 | $0 \cdot 30$ | 15. | 22.7 |
| $0 \cdot 25$ | 0.38 | 20. | $30 \cdot 3$ |
| $0 \cdot 30$ | $0 \cdot 45$ | 25. | $37 \cdot 9$ |
| 0.40 | 0.60 | $30^{\circ}$ | $45 \cdot 4$ |
| $0 \cdot 50$ | 0.76 | 33. | $50 \cdot 0$ |
| $0 \cdot 60$ | 0.91 | 35. | 53.0 |
| 0.70 | $1 \cdot 06$ | $40^{\circ}$ | $60 \cdot 6$ |
| 0.75 | 1.13 | 45. | 68.2 |
| 0.80 | $1 \cdot 21$ | 50. | 75.8 |
| 0.90 | 136 | $55^{\circ}$ | $83 \cdot 3$ |
| $1 \cdot 00$ | 1.52 | 60. | $90 \cdot 9$ |
| 2. | 30 | $65^{\circ}$ | $98 \cdot 5$ |
| 3. | $4 \cdot 5$ | $70^{\circ}$ | $106 \cdot 1$ |
| 4. | $6 \cdot 1$ | $75^{\circ}$ | 113.6 |
| 5. | $7 \cdot 6$ | 80. | 121.2 |
| 6 | $9 \cdot 1$ | 85. | 128.8 |
| 7 | $10 \cdot 6$ | $90^{.}$ | 136.4 |
| 8 | $12 \cdot 1$ | 95 | $143 \cdot 9$ |
| $9 \cdot$ | $13 \cdot 6$ | $100^{\circ}$ | $151 \cdot 5$ |

[^2]To reduce links to feet, subtract from the number of links as many units as it contains hundreds ; multiply the remainder by 2 and divide by 3 .

To reduce feet to links, add to the given number half of itself, and add one for each hundred (more exactly, for each ninety-nine) in the sum.

The chain is liable to be lengthened by its rings being pulled open, and to be shortened by its links being bent. It should therefore be frequently tested by a carefully measured length of 66 feet, set out by a standard measure on a flat surface, such as the top of a wall, or on smooth level ground between two stakes, their centers being marked by small nails. It may be left a little longer than the true length, since it can seldom be stretched so as to be perfectly horizontal and not hang in a curve, or be drawn out in a perfectly straight line.* Distances measured with a perfectly accurate chain will always and unavoidably be recorded as longer than they really are. To insure the chain being always strained with the same force, a spring, like that of a spring-balance, is sometimes placed between one handle and the rest of the chain.

If a line has been measured with an incorrect chain, the true length of the line will be obtained by multiplying the number of chains and links in the measured distance by 100, and dividing by the length of the standard distance, as given by measurement of it with the incorrect chain. The proportion here employed is this: $A s$ the length of the standard given by the incorrect chain is to the true length of the standard, so is the length of the line given by the measurement to the true length. Thus, suppose that a line has been measured with a certain chain, and found by it to be ten chains long, and that the chain is afterward found to hare been so stretched that the standard distance measured by it appears to be only 99 links long. The measured line is therefore longer than it had been thought to be, and its true length is obtained by multiplying 10 by 100 , and dividing by 99 .

[^3]17. Pins. Ten iron pins, or " arrows," usually accompany the chain.* They are about a foot long, and are made of stout iron wire, sharpened at one end, and bent into a ring at the other. Pieces of red and white cloth should be tied to their heads, so that they can be easily found in grass, dead leaves, etc.

They should be strung on a ring, which has a spring-catch to retain them. Their usual form is shown in Fig. 9. Fig. 10 shows another form, made very large, and therefore
 very heavy near the point, so that, when held by the top and dropped, it may fall vertically. The uses of this will be seen presently.

On irregular ground, two stout stakes, about six feet long, are needed to put the forward chain-man in line, and to enable whichever of the two is lowest to raise his end of the chain in a truly vertical line, and to strain the chain straight.
A number of long and slender rods are also necessary for "ranging out." lines between distant points.
18. How to Chain. Two men are required-a forward chainman and a hind chain-man, or leader and follower. The latter takes the handles of the chain in his left hand, and the chain itself in his right hand, and throws it out in the direction in which it is to be drawn. The former takes a handle of the chain and one pin in his right hand, and the other pins (and the staff, if used), in his left hand, and draws out the chain. The follower then walks beside it, examining carefully that it is not twisted or bent. He then returns to its hinder end, which he holds at the beginning of the line to be measured, puts his eye exactly orer it and. by the words "Right," "Left," directs the leader how to put his staff, or the pin which he holds up, "in line," so that it may seem to cover and hide the flag-staff, or other object at the end of the line. The leader all the while keeps the chain tightly stretched, and his

[^4]end of it touching his staff. Every time he moves the chain, he should straighten it by an undulating shake. When the staff (or pin) is at last put "in line," the follower says "Down." The leader then puts in the single pin precisely at the end of the chain, and replies "Down." The follower then (and never before hearing this signal that the point is fixed) loosens his end of the chain, retaining it in his hand. The leader draws on the chain, making a step to one side of the pin just set, to avoid dragging it out. He should keep his eye steadily on the object ahead, or, in a hollow, should line himself approximately by looking back. The follower should count his steps, so as to know where to look for the pin in high grass, etc. As he approaches the pin, he calls "Halt." On reaching it, he holds the handle of the chain against it, pressing his knee against both to keep the pin firm. He then, with his eye over the pin, "lines" the leader as before. When the "Down" has been again called by the follower, and answered by the leader, the former pulls out the pin with the chain-hand, and carries it in his other hand, and they go on as before.* The operation is repeated till the leader has arrived at the end of the line, or has put down all his pins.

When the leader has put down his tenth pin, he draws on the chain its length farther, and, after being lined, puts his foot on the handle to keep it firm, and calls "Tally." The follower then drops his end of the chain, goes up to the leader and gives him back all the pins, both counting them to make sure that none have been lost. One pin is then put down at the forward end of the chain, and they go on as before.

Some surveyors cause the leader to call "tally" at the tenth pin, and then exchange pins; but then the follower has only the hole made by the pin, or some other indefinite mark, to measure from.

Eleven pins are sometimes preferred, the eleventh being of brass, or otherwise different from the rest, and being used to mark

[^5]the end of the elerenth chain ; another being substituted for it before the leader goes on.

The two chain-men may change duties at each change of pins, if they are of equal skill, but the more careful and intelligent of two laborers should generally be made "follower."

When the leader reaches the end of the line, he stops, and holds his end of the chain against it. The follower drops his end and counts the links beyond the last pin, noting carefully on which side of the "fifty" mark it comes. Each pin now held by the follower, including the one in the ground, represents one chain; each time "tally" has been called, and the pins exchanged, represents ten chains, and the links just counted make up the total distance.
19. Tallies. In chaining very long distances, there is danger of miscounting the number of "tallies," or tens. To aroid mistakes, pebbles, etc., may be changed from one pocket into another at each change of pins; or bits of leather on a cord may be slipped from one side to the other ; or knots tied on a string ; but the best plan is the following : Instead of ten iron pins, use nine iron pins, and four, or eight, or ten pins of brass, or very much longer than the rest. At the end of the tenth chain, the iron pins being exhausted, a brass pin is put down by the leader. The follower then comes up, and returns the nine iron pins, but retains the brass one, with the additional advantage of having this pin to measure from. At the end of the twentieth chain, the same operation is repeated ; and so on. When the measurement of the line is completed, each brass pin held by the follower counts ten chains, and each iron pin one, as before.
20. Chaining on Slopes. All the distances employed in Landsurveying must be measured horizontally, or on a lerel. When the ground slopes, it is therefore necessary to make certain allowances or corrections. If the slope be gentle, hold the up-hill end of the chain on the ground, and raise the down-hill end till the chain is level. To insure the elcrated end being exactly orer the desired spot, raise it along a staff kept rertical, or drop a pin held
by the point with the ring downward (if you have not the heavy pointed ones shown in Fig. 10), or, which is better, use a plumbline. A person standing beside the chain, and at a little distance from it, can best tell if it be nearly level. If the hill be so steep that a whole chain can not be held up level, use only half or quarter of it at a time.

Fig. 11.
 Great care is necessary in this operation. To measure down a steep hill, stretch the whole chain in line. Hold the upper end fast on the ground. Raise up the 20 or 30 . link-mark, so that that portion of the chain is level. Drop a plumb-line or pin. Then let the follower come forward and hold down that link on this spot, and the leader hold up another short portion, as before. Chaining down a slope is more accurate than chaining up it, since in the latter case the follower can not easily place his end of the chain exactly over the pin.

A more accurate, though more troublesome, method, is to measure the angle of the slope, and make the proper allowance by calculation, or by a table, previously prepared. The correction being found, the chain may be drawn forward the proper number of links, and the correct distance of the various points to be noted will thus be obtained at once, without any subsequent calculation or reduction. If the survey is made with the Transit provided with a vertical circle, the slope of the ground can be measured directly. A "Tangent Scale," for the same purpose, may be formed on the sides of the sights of a Compass. It will be described when the instrument is explained.

In the following table, the first column contains the angle which the surface of the ground makes with the horizon; the second column contains its slope, named by the ratio of the perpendicular to the base ; and the third, the correction in links for each chain measured on the slope, i. e., the difference between the hypothenuse, which is the distance measured, and the horizontal base, which is the distance desired.

TABLE FOR CHAINING ON SLOPES.

| Angle. | Slope. | Correction in links. | Angle. | Slope. | Correction in links. |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $3^{\circ}$ | 1 in 19 | $0 \cdot 14$ | $13^{\circ}$ | 1 in $4 \frac{1}{2}$ | $2 \cdot 56$ |
| $4^{\circ}$ | 1 in 14 | $0 \cdot 24$ | $14^{\circ}$ | 1 in 4 | $2 \cdot 97$ |
| $5^{\circ}$ | 1 in $11 \frac{1}{2}$ | $0 \cdot 38$ | $15^{\circ}$ | 1 in 4 | $3 \cdot 41$ |
| $6^{\circ}$ | 1 in $9 \frac{1}{2}$ | $0 \cdot 55$ | $16^{\circ}$ | 1 in $3 \frac{3}{4}$ | $3 \cdot 87$ |
| $7{ }^{\circ}$ | 1 in 8 | 0.75 | $17^{\circ}$ | 1 in $3 \frac{1}{2}$ | $4 \cdot 37$ |
| $8^{\circ}$ | 1 in 7 | 0.97 | $18^{\circ}$ | 1 in $3 \frac{1}{4}$ | $4 \cdot 89$ |
| $9^{\circ}$ | 1 in $6 \frac{1}{2}$ | 1-23 | $19^{\circ}$ | 1 in 3 | $5 \cdot 45$ |
| $10^{\circ}$ | 1 in 6 | $1 \cdot 53$ | $20^{\circ}$ | 1 in $2 \frac{3}{4}$ | $6 \cdot 03$ |
| $11^{\circ}$ | 1 in $5 \frac{1}{4}$ | $1 \cdot 84$ | $25^{\circ}$ | 1 in 2 | $9 \cdot 37$ |
| $12^{\circ}$ | 1 in $4 \frac{3}{4}$ | $2 \cdot 19$ | $30^{\circ}$ | 1 in $1 \frac{3}{4}$ | $13 \cdot 40$ |

21. Chaining is the fundamental operation in all kinds of Surveying. It has for this reason been rery minutely detailed. The "follower" is the most responsible person, and the surveyor will best insure his accuracy by taking that place himself. If he has to employ inexperienced laborers, he will do well to cause them to measure the distance between any two points, and then remeasure it in the opposite direction. The difference of their two results will impress on them the necessity of great carefulness.

To "do up" the chain, take the middle of it in the left hand, and with the right hand take hold of the doubled chain just beyond the second link; double up the two links between your hands, and continue to fold up two double links at a time, laying each pair obliquely across the others, so that when it is all folded up the handles will be on the outside, and the chain will have an hour-glass shape, easy to strap up and to carre.
22. Tape. Though the chain is most usually emplosed for the principal measurements of Survering, a tape-line, dirided on one side into links, and on the other into feet and inches, is more convenient for some purposes. It should be tested rery frequently, particularly after getting wet, and the correct length marked on it at every ten feet. A "Metallic Tape," less liable to stretch, is manufactured, in which fine wires form its warp. When the tape is being wound up, it should be passed between two fingers to prevent its twisting in the box, which would make it necessary to unscrem its nut to take it out and untwist it. While in use, it
should be made portable by being folded up by arm's lengths, instead of being wound up.

A "Steel Tape," made of a thin ribbon of steel, with the divisions and numbers etched on it, is one of the most accurate measuring instruments. Those intended for accurate measurement have at one end an arrangement for shortening and lengthening the tape to provide for variations in length, due to changes of temperature, and at the other end a level and a spring-balance, so that when measuring the ends of the tape may be held at the same height, and always with the same tension. For methods employed. in making accurate measurements, see Part IV.
23. Substitutes for a chain or a tape may be found in leather driving-lines, marked off with a carpenter's rule, or in a cord knotted at the length of every link. A well-made rope (such as a "patent wove line," woven circularly with the strands always straight in the line of the strain), when once well stretched, wetted, and allowed to dry with a moderate strain, will not vary from a chain more than one foot in two thousand, if carefully used.
24. Rods. When unusually accurate measurements are required, rods are employed. They may be of well-seasoned wood, of glass, of iron, etc. They must be placed in line very carefully end to end, or made to coincide in other ways, as will be explained under "Triangular Surveying," in which the peculiarly accurate measurement of one line is required, as all the others are founded upon it.
25. Pacing, sound, and other approximate means, may be used for measuring the length of a line. The Stadia and Gradienter will be described in Chapter IV.
26. A Perambulator, or "Measuring-Wheel," is sometimes used for measuring distances, particularly roads. It consists of a wheel which is made to roll over the ground to be measured, and whose motion is communicated to a series of toothed wheels within the machine. These wheels are so proportioned that the index-wheel registers their revolutions, and records the whole distance passed
over. If the diameter of the wheel be $31 \frac{1}{2}$ inches, the circumference, and therefore each rerolution, will be $8 \frac{1}{4}$ feet, or half a rod. The roughnesses of the road and the slopes necessarily cause the registered distances to exceed the true measure.

The Odometer is an instrument designed to register the number of revolutions of a wagon-wheel. Knowing the circumference of the wheel to which it is attached, and determining the number of revolutions by the odometer, the distance over which the wheel has passed may be approximately determined.

## Measuring Angles.

27. The angle made by any two lines-that is, the difference of their directions-may be obtained by a great rariety of instru-

Fig. 12.
 ments. All of them are in substance mere modifications of the rery simple one which will now be described, and which any one can make for himself :

Proride a circular piece of rood, and divide its circumference (by any of the methods of Geometrical Drafting) into three hundred and sixty equal parts, or " degrees," and num-
ber them as in the figure. The divisions will be like those of a watch-face, but six times as many. These divisions are termed graduations. The figure shows only erery fifteenth one. In the center of the circle fix a needle, or sharp-pointed wire, and upon this fix a straight stick, or thin ruler placed edgewise (called an alidade), so that it may turn freely on this point and nearly touch the graduations of the circle. Fasten the circle on a staff, pointed at the other end, and long enough to bring the alidade to the height of the eyes. The instrument is now complete. It may be called a Goniometer, or Angle-measurer.

Now let it be required to measure the angle between the lines A B and A C. Fix the staff in the ground, so that its center shall be exactly orer the intersection of the two lines. Turn the alidade so that it points (as determined by sighting along it) to a rod, or
other mark at B , a point on one of the lines, and note what degree it covers-i. e., "The Reading." Then, without disturbing the circle, turn the alidade till it points to C , a point on the other line. Note the new reading. The difference of these readings (in the figure, 45 degrees) is the difference in the directions of the two lines, or is the angle which one makes with the other. If the dis-

Fig. 13.
 tance from A to C be now measured, the point $C$ is "determined," with respect to the points $A$ and $B$, on the Third Principle. Any number of points may be thus determined.

Instead of the very simple and rude alidade, which has been supposed to be used, needles may be fixed on each end of the alidade ; or sights may be added ; or a small straight tube
Fig. 14. - $\oplus$ may be used, one end being covered with a piece of pasteboard in which a very small eye-hole is pierced, and threads, called " cross-bairs," being stretched across the other end of it, as in the figure, so that their intersection may give a more precise line for determining the direction of any point.

When a telescope is substituted for this tube, and supported in such a way that it can turn over, so as to look both backward and forward, the instrument (with various other additions, which, however, do not affect the principle) is called a Transit.
28. Chain Angles. The angle made by any two lines can also be determined without the aid of an angle-measurer. Let it be required to find the angle made by the two lines A B and A C, Fig. 15. Measure off equal distances from $A$ to $B$ and $C$, and also the "tieline " B C. It is evident that the tie-line is the chord of the angle to a radius equal to one of the equal distances measured on the sides. Therefore, divide the length of the tie-line by the length of this distance. The quotient will be the chord of the angle to a radius of one. In the Table of Chords, at the end of this volume, find this quotient, and the number of degrees and minutes corre-
sponding to it gives the angle required. Otherwise, since the chord of any angle equals twice the sine of half the angle, we have this rule : Divide half the tie-line by the measured distance, find in a table of natural sines the angle corresponding to the quotient, and multiply this angle by two, to get the angle desired.

## Surveying without Instruments.

29. Distances by Pacing. Quite an accurate measurement of a line of ground may be made by walking over it at a uniform pace, and counting the steps taken. But the art of walking in a straight line must first be acquired. To do this, fix the eye on two objects in the desired line, such as two trees, or bushes, or stones, or tufts of grass. Walk forward, keeping the nearest of these objects steadily covering the other. Before getting up to the nearest object, choose a new one in line farther ahead, and then proceed as before, and so on. It is better not to attempt to make each of the paces three feet, but to take steps of the natural length, and to ascertain the value of each by walking over a known distance, and dividing it by the number of paces required to traverse it. Erery person should thus determine the usual length of his own steps, repeating the experiment sufficiently often. The French "geographical engineers" accustom themselres to take regular steps of eight tenths of a metre, equal to two feet seven and a half inches. The United States military pace is two feet and six inches. This is regarded as a usual arerage. Quick pacing of 120 such paces per minute gires $3 \cdot 41$ miles per hour. Slow paces, of three feet each and sisty per minute, give 2.04 miles per hour.*

The Pedometer is an instrument which counts the steps taken by one wearing it, without any attention on his part. It is made in the form of a watch, and carried in the pocket. The number of the steps given by the pedometer, multiplied by the length of the step, will give approximately any distance walked over. In one form of this instrument the number of steps is registered on a dial up to 2,500 .

In another form the instrument is intended to be regulated ac-

[^6]cording to the length of step of the person carrying it, and then the distance is registered on the dial in miles.
30. Distances by Visual Angles. Prepare a scale, by marking off on a pencil what length of it, when it is held off at arm's length, a man's height appears to cover at different distances (previously

## Fig. 16.


measured with accuracy) of $100,500,1,000$ feet, etc. To apply this, when a man is seen at any unknown distance, hold up the pencil at arm's length, making the top of it come in the line from the eye to his head, and placing the thumb-nail in the line from the eye to his feet, as in Fig. 16. The pencil having been previously graduated by the method above explained, the portion of it now intercepted between these two lines will indicate the corresponding distance.

If no previous scale have been prepared, and the distance of a man be required, take a foot-rule, or any measure minutely divided, hold it off at arm's length as before, and see how much a man's height covers. Then, knowing the distance from the eye to the rule, a statement by the rule of three (on the principle of similar triangles) will give the distance required. Suppose a man's height, of 70 inches, covers one inch of the rule. He is then seventy times as far from the eye as the rule, and, if its distance be two feet, that of the man is 140 feet. Instead of a man's height, that of an ordinary house, of an apple-tree, the length of a fence-rail, etc., may be taken as the standard of comparison.

To keep the arm immovable, tie a string of known length to the pencil, and hold between the teeth a knot.tied at the other end of the string.
31. Distances by Visibility. The degree of visibility of various well-known objects will indicate approximately how far distant they
are. Thus, by ordinary eyes, the windows of a large house can be counted at a distance of about 13,000 feet, or $2 \frac{1}{2}$ miles ; men and horses will be perceived as points at about half that distance, or $1 \frac{1}{4}$ mile ; a horse can be cleaily distinguished at about 4,000 feet ; the movements of men at 2,600 feet, or half a mile ; and the head of a man, occasionally, at 2,300 feet, and very plainly at 1,300 feet, or a quarter of a mile. The Arabs of Algeria define a mile as "the distance at which you can no longer distinguish a man from a woman." These distances of visibility will of course rary somewhat with the state of the atmosphere, and still more with indiridual acuteness of sight, but each person should make a corresponding scale for himself.
32. Distances by Sound. Sound passes through the air with a moderate and known velocity ; light passes almost instantaneously. If, then, two distant points be risible from each other, and a gun be fired at night from one of them, an observer at the other, noting by a stop-watch the time at which the flash is seen, and then that at which the report is heard, can tell by the intervening number of seconds how far apart the points are, knowing how far sound trarels in a second. Sound mores about 1,098 feet per second in dry air, with the temperature at the freezing-point, $32^{\circ}$ Fahr. For higher or lower temperatures add or subtract $1 \frac{1}{7}$ foot for each degree of Fahrenheit. If a wind blows with or against the morement of the sound, its velocity must be added or subtracted. If it blows obliquely, the correction will evidently equal its relocity multiplied by the cosine of the angle which the direction of the wind makes with the direction of the sound. If the gun be fired at each end of the base in turn, and the means of the times taken, the effect of the wind will be eliminated.

If a watch is not at hand, suspend a pebble to a string (such as a thread drawn from a handkerchief) and count its ribrations. If it be $39 \frac{1}{8}$ inches long, it will vibrate in one second ; if $9 \frac{3}{4}$ inches long, in half a second, etc. If its length is unknown at the time, still count its vibrations; measure it subsequently; and then will the time of its vibration, in seconds, $=\int /\left(\frac{\text { length of string }}{39 \frac{1}{3}}\right)$.
33. Angles. Right angles are those most frequently required in this kind of survey, and they can be estimated by the eye with much accuracy. If other angles are desired, they will be determined by measuring equal distances along the lines which make the angle, and then the line, or chord, joining the ends of these distances, thus forming chain-angles, explained in Article 28.

## Noting the Measurements.

34. The measurements which have been made, whether of lines or of angles, require to be very carefully noted and recorded. Clearness and brevity are the points desired. Different methods of notation are required for each of the systems of surveying which are to be explained, and will therefore be given in their appropriate places.

## DRAWING THE MIAP.

35. A Map of a survey represents the lines which bound the surface surveyed, and the objects upon it, such as fences, roads, rivers, houses, woods, hills, etc., in their true relative dimensions and positions. It is a miniature copy of the field, farm, etc., as it would be seen by an eye moving over it ; or as it would appear, if, from every point of its irregular surface, plumb-lines were dropped to a level surface under it, forming what is called, in geometrical language, its horizontal projection.
36. Platting. A plat of a survey is a skeleton, or outline map. It is a figure "similar" to the original, having all its angles equal and its sides proportional. Every inch on it represents a foot, a yard, a rod, a mile, or some other length, on the ground ; all the measured distances being diminished in exactly the same ratio.

Platting is repeating on paper, to a smaller scale, the measurements which have been made on the ground.

Its various operations may there-
 fore be reduced, in accordance with the principles established in this chapter, to two, viz.: drawing a straight line in a given direction and of a given length;
and describing an arc of a circle with a radius whose length is also given. The only instruments absolutely necessary for this are a straight ruler and a pair of "dividers" or "compasses." Others, however, are often convenient, and will be now briefly noticed.
37. Straight Lines. These are usually drawn by the aid of a straight-edged ruler. But to obtain a very long straight line upon paper, stretch a fine silk thread between any two distant points, and mark in its line various points near enough together to be afterward connected by a common ruler. The thread may also be blackened with burned cork and snapped on the paper, as a carpenter snaps his chalk-line ; but this is liable to inaccuracies, from not raising the line vertically.
38. Arcs. The arcs of circles used in fixing the position of a point on paper are usually described with compasses, one leg of which carries a pencil-point. A convenient substitute is a strip of pasteboard, through one end of which a fine needle is thrust into the given center, and through a hole in which, at the desired distance, a pencil-point is passed, and can thus describe a circle about the center, the pasteboard keeping it always at the proper distance. A string is a still readier, but less accurate, instrument.
39. Parallels. The readiest mode of drawing parallel lines is by the aid of a triangular piece of wood and a ruler. Let A B

Fig. 18.
 be the line to which a parallel is to be drawn, and $C$ the point through which it must pass. Place one side of the triangle against the line, and place the ruler against another side of the triangle. Hold the ruler firm and immorable, and slide the triangle along it till the side of the triangle which had coincided with the given line passes through the giren point. This side will then be parallel to that giren line, and a line drawn by it will be the line required.

Another easy method of drawing parallels is by means of a T-
square, an instrument very valuable for many other purposes. It is nothing but a ruler let into a thicker piece of wood, very truly at right angles to it. For this use of it, one side of the cross-piece must be even or "flush" with the ruler. To use it, lay it on the paper so that one edge of the ruler coincides with the given line A B. Place another ruler against the cross-piece, hold it firm, and slide the T -square along till its edge passes through the given point C, as shown by the lower part of the

Fig. 19.
 figure. Then draw by this edge the desired line parallel to the given line.
40. Perpendiculars. These may be drawn by the various problems given in Geometry, but more readily by a triangle which has one right angle. Place the longest side of the triangle on the given line, and place a ruler against a

Fig. 20.
 second side of the triangle. Hold the ruler fast, and turn the triangle so as to bring its third side against the ruler. Then will the long side be perpendicular to the given line. By sliding the triangle along the ruler, it may be used to draw a perpendicular from any point of the line, or from any point to the line.
41. Angles. These are most easily set out with an instrument called a Protractor. This is usually a semicircle of brass, as in the figure, with its semi-circumference divided into 180 equal parts, or degrees, and numbered in both directions. It is, in fact, a miniature of the instrument (or of half of it) with which the angles have been measured. To lay off any angle at any point of a straight line, place the protractor so that its straight side, the diameter of the semicircle, is on the given line, and the middle of this diameter, which is marked by a notch, is at the given point. With a
needle or sharp pencil make a mark on the paper at the required number of degrees, and draw a line from the mark to the giren point.

Sometimes the protractor has an arm turning on its center and
Fig. 21.

extending beyond its circumference, so that a line can be at once drawn by it when it is set to the desired angle. A Vernier scale is sometimes added to it to increase its precision.

A Rectangular Protractor is sometimes used, the dirisions of degrees being engrared along three edges of a plane scale. The

Fig. 22.

semicircular one is preferable. The objection to the rectangular protractor is that the division corresponding to a degree is rery unequal on different parts of the scale, being usually tro or three times as great at its ends as at its middle."

A Protractor embracing an entire circle, with arms carrying rerniers, is also sometimes employed, for the sake of greater accuracs.
42. Drawing to Scale. The operation of drawing on paper lines whose length shall be a half, a quarter, a tenth, or any other fraction of the lines measured on the ground, is called "Drawing to Scale."

To set off on a line any given distance to any required scale, determine the number of chains or links which each division of the scale of equal parts shall represent. Divide the given distance by this number. The quotient will be the number of equal parts to be taken in the dividers and to be set off.

For example, suppose the scale of equal parts to be a common carpenter's rule divided into inches and eighths. Let the given distance be twelve chains, which is to be drawn to a scale of two chains to an inch. Then six inches will be the distance to be set off. If the given distance had been twelve chains and seventy-five links, the distance to be set off would have been six inches and three eighths, since each eighth of an inch represents twenty-five links.

If the desired scale were three chains to an inch, each eighth of an inch would represent $3 \% \frac{1}{2}$ links; and the distance of $1,2 \% 5$ links would be represented by thirty-four eighths of an inch, or $4 \frac{1}{4}$ inches.

A similar process will give the correct length to be set off for any distance to any scale.

If the scale used had been divided into inches and tenths, as is much the most convenient, the above distances would have become on the former scale $6 \frac{37}{100}$ inches, or nearly $6 \frac{4}{10}$ inches; and on the latter scale $\frac{25}{100}$ inches, coming midway between the second and third tenth of an inch.

Conversely, to find the real length of a line drawn on paper to any known scale, reverse the preceding operation. Take the length of the line in the dividers, apply it to the scale, and count how many equal parts it includes. Multiply their number by the number of chains or links which each represents, and the product will be the desired length of the line on the ground.

This operation and the preceding one are greatly facilitated by the use of the scales to be described in Art. 4\%.
43. Scales. The choice of the scale to which a plat should be drawn-that is, how many times smaller its lines shall be than those which have been measured on the ground-is determined by sereral considerations. The chief one is that it shall be just large enough to express clearly all the details which it is desirable to know. A Farm Survey would require its plat to show every field and building. A State Survey would show only the towns, rivers, and leading roads. The size of the paper at hand will also limit the scale to be adopted. If the content is to be calculated from the plat, that will forbid it to be less than 3 chains to 1 inch.

Scales are named in various ways. They should always be expressed fractionally-i. e., they should be so named as to indicate what fractional part of the real line measured on the ground, the representative line drawn on the paper, actually is. When custom requires a different way of naming the scale, both should be given. It would be still better if the denominator could always be some power of 10 , or at least some multiple of 2 and 5 , such as $\frac{1}{500}$, $\frac{1}{1000}, \frac{1}{2000}, \frac{1}{2500}$, etc. For convenience in printing, these may be written thus: $1: 500,1: 1,000,1: 2,000,1: 2,500$, etc.

Plats of Farm Surveys are usually named as being so many chains to an inch.

Maps of Surveys of States are generally named as being made to a scale of so many miles to an inch.

Maps of Railroad Surveys are said to be so many feet to an inch, or so many inches to a mile.
44. Farm Surveys. If these are of small extent, two chains to one inch (which is $=\frac{1}{2 \times 66 \times 12}=\frac{1}{1554}=1: 1,581$ ) is convenient.

A scale of one chain to one inch (1: 792) is useful for plans of buildings. Three chains to one inch $(1: 2,376)$ is suitable for larger farms. It is the scale prescribed by the English Tithe Commissioners for their first-class maps.

In France, the Cadastre Survess are lithographed on a scale about equiralent to this, being $1: 2,500$. The original plans are drawn to a scale of $1: 5,000$. Plans for the division of property are made on the former scale. When the district exceeds 3,000 acres, the scale is $1: 10,000$. When it exceeas 7,500 acres, the scale is $1: 20,000$. A common scale in France for small surveys is $1: 1,000$, about $1 \frac{1}{4}$ chain to 1 inch.
45. State Surveys. On these surveys smaller scales are necessarily employed.

On the United States Coast and Geodetic Survey all the scales are expressed fractionally and decimally. "The surveys are generally platted originally on a scale of one to ten or twenty thousand, but in some instances the scale is larger or smaller.
"These original surveys are reduced for engraving and publication, and, when issued, are embraced in three general classes: 1, small harbor-charts; 2, charts of bays and sounds ; and, 3 , the General Coast Charts.
"The scales of the first class vary from $1: 10,000$ to $1: 60,000$, according to the nature of the harbor and the different objects to be represented.
"Where there are many shoals, rocks, or other objects, as in Nantucket Harbor and Hell Gate, or where the importance of the harbor makes it necessary, a larger scale of $1: 5,000,1: 10,000$, and $1: 20,000$ is used. But where, from the size of the harbor or its ease of access, a smaller one will point out every danger with sufficient exactness, the scales of $1: 40,000$ and $1: 60,000$ are used, as in the case of New Bedford Harbor, Cat and Ship Island Harbor, New Haven, etc.
"The scale of the second class, in consequence of the large areas to be represented, is usually fixed at $1: 80,000$, as in the case of New York Bay, Delaware Bay and River. Preliminary charts, however, are issued of various scales from 1: 80,000 to $1: 200,000$.
"Of the third class, the scale is fixed at $1: 400,000$ for the General Chart of the Coast from Gay Head to Cape Henlopen, although considerations of the proximity and importance of points on the coast may change the scales of charts of other portions of our exiended coast."

The National Survey of Great Britain is called, from the corps employed on it, the "Ordnance Survey."

The "Ordnance Survey" of the southern counties of England was platted on a scale of 2 inches to 1 mile ( $1: 31,680$ ), and reduced for publication to that of 1 inch to a mile $(1: 63,360)$. The scale of 6 inches to a mile ( $1: 10,560$ ) was adopted for the northern counties of England and for the southern counties of Scotland. The same scale was employed for platting and engraving in outline the "Ordnance Survey" of Ireland. But a map on a scale of 1 inch to 1 mile $(1: 63,360)$ is now published, the former scale rendering the maps too unwieldy and cumbrous for consultation.

The Ordnance Survey of Scotland was at first platted on a scale of 6 inches to 1 mile ( $1: 10,560$ ). That scale has since been abandoned, and it is now platted on a scale of 2 inches to 1 mile ( $1: 31,680$ ), and the general maps are made to only half that scale.

The Ordnance Survey scale for the maps of London and other large towns is 5 feet to 1 mile ( $1: 1,056$ ), or $1 \frac{1}{8}$ chain to 1 inch.

In the "Surveys under the Public Health Act" of England, the scale for the general plan is 2 feet to 1 mile ( $1: 2,640$ ) ; and for the detailed plan 10 feet per mile ( $1: 528$ ), or $\frac{2}{8}$ of a chain per inch.

The Government Survey of France is platted to a scale of $1: 20,000$. Copies are made to $1: 40,000$ : and the maps are engraved to a scale of $1: 80,000$, or about $\frac{3}{4}$ of an inch to 1 mile.

Cassini's famous map of France was on a scale of $1: 86,400$.
The French War Department employ the scales of $1: 10,000,1: 20,000$, $1: 40,000$, and $1: 80,000$ for the topography of France.
46. Railroad Surveys. For these the New York Railroad Law of 1880 directs the horizontal scale of maps which are to be filed in the State Engineer's Office to be 500 feet to $\frac{1}{10}$ of a foot ( $=1: 5,000$ ), and vertical scale for profiles to be 100 feet to $\frac{1}{10}$ of a foot $(=1: 1,000)$.

For the New York Canal Maps a horizontal scale of 2 chains to 1 inch ( $1: 1,584$ ), and a vertical scale of 20 feet to 1 inch, are employed.

The parliamentary "standing orders" prescribe the plans of railroads, prepared for parliamentary purposes, to be made on a scale of not less than 4 inches to the mile ( $1: 15,840$ ); and the enlarged portions (as of gardens, court-yards, etc.) to be on a scale not smaller than 400 feet to the inch ( $1: 4,800$ ). Accordingly, the practice of English railway engineers is to draw the whole plan to a scale of 6 chains, or 393 feet to the inch ( $1: 4,752$ ), as being just within the parliamentary limits.

In France, the engineers of "Bridges and Roads" (Corps des Ponts et Chaussées) employ for the general plan of a road a scale of $1: 5,000$, and for appropriations, 1 : 500.

In the United States Engineer Service the following plans are prescribed: General plans of buildings, 1 inch to 10 feet ( $1: 120$ ).
Maps of grounds, with horizontal curves one foot apart, 1 inch to 50 feet (1:600).
Topographical maps, one mile and a half square, 2 feet to 1 mile ( $1: 2,640$ ). Do., comprising three miles square, 1 foot to one mile ( $1: 5,280$ ).
Do., between four and eight miles square, 6 inches to one mile ( $1: 10,560$ ). Do., comprising nine miles square, 4 inches to one mile ( $1: 15,840$ ). Maps not exceeding 24 miles square, 2 inches to one mile ( $1: 31,680$ ). Maps comprising 50 miles square, 1 inch to one mile ( $1: 63,360$ ).
Maps comprising 100 miles square, $\frac{1}{3}$ inch to one mile ( $1: 126,720$ ). Surveys of roads, canals, etc., 1 inch to 50 feet ( $1: 600$ ).
47. The most conrenient scales of equal parts are those of boxwood, or ivory, which have a fiducial or feather edge, along which they are divided, so that distances can be at once marked off from this edge, without requiring to be taken off with the dividers ; or the length of a given line can be at once read off. Bor-wood is preferable to ivory, as much less liable to warp, or to rary in length with changes in the moisture in the air.

The student can, howerer, make for himself platting-scales of drawing-paper, or Bristol board. Cut a straight strip of this material, about an inch wide. Draw a line through its middle, and set
off on it a number of equal parts, each representing a chain to the desired scale. Subdivide the left-hand division into ten equal

Fig. 23.

parts, each of which will therefore represent ten links to this scale. Through each point of division on the central line, draw (with the T-square) perpendiculars extending to the edges, and the scale is made. It explains itself. The above figure is a scale of 2 chains to 1 inch. On it the distance 220 links would extend between the arrow-heads above the line in the figure ; 560 links extend between the lower arrow-heads, etc.

A paper scale has the great advantage of varying less from a plat which has been made by it, in consequence of changes in the weather, than any other. The mean of many trials showed the difference between such a scale and drawing-paper, when exposed alternately to the damp open atmosphere, and to the air of a warm dry room, to be equal to 005 , while that between box-wood scales and the paper was $\cdot 012$, or nearly $2 \frac{1}{2}$ times as much. The difference with ivory would have been even greater.

Some of the more usual platting-scales are here given in their actual dimensions.

In these five figures, different methods of drawing the scales
Fig. 24.-Scale of 1 chain to 1 inch.

have been given, but each method may be applied to any scale. The first and second, being the most simple, are generally the best. In the third the subdivisions are made by a diagonal line : the dis-

Fig. 25.-Scale of 2 chains to 1 inch.

tances between the various pairs of arrow-heads, beginning with the uppermost, are respectively 310 , 540 , and 270 links.

In the fourth figure, the distances between the arrow-heads are respectively $310,2 \%$, and 540 links.

Fig. 26.-Scale of 3 chains to 1 inch.


In the fifth figure, the scale of 5 chains to 1 inch is subdirided diagonally to only every quarter-chain, or 25 links. The distance

$$
\text { Fig. } 26^{1} \text {.-Scale of } 4 \text { chains to } 1 \text { inch. }
$$


between the upper pair of arrow-heads on it is $12 \pm$ chains, or $12 \cdot 25$; between the next pair of arrow-heads it is 6.50 ; and between the lower pair $14 * \%$.

Fig. 27.-Scale of 5 chains to 1 inch.


A diagonal scale for dividing an inch, or half an inch, into 100 equal parts, is formd on the "plain scale" in every case of instruments.
48. Vernier Scale. This is an ingenious substitute for the diagonal scale. The one given in the following figure dirides an inch into 100 equal parts, and, if each inch be supposed to represent a chain, it gives single links.

Make a scale of an inch divided into teuths, as in the upper scale of the above figure. Take in the dividers eleren of these divisions, and set off this distance from the 0 of the scale to the
left of it. Divide the distance thus set off into 10 equal parts. Each of them will be one tenth of eleven tenths of one inch, i. e.,

Fig. 28.

eleven hundredths, or a tenth and a hundredth, and the first division on the short, or vernier scale, will overlap, or be longer than the first division on the long scale, by just one hundredth of an inch ; the second division will overlap two hundredths, and so on. The principle will be more fully developed in treating of "Verniers."

Now, suppose we wish to take off from this scale $2 \% 5$ hundredths of an inch. To get the last figure, we must take five divisions on the lower scale, which will be 55 hundredths, for the reason just given; 220 will remain, which are to be taken from the upper scale, and the entire number will be obtained at once by extending the dividers between the arrow-heads in the figure from 220 on the upper scale (measuring along its lower side) to 55 on the lower scale ; 254 would extend from 210 on the upper scale to 44 on the lower; 318 would extend from 230 on the upper scale to 88 on the lower. Always begin then with subtracting 11 times the last figure from the given number; find the remainders on the upper scale, and the number subtracted on the lower scale.
49. A plat is sometimes made by a nominally reduced scale in the following manner: Suppose that the scale of the plat is to be ten chains to one inch, and that a diagonal scale of inches, divided into tenths and hundredths, is the only one at hand. By dividing all the distances by ten, this scale can then be used without any further reduction. But if the content is measured from the plat to the same scale, in the manner explained in the next chapter, the result must be multiplied by 10 times 10 . This is called by old surveyors "raising the scale," or "restoring true measure."
50. Sectoral Scales. The Sector (called by the French "Compass of Proportion") is an instrument sometimes convenient for obtaining a scale of equal parts. It is in two portions, turning on a hinge, like a carpenter's pocket-rule. It contains a great number of scales, but the one intended for this use is lettered at its

Fig. 29.
 ends L in English instruments, and consists of two lines running from the center to the ends of the scale, and each divided into ten equal parts, each of which is again subdivided into ten, so that each leg of the scale contains 100 equal parts. To illustrate its use, suppose that a scale of 7 chains to 1 inch is required. Take 1 inch in the dividers, and open the sector till this distance will just reach from the 7 on one leg to the 7 on the other. The sector is then "set" for this scale, and the angle of its opening must not be again changed. Now let a distance of oั 80 links be required. Open the diriders till they reach from 58 to 58 on the two legs, as in the dotted line in the figure, and it is the required distance. Again, suppose that a scale of $2 \frac{1}{2}$ chains to 1 inch is desired. Open the sector so that 1 inch shall extend from 25 to 25 . Any other scale may be obtained in the same manner.

Conversely, the length of any known line to any desired scale can thus be readily determined.
51. Whaterer scale may be adopted for platting the surrey, it should be drawn on the map, both for conrenience of reference and in order that the contraction and expansion caused by changes in the quantity of moisture in the atmosphere may affect the scale and the map alike. When the draming-paper has been wet and glued to a board, and cut off when the map is completed, its contractions hare been found by many obserrations to arerage from one fourth to one half per cent on a scale of 3 chains to an inch
(1:2,376), which would therefore require an allowance of from one half perch to one perch per acre.

A scale made as directed in Art. 47, if used to make a plat on unstretched paper, and then kept with the plat, will answer nearly the same purpose.

Such a scale may be attached to a map by slipping it through two or three cuts in the lower part of the sheet, and will be a very convenient substitute for a pair of dividers in measuring any distance upon it.
52. Scale omitted. It may be required to find the unknown scale to which a given map has been drawn, its superficial content being known. Assume any convenient scale, measure the lines of the map by it, and find the content by the methods to be given in the next chapter, proceeding as if the assumed scale were the true one. Then make this proportion, founded on the geometrical principle that the areas of similar figures are as the squares of their corresponding sides : $A s$ the content found is to the given content, so is the square of the assumed scale to the square of the true scale.

## CALCULATING THE CONTENT.

53. The Content of a piece of ground is its superficial area, or the number of square feet, yards, acres, or miles which it contains.
54. Horizontal Measurement. All ground, however inclined or uneven its surface may be, should be measured horizontally, or as if brought down to a horizontal plane, so that the surface of a hill, thus measured, would give the same content as the level base on which it may be supposed to stand, or as the figure which would be formed on a level surface beneath it by dropping plumb-lines from every point of it.

This method of procedure is required for both geometrical and social reasons.

Geometrically, it is plain that this horizontal measurement is absolutely necessary for the purpose of obtaining a correct plat. In Fig. 30, let A B CD and B C E F be two square lots of ground,
platted horizontally. Suppose the ground to slope in all directions from the point $C$, which is the summit of a hill. Then the lines

Fig. 30.
 $\mathrm{BC}, \mathrm{D} C$, measured on the slope, are longer than if measured on a level, and the field ABCD , of Fig. 30, platted with these long lines, would take the shape ABGD in Fig. 31; and the field BCEF , of Fig. 30, would become B HEF, of Fig. 31. The two adjoining fields would thus overlap each other ; and the same difficulty would occur in every case of platting any two adjoining fields by the measurements made on the slope.
Let us suppose another case, more simple than would erer occur in practice, that of a three-sided field, of equal sides, and composed of three portions, each sloping down uniformly (at the rate of one to one) from one point in the center, as in Fig. 32. Each slope being accurately platted, the three could not come together, but


Fig. 33.
 would be separated as in Fig. 33.

We have here taken the most simple cases, those of uniform slopes. But with the common irregularities of uneren ground, to measure its actual surface would not only be improper, but impossible.

In the social aspect of this question, the horizontal measurement is justified by the fact that no more houses can be built on a hill than could be built on its flat base:

Fig. 34.
 and that no more trees, corn, or other plants, which shoot up rertically, can grow on it; as is represented by the rertical lines in the figure.* Eren if a side-hill

[^7]should produce more of certain creeping plants, the increased difficulty in their cultivation might perhaps balance this. For this reason the surface of the soil thus measured is sometimes called the productive base of the ground.

Again, a piece of land containing a hill and a hollow, if measured on the surface, would give a larger content than it would after the hollow had been filled up by the hill, while it would yet really be of greater value than before.

Horizontal measurement is called the "Method of Cultellation," and superficial measurement the " Method of Development." *

An act of the State of New York prescribes that " the acre, for land-measure, shall be measured horizontally."
55. Unit of Content. The Acre is the unit of land-measurement. It contains 4 Roods. A Rood contains 40 Perches. A Perch is a square Rod; otherwise called a Pole. A Rod is $5 \frac{1}{2}$ yards, or $16 \frac{1}{2}$ feet.

Hence, 1 Acre $=4$ Roods $=160$ Perches $=4,840$ square yards $=43,560$ square feet.

One square mile $=5,280 \times 5,280$ feet $=640$ acres.
Since a chain is 66 feet long, a square chain contains 4,356 square feet ; and, consequently, ten square chains make one acre. $\dagger$

The French units of land-measure are the Are $=100$ square Metres $=0.0247$ acre $=$ one fortieth of an acre, nearly ; and the Hectare $=100$ Ares $=2.47$ acres, or nearly two and a half. Their old land-measures were the "Arpent of Paris," containing 36,800 square feet ; and the "Arpent of Waters and Woods," containing 55,000 square feet.
56. When the content of a piece of land (obtained by any of the methods to be explained presently) is given in square links, as is
of armies, imagine that unequal and hilly ground will contain more houses than a surface which is flat and level. This, however, is not the truth. For, the houses, being raised in a vertical line, form right angles, not with the declivity of the ground, but with the flat surface which lies below, and upon which the hills themselves also stand."

* The former from cultellum, a knife, as if the hills were sliced off ; the latter so named because it strips off or unfolds, as it were, the surface.
$\dagger$ Let the young student beware of confounding 10 square chains with 10 chains square. The former make one acre ; the latter space contains ten acres.
customary, cut off four figures on the right (i. e., divide by 10,000 ) to get it into square chains and decimal parts of a chain ; cut off the right-hand figure of the square chains, and the remaining figures will be Acres. Multiply the remainder by 4, and the figure, if any, outside of the new decimal-point will be Roods. Multiply the remainder by 40 , and the outside figures will be Perches. The nearest round number is usually taken for the Perches ; fractions less than a half-perch being disregarded.*

Thus, $86 \cdot 22$ square chains $=8$ Acres 2 Roods 20 Perches.

$$
\begin{array}{clllll}
\text { Also, } & 64 \cdot 1818 & \text { do. } & =6 \mathrm{~A} . & 1 \mathrm{R} . & 2 \% \mathrm{P} . \\
\text { " } & 43 \cdot \% 664 & \text { do. } & =4 \mathrm{~A} . & 1 \mathrm{R} . & 20 \mathrm{P} .
\end{array}
$$

57. Chain Correction. When a surrey has been made, and the plat has been drawn, and the content calculated; and aftermard the chain is found to have been incorrect, too short or too long, the true content of the land may be found by this proportion : $A s$ the square of the length of the standard given by the incorrect chain is to the square of the true length of the standard, so is the calculated content to the true content. Thus, suppose that the chain used had been so stretched that the standard distance measured by it appears to be only 99 links long ; and that a square field had been measured by it, each side containing 10 of these long chains, and that it had been so platted. This plat, and therefore the content calculated from it, will be smaller than it should be, and the correct content will be found by the proportion $99^{2}$ : $100^{2}$ : : 100 square chains : $102 \cdot 03$ square chains. If the chain had been stretched so as to be 101 true links long, as found by comparing it with a correct chain, the content would be given by this proportion : $100^{2}: 101^{2}:$ : 100 square chains : 102.01 square chains. In the former case, the elongation of the chain was $1_{99} \frac{1}{9}$ true links ; and $100^{2}:\left(101 \frac{1}{99}\right)^{2}:: 100$ square chains $: 102 \cdot 03$ square chains.
58. Boundary-Lines. The lines which are to be considered as bounding the land to be surreyed are often rery uncertain, unless specified by the title-deeds.
[^8]If the boundary be a brook, the middle of it is usually the boundary-line. On tide-waters, the land is usually considered to extend to low-water mark.

Where hedges and ditches are the boundaries of fields, as is almost universally the case in England, the dividing line is generally the top edge of the ditch farthest from the hedge, both hedge and ditch belonging to the field on the hedge side. This varies, however, with the customs of the locality. From three to six feet from the roots of the quick-wood of the hedges are allowed. for the ditches.

## Methods of Calculation.

59. The various methods employed in calculating the content of a piece of ground may be reduced to four, which may be called Arithmetical, Geometrical, Instrumental, and Trigonometrical.
60. First Method.-Arithmetically. From direct measurements of the necessary lines on the ground.

The figures to be calculated by this method may be either the shapes of the fields which are measured, or those into which the fields can be divided by measuring various lines across them.

The familiar rules of mensuration for the principal figures which occur in practice will be now briefly enunciated.
61. Rectangles. If the piece of ground be rectangular in shape, its content is found by multiplying its length by its breadth.
62. Triangles. When the given quantities are one side of a triangle and the perpendicular distance to it from the opposite angle, the content of the triangle is equal to half the product of the side and the perpendicular.

When the given quantities are the three sides of the triangle, add together the three sides and divide the sum by 2 ; from this half sum subtract each of the three sides in turn; multiply together the half sum and the three remainders; take the square root of the product; it is the content required. If the sides of the triangle be designated by $a, b, c$, and their sum

Fig. 35.

by $s$, this rule will give its area $=\sqrt{ }\left[\frac{1}{9} s\left(\frac{1}{2} s-a\right)\right.$ $\left.\left(\frac{1}{2} s-b\right)\left(\frac{1}{2} s-c\right)\right]$.

When two sides of a triangle and the included angle are given, its content equals half the product of its sides into the sine of the included angle. Designating the angles of the triangle by
the capital letters $A, B, C$, and the sides opposite them by the corresponding small letters $a, b, c$, the area $=\frac{1}{2} b c \sin$. A.

When one side of a triangle and the adjacent angles are given, its content equals the square of the given side multiplied by the sines of each of the given angles, and divided by twice the sine of the sum of these angles. Using the same symbols as before, the area $=a^{2} \frac{\sin . \mathrm{B} \cdot \sin . \mathrm{C}}{2 \sin \cdot(\mathrm{~B}+\mathrm{C})}$.

When the three angles of a triangle and its altitude are given, its area, referring to the above figure, $=\frac{1}{2} \mathrm{BD}^{2} \cdot \frac{\sin \cdot \mathrm{~B}}{\sin \cdot \mathrm{~A} \cdot \sin . \mathrm{C}}$.
63. Parallelograms, or four-sided figures whose opposite sides are parallel. The content of a Parallelogram equals the product of one of its sides by the perpendicular distance between it and the side parallel to it.
64. Trapezoids, or four-sided figures, two opposite sides of which are parallel. The content of a Trapezoid equals half the product of the sum of the parallel sides by the perpendicular distance between them.

If the given quantities are the four sides $a, b, c, d$, of which $b$ and $d$ are parallel ; then, making $q=\frac{1}{2}(a+b+c-d)$, the area of the trapezoid will $=\frac{b+d}{b-d} \sqrt{ }[q(q-a)(q-c)(q-b+d)]$.

When tro parallel sides, $b$ and $d$, and a third side, $a$, are giren, and also the angle $C$, which this third side makes with one of the parallel sides, then the content of the trapezoid $=\frac{b+d}{2}, a . \sin . C$.
65. Trapeziums; four-sided figures, none of whose sides are parallel.

A very gross error, often committed as to this figure, is to take the arerage, or half sum of its opposite sides, and multiply them together for the area: thus, assuming the trapezium to be equivalent to a rectangle with these averages for sides.

In practical surveying, it is usual to measure a line across it from corner to corner, thus dividing it into two triangles, whose sides are known, and which can therefore be calculated by Art. 62.

When two opposite sides, and all the angles are given, take one side and its adjacent angles (or their supplements, when their sum exceeds $180^{\circ}$ ), consider them as belonging to a triangle, and find its area be the second formula in Art. 62. Do the same with the other side and its adjacent angles. The difference of the two areas will be the area of the quadrilateral.

When three sides and their two included angles are given, multiply together the sine of one given angle and its adjacent sides. Do the same with the sine of the other given angle and its adjacent sides. Mrultipls together the two opposite sides and the sine of the supplement of the sum of the giren angles. Add together the first two products, and.add also the last product, if the sum of the given angles is more than $180^{\circ}$, or subtract it if this sum be less, and take half the result. Calling the giren sides $p, q, r$, and the angle
between $p$ and $q=\mathrm{A}$; and the angle between $q$ and $r=\mathrm{B}$; the area of the quadrilateral

$$
=\frac{1}{3}\left[p \cdot q \sin \cdot \mathrm{~A}+q \cdot r \cdot \sin \mathrm{~B} \pm p \cdot r \sin .\left(180^{\circ}-\mathrm{A}-\mathrm{B}\right)\right] .
$$

When the four sides and the sum of any two opposite angles are given, proceed thus: Take half the sum of the four given sides, and from it subtract each side in turn. Multiply together the four remainders, and reserve the product. Multiply together the four sides. Take half their product, and multiply it by the cosine of the given sum of the angles increased by unity. Regard the sign of the cosine. Subtract this product from the reserved product, and take the square root of the remainder. It will be the area of the quadrilateral.

When the four sides and the angle of intersection of the diagonals of the quadrilateral are given, square each side; add together the squares of the opposite sides; take the difference of the two sums; multiply it by the tangent of the angle of intersection, and divide by four. The quotient will be the area.

When the diagonals of the quadrilateral and their included angle are given, multiply together the two diagonals and the sine of their included angle, and divide by two. The quotient will be the area.
66. Second Method.-Geometrically. From measurements of the necessary lines upon the plat.
67. Division into Triangles. The plat of a piece of ground having been drawn from the measurements made by any of the methods which will be hereafter explained, lines may be drawn upon the plat so as to divide it into a number of triangles. Four ways of doing this are shown in the figures, viz.: by drawing lines

## Fig. 36.



Fig. 37.


Fig. 38.


Fig. 39.

from one corner to the other corners; from a point in one of the sides to the corners ; from a point inside of the figure to the corners ; and from various corners to other corners. The last method is usually the best. The lines ought to be drawn so as to make the triangles as nearly equilateral as possible.

One side of each of these triangles, and the length of the perpendicular let fall upon it, being then measured, the content of
these triangles can be at once obtained by multiplying their base by their altitude, and diriding by two.

The easiest method of getting the length of the perpendicular, without actually drawing it, is to set one point of the dividers at the angle from which a perpendicular is to be let fall, and to open and shat their legs till an are described by the other point will just touch the opposite side.

Otherwise, a platting scale may be placed so that the zero-point of its edge coincides with the angle, and one of its cross-lines coincides with the side to which a perpendicular is to be drawn. The length of the perpendicular can then at once be read off.

The method of dividing the plat into triangles is the one most commonly employed by surveyors for obtaining the content of a survey, because of the simplicity of the calculations required. Its correctness, however, is dependent on the accuracy of the plat, and on its scale, which should be as large as possible. Three chains to an inch is the smallest scale allowed by the English Tithe Commissioners for plats from which the content is to be determined.

In calculating in this way the content of a farm, and also of its separate fields, the sum of the latter ought to equal the former. A difference of one three-hundredth ( $\frac{1}{300}$ ) is considered allowable.

Some surveyors measure the perpendiculars of the triangles by a scale half of that to which the plat is made. Thus, if the scale of the plat be two chains to the inch, the perpendiculars are measured with a scale of one chain to the inch. The product of the base by the perpendicular thus measured. gives the area of the triangle at once, without its requiring to be dirided by two.

Another way of attaining the same end, with less danger of mistakes, is to construct a new scale of equal parts, longer than those by which the plat was made in the ratio $\sqrt{ } 2: 1$; or $1 \cdot 414: 1$. When the base and perpendicular of a triangle are measured by this new scale, and then multiplied together. the product will be the content of the triangle, without any division by tro, In this method there is the additional adrantage of the greater size and consequent greater distinctness of the scale.

When the measurement of a plat is made some time after it has been drawn, the paper will rery probably hase contracted or expanded so that the scale used will not exactly apply. In that case a correction is necessary. Measure very precisely the present length of some line on the plat, of known length originally. Then make this proportion : As the square of the present length of this line is to the square of its original length, so is the content obtained by the present measurement to the trive content.
68. Graphical Multiplication. Prepare a strip of drawing-paper, of a width exactly equal to two chains on the scale of the plat; i. e., one inch wide, as in the figure, for a scale of two chains to one inch; tro thirds of an inch wide for a scale of three chains; half an inch for four chains, and so on. Draw perpendicular lines across the paper at distances representing one tenth
of a chain on the scale of the triangle to be measured, thus making a platting scale. Apply it to the triangle so that one edge of the scale shall pass through one corner, A, of the triangle, and the other edge through another corner, B ; and note very precisely what divisions of the scale are at these points. Then slide the scale in such a way that the points of the scale which had coincided with A and B shall always remain on the line B A produced, till the edge arrives at the point C . Then will $\mathrm{A}^{\prime} \mathrm{C}$-that is, the distance, or

Fig. 40.

number of divisions on the scale, from the point to which the division A on the scale has arrived, to the third corner of the triangle-express the area of the triangle A B C in square chains.

For, from C draw a parallel to $\mathrm{A} B$, meeting the edge of the scale in $\mathrm{C}^{\prime}$, and draw $\mathrm{C}^{\prime} \mathrm{B}$. Then the given triangle $\mathrm{ABC}=\mathrm{ABC}$. But the area of this last triangle $=\mathrm{A}^{\prime}$ multiplied by half the width of the scale, i. e., $=\mathrm{AC}^{\prime} \times 1=\mathrm{AC}^{\prime}$. But, because of the parallels, $\mathrm{A}^{\prime} \mathrm{C}=\mathrm{A} \mathrm{C}^{\prime}$, therefore the area of the given triangle $A B C=A^{\prime} C$; i. e., it is equal in square chains to the number of linear chains read off from the scale. This ingenious operation is due to M. Cousinery.
69. Division into Trapezoids. A line may be drawn across the field, as in Fig. 41, and perpendiculars drawn to it. The field

Fig. 41.


Fig. 42.

will thus be divided into trapezoids (excepting a triangle at each end), and their content can be calculated by Art. 64.

Otherwise : a line may be drawn outside of the figure, and perpendiculars to it be drawn from each angle. In that case the difference between the trapezoids formed by lines drawn to the outer angles of the figure, and those drawn to the inner angles, will be the content.
70. Division into Squares. Two sets of parallel lines, at right angles to each other, one chain apart (to the scale of the plat) may be drawn over the plat, so as to divide it into

Fig. 43.
 squares, as in the figure. The number of squares which fall within the plat represent so many square chains; and the triangles and trapezoids which fall outside of these may then be calculated and added to the entire square chains which have been counted.

Instead of drawing the parallel lines on the plat, they may better be drawn on a piece of transparent "tracing-paper," which is simply laid upon the plat, and the squares counted as before. The same paper will answer for any number of plats drawn to the same scale. This method is a valuable and easy check on the results of other calculations.

To calculate the fractional parts, prepare a piece of tracing-paper, or glass, by drawing on it one square of the same size as a square of the plat, and subdividing it, by two sets of ten parallels at right angles to each other, into hundredths. This will measure the fractions remaining from the former measurement, as nearly as can be desired.
71. Division into Parallelograms. Draw a series of parallel lines across the plat at equal distances depending on the scale. Thus, for a plat made to a scale of 2 chains to 1 inch, the distance between the parallels should be $2 \frac{1}{2}$ inches; for a scale of 3 chains to 1 inch, $1 \frac{1}{9}$ inch; for a scale of 4 chains to $1 \mathrm{inch}, \frac{5}{8} \mathrm{inch}$; for a scale of 5 chains to 1 inch, $\frac{4}{10}$ inch; and for any scale, make the distance between the parallels that fraction of an inch which would be expressed by 10 dirided by the square of the number of chains to the inch. Then apply a common inch scale, divided on the edge into tenths, to these parallels; and every inch

Fig. 44.

in length of the spaces included between each pair of them will be an acre, and every tenth of an inch will be a square chain.

For, calling the number of chains to the inch, $=n$, and making the width between the parallels $\frac{10}{n^{2}}$ inch, this width will represent $\frac{10}{n^{2}} \times n=\frac{10}{n}$ chains ; and as the inch length represents $n$ chains, their product, $\frac{10}{n} \times n=10$ square chains $=1$ acre .

To measure the triangles at the ends of the strips between the parallels, prepare a piece of glass, or stout tracing-paper, of a width equal to the width between the parallels, and draw a line through its middle longitudinally. Apply it to the oblique line at the end of the space between two parallels, and it will bisect the line, and thus reduce the triangle to an equivalent rectangle, as at A in the figure. When an angle occurs between two parallels, as at $B$ in the figure, the fractional part may be measured by any of the preceding methods.

A somewhat similar method is much used by some surveyors, particularly in Ireland-the plat being made on a scale of 5 chains to 1 inch, parallel lines being drawn on it, half an inch apart, and the distances along the parallels being measured by a scale, each large division of which is $\frac{8}{10}$ inch in length. Each dirision of this scale indicates an acre; for it represents 4 chains, and the distance between the parallels is $2 \frac{1}{2}$ chains. This scale is called the "Scale of Acres."
72. Addition of Widths. When the lines of the plat are very irregularly curved, as in the figure, draw across it a number of equidistant lines, as near together as the case may seem to require. Take a straightedged piece of paper, and apply one edge of it to the middle of the tirst space, and mark its length from one end; apply the same edge to the middle of the next space, bringing

Fra. 45.
 the mark just made to one end, and making another mark at the end of the additional length; so go on, adding the length of each space to the previous ones. When all have been thus measured, the total length, multiplied by the uniform width, will give the content.

## 73. Third Method,-Instrumentally. By performing certain instrumental operations on the plat.

74. Reduction of a many-sided figure to a single equivalent triangle. Any plane figure bounded by straight lines may be reduced to a single triangle, which shall have the same content. This can be done by any instrument for drawing parallel lines.

Let the trapezium, or four-sided figure, shown in Fig. 46, be required to be reduced to a single equivalent triangle. Produce one side of the figure, as 4-1.
 Draw a line from the first to the third angle of the figure. From the second angle draw a parallel to the line just drawn, cutting the produced side in a point $1^{\prime}$. From the point 1' draw a line to the third angle. A triangle ( $1^{\prime}$ -$3-4$ in the figure) will thus be formed, which will be equivalent to the original trapezium.

For, the triangle $1-2-3$ taken away from the original figure is equivalent to the triangle $1^{\prime}-1-3$ added to it ; because both these triangles have the same base and also the same altitude, since the vertices of both lie in the same line parallel to the base.

The content of this final triangle can then be found by measuring its perpendicular, and taking half the product of this perpendicular by thebase.

Let the given figure

Fig. 47.
 have five sides, as in Fig. 47. For brevity, the angles of the figure will be named as numbered in the engraving. Produce $5-1$. Join $1-3$. From 2 draw a parallel to $1-3$, cutting the produced base in $1^{\prime}$. Join $1^{\prime}-4$. From 3 draw a parallel to it, cutting the base in $2^{\prime}$. Join $2^{\prime}-4$. Then will the triangle $2^{\prime}-4-5$ be equivalent to the five-sided figure $1-2-3-4-5$, for similar reasons to those of the preceding case.

Let the given figure be $1-2-3-4-5-6-7-8$, as shown in Fig. 48. All the operations are shown by dotted lines, and the finally resulting triangle, $5-7-8$, is equivalent to the original figure of eight sides.

It is best, in choosing the side to be produced, to take one which has a long side adjoining it on the end not produced; so that this long side may form one side of the final triangle, the base of which will therefore be shorter, and will not be cut so acutely by the final line drawn, as to make the point of intersection too indefinite.

75. General Rule. When the given figure has many sides, with angles sometimes salient and sometimes re-entering, the operations of reduction are very liable to errors if the draughtsman attempts to reason out each step. All difficulties, however, will be removed by the following General Rule:

1. Produce one side of the figure, and call it a base. Call one of the angles at the base the first angle, and number the rest in regular succession around the figure.
2. Draw a line from the 1 st angle to the 3 d angle. Draw a parallel to it from the 2 d angle. Call the intersections of this parallel with the base the 1st mark.
3. Draw a line from the 1st mark to the 4th angle. Draw a parallel to it from the 3 d angle. Its intersection with the base is the 2 d mark.
4. Draw a line from the 2 d mark to the 5 th angle. Draw a parallel to it from the 4th angle. Its intersection with the base is the 3 d mark.
5. In general terms, which apply to every step after the first, draw a line from the last mark obtained to the angle whose number is greater by three than the number of the mark. Draw a parallel to it through the angle whose number is greater by two than that of the mark. Its intersection with the base will be a mark whose number is greater by one than that of the preceding mark.

In the concise language of algebra, draw a line from the $n$th
mark to the $n+3$ angle. Draw a parallel to it through the $n+2$ angle, and the intersection with the base will be the $n+1$ mark.
6. Repeat this process for each angle, till you get a mark whose number is such that the angle haring a number greater by three is the last angle of the figure-i. e., the angle at the other end of the base. Then join the last mark to the angle which precedes the last angle in the figure, and the triangle thus formed will be the equivalent triangle required.

In practice it is unnecessary to actually draw the lines joining the successive angles and marks, but the parallel ruler is merely laid on so as to pass through them, and the points where the parallels cut the base are alone marked.
76. It is generally more convenient to reduce half of the figure on one side and half on the other, as is shown in

Fig. 49.
 Fig. 49, which represents the same field as Fig. 47. The equivalent triangle is here $1^{\prime}-3-2$.

When the figure has many angles, they should not be numbered consecutively all the way around, but, after the numbers have gone around as far as the angle where it is intended to hare the vertex of the final triangle, the numbers should be continued from the other angle of the base, as is shown in Fig. г0. In it only the intersections are marked.
A figure with curred boundaries may be reduced to a triangle in a similar manner. Straight lines must be drawn about the figure, so as to be partly in Fig. 50.

it and partly out, giving and taking about equal quantities, so that the figure which these lines form shall be about equivalent to the curved figure. This

Fig. 51.

having been done, the equivalent straight-lined figure is reduced by the above method.

It is sometimes more convenient not to produce one of the sides of the figure, but to draw at one end of $i t$, as at the point 1 in Fig. 51, an indefinite line, usually a perpendicular, to a line joining two distant angles of the figure, and make this line the base of the equivalent triangle desired. The operation is shown by the dotted lines in the figure. The same General Rule applies to it as to the previous figures.
77. Special Instruments. A variety of instruments have been invented for the purpose of determining areas rapidly and correctly.

One of the simplest is the "Computing Scale," which is on the same principles as the Method of Art. 71. It is represented in Fig. 52. It consists of a scale divided for its whole length from the zero-point into divisions, each representing $2 \frac{1}{2}$ chains to the scale of the plat. The scale carries a slider, which moves along it, and has a wire drawn across its center at right angles to the edges of the scale. On each side of this wire a portion of the slider, equal in length to one of the primary, or $2 \frac{1}{2}$ chain, divisions of the scale, is laid off and divided into 40 equal parts.

This instrument is used in connection with a shcet of transparent paper, ruled with parallel lines at distances apart each equal to one chain on the scale of the plat. It is plain that when the instrument is laid on this paper, with its edge on one of the parallel lines, and the slider is moved over one of the divisions of $2 \frac{1}{2}$ chains, that one rood, or a quarter of an acre, has been measured between two of the parallel lines on the paper (since 10 square chains make one acre) ; and that one of the smaller divisions measures one perch between the same parallels. Four of the larger divisions give one acre. The scale is generally made long enough to measure at once five acres.

Fig. 52.


To apply this to the plat of a field, or farm, lay the transparent paper orer it in such a position that two of the ruled lines shall touch two of the exterior points of the boundaries, as at A and B. Lay the scale, with the slide set to zero, on the paper, in a direction parallel to the ruled lines, and so that the wire of the slide cuts the left-hand oblique line so as to make the spaces $c$ and $d$ about equal. Hold the scale firm, and move the slider till the wire cuts the right-hand oblique line in such a way as to equalize the spaces $e$ and $f$. Without changing the slide, move the scale doivn the width of a space and to the left-hand end of
 the next space ; begin there again, and proceed as before.

So go on, till the whole length of the scale is run out (fire acres having been measured), and then begin at the right-hand side and work backward to the left, reading the lower divisions, which run up to 10 acres. By continuing this process, the content of plats of any size can be obtained.

A still simpler substitute for this is a scale similarly divided, but without an attached slide. In place of it there is used a piece of glass, having a line drawn across it and riveted to the end of a short scale of box-wood, dirided like the former slide. It is used like the former, except that, at starting, the zero of the short scale and not the line on the glass is made to coincide with the zero of the long scale. The slide is to be held fast to the instrument when this is mored.
78. Planimeters. These determine the area of any figure, whether bounded by straight lines or curred, by merely moring a point around the outline of the surface. This causes motion in a train of wheel-work, which registers the algebraic sum of the product of ordinates to every point in that perimeter, by the increment of their abscissas, and therefore measures the included space.

There are several varieties of these instruments. One of the best of them is Amsler's Polar Planimeter. (For descriptions and theory of Planimeters, see " Mechanical Integrators," by Henry S. H. Shaw.)
79. A purely mechanical means of determining the area of any surface by means of its weight, may be placed here. The plat is cut out of paper and weighed by a delicate balance. The weight of a rectangular piece of the same paper containing just one acre
is also found; and the "Rule of Three" gives the content. A modification of this is to paste a tracing of the plat on thin sheetlead, cut out the lead to the proper lines and weigh it.
80. Fourth Method.-Trigonometrically. By calculating, from the observed angles of the boundaries of the piece of ground, the lengths of the lines needed for calculating the content.

This method is employed for surveys made with angular instruments, as the compass, etc., in order to obtain the content of the land surveyed, without the necessity of previously making a plat, thus avoiding both that trouble and the inaccuracy of any calculations founded upon it. It is therefore the most accurate method; but will be more appropriately explained in Part I, Chapter III, under the head of "Compass Surveying."

## CHAPTER II.

## CHAIN-SURVEYING; BY THE FIRST AND SECOND METHODS : OR DIAGONAL AND PERPENDICULAR SURVEYING.

81. The chain alone is abundantly sufficient, without the aid of any other instrument, for making an accurate survey of any surface, whatever its shape or size, particularly in a district tolerably level and clear. Moreover, since a chain, or some substitute for it, formed of a rope, of leather driving-reins, etc., can be obtained by any one in the most secluded place, this method of surveying deserves more attention than has usually been given to it.

## SURVEYING BY DIAGONALS: OR BY THE FIRST METHOD.

82. Surveying by Diagonals is an application of the First Method of determining the position of a point, given in Art. 3, to which the student should again refer. Each corner of the field or farm which is to be surveyed is "determined" by measuring its distances from two other points. The field is then "platted" by repeating this process on paper, for each corner, in a contrary order, and the "content" is obtained by some of the methods explained in Chapter I.

The lines which are measured in order to determine the corners of the field are usually sides and diagonals of the irregular polygon which is to be surveyed. They therefore divide it up into triangles; whence this mode of surveying is sometimes called "Chain Triangulation."

A few examples will make the principle and practice perfectly clear. Each will be seen to require the three operations of measuring, platting, and calculating.

## A three-sided field; as Fig. 54.

Field-work. Measure the three sides, A B, B C, and C A. Measure also, as a proof-line, the distance from one of the corners, as $C$, to some point in the opposite side, as D , at which a mark should have been left, when measuring from A to B , at a known distance from A. A stick or twig, with a slit in its top, to receive a piece of paper with the distance from A marked on it, is the most convenient mark.


Platting. Choose a suitable scale. Then draw a line equal in length, on the chosen scale, to one of the sides; $A B$, for example. Take in the compasses the length of another side, as A C, to the same scale, and with one $\operatorname{leg}$ in $A$ as a center, describe an arc of a circle. Take the length of the third side, B C, and, with B as a center, describe another are, intersecting the first are in a point which will be the third corner C. Draw the lines AC and BC ; and ABC will be the plat, or miniature copy of the field surveyed.

Instead of describing two arcs to get the point C, two pairs of compasses may be conveniently used. Open them to the lengths, respectively, of the last two sides. Put one foot of each at the ends of the first side, and bring their other feet together, and their point of meeting will mark the desired third point of the triangle.

To "prove" the accuracy of the work, fix the point D, by setting off from A the proper distance, and measure the length of the line D C. If its length on the plat corresponds to its measurement on the ground, the work is correct.

It is a universal principle, in all surveying operations, that the work must be tested by some means independent of the original process, and that the same result must be arrived at by two different methods. The necessary length of this proof-line can also easily be calculated by the principles of trigonometry.

Calculation. The content of the field may now be found, either from the three sides, or more easily though not so accurately, by measuring on the plat, the length of the perpendicular CE, let fall from any angle to the opposite side, and taking half the product of these two lines.

Example 1. Fig. 54 is the plat, on a scale of two chains to one inch, of a field, of which the side A B is 200 links, B C is 100 links, and A C is 150 links. Its content, by the rule of Art. 62, is 0.726 of a square chain, or 0 A .0 R. 12 P . If the perpendicular CE be accurately measured, it will be found to be $72 \frac{1}{2}$ links. Half the product of this perpendicular by the base will be found to give the same content.

Ex. 2. The three sides of a triangular field are respectively $89 \cdot 39,54 \cdot 08$, and $45 \cdot 98$. Required its content.

Ans. 100 A. 0 R. 10 P .

## A four-sided field; as Fig. 55.

Field-work. Measure the four sides. Measure also a diagonal, as A C, thus dividing the four-sided field into two triangles. Measure also the other diagonal, or B D, for a " proof-line."

Platting. Draw a line, as A C, equal in length to the diagonal, to any
scale. On each side of it construct a triangle with the sides of the field, as directed above.

To prove the accuracy of the work, measure on the plat the length of the "proof-line," B D, and if it

Fig. 55.
 agrees with the length of the same line measured on the ground, the field-work and platting are both proved to be correct.

Calculation. Find the content of each triangle separately, as in the preceding case, and add them together; or, more briefly, multiply either diagonal (the longer one is preferable) by the sum of the two perpendiculars, and divide the product by two.

Otherwise: reduce the four-sided figure to one triangle, as in Art. 74 ; or, use any of the methods of the preceding chapter.

Ex. 3. In the field drawn in Fig. 55, on a scale of 3 chains to the inch, $\mathrm{AB}=588$ links, $\mathrm{BC}=210, \mathrm{CD}=430, \mathrm{D} \mathrm{A}=274$, the diagonal $\mathrm{A} \mathrm{C}=626$, and the proof diagonal $\mathrm{BD}=500$. The total content will be 1 A .0 R .17 P .

Ex. 4. The sides of a four-sided field are $\mathrm{AB}=12 \cdot 41, \mathrm{BC}=5 \cdot 86, \mathrm{CD}$ $=8.25, \mathrm{D} \mathrm{A}=4.24$; the diagonal $\mathrm{BD}=11.05$, and the proof-line AC $=11 \cdot 04$. Required the content. Ans. 4 A .2 R .38 P .

Ex. 5. The sides of a four-sided field are as follows: $\mathrm{AB}=8.95, \mathrm{~B} \mathrm{C}$ $=5 \cdot 33, \mathrm{CD}=10 \cdot 10, \mathrm{DA}=6.54$; the diagonal from A to C is $11 \cdot 52$; the proof diagonal from B to D is 10.92 . Required the content. Ans.

Ex. 6. In a four-sided field, $\mathrm{AB}=7 \cdot 68, \mathrm{~B} \mathrm{C}=4 \cdot 09, \mathrm{CD}=10 \cdot 64$, D A $=7 \cdot 24, \mathrm{AC}=10 \cdot 32, \mathrm{BD}=10^{\circ} 74$. Required the content. Ans.

A many-sided field, as Fig. 56.
Field-work. Measure all the sides of the field. Measure also diagonals
Fig. 56.

enough to divide the field into triangles, of which there will always be two less than the number of sides. Choose such diagonals as will divide the field into triangles as nearly equilateral as possible. Measure also one or more diagonals for "proof-lines." It is well for the surveyor himself to place stakes in advance at all the corners of the field, as he can then select the best mode of division.

Platting. Begin with any diagonal and plat one triangle. Plat a second triangle adjoining the first one. Plat another adjacent triangle, and so proceed till all have been laid down in their proper places. Measure the proof-lines as before.

Calculation. Proceed to calculate the content of the figure, precisely as directed for the four-sided field, measuring the perpendiculars and calculating the content of each triangle in turn; or taking in pairs those on opposite sides of the same diagonal ; or using some of the other methods which have been explained.
$E x .7$. The six-sided field, shown in Fig. 56, has the lengths of its lines, in chains and links, written upon them, and is divided into four triangles, by three diagonals. The diagonal BE is a "proof-line." The figure is drawn to a scale of 4 chains to the inch. The content of the field is 5 A .3 R .22 P .

Ex. 8. In a five-sided field, the lengths of the sides are as follows: A B $=2 \cdot 69, \mathrm{BC}=1 \cdot 22, \mathrm{C}=2 \cdot 32, \mathrm{DE}=3 \cdot 55, \mathrm{E} \mathrm{A}=3 \cdot 23$. The diagonals are $\mathrm{AD}=4 \cdot 81, \mathrm{BD}=3 \cdot 33$. Required its content. Ans.

A field may be divided up into triangles, not only by measuring diagonals as in the last figure, but by any of the methods shown in the four figures of Art. 67. The one which we have been employing corresponds to the last of those figures.

Still another mode may be used when the angles can not be seen from one another, or from any one point within. Take two or more convenient points within the field, and measure from them to the corners, and thus form different sets of triangles.

## Keeping the Field-Notes.

83. By Sketch. The most simple method is to make a sketch of the field, as nearly correct as the unassisted hand and eye can produce, and note down on it the lengths of all the lines, as in Fig. 56. But when many other points require to be noted, such as where fences, or roads, or streams are crossed in the measurement, or any other additional particulars, the sketch would become confused, and be likely to lead to mistakes in the subsequent platting from it. The following is therefore the usual method of keeping the field-notes. A long, narrow book is most convenient for it.
84. In Columns. Draw two parallel lines, about an inch apart from the bottom to the top of the page of the field-book, as
in the margin. This column, or pair of lines, may be conceived to represent the measured line, split in two, its two halves being then separated, an inch apart, merely for convenience, so that the distances measured along the line may be written between these halves.

Hold the book in the direction of the measurement. At the bottom of the page write down the name, or number, or letter, which represents the station at which the surrey is to begin.

A "station" is marked with a triangle or circle, as in the margin. The latter is more easily made.

In the complicated cases, which will be hereafter explained, and in which one long base-line is measured, and also many other subordinate lines, it will be mell, as a help to the memory, to mark the stations on the base-line with a triangle, and the stations on the other lines with the ordinary circle.

The station from which the measurements are made is usually put on the left of the column ; and the station which is measured to, is put on the right.

From A

But it is more compact, and aroids interfering with the notes of "offsets" (to be explained hereafter), to write the name or number of the station in B 562 the column, as in the margin.

The measurements to different points of a line are written abore one another. The numbers all refer to the beginning of the line, and are counted from it.

The end of a measured line is marked by a line dramn across the page abore the numbers which indicate the measurements which have been made.

If the chaining does not continue along the adjoining line, but the chain-men go to some other part of the field to begin another measurement, two lines are drawn across the page.

When a line has been measured, the marks $\Gamma$ or 7 are made to show whether the following line turns to the right or to the left.

A line is named, either by the names of the stations between which it is measured, as the line AB ; or by its length, a line 562 links long, being called the line 562 ; or it is recorded as Line No. 1, Line No. 2, etc. ; or as Line on page 1, 2, etc., of the field-book.

When a mark is left at any point of a line, as at D, in Fig. 49, with the intention of coming back to it again, in order to measure to some other point, the place marked is called a False Station, and is marked in the field-book "F. S." ; or has a line drawn around it, to distinguish it ; or has a station mark $\Delta$ placed outside of the column, to the right or left, according to the direction in which the measurement from it is to be made. Examples of these thres modes are given in the margin.

A false station is named by its position on the line where it belongs ; as thus-" 200 on 562."

When a gate occurs in a measured line, the distance from the beginning of the line to the side of the gate first reached is the one noted.

When the measured line crosses a fence, brook, road, etc., they are drawn on the field-notes in their true direction, as nearly as possible, but not in a continuous line across the column, as in the first figure in the margin, but as in the second figure, so that the two parts would form a continuous straight line, if the halves of the "split line " were brought together.

It is convenient to name the lines in the mar-
 gin as being Sides, Diagonals, Proof-lines, etc.
85. The field-notes of the triangular field plated in Fig. 54 are given below, according to both the methods mentioned in the preceding article.

In the field-notes in the column on the right hand, it is not absolutely necessary to repeat the B and C .


|  | From | C <br> 89 <br> 80 | on 200 |
| :---: | :---: | :---: | :---: |
| $\stackrel{\text { 圌 }}{ }$ | 7 | $\begin{gathered} \mathrm{A} \\ 1000 \\ \mathrm{C} \end{gathered}$ |  |
| $\stackrel{\text { a }}{\text { ¢ }}$ | 7 | $\begin{gathered} \mathrm{C} \\ 100 \\ \mathrm{~B} \end{gathered}$ |  |
| $\frac{\text { ù }}{\text { ì }}$ |  | B 200 $\left(\frac{80}{A}\right)$ |  |

86. The field-notes of the survey platted in Fig. 56 are giren below. They begin at the bottom of the left-hand column.


## SURVEYING BY TIE-LINES.

87. Surveying by Tie-lines is a modification of the method explained in the last chapter. It frequently happens that it is impossible to measure the diagonals of a field of many sides, in consequence of obstacles to measurements, such as woods, water, houses, etc. In such cases, "tie-lines" (so called because they tie the sides together) are employed as substitutes for diagonals.

Thus, in the four-sided field shown in the figure, the diagonals can not be measured because of woods intervening. As a substitute, measure off from any convenient corner of the field, as B , any distances, $\mathrm{BE}, \mathrm{B} \mathrm{F}$, along the sides of the field. Measure also

Fig. 57.
 the "tie-line" EF. Measure all the sides of the field as usual.

To plat this field, construct the triangle BEF, as in Art. 82. Produce the sides BE and BF , till they become respectively equal to $B A$ and $B C$, as measured on the ground. Then, with $A$ and $C$ as centers, and with radii respectively equal to AD and CD , describe arcs, whose intersection will be D , the remaining corner of the field.
88. It thus appears that one tie-line is sufficient to determine a four-sided field, two a five-sided field, and so on. But, as a check on errors, it is better to measure a tie-line for each angle, and the agreement, in the plat, of all the measurements will prove the accuracy of the whole work.

Since any inaccuracy in the length of a tie-line is increased in proportion to the greater length of the sides which it fixes, the tielines should be measured as far from the point of meeting of these sides as possible-that is, they should be as long as possible.

The radical defect of the system is that it is "working from less

Fig. 58.
 to greater" (which is the exact converse of the true principle), thus magnifying inaccuracies at every step.

A tie-line may also be employed as a " proofline," in the place of a diagonal, and tested in the same manner.

If any angle of the field is re-entering, as at B in the figure, measure a tie-line across the salient angle A B C.
89. Chain-Angles. It is convenient, though not necessary, to measure equal distances along the sides : B E, B F, in Fig. 5\%, and B A, B C, in Fig. 58. "Chain-angles" are thus formed. To reduce "chain-angles" to degrees and minutes, see Art. 28.
90. Inaccessible Areas. The method of tie-lines can be applied to measuring fields which can not be entered.

Thus, in the figure, ABCD is an inaccessible wooded field, of

## Fig. 59.

 four sides. To survey it, measure all the sides, and at any corner, as D , measure any distance DE , in the line of AD produced. Measure also another distance D F in the line of CD produced. Measure the tie-line EF, and the figure can be platted as in the case of the field of Fig. 5\% the sides of the triangle being produced in the contrary direction.

The same end would be attained by prolonging only one side, as shown at the angle $A$ of the same figure, and measuring $A G, A H$, and GH. It is better, in both cases, to tie all the angles in a similar manner.

This method may be applied to a figure of any number of sides by prolonging as many of them as are necessary; all of them, if possible.
91. If the sides CD and AD were prolonged by their full length, the content of the figure could be calculated without any plat ; for the new triangle DEF would equal the triangle DAC , and the sides of the triangle A C B would then be known.

This principle may be extended still further. For a five-sided field, as in Fig. 60, produce two pairs of sides, a distance equal to

their length, forming two new triangles, as shown by the dotted lines, and measure the sides $\mathrm{B}^{\prime} \mathrm{D}^{\prime}$ and $\mathrm{A}^{\prime} \mathrm{D}^{\prime \prime}$. The three sides of each of these triangles will thus be known, and also the three sides of the triangle BAD , since $A D=A^{\prime} D^{\prime \prime}$, and $B D=B^{\prime} D^{\prime}$.

The method of this article may be employed for a figure of six sides, as shown in Fig. 61 (in which the dotted lines within the wooded field have their lengths deter-

Fig. 61.
 mined by the triangles formed outside of it), but not for figures of a greater number of sides.

## SURVEYING BY PERPENDICULARS: OR BY THE SECOND METHOD.

92. The method of Surveying by Perpendiculars is founded on the Second Method of determining the position of a point, explained in Art. 4. It is applied in two ways, either to making a complete survey by "Diagonals and Perpendiculars," or to measuring a crooked boundary by "Offsets." Each will be considered in turn.

The best method of getting perpendiculars on the ground must, however, be first explained.
Fig. 62.


## To set out Perpendiculars.

93. Surveyor's Cross. The simplest instrument for this purpose is the Surveyor's Cross, or Cross-Staff, shown in the figure. It consists of a block of wood, of any shape, having in it two saw-cuts, made very precisely at right angles to each other, about half an inch deep, and with center-bit holes made at the bottom of the cuts to assist in finding the objects. This block is fixed on a pointed staff, on which it can turn freely, and which should be precisely 8 links ( $63 \frac{1}{3}$ inches) long, for the convenience of short measurements.

To use the cross-staff to erect a perpendicular, set it at the point of the line at which a perpendicular is wanted. Turn its head till, on looking through one
sar-cut, you see the ends of the line. Then will the other sawcut point out the direction of the perpendicular, and thus guide the measurement desired.

To find where a perpendicular to the line, from some object, as a corner of a field, a tree, etc., would meet the line, set up the cross-staff at a point of the line which seems to the eye to be about the spot. Note about how far from the object the perpendicular at this point strikes, and move the cross-staff that distance ; and repeat the operation till the correct spot is found.
94. To test the accuracy of the instrument, sight through one slit to some point A, and place a stake B
 in the line of sight of the other slit. Then turn its head a quarter of the way around, so that the second slit, looked through, points to A. Then see if the other slit covers B again, as it will if correct. If it does not do so, but sights to some other point, as $\mathrm{B}^{\prime}$, the apparent error is double the real one, for it now points as far to the right of the true point C as it did before to its left.

This is the first example we have had of the invaluable principle of Reversion, which is used in almost every test of the accuracy of surveying and astronomical instruments, its peculiar merit being that it doubles the real error, and thus makes it twice as easy to perceive and correct it.

The instrument, in its most finished form, is made of a hollow brass cylinder, which has two pairs of slits exactly opposite to each other, one of each pair being narrow and the other wide, with a horse-hair stretched from the top to the bottom of the latter. It is also, sometimes, made with eight faces, and two more pairs of slits added, so as to set off half a right angle.

Another form is a hollow brass sphere, as in the figure. This enables the surveyor to set off perpendiculars on very steep slopes.

Another form of the surreyor's cross consists of

two pairs of plain "Sights," each shaped as in the figure, placed at the ends of two bars at right angles to each other. The slit, and the opening with a hair stretched from its top to its bottom, are respectively at the top of one sight and at the bottom of the opposite sight.* This is used in the same manner as the preced-
 ing form, but is less portable, and more liable to get out of order.

A temporary substitute for these instruments may Fra. 66. be made by sticking four pins into the corners of a
 square piece of board, and sighting across them, in the direction of the line and at right angles to it.
95. Optical Square. The most convenient and accurate instrument is, however, the Optical Square. The figures give a perspective view of it, and also a plan with the lid removed. It is a small circular box, containing a strip of looking-glass, from the upper half of which the silvering is removed. This glass is placed so as to make precisely half a right angle with the line of sight, which passes through a slit on one side of the box, and a vertical hair stretched across the opening on the other side, or a mark on the glass. The box is held in the hand over the spot where the perpendicular is desired (a plumb-line in the hand will give perfect accuracy), and the observer applies his eye to the slit A, looking through the upper or unsilvered part of the glass, and
 turns the box till he sees the other end of the line B, through the opening C. The assistant, with a rod, moves along in the direction where the perpendicular is desired, being seen in the silvered

[^9]parts of the glass, by reflection through the opening D , till his rod, at E , is seen to coincide with, or to be exactly under, the object B . Then is the line DE at right angles to the line AB , by the optical principle of the equality of the angles of incidence and reflection.

To find where a perpendicular from a distant object would strike the line, walk along the line, with the instrument to the eye, till the image of the object is seen, in the silvered part of the glass, to coincide with the direction of the line seen through the unsilvered part.

The instrument may be tested by sighting along the perpendicular, and fixing a point in the original line, on the principle of "reversion."

The surveyor can make it for himself, fastening the glass in the box by four angular pieces oi cork, and adjusting it by cutting away the cork on one side, and introducing wedges on the other side. The box should be blackened inside.

Another form of the optical square contains two glasses, fixed at an angle of $45^{\circ}$, and giving a right angle on the principle of the sextant.

Perpendiculars may be set out with the chain alone, by a rariety of methods. These methods generally consist in performing on the ground, the operations executed on paper in practical geometry, the chain being used, in the place of the compasses, to describe the necessary arcs.

As these operations, however, are less often used for the method of surveying now to be explained, than for overcoming obstacles to measurement, it will be more convenient to consider them in that connection.

## Diagonals and Perpendiculars.

96. We have seen in the preceding pages that plats of surrevs, made with the chain alone, have their contents most easily determined by measuring, on the plat, the perpendiculars of each of the triangles, into which the diagonals measured on the ground have divided the field. In the Method of Surveying by Diagonals and Perpendiculars, now to be explained, the perpendiculars are measured on the ground. The content of the field can, therefore,
be found at once (by adding together the half products of each perpendicular by the diagonal on which it is let fall), without the necessity of previously making a plat, or of measuring the sides of the field. This is, therefore, the most rapid and easy method of surveying when the content alone is required, and is particularly applicable to the measurement of the ground occupied by crops, for the purpose of determining the number of bushels grown to the acre, the amount to be paid for mowing by the acre, etc.

A Three-sided Field. Measure the longest side, as $\Lambda \mathrm{B}$, and the perpendicular, CD, let fall on it from the opposite angle C. Then the content is equal to half the product of the side by the perpendicular. If obstacles prevent this, find the point, where a perpendicular let fall from an angle, as A , to the opposite side produced, as B C, would meet it, as at E in the figure. Then half the product of $\mathrm{A} E$ by CB is the content

Fig. 68.
 of the triangle.

A Four-sided Field. Measure the diagonal A C. Leave marks at the points on this diagonal at which perpendiculars from B and from D would meet it, finding these points by trial,
 as previously directed. The best marks at these "false stations" have been described. Return to these false stations and measure the perpendiculars. When these perpendiculars are measured before finishing the measurement of the diagonal, great care is necessary to avoid making mistakes in the length of the diagonal, when the chainmen return to continue its measurement. One check is to leave at the mark as many pins as have been taken up by the hind-chainman in coming to that point from the beginning of the line.

Ex. 9. Required the content of the field of Fig. 69. Ans. 0 A. 2 R. 29 P.
The field may be platted from these measurements, if desired, but with more liability to inaccuracy than in the first method, in which the sides are measured. The plat of the figure is three chains to one inch.

The field-notes may be taken by writing the measurements on a sketch, as in the figure; or, in more complicated cases, by the column method, as below. A new symbol may be employed, this mark, $\vdash$, or -1 , to show the false station, from which a perpendicular is to be measured.

| $\begin{aligned} & \text { 芯 } \\ & \text { 岂 } \end{aligned}$ | From 200 on 480 | $\begin{gathered} 110 \\ \text { F.S. } \end{gathered}$ | $\stackrel{\text { to } B}{-1}$ |
| :---: | :---: | :---: | :---: |
| $\begin{aligned} & \text { O} \\ & \underset{\sim}{\sim} \end{aligned}$ | From 280 on 480 | $\begin{aligned} & 175 \\ & \text { F.S. } \end{aligned}$ | $\begin{aligned} & \text { to } \mathrm{D} \\ & \vdash \end{aligned}$ |
|  | From ${ }^{-1}$ | $\begin{gathered} 480 \\ 280 \\ 200 \\ \odot \end{gathered}$ | $\stackrel{t o}{\vdash}$ |

## Ex. 10. Calculation.

$$
\begin{array}{r}
\text { sq. lks. } \\
\mathrm{ABC}=\frac{1}{2} \times 480 \times 110=26400 \\
\mathrm{ADC}=\frac{1}{2} \times 480 \times 175=\frac{42000}{\text { sq. chains }} \mathbf{6 . 8 4 0 0} \\
\text { Acres } 0.684
\end{array}
$$

It is still easier to take the tro triangles together; multiplying the diagonal by the sum of the perpendiculars and dividing by two.

A Many-sided Field. Fig. 70 and the accompanying field-notes represent the field which was surreyed by the first method and platted in Fig. 56.

|  |  |  |
| :--- | :---: | :---: | :---: |
| From 5.07 on 7.37 | F.S. |  |
| From 1.60 on 7.75 | 2.53 | to D |

Ex. 11. Calculation. The content of the triangles may be expressed thus:
sq. 7ks.

$$
\begin{array}{r}
\mathrm{ABC}=\frac{1}{2} \times 1142 \times 26 \uparrow=152457 \\
\mathrm{AEC}=\frac{1}{2} \times 1142 \times 493=281503 \\
\mathrm{CDE}=\frac{1}{2} \times 775 \times 253=98037 \\
\mathrm{AEF}=\frac{1}{2} \times 737 \times 154=\frac{56749}{} \\
\text { 8q. chains } 58.8746 \\
\text { Acres } 5.88746
\end{array}
$$

or, $๊$ A. 3 R. 22 P.
The first two triangles might hare been taken together, as in the prerious field.

Content calculated from the perpendiculars will generally rary slightly from that obtained by measuring on the plat.

A small field which has many sides may sometimes be conreniently surreyed by taking one diagonal and measuring the perpendiculars let fall on it from each angle of the field, and thus dividing the whole area into triangles and trapezoids, as in Fig. 41.

The line on which the perpendiculars are to be let fall may also be outside of the field, as in Fig. 42.

Such a surrey can be platted very readily, but the length of the perpendiculars renders the plat less accurate.

This procedure supplies a transition to the method of "otfsets," which is explained in the next article.

Fig. 70.


## Offsets.

97. Offsets are short perpendiculars, measured from a straight line, to the angles of a crooked or zigzag line near which the
straight line runs. Thus, in the figure, let A C D B be a crooked fence, bounding one side of a field. Chain along the straight line AB,

Fig. 71.
 which runs from one end of the fence to the other, and, when opposite each corner, note the distance from the beginning, or the point A, and also measure and note the perpendicular distance of each corner C and D from the line. These corners will then be "determined" by the Second Method, Art. 4.

The field-notes, corresponding to Fig. 71, are as in the margin. The measurements along the line are written in the column, as before, counting from the beginning of the line, and the offsets are written beside it, on the right or left, opposite the distance at which they are taken. A sketch of the crooked line is also usually made in the field-notes, though not absolutely necessary in so simple a case as this. The letters C and D would not be
used in practice, but are here inserted to show the connection between the field-notes and the plat.

In taking the field-notes, the widths of the offsets should not be drawn proportionally to the distances between them, but the breadths should be greatly exaggerated in proportion to the lengths.

A more extended example, with a little different notation, is given below. In the figure, which is on a scale of eight chains to one inch for the distances along the line, the breadths of the offsets are exaggerated to four times their true proportional dimensions.


The plat and field-notes of the position of two houses, determined by offsets, are given above on a scale of two chains to one inch :

Double offsets are sometimes convenient; and sometimes triple and quadruple ones. Below are given the notes and the plat, one chain to one inch, of a road of rarying width, both sides of which are determined by double offsets. It will be seen that the line A B crosses one side of the road at 160 links from A, and the other side of it at 220 .

Two methods of keeping the field-notes are given. In the first form, the offsets to each side of the road are given separately and connected by the sign + . In the second form, the total distance of the second offset is giren, and the two measurements connected by the word "to." This is easier both for measuring and platting.


|  | $\begin{gathered} \mathrm{B} \\ 260 \end{gathered}$ | $30+60$ | B 260 | 30 to 90 |
| :---: | :---: | :---: | :---: | :---: |
|  | 240 | $10+70$ | 240 | 10 to 80 |
| 0 | 220 | 50 0 | 220 | 50 |
| 20 | 200 | $30 \quad 20$ | 200 | 30 |
| 40 | 180 | 1040 | 180 | 10 |
| 45 | 160 | $0 \quad 45$ | 160 | 0 |
| $50+0$ | 140 | 50 to 0 | 140 |  |
| $55+5$ | 120 | 60 to 5 | 120 |  |
| $50+20$ | 100 | 70 to 20 | 100 |  |
| $45+15$ | 80 | 60 to 15 | 80 |  |
| $50+10$ | 60 | 60 to 10 | 60 |  |
| $50+20$ | 40 | 70 to 20 | 40 |  |
| $55+20$ | 20 | 75 to 20 | 20 |  |
| $60+0$ | A | 60 to 0 | A |  |

Offsets may generally be taken with sufficient accuracy by measuring them as nearly at right angles to the base-line as the eye can estimate. The surveyor should stand by the chain, facing the fence, at the place which he thinks opposite to the corner to which he wishes to take an offset, and measure "square" to it by the eye, which a little practice will enable him to do with much correctness.

The offsets may be measured, if short, with an Offset-staff, a light stick, 10 or 15 links in length, and divided accordingly ; or, if they are long, with a tape. They are generally but a few links in length. A chain's length should be the extreme limit, as laid down by the English "Tithe Commissioners," and that should be employed only in exceptional cases. When the "cross-staff" is in use, its divided length of 8 links renders the offset-staff needless.

When offsets are to be taken, the method of chaining to the end of a line (described in Art. 18) is somewhat modified. After the leader arrives at the end of the line, he should draw on the chain
till the follower, with the back end of the chain, reaches the last pin set. This facilitates the counting of the links to the places at which the offsets are taken.

The offsets are to be taken to every angle of the fence or other crooked line; that is, to every point where it changes its direction. These angles or prominent bends can be best found by one of the party walking along the crooked fence and directing another at the chain what points to measure opposite to. If the line which is to be thus determined is curved, the offsets should be taken to points so near each other that the portions of the curved line lying between them may, without much error, be regarded as straight. It will be most convenient, for the subsequent calculations, to take the offsets at equal distances apart along the straight line from which they are measured.

In the case of a crooked brook, such as is shown in the figure given below, offsets should be taken to the most prominent angles, such as are marked $a a a$ in the figure, and the intermediate bends may be merely sketched by the eye.

Fig. 75.


When offsets from lines measured around a field are taken inside of these bounding lines, they are sometimes distinguished as insets.
98. Platting. The most rapid method of platting the offsets is by the use of a Platting Scale (described in Art. 47) and an O.ffset Scate, which is a short scale divided on its edges like a platting scale, but having its zero in the middle, as in the figure.

The platting scale is placed parallel to the line, with its zeropoint opposite to the beginning of the line. The offset scale is slid along the platting scale, till its edge comes to a distance on the latter at which an offset had been taken, the length of which is marked off with a needle-point from the offset scale. This is then
slid on to the next distance, and the operation is repeated. If one person reads off the field-notes, and another plats, the operation

Fig. 76.

will be greatly facilitated. The points thus obtained are joined by straight lines, and a miniature copy of the curved line is thus obtained ; all the operations of the platting being merely repetitions of the measurements made on the ground.

If no offset scale is at hand, make one of a strip of thick draw-ing-paper, or pasteboard ; or use the platting scale itself, turned crossways, having previously marked off from it the points from which the offsets had been taken.

In plats made on a small scale, the shorter offsets are best estimated by the eye.

On the ordnance survey of Ireland, the platting of offsets was facilitated by the use of a combination of the offset scale and the platting scale, the former being made to slide in a groove in the latter, at right angles to it.
99. Calculating Content. When the crooked line determined by offsets is the boundary of a field, the content, inclosed between it and the straight line surveyed, must be determined, that it may be added to, or subtracted from, the content of the field bounded by the straight lines. There are various methods of effecting this.

The area inclosed between the straight and the crooked lines is divided up by the offsets into triangles and trapezoids, the content of which may be calculated separately and then added together. The content of the plat on page 65 will, therefore, be $1500+$ $4125+625=6250$ square links $=0.625$ square chain. The con-
tent of the plat on page 66 will in like manner be found to be, on the left of the straight line, 30,000 square links, and on its right, 5,000 square links.
100. When the off sets have been taken at equal distances, the content may be more easily obtained by adding together half of the first and of the last offset, and all the intermediate ones, and multiplying the sum by one of the equal distances between the offsets. This rule is merely an abbreviation of the preceding one.

Thus, in the plat of page 66, the distances being equal, the content of the offsets on the left of the straight line will be $120 \times 250$ $=30,000$ square links, and on the right, $20 \times 250=5,000$ square links; the same results as before.

When the line determined by the offsets is a curved line, "Simpson's rule" gives the content more accurately. To employ it, an even number of equal distances must have been measured in the part to be calculated. Then add together the first and last offset, four times the sum of the even offisets (i. e., the 2d, 4 th, 6 th, etc.), and twice the sum of the odd offsets (i. e., the $3 \mathrm{~d}, 5$ th, \% th , etc.), not including the first and the last. Multiply the sum by one of the equal distances between the offsets, and divide by 3 . The quotient will be the area.

Ex. 12. The offsets from a straight line to a curred fence were $8,9,11,15,16,14,9$, links, at equal distances of 5 links. What was the content included between the curved fence and the straight line?

Ans. $3 \approx 1 \cdot 666$.
101. Equalizing, or giving and taking, is an approximate mode of calculation much used by practical surveyors. A crooked line,

Fig. 77.

determined by offsets, having been platted, a straight line is dramn on the plat, across the crooked line, leaving as much space outside
of the straight line as inside of it, as nearly as can be estimated by the eye, "equalizing" it, or "giving and taking" equal portions. The straight line is best determined by laying across the irregular outline the straight edge of a piece of transparent horn, or tracingpaper, or glass, or a fine thread or horse-hair stretched straight by a light bow of whalebone. In practical hands, this method is sufficiently accurate in most cases. The student will do well to try it on figures, the content of which he has previously ascertained by perfectly accurate methods.

## SURVEYiNg by the Preceding methods Combined.

102. All the methods which have been explained in the preceding sections-ṣurveying by Diagonals, by Tie-lines, and by Perpendiculars, particularly in the form of offsets-are frequently required in the same survey. The method by Diagonals should be the leading one ; in some parts of the survey obstacles to the measurements of diagonal may require the use of Tie-lines; and, if the fences are crooked, straight lines are to be measured near them, and their crooks determined by Offsets.

Offsets are necessary additions to almost every other method of surveying. In the smallest field surveyed by diagonals, unless all the fences are perfectly straight lines, their bends must be determined by offsets. The plat (scale of one chain to one inch) and field-notes of such a case are given below. A sufficient number of the sides, diagonals, and proof-lines, to prove the work, should be platted before platting the offsets.

Fig. 78.



Ex. 13. Required the content of the above field. Ans.

103. Field-Books. The difficulty and the importance of keeping the field-notes clearly and distinctly increase with each new combination of methods. For this reason, three different methods of keeping the field-notes of the same surrey will now be given (from Bourn's "Surveying"), and a careful comparison by the student of the corresponding portions of each will be rery profitable to him :

Field-Book No. 1.
Fig. 79,


Field-Book No. 1 (Fig. 79) shows the Sketch method, explained in Art. 83.

Field-Book No. 2.
Fig. 80.


Field-Book No. 2 (Fig. 80) shows the Column method, explained in Art. 84.

Field-Book No. 3.
Fig. 81.


Field-Book No. 3 (Fig. 81) is a convenient combination of the two preceding methods. The bottom of the book is at the side of this figure, at A.
104. Inaccessible Areas. A combination of offsets and tie-lines supplies an easy method of surveying an inaccessible area, such as a pond, swamp, forest, block of houses, etc., as appears from the figure, in which external bounding lines are taken at will and measured, and tied by "tie-lines" measured between these lines, prolonged when necessary, while offsets from them determine the irregularities of the actual boundaries of the pond, etc.

These offsets are insets, and their content is, of course, to be subtracted from the content of the principal figure.

Even a circular field might thus be approximately measured from the outside.

If the shape of the field admits of it, it will be preferable to measure four lines about the field in such directions as to inclose it in a rectan-

Fig. 83.
 gle, and to measure offsets from the sides of this to the angles of the field.

## OBSTACLES TO MEASUREMENT IN CHAIN-SURVEYING.

105. In the practice of the rarious methods of surreying which have been explained, the hills and valleys which are to be crossed, the sheets of water which are to be passed over, the woods and houses which are to be gone through-all these form obstacles to the measurement of the necessary lines which are to join certain points, or to be prolonged in the same direction. Many special precautions and contrivances are therefore rendered necessary ; and the best methods to be employed, when the chain alone is to be used, will now be given.

These methods for overcoming the rarious obstacles met with in practice constitute a Land-Geonetry. Its problems are per-
formed on the ground instead of on paper ; its compasses are a chain fixed at one end and free to swing around with the other ; its scale is the chain itself; and its ruler is the same chain stretched tight. Its advantages are that its single instrument (or a substitute for it, such as a tape, a rope, etc.) can be found anywhere ; and its only auxiliaries are equally easy to obtain, being a few straight and slender rods, and a plumb-line, for which a pebble suspended by a thread is a sufficient substitute.

Many of these problems require the employment of perpendicular and parallel lines. For this reason we will commence with this class of problems.

The demonstrations of most of these problems will be left as exercises for the student.

The elegant "Theory of Transversals" (Appendix B) will be an important element in some of these demonstrations.

## Problems on Perpendiculars.*

Problem 1. To erect a perpendicular at any point of a line.
106. First Method. Let A be the point at which a perpendicular to the line is to be set out. Measure off equal distances $\mathrm{AB}, \mathrm{AC}$, on each side of the point. Take a portion of the chain not quite $1 \frac{1}{2}$ time as long as AB or A C, fix one end of this at B, and describe an are with the other end. Do the same from C. The intersection of these ares will
 fix a point D . AD will be the perpendicular required. Repeat the operation on the other side of the line.

[^10]If that is impossible, repeat it on the same side with a different length of chain.
107. Second Method. Measure off as before, equal distances AB, A C, but each about only one third of the chain. Fasten
Fig. 85.
 the ends of the chain with two pins at $B$ and $C$. Stretch it out on one side of the line and put a pin at the middle of it, D. Do the same on the other side of the line, and set a pin at E. Then is DE a perpendicular to BC . If it is impossible to perform the operation on both sides of the line, repeat it on the same side with a different length of chain, as shown by the lines $B F$ and $C F$ in the figure, so as to get a second point.
108. Other Methods. All the methods to be given for the nest problem may be applied to this.

Problem 2. To erect a perpendicular to a line at a given point, when the point is at or near the end of the line.
109. First Method. Measure 40 links along the line. Let one assistant hold one end of the chain at that point; let a second hold the 20 link mark which is nearest the other end, at the given "point A, and let a third take the $50-\mathrm{link}$ mark, and tighten the chain, drawing equally on both portions of it. Then will the 50 link mark be in the perpendicu-
 lar desired. Repeat the operation on the other side of the line so as to test the work.

The above numbers are the most easily remembered, but the longer the lines measured the better ; and nearly the whole chain may be used ; thus: Fix down the 36th link from one end at A, and the 4 th link from the same end on the line at B. Fix the other end of the chain also at B. Take the 40 th-link mark from this last end, and draw the chain tight, and this mark will be in the perpendicular desired. The sides of the triangle formed by the chain will be 24,32 , and 40 .
110. Otherwise : using a 50 -feet tape, hold the 16 -feet mark at A; hold the 48 -feet mark and the ringend of the tape together on the line; take the 28 -feet mark of the tape, and draw it tight; then will the 28 -feet mark be in the perpendicular desired.

Fig. 87.

111. Second Method. Hold one end of the chain at A and fix

Fig. 88.
 the other end at a point B, taken at will. Swing the chain around $B$ as a center, till it again meets the line at $C$. Then carry the same end around (the other end remaining at $B$ ) till it comes in the line of $C B$ at $D$. AD is the perpendicular required.

Problem 3. To erect a perpendicular to an inaccessible line, at a given point of it.
112. First Method. Get points in the direction of the inaccessible line prolonged, and from them set out a parallel to the line, by methods which are given in Art. 121, etc. Find by trial the point in which a perpendicular to this second line (and therefore to the first line) will pass through the required point.

Problem 4. To let fall a perpendicular from a given point to a given line.
113. First Method. Let P be the given point, and AB the given line. Measure some distance, a chain or less, from $C$ to $P$, and then fix one end of the chain at $P$, and swing it around till the same distance meets the line at some point $D$. The middle point $E$ of the distance $C D$ will be the required point, at which
 the perpendicular from P would meet the line.
114. Second Method. Stretch a chain, or a portion of it,
 from the given point $P$, to some point, as $A$, of the given line. Hold the end of the distance at $A$, and swing round the other end of the chain from $P$, so as to set off the same distance along the given line from A to some point B. Measure B P. Then will the distance $B C$ from $B$ to the foot of the desired perpendicular $=\frac{B P^{2}}{2 \mathrm{AB}}$.

Problem 5. To let fall a perpendicular to a line from a point nearly opposite to the end of the line.
115. First Method. Stretch a chain from the giren point $P$, to some point, as $A$, of the given line. Fix to the ground the middle point $B$ of the chain AP, and swing around the end which was at P , or at $A$, till it meets the given line in a point $C$, which will be the foot of the required

Fig. 91.
 perpendicular.

Fig. 92.


Fig. 93.

116. Second Method. At any conrenient point, as $A$ of the giren line, erect a perpendicular of any convenient length, as A B, and mark a point $C$ on the giren line in the line of $P$ and $B$. Measure $C A, C B$, and $C$ P. Then the distance from $C$ to the foot of the perpendicular, i. e., $\mathrm{CD}=\frac{\mathrm{CA} \times \mathrm{C} \mathrm{P}}{\mathrm{CB}}$.

Problem 6. To let fall a perpendicular to a line from an inaccessible point.
117. First Method. Let P be the given point. At any point $A$, on the giren line, set out a perpendicular, $A B$, of any conrenient length. Prolong it on the other side of
the line the same distance. Mark on the given line a point D in the line of PB , and a point E in the line of PC. Mark the point F at the intersection of D C and BE prolonged. The line F P is the line required, being perpendicular to the given line at the point G.
118. Second Method. Let A and B be two points of the given line. From A let fall a perpendicular, AC , to the visual line, B P ; and from B let fall a perpendicular, B D , to the visual line, A P.

Fig. 94.
 Find the point at which these perpendiculars intersect, as at E , and the line PE , prolonged to F , will give the perpendicular required.

Problem 7. To let fall a perpendicular from a given point to an inaccessible line.
119. First Method. Let P be the given point, and AB the

Fig. 95.
 given line. By the preceding problem let fall perpendiculars from A to B P at C ; and from B to AP at D; the line PE , passing from the given point to the intersection of these perpendiculars, is the desired perpendicular to the inaccessible line AB .
This method will apply when only two points of the line are visible.

The proof of 118 and 119 is found in the "Theory of Transversals," Corollary 3.
120. Second Method. Through the given point set out a line parallel to the inaccessible line. At the given point erect a perpendicular to the parallel line, and it will be the required perpendicular to the inaccessible line.

## Problems on Parallels.

Problem 1. To run a line from a given point parallel to a given line.
121. First Method. Let fall a perpendicular from the point to the line. At another point of the line, as far off as possible, erect a perpendicular equal in length to the one just let fall. The line joining the end of this line to the given point will be the parallel required.
${ }^{\text {Fig. }} 96$.

122. Second Method. Measure from P to any point, as C of the given line, and put a mark at the middle point $D$ of that line. From any point, as E of the given line, measure a line to the point D , and continue it till $\mathrm{DF}=\mathrm{DE}$. Then will the line P F be parallel to A B.
123. Third Method. From any point, as C of the line, set off equal distances along the line to D and E. Take a point F , in the line of PD. Stake out the lines F C and FE , and also the line EP, crossing the line CF in the point G. Lastly , prolong the line DG till it meets the line EF in the point H. PH is
 the parallel required.

The proof is found in Corollary 4 of "Transrersals."
Problem 2. To run a line from a given point parallel to an inaccessible line.
124. First Method. Let A B be the given line, and P the given point. Set a stake at $C$, in the line of PA, and another at any convenient point, D. Through P set out, by the preceding problem, a parallel to D A, and set a stake at the point, as E , where this parallel intersects D C prolonged. Through E set out a parallel to B D, and set a stake at the point F , where this parallel intersects B C prolonged. PF is the parallel required.
125. Second Method. Set a stake at any point $C$ in the line of $A P$, and another at any convenient place, as at D . Through P set out a parallel to AD, intersecting C D in E. Through E set out a parallel to DB , intersecting CB in F . The line P F will be the parallel required.

Fig. 99.

126. Alinement and Measurement. We are now prepared, having secured a variety of methods for setting out Perpendiculars and Parallels in every probable case, to take up the general subject of overcoming Obstacles to Measurement.

Before a line can be measured its direction must be determined. This operation is called Ranging the line, or Alining it, or Boning it.* The word alinement $\dagger$ will be found very convenient for expressing the direction of a line on the ground, whether between two points or in their direction prolonged.

This branch of our subject naturally divides itself into two parts, the first of which is preliminary to the second, viz. :
I. Of Obstacles to Alinement; or how to establish the direction of a line in any situation.
II. Of Obstacles to Measurement; or how to find the length of a line which can not be actually measured.

## 1. Obstacles to Alinement.

127. All the cases which can occur under this head may be reduced to two, viz. :
A. To find points in a line beyond the given points, i. e., to prolong the line.
B. To find points in a line between two given points of it, i. e., to interpolate points in the line.

## A. To Prolong a Line.

128. By ranging with Rods. When two points in a line are given, and it is desired to prolong the line by ranging it out with

[^11]rods, three persons are required, each furnished with a straight, slender rod, and with a plumb-line, or other means of keeping their rods vertical. One holds his rod at one of the given points, $A$ in the figure, and another at B. A third, C, goes forward as far as he can without losing sight of the first two rods, and then, looking back, puts himself "in line" with A and B-i. e., so that when his eye is placed at $C$ the rod at $B$ hides or covers the rod at $A$. This he can do most accurately by holding a plumb-line before his eye, so that it shall cover the first two rods. The lower end of the plumb-bob will then indicate the point where the third rod should be placed, and so with the rest. The first man, at A, is then signaled and comes forward, passes both the others, and puts himself at D, "in line" with C and B. The man at B then goes on to E, and " lines" himself with D and C ; and so they proceed, in this "hand-over-hand" operation, as far as is desired. Stakes are driven at each point in the line as soon as it is determined.
129. The rods should be perfectly straight, either cylindrical or polygonal, and as slender as they can be without bending. They should be painted in alternate bands of red and white, each a foot or link in length. Their lower ends should be pointed with iron, and a projecting bolt of iron will enable them to be pressed down by the foot into the earth, so that they can stand alone. When this is done, one man can range out a line. A rod can be set perfectly vertical by holding a plumb-line before the eye at some distance from the rod, and adjusting the rod so that the plumbline covers it from top to bottom, and then repeating the operation in a direction at right angles to the former. A stone dropped from top to bottom of the rods will approximately attain the same end.

When the lines to be ranged are long, and great accuracy is required, the rods may have attached to them plates of tin with openings cut out of them, and black horse-hairs stretched from top to bottom of the openings.


A small telescope must then be used for ranging these hairs in line. In a hasty survey, straight twigs, with their tops split to receive a paper folded as in the figure, may be used.
130. By Perpendiculars. The straight line, A B in the figure, is supposed to be stopped by a tree, a house, or other obstacle, and it is desired to prolong the line beyond this obstacle. From any two points, as $A$ and $B$ of the line, set off (by some of

Fig. 102.
 the methods which have been given) equal perpendiculars, AC and BD , long enough to pass the obstacle. Prolong this line beyond the obstacle, and from any two points in it, as E and F , measure the perpendiculars EG and FH equal to the first two, but in a contrary direction. Then will G and H be two points in the line AB prolonged which can be continued by the method of the last article. The points A and B should be taken as far apart as possible, as should also the points E and F . Three or more perpendiculars on each side of the obstacle may be set off, in order to increase the accuracy of the operation. The same thing may also be done on the other side of the line, as another confirmation or test of the accuracy of the prolonged line.
131. By Equilateral Triangles. The obstacles noticed in the last article may also be overcome by means of three equilateral trian-

Fig. 103.
 gles formed by the chain. Fix one end of the chain, and also the end of the first link from its other end, at B; fix the end of the 33d link at A; take hold of the 66th link and draw the chain tight, pulling equally on each part, and put a pin at the point thus found, C in the figure. An equilateral triangle will thus be formed, each side being 33 links. Prolong the line AC past the obstacle to some
point, as D. Make another equilateral triangle, DEF, as before, and thus fix the point F . Prolong DF to a length equal to that of $A D$, and thus fix a point, $G$. At $G$ form a third equilateral triangle, GHK, and thus fix a point, K. Then will K G give the direction of AB prolonged.
132. By Symmetrical Triangles. Let A B be the line to be pro-

Fig. 104.
 longed. Take any convenient point, as C. Kange out the line, $\mathrm{A} C$, to a point $\mathrm{A}^{\prime}$, such that $\mathrm{CA}^{\prime}=\mathrm{CA}$. Range out $C B$, so that $C B^{\prime}=C B$. Range backward $\mathrm{A}^{\prime} \mathrm{B}^{\prime}$ to some point D, such that D C prolonged will pass the obstacle. Find, by ranging, the inter-
section at E of $\mathrm{D} B$ and AC . From $C$ measure, on $\mathrm{CA}^{\prime}$, the distance $\mathrm{C} \mathrm{E}^{\prime}=\mathrm{CE}$. Then range out DC and $\mathrm{B}^{\prime} \mathrm{E}^{\prime}$ to their intersection in $P$, which will be a required point in the direction of $A B$ prolonged. The symmetrical points are marked by corresponding letters. Several other points should be obtained in the same manner.

In this, as in all similar operations, very acute intersections should be avoided as far as possible.
133. By Transversals. Let AB be the given line. Take any two points $C$ and $D$, such that the line $C D$ will pass the obstacle. Take another point, $E$, in the intersection of $C \mathrm{~A}$ and DB . Measure $\mathrm{A} E, \mathrm{AC}, \mathrm{CD}, \mathrm{BD}$, and BE . Then the distance from D to P , a point in the required prolongation, will be $D P=\frac{C D \times B D \times A E}{B E \times A C-B D \times A E}$.

Other points in the prolongation may be obtained in the same manner, by merely moving the single point C in

the line of EA ; in which case the new distances, CA and CD, will alone require to be measured.

If $A E$ be made equal to $A C$, then is $D P=\frac{C D \times B D}{B E-B D}$.
If $B E$ be made equal to $B D$, then is $D P=\frac{C D \times A E}{A C-A E}$
The minus sign in the denominators must be understood as only meaning that the difference of the two terms is to be taken, without regard to which is the greater.
134. By Harmonic Conjugates. Let A B be the given line. Set a stake at any point C. Set stakes at points D , on the line CA , and at E , on the line CB ; these points, D and E , being so chosen that the line DE will pass beyond the obstacle. Set a fourth stake, F, at the intersection of the lines AE and DB. Set a fifth stake, G, anywhere in the line C F ; a sixth stake, H , at the intersection of C B and D G prolonged ; and a seventh, K, at the intersection of CA and EG prolonged. Finally, range out the lines DE and KH, and their intersection at P will be in the line A B prolonged.
135. By the Complete Quadrilateral. Let A B be the given line. Take any convenient point
 - C ; measure from it to B , and onward, in the same line prolonged, an equal distance to D. Take any other convenient point, E , such that CE and DE produced will clear the obstacle. Measure from E to A , and onward, an equal distance, to F. Range out the lines FC and DE to their intersection in $G$.

Range out FD and CE to intersect in H. Measure G H. Its middle point, P , is the required point in the line of A B prolonged. The unavoidable acute intersections in this construction are objectionable.

## B. To interpolate Points in a Line.

136. The most distant given point of the line must be made

Fig. 108.
Front View.


Back Viero.
 as conspicuous as possible by any efficient means, such as placing there a staff bearing a flag : red and white, if seen against woods or other dark background ; and red and green, if seen against the sky.

A convenient
and portable signal is shown in the figure.
The figure represents a disk of tin about six inches in diameter, painted white and hinged in the middle, to make it more portable. It is kept open by the bar, B, being turned into the catch, C. A screw, S , holds the disk in a slit in the top of the pole.

Another contrivance is a strip of tin, which has its ends bent horizontally in contrary directions. As the wind will take strongest hold of the side which is concare toward it, the bent strip will continually revolve, and thus be very conspicuous. Its upper half should be painted red, and its lower half white.

A bright tin cone set on the staff can be seen at a great distance when the sun is shining.
137. Ranging to a point thus made conspicuous is rery simple when the ground is level. The surreyor places his eye at the nearest end of the line, or stands a little behind a rod placed on it, and by signs moves an assistant, holding a rod at some point as nearly
in the desired line as he can guess, to the right or left, till his rod appears to cover the distant point.
138. Across a Valley. When a valley or low spot intervenes between the two ends of the line, A and Z in the figure, a rod held in the low place, as at $B$, would seldom be high enough to be seen from A , to cover the distant rod at Z. In such a

Fig. 109.
 case, the surveyor at A should hold up a plumb-line over the point, at arm's length, and place his eye so that the plumb-line covers the rod at Z. He should then direct the rod held at $B$ to be moved till it, toa, is covered by the plumb-line. The point B is then said to be "in line" between A and Z. In geometrical language, B has now been placed in the vertical plane determined by the vertical plumbline and the point Z. Any number of intermediate points can thus be "interpolated," or placed in line between A and Z.
139. Over a Fill. When a hill rises between two points and prevents one being seen from the other, as in the figure (the upper

Fig. 110.
 part of which shows the hill in "elevation," and the lower part in " plan"), two observers, B and C , each holding a rod, may place themselves on the ridge, in the line between the two points, as nearly as they can guess, and so that each can at once see the other and the point
beyond him. B looks to Z, and by signals puts C "in line." $C$ then looks to $A$, and puts $B$ in line at $B^{\prime}$. B repeats his operation from $B^{\prime}$, putting $C$ at $C^{\prime}$, and is then himself mored to $B^{\prime \prime}$, and so they alternately "line" each other, continually approximating to the straight line between A and Z , till they at last find themselves both exactly in it, at $\mathrm{B}^{\prime \prime \prime}$ and $\mathrm{C}^{\prime \prime \prime}$.
140. A single person may put himself in line between two points, on the same principle, by laying a straight stick on some support, going to each end of it in turn, and making it point successively to each end of the line. The "Surveyor's Cross," Art. 93 , is convenient for this purpose, when set up between the two given points and moved again and again, until, by repeated trials, one of its slits sights to the given points when looked through in either direction.
141. On Water. A simple instrument for the same object is represented in the figure. AB and CD

Fig. 111.
 are two tubes, about $1 \frac{1}{2}$ inch in diameter, connected by a smaller tube, EF. A piece of looking-glass, G H, is placed in the lower part of the tube $A B$, and another, K L , in the tube CD . The planes of the two mirrors are at right angles to each other. The eye is placed at $A$, and the tube $A B$ is directed to any distant object, as X , and any other object behind the observer, as Z, will be seen, apparently under the first object in the mirror G H, by reflection from the mirror K L, when the observer has succeeded in getting in line between the tro objects. M N are screws by which the mirror K L may be adjusted. The distance between the two tubes will cause a small parallax, which will, however, be insensible except when the two objects are near together.
142. Through a Wood. When a wood intervenes between any two given points, preventing one from being seen from the other, as in the figure, in which $\mathbf{A}$ and Z are the given points, proceed thus: Hold a rod at some point $\mathrm{B}^{\prime}$ as nearly in the desired line from A as can

Fig. 112.
 be guessed at, and as far from A as possible. To approximate to the proper direction, an assistant may be sent to the other end of the line, and his shouts will indicate the direction; or a gun may be fired there; or, if rery distant, a rocket may be sent up after dark. Then range out the "random line" A B', by the method given in Art. 128, noting also the distance from $A$ to each point found, till you arrive at a point $Z^{\prime}$, opposite to the point Z-i. e., at that point of the line from which a perpendicular there erected would strike the point Z. Measure Z' Z. Then move each of the stakes, perpendicularly from the line $A Z^{\prime}$, a distance proportional to their distances from A. Thus, if $A Z^{\prime}$ be 1,000 links, and $Z^{\prime} Z$ be 10 links, then a stake $B^{\prime}$, 200 links from $A$, should be moved 2 links to a point $B$, which will be in the desired straight line $A Z$; if $\mathrm{C}^{\prime}$ be 400 links from A , it should be moved 4 links to C , and so with the rest. The line should then be cleared, and the accuracy of the position of these stakes tested by ranging from A to Z .
143. To an Invisible Intersection. Let $A B$ and $C D$ be two lines, which, if pro-

Fig. 113.
 given points, A, B, C, D, P. Set a sixth stake at E, in the 7
alinements of $A D$ and $C P$; and a seventh stake at $F$, in the alinements of BC and A P. Then set an eighth stake at G, in the alinements of BE and DF . PG will be the required line. This is an application of the "Theory of Transversals."

Otherwise: Through P range out a parallel to the line B D. Note the points where this parallel meets AB and CD, and call these points $Q$ and $R$. Then the distance from $B$, on the line $B D$, to a point which shall be in the required line running from P to the invisible point, will be $=\frac{B D \times Q P}{Q R}$.

## II. Obstacles to Measurement.

144. The cases in which the direct measurement of a line is prevented by various obstacles may be reduced to three:
A. When both ends of the line are accessible.
B. When one end of it is inaccessible.
C. When both ends of it are inaccessible.
A. When Both Ends of the Line are accessible.
145. By Perpendiculars. On reaching the obstacle, as at A in the figure, set off a perpendicular,

Fig. 114.
 AB ; turn a second right angle at B , and measure past the obstacle ; turn a third right angle at C , and measure to the original line at D . Then will the measured distance, BC , be equal to the desired distance, AD .

If the direction of the line is also unknown, it will be most easily obtained by the additional perpendiculars shown in Fig. 102 of Art. 130.
146. By Equilateral Triangles. The method given in Art. 131 for determining the direction of a line through an obstacle will also gise its length ; for

Fig. 115.
 in Fig. 115 the desired distance A G is equal to the measured distances AD or DG .
147. By Symmetrical Triangles. Let AB be the distance required. Measure from A obliquely to some point C past the obstacle. Measure onward, in the same line, till CD is as long as A C. Place stakes at C and D. From B measure to C , and from C measure onward, in the same line, till CE is equal to CB. Measure E D, and
 it will be equal to $\mathrm{A} B$, the distance required. If more convenient, make CD and CE equal, respectively, to half of AC and C B ; then will AB be equal to twice D E.
148. By Transversals. Let A B be the required distance. Set

Fig. 117.
 a stake, C , in the line prolonged; set another stake, D, so that C and B can be seen from it ; and a third stake, E , in the line of BD prolonged, and at a distance from D equal to the distance from D to B. Set a fourth stake, F, at the intersection of EA and CD .
Measure AC, AF, and FE. Then is $A B=\frac{A C}{A F}(F E-A F)$.
B. When One End of the Line is inaccessible.
149. By Perpendiculars. This principle may be applied in a variety of ways. In Fig. 118 let AB be the required distance. At the point A set off AC perpendicular to AB , and of any convenient length. At C set off a perpendicular to C B. and continue it to a point, D , in the line of A and B. Measure D A. Then is A B $=\frac{\mathrm{A} \mathrm{C}^{2}}{\mathrm{AD} \text {. }}$

Fig. 118.

150. Otherwise: At the point A, in Fig. 119, set off a perpendicular, A C. At C set off another perpendicu-

Fig. 119.

lar, CD. Find a point, E, in the line of $A C$ and $B D$. Measure $A E$ and EC. Then is $A B=\frac{A E \times C D}{C E}$.

If EC be made equal to AE , and D be set in the line of $B E$, and also in the perpendicular from C , then will CD be equal to AB.

$$
\text { If } \mathrm{EC}=\frac{1}{2} \mathrm{AE} \text {, then } \mathrm{CD}=\frac{1}{2} \mathrm{AB} \text {. }
$$

151. Otherwise: At A, in Fig. 120, measure a perpendicular, $A C$, to the line $A B$; and at any point, as $D$ in this line, set off a perpendicular to $\mathrm{D} B$, and continue it to a point E , in the line of CB. Measure DE and also D A.
Then is $A B=\frac{A C \times A D}{D E-A C}$.


Fig. 121.

152. By Parallels. From A measure AC in any conrenient direction. From a point D , in the line of BC , measure $a$ line parallel to $C A$, to a point $E$ in the line of AB . Measure also A E.

Then is $A B=\frac{A C \times A E}{D E-A C}$.
153. By a Parallelogram. Set a stake, C , in the line of A and B , and set another stake, D , wherever conrenient. With a distance equal to CD , describe from A an arc on the ground; and, with a distance equal to AC , describe another are from D intersecting the first are in E . Or, take A C and CD so that together they make one chain ; fix the ends of the chain at A and D ; take hold of the chain at such a link that one part of it equals AC and the other CD, and draw
 it tight to fix the point E . Set a stake at F in the intersection
of AE and D B. Measure AF and EF. Then is AB= $\mathrm{AC}_{-\mathrm{AF}}^{\mathrm{EF}}$; or $\mathrm{CB}=\frac{\mathrm{AC} \times \mathrm{CD}}{\mathrm{EF}}$.
154. By Symmetrical Triangles. Let AB be the required distance. From A measure a line in any convenient direction, as A C, and measure onward, in the same direction, till $\mathrm{CD}=\mathrm{AC}$. Take any point E in the line of A and B . Measure from E to C , and onward in the same line, till CF $=C E$. Then find by trial a point G, which shall be at the same time in the line of C and B , and in the line of D and F. Measure the distance from G to D , and it will be equal to the required distance from A to B . If more convenient, make $\mathrm{CD}=$ $\frac{1}{2} \mathrm{AC}$, and $\mathrm{CF}=\frac{1}{2} \mathrm{CE}$, as shown by the finely dotted lines in the figure. Then will $\mathrm{D} G=\frac{1}{2} \mathrm{AB}$.
155. Otherwise: Prolong BA to some point C. Range out any convenient line $\mathrm{CA}^{\prime}$, and measure

Fig. 124.


Fig. 123.
 $\mathrm{CA}^{\prime}=\mathrm{CA}$. The triangle $\mathrm{CAA}^{\prime} \mathrm{B}$ is now to be reproduced in a symmetrical triangle situated on the accessible ground. For this object take, on A C, some point D and measure $\mathrm{CD}^{\prime}=$ $C D$. Find the point $E$ at the intersection of $\mathrm{AD}^{\prime}$ and $\mathrm{A}^{\prime} \mathrm{D}$. Find the point $F$ at the intersection of $A^{\prime} B$ and CE. Lastly, find the point $B^{\prime}$ at the intersection of $A F$ and $C A^{\prime}$. Then will $\mathrm{A}^{\prime} \mathrm{B}^{\prime}=\mathrm{A} B$. The symmetrical points have corresponding letters affixed to them.
156. By Transversals. Set a stake, C , in the alinement of BA ; a second, D , at any convenient point ; a third, E , in the line CD ; and a fourth, F , at the intersection of the aline-

Fig. 125.

ments of D A and EB. Measure A C, $\mathrm{CE}, \mathrm{ED}, \mathrm{DF}$, and FA. Then is $\mathrm{AB}=$ AC $\times \mathrm{AF} \times \mathrm{DE}$ $\overline{\mathrm{CE} \times D F-A F \times D E}$

If the point E be taken in the middle of $C D$ (as it is in the figure), then $A B=$ $\frac{A C \times A F}{D F-A F}$.
If the point F be taken in the middle of AD , then $\mathrm{AB}=$ AC $\times D E$ $\overline{\mathrm{CE}} \mathrm{E} \mathrm{DE}$.

The minus signs must be interpreted as in Art. 121.
15\%. By Harmonic Division. Set stakes, C and D, on each side of $A$, and so that the three are in the same straight line. Set a third stake at any point, $E$, of the line A B. Set a fourth, $F$, at the intersection of $C B$ and $D E$; and a fifth, $G$, at the inter. section of D B and CE. Set a sixth stake, $H$, at the intersection of AB and FG. Measure AE and EH. Then is $\mathrm{AB}=\frac{\mathrm{AE} \times \mathrm{AH}}{\mathrm{AE}-\mathrm{EH}}$.

Fig. 126.

158. To an Inaccessible Line. The shortest distance, C D, from

Fig. 127.
 a giren point, C , to an inaccessible straight line A B , is required. From C let fall a perpendicular to $\mathrm{A} B$, by the method of Art 119. Then set a stake at any point, E , on the line AC ; set a second, F , at the intersection of $E B$ and $C D$; a third, $G$, at the intersection of AF and CB; and a fourth, H , at the intersection of EG and CD .
Measure CH and HF. Then is $C D=\frac{C H \times C F}{C H-H F}$; or CD
$=C H \cdot \frac{C H+H F}{C H-H F}$; or $C D=\frac{C H \times C F}{2 C H-C F}$.
159. To an Inaccessible Intersection. When two lines (as $A B$, CD, in the figure) meet in a river, a building, or any other inaccessible point, the distance from any point of either to their intersection, D E, for example, may be found thus : From any point B, on one line, measure BD , and continue it

Fig. 128.
 till $D F=D B$. From any other point $G$ of the former line measure $G D$, and continue the line till $D H=G D$. Continue $H F$ to meet $D C$ in some point K. Measure K D. K D will be equal to the desired distance D E. $B E$ can be found by measuring FK , which is equal to it.
If DF and DH be made respectively equal to one half or one third, etc., of $D B$ and $D G$, then will $K D$ and $K F$ be respectively equal to one half or one third, etc., of DE and BE .

## C. When Both Ends of the Line are inaccessible.

160. By Similar Triangles. Let AB be the inaccessible distance. Set a stake at any convenient point

Fig. 129.
 C , and find the distances CA and CB by any of the methods just given. Set a second stake at any point, D , on the line C A . Measure a distance equal to $\frac{\mathrm{CB} \times \mathrm{CD}}{\mathrm{CA}}$, from $C$, on the line $C B$, to some point E . Measure D E. Then is $A B=\frac{A C \times D E}{C D}$.
If more convenient, measure $\mathrm{C} D$ in the contrary direction from the river, as in Fig. 130, instead of toward it, and in other respects proceed as before.
161. By Parallels. Let $\mathrm{A} B$ be the inaccessible distance. From any point, as C, range out a parallel to A B, as in Art. 124, etc. Find the distance C A by Art. 149 , etc. Set a stake at the point E , the

Fig. 130.


Fig. 131.

intersection of CA and DB , and measure $C E$. Then is $A B=\frac{C D \times(A C-C E)}{C E}$.
162. By a Parallelogram. Set a stake at any convenient point C. Set stakes D and E anywhere in the alinements CA and CB . With D as a center, and a length of the chain equal to CE, describe an arc ; and with E as a center, and a length of the chain equal to CD , describe another arc, intersecting the former one at F. A parallelogram, CDEF, will thus be formed. Set stakes at G and H , where the aline-
 ments DB and EA intersect the sides of this parallelogram. Measure C D, D F, GF, F H, and H G. The inaccessible distance $A B=\frac{C D \times D F \times G H}{F G \times F H}$.

If $\mathrm{CD}=\mathrm{CE}$, then $\mathrm{AB}=\frac{\mathrm{CD}^{2} \times \mathrm{GH}}{\mathrm{FG} \times \mathrm{FH}}$.
163. By Symmetrical Triangles. Take any conrenient point, as

FIG. 133.
 C. Set stakes at two other points, D and $\mathrm{D}^{\prime}$, in the same line, and at equal distances from C. Take a point E, in the line of AD ; measure from it to C , and onward till $\mathrm{CE} \mathrm{E}^{\prime}=\mathrm{CE}$. Take a point F in the line of BD ; measure from it to $C$, and onward till $C F^{\prime}=C F$. Range out the lines AC and $\mathrm{E}^{\prime} \mathrm{D}^{\prime}$, and set a stake at their intersection, $\mathrm{A}^{\prime}$. Range out the lines $B C$ and $F^{\prime} \mathrm{D}^{\prime}$; and set a stake at their intersection, $\mathrm{B}^{\prime}$. Measure $\mathrm{A}^{\prime} \mathrm{B}^{\prime}$. It will be equal to the desired distance AB.
164. Otherwise: Take any convenient point, as C, and set off equal distances on each side of it, in the line of CA, to D and $\mathrm{D}^{\prime}$. Set off the same distances from $C$, in the line of $C B$, to $E$ and $E^{\prime}$. Through C set out a parallel to DE or $\mathrm{D}^{\prime} \mathrm{E}^{\prime}$, and set stakes at the points F and $\mathrm{F}^{\prime}$ where this parallel intersects $A \mathrm{E}^{\prime}$ and $\mathrm{B} \mathrm{D}^{\prime}$. Range out the lines $A D^{\prime}$ and EF', and set a stake

Fig. 134.
 at their intersection $\mathrm{A}^{\prime}$. Range out the lines $B E^{\prime}$ and $D F$, and set a stake at their intersection $B^{\prime}$. Measure $A^{\prime} B^{\prime}$, and it will be equal to the desired distance A B.

## CHAPTER III.

## COMPASS-SURVEYING; OR BY THE THIRD METHOD.

165. Angular Surveying determines the relative positions of points, and therefore of lines, on the Third Principle, as explained in Art. 5.

Either the compass or the transit may be employed in angular surveying.
166. Surveying with the compass is a less direct operation than surveying with the transit. But as the use of the compass is much more rapid and eass, for this reason, as well as for its smaller cost, it is the instrument most commonly employed in land-surveying in spite of its imperfections and inaccuracies.

The method of Polar Surveying (or surrering by the third method) embraces two minor methods. The most usual one consists in going around the field with the instrument, setting it at each corner, and measuring there the angle which each side makes with its neighbor, as well as the length of each side. This method is called by the French the method of Cheminement. It has no special name in English, but may be called (from the American verb, to progress) the Method of Progression. The other system, the Method of Radiation, consists in setting the instrument at one point and thence measuring the direction and distance of each corner of the field or other object. The corresponding name of what we have called triangular surrering is the Method of Intersections, since it determines points by the intersections of straight lines.

16\%. When the two lines which form an angle lie in the same horizontal or level plain, the angle is called a horizontal angle.*

When these lines lie in a plane perpendicular to the former, the angle is called a vertical angle.

When one of the lines is horizontal, and the other line from tho eye of the observer passes above the former, and in the same vertical plane, the angle is called an angle of elevation.

When the latter line passes below the horizontal line,


## THE COMPASS.

168. The Needle. The most essential part of the compass is the magnetic needle. It is a slender bar of steel, usually five or six inches long, strongly magnetized, and balanced on a pivot, so that it may turn freely, and thus be enabled to continue pointing in the same direction (that of the " magnetic meridian," approximately north and south) however much the "compass-box," to which the pivot is attached, may be turned around.

As it is important that the needle should move with the least

[^12]possible friction, the pivot should be of the hardest steel ground to a very sharp point; and in the center of the needle, which is to rest on the pivot, should be inserted a cap of agate, or other hard material. Iridium for the pivot, and ruby for the cap, are still better.

If the needle be balanced on its pivot before being magnetized, one end will sink, or "dip," after the needle is magnetized. To bring it to a level, several coils of wire are wound around the needle so that they can be slid along it, to adjust the weight of its two ends and balance it more perfectly.

The north end of the needle is usually cut into a more ornamental form than the south end for the sake of distinction.

The principal requisites of a compass-needle are intensity of directive force and susceptibility. Beyond a certain limit, say fire inches, no additional power is gained by increasing the length of the needle. On the contrary, longer ones are apt to have their strength diminished by several consecutire poles being formed. Short needles, made very hard, are therefore to be preferred.

The needle should not come to rest very quickly. If it does, it indicates either that it is weakly magnetized, or that the friction on the pivot is great. Its sensitiveness is indicated by the number of vibrations which it makes in a small space before coming to rest.

A screw, with a milled head, on the under side of the plate which supports the pirot, is used to raise the needle off this pirot when the instrument is carried about, to prevent the point being dulled by unnecessary friction.
169. The Sights. Next after the needle, which gives the direction of the fixed line whose angles with the lines to be surveyed are to be measured, should be noticed the sights, which show the directions of these last lines. At each end of a line passing through the pivot is placed a "sight," consisting of an upright bar of brass, with openings in it of various forms-usually slits, with a circular aperture at their top and bottom; all these arrangements being intended to enable the line of sight to be directed to any desired object with precision.

A telescope which can move up and down in a vertical plane, i. e., a plunging telescope, or one which can turn completely over, is sometimes substituted for the sights. It has the great advantage of giving more distinct vision at long distances, and of admitting of sights up and down very steep slopes. Its accuracy of vision is, however, rendered nugatory by the want of precision in the readings of the needle. If a telescope be applied to the compass, a graduated circle with vernier should be added, thus converting the compass into a "transit."
170. The Divided Circle. We now have the means of indicating the directions of the two lines whose angle is to be measured. The number of degrees contained in it is to be read from a circle divided into degrees, in the center of which is fixed the pivot bearing the needle. The graduations are usually made to half a degree, and a quarter of a degree or less can then be "estimated." The pivot and needle are sunk in a circular box, so that its top may be on a level with the needle. The graduations are usually made on the top of the surrounding rim of the box, but should also be continued down its inside circumference so that it may be easier to see with what division the ends of the needle coincide.

The degrees are not numbered consecutively from $0^{\circ}$ around to $360^{\circ}$, but run from $0^{\circ}$ to $90^{\circ}$, both ways from the two diametrically opposite points at which a line, passing through the slits in the middle of the sights, would meet the divided circle.

The lettering of the surveyor's compass has one important difference from that of the mariner's compass.

When we stand facing the north, the east is on our right hand, and the west on our left. The graduated card of the mariner's compass, which is fastened to the needle and turns with it, is marked accordingly. But, in the surveyor's compass, one of the 0 points being marked N. or north (or indicated by a fleur-de-lis), and the opposite one S . or south, the 90 -degrees-point on the right of this line, as you stand at the S. end and look toward the N., is marked W. or west ; and the left hand 90 -degrees-point is marked E. or east. The reason of this will be seen when the method of using the compass comes to be explained.
171. The Points. In ordinary land-surresing only four points of the compass hare names, viz., north, south, east, and west;
 the direction of a line being described by the angle which it makes with a north and south line to its east or to its west. But, for nautical purposes, the circle of the compass is divided in to thirty-two points, the names of which are shown in the figure. Two rules embrace all the cases : 1. When the letters indicating two points are joined together, the point half-way between the tro meant; thus, N. E. is half-way between north and east; and N. N. E. is half-way between north and northeast. 2. When the letters of two points are joined together with the intermediate word $b y$, it indicates the point which comes next after the first in going toward the second ; thus, N. by E. is the point which follows north in going toward the east ; S. E. by S. is the next point from southeast going toward the south.
172. Eccentricity. The center-pin, or pirot of the needle, ought to be exactly in the center of the graduated circle; the needle ought to be straight, and the line of the sights ought to pass exactly through this center and through the 0 points of the circle. If this is not the case, there will be an error in every observation. This is called the error of eccentricity.

When the maker of a compass is about to fix the pirot in place, he is in doubt of two things: whether the needle is perfectly straight, and whether the pirot is exactly in the center. In Figs. $13 \%$ and 138 both of these are represented as being excessirely in error.

First, to examine if the needle be straight. Fix the pirot temporarily so that the ends of the needle may cut opposite de-grees-i. e., degrees differing by $180^{\circ}$. The condition of things at
this stage of progress will be represented by Fig. 13\%. Then turn the compass-box half-way around. The error will now be doubled,

Fig. 137.


Fig. 138.


Fig. 139.

as is shown by Fig. 138, in which the former position of the needle is indicated by a dotted line.* Now bend the needle, as in Fig. 139 , till it cuts divisions midway betwen those cut by it in its present and in its former position. This makes it certain that the needle is straight, or that its two ends and its center lie in the same straight line.

Second, to put the pivot in the center. Move it till the straightened needle cuts opposite divisions. It is then certain that the direction of the needle passes through the center. Turn the compass-box one quarter around, and, if the needle does not then cut opposite divisions, move the pivot till it does. Repeat the operation in various positions of the box. It will be a sufficient test if it cuts the opposite divisions of $0^{\circ}, 45^{\circ}$, and $90^{\circ}$.

To fix the sights precisely in line, draw a hair through their slits and move them till the hair passes over the 0 points on the circle.

The surveyor can also examine for himself, by the principle of reversion, whether the line of the sights passes through the center or not. Sight to any very near object. Read off the number of degrees indicated by one end of the needle. Then turn the compass half around, and sight to the same object. If the two readings do not agree, there is an error of eccentricity, and the arithmetical mean, or half sum of the two readings, is the correct one.

In Fig. 140 the line of sight A B is represented as passing to

[^13]one side of the center, and the needle as pointing to $46^{\circ}$. In Fig. 141 the compass is supposed to have been turned half around, and

Fig. 140.


Fig. 141.

the other end of the sights to be directed to the same object. Suppose that the needle would hare pointed to $45^{\circ}$ if the line of sight had passed through the center; the needle will now point to $44^{\circ}$, the error being doubled by the reversion, and the true reading being the mean.

This does not, however, make it certain that the line of the sights passes through the 0 points, which can only be tested by the hair, as mentioned abore.
173. Levels. On the compass-plate are tro small spirit-levels. They consist of glass tubes slightly curved upward, and nearly filled with alcohol, leaving a bubble of air within them. They are so adjusted that, when the bubbles are in the centers of the tubes, the plate of the compass shall be level. One of them lies in the direction of the sights, and the other at right angles to this direction.

On the compass-plate, and between the vernier and the lefthand sight in the figure, is the Outkeeper, for keeping tally of the chains in any distance.
174. Tangent Scale. This is a conrenient, though not essential, addition to the compass, for the purpose of measuring the slopes of ground, so that the proper allorance in chaining may be made. In the figure of the compass may be seen, on the edge of
the left-hand sight, a small projection of brass with a hole through it. On the edge of the other sight are engraved lines numbered from $0^{\circ}$ to $20^{\circ}$, the $0^{\circ}$ being of the same height above the compassplate that the eye-hole is. To use this, set the compass at the bottom of a slope, and at the top set a signal of exactly the height of the eye-hole from the ground. Level the compass very carefully, particularly by the level which lies lengthwise, and, with the eye at the eye-hole, look to the signal and note the number of the division on the farther sight which is cut by the visual ray. That will be the angle of the slope ; the distances of the engraved lines from the $0^{\circ}$ line being tangents (for the radius equal to the distance between the sights) of the angles corresponding to the numbers of the lines.
175. Vernier. The compass-box is connected with the plate which carries it and the sights, so that it can turn around on this plate. This motion is given to it by a slow motion or tangent screw, shown on the left of the compass-box in the figure. The space through which the compass-box is moved is indicated by a vernier. For description of a vernier, and method of reading it, see subject Verniers under Transit-Surveying.
176. Tripod. The compass, like Fig. 142.

Fig. 143. most surveying instruments, is usually supported on a tripod, consisting of three legs, shod with iron, and so connected at top as to be movable in any direction. There are many forms of these. Lightness and stiffness are the qualities desired. The most usual form is shown in the figures of the transit and the level. Of the two represented in Figs. 142 and 143 the first has the advantage of being very easily and cheaply made; and the second that of being light and yet capable of very firmly resisting horizontal torsion.


The joints by which the instrument is connected with the tripod are also various. Fig. 144 is the "ball-and-socket joint," most usual in this country. It takes its name from the ball in which terminates the covered spindle which enters a corresponding cavity under the compass-plate and the socket in which this ball turns. It admits of motion in any direction, and can be tightened or loosened by turning the upper half of the hollow piece inclosing

it, which is screwed on the lower half. Fig. 145 is called the "shell-joint." In it the two shell-shaped pieces inclosing the ball are tightened by a thumb-screw. Fig. 146 is "Cugnot's joint." It consists of two cylinders placed at right angles to each other, and through the axes of which pass bolts, which turn freely in the cylinder, and can be tightened or loosened by thumb-screws at their ends. The combination of the two motions which this joint permits enables the instrument which it carries to be placed in any desired direction. This joint is much the most stable of the three.
177. Jacob's Staff. A single leg, called a "Jacob's staff," has some advantages, as it is lighter to carry in the field, and can be made of any wood on the spot where it is to be used, thus saring the expense of a tripod and the trouble of its transportation. Its upper end is fitted into the lower end of a brass head which has a ball-and-socket joint and axis above. Its lower end should be shod
with iron, and a spike running through it is useful for pressing it into the ground with the foot. Of course, it can not be conveniently used on frozen ground or on parements. It may, however, be set before or behind the spot at which the angle is to be measured, provided that it is placed very precisely in the line of direction from that station to the one to which a sight is to be taken.
178. The Prismatic Compass. The peculiarity of this instrument (often called Schmalcalder's) is that a glass triangular prism is substituted for one of the sights. Such a prism has this peculiar property that at the same time it can be seen through, so that a sight can be taken through it, and that its upper surface reflects like a mirror, so that the numbers of the degrees immediately under it can be read off at the same time that a sight to any object is taken. Another peculiarity necessary for profiting by the last one is that the divided circle is not fixed, but is a card fastened to the needle and moving around with it, as in the mariner's compass. The minute description which follows is condensed from Simms.

In the figure, A represents the compass-box and B the card, which, being attached to the magnetic needle, mores as $i t$ moves

Fig. 147.
 around the agate center $a$, on which it is suspended. The circumference of the card is usually divided to $\frac{1}{4}$ or $\frac{1}{2}$ of a degree. © is a prism which the observer looks through. The perpendicular thread of the sight-vane, $E$, and the divisions on the card appear together on looking through the prism, and the division with which the thread coincides when the needle is at rest, is the "bearing" of whatever object the thread may bisect-i. e., is the angle which the line of sight makes with the direction of the needle. The
prism is mounted with a hinge-joint, $D$. The sight-vane has a fine thread stretched along its opening in the direction of its length, which is brought to bisect any object by turning the box around horizontally. F is a mirror made to slide on or off the sight-rane, E ; and it may be reversed at pleasure-that is, turned face downward; it can also be inclined at any angle by means of its joint, $d$; and it will remain stationary on any part of the rane by the friction of its slides. Its use is to reflect the image of an object to the eye of an observer when the object is much above or below the horizontal plane. The colored glasses represented at $G$ are intended for observing the sun. At $e$ is shown a spring, which, being pressed by the finger at the time of observation and then released, checks the vibrations of the card, and brings it more speedily to rest. A stop is likewise fixed to the other side of the box, by which the needle may be thrown off its center.

The method of using this instrument is very simple: First raise the prism in its socket, $b$, until you obtain a distinct riew of the divisions on the card. Then, standing over the point where the angles are to be taken, hold the instrument to the eye, and, looking through the slit, $C$, turn around till the thread in the sight-rane bisects one of the objects whose bearing is required; then, by touching the spring, $e$, bring the needle to rest, and the division on the card which coincides with the thread on the rane will be the bearing of the object from the north or south points of the magnetic meridian. Then turn to any other object and repeat the operation; the difference between the bearing of this object
 and that of the former will be the angular distance of the objects in question. Thus, suppose the former bearing to be $40^{\circ} 30^{\prime}$, and the latter $10^{\circ} 15^{\prime}$, both east or both west, from the north or south, the angle will be $30^{\circ} 15^{\prime}$. The divisions are generally numbered $5^{\circ}, 10^{\circ}, 15^{\circ}$, etc., around the circle to $360^{\circ}$.

The figures on the compass-
card are reversed or written upside down, as in the figure (in which only every fifteenth degree is marked), because they are again reversed by the prism.

The prismatic compass is generally held in the hand, the bearing being caught, as it were, in passing ; but more accurate readings would, of course, be obtained if it rested on a support, such as a stake cut flat on its top.

In the former mode, the needle never comes completely to rest, particularly in the wind. In such cases, observe the extreme divisions between which the needle vibrates, and take their arithmetical mean.
179. Defects of the Compass. The compass is deficient in both precision and correctness.*

The former defect arises from the indefiniteness of its mode of indicating the part of the circle to which it points. The point of the needle has considerable thickness; it can not quite touch the divided circle ; and these divisions are made only to whole or half degrees, though a fraction of a division may be estimated or guessed at. The vernier does not much better this, as we shall see when explaining its use. Now, an inaccuracy of one quarter of a degree in an angle-i. e., in the difference of the directions of two lines-causes them to separate from each other $5 \frac{1}{4}$ inches at the end of 100 feet; at the end of 1,000 feet, nearly $4 \frac{1}{2}$ feet; and, at the end of a mile, 23 feet. A difference of only one tenth of a degree, or six minutes, would proance a difference of $1 \frac{3}{4}$ foot at the end of 1,000 feet; and $9 \frac{1}{4}$ feet at the distance of a mile. Such are the differences which may result from the want of precision in the indications of the compass.

But a more serious defect is the want of correctness in the compass. Its not pointing exactly to the true north does not, indeed, affect the correctness of the angles measured by it. But it does not point in the same or in a parallel direction during even the

[^14]same day, but changes its direction between sunrise and noon nearly a quarter of a degree, as will be fully explained hereafter. The effect of such a difference we have just seen. This direction may also be greatly altered in a moment, without the knowledge of the surveyor, by a piece of iron being brought near to the compass, or by some other local attraction, as will be noticed in Art. 186. This is the weak point in the compass.

Notwithstanding these defects, the compass is a very valuable instrument, from its simplicity, rapidity, and convenience in use ; and, though never precise, and seldom correct, it is generally not very wrong.

## THE FIELD-WORK.

180. Taking Bearings. The "bearing" of a line is the angle which it makes with the direction of the needle. The bearing and length of a line are named collectively the Course.

To take the bearing of any line, set the compass exactly over any point of it by a plumb-line suspended from beneath the center of the compass, or, approximately, by dropping a stone. Level the compass by bringing the air-bubbles to the middle of the level tubes. Direct the sights to a rod held truly rertical or "plumb" at another point of the line, the more distant the better. The tro ends are usually taken. Sight to the lowest risible point of the rod. When the needle comes to rest, note what division on the circle it points to ; taking the one indicated by the north end of the needle, if the north point on the circle is farthest from you, and vice versa.

In reading the division to which one end of the needle points, the eye should be placed over the other end, to avoid the error which might result from the "parallax," or apparent change of place of the end read from, when looked at obliquely.

The bearing is read and recorded by noting between what letters the end of the needle comes, and to what number ; naming, or writing down, first, that letter, N. or S., which is at the $0^{\circ}$ point nearest to that end of the needle from which you are reading; second, the number of degrees to which it points; and, third, the letter E . or W . of the $90^{\circ}$ point which is nearest to the
same end of the needle. Thus, in the figure, if when the sights were directed along a line (the north point of the compass being most distant from the observer) the north end of the needle was at the point $A$, the bearing of the line sighted on would be north $45^{\circ}$ east ; if the end of the needle was at $B$, the bearing would be east; if at
 C, S. $30^{\circ}$ E. ; if at D, south ; if at $\mathrm{E}, \mathrm{S} .60^{\circ} \mathrm{W}$. ; if at F , west ; if at $\mathrm{G}, \mathrm{N} .60^{\circ} \mathrm{W}$. ; if at H , north.
181. We can now understand why W. is on the right hand of the compass-box and E. on the left. Let the direction from the center of the compass to the point

Fig. 150.
 $B$ in the figure be required, and suppose the sights in the first place to be pointing in the direction of the needle, S. N., and the north sight to be ahead. When the sights (and the circle to which they are fastened) have been turned so as to point in the direction of $B$, the point of the circle marked E. will have come round to the north end of the needle (since the needle remains immovable), and the reading will therefore be "east," as it should be. The effect on the reading is the same as if the needle had moved to the left the same distance which the sights have moved to the right, and the left side is therefore properly marked "east," and vice versa. So, too, if the bearing of the line to C be desired half-way between north and east-i. e., N. $45^{\circ}$ E.; when the sights and the circle have turned $45^{\circ}$ to the right, the needle, really standing still, has apparently arrived at the point half-way between N . and E., i. e., N. $45^{\circ} \mathrm{E}$.

Some surveyors' compasses are marked the reverse of this, the E. on the right and the W. on the left. These letters must
then be reversed in the mind before the bearing is noted down.
182. Reading with Vernier. When the needle does not point precisely to one of the division-marks on the circle, the fractional part of the smallest space is usually estimated by the eye, as has been explained. But this fractional part may be measured by the vernier as follows : Suppose the needle to point lotween N. $31^{\circ} \mathrm{E}$. and N. $31 \frac{1}{2}^{\circ}$ E. Turn the tangent-screw which mores the com-pass-box till the smaller division (in this case $31^{\circ}$ ) has come round to the needle. The vernier will then indicate through what space the compass-box has moved, and therefore how much must be added to the reading of the needle. Suppose it indicates ten minutes of a degree. Then the bearing is $\mathrm{N} .31^{\circ} 10^{\prime} \mathrm{E}$. It is, however, so difficult to move the vernier without disturbing the whole instrument, that this is seldom resorted to in practice. The chief use of the vernier is to set the instrument for running lines and making an allowance for the variation of the needle, as will be explained in the proper place. A vernier arc is sometimes attached to one end of the needle and carried around by it.
183. Practical Hints. Mark erery station or spot at which the compass is set by driving a stake, or digging up a sod, or piling up stones, or otherwise, so that it can be found if any error or other cause makes it necessary to repeat the surrey.

Very often, when the line of which the bearing is required is a fence, etc., the compass can not be set upon it. In such cases, set the compass so that its center is a foot or two from the line, and set the flag-staff at precisely the same distance from the line at the other end of it. The bearing of the flag-staff from the compass will be the same as that of the fence, the two lines being parallel. The distances should be measured on the real line. If more convenient, the compass may be set at some point on the line prolonged, or at some intermediate point of the line, "in line" between its extremities.

In setting the compass lerel, it is more important to hare it level crosswise of the sights than in their direction : since, if it be
not so, on looking up or down hill through the upper part of one sight and the lower part of the other, the line of sight will not be parallel to the N . and S . or zero line on the compass, and an incorrect bearing will therefore be obtained.

The compass should not be lereled by the needle, as some books recommend-i. e., so leveled that the ends of the needle shall be at equal distances below the glass. The needle should be brought so originally by the maker, but, if so adjusted in the morning, it will not be so at noon, owing to the daily variation in the dip. If, then, the compass be leveled by it, the lines of sight will generally be more or less oblique, and therefore erroneous. If the needle touches the glass when the compass is leveled, balance it by sliding the coil of wire along it.

The same end of the compass should always go ahead. The north end is preferable. The south end will then be nearest to the observer. Attention to this and to the caution in the next paragraph will prevent any confusion in the bearings.

Always take the readings from the same end of the needle; from the north end, if the north end of the compass goes ahead, and vice versa. This is necessary, because the tro ends will not always cut opposite degrees. With this precaution, however, the angle of two meeting lines can be obtained correctly from either end, provided the same one is used in taking the bearings of both the lines.

Guard against a very frequent source of error with beginners in reading from the wrong number of the two between which the needle points, such as reading $34^{\circ}$ for $26^{\circ}$ in a case like that in the figure.

Check the vibrations of the needle

Fig. 151.
 by gently raising it off the pivot so as to touch the glass, and letting it down again by the screw on the under side of the box.

The compass should be smartly tapped after the needle has settled, to destroy the effect of any adhesion to the pivot or friction of dust upon it.

All iron, such as the chain, etc., must be kept at a distance
from the compass, or it will attract the needle, and cause it to deviate from its proper direction.

The surveyor is sometimes troubled by the needle refusing to traverse and adhering to the glass of the compass after he has briskly wiped this off with a silk handkerchief, or it has been carried so as to rub against his clothes. The cause is the electricity excited by the friction. It is at once discharged by applying a wet finger to the glass.

A compass should be carried with its face resting against the side of the surreyor, and one of the sights hooked orer his arm.

In distant surveys an extra center-pin should be carried (as it is rery liable to injury, and its perfection is most essential), and also an extra needle. When two such are carried they should be placed so that the north pole of one rests against the south pole of the other.
184. When the magnetism of the needle is lessened or destroyed by time, it may be renewed as follows: Obtain tro bar magnets. Provide a board with a hole to admit of the axis, so that its collar may fit fairls, and that the needle may rest flat on it without bearing at the center. Place the board before you with the north end of the needle to your right. Take a magnet in each hand, the left holding the north end of the bar, or that which has the mark across, downward, and the right bolding the same mark upward. Bring the bars orer the axis, about a foot abore it, without approaching each other within two inches; bring them down vertically on the needle (the marks as directed) about an inch on each side of its axis ; slide them outward to its ends with slight pressure ; raise them up; bring them to their former position, and repeat this a number of times.
185. Back-Sights. To test the accuracy of the bearing of a line taken at one end of it, set up the compass at the other end or point sighted to, and look back to a rod held at the first station or point where the compass had been placed originally. The reading of the needle should now be the same as before.

If the position of the sights had been reversed, the reading
would be the Reverse Bearing; a former bearing of N. $30^{\circ}$ E. would then be S. $30^{\circ} \mathrm{W}$., and so on.
186. Local Attraction. If the back-sight does not agree with the first or forward sight, this latter must be taken over again. If the same difference is again found, this shows that there is local attraction at one of the stations-i. e., some influence, such as a mass of iron-ore, ferruginous rocks, etc., under the surface, which attracts the needle, and makes it deviate from its usual direction. Any high object, such as a house, a tree, etc., has been found to produce a similar effect.

To discover at which station the attraction exists, set the compass at several intermediate points in the line which joins the two stations, and at points in the line prolonged, and take the bearing of the line at each of these points. The agreement of several of these bearings, taken at distant points, will prove their correctness. Otherwise, set the compass at a third station, sight to each of the two doubtful ones, and then from them back to this third station. This will show which is correct.

When the difference occurs in a series of lines, such as around a field or along a road, proceed thus: Let C be the station at which the back-sight to B differs from the fore-sight from B to C. Since the back-sight from B to A is supposed to have agreed with the fore-sight

Fig. 152.
 from A to B, the local attrac-

Fig. 153.

tion must be at C , and the forward bearing must be corrected by the difference just found between the foreand back-sights, adding or subtracting it, according to circumstances. An easy method is to draw a figure for the case, as in Fig. 153. In it, suppose the true bearing of BC , as given by a fore-sight from B to C , to be $\mathrm{N} .40^{\circ}$ E., but that there is local attraction
at $C$, so that the needle is drawn aside $10^{\circ}$, and points in the direction $S^{\prime} \mathrm{N}^{\prime}$ instead of SN . The back-sight from $C$ to $B$ will then give a bearing of $N .50^{\circ}$ E. ; a difference or correction for the next fore-sight of $10^{\circ}$. If the next fore-sight, from C to D , be N. $70^{\circ}$ E., this $10^{\circ}$ must be subtracted from it, making the true fore-sight N. $60^{\circ} \mathrm{E}$.

A general rule may also be given. When the back-sight is greater than the fore-sight, as in this case, subtract the difference from the next fore-sight, if that course and the preceding one have both their letters the same (as in this case, both being $N$. and E.), or both their letters different ; or add the difference if either the first or last letters of the two courses are different. When the back-sight is less than the fore-sight, add the difference in the case in which it has just been directed to subtract it, and subtract it where it was before directed to add it.
187. Angles of Deflection. When the compass indicates much local attraction, the difference between the directions of two meeting lines (or the "angle of deflection" of one from the other) can still be correctly measured by taking the difference of the bearings of the two lines, as observed at the same point. For the error caused by the local attraction, whaterer it may be, affects both bearings equally, inasmuch as a "bearing" is the angle which a line makes with the direction of the needle, and that here remains fixed in some one direction, no matter what, during the taking of the two bearings. Thus, in Fig. 153, let the true bearing of BCi. e., the angle which it makes with the line S N -be, as before, N. $40^{\circ} \mathrm{E}$., and that of CD,N. $60^{\circ} \mathrm{E}$. The true "angle of deflection" of these lines, or the angle $\mathrm{B}^{\prime} \mathrm{CD}$, is therefore $20^{\circ}$. Now, if local attraction at $C$ causes the needle to point in the direction of $\mathrm{S}^{\prime} \mathrm{N}^{\prime}, 10^{\circ}$ to the left of its proper direction, B C will bear N . $50^{\circ} \mathrm{E}$., and CD N. $70^{\circ} \mathrm{E}$., and the difference of these bearingsi. e., the angle of deflection-will be the same as before.
188. Angles between Courses. To determine the angle of deflection of two courses meeting at any point, the following simple rules, the reasons of which will appear from the accompanying figures, are sufficient:

Case 1. When the first letters of the bearing are alike (i. e., both N. or both S.), and the last letters also alike (i. e., both E. or both W.), take the difference of the bearings. Example: If AB bears N. $30^{\circ}$ E., and B C bears N. $10^{\circ}$ E., the angle of deflection $\mathrm{CBB} \mathrm{B}^{\prime}$ is $20^{\circ}$.

Case 2. When the first letters are alike and the last letters different, take the sum of the bearings. Ex.: If AB bears N. $40^{\circ}$ E. and B C bears N. $20^{\circ}$ W., the angle $\mathrm{CBB} \mathrm{B}^{\prime}$ is $60^{\circ}$.

Case 3. When the first letters are
 different and the last letters alike, sub-

Fig. 155.
 tract the sum of the bearings from $180^{\circ}$. Ex. : If AB bears N. $30^{\circ}$ E. and B C bears S. $40^{\circ} \mathrm{E}$., the angle $\mathrm{C} \mathrm{B} \mathrm{B}{ }^{\prime}$ is $110^{\circ}$.

Case 4. When both the first and last letters are different, subtract the difference of the bearings from $180^{\circ}$. Ex.: If A B bears S. $30^{\circ} \mathrm{W}$. and BC bears N. $70^{\circ} \mathrm{E}$., the angle C B B' is $140^{\circ}$.

If the angles included between the courses are desired, they will be at once found by reversing one bearing and then applying the above rules; or by subtracting the results obtained as above from $180^{\circ}$; or an analogous set of rules could be formed for them.

Fig. 156.


Fig. 157.

189. To change Bearings. It is convenient in certain calculations to suppose one of the lines of a surrey to change its direction so as to become due north and south; that is, to become a new meridian line. It is, then, necessary to determine what the bearings of the other lines will be, supposing them to change with it. The subject may be made plain by supposing the survey to be platted in the usual way, with the north uppermost, and the plat to be then turned around till the line to be changed is in the desired direction. The effect of this on the other lines will be readily seen. A general rute can also be formed :

Take the difference between the original bearing of the side which becomes a meridian, and each of those bearings which have both their letters the same as it, or both different from it. The changed bearings of these lines retain the same letters as before, if they were originally greater than the original bearing of the new meridian line ; but, if they were less, they are thrown on the other side of the N . and S . line, and their last letters are changed, E . being put for W., and W. for E.

Take the sum of the original bearing of the new meridian line, and each of those bearings which have one letter the same as one letter of the former bearing and one different. If this sum exceeds $90^{\circ}$, this shows that the line is thrown on the other side of the east or west point, and the difference between this sum and $180^{\circ}$ will be the new bearing, and the first letter will be changed, N . being put for S. and S. for N.

Example: Let the bearings of the sides of a field be as follows: N. $32^{\circ}$ E. ; N. $80^{\circ}$ E. ; S. $48^{\circ}$ E. ; S. $18^{\circ}$ W. ; N. $73 \frac{1^{\circ}}{}{ }^{\circ} \mathrm{T}$. ; North. Suppose the first side to become due north ; the changed bearings will then be as follows : North ; N. $48^{\circ}$ E. ; S. $80^{\circ}$ E. ; S. $14^{\circ}$ E. ; S. ${ }^{7} 44^{\circ}{ }^{\circ} \mathrm{W} . ;$ N. $32^{\circ} \mathrm{W}$.

To apply the rule to the "North" course, as above, it must be called N. $0^{\circ} \mathrm{W}$. ; and then, by the rule, $32^{\circ}$ mutst be added to it.

The true bearings can, of course, be obtained from the changed bearings by reversing the operation, taking the sum instead of the difference, and vice versa.
190. Line-Surveying. This name may be given to surveys of lines, such as the windings of a brook, the curves of a road, ctc., by way of distinction from Farm-Surveying, in which the lines surveyed inclose a space.

To survey a brook, or any similar line, set the compass at or near one end of it, and take the bearing of an imaginary or visual

line running in the general average direction of the brook, such as A B in the figure. Measure this line, taking offsets to the various bends of the brook, as explained in Art. 9\%. Then set the compass at $B$, and take a back-sight to $A$, and, if they agree, take a fore-sight to C , and proceed as before, noting particularly the points where the line crosses the brook.

To survey a road, take the bearings and lengths of the lines
Fig. 159.

which can be most conveniently measured in the road, and measure offsets on each side to the outside of the road.

When the line of a new road is surveyed, the bearings and lengths of the various portions of its intended center-line should be measured, and the distance which it runs through each man's land should be noted. Stones should be set in the ground at recorded distances from each angle of the line, or in each line prolonged a known distance, so as not to be disturbed in making the road.

In surveying a wide river, one bank may be surveyed by the method just given, and points on the opposite banks, as trees, etc., may be fixed by the method of intersections founded on the fourth method of determining the position of a point.
191. Checks by Intersecting Bearings. At each station at which the compass is set take bearings to some remarkable object, such as a church-steeple, a distant house, a high tree, etc. At least three bearings should be taken to each object to make it of any use, since two are necessary to determine it (by our fourth method), and, till thus determined, it can be no check. When the line is platted, by the methods to be explained hereafter, plat also the lines given by these bearings. If those taken to the same object from three different stations intersect in the same point, this proves that there has been no mistake in the survey or platting of those stations.

If any bearing does not intersect a point fixed by previous bearings, it shows that there has been an error, either between the last station and one of those which fixed the point, or in the last bearing to the point. To discover which it was, plat the following line of the survey, and, at its extremity, set off the bearing from it to the point, and, if the line thus platted passes through the point, it proves that there was no error in the line, but only in the bearing to the point. If otherwise, the error was somewhere in the line between the stations from which the bearings to that point were taken.
192. Keeping the Field-Notes. The simplest and easiest method for a beginner is to make a rough sketch of the surrey by ese, and write down on the lines their bearings and lengths.

An improvement on this is to actually lay down the precise bearings and lengths of the lines in the field-book in the manner to be explained in the section on Platting, Art. 209.
193. A second method is to draw a straight line up the page of the field-book, and to write on it the bearings and lengths of the lines. The only advantage of this method is that the line will not run off the side of the page, as it is apt to do in the preceding method.
194. A third method is to represent the line surresed by a double column, as in Art. 84, which should be now referred to. The bearings are written obliquely up the columns. At the end of
each course its length is written in the column, and a line drawn across it. Dotted lines are drawn across the column at any intermediate measurement. Offsets are noted as explained in Art. 9\%.

The intersection bearings, described in Art. 191, should be entered in the field-book before the bearings of the line, in order to avoid mistakes of platting in setting off the measured distances on the wrong line.
195. A fourth method is to write the stations, bearings, and distances in three columns. This is compact, and has the advantage, when applied to farm-surveying, of presenting a form suitable for the subsequent calculations of content, but does not give facilities for noting offsets.

Examples of these four methods are given in Art. 199, which contains the field-notes of the lines bounding a field.
196. New York Canal-Maps. The following is a description of the original maps of the survey of the line of the New York Erie Canal, as published by the Canal Commissioners. The figure represents a portion of such a map, but, necessarily, with all its lines black, red and blue lines being used on the real map:
"The Red Line described along the inner edge of the towingpath is the base-line, upon which all the measurements in the di-

Fig. 160.

rection of the length of the canal were made. The bearings refer to the magnetic meridian at the time of the survey. The lengths of the several portions are inserted at the end of each in chains and links. The offsets at each station are represented by red lines drawn across the canal in such a direction as to bisect the angles
formed by the two contiguous portions of the red or base line upon the towing-path. The intermediate offsets are set off at right angles to the base-line, and the distances on both are given from it in links. The intermediate offsets are represented by red dotted lines, and the distances to them upon the base-line are reckoned, in each case, from the last preceding station. The same is likewise done with the other distances upon the base-line; those to the bridges being taken to the lines joining the nearest angles or corner posts of their abutments; those to the locks extending to the lines passing through the centers of the two nearest quoin-posts; and those to the aqueducts to the faces of their abutments. The space inclosed by the Blue Lines represents the portion embraced within the limits of the survey as belonging to the State; and the names of the adjoining proprietors are given as they stood at the time of executing the survey. The distances are projected upon a scale of two chains to the inch."
197. Farm-Surveying. A farm or field or other space included within known lines is usually surveyed by the compass thus : Begin by walking around the boundary-lines and setting stakes at all the corners, which the flag-man should specially note, so that he may readily find them again. Then set the compass at any corner, and send the flag-man to the next corner. Take the bearing of the bounding-line running from corner to corner, which is usually a fence. Measure its length, taking offsets if necessary. Note where any other fence, or road, or other line crosses or meets it, and take their bearings. Take the compass to the end of this first bounding-line; sight back, and, if the back-sight agrees, take the bearing and distance of the next bouinding-line ; and so proceed till you have got back to the point of starting.
198. Where speed is more important than accuracy in a surrey, whether of a line or a farm, the compass need be set only at every other station, taking a forward sight from the first station to the second; then, setting the compass at the third station, taking a back-sight to the second station (but with the north point of the compass always ahead), and a fore-sight to the fourth : then going to the fifth, and so on. This is, howerer, not to be recommended.

199．Field－Notes．
The field－notes of a farm－survey may be kept by any of the methods which have been described with reference to a line－（1） survey．Below are given the field－notes of the same field re－ corded by each of the methods．


Second Method．Third Method．＊

| （1） | －（1）－ |
| :---: | :---: |
|  | $3 \cdot 23$ |
|  | $\pm$ |
|  | －18 |
|  | 20 |
|  | 立 |
| $\geqslant$ | －（5）－ |
|  | $3 \cdot 55$ |
|  | $\geqslant$ |
|  | ＋ |
| $\dot{\sim}$ | $\stackrel{7}{6}$ |
| $\bigcirc(4)$ | ๗் |
| 乵 | －（4）－ |
| $\bigcirc$ | $2 \cdot 22$ |
| 20 or | 1 |
| $\odot(3)$ | $\stackrel{ }{ }$ |
|  | 2 |
| 込 | $\stackrel{\sim}{\circ}$ |
| $\stackrel{\circ}{\circ}$ | －（3）－ |
|  | $1 \cdot 29$ |
| $\infty$ | $\stackrel{\square}{2}$ |
| $\bigcirc(2)$ | 8 |
| $\pm$ | $\stackrel{8}{2}$ |
| ${ }^{\circ} \mathrm{c}$ ¢ | （2） |
|  | 2.70 |
|  | 凩 |
| －（1） | ${ }^{2}$ |
|  | z |
|  | －（1）－ |

（5）
Fourth Method．

| stations． | bearings． | distances． |
| :---: | :---: | :---: |
| 1 | N． $35^{\circ}$ E． | $2 \cdot 70$ |
| 2 | N． $83 \frac{1}{2}^{\circ} \mathrm{E}$ ． | $1 \cdot 29$ |
| 3 | S． $57^{\circ} \mathrm{E}$ ． | $2 \cdot 22$ |
| $\stackrel{4}{5}$ | S． $34^{\frac{1}{4}}{ }^{\circ} \mathrm{W}$ ． | 3.55 |
| 5 | N． $56 \frac{1}{2}{ }^{\circ} \mathrm{W}$ ． | $3 \cdot 23$ |



[^15]200. The field-notes of a field in which offsets occur may be most easily recorded by the third method, as in Fig. 162.

When the field-notes are recorded by the fourth method, the offsets may be kept in a separate table, in which the first column will contain the stations from which the measurements are made, the second column the distances at which they occur, the third column the lengths of the offsets, and the fourth column the side of the line, "right" or "left," on which they lie.

For calculation, four more columns may be added to the table, containing the intervals between the offsets, the sums of the adjoining pairs, and the products of the numbers in the two preceding columns, separated into right and left, one being additire to the field, and the other subtractive.
201. Tests of Accuracy. 1. The check of intersections described in Art. 191 may be employed to great advantage when some conspicuous object near the center of the farm can be seen from most of its corners.
2. When the surrey is platted, if the last course meets the starting-point, it proves the work, and the survey is then said to "close."
3. Diagonal lines running from corner to corner of the farm, like the "proof-lines" in chain-surreying, may be measured and their bearings taken. When these are laid down on the plat, their meeting the points to which they had been measured proves the work.
4. The only certain and precise test is, however, that by "latitudes and departures."
202. Method of Radiation. A field may be surveyed from one station, either within it or without it, by taking the bearings and the distances from that point to each of the corners of the field. These corners are then "determined" by the third method, Art. 5. This modification of that method is called the Method of Radiation. All our preceding survers with the compass hare been by the Method of Progression.

The compass may be set at one corner of the field, or at a point
in one of its sides, and the same method of radiation employed.

This method is seldom used, however, since, unlike the method of progression, its operations are not checks upon each other.
203. Method of Intersection. A field may also be surveyed by measuring a base-line, either within it or without it, setting the compass at each end of the base-line, and taking from each end the bearings of each corner of the field, which will then be fixed and determined by the fourth method, Art. 6. This mode of surveying is the Method of Intersections, noticed in Art. 166.
204. Running out Old Lines. The original surveys of lands in the older States of the American Union were exceedingly deficient in precision. This arose from two principal causes: the small value of land at the period of these surveys, and the want of skill in the surveyors. The effect at the present day is frequent dissatisfaction and litigation. Lots sometimes contain more acres than they were sold for, and sometimes less. Lines which are straight in the deed and on the map are found to be crooked on the ground. The recorded surveys of two adjoining farms often make one overlap the other, or leave a gore between them. The most difficult and delicate duty of the land-surveyor is to run out these old boundary-lines. In such cases, his first business is to find monuments, stones, marked trees, stumps, or any other old "corners" or landmarks. These are his starting-points. The owners whose lands join at these corners should agree on them. Old fences must generally be accepted by right of possession, though such questions belong rather to the lawyer than to the surveyor.* His business is to mark out on the ground the lines given in the deed. When the bounds are given by compass-bearings, the surveyor must be reminded that these bearings are very far from being the same now as originally, having been changing every year. The method of

[^16]determining this important change, and of making the proper allowance, will be found under "Declination of the Magnetic Needle."

## PLATtING THE SURVEY.

205. The platting of a surrey made with the compass consists in drawing on paper the lines and the angles which have been measured on the ground. The angles are laid off and the lines are drawn "to scale," as has been explained in Chapter I.
206. Platting Bearings. Since "bearings" taken with the compass are the angles which the rarious lines make with the magnetic meridian, or the direction of the compass-needle, which, as we have seen, remains always (approximately) parallel to itself, it is necessary to draw these meridians through each station before laying off the angles of the bearings.

The $T$-square is the most convenient instrument for this purpose. The paper on which the plat is to be made is fastened on the board so that the intended direction of the north and south line may be parallel to one of the sides of the board. The inner side of the stock of the T-square being pressed against one of the other sides of the board and slid along, the edge of the long blade of the square will always be parallel to itself and to the first-named side of the board, and will thus represent the meridian passing through any station.

If a straight-edged drawing-board or table can not be procured, nail down on a table of any shape a

Fig. 163.
 straight-edged ruler, and slide along against it the outside of the stook of a $T$-square, one side of the stock being flush with the blade.

A parallel ruler may also be used, one part of it being scretred down to the board in the proper position.

If none of these means are at hand, approximately parallel meridians may be drawn be the edges of a common ruler at distances apart equal to its width, and the diameter of the protractor made parallel to them by measuring equal distances between it and them.

20\%. To plat a survey with these instruments, mark with a fine point inclosed in a circle a convenient spot in the paper to represent the first station, 1 in the figure. Its place must be so chosen that the plat may not "run off" the paper. With the T-square draw a meridian through it. The top of the paper is usually, though not necessarily, called north. With the protractor lay off the angle of the first bearing. Set off the length of the first line to the

Fig. 164.
 desired scale from 1 to 2. The line $1---2$ represents the first course.

Through 2 draw another meridian, lay off the angle of the second course, and set off the length of this course from 2 to 3.

Proceed in like manner for each course. When the last course is platted, it should end precisely at the starting-point, as the survey did, if it were a closed survey, as of a field. If the plat does not "close" or "come together," it shows some error or inaccuracy either in the original survey, if that have not been "tested" by latitudes and departures, or in the work of platting. The plat here given is the same as that of Fig. 161.

This manner of laying down the directions of lines by the angles which they make with a meridian line has a great advantage, in both accuracy and rapidity, over the method of platting lines by the angles which each makes with the line which comes before it. In the latter method, any error in the direction of one line makes all that follow it also wrong in their directions. In the former, the direction of each line is independent of the preceding line, though its position would be changed by a previous error.

Instead of drawing a meridian through each station, sometimes only one
is drawn, near the middle of the sheet, and all the bearings of the surrey are laid off from some one point of it, as shown in the figure, and numbered to correspond with the stations from which these bearings were taken. The circular protractor is convenient for this. They are then transferred to the places where they are wanted by a triangle or other parallel ruler. Fig. 165 represents the same field platted by this method.

A semicircular protractor is sometimes attached to the stock end of the T-square so that its blade may be set at any desired angle with the meridian, and any bearing be thus protracted without drawing a meridian. It has some inconveniences.

The compass itself may be used to plat bearings. For this purpose it must be at-
 tached to a square board so that the N and $S$ line of the compass-box may be parallel to two opposite edges of the board. This is placed on the paper, and the box is turned till the needle points as it did when the first bearing was taken. Then a line drawn by one edge of the board will be in a proper direction. Mark off its length, and plat the next and the succeeding bearings in the same manner.
208. When the plat of a surrey does not "close," it may be corrected as follows: Let A B CDE be the boundary-lines platted according to the given bearings and distances, and suppose that the last course comes to E instead of ending at A, as it should. Suppose also that there is no reason to suspect any single great error, and that no one of the lines was measured over rery rough ground, or was specially uncertain in its direction when obserred. The inaccuracy must then be distributed among all the lines in proportion to their length. Each point in the figure, B, C, D, E, must
be moved in a direction parallel to EA by a certain distance which is obtained thus: Multiply the distance EA by the distance A B, and divide by the sum of all the courses. The quotient will be the distance B B'. To get C C', multiply E A by $A B+B C$, and divide the product by the same sum of all the courses. To get
 D D', multiply EA by $A B+B C+C D$, and divide as before. So for any course, multiply by the sum of the lengths of that course and of all those preceding it, and divide as before. Join the points thus obtained, and the closed polygon $\mathrm{A} \mathrm{B}^{\prime} \mathrm{C}^{\prime} \mathrm{D}^{\prime} \mathrm{A}$ will thus be formed, and will be the most probable plat of the given surves.*

The method of latitudes and departures, to be explained hereafter, is, however, the best for effecting this object.
209. Field Platting. It is sometimes desirable to plat the courses of a survey in the field as soon as they

Fig. 167.
 are taken, as was mentioned in Art. 192 under the head of "Keeping the Field-Notes." One method of doing this is to have the paper of the field-book ruled with parallel lines at unequal distances apart, and to use a rectangular protractor (which may be made of Bristol-board or other stout drawing-paper) with lines ruled across it at equal distances of some fraction of an inch. A bearing having been taken and noted, the protractor is laid on the paper and its center placed at the station where the bearing is to be laid off. It is then turned till one of its cross-lines coincides with some one of the lines on the paper, which represent east and west lines. The long side of the protractor will then be on a meridian, and the proper angle ( $40^{\circ}$ in the figure) can be at once marked off. The length of the course can also be set off by the equal spaces between the cross-lines, letting each space represent any convenient number of links.

[^17]210. A common rectangular protractor without any cross-lines, or a semicircular one, can also be used

Fig. 168.
 for the same purpose. The parallel lines on the paper (which, in this method, may be equidistant, as in common ruled writing-paper) will now represent meridians. Place the center of the protractor on the meridian nearest to the station at which the angle is to be laid off, and turn it till the given number of degrees is cut by the meridian. Slide the protractor up or down the meridian (which must continue to pass through the center and the proper degree) till its edge passes through the station, and then draw by this edge a line, which will have the bearing required.
211. Paper ruled into squares (as are sometimes the right-hand pages of surveyors' field-books) may be used for platting bearings in the field. The lines running up the page may be called north and south lines, and those running across the page will then be east and west lines. Any course of the survey will be the hypotenuse of a right-angled triangle, and the ratio of its other two sides will determine the angle. Thus, if the ratio of the trio sides of the rightangled triangle, of which the line AB in the figure is the hypotemuse, is 1 , that line makes an angle of $45^{\circ}$ with the meridian. If the ratio of the long to the short side

Fig. 169.
 of the right-angled triangle, of which the line AC is the hypotenuse, is 4 to 1 , the line $\mathrm{A} C$ makes an angle of $14^{\circ}$ with the meridian. The line AD, the hypotenuse of an equal triangle which has its long side lying east and west, makes likewise an angle of $14^{\circ}$ with that side, and therefore makes an angle of $76^{\circ}$ with the meridian.
212. With a Paper Protractor. Engraved paper protractors mar be obtained from the instrument-makers, and are very convenient. A circle of large size, divided into degrees and quarters, is engraved on copper, and imppressions from it are taken on drawing-paper. The divisions are not nom-
bered. Draw a straight line to represent a meridian through the center of the circle in any convenient direction. Number the degrees from $0^{\circ}$ to $90^{\circ}$ each way from the ends of this meridian, as on the compass-plate. The protractor is now ready for use. Choose a convenient point for the first station. Suppose the first bearing to be N. $30^{\circ} \mathrm{E}$. The line passing through the center of the circle and through the opposite points N. $30^{\circ}$ E. and S. $30^{\circ}$ W. has the bearing required. But it does not pass through the station 1. Transfer it thither by drawing through station 1 a line
 parallel to it, which will be the course required, its proper length being set off on it from 1 to 2. Now, suppose the bearing from 2 to be $\mathrm{S} .60^{\circ} \mathrm{E}$. Draw through 2 a line parallel to the line passing through the center of the circle and through the opposite points $\mathrm{S} .60^{\circ} \mathrm{E}$. and $\mathrm{N} .60^{\circ} \mathrm{W}$., and it will be the line desired. On it set off the proper length from 2 to 3 , and so proceed.

When the plat is completed, the engraved sheet is laid on a clean one and the stations "pricked through," and the points thus obtained on the clean sheet are connected by straight lines. The penciled plat is then rubbed off from the engraved sheet, which can be used for a great number of plats.

If the central circle be cut out, the plat, if not too large, can be made directly on the paper where it is to remain.

The surveyor can make such a paper protractor for himself with great ease by means of the Table of Chords at the end of this volume, the use of which is explained in Art. 215. The engraved ones may have shrunk after being printed.

Such a circle is sometimes drawn on the map itself. This will be particularly convenient if the bearings of any lines on the map not taken on the ground are likely to be required. If the map be very long, more than one may be needed.
213. Drawing-Board Protractor. Such a divided circle as has just been described, or a circular protractor, may be placed on a drawing-board near its center, and so that its $0^{\circ}$ and $90^{\circ}$ lines are parallel to the sides of the drawing-board. Lines are then to be drawn through the center and opposite divisions by a ruler long enough to reach the edges of the drawing-board on which they are to be cut in and numbered. The drawing-board thus becomes, in fact, a double rectangular protractor. A strip of white paper may have previously been pasted on the edges, or a narrow strip of white wood inlaid. When this is to be used for platting, a sheet of paper is put on the board as usual, and lines are drawn by a ruler laid across the $0^{\circ}$ points and the $90^{\circ}$ points, and the center of the circle is at once found, and should be marked $\odot$. The bearings are then platted as in the last method.
214. With a Scale of Chords. On the plane scale contained in cases of mathematical drawing instruments will be found a series of divisions numbered from 0 to 90 , and marked CH or C. This is a

Fig. 171.
 scale of chords, and gives the lengths of the chords of any are for a radius equal in length to the chord of $60^{\circ}$ on the scale. To lay off an angle with this scale, as, for example, to draw a line making at $A$ an angle of $40^{\circ}$ with A B, take, in the dividers, the distances from 0 to 60 on the scale of chords; with this for radius and A for center, describe an indefinite arc CD. Take the distance from 0 to 40 on the same scale, and set it off on the arc as a chord from C to some point D . Join $\mathrm{A} D$ and prolong it. BA E is the angle required.

The sector, Fig. 23, supplies a modification of this method sometimes more conrenient. On each of its legs is a scale marked C or CH. Open it at pleasure ; extend the compass from 60 to 60 , one on each leg, and with this radius describe an arc. Then extend the compasses from 40 to 40 , and the distance will be the chord of $40^{\circ}$ to that radius. It can be set off as above.

The smallness of the scale renders the method with a scale of chords practically deficient in exactness, but it serves to illustrate the next and best method.
215. With a Table of Chords. At the end of this volume mill be found a table of the lengths of the chords of arcs for every degree and minute of the quadrant calculated for a radius equal to 1 .

To use it, take in the compasses one inch, one foot, or any other conrenient distance (the longer the better), divided into tenths and hundredths br a diagonal scale or otherwise. With this as radius describe an arc as in the last case. Find in the table of chords the length of the chord of the desired angle. Take it from the scale just used to the nearest decimal part which the scale will give. Set it off as a chord, as in the last figure, and join the point thus obtained to the starting-point. This gires the angle desired.

The superiority of this method to that which emplors a protractor is due to the greater precision with which a straight line can be dirided than can a circle.

A slight moditication of this method is to take in the compasses ten equal parts of any convenient length, inches, half inches, quarter inches, or any other at hand, and with this radius describe an are as before, and set off a chord ten times as great as the one found in the table-i. e., imagine the decimal-point mored one place to the right.

If the radius be 100 or 1,000 equal parts, imagine the decimal-point mored two or three places to the right.

Whatever radius may be taken or giren, the product of that radius into a chord of the table will give the chord for that radius.

This gires an easy and exact method of getting a right angle br describing an are with a radius of 1 , and setting off a chord equal to $1 \cdot 4142$.

If the angle to be constructed is more than $90^{\circ}$, construct on the other side of the given point upon the given line prolonged an angle equal to what the given angle wants of $180^{\circ}-\mathrm{i}$. e., its supplement, in the language of trigonometry.

This same table gives the means of measuring any angle. With the angular point for a center, and 1 or 10 for a radius, describe an arc. Measure the length of the chord of the arc between the legs of the angle, find this length in the table, and the angle corresponding to it is the one desired.

This table will also serve to find the natural sine or cosine of any angle. Multiply the given angle by two; find in the table the chord of this domble angle; and half of this chord will be the natural sine required. For the chord of any angle is equal to twice the sine of half the angle. To find the cosine, proceed as above, with the angle which, added to the given angle, would make $90^{\circ}$.

Another use of this table is to inscribe regular polygons in a circle by setting off the chords of the arcs which their sides subtend.

Still another use is to divide an arc or angle into any number of equal parts by setting off the fractional arc or angle.
216. With a Table of Natural Sines. In the absence of a table of chords, heretofore rare, a table of natural sines, which can be found anywhere, may be used as a less convenient substitute. Since the chord of any angle equals twice the sine of half the angle, divide the given angle by two: find in the table the natural sine of this half angle ; double it, and the product is the chord of the whole angle. This can then be used precisely as was the chord in the preceding article.

An ingenious modification of this method has been much used. Describe an arc from the given point as center, as in the last two articles, but with a radius of five equal parts. Take from a table the length of the natural sine of half the given angle to a radius of ten. Set off this length as a chord on the are just described, and join the point thus obtained to the given point.

The reason of this is apparent from the figure. D E is the sine of half the angle BAC to a radius of ten equal parts, and BC is the chord directed to be set off to a radius of five equal parts. BC is equal to DE , for $\mathrm{BC}=2 \cdot \mathrm{BF}$ by trigonometry, and $\mathrm{DE}=2 \cdot \mathrm{BF}$ by similar triangles; hence $\mathrm{BC}=\mathrm{D} \mathrm{E}$.
217. By Latitudes and Departures. When the latitudes and departures of a survey have been obtained and corrected, either to test its accuracy or to obtain its content, they afford the easiest and best means of platting it. The description of this method will be given in Art. 246.

Fig. 172.


## COFYING PLATS.

218. The plat of a survey necessarily has many lines of construction drawn upon it which are not needed in the finished map. These lines and the marks of instruments so disfigure the paper that a fair copy of the plat is usually made before the map is finished. The various methods of copying plats, etc., whether on the same scale, or reduced, or enlarged, will therefore now be described.
219. Stretching the Paper. If the map is to be colored, the paper must first be wetted and stretched, or the application of the wet colors will cause its surface to swell or blister and become uneven. Therefore, with a soft sponge and clean water, wet the back of the paper, working from the center outward in all directions. The " water-mark" reads correctly only when looked at from the front side, which it thus distinguishes. When the paper is thoroughly wet and thus greatly expanded, glue its edges to the draw-ing-board for half an inch in width, turning them up against a ruler, passing the glue along them, and then turning them down and pressing them with the ruler. Some prefer gluing down opposite edges in succession, and others adjoining edges. The paper must be moderately stretched smooth during the process. Hot glue is best. Paste or gum may be used, if the paper be kept wet by a damp cloth, so that the edges may dry first. "Mouth-glue" may be used by rubbing it (moistened in the mouth or in boiling water) along the turned-up edges, and then rubbing them dry by an ivory folder, a piece of dry paper being interposed. As this is a slower process, the middle of each side should first be fastened down, then the four angles, and lastly the intermediate portions. When the paper becomes dry, the creases and puckerings will have disappeared, and it will be as smooth and tight as a drum-head.
220. Copying by Tracing. Fix a large pane of clear glass in a frame so that it can be supported at any angle before a window, or, at night, in front of a lamp. Place the plat to be copied on this glass, and the clean paper upon it. Connect them by pins, etc. Trace all the desired lines of the original with a sharp pencil as
lightly as they can be easily seen. Take care that the paper does not slip. If the plat is larger than the glass, copy its parts successively, being very careful to fix each part in its true relative position. Ink the lines with India ink, making them very fine and pale if the map is to be afterward colored.
221. Copying on Tracing-Paper. A thin transparent paper is prepared expressly for the purpose of making copies of maps and drawings, but it is too delicate for much handling. It may be prepared by soaking tissue-paper in a mixture of turpentine and Canada balsam or balsam of fir (two parts of the former to one of the latter), and drying very slowly. Cold-drawn linseed-oil will answer tolerably, the sheets being hung up for some weeks to dry. Linen is also similarly prepared, and sold under the name of "vellum tracing-paper." It is less transparent than the tracing-paper, but is very strong and durable. Both of these are used rather for preserving duplicates than for finished maps.
222. Copying by Photography. This may be used for copying drawings, and is especially applicable when the drawings are to be very much reduced in size.
223. Copying by Blue Prints. Dissolve one ounce of ferricyanide of potassium in ten ounces of pure water. Also dissolve two ounces of ammonia citrate of iron in ten ounces of water. Mix the two solutions in a cup, and with a brush cover the surface of the paper on which the print is to be made with the mixture.

The surface should be thoroughly covered, but no more of the mixture should be applied than the paper will take up. The paper should become limp and moist but not wet. The work should be done in a room lighted with a lamp, and when the paper is dry it should be kept in a dark place.

To make a blue-print copy, a tracing of the drawing should first be made. Put the tracing over a sheet of the prepared paper and a sheet of glass over the tracing, in order to keep the tracing in contact with the prepared paper. Expose the paper to the sunlight, with the glass toward the sun, until the lines of the drawing are plainly seen on the prepared paper. Wash the paper until the
water running off is no longer colored yellow. When dried, the lines of the drawing will be white upon a blue ground. The prepared paper for blue prints can be bought of dealers in engineers' supplies.

There are several similar methods of making prints, differing in the chemicals used, and in the color of the lines and background.
224. Copying by Transfer-Paper. This is thin paper, one side of which is rubbed with black-lead, etc., smoothly spread by cotton. It is laid on the clean paper, the blackened side downward, and the plat is placed upon it. All the lines of the plat are then gone over with moderate pressure by a blunt point, such as the eye-end of a small needle. A faint tracing of these lines will then be found on the clean paper, and can be inked at leisure. If the original can not be thus treated, it may first be copied on tracing-paper, and this copy be thus transferred. If the transfer-paper be prepared by rubbing it with lampblack ground up with hard soap, its lines will be ineffaceable. It is then called "Camp-paper."
225. Copying by Punctures. Fix the clean paper on a drawing-board and the plat over it. Prepare a fine needle with a sealing-was head. Hold it very truly perpendicular to the board, and prick through every angle of the plat, and every corner and intersection of its other lines, such as houses, fences, etc., or at least the two ends of every line. For circles, the center and one point of the circumference are sufticient. For irregular curres, such as rivers, etc., enough points must be pricked to indicate all their sinuosities. Work with system, finishing up one strip at a time, so as not to omit any necessary points nor to prick through any twice, though the latter is safer. When completed, remove the plat. The copy will present a wilderness of fine points. Select those which deternine the leading lines, and then the rest will be easily recognized. A beginner should first pencil the lines lightly, and then ink them. An experienced draughtsman will omit the penciling. Two or three copies may be thus pricked through at once. The holes in the original plat may be made nearly invisible by rubbing them on the back of the sheet with a paper-folder, or the thumb-nail.
226. Copying by Intersections. Draw a line on the clean paper equal in lengtl to some important line of the original. Two starting-points are thus obtained. Take in the dividers the distance from one end of the line on the original to a third point. From the corresponding end on the cops, describe an arc with this distance for radius and about where the point will come. Take the distance on the original from the other end of the line to the point, and describe a corresponding are on the copy to intersect the former are in a point which will be that desired. The principle of the operation is that of our "First Method" (Art. 3). Tro pairs of dividers may be used, as explained in Art. 82. "Triangular compasses," haring three legs, are used by fixing two of their legs on the tiro giren points of the original,
and the third leg on the point to be copied, and then transferring them to the copy. All the points of the original can thus be accurately reproduced. The operation is, however, very slow. Only the chief points of a plat may be thus transferred, and the details filled in by the following method:
227. Copying by Squares. On the original plat draw a series of parallel and equidistant lines. The T-square does this most readily. Draw a similar series at right angles to these. The plat will then be covered with squares, as in Fig. 43. On the clean paper draw a similar series of squares. The important points may now be fixed as in the last article, and the rest copied by eye, all the points in each square of the original being properly placed in the corresponding square of the copy, noticing whether they are near the top or bottom of each square, on its right or left side, etc. This method is rapid, and in skillful hands quite accurate.

Instead of drawing lines on the original, a sheet of transparent paper containing them may be placed over it; or an open frame with threads stretched across it at equal distances and at right angles.

This method supplies a transition to the Reduction and Enlargement of plats in any desired ratio; under which head Copying by the Pantagraph and Camera Lucida will be noticed.
228. Reducing by Squares. Begin, as in the preceding article, by drawing squares on the original, or placing them over it. Then on the clean paper draw a similar set of squares, but with their sides one half, one third, etc. (according to the desired reduction), of those of the original plat. Then proceed as before to copy into each small square all the points and lines found in the large square of the plat in their true positions relative to the sides and corners of the square, observing to reduce each distance, by eye, or as directed in the following article, in the given ratio.
229. Reducing by Proportional Scales. Many graphical methods of finding the proportionate length of the copy, of any line of the original, may be used. The " angle of reduction" is constructed thus: Draw any line A B. With it for radins and A for center, describe an indefinite arc. With B for center and a radius equal to one half, one third, etc., of A B according to the desired reduction, describe another are intersecting the former arc in C. Join

Fig. 173.
 A C. From A as center describe a series of arcs. Now, to reduce any distance, take it in the dividers, and set it off from $A$ on $A B$, as to $D$. Then the distance from $D$ to $E$, the other end of the arc passing through D , will be the proportionate length to be set off on the copy, in the manner directed in Art. 226.

The sector, or "compass of proportion," described in Art. 50, presents such an "angle of reduction," always ready to be used in this manner.

Fig. 174.


The " angle of reduction" may be simplified thus: Draw a line, AB , parallel to one side of the draw-ing-board, and another, B C, at right angles to it, and one half, etc., of it, as desired. Join A C. Then let A D be the distance required to be reduced. Apply a T-square so as to pass through D. It will meet A C in some point E , and D E will be the reduced length required.

Another arrangement for the same object is shown in Fig. 175. Draw two lines, A B, A C, at any angle, and describe a series of arcs from their intersection, $A$, as in the figure. Suppose the reduced scale is to be half the original scale. Divide the outermost arc into three equal parts, and draw a line from A to one of the points of division, as D. Then each are will be divided into parts, one of which is twice the other. Take any distance on the original scale, and find by trial which of the ares on the right-hand side of the figure it corresponds to. The other part of that are will be half of it, as desired.
"Proportional compasses," being properly set, reduce lines in any desired ratio. A simple form of them, known as " wholes and halves," is often useful. It consists of two slender bars, pointed at each end, and united

Fig. 175.
 by a pirot which is twice as far from one pair of the points as from the other pair. The long ends being set to ans distance, the short ends will give precisely half that distance.
230. Reducing by a Pantagraph. This instrument consists of two long and two short rulers, connected so as to form a parallelogram, and capable of being so adjusted that when a tracing-point attached to it is mored over the lines of a map, etc., a pencil attached to another part of it will mark on paper a precise copy, reduced on any scale desired. It is made in various forms. It is troublesome to use, though rapid in its work.
231. Reducing by a Camera Lucida. This is used in the Coast Survey Office. It can not reduce smaller than one fourth, without losing distinctness, and is very trying to the eyes. Squares drawn on the original are brought to apparently coincide with squares on the reduction, and the details are then filled in with the pencil, as seen through the prism of the instrument.
232. Enlarging Plats. Plats may be enlarged by the principal methods which have been given for reducing them, but this should be done as seldom as possible, since erery inaccuracy in the original becomes magnified in the copy. It is better to make a ner plat from the original data.
233. Conventional Signs. Tarious conventional signs or marks hare been adopted, more or less generally, to represent on maps the inequalities of
the surface of the ground, its different kinds of culture or natural products, and to objects upon it, so as not to encumber and disfigure it with much writing or many descriptive legends. This is the purpose of what is called Topographical Mapping. (See Part III, Topography.)
234. Orientation. The map is usually so drawn that the top of the paper may represent the north. A meridian line should also be drawn, both true and magnetic, as in Fig. 186. The number of degrees and minutes in the variation, if known, should also be placed between the two north points. Sometimes a compass-star is drawn and made very ornamental.
235. Lettering. The style in which this is done very much affects the general appearance of the map. The young surveyor should give it much attention and careful practice. It must all be in imitation of the best printed models. No writing, however beautiful, is admissible. The usual letters are the ordinary ROMAN CAPITALS, Small Roman, ITALIC CAPITALS, Small Italic, and GOTHIC OR EGYPTIAN. This last, when well done, is very effective. For the titles of maps, various fancy letters may be used. For very large letters, those formed only of the shades of the letters regarded as blocks (the body being rubbed out after being penciled as a guide to the placing of the shades) are most easily made to look well. The simplest lettering is generally the best. The sizes of the names of places, etc., should be proportional to their importance. Elaborate tables for various scales have been published. It is better to make the letters too small than too large. They should not be crowded. Pencil-lines should always be ruled as guides. The lettering should be in lines parallel to the bottom of the map, except the names of rivers, roads, etc., whose general course should be followed.
236. Borders. The Border may be a single heavy line, inclosing the map in a rectangle, or such a line may be relieved by a finer line drawn parallel and near to it. Time should not be wasted in ornamenting the border. The simplest is the best.
237. Joining Paper. If the map is larger than the sheets of paper at hand, they should be joined with a feather-edge, by proceeding thus: Cut, with a knife guided by a ruler, about one third through the thickness of the paper, and tear off, on the under side, a strip of the remaining thickness, so as to leave a thin, sharp edge. Treat the other sheet in the same way on the other side of it. When these two feather-edges are then put together (with paste, glue, or varnish), they will make a reat and strong joint. The sheet which rests upon the other must be on the right-hand side, if the sheets are joined lengthwise, or below if they are joined in that direction, so that the thickness of the edge may not cast a shadow when properly placed as to the light. The sheets must be joined before lines are drawn across them, or the lines will become distorted. Drawing-paper is now made in rolls of great length, so as to render this operation unnecessary.
238. Mounting Maps. A map is sometimes required to be mountedi. e., backed with canvas or muslin. To do this, wet the muslin and stretch
it strongly on a board by tacks driven very near together. Cover it with strong paste, beating this in with a brush to fill up the pores of the muslin. Then spread paste over the back of the paper, and when it has soaked into it apply it to the muslin, inclining the board, and pasting first a strip, about two inches wide, along the upper side of the paper, pressing it down with clean linen in order to drive out all air-bubbles. Press down another strip in like manner, and so proceed till all is pasted. Let it dry very gradually and thoroughly before cutting the muslin from the board.

Maps may be varnished with picture-varnish, or by applying four or five coats of isinglass-size, letting each dry well before applying the nest, and giving a full, flowing coat of Canada balsam diluted with the best oil of turpentine.

## LATITUDES AND DEPARTURES.

239. Definitions. The Latitude of a point is its distance north or south of some " Parallel of Latitude," or line running east or west. The Longitude of a point is its distance east or west of some "Meridian," or line running north and south. In compasssurveying, the magnetic meridian-i. e., the direction in which the magnetic needle points-is the line from which the longitudes of points are measured or reckoned.

The distance which one end of a line is due north or south of the other end is called the Difference of Latitude of the two ends of the line ; or its northing or southing; or simply its latitude.

The distance which one end of the line is due east or west of the other is here called the Difference of Longitude of the two ends of the line ; or its easting or westing; or its departure.

Latitudes and Departures are the most usual terms, and will be generally used hereafter, for the sake of brerity.

This subject may be illustrated geographically, by noticing that a traveler, in going from New York to Buffalo in a straight line, would go about one hundred and fifty miles due north, and tro hundred and fifty miles due west. These distances would be the differences of latitude and of longitude between the two places, or his northing and westing. Returning from Buffalo to Nem York, the same distances would be his southing and easting.*

[^18]In mathematical language, othe operation of finding the latitude and longitude of a line, from its bearing and length, would be called the transformation of Polar Coordinates into Rectangular Co-ordinates. It consists in determining, by our Second Principle, the position of a point which had originally been determined by the

## Fig. 176.

 Third Principle. Thus, in the figure (which is the same as that of Art. 7), the point $S$ is determined by the angle S A C and by the distance A S. It is also determined by the distances AC and CS, measured at right angles to each other ; and then, supposing $C S$ to run due north and south, CS will be the latitude, and A C the departure of the line A S.
240. Calculation of Latitudes and Departures. Let AB be a
 given line, of which the length AB , and the bearing (or angle, B A C, which it makes with the magnetic meridian), are known. It is required to find the differences of latitude and of longitude between its two extremities $A$ and $B$-that is, to find $A C$ and C B ; or, what is the same thing, B D and D A.

It will be $a\llcorner$ once seen that $A B$ is the hypotenuse of a right-angled triangle, in which the "Latitude" and the "Departure" are the sides about the right angle. We therefore know, from the principles of trigonometry, that

$$
\begin{aligned}
& \mathrm{AC}=\mathrm{AB} \cdot \cos \cdot \mathrm{BAC}, \\
& \mathrm{BC}=\mathrm{AB} \cdot \sin \cdot \mathrm{BAC}
\end{aligned}
$$

Hence, to find the latitude of any course, multiply the natural cosine of the bearing by the length of the course ; and to find the departure of any course, multiply the natural sine of the bearing by the length of the course.

If the course be northerly, the latitude will be north, and will be marked with the algebraic sign + , plus, or additive; if it be
southerly, the latitude will be south, and will be marked with the algebraic sign -, minus, or subtractive.

If the course be easterly, the departure will be east, and marked + , or additive; if the cuurse be westerly, the departure will be west, and marked -, or subtractive.
241. Formulas. The rules of the preceding article may be expressed thus :

$$
\begin{aligned}
\text { Latitude } & =\text { distance } \times \text { cos. bearing, } \\
\text { Departure } & =\text { distance } \times \text { sin. bearing. }
\end{aligned}
$$

From these formulas may be obtained others, by which, when any two of the above four things are given, the remaining two can be found.

When the Bearing and Latitude are given;
Distance $=\frac{\text { latitude }}{\text { cos. bearing }}=$ latitude $\times$ sec. bearing,
Departure $=$ latitude $\times$ tang. bearing.
When the Bearing and Departure are given;
Distance $=\frac{\text { departure }}{\text { sin. bearing }}=$ departure $\times$ cosec. bearing.
Latitude $=$ departure $\times$ cotang. bearing.
When the Distance and Latitude are given;

$$
\text { Cos. bearing }=\frac{\text { latitude }}{\text { distance }}
$$

Departure $=$ latitude $\times$ tang. bearing.
When the Distance and Departure are given;

$$
\text { Sin. bearing }=\frac{\text { departure }}{\text { distance }}
$$

Latitude $=$ departure $\times$ cotang. bearing.
When the Latitude and Departure are given;
Tang. of bearing $=\frac{\text { departure }}{\text { latitude }}$,
Distance $=$ latitude $\times$ sec. bearing.
Still more simply, any two of these three-distance, latitude, and departure-being given, we hare

$$
\begin{aligned}
\text { Distance } & =\sqrt{ }\left(\text { latitu }^{3} e^{2}+\text { departure }^{2}\right) \\
\text { Latitude } & =\sqrt{ }\left(\text { distance }^{2}-\text { departure }^{2}\right) \\
\text { Departure } & =\sqrt{ }\left(\text { distance }^{2}-\text { latitude }^{2}\right)
\end{aligned}
$$

[^19]242. Traverse-Tables. The latitude and departure of any distance, for any bearing, could be found by the method given in Art. 240, with the aid of a table of natural sines. But to facilitate these calculations, which are of so frequent occurrence and of so great use, traverse-tables have been prepared, originally for navigators (whence the name traverse), and subsequently for surreyors.*

The traverse-table at the end of this volume gives the latitude and departure for any bearing, to each quarter of a degree, and for distances from 1 to 9.

To use it, find in it the number of degrees in the bearing, on the left-hand side of the page, if it be less than $45^{\circ}$, or on the right-hand side if it be more. The numbers on the same line, running across the page, $\dagger$ are the latitudes and departures for that bearing, and for the respective distances- $1, \cdot 2,3,4,5,6,7,8,9$ -which are at the top and bottom of the page, and which may represent chains, links, rods, feet, or any other unit. Thus, if the bearing be $15^{\circ}$, and the distance 1 , the latitude would be 0.966 and the departure 0.259 . For the same bearing, but a distance of 8 , the latitude would be $\gamma \cdot \% 2 \%$ and the departure $2 \cdot 0 \% 1$.

Any distance, however great, can have its latitude and departure readily obtained from this table ; since, for the same bearing, they are directly proportional to the distance, because of the similar triangles which they form. Therefore, to find the latitude or departure for 60 , multiply that for 6 by 10 , which merely moves the decimal-point one place to the right; for 500 , multiply the numbers found in the table for 5 , by $100-\mathrm{i}$. e., move the decimalpoint two places to the right, and so on. Merely moving the deci-mal-point to the right, one, two, or more places, will therefore enable this table to give the latituce and departure for any decimal multiple of the numbers in the table.

[^20]For compound numbers, such as $8 \% 3$, it is only necessary to find separately the latitudes and departures of 800 , of 70 , and of 3 , and add them together. But this may be done, with scarcely any risk of error, by the following simple rule :

Write down the latitude and departure for the first figure of the given number, as found in the table, neglecting the decimalpoint; write under them the latitude and departure of the second figure, setting them one place farther to the right; under them write the latitude and departure of the third figure, setting them one place farther to the right ; and so proceed with all the figures of the given number. Add up these latitudes and departures, and cut off the three right-hand figures. The remaining figures will be the latitude and departure of the given number in links, or chains, or feet, or whatever unit it was given in.

For example : Let the latitude and departure of a course haring a distance of 873 links, and a bearing of $20^{\circ}$, be required. In the table find $20^{\circ}$, and then take out the latitude and departure for 8 , 7 , and 3 , in turn, placing them as above directed, thus:

| Distances. | Latitudes. | Departures. |
| :---: | :---: | :---: |
| 800 | 7518 | 2736 |
| 70 | 6578 | 2394 |
| 3 | $\underline{2819}$ | 1026 |
| 873 | $\overline{820.399}$ | $\underline{298.566}$ |

Taking the nearest whole numbers and rejecting the decimals, we find the desired latitude and departure to be $\delta 20$ and 299.*

When a 0 occurs in the given number, the next figure must be set two places to the right, the reason of which will appear from the following example, in which the 0 is treated like any other number :

Given a bearing of $35^{\circ}$, and a distance of 3048 links.

| Distances. | Latitudes. | Departures. |
| :---: | :---: | :---: |
| 3000 | $245 \%$ | $1 \% 21$ |
| 000 | 0000 | 0000 |
| 40 | $32 \% \%$ | 2294 |
| 8 | 65533 | 4589 |
| 3048 | $\overline{2496.323}$ | $\overline{1 \% 48.529}$ |

[^21]Here the latitudes and departures are 2496 and 1749 links.
When the bearing is over $45^{\circ}$, the names of the columns must be read from the bottom of the page, the latitude of any bearing, as $50^{\circ}$, being the departure of the complement of this bearing, or $40^{\circ}$, and the departure of $40^{\circ}$ being the latitude of $50^{\circ}$, etc. The reason of this will be at once seen on inspecting Fig. 17\%, and imagining the east and west line to become a meridian. For, if A C be the magnetic meridian, as before, and therefore BAC be the bearing of the course A B , then is AC the latitude, and CB the departure of that course. But if AE be the meridian and BAD (the complement of BAC ) be the bearing, then is AD (which is equal to CB ) the latitude, and DB (which is equal to A C) the departure.

As an example of this, let the bearing be $63 \frac{1}{4}^{\circ}$, and the distance 3,469 links. Proceeding as before, we have-

| Distances. | Latitudes. | Departures. |
| :---: | :---: | :---: |
| 3000 | 1350 | 2679 |
| 400 | 1800 | $35 \% 2$ |
| 60 | 2701 | 5358 |
| 9 | 4051 | $803 \%$ |
| 3469. | $\overline{1561 \cdot 061}$ | $\overline{309 \% \cdot 81 \%}$ |

The required latitude and departure are 1561 and 3098 links.
In the few cases occurring in compass-surveying, in which the bearing is recorded as somewhere between the fractions of a degree given in the table, its latitude and departure may be found by interpolation. Thus, if the bearing be $103^{\circ}{ }^{\circ}$, take the half sum of the latitudes and departures for $10 \frac{1}{4}^{\circ}$ and $10 \frac{1}{2}^{\circ}$. If it be $10^{\circ} 20^{\prime}$, add one third of the difference between the latitudes and departures for $10 \frac{1}{4}^{\circ}$ and for $10 \frac{1}{2}^{\circ}$, to those opposite to $101^{\circ}$; and so in any similar case.

The uses of this table are very varied. The principal applications of it, which will now be explained, are to testing the accu.racy of surveys; to supplying omissions in them; to platting them ; and to calculating their content.*

[^22]243. Application to testing a Survey. It is self-erident that, when the surveyor has gone completely around a field or farm, taking the bearings and distances of each boundary-line, till he has got back to the starting-point, he has gone precisely as far south as north, and as far west as east. But the sum of the north latitudes tells how far north he has gone, and the sum of the south latitudes how far south he has gone. Hence these two sums will be equal to each other, if the survey has been correctly made. In like manner, the sums of the east and of the west departures must also be equal to each other.

We will apply this principle to testing the accuracy of the survey of which Fig. 61 is a plat. Prepare seven columns, and head them as below. Find the latitude and departure of each course to the nearest link, and write them in their appropriate columns. Add up these columns. Then will the difference between the sums of the north and south latitudes, and between the sums of the east and west departures, indicate the degree of accuracy of the survey.

| stations. | bearings. | distances. | latitudes. |  | departires. |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | N. | s. | E. | w. |
| $\begin{aligned} & 1 \\ & 2 \\ & 3 \\ & 4 \\ & 5 \end{aligned}$ | $\begin{aligned} & \text { N. } 35^{\circ} \text { E. } \\ & \text { N. } 83 \frac{1}{3}^{\circ} \mathrm{E} . \\ & \text { S. } 57^{\circ} \\ & \text { S. } \\ & \text { S. } 341^{\circ} \mathrm{W} \\ & \text { N. } 561_{2}^{\circ} \end{aligned}$ | $\begin{aligned} & 2 \cdot 70 \\ & 1 \cdot 29 \\ & 2 \cdot 29 \\ & 3 \cdot 55 \\ & 3 \cdot 23 \end{aligned}$ | 2'21 |  | 1.55 |  |
|  |  |  | $\cdot 15$ |  | $1 \cdot 28$ |  |
|  |  |  |  | $1 \cdot 21$ |  |  |
|  |  |  |  | $2 \cdot 93$ |  |  |
|  |  |  | 1.78 |  |  | $2 \cdot 69$ |
|  |  |  | $4 \cdot 14$ | $4 \cdot 14$ | $4 \cdot 69$ | $4 \cdot 69$ |

The entire work of the above example is given on the following page.
for solving, approximately, any right-angled triangle by mere inspection, the bearing being taken for one of the acute angles; the latitude being the side adjacent, the departure the side opposite, and the distance the hypotenuse. Any two of these being given, the others are given by the table. The table will therefore serve to show the allowance to be made in chaining on slopes (see Art. 20). Look in the column of bearings for the slope of the ground-i. e., the angle it makes with the horizon, find the given distance, and the latitude corresponding will be the desired horizontal measurement, and the difference betreen it and the distance will be the allorance to be made.

| $35^{\circ}$ | $\begin{aligned} & 1638 \\ & 57340 \end{aligned}$ | $\begin{aligned} & 1147 \\ & 40150 \end{aligned}$ | $344^{\circ}$ | $\begin{gathered} 2480 \\ 4133 \end{gathered}$ | $\begin{gathered} 1688 \\ 2814 \end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 270 - | 221-140 | $154 \cdot 850$ |  | 4133 | 2814 |
|  |  |  | 355. | $293 \cdot 463$ | $199 \cdot 754$ |
| $83 \frac{1}{2}^{\circ}$ | 113 | 994 | $56 \frac{1}{2}^{\circ}$ | 16.56 | 2502 |
|  | 226 | 1987 |  | 1104 | 1668 |
|  | 1019 | 8942 |  | 1656 | 2502 |
| 129. | $\overline{14 \cdot 579}$ | $\overline{128 \cdot 212}$ | 323. | $\overline{178 \cdot 296}$ | $\overline{269 \cdot 382}$ |
| $57^{\circ}$ | 1089 | 1677 | The nearest link is taken to be inserted in the table, and the remaining decimals are neglected. |  |  |
|  | 1089 | 1677 |  |  |  |
|  | 1089 | 1677 |  |  |  |
| 222. | $120 \cdot 879$ | $186 \cdot 147$ |  |  |  |

In the preceding example the respective sums were found to be exactly equal. This, however, will rarely occur in an extensive survey. If the difference be great, it indicates some mistake, and the survey must be repeated with greater care ; but if the difference be small it indicates, not absoluţe errors, but only inaccuracies, unavoidable in surveys with the compass, and the survey may be accepted.

How great a difference in the sums of the columns may be allowed, as not necessitating a new survey, is a dubious point. Some surveyors would admit a difference of 1 link for every 3 chains in the sum of the courses; others only 1 link for every 10 chains. One writer puts the limit at 5 links for each station ; another at 25 links in a survey of 100 acres. But every practical surveyor soon learns how near to an equality his instrument and his skill will enable him to come in ordinary cases, and can therefore establish a standard for himself, by which he can judge whether the difference, in any survey of his own, is probably the result of an error, or only of his customary degree of inaccuracy, two things to be very carefully distinguished.*
244. Application to supplying Omissions. Any two omissions in the field-notes can be supplied by a proper use of the method of latitudes and departures ; as will be explained in Chapter V, which treats of "Obstacles to Measurement," under which head this

[^23]subject most appropriately belongs. But a knowledge of the fact that any two omissions can be supplied, should not lead the joung surveyor to be negligent in making erery possible measurement, since an omission renders it necessary to assume all the notes taken to be correct, the means of testing them no longer existing.
245. Balancing a Survey. The subsequent applications of this method require the surrey to be preriously balanced. This operation consists in correcting the latitudes and departures of the courses, so that their sums shall be equal, and thus "balance." This is usually done by distributing the differences of the sums among the courses in proportion to their length; saring, as the sum of the lengths of all the courses is to the whole difference of the latitude, so is the length of each course to the correction of its latitude. A similar proportion corrects the depaitures.*

It is not often necessary to make the exact proportion, as the correction can usually be made, with sufficient accuracy, by noting how much per chain it should be, and correcting accordingly.

In the example giren below, the differences hare purposely been made considerable. The corrected latitudes and departures hare been here inserted in four additional columns, but in practice they should be written in red ink orer the original latitudes and departures, and the latter crossed out with red ink.

|  | bearings. |  | latitcdes. |  | departires. |  | CORrected |  | CORRECTED departtres. |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | N. + | s. - | E. + | W. - | N. | S. - | E. + | $\pi$. |
| 1 | N. $52^{\circ} \mathrm{E}$. | $10 \cdot 63$ | 6.54 |  | $8 \cdot 38$ |  | 6.58 |  | $8 \cdot 34$ |  |
| 2 | S. $293^{\circ} \mathrm{E}$. | $4 \cdot 10$ |  | $3 \cdot 56$ |  |  |  | $3 \cdot 55$ | 2.01 |  |
| 4 | N. $611^{\text {S }}$ W. | \%-13 | $3 \cdot 46$ |  |  | $\begin{aligned} & 4.05 \\ & 6.24 \end{aligned}$ | $3 \cdot 48$ |  |  | $\begin{aligned} & 4 \cdot 08 \\ & 6 \cdot 27 \end{aligned}$ |
|  |  | 29:55 | $10 \cdot 00$ | $10 \cdot 10$ | $10 \cdot 41$ | $10 \cdot 29$ | 10.06 | 10.06 | 10.35 | $10 \cdot 35$ |

The corrections are made by the following proportions; the nearest whole numbers being taken :

[^24]| For the Latitudes. | For the Dcparturcs. |  |
| :---: | :---: | :---: |
| $29 \cdot 55: 10 \cdot 63:: 10: 4$ | $29 \cdot 55: 10 \cdot 63:: 12: 4$ |  |
| $29 \cdot 55:$ | $4 \cdot 10: 10: 1$ |  |

This rule is not always to be strictly followed. If one line of a survey has been measured over very uneven and rough ground, or if its bearing has been taken with an indistinct sight, while the other lines have been measured over level and clear ground, it is probable that most of the error has occurred on that line, and the correction should be chiefly made on its latitude and departure.

If a slight change of the bearing of a long course will favor the balancing, it should be so changed, since the compass is much more subject to error than the chain. So, too, if shortening any doubtful line will favor the balancing, it should be done, since distances are generally measured too long.
246. Application to Platting. Rule three columns; one for stations, the next for total latitudes, and the third for total departures. Fill the last two columns by beginning at any convenient station (the extreme east or west is best) and adding up (algebraically) the latitudes of the following stations, noticing that the south latitudes are subtractive. Do the same for the departures, observing that the westerly ones are also subtractive.

Taking the example given in Art. 243, and beginning with station 1 , the following will be the results:

| stations. | total latitudes from station 1. | total departures from station 1. |
| :---: | :---: | :---: |
| 1 | $0 \cdot 00$ | $0 \cdot 00$ |
| 2 | $+2 \cdot 21 \mathrm{~N}$. | + 1.55 E . |
| 3 | +2.36 N . | + 2.83 E . |
| 4 | +1.15 N . | + 4.69 E. |
| 5 | -1.78 S . | +2.69 E. |
| 1 | $0 \cdot 00$ | $0 \cdot 00$ |

It will be seen that the work proves itself, by the total latitudes and departures for station 1, again coming out equal to zero.

To use this table, draw a meridian through the point taken for
station 1, as in Fig. 178. Set off, upward from this, along the meridian, the latitude, 221 links, to $A$, and from $A$, to the right perpendicularly, set off the departure, 155 links.* This gires the point 2. Join 1....2. From 1 again, set off, upward, 236 links, to $B$, and from $B$, to the right, perpendicularly, set off 283 links, which will fix the point 3. Join $2 . . .3$; and so proceed, setting off north latitudes along the meridian upward, and south latitudes along it downward ; east departures perpendicularly to the right, and west departures perpendicularly to the left.

The adrantages of this method are its rapidity, ease, and accuracy ; the impossibility of any error in platting any one course affecting the following points ; and the certainty of the plat "coming together," if the latitudes and departures hare been "balanced."

## CALCULATING THE CONTENT.

247. Methods. When a field has been platted, by* whaterer method it may have been surreyed, its content can be obtained from its plat by dividing it up into triangles, and measuring on the plat their bases and perpendiculars ; or by any of the other means explained in Chapter II.

But these are only approximate methods, their degree of accu-

[^25]racy depending on the largeness of scale of the plat and the skill of the draughtsman. The invaluable method of latitudes and departures gives another means, perfectly accurate, and not requiring the previous preparation of a plat. It is sometimes called the rectangular, or the Pennsylvania, or Rittenhouse's method of calculation.*
248. Definitions. Imagine a meridian line to pass through the extreme east or west corner of a field. According to the definitions established in Art. 239 (and here recapitulated for convenience of reference), the perpendicular distance of each station from that meridian is the Longitude of that station ; additive, or plus, if east; subtractive, or minus, if west. The distance of the middle of any line, such as the side of the field, from the meridian; is called the longitude of that side. $\dagger$ The difference of the longitudes of the two ends of a line is called the Departure of that line. The difference of the latitudes of the two ends of a line is called the Latitude of the line.
249. Longitudes. To give more definiteness to the development of this subject, the figure in the margin will be referred to, and may be considered to represent any space inclosed by straight lines.

Fig. 179.


Let N S be the meridian passing through the extreme westerly station of the field ABCDE.

[^26]From the middle and ends of each side draw perpendiculars to the meridian. These perpendiculars will be the longitudes and departures of the respective sides. The longitude, FG, of the first course, A B , is evidently equal to half its departure, HB . The longitude, JK, of the second course, B C, is equal to JL $+\mathrm{LM}+\mathrm{MK}$, or equal to the longitude of the preceding course, plus half its departure, plus half the departure of the course itself. The longitude, Y Z, of some other course, as E A, taken anywhere, is equal to $\mathrm{WX}-\mathrm{VX}-\mathrm{UV}$, or equal to the longitude of the preceding course, minus half its departure, minus half the departure of the course itself-i. e., equal to the algebraic sum of these three parts, remembering that westerly departures are negative, and therefore to be subtracted when the directions are to make an algebraic addition.

To avoid fractions it will be better to double each of the preceding expressions. We shall then have a

## General Rele for finding Dotble Longitcdes.

The double longitude of the FIRST COURSE is equal to its departure.

The double longitude of the SECOND COTRSE is equal to the double longitude of the first course, plus the departure of that course, plus the departure of the second course.

The double longitude of the THIRD Cocrse is equal to the double longitude of the second course, plus the departure of that course, plus the departure of the course itself.

The double longitude of ANY course is equal to the double longitude of the preceding course, plus the departure of that course, plus the cleparture of the course itself.*

The double longitude of the last course (as well as of the first) is equal to its departure. Its "coming out" so, when obtaincd by the above rule, proves the accuracy of the calculation of all the preceding double longitudes.
250. Areas. We will now proceed to find the area or content of a field, by means of the "double longitudes" of its sides, Thich

[^27]can be readily obtained by the preceding rule, whatever their number.
251. Beginning with a three-sided field, A B C in the figure, draw a meridian through A, and draw perpendiculars to it as in the last figure. It is plain that its content is equal to the difference of the areas of the trapezoid DBCE, and of the triangles ABD and ACE.

The area of the triangle ABD is equal to the product of AD by half of DB , or to the product of AD by FG ; i. e., equal to the product of the latitude of the first course by its longitude.

The area of the trapezoid DBCE is equal to the product of DE by half the sum of DB and CE, or by HJ ; i. e., to the product of the latitude of the sec-

Fig. 180.
 ond course by its longitude.

The area of the triangle ACE is equal to the product of AE by half EC , or by KL; i. e., to the product of the latitude of the third course by its longitude.

Fig. 181.


Calling the products in which the latitude was north, North Products, and the products in which the latitude was south, South Products, we shall find the area of the trapezoid to be a south product, and the areas of the triangles to be north products. The difference of the north products and the south products is therefore the desired area of the three-sided field ABC.

Using the double longitudes (in order to avoid fractions) in each of the preceding products, their difference will be the double area of the triangle ABC.
252. Taking now a four-sided field,

ABCD in the figure, and drawing a meridian and longitudes as
before, it is seen, on inspection, that its area would be obtained by taking the two triangles, $\mathrm{ABE}, \mathrm{ADG}$, from the figure EBCDGE, or from the sum of the two trapezoids EBCF and FCDG.

The area of the triangle AEB will be found, as in the last article, to be equal to the product of the latitude of the first course by its longitude. The product will be North.

The area of the trapezoid EBCF will be found to equal the latitude of the second course by its longitude. The product will be South.

The area of the trapezoid FCD G will be found to equal the product of the latitude of the third course by its longitude. The product will be South.

The area of the triangle A D G will be found to equal the product of the latitude of the fourth course by its longitude. The product will be North.

The difference of the north and south products will therefore be the desired area of the four-sided field A B CD.

Using the double longitude as before, in each of the preceding products, their difference will be double the area of the field.
253. Whatever the number or directions of the sides of a field, or of any space inclosed by straight lines, its area will always be equal to half of the difference of the north and south products arising from multiplying together the latitude and double longitude of each course or side.

We have, therefore, the following

## General Rele for finding Areas.

1. Prepare ten columns, headed as in the example below, and in the first three write the stations, bearings, and distances.
2. Find the latitudes and departures of each course, by the traverse-table, as directed in Art. 242, placing them in the four following columns.
3. Balance them, as in Art. 245, correcting them in red ink.
4. Find the double longitudes, as in Avt. 249, with reference
to a meridian passing through the extreme east or west station, and place them in the eighth column.
5. Multiply the double longitude of each course by the corrected latitude of that course, placing the north products in the ninth column, and the south products in the tenth column.
6. Add up the last two columns, subtract the smaller sum from the larger, and divide the difference by two. The quotient will be the content desired.
7. To find the most easterly or westerly station of a survey, without a plat, it is best to make a rough hand-sketch of the survey, drawing the lines in an approximation to their true directions, by drawing a north and south, and east and west lines, and considering the bearings as fractional parts of a right angle, or $90^{\circ}$; a course N. $45^{\circ}$ E., for example, being drawn about half-way between a north and an east direction ; a course N. $28^{\circ} \mathrm{W}$. being not quite one third of the way around from north to west ; and so on, drawing them of approximately true proportional lengths.
8. Example 1, given below, refers to the five-sided field, of which a plat is given in Fig. 161, and the latitudes and departures of which were calculated in Art. 243. Station 1 is the most westerly station, and the meridian will be supposed to pass through it. The double longitudes are best found by a continual addition and subtraction, as in the margin, where they are marked D. L. The double longitude of the last course comes out equal to its departure, thus proving the work.

The double longitudes being thus obtained, are multiplied by the corresponding latitudes, and the content of the field ob-

| $\begin{aligned} & \text { STA- } \\ & \text { TIONS. } \end{aligned}$ |  |
| :---: | :---: |
| 1 | $\begin{aligned} & +1 \cdot 55 \text { D. L. } \\ & +1 \cdot 55 \\ & +1 \cdot 28 \end{aligned}$ |
| 2 | $\begin{aligned} & +4 \cdot 38 \text { D. L. } \\ & +1 \cdot 28 \\ & +1.86 \end{aligned}$ |
| 3 | $\begin{aligned} & +7.52 \mathrm{D} . \mathrm{L} . \\ & +1.86 \\ & -2.00 \end{aligned}$ |
| 4 | $\begin{aligned} & +7.38 \mathrm{D} . \mathrm{L} . \\ & -2.00 \\ & -2.69 \end{aligned}$ |
| 5 | + 2.69 D. L. | tained as directed in the General Rule.

This example may serve as a pattern for the most compact manner of arranging the work.

|  | bearings. |  | latitudes. |  | departures. |  | $\begin{aligned} & \text { DOUBLE } \\ & \text { LONGI- } \\ & \text { TUDES. } \end{aligned}$ |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | N. + | S. - | E. + | W.- |  | docble areas. |  |
| $\begin{aligned} & 1 \\ & 2 \\ & 3 \\ & 4 \\ & 4 \end{aligned}$ | N. $35^{\circ} \mathrm{E}$.N. $83 \frac{1}{1}^{\circ} \mathrm{E}$.S.S. $57^{2}$S.S.N. $561^{\circ}{ }^{\circ} \mathrm{W}$W. | $2 \cdot 70$ | $\begin{array}{r} 2 \cdot 21 \\ .15 \end{array}$ |  | 1.55 |  | $\begin{aligned} & \hline+1.55 \\ & +4.38 \\ & +7.52 \\ & +7.38 \\ & +2.69 \end{aligned}$ | $3 \cdot 4255$ |  |
|  |  | $1 \cdot 29$ |  |  | 1.28 |  |  | $0 \cdot 6570$ |  |
|  |  | $2 \cdot 22$ |  | $1 \cdot 21$ | $1 \cdot 86$ |  |  |  | $9 \cdot 0992$ |
|  |  | $3 \cdot 55$ |  | $2 \cdot 93$ |  | $2 \cdot 00$ |  |  | $21 \cdot 6234$ |
|  |  | $3 \cdot 23$ | 1.78 |  |  | $2 \cdot 69$ |  | 4.7882 |  |
|  |  |  | $\overline{4 \cdot 14}$ | $4 \cdot 14$ | $4 \cdot 69$ | 4•69 |  | 8.8707 | $\begin{array}{r} 30.7226 \\ 8.8707 \end{array}$ |
|  | Cont | $n t=$ | A. 0 | R. 15 |  |  |  |  | $21 \cdot 8519$ |
|  |  |  |  |  |  |  | square | chains, | 10.9259 |


| $\begin{gathered} \text { sTA- } \\ \text { TIONS. } \end{gathered}$ |  |
| :---: | :---: |
| 4 | $\begin{aligned} & -2.00 \mathrm{D} . \mathrm{L} . \\ & =2.00 \\ & -2.69 \end{aligned}$ |
| 5 | $\begin{aligned} & -6.69 \text { D. L. } \\ & -2.69 \\ & +1.55 \end{aligned}$ |
| 1 | $\begin{aligned} & -7.83 \text { D. L. } \\ & +1.55 \\ & +1.28 \end{aligned}$ |
| 2 | $\begin{aligned} & -5.00 \text { D. L. } \\ & +1.28 \\ & +1.86 \end{aligned}$ |
| 3 | $-1.86$ |

256. The meridian might equally well have been supposed to pass through the most easterly station, 4 in the figure. The double longitudes could then have been calculated as in the margin. They will, of course, be all west, or minus. The products being then calculated, the sum of the north products will be found to be 29.9625 , and of the south products $8 \cdot 1106$, and their difference to be 21.8519 , the same result as before.
257. A number of examples, with and without answers, will now be given as exercises for the student, who should plat them by some of the methods given in the chapter on platting, using each of them at least once. He should then calculate their content by the method just given, and check it, by also calculating the area of the plat by some of the geometrical or instrumental methods given in Chapter
 I; for no single calculation is
ever reliable. All the examples (except the last) are from the author's actual surveys.

Example 2, given below, is also fully worked out, as another pattern for the student, who need have no difficulty with any possible case if he strictly follows the directions which have been given. The plat is on a scale of 2 chains to 1 inch ( $=1: 1584$ ).

|  | bearings. |  | latitudes. |  | departures. |  | double LONGItUDES | double areas. |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | N. + | S. - | + | W.- |  | N. + | S. - |
| 1 | N. $121^{\circ} \mathrm{E}$. | $2 \cdot 81$ | $2 \cdot 75$ |  | $\cdot 60$ |  | $+6.56$ | $18 \cdot 0400$ |  |
| 2 | N. $76{ }^{\circ} \mathrm{W}$. | $3 \cdot 20$ | $\cdot 77$ |  |  | $3 \cdot 11$ | +4.05 | $3 \cdot 1185$ |  |
| 3 | S. $241^{\circ}{ }^{\circ} \mathrm{W}$. | $1 \cdot 14$ |  | 1.04 |  | $\cdot 47$ | + 47 |  | -4888 |
| 4 | S. $48^{\circ} \mathrm{E}$. | 1.53 |  | $1 \cdot 02$ | $1 \cdot 14$ |  | $+1 \cdot 14$ |  | $1 \cdot 1628$ |
| 5 | S. $121^{\circ} \mathrm{E}$. | $1 \cdot 12$ |  | 1.09 | . 24 |  | +2.52 |  | 2.7468 |
| 6 | S. $77^{\circ} \mathrm{E}$. | $1 \cdot 64$ |  | $\cdot 37$ | $1 \cdot 60$ |  | +4.36 |  | $1 \cdot 6132$ |
|  |  |  | 3.52 | $3 \cdot 52$ | $3 \cdot 58$ | 3.58 |  | $\begin{array}{r} 21 \cdot 1585 \\ 6.0116 \end{array}$ | $6 \cdot 0116$ |
|  | Cont | t $=$ | A. 3 | 2. 1 P |  |  |  | $15 \cdot 1469$ |  |
|  |  |  |  |  |  | quare | chains, | $7 \cdot 5734$ |  |

Example 3.

| stations. | bearings. | distances. |
| :---: | :---: | :---: |
| 1 | N. $52^{\circ}$ E. | 10.64 |
| 2 | S. $293^{\circ} \mathrm{E}$. | $4 \cdot 09$ |
| 3 | S. $31 \frac{3}{4}^{\circ} \mathrm{W}$. | $7 \cdot 68$ |
| 4 | N. $61{ }^{\circ} \mathrm{W}$. | $7 \cdot 24$ |

Ans. 4 A. 3 R. 28 P.

Example 5.

| stations. | bearings. | distances |
| :---: | :---: | :---: |
| 1 | N. $34{ }^{10} \mathrm{E}$. | 2.73 |
| 2 | N. $85^{\circ} \mathrm{E}$. | $1 \cdot 28$ |
| 3 | S. $56 \frac{3}{4}^{\circ} \mathrm{E}$. | $2 \cdot 20$ |
| 4 | S. $34 \frac{1}{4}^{\circ} \mathrm{W}$. | $3 \cdot 53$ |
| 5 | N. $56 \frac{1}{2}^{\circ} \mathrm{W}$. | 3.20 |

Ans. 1 A. 0 R. 14 P.

Example 4.

| stations. | bearings. |  | distances. |
| :---: | :---: | :---: | :---: |
| 1 | S. $21^{\circ}$ | W. | 12.41 |
| 2 | N. $831^{\circ}$ | E. | 5.86 |
| 3 | N. $12^{\circ}$ | E. | 8.25 |
| 4 | N. $47^{\circ}$ | W. | 4.24 |

Ans. 4 A. 2 R. 37 P.

Example 6.

| stations. | bearings. | distances. |
| :---: | :---: | :---: |
| 1 | N. $35^{\circ}$ E. | $6 \cdot 49$ |
| 2 | S. $564^{\circ} \mathrm{E}$. | $14 \cdot 15$ |
| 3 | S. $34^{\text {² }} \mathrm{W}$. | $5 \cdot 10$ |
| 4 | N. $56^{\circ} \mathrm{W}$. | $5 \cdot 84$ |
| 5 |  | $2 \cdot 52$ |
| 6 | N. $484^{\circ} \mathrm{W}$. | $8 \cdot 73$ |

Example 7.

| stations. | bearings. | distances. |
| :---: | :---: | :---: |
| 1 | S. $211^{\circ}{ }^{\circ} \mathrm{W}$. | 17.62 |
| 2 | S. $34^{\circ} \mathrm{W}$. | 10.00 |
| 3 | N. $\check{5} 6^{\circ} \mathrm{W}$. | $14 \cdot 15$ |
| 4 | N. $34^{\circ} \mathrm{E}$. | $9 \cdot 76$ |
| 5 | N. $67^{\circ}$ E. | $2 \cdot 30$ |
| 6 | N. $23^{\circ} \mathrm{E}$. | $7 \cdot 03$ |
| 7 | N. $18 \frac{1}{2}^{\circ} \mathrm{E}$. | $4 \cdot 43$ |
| 8 | S. $76 \frac{1}{2}^{\circ} \mathrm{E}$. | $12 \cdot 41$ |

Example 9.

| stations. | bearings. | distances. |
| :---: | :---: | :---: |
| 1 | S. $57{ }^{\circ} \mathrm{E}$. | 5.77 |
| 2 | S. $36 \frac{1}{0}^{\circ} \mathrm{W}$. | $2 \cdot 25$ |
| 3 | S. $399^{\circ} \mathrm{W}$. | 1.00 |
| 4 | S. $701^{\circ} \mathrm{W}$. | 1.04 |
| 5 | N. $683^{\frac{3}{\circ}}{ }^{\circ} \mathrm{W}$. | $1 \cdot 23$ |
| 6 | N. $56^{\circ} \mathrm{W}$. | $2 \cdot 19$ |
| 7 | N. $331^{\circ} \mathrm{E}$. | 1.05 |
| 8 | N. $561^{\circ}{ }^{\circ} \mathrm{W}$. | 1.54 |
| 9 | N. $33 \frac{1}{2}^{\circ} \mathrm{E}$. | $3 \cdot 18$ |

Ans. 2 A. 0 R. 32 P.

Example 11.

| stations. | bearings. | distances. |
| :---: | :---: | :---: |
| 1 | N. $183^{\circ} \mathrm{E}$. | $1 \cdot 93$ |
| 2 | N. $9^{\circ} \mathrm{W}$. | $1 \cdot 29$ |
| 3 | N. $14^{\circ} \mathrm{W}$. | $2 \cdot 71$ |
| 4 | N. $74^{\circ} \mathrm{E}$. | $0 \cdot 95$ |
| 5 | S. $481^{\circ} \mathrm{E}$. | 1.059 |
| 6 | S. $14 \frac{1}{\circ}^{\circ} \mathrm{E}$. | $1 \cdot 14$ |
| 7 | S. $19 \frac{1}{3}^{\circ} \mathrm{E}$. | $2 \cdot 15$ |
| 8 | S. $23 \frac{1}{3}^{\circ} \mathrm{W}$. | $1 \cdot 22$ |
| 9 | S. $5^{\circ} \mathrm{W}$. | $1 \cdot 40$ |
| 10 | S. $30^{\circ} \mathrm{W}$. | $1 \cdot 02$ |
| 11 | S. $811^{\circ} \mathrm{W}$. | $0 \cdot 69$ |
| 12 | N. $32 \frac{1}{3}^{\circ} \mathrm{W}$. | $1 \cdot 98$ |

Example 8.

| stations. | bearings. | distances. |
| :---: | :---: | :---: |
| 1 | S. $65 \frac{1}{10}^{\circ} \mathrm{E}$. | $4 \cdot 98$ |
| 2 | S. $58^{\circ} \mathrm{E}$. | 8.56 |
| 3 | S. $141^{\circ} \mathrm{W}$. | $20 \cdot 69$ |
| 4 | S. $47^{\circ} \mathrm{W}$. | $0 \cdot 60$ |
| 5 | S. $577^{\circ} \mathrm{W}$. | $8 \cdot 98$ |
| 6 | N. $56^{\circ} \mathrm{W}$. | $12 \cdot 90$ |
| 7 | N. $34^{\circ} \mathrm{E}$. | $10 \cdot 00$ |
| 8 | N. $214^{\circ} \mathrm{E}$. | $17 \cdot 62$ |

Example 10.

| STA- TIONS. | bearings. | distaices. |
| :---: | :---: | :---: |
| 1 | N. $63^{\circ} 51^{\prime}$ W. | 6.91 |
| 2 | N. $63^{\circ} 44^{\prime} \mathrm{W}$. | $7 \cdot 26$ |
| 3 | N. $69^{\circ} 35^{\prime} \mathrm{W}$. | $3 \cdot 34$ |
| $\pm$ | N. $77^{\circ} 50^{\prime} \mathrm{W}$. | 6.54 |
| 5 | N. $31^{\circ} 24^{\prime} \mathrm{E}$. | 14.38 |
| 6 | N. $31^{\circ} 18^{\prime} \mathrm{E}$. | 16.81 |
| 7 | S. $68^{\circ} 55^{\prime} \mathrm{E}$. | $13 \cdot 64$ |
| 8 | S. $68^{\circ} 42^{\prime} \mathrm{E}$. | 11.54 |
| 9 | S. $33^{\circ} 45^{\prime} \mathrm{W}$. | 31.55 |

Ans. 74 acres.

Example 12.

| stations. | bearings. | distances |
| :---: | :---: | :---: |
| 1 | N. $723^{\circ} \mathrm{E}$. | 0.88 |
| 2 | S. $20 \frac{1}{3}^{\circ} \mathrm{E}$. | $0 \cdot 22$ |
| 3 | S. $63^{\circ} \mathrm{E}$. | $0 \cdot 75$ |
| 4 | N. $51^{\circ} \mathrm{E}$. | $2 \cdot 35$ |
| 5 | N. $44^{\circ} \mathrm{E}$. | $1 \cdot 10$ |
| 6 | N. $251^{\circ} \mathrm{T}$. | $1 \cdot 96$ |
| 7 | N. $8 \frac{1}{}{ }^{\circ} \mathrm{T}$. | $1 \cdot 05$ |
| 8 | S. $29^{\circ} \mathrm{T}$. | $1 \cdot 63$ |
| 9 | N. $711^{\circ}{ }^{\circ} \mathrm{W}$. | $0 \cdot 81$ |
| 10 | N. $13{ }^{\frac{1}{2}}{ }^{\circ} \mathrm{T}$. | $1 \cdot 17$ |
| 11 | N. $63^{\circ} \mathrm{W}$. | $1 \cdot 28$ |
| 12 | West. | 1.68 |
| 13 | N. $49^{\circ} \mathrm{W}$. | $0 \cdot 80$ |
| 14 | S. $191_{3}^{\circ} \mathrm{E}$. | $6 \cdot 0$ |

Example 13. A farm is described in an old deed as bounded thus: Beginning at a pile of stones, and running thence trents-
seven chains and seventy links southeasterly sixty-six and a half degrees to a white-oak stump ; thence eleven chains and sixteen links northeasterly twenty and a half degrees to a hickory-tree ; thence two chains and thirty-five links northeasterly thirty-six degrees to the southeasterly corner of the homestead; thence nineteen chains and thirty-two links northeasterly twenty-six degrees to a stone set in the ground ; thence twenty-eight chains and eighty links northwesterly sixty-six degrees to a pine-stump; thence thirty-

Fig. 183.
 three chains and nineteen links southwesterly twenty-two degrees to the place of beginning, containing ninety-two acres, be the same more or less. Required the exact content.
258. Mascheroni's Theorem. The surface of any polygon is equal to half the sum of the products of its Fig. 184. sides (omitting any one side) taken two and
 two, into the sines of the angles which those sides make with each other.

Thus, take any polygon, such as the fivesided one in the figure. Express the angle which the directions of any two sides, as AB , $C D$, make with each other, thus ( $\mathrm{AB} \wedge \mathrm{CD}$ ). Then will the content of that polygon be, as below :

$$
\begin{aligned}
= & \frac{1}{2}[A B \cdot B C \cdot \sin (A B \wedge B C)+A B \cdot C D \cdot \sin (A B \wedge C D) \\
& +A B \cdot D E \cdot \sin (A B \wedge D E)+B C \cdot C D \cdot \sin (B C \wedge C D) \\
& +B C \cdot D E \cdot \sin (B C \wedge D E)+C D \cdot D E \cdot \sin (C D \wedge D E)]
\end{aligned}
$$

The demonstration consists merely in dividing the polygon into
triangles by lines drawn from any angle (as A) ; then expressing the area of each triangle by half the product of its base and the perpendicular let fall upon it from the above-named angle; and finally separating the perpendicular into parts which can each be expressed by the product of some one side into the sine of the angle made by it with another side.

Fig. 185.
 The sum of these triangles equals the polygon.

The expressions are simplified by dividing the proposed polygon into two parts by a diagonal, and computing the area of each part separately, making the diagonal the side omitted.

## A New Method of calculating: Areas.

259. In Fig. 185, let the total latitudes (Art. 246) of the stations 1, 2, 3, and 4 be represented by $l_{1}, l_{2}, l_{3}$, and $l_{\ddagger}$, respectively.
Let the departures of each course separately be represented by $d_{1}, d_{2}, d_{3}$, and $d_{4}$, respectively.

The double area of A B 23

$$
\begin{aligned}
& =\mathrm{A} \mathrm{~B}(\mathrm{~A} 2+\mathrm{B} 3) \\
& =\left(l_{2}-l_{3}\right)\left(d_{1}+d_{1}+d_{2}\right) \\
& =l_{2} d_{1}+l_{2} d_{1}+l_{2} d_{2}-l_{3} d_{1}-l_{3} d_{1}-l_{3} d_{2 .} .
\end{aligned}
$$

The double area of C B34

$$
\begin{align*}
& =\mathrm{CB}(\mathrm{~B} 3+\mathrm{C} 4) \\
& =\left(l_{3}+l_{4}\right)\left(d_{4}+d_{4}+d_{3}\right)  \tag{3.}\\
& =l_{3} d_{4}+l_{3} d_{4}+l_{3} d_{3}+l_{4} d_{4}+l_{4} d_{4}+l_{4} d_{3} . \tag{2.}
\end{align*}
$$

The double area of $12 \mathrm{~A}=\mathrm{A} 1(\mathrm{~A} 2)=l_{2} d_{1}$.
The double area of $14 \mathrm{C}=\mathrm{C} 1(\mathrm{C} 4)=l_{1} d_{4}$.
Now, the double area of the figure 1234 is equal to the sum of [1] and [2] - the sum of [3] and [4].

Combining and reducing, we hare :
Double area of $1234=l_{2}\left(d_{1}+d_{2}\right)+l_{3}\left(d_{4}+d_{4}+d_{3}-d_{1}-d_{3}\right.$ $\left.-d_{2}\right)+l_{4}\left(d_{3}+d_{4}\right)$.

Noting that $d_{4}+d_{3}=d_{1}+d_{2}$, we have,
Double area of $1234=l_{2}\left(d_{1}+d_{2}\right)+l_{3}\left(d_{2}-d_{3}\right)+l_{4}\left(d_{3}+d_{4}\right)$.
Putting this in the form of a rule, we have: Multiply the total latitude of each station by the algebraic sum of the departures of the two adjacent courses. One half of the algebraic sum of the products will be the area.

As an exercise for the student, let him find, by the above method, an expression for the area of figures having five and six sides.

The following example, worked out by the method of double longitudes (on page 158), and below, by the new method, will show the difference between the two methods:

|  | bearings. | 風寅 | latitudes. |  | departures. |  | totalLati-LUDES tudes. | $\begin{array}{\|c\|} \hline \text { ADJA- } \\ \text { CENT } \\ \text { DEPART- } \\ \text { URES. } \end{array}$ | ¢ $\begin{gathered}\text { double } \\ \text { Areas. }\end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | N. + | S. - | E. + | w.- |  |  |  |
| 1 | N. $35^{\circ} \mathrm{E}$. | 2.70 | $2 \cdot 21$ |  | 1.55 |  |  |  |  |
| 2 | N. $831^{\circ}{ }^{\circ} \mathrm{E}$. | $1 \cdot 29$ | $\cdot 15$ |  | $1 \cdot 28$ |  | $2 \cdot 21$ | $2 \cdot 83$ | 6.2543 |
| 3 | S. $57^{\circ} \mathrm{E}$. | $2 \cdot 22$ |  | 121 | $1 \cdot 86$ |  | $2 \cdot 36$ | $3 \cdot 14$ | $7 \cdot 4104$ |
| 4 | S. $344^{\circ} \mathrm{W}$. | $3 \cdot 55$ |  | $2 \cdot 93$ |  | 2.00 | $1 \cdot 15$ | -0.14 | $-0 \cdot 1610$ |
| 5 | N. $56 \frac{1}{3}^{\circ} \mathrm{W}$. | $3 \cdot 23$ | $1 \cdot 78$ |  |  | $2 \cdot 69$ | $-1.78$ | $-4.69$ | 8.3482 |
|  |  |  | 4•14 | $4 \cdot 14$ | $\stackrel{ }{4 \cdot 69}$ | 4.69 |  |  | 21.8519 |
|  |  |  |  |  |  |  | square | chains, | 10.9259 |

In computing the total latitudes, if the total latitude of the last station equals the latitude of the last course with sign changed, the total latitudes may be considered correct.

The station through which the meridian of the survey is supposed to pass, and from which the total latitude is reckoned, will have no latitude, and hence the product of its latitude and adjacent departures will be zero. There will therefore be one less product than there are stations.

Any station may be taken as the starting-point.
To verify the area obtained in any case, calculate a second time, using a different station as the starting-point.

This method was first published by J. Woodbridge Davis, C. E., Ph. D., in Van Nostrand's "Engineering Magazine," for April, 1879, where a general discussion of the method is given.

## THE DECLINATION OF THE MAGNETIC NEEDLE.

260. Definitions. The magnetic meridian is the direction indicated by the magnetic needle. The true meridian is a true north and south line, which, if produced, would opass through
Fig. 186.
 the poles of the earth. The declination of the needle is the angle which one of these lines makes with the other.

In the figure, if $\mathrm{N} S$ represent the direction of the true meridian, and $\mathrm{N}^{\prime} \mathrm{S}^{\prime}$ the direction of the magnetic meridian at any place, then is the angle $\mathrm{N}^{\top} \mathrm{A} \mathrm{N}^{\prime}$ the declination of the needle at that place.
261. Direction of the Needle. The directions of these

傻 $\mathbf{S}^{\prime}$ two meridians do not generally coincide, but the needle in most places points to the east or to the west of the true north, more or less according to the locality. Observations of the amount and the direction of this declination have been made in nearly all parts of the world. In the United States the declination in the Eastern States is westerly, and in the Western States is easterly, as will be given in detail, after the methods for determining the true meridian, and consequently the declinations, at any place have been explained.

## To determine the True Meridian.

262. By Equal Shadows of the Sun. On the south side of any level surface erect an upright staff, shown in horizontal projection at S . Two or three hours before noon, mark the extremity, A, of its shadow. Describe an arc of a circle with $S$, the foot of the staff, for center, and SA , the distance to the extremity of the shadow,
 for radius. About as many hours after noon as it had been before noon when the first mark was made, watch for the moment when the end of the shadow
touches the are at another point, B. Bisect the are AB at N. Draw S N, and it will be the true meridian, or north and south line required.

For greater accuracy, describe several arcs beforehand, mark the points in which each of them is touched by the shadow, bisect each, and adopt the average of all. The shadow will be better defined if a piece of tin with a hole through it be placed at the top of the staff, as a bright spot will thus be substituted for the less definite shadow. Nor need the staff be vertical, if from its summit a plumb-line be dropped to the ground, and the point which this strikes be adopted as the center of the arcs.

This method is a very good approximation, though perfectly correct only at the time of the solstices, about June 21st and December 22d. It was employed by the Romans in laying out cities.

To get the declination, set the compass at one end of the true meridian line thus obtained, sight to the other end of it, and take the bearing as of any ordinary line. The number of degrees in the reading will be the desired declination of the needle.
263. By the North Star, when in the Meridian. The north star, or pole star (called by astronomers Alpha Ursce Minoris, or Polaris), is not situated precisely at the north pole of the heavens. If it were, the meridian could be at once determined by sighting to it, or placing the eye at some distance behind a plumb-line so that this line should hide the star. But the north star is about $11_{2}^{\circ}$ from the pole. Twice in twenty-four
 hours, however (more precisely, twentythree hours fifty-six minutes), it is in the meridian, being then exactly above or below the pole, as at A and C in the figure. To know when it is so, is rendered easy by the aid of another star, easily identified, which at these times is almost exactly above or below the north star-i. e., situated in the same vertical plane. If, then, we watch for the moment at which a suspended plumbline will cover both these stars, they will then be in the meridian.

The other star is in the well-known constellation of the Great Bear, called also the Plow, or the Dipper, or Charles's Wain.

Fig. 189.
 Two of its five bright stars (the right-hand ones in Fig. 189) are known as the "Pointers," from their pointing near to the north star, thus assisting in finding it. The star in the tail or handle, nearest to the four which form a quadrilateral, is the star which comes to the meridian at the same time with the north star, twice in twenty-four hours, as in Fig. 189 or 190. It is known as Alioth, or Epsilon Ursce Majoris.*

To determine the meridian by this method, suspend a long plumb-line from some elerated point, such as a stick projecting from the highest window of a house suitably situated. The plumbbob may pass into a pail of water to lessen its ribrations. South of this set up the compass, at such a distance from the plumb-line that neither of the stars will be seen abore its highest point-i. e., in latitudes of $40^{\circ}$ or $50^{\circ}$, not quite as far from the plumb-line as it is long. Or, instead of a compass, place a board on two stakes, so as to form a sort of bench, running east and west, and on it place one of the compass-sights, or anything having a small hole in it to look through. As the time approaches for the north star to be on the meridian (as taken from the table given below) place the compass, or the sight, so that, looking through it, the plumb-line shall seem to cover or hide the north star. As the star mores one way, move the eye and sight the other way, so as to constantly keep the star behind the plumb-line. At last Alioth, too, will be

[^28]covered by the plumb-line. At that moment the eye and the plumb-line are (approximately) in the meridian. Fasten down the sight on the board till morning, or with the compass take the bearing at once, and the reading is the declination.

Instead of one plumb-line and a sight, two plumb-lines may be suspended at the end of a horizontal rod, turning on the top of a pole.

The line thus obtained points to the east of the true line when the north star is above Alioth, and vice versa. The north star is exactly in the meridian about twenty-five minutes after it has been in the same vertical plane with Alioth, and may be sighted to, after that interval of time, with perfect accuracy.

Another bright star, which is on the opposite side of the pole, and is known to astronomers as Gamma Cassiopeice, also comes on the meridian nearly at the same time as the north star, and will thus assist in determining its direction.
264. The time at which the north star passes the meridian above the pole, for every tenth day in the year, is given in the following table, in common clock-time.* The upper transit is the most convenient, since at the other transit Alioth is too high to be conveniently observed :

| E. | Monthe. | 1st Day. | 11th Day. | 21st Day. |
| :---: | :---: | :---: | :---: | :---: |
| - |  | н. м. | н. м. | н. м. |
| \% | Jannary | 630 P. м. | 551 Р. м. | 511 р. м. |
| $\approx$ | February | 428 " | 348 " | 309 " |
| . | March | 237 " | 158 " | 118 " |
| 莍 | April | 035 " | $1156 \mathrm{~A} . \mathrm{m}$. | $1116 \mathrm{~A} . \mathrm{M}$. |
| 2 | May | $1033 \mathrm{~A} . \mathrm{m}$. | 954 " | 915 " |
| 気 | June | 832 " | 752 " | 713 " |
|  | July | 634 " | 555 " | 515 " |
| \% | August. | 433 " | 353 " | 314 " |
| \% | September | 231 " | 152 " | 112 " |
| B | October. | 034 " | 1150 P. м. | 1111 p. M. |
| \% | November | 1028 р. м. | 948 " | 909 " |
|  | December | 830 " | 750 " | 711 " |

[^29]To find the time of the star's passage of the meridian for other days than those given in the table, take from it the time for the day most nearly preceding that desired, and subtract from this time four minutes for each day from the date of the day in the table to that of the desired day; or, more accurately, interpolate by saying : " $A s$ the number of days between those giren in the table is to the number of days from the next preceding day in the table to the desired day, so is the difference between the times giren in the table for the days next preceding and following the desired day to the time to be subtracted from that of the next preceding day."

The north star passes the meridian later every year. In 1890 it will pass the meridian about two minutes later than in 1885 ; in 1895 six minutes, and in 1900 ten minutes later than in 1885 , the year for which the preceding table has been calculated.

The times at which the north star passes the meridian below the pole in its lower transit can be found by adding eleven hours and fifty-eight minutes to the time of the upper transit, or by subtracting that interval from it.*
265. By the North Star at its Extreme Elongation. When the north star is at its greatest apparent angular distance east or west of the pole, as at B or D in Fig. 18S, it is said to be at its extreme eastern or extreme western elongation. If it be observed at either of these times, the direction of the meridian can be easily obtained
sion of the star, and from it (increased by twenty-four hours if necessary to render the subtraction possible) subtract the right ascension of the sun at mean noon, or the sidereal time at mean noon, for the giren day, as found in the "ephemeris of the sun" in the same almanac. From the remainder subtract the acceleration of sidereal on mean time corresponding to this remainder ( 3 m .56 s . for 24 hours), and the new remainder is the required mean solar time of the upper passage of the star across the meridian, in "astronomical" reckoning, the astronomical day beginning at noon of the common civil day of the same date.

* The north star, which is nor about $1^{\circ} 18^{\prime}$ from the pole, was $12^{\circ}$ distant from it when its place was first recorded. Its distance is now diminishing at the rate of about a third of a minute in a year, and will continue to do so till it approaches to within half a degree, when it will again recede. The brightest star in the northern hemisphere, Alpha Lyrce, will be the pole-star in about 12,000 rears, being then within about $5^{\circ}$ of the pole, though now more than $51^{\circ}$ distant from it.
from the observation. The great advantage of this method over the preceding is that then the star's motion apparently ceases for a short time.

MEAN TLME OF THE ELONGATIONS OF POLARIS FOR 18S5, LATITUDE $40^{\circ}$ NORTII.*

| date. | eastern elongation. | western elongation. |
| :---: | :---: | :---: |
| $\begin{array}{\|cc} \text { January } \\ \text { " } & 1885 \ldots \ldots \\ \hline \end{array}$ | $\begin{array}{ll} \text { н. м. } \\ 1235 \cdot 3 \text { р. м. } \\ 11 & 36 \cdot 1 \text { А. м. } \end{array}$ | н. м. <br> $1224 \cdot 6$ A. м. <br> $1129 \cdot 3$ р. м. |
|  | $\begin{array}{rrr} 10 & 29 \cdot 0 & \text { " } \\ 9 & 33 \cdot 7 & \text { " } \\ \hline \end{array}$ | $\begin{array}{rrr} 10 & 22 \cdot 2 \\ 9 & 27 \cdot 0 & \text { " } \end{array}$ |
|  | $\begin{array}{lll} \hline 8 & 38.5 & 6 \\ 7 & 43.4 & 6 \\ \hline \end{array}$ | $\begin{array}{lll} 8 & 31 \cdot 8 & \text { " } \\ 7 & 36.6 & \\ \hline \end{array}$ |
| April 1,    <br> "6 15, 6 $\cdots$ | $\begin{array}{lll} \hline 6 & 36.4 & 6 \\ 5 & 41.4 & 6 \\ \hline \end{array}$ | $\begin{array}{lll} 6 & 29 \cdot 7 & \text { " } \\ 5 & 34 \cdot 7 & \\ \hline \end{array}$ |
| $\begin{array}{\|ccc} \hline \text { May } & 1, & \text { " } \\ \text { "6 } & 15, & \\ \hline \end{array}$ | $\begin{array}{lll} 4 & 38 \cdot 6 & \text { " } \\ 3 & 43 \cdot 7 & \text { " } \end{array}$ | $\begin{array}{lll} \hline 431.8 & \text { " } \\ 3 & 36.9 & \\ \hline \end{array}$ |
|  | $\begin{array}{lll} 2 & 37 \cdot 1 & \text { " } \\ 1 & 42 \cdot 2 & 6 \\ \hline \end{array}$ | $\begin{array}{lll} 2 & 30 \cdot 3 & \text { " } \\ 1 & 35 \cdot 4 & \text { " } \\ \hline \end{array}$ |
| July 1, " $\ldots \ldots$. <br> " 15, " $\ldots \ldots$. | $\begin{array}{ll} 12 & 39 \cdot 6 \\ 11 & 44 \cdot 7 \\ \text { р. м. } \end{array}$ | $\begin{array}{ll} 12 & 32 \cdot 8 \quad \text { " } \\ 11 & 34 \cdot 0 \text { А. м. } \end{array}$ |
|  | $\begin{array}{rll} 10 & 38 \cdot 2 & \text { " } \\ 9 & 43 \cdot 3 & \end{array}$ | $\begin{array}{rrr} 10 & 27.5 & \text { " } \\ 9 & 32 \cdot 6 & \text { " } \end{array}$ |
| September " 15, , " " | $\begin{array}{lll} 8 & 36 \cdot 6 & \text { " } \\ 7 & 41 \cdot 7 & \end{array}$ | $\begin{array}{lll} 8 & 26 \cdot 0 & \text { " } \\ 7 & 31.1 & \text { " } \end{array}$ |
| October 1, " $\ldots \ldots$. <br>  15, " $\ldots \ldots$. | $\begin{array}{lll} 6 & 38 \cdot 9 & \text { " } \\ 5 & 43 \cdot 9 & \end{array}$ | $\begin{array}{lll} 6 & 28 \cdot 2 & \text { " } \\ 5 & 33 \cdot 2 & \end{array}$ |
| $\begin{array}{cccc} \hline \text { November } & 1, & & \ldots \ldots \\ 6 & 15, & & \ldots \ldots \\ \hline \end{array}$ | $\begin{array}{ll} 437 \cdot 0 & \text { " } \\ 3 & 41 \cdot 9 \\ \hline \end{array}$ | $\begin{array}{lll} 4 & 26 \cdot 4 & 6 \\ 3 & 31 \cdot 3 & 6 \end{array}$ |
| $$ | $\begin{array}{lll} 2 & 38 \cdot 9 & \text { " } \\ 1 & 43.6 \end{array}$ | $\begin{array}{lll} 2 & 28 \cdot 2 & \text { " } \\ 1 & 33 \cdot 0 & \text { " } \end{array}$ |
| January 1, 1886...... | 1235.0 " | $1224 \cdot 3$ " |

For any other days than those given in the table, interpolate directly, or subtract 3.94 minutes for every day elapsed. For any other year add 0.35 minute for every year. Also add one minute

[^30]if the year is the second after leap-year ; add two minutes if it is the third after leap-year ; add three minutes if it is leap-year before March 1st, and subtract one minute if it is leap-year after March 1st.

For any other latitude than $40^{\circ}$ north (between $20^{\circ}$ and $50^{\circ}$ ) add 0.14 minute for each degree of latitude south of $40^{\circ}$, or subtract 0.18 minute for each degree of latitude north of $40^{\circ}$.
266. Observations. Knowing from the preceding table the hour and minute of the extreme elongation on any day, a little before that time suspend a plumb-line, precisely as in Art. 263, and place yourself south of it as there directed. As the north star moves one way, move your eje the other, so that the plumb-line shall continually seem to corer the star. At last the star will appear to stop moring for a time, and then begin to move backward. Fix the sight on the board (or the compass, etc.) in the position in which it was when the star ceased moring ; for the star was then at its extreme apparent elongation, east or west, as the case may be.

The eastern elongations from October to March, and the western elongations from April to September, occurring in the daytime, they will generally not be visible except with the aid of a powerful telescope.
267. Azimuths. The angle which the line from the eye to the plumb-line makes with the true meridian-i. e., the angle between the meridian plane and the rertical plane passing through the eye and the star-is called the Azimuth of the star. It is given in the following table for different latitudes, and for a number of years to come. For the intermediate latitudes it can be obtained by a simple proportion, similar to that explained in detail in Art. 264.*

[^31]| 8\% \% $\sim$ |  <br>  <br>  |
| :---: | :---: |
| \%ั่ |  <br>  <br>  <br> - -1 |
| O |  <br>  <br>  <br> - - |
| \% |  <br>  <br>  <br> - $\quad$ - |
| 8 |  <br>  <br>  <br> - - |
| 0 | ¢ <br>  <br>  <br> - -1 |
| ホ̇ |  <br>  <br>  <br> - -1 |
| 凩 |  <br>  <br> - -1 |
| ®i ¢ - |  <br>  <br>  <br> - $\quad \rightarrow$ |
| $\xrightarrow{\circ}$ |  <br>  <br>  |
| 808 |  <br>  <br>  <br> - - |
| a <br> $\substack{0 \\ \sim \\ \sim \\ \hline}$ |  <br>  <br>  |
| $\infty$ $\infty$ $\infty$ $\sim$ |  <br>  <br>  <br> - - |
| $\xrightarrow{\substack{0 \\ \sim \\ \sim}}$ |  <br>  <br> - -1 <br> - |
| - <br> $\sim$ <br> $\sim$ |  <br>  <br>  <br> - - <br> - |
| 10 $\substack{0 \\ \sim \\ \sim}$ |  <br>  <br>  <br> - $\quad$ |
| $\stackrel{\mathrm{E}}{4}$ |  |

268. Setting out a Meridian. When two points in the direction of the north star at its extreme elongation have been ob-
Fig. 191. tained, as in Art. 266, the true meridian can be found thus: Let A and B be the two points. Multiply the natural tangent of the azimuth given in the table by the distance A B. The product will be the length of a line which is to be set off from $B$, perpendicular to $A B$, to some point C . A and C will then be points in the true meridian. This operation may be postponed till morning.

If the directions of both the extreme eastern and extreme western elongations be set out, the line lying midway between them will be the true meridian.
269. Determining the Declination. The declination would, of course, be given by taking the bearing of the meridian thus obtained, but it can also be determined by taking the bearing of the star at the time of the extreme elongation, and applying the following rules :

When the azimuth of the star and its magnetic bearing are one east and the other west, the sum of the two is the magnetic declination, which is of the same name as the azimuth-i. e., east, if that be east, and west, if it be west.

When the azimuth of the star and its magnetic bearing are both east or both west, their difference is the declination, which will be of the same name as the azimuth and bearing, if the azimuth be the greater of the two, or of the contrary name if the azimuth be the smaller.

All these cases are presented together in the figure, in which P is the north pole, Z the place of the observer, Z P the true meridian, S the star at its greatest eastern elongation, and $\mathrm{ZN}, \mathrm{Z} \mathrm{N}^{\prime}$, $Z \mathrm{~N}^{\prime \prime}$ rarious supposed directions of the needle.

Call the azimuth of the star-i. e., the angle P Z S— $2^{\circ}$ east.

Suppose the needle to point to N , and the

Fig. 192.

bearing of the star-i. e., S Z N-to be $5^{\circ}$ west of magnetic north. The declination PZN will evidently be $7^{\circ}$ east of true north.

Suppose the needle to point to $\mathrm{N}^{\prime}$, and the bearing of the star --i. e., $N^{\prime} Z S$-to be $1_{4}^{1^{\circ}}$ east of magnetic north. The declination will be $\frac{3^{\circ}}{}{ }^{\circ}$ east of true north, and of the same name as the azimuth, because that is greater than the bearing.

Suppose the needle to point to $\mathrm{N}^{\prime \prime}$, and the bearing of the star -i. e., $\mathrm{N}^{\prime \prime} \mathrm{Z} \mathrm{S}$-to be $10^{\circ}$ east of magnetic north. The declination will be $8^{\circ}$ west of true north, of the contrary name to the azimuth, because that is the smaller of the two.*

If the star were on the other side of the pole, the rules would apply likewise.
270. Other Methods. Many other methods of determining the true meridian are employed ; such as by equal altitudes and azimuths of the sun, or of a star; by one azimuth, knowing the time ; by observations of circumpolar stars at equal times before and after their culmination, or before and after their greatest elongation, etc.

All these methods, however, require some degree of astronomical knowledge ; and those which have been explained are abundantly sufficient for all the purposes of the ordinary land-surveyor.
"Burt's Solar Compass" is an instrument by which, "when adjusted for the sun's declination and the latitude of the place, the azimuth of any line from the true north and south can be read off, and the difference between it and the bearing by the compass will then be the variation."
271. Magnetic Declination in the United States. The declination in any part of the United States can be approximately obtained by mere inspection of the map at the beginning of this volume. $\dagger$ Through all the places at which the needle, in 1885, pointed to the true north, a line is drawn on the map, and called

[^32]the line of no declination. It will be seen to pass a little east of Charleston, South Carolina, thence in a northwesterly direction, passing near Zanesville, Ohio, through the west end of Lake Erie, passing a little west of Detroit, and up through the east end of Lake Superior. This line is now slowly moving westward.

At all places situated to the east of this line (including the New England States, New York, New Jersey, Delaware, Maryland, Pennsylvania, most of Virginia, and the east half of North Carolina and Ohio) the declination is westerly-i. e., the north end of the needle points to the west of the true north. At all places situated to the west of this line (including the Western and Soathern States) the declination is easterly-i. e., the north end of the needle points to the east of the true north. This declination increases in proportion to the distance of the place on either side of the line of no variation, reaching $23^{\circ}$ of easterly declination in Washington Territory, and $21^{\circ}$ of westerly declination in Maine.

Isogonics, or lines of equal declination, are lines drawn through all the places which have the same declination. On the map they are drawn for each degree. All the places situated on the line marked $5^{\circ}$, east or west, have $5^{\circ}$ declination ; those on the $10^{\circ}$ line have $10^{\circ}$ declination, etc. The declination at the intermediate places can be approximately estimated by the eye. These lines all refer to 1885.

The sign + indicates west declination, and the sign - indicates east declination. The aunual change in the secular rariation for stations is given in minutes and decimals, a + indicating increasing west declination or decreasing east declination, and a sign indicating increasing east and decreasing west declination.
272. To correct Magnetic Bearings. The declination at any place and time being known, the magnetic bearings taken there and then may be reduced to their true bearings by these rules :

Rule 1. When the declination is west, as it is in the Northeastern States, the true bearing will be the sum of the declination, and a bearing which is north and west, or south and east ; and the difference of the declination and a bearing which is north and east, or south and west. To apply this to the cardinal points, a
north bearing must be called $\mathrm{N} .0^{\circ}$ west, an east bearing N. $90^{\circ}$ E., a south bearing S. $0^{\circ}$ E., and a west bearing S. $90^{\circ} \mathrm{W}$.; counting around from $\mathrm{N}^{\prime}$ to N , in the figure, and so onward, "with the sun."

The reasons for these corrections are apparent from the figure, in which the dotted lines and the accented letters represent the direction
 of the needle, and the full lines and the unaccented letters represent the true north and south and east and west lines.

When the sum of the declination and the bearing is directed to be taken, and comes to more than $90^{\circ}$, the supplement of the sum is to be taken, and the first letter changed. When the difference is directed to be taken, and the declination is greater than the bearing, the last letter must be changed. A diagram of the case will remove all doubts. Examples of all these cases are given below for a declination of $8^{\circ}$ west:

| MAGNETIC BEARINGS. | $\begin{gathered} \text { TRUE } \\ \text { BEARINGS. } \end{gathered}$ | MAGNETIC BEARINGS. | ${ }_{\text {- }}^{\text {bearinge }}$ tree |
| :---: | :---: | :---: | :---: |
| North. | N. $8^{\circ} \mathrm{W}$. | South. | S. $8^{\circ} \mathrm{E}$. |
| N. $1^{\circ} \mathrm{E}$. | N. $7^{\circ} \mathrm{W}$. | S. $2^{\circ} \mathrm{W}$. | S. $6^{\circ} \mathrm{E}$. |
| N. $40^{\circ} \mathrm{E}$. | N. $32^{\circ} \mathrm{E}$. | S. $60^{\circ} \mathrm{W}$. | S. $52^{\circ} \mathrm{W}$. |
| East. | N. $82^{\circ} \mathrm{E}$. | West. | S. $82^{\circ} \mathrm{W}$. |
| S. $50^{\circ} \mathrm{E}$. | S. $58^{\circ} \mathrm{E}$. | N. $70^{\circ} \mathrm{W}$. | N. $78^{\circ} \mathrm{W}$. |
| S. $89^{\circ} \mathrm{E}$. | N. $83^{\circ} \mathrm{E}$. | N. $83^{\circ} \mathrm{W}$. | S. $89^{\circ} \mathrm{W}$. |

Fig. 194.


Rule 2. When the declination is east, as in the Western and Southern States, the preceding directions must be exactly reversed -i. e., the true bearing will be the difference of the declination, and a bearing which is north and west or south and east; and the sum of the declination and a bearing which is north and east or south and west. A north bearing
must be called N. $0^{\circ}$ E., a west bearing N. $90^{\circ}$ W., a south bearing S. $0^{\circ}$ W., and an east bearing S. $90^{\circ}$ E., counting from $\mathrm{N}^{\prime}$ to N , and so onward, "against the sun." The reasons for these rules are seen in the figure. Examples are given below for a declination of $5^{\circ} \mathrm{E}$.:

| $\underbrace{\text { Bearives }}_{\text {Magnetic }}$ |  |  | $\underset{\text { bearings. }}{\text { Trem }}$ |
| :---: | :---: | :---: | :---: |
| North. | N. $5^{\circ} \mathrm{E}$. |  |  |
| $\begin{aligned} & \text { N. } 40^{\circ} \mathrm{E} . \\ & \text { N. } 80^{\circ} \mathrm{E} . \end{aligned}$ | N. $45^{\circ} \mathrm{E}$. S. 86 | S. $60^{\circ} \mathrm{W}$. <br> S. $87^{\circ} \mathrm{W}$ | S. ${ }_{\text {S. }} 65^{\circ}{ }^{\circ} \mathrm{W}$ W. |
| N. ${ }^{\text {East. }}$. | S. $85^{\circ} \mathrm{E}$. | West. | N. $85^{\circ} \mathrm{W}$. |
| S. $1^{1}{ }^{\circ} \mathrm{E}$. | S. $4^{\circ} \mathrm{W}$. | N. $70^{\circ} \mathrm{W}$ W. | N. $65^{\circ} \mathrm{W}$. |
| S. $50{ }^{\circ} \mathrm{E}$. | S. $45^{\circ} \mathrm{E}$. | N. $2^{\circ} \mathrm{W}$. | N. $3^{\circ} \mathrm{E}$. |

273. To survey a Line with True Bearings. The compass may be set, or adjusted, by means of the vernier, according to the declination in any place, so that the bearings of any lines then taken with it will be their true bearings. To effect this, turn aside the compass-plate by means of the tangent-screw which moves the vernier a number of degrees equal to the declination, moving the south end of the compass-box to the right (the north end being supposed to go ahead) if the declination be westerly, and vice versa; for that moves the north end of the compass-box in the contrary direction, and thus makes a line which before was N. by the needle, now read, as it should truly, north, so many degrees west if the declination was west; and similarly in the reverse case.

## Variations of Magnetic Declination.

274. The variations of the declination are of more practical importance than its absolute amount. They are of four kinds : Irregular, diurnal, annual, and secular.
275. Irregular Variation. The needle is subject to sudden and violent changes, which hare no known law. They are sometimes coincident with a thunder-storm, or an aurora borealis
(during which changes of nearly $1^{\circ}$ in one minute, $2 \frac{1}{2}^{\circ}$ in eight minutes, and $10^{\circ}$ in one night, have been observed), but often have no apparent cause, except an otherwise invisible " magnetic storm."
276. The Diurnal Variation. On continuing observations of the direction of the needle throughout an entire day, it will be found, in the northern hemisphere, that the north end of the needle moves westward from about 8 A. M. till about $1 \frac{1}{2}$ P. M., over an arc of from $5^{\prime}$ to $15^{\prime}$, and then gradually returns to its former position. A similar but smaller movement takes place during the night. At Philadelphia, the most easterly deflection of the needle is at about ${ }^{7} \frac{8}{4} \mathrm{~A}$. m. The north end of the needle then begins to move toward the west, crossing the mean magnetic meridian about $10 \frac{1}{2}$ A. M., and reaching its extreme western position about $1 \frac{1}{2}$ P. m. The total angular range averages about $8^{\prime}$, being $10 \frac{1}{2}^{\prime}$ in August, and $6^{\prime}$ in November. ${ }^{*}$ The period of this change being a day, it is called the Diurnal Variation. Its effect on the permanent variation is necessarily to cause it, in places where it is west, to attain its maximum at about $1 \frac{1}{2}$ P. M., and its minimum at about 8 A. M. ; and the reverse where the declination is east.

This diurnal variation adds a new element to the inaccuracies of the compass, since the bearings of any line taken on the same day, at a few hours' interval, might vary a quarter of a degree, which would cause a deviation of the end of the line, amounting to nearly half a link at the end of a chain, and to 35 links, or 23 feet, at the end of a mile. The hour of the day at which any important bearing is taken should therefore be noted.

2\%\%. The Annual Variation. If the observations be continued. throughout an entire year, it will be found that the diurnal changes vary with the seasons, being greater in summer than in winter. The period of this variation being a year, it is called the Annual Variation.

[^33]278. The Secular Variation. When accurate observations on the declination of the needle in the same place are continued for several years, it is found that there is a continual and tolerably regular increase or decrease of the declination, continuing to proceed in the same direction for so long a period, that it may be called the Secular Variation of the declination.

The most ancient observations are those taken in Paris. In the year 1541 the needle pointed $7^{\circ}$ east of north ; in 1580 the declination had increased to $11_{\frac{1}{2}}{ }^{\circ}$ east, being its maximum ; the needle then began to more westward, and in 1666 it had returned to the meridian ; the declination then became west, and continued to increase till in 1814 it attained its maximum, being $22^{\circ} 34^{\prime}$ west of north. It is now decreasing, and, January 1, $18 \div 9$, it was $16^{\circ} \check{\circ} 6^{\prime}$ west.

In this country the north end of the needle was moring eastward at the earliest recorded observations, and continued to do so till about the year 1810 (rariously recorded as from $1 \% 65$ to 1819), when it began to more westward, which it has ever since continued to do. Thus, in Boston, from $1 ; 00$ to 1807, the declination changed from $10^{\circ}$ west to $6^{\circ} 5^{\prime}$ west, and, from $180 \%$ to 1879 , it changed from $6^{\circ} 5^{\prime}$ west to $11^{\circ} 36^{\prime}$ west.

In Philadelphia, from $1 \% 01$ to 1802 , the declination changed from $8^{\circ} 30^{\prime}$ west to $1^{\circ} 30^{\prime}$ west, and, from 1802 to $18 \div \%$, it changed from $1^{\circ} 30^{\prime}$ west to $6^{\circ} 2^{\prime}$ west.

Extensive tables of the declination, at more than tro thousand stations, in various parts of the United States, are given in the "Report of the United States Coast and Geodetic Surrer." 1882, Appendix XIII, by Charles A. Schott. The secular variation is noted on the declination-map in this rolume.

An examination of the abore-mentioned tables will show that the secular rariation often differs greatly in places not far apart, and that it varies in amount at the same place from year to year :

TABLE OF COMPUTED ANNUAL CHANGES IN DECLINATION.

| LOCALITY. | annual change. |  |  |
| :---: | :---: | :---: | :---: |
|  | 1870. | 1880. | 1885. |
| Portland, Me. | $+2 \cdot 4^{\prime}$ | $+1 \cdot 6^{\prime}$ | $+1 \cdot 2^{\prime}$ |
| Burlington, Vt. | $+5 \cdot 0$ | +6.0 | $+5 \cdot 8$ |
| Portsmouth, N. H. | $+4 \cdot 4$ | $+3 \cdot 7$ | +3.3 |
| Boston, Mass. | $+3 \cdot 4$ | $+2 \cdot 9$ | $+2 \cdot 5$ |
| Hartford, Conn | $+3 \cdot 8$ | $+3 \cdot 7$ | $+3 \cdot 6$ |
| Albany, N. Y.. | $+4 \cdot 3$ | $+3 \cdot 7$ | $+3 \cdot 4$ |
| New York, N. Y | $+2 \cdot 4$ | $+2 \cdot 5$ | $+2 \cdot 6$ |
| Buffalo, N. Y. | $+5 \cdot 1$ | $+5 \cdot 0$ | $+4 \cdot 8$ |
| Philadelphia, Pa | $+4 \cdot 9$ | +4.9 | $+5 \cdot 3$ |
| Baltimore, Md | $+3 \cdot 9$ | $+3 \cdot 6$ | $+3 \cdot 2$ |
| Washington, D. C. | $+3 \cdot 5$ | $+3 \cdot 2$ | +3.0 |
| Cleveland, Ohio. | $+2 \cdot 8$ | $+2.5$ | +2.2 |
| Detroit, Mich. | $+3 \cdot 4$ | +3.0 | +2.8 |
| St. Louis, Mo. | $+3 \cdot 4$ | $+3 \cdot 2$ | $+3 \cdot 0$ |
| Cape Henry, Va. | +3. 8 | $+3 \cdot 7$ | $+3 \cdot 6$ |
| Charleston, S. C | $+3 \cdot 5$ | $+3 \cdot 0$ | $+2 \cdot 7$ |
| Savannah, Ga.. | $+3 \cdot 6$ | $+3 \cdot 5$ | +3.3 |
| Key West, Fla | $+4 \cdot 3$ | $+4 \cdot 2$ | $+4 \cdot 1$ |
| Mobile, Ala. | $+2 \cdot 8$ | +3.4 | $+3 \cdot 7$ |
| New Orleans, La. | $+3 \cdot 1$ | $+3 \cdot 5$ | $+3 \cdot 7$ |
| San Francisco, Cal............. | $-1 \cdot 0$ | $-0 \cdot 5$ | $-0 \cdot 3$ |
| Cape Disappointment, W. Ter. . | $-3 \cdot 4$ | $-3 \cdot 1$ | $-2 \cdot 7$ |
| Sitka, Alaska.. | $+1 \cdot 0$ | $+2 \cdot 1$ | $+2 \cdot 5$ |

279. Determination of the Change, by Interpolation. To determine the change at any place and for any interval not found in the recorded observations, an approximation, sufficient for most purposes of the surveyor, may be obtained by interpolation (by a simple proportion) between the places given on the map, assuming the movements to have been uniform between the given dates, and also assuming the change at any place not found on the map to have been intermediate between those of the lines of equal variation, which pass through the places of recorded observations on each side of it, and to have been in the ratio of its respective distances from those two lines ; for example, taking their arithmetical mean, if the required place is midway between them ; if it be twice as near one as the other, dividing the sum of twice the change of the nearest line, and once the change of the other, by three ; and so in other cases-i. e., giving the change at each place, a
"weight" inversely as its distance from the place at which the change is to be found.
280. Determination of the Change by Old Lines. When the former bearing of any old line, such as a farm-fence, etc., is recorded, the change in the declination from the date of the original observation to the present time can be at once found by setting the compass at one end of the line and sighting to the other. The difference of the two bearings is the required change.

If one end of the old line can not be seen from the other, as is often the case when the line is fixed only by a "corner" at each end of it, proceed thus : Run a line from one corner with the old bearing and with its distance. Measure the distance from the end of this line to the other corner, to which it will be opposite. Multiply this distance by $5 \% \cdot 3$, and divide by the length of the line. The quotient will be the change of rariation in degrees.*

For example, a line 63 chains long, in $182 \%$ had a bearing of north $1^{\circ}$ east. In $184 \%$ a trial line was run from one end of the former line with the same bearing and distance, and its other end was found to be 125 links to the west of the true corner. The change of declination was therefore $\frac{1 \cdot 25 \times 5 \% \cdot 3}{63}=1 \cdot 137^{\circ}=1^{\circ} 8^{\prime}$ westerly.
281. Effects of the Secular Change. These are exceedingly important in the resurrey of farms by the bearings recorded in old deeds. Let SN denote the direction of the needle at the time of the original surrey, and $S^{\prime} N^{\prime}$ its direction at the time of the resurvey, a number of years later. Suppose the change to have been

Fig. 195.

> * Let AB be the original line; AC the trial line, and $B C$ the distance between their extremities. $A . B$ and $A C$ may be regarded as radii of a circle and B C as a chord of the are which subtends their angle. Assuming the chord and arc to coincide (which they will, nearly, for small angles), we have this proportion: Whole circumference : are BC: $360^{\circ}:$ BAC $:$ or, $2 \times \mathrm{AC} \times 3 \cdot 1416:$ BC $: 360^{\circ}:$

B AC, whence B A C $=\frac{\mathrm{BC}}{\mathrm{AB}} \times 57.3$; or, more precisely, 57.29578 .
$3^{\circ}$, the needle pointing so much farther to the west of north. The line SN, which before was due north and south by the needle, will now bear N. $3^{\circ}$ E. and S. $3^{\circ} \mathrm{W}$. ; the line A B, which before was N. $40^{\circ} \mathrm{E}$. will now bear N. $43^{\circ}$ E. ; the line D F, which before was N. $40^{\circ}$ W., will now bear N. $37^{\circ} \mathrm{W}$., and the line W E, which before was due east and west, will now bear S. $87^{\circ} \mathrm{E}$.

Fig. 196.
 and N. $87^{\circ} \mathrm{W}$. Any line is similarly changed. The proof of this is apparent on inspecting the figure.

Suppose, now, that a surveyor, ignorant or neglectful of this change, should attempt to run out a farm by the old bearings of the deed, none of the old fences or corners remaining. The full

Fig. 197.
 lines in the figure represent the original bounds of the farm, and the dotted lines those of the new piece of land which, starting from A, he would unwittingly run out. It would be of the same size and the same shape as the true one, but it would be in the wrong place. None of its lines would agree with the true ones, and in some places it would encroach on one neighbor, and in other places would leave a gore, which belongs to it, between itself and another neighbor. Yet this is often done, and is the source of a great part of the litigation among farmers respecting their " lines."
282. To run out Old Lines. To succeed in retracing old lines, proper allowance must be made for the change in the variation
since the date of the original survey. That date must first be accurately ascertained; for the survey may be much older than the deed, into which its bearings may have been copied from an older one. The amount and direction of the change is then to be ascertained by the methods of Art. 279 or 280. The bearings may then be corrected by the following Rules :

When the north end of the needle has been moring westerly, the present bearings will be the sums of the change and the old bearings which were northeasterly or southwesterly, and the differences of the change and the old bearings which were northwesterly or southeasterly.

If the change has been easterly, reverse the preceding rules, subtracting where it is directed to add, and adding where it is directed to subtract.

Run out the lines with the bearings thus corrected.
It will be noticed that the process is precisely the reverse of that in Art. 2\%\%. The rules, there given in more detail, may therefore be used : Rule 1. "When the declination is west," being employed when the change has been a movement of the N. end of the needle to the east ; and Rule 2, "when the declination is east," being employed when the N . end of the needle has been moving to the west.

If the compass has a vernier, it can be set for the change, once for all, precisely as directed in Art. 273, and then the courses can be run out as given in the deed, the correction being made by the instrument.

Example. The following is a remarkable case which came before the Supreme Court of New York: The north line of a large estate was fixed by a royal grant, dated in $1 \% 04$, as a due east and west line. It was run out in 1\%15, by a surveyor, whom we will call Mr. A. It was again surreyed in $1: 65$, by Mr. B., who ran a course N. $8 \%^{\circ} 30^{\prime}$ E. It was run out for a third time in 1759 , by Mr. C., who adopted the course N. $86^{\circ} 18^{\prime}$ E. In 1845 it was surreyed for the fourth time by Mr. D., with a course of N. $88^{\circ} 30^{\prime} \mathrm{E}$. He found old "corners," and "blazes" of a former surrey, on his line. They are also found on another line, south of his. Which
of the preceding courses were correct, and where does the true line lie?

The question was investigated as follows: There were no old records of variation at the precise locality, but it lies between the lines of equal variation which pass through New York and Boston, its distance from the Boston line being about twice its distance from the New York line. The records of those two cities (referred to in Art. 2\%8) could therefore be used in the manner explained in Art. 2\%9. For the later dates, obserrations at New Haven could serve as a check. Combining all these, the author inferred the variation at the desired place to have been as follows :
In 1715 , variation $8^{\circ} 02^{\prime}$ west.
In 1765 , " $5^{\circ} 32^{\prime} " \quad$ Decrease since $1715,2^{\circ} 30^{\prime}$.
In 1789, " ${ }^{\circ} 05^{\prime} "$ Decrease since $1765,0^{\circ} 27^{\prime}$. In 1845, " ${ }^{2} \quad 7^{\circ} 23^{\prime}{ }^{\prime} \quad$ Increase since $1789,2^{\circ} 18^{\prime}$.
We are now prepared to examine the correctness of the allowances made by the old surveyors.

The course run by Mr. B. in 1765 , N. $8 \%^{\circ} 30^{\prime}$ E., made an allowance of $2^{\circ} 30^{\prime}$ as the decrease of variation, agreeing precisely with our calculation. The course of Mr. C. in $1 \% 89$, N. $86^{\circ} 18^{\prime}$ E., allowed a change of $1^{\circ} 12^{\prime}$, which was wrong by our calculation, which gives only about $27^{\prime}$, and was deduced from three different records. Mr. D., in 1845, ran a course of N. $88^{\circ} 30^{\prime}$ E., calling the increase of variation since $1789,2^{\circ} 12^{\prime}$. Our estimate was $2^{\circ}$ $18^{\prime}$, the difference being comparatively small. Our conclusion, then, is this : The second surveyor retraced correctly the line of the first; the third surveyor ran out a new and incorrect line; and the fourth surveyor correctly retraced the line of the third, and found his marks, but this line was wrong originally, and therefore wrong now. All the surveyors ran their lines on the supposition that the original "due east and west line" meant east and west as the needle pointed at the time of the original survey.

The preponderance of the testimony as to old landmarks agreed with the results of the above reasoning, and the decision of the court was in accordance therewith.

In the figure below, the horizontal and rertical lines represent true east and north lines ; and the two upper lines running from left to right represent the two lines set out by
 the surreyors, and in the jears there named.
283. Remedy for the Evils of the Secular Change. The only complete remedy for the disputes, and the uncertainty of bounds, resulting from the continued change in the declination, is this: Let a meridian-i. e., a true north and south line-be established in every town or county, by the authority of the State ; monuments, such as stones, set deep in the ground, being placed at each end of it. Let every surveyor be obliged by law to test his compass by this line, at least once in each year, at a given hour in the day. This he could do as easily as in taking the bearing of a fence, by setting his instrument on one monument, and sighting to a staff held on the other. Let the rariation thus ascertained be inserted in the notes of the surrey, and recorded in the deed. Another surveyor, years or centuries afterward, could test his compass by taking the bearing of the same monuments, and the difference between this and the former bearing would be the change of declination. He could thus determine with entire certainty the proper allowance to be made (as in Art. 282) in order to retrace the original line, no matter how much, or how irregularly, the declination may hare changed, or how badly adjusted was the compass of the original surver. Any permanent line employed in the same manner as the meridian line would answer the same purpose, though less conveniently, and every surveyor should have such a line, at least for his own use.*

[^34]
## CHAPTER IV.

## TRANSIT-SURVEYING-BY THE THIRD METHOD.

## THE INSTRUMENTS.

284. The Transit is a Goniometer, or Angle-Measurer. It consists, essentially, of a circular plate of metal, supported in such a manner as to be horizontal, and divided on its outer circumference into degrees and parts of degrees. Through the center of this plate passes an upright axis, and on it is fixed a second circular plate, which nearly touches the first plate, and can turn freely around to the right and to the left. This second plate carries a telescope, which rests on upright standards firmly fixed to the plate, and which can be pointed upward and downward. By the combination of this motion and that of the second plate around its axis, the telescope can be directed to any object. The second plate has some mark on its edge, such as an arrow-head, which serves as a pointer or index for the divided circle, like the hand of a clock. When the telescope is directed to one object, and then turned to the right or to the left, to some other object, this index which moves with it, and passes around the divided edge of the other plate, points out the arc passed over by this change of direction, and thus measures the horizontal angle made by the lines imagined to pass from the center of the instrument to the two objects.

The great value of this instrument, and the accuracy of its measurements of angles, are due chiefly to two things : to the telescope with its cross-hairs, by which great precision in sighting to a point is obtained ; and to the vernier scale, which enables minute portions of any are to be read with ease and correctness. The
former assists the eye in directing the line of sight, and the latter aids it in reading off the results. Arrangements for giving slow

Fig. 199.

and steady motion to the movable parts of the instrument add to the ralue of the above. A contrivance for repeating the obserra-
tion of angles still further lessens the unaroidable inaccuracies of these observations.
285. The Surveyor's Transit (Fig. 199). In this instrument the telescope takes the place of the plain sights of the surveyor's compass, and the angles are read on the graduated limb to single minutes by the vernier.

A level is attached to the telescope, and a vertical circle is attached to the telescope-axis inside of the left-hand standard. The vertical angles through which the telescope is moved may be read off from the vernier attached to the left-hand standard, and shown below the vertical circle. The slow-motion screw for the vertical circle is shown attached to the right-hand standard. The clamp for the axis is hidden by the telescope. The standards upon which the telescope-axis rests are fastened to the upper plate (the vernier-plate). This plate also carries the compass-circle. The compass-circle with its accessories is similar to that already explained in the Surveyor's Compass. The compass-circle can be turned on its center, so that the declination of the needle can be set off, and lines can be run with their true bearings. The vernierplate covers the lower plate (the divided limb), so that only two short ares of the divided limb are seen through openings where the verniers are placed. The screw which clamps the vernier-plate to the divided limb is shown on the right of the plate, together with the slow-motion screw. The lower clamp and the slow-motion screw are attached to the upper parallel plate.
286. As the value of this instrument depends greatly on the accurate fitting and bearings of the two concentric vertical axes, and as their connection ought to be thoroughly understood, a vertical section through the body of the instrument is given in Fig. 200.

The upper plate, or vernier-plate, A, A, carries the verniers, compass-box, and telescope. It is attached to its socket by the flange, K. This socket is fitted to the outside, conical surface of the main socket, C . The main socket, to which is attached the divided limb, B, B, is fitted to the conical spindle H , and held
on the spindle by the spring-catch S . A screw holds the conical center, whose upper flange keeps the sockets of the two plates

Fig. 200.

together. The clamp is at F. Two of the four leveling screws are shown in section. The spindle, H, passes through the upper parallel plate, and is attached to a movable section of the lower parallel plate by a ball-and-socket joint. The leveling screws pass through the upper parallel plate, and rest in cups on the lower parallel plate. As the leveling screws are morable on the lower parallel plate, the movable section of this plate enables the upper part of the instrument to be mored from side to side, so as to bring the center of the instrument precisely orer any desired point. This arrangement is called a "shifting center." At the lower end of the spindle is a loop, P , from which the plumb-bob is suspended.
287. The Telescope. This is a combiuation of leuses, placed in a tube, and so arranged, in accordance with the laws of optical science, that an image of any object to which the telescope may be directed, is formed within the tube (by the rays of light coming from the object and bent in passing through the object-glass), and
there magnified by an eye-glass, or eye-piece, composed of several lenses. The arrangement of these lenses is very various. Those two combinations, which are preferred for surveying instruments, will be here explained :

Fig. 201 represents a telescope which inverts objects. Any object is rendered visible by every point of it sending forth rays of light in every direction. In this figure the highest and lowest points of the object, which here is an arrow, A, are alone considered. Those of the rays proceeding from them, which meet the object-glass, 0 , form a cone. The center line of each cone, and its extreme upper and lower lines, are alone shown in the figure. It will be seen that these rays, after passing through the object-glass, are refracted or bent by it, so as to cross one another, and thus to form at B an inverted image of the object. This would be rendered visible, if a piece of groundglass, or other semi-transparent substance, were placed at the point B , which is called the focus of the objectglass. The rays which form this image continue onward and pass through the two lenses C and D , which act like one magnifying-glass, so that the rays, after being refracted by them, enter the eye at such angles as to form there a magnified and inverted image of the object. This combination of the two plano-convexlenses, C and D, is known as " Ramsden's Eye-piece."

This telescope, inverting objects, shows them upside down, and the right side on the left. They can be shown erect by adding one or two more lenses, as in the marginal figure. But as these lenses absorb light and lessen the distinctness of vision, the former arrangement is sometimes preferred. A little practice makes it equally convenient for the observer, who soon becomes accustomed to seeing his flagmen standing on their heads, and soon learns to motion them to the right when he wishes them to go to the left, and vice versa.

Fig. 201


Fig. 202.


Fig. 202 represents a telescope which shows objects erect. Its eye-piece has four lenses. The eye-piece of the common terrestrial telescope, or spy-glass, has three. Many other combinations may ke used, all intended to show the object achromatically, or free from false coloring, but the one here shown is that most generally preferred at the present day. It will be seen that an inverted image of the object $A$ is formed at $B$, as before, but that the rays continuing onward are so refracted in passing through the lens C as to again cross, and thus, after further refraction by the lenses D and E , to form, at F , an erect image, which is magnified by the lens G.

In both these figures, the limits of the page render it necessary to draw the angles of the rays very much out of proportion.
288. Cross-Hairs. Since a considerable field of viem is seen in looking through the telescope, it is necessary to provide means for directing the line of sight to the precise point which is to be observed. This could be effected by placing a very fine point, such as that of a needle, within the telescope, at some place where it could be distinctly seen. In practice, this fine point is obtained by the intersection of two very fine lines, placed in the common focus of the object-glass and of the eye-piece. These lines are called the crosshairs, or cross-wires. Their intersection can be seen through the eye-piece, at the same time, and apparently at the same place, as the image of the distant object. The magnifying powers of the ere-piece will then detect the slightest deviation from perfect coincidence. "This application of the telescope may be considered as completely annihilating that part of the error of observation which might otherwise arise from an erroneous estimation of the direction in which an object lies from the observer's eye, or from the center of the in-
strument. It is, in fact, the grand source of all the precision of modern astronomy, without which all other refinements in instrumental workmanship would be thrown away." What Sir John Herschel here says of its utility to astronomy is equally applicable to surveying.

The imaginary line which passes through the intersection of the cross-hairs and the optical center of the object-glass is called the line of collimation of the telescope.*

The cross-hairs are attached to a ring, or short, thick tube of brass, placed within the telescope - tube, through holes in which pass loosely four screws, whose threads enter and take hold of the ring, behind or in front of the crosshairs, as shown (in front view and in section) in the two figures in the margin. Their movements will be explained in "AdJust-

Fig. 203.


Fig. 204.
 ments."

Usually, one cross-hair is horizontal, and the other vertical, as in Fig. 203, but sometimes they are arranged as in

Fig. 205.
 Fig. 205, which is thought to enable the object to be bisected with more precision. A horizontal hair is sometimes added.

The cross-hairs are best made of platinum wire, drawn out very fine by being previously inclosed in a larger wire of silver, and the silver then removed by nitric acid. Silk threads from a cocoon are sometimes used. Spiders' threads are, however, the most usual. If a crosshair is broken, the ring must be taken out by removing two opposite screws, and inserting a wire with a screw cut on its end, or a stick of suitable size, into one of the holes thus left open

[^35]in the ring, it being turned sidewise for that purpose, and then removing the other screws. The spiders' threads are then stretched across the notches seen in the end of the ring, and are fastened by gum, or varnish, or beesrax. The operation is a very delicate one. The following plan has been employed : A

Fig. 206.
 piece of wire is bent, as in the figure, so as to leave an opening a little wider than the ring of the cross-hairs. A cobweb is chosen, at the end of which a spider is hanging, and it is wound around the bent wire, as in the figure, the weight of the insect keeping it tight and stretching it ready for use, each part being made fast by gum, etc. When a cross-hair is wanted, one of these is laid across the ring and there attached. One method is to draw the thread out of the spider, persuading him to spin, if he sulks, by tossing him from hand to hand. Another method is to unwind the spider-web from the cocoons, frequently to be found in spiderwebs. A stock of such threads must be obtained in warm weather for the winter's wants. A piece of thin glass, with a horizontal and a vertical line etched on it, may be made a substitute.
289. Instrumental Parallax. This is an apparent morement of the cross-hairs about the object to which the line of sight is directed, taking place on any slight morement of the eye of the observer. It is caused by the image and the cross-hairs not being precisely in the common focus, or point of distinct rision of the eye-piece and the object-glass. To correct it, more the eye-piece out or in till the cross-hairs are seen clearly and sharply defined against any white object. Then move the object-glass in or out till the object is also distinctly seen. The cross-hairs will then seem to be fixed to the object, and no morement of the eye will cause them to appear to change their place.
290. A milled-headed screw (on the farther side of the telescope, and not shown in the figure) passes into the telescope, and has a pinion at its other end entering a toothed rack (Fig.

207 ), and is used to move the object-glass, $O$, out and in, according as the object looked at is nearer or farther than the one last observed. Short distances require a long tube; long dis-

Fig. 207.
 tances a short tube.

The eye-piece is moved in and out by a similar arrangement to the preceding. This movement is necessary in order to obtain a distinct view of the cross-hairs. Short-sighted persons require the eye-piece to be pushed farther in than persons of ordinary sight, and old or long-sighted persons to have it drawn farther out.
291. Supports. The telescope of the transit is supported by a hollow axis at right angles to it, which itself rests, at each end, on two upright pieces, or standards, spreading at their bases so as to increase their stability.

One end of the axis rests upon a movable block, which can be raised or lowered by a capstan-screw. The use of this will be shown in "Adjustments."
292. The Indexes. The supports, or standards, of the telescope just described are attached to the upper or index-carrying circle. This, as has been stated, can turn freely on the lower or graduated circle, by means of its conical axis moving in the hollow conical axis of the latter circle. This upper circle carries the index, V,
 which is an arrow-head or other mark on its edge, or the zero-point of a vernier scale. There are usually two of these, situated exactly opposite to each other, or at the extremities of a diameter of the upper circle, so that the readings on the graduated circle pointed out by them differ, if both are correct, exactly $180^{\circ}$. The object of this arrangement is to correct any error of eccentricity, arising from the center of the axis which carries the upper circle (and with which it and its index-pointers
turn), not being precisely in the center of the graduated circle. In the figure, let C be the true center of the graduated circle, but $\mathrm{C}^{\prime}$ the center on which the plate carrying the indexes turns. Let A $\mathrm{C}^{\prime} \mathrm{B}$ represent the direction of a sight taken to one object, and $\mathrm{D}^{\prime} \mathrm{C}^{\prime} \mathrm{E}^{\prime}$ the direction when turned to a second object. The angle subtended by the two objects at the center of the instrument is required. Let DE be a line passing through C , and parallel to $\mathrm{D}^{\prime} \mathrm{E}^{\prime}$. The angle ACD equals the required angle, which is therefore truly measured by the arc AD or B E. But if the arc shown by the index is read, it will be $\mathrm{AD}^{\prime}$ on one side, and $B E^{\prime}$ on the other ; the first being too small by the arc $\mathrm{D}^{\prime}$, and the other too large by the equal arc $\mathrm{E} \mathrm{E}^{\prime}$. If, however, the half-sum of the two arcs $\mathrm{AD}^{\prime}$ and $\mathrm{BE}^{\prime}$ be taken, it will equal the true arc, and therefore correctly measure the angle. Thus, if $\mathrm{A} \mathrm{D}^{\prime}$ was $19^{\circ}$, and $\mathrm{BE}^{\prime} 21^{\circ}$, their half-sum, $20^{\circ}$, would be the correct angle.

Three indexes, $120^{\circ}$ apart, are sometimes used. They hare the advantage of averaging the unaroidable inaccuracies and inequalities of graduation on different parts of the limb, and thus diminishing their effect on the resulting angle.
293. The Graduated Circle. This is divided into three hundred and sixty equal parts, or degrees, and each of these is subdivided into two or three parts or more, according to the size of the instrument. In the first case, the smallest dirision on the circle will of course be $30^{\prime}$; in the second case 20'. More precise reading, to single minutes or even less, is effected by means of the rernier of the index, all the rarieties of which will be fully explained under "Verniers." The numbers run from $0^{\circ}$ around to $360^{\circ}$, which number is necessarily at the same point as the 0 , or zeropoint. In most instruments there is another concentric circle, on which the degrees are also numbered from $0^{\circ}$ to $90^{\circ}$, as on the com-pass-circle. Each tenth degree is usually numbered, each fifth degree is distinguished by a longer line of dirision, and each de-gree-division line is longer than those of the subdivisions. A mag-nifying-glass is needed for reading the dirisions with ease. In large instruments it is attached to each rernier.
294. Movements. When the line of sight of the telescope is directed to a distant, well-defined point, the unaided hand of the observer can not move it with sufficient delicacy and precision to make the intersection of the cross-hairs exactly cover or "bisect" that point. To effect this, a clamp, and a tangent, or slow-motion, screw are required. This arrangement, as usually applied to the movement of the upper, or vernier plate, consists of a short post of brass, which is attached to the vernier-plate, and through which passes a long and fine-threaded " tangent-screw." The other end of this screw enters into and carries the clamp. This consists of two pieces of bbrass, which, by turning the clamp-screw, which passes through them on the outside, can be made to take hold of and pinch tightly the edge of the lower circle, which lies between them on the inside. The upper circle is now prevented from moving on the lower one, for the tangent-screw keeps them at a fixed distance apart, so that they can not move to or from one another, nor consequently the two circles to which they are respectively made fast. But when this tangent-screw is turned by its milled head, it gives the clamp and with it the upper plate a smooth and slow motion, backward or forward, whence it is called the "slow-motion screw," as well as "tangent-screw," from the direction in which it acts. Another form of clamp is shown in Fig. 200.

A little different arrangement is employed to give a similar motion to the lower circle on the body of the instrument. Its axis is embraced by a brass ring, into which enters a clamp-screw. The clamp-screw causes the ring to pinch and hold immovably the axis of the lower circle, while a turn of the tangent-screw will slowly more the clamp-ring itself, and therefore with it the lower circle. When the clamp is loosened, the lower circle, and with it everything above it, has a perfectly free motion.
295. Levels. Since the object of the instrument is to measure horizontal angles, the circular plate on which they are measured must itself be made horizontal. Whether it is so or not is known by means of two small levels placed on the plate at right angles to each other. Each consists of a glass tube, slightly curved upward in its middle, and so nearly filled with alcohol that only a small
bubble of air is left in the tube. This always rises to the highest part of the tubes. They are so "adjusted" that when this bubble of air is in the middle of the tubes, or its ends equidistant from the central mark, the plate on which they are fastened shall be level, which way soever it may be turned. One of the levels is sometimes fixed between the standards above one of the verniers, and the other on the plate at the north end of the compass-box.
296. Parallel Plates. To raise or lower either side of the circle, so as to bring the bubbles into the centers of the tubes, requires more gentle and steady movements than the unaided hands can give, and is attained by the parallel plates, and their four milled-headed screws, which hold the plates firmly apart, and, by being turned in or out, raise or lower one side or the other of the upper plate, and thereby of the graduated circle. The two plates are held together by a ball-and-socket joint. To level the instrument, loosen the lower clamp and turn the circle till each level is parallel to the vertical plane passing through a pair of opposite screws. Then take hold of two opposite screws and turn

## Fig. 299.


them simultaneously and equally, but in contrary directions, screwing one in and the other out, as shown by the arrows in the figures. A rule easily remembered is that both thumbs must turn in, or both out. The movements represented in the first of these figures would raise the left-hand side of the circle and lower the righthand side. The morements of the second figure would produce the reverse effect. Care is needed to turn the opposite screws equally, so that they shall not become so loose that the instrument will rock, or so tight as to be cramped. When this last occurs, one of the other pair should be loosened.

Sometimes one of each pair of the screws is replaced by a strong spring, against which the remaining screws act.

The French and German instruments, and most large instruments, are usually supported by only three screws. In such cases, one level is brought parallel to one pair of screws and leveled by them, and the other level has its bubble brought to its center by the third screw. If there is only one level on the instrument, it is first brought parallel to one pair of screws and leveled, and is then turned one quarter around so as to be perpendicular to them and over the third screw, and the operation is repeated.
297. Watch-Telescope. A second telescope is sometimes attached to the lower part of the instrument. When a number of angles are to be observed from any one station, direct the upper and principal telescope to the first object, and then direct the lower one to any other well-defined point. Then make all the desired observations with the upper telescope, and, when they are finished, look again through the lower one, to see that it and therefore the divided circle have not been moved by the movements of the vernier-plate. The French call this the Witness-Telescope (Lunette témoin).
298. The Compass. Upon the upper plate is fixed a compass. It has been fully explained in Chapter III. It is little used in connection with the transit, which is so incomparably more accurate, except as a "check," or rough test of the accuracy of the angles taken, which should about equal the difference of the magnetic bearings.
299. The Reflector. In making observations on Polaris at night, or in surveying mines, a reflector (Fig. 210) is used. This is a silvered plate with a hole in it for observing through with the telescope, while a light, held near the silvered surface, illuminates the cross-hairs. The reflector is attached to a ring, fitted to the object-glass slide, and is in-

Fig. 210.
 clined at an angle of $45^{\circ}$ to the ring.
300. The Diagonal Prism (Fig. 211). This is a prism attached to the eye-piece of the telescope, so that the

Fig. 211.
 rays of light, coming from the object sighted to, and passing through the telescope, are reflected to the eye at an angle of $90^{\circ}$ to the line of sight of the telescope. The prism is attached to a movable plate so that it can be turned to suit the position of the observer. This prism enables larger vertical angles to be measured than would be possible without it.

## The Transit.

301. The Engineer's Transit (Fig. 213). This instrument is similar in general construction to that shown in Fig. 199, but differs from it in several important particulars. The sockets for the axes of the plates are longer and differently arranged. These are shown in Fig. 212.

Both levels are attached to the upper plate. The verniers, instead of being placed at the sides between the legs of the standards,

Fig. 212.

as is usual, are placed near the north and south points of the compass-circle, so that the observer can read the rernier without stepping to the side of the instrument. The slow motion, both of the upper and lower plate, is given by one tangent-screw. In each
case an opposing spiral spring prevents any shake in the tangentscrew.

The vertical are is attached to the axis of the telescope by a clamp-screw, shown in the figure. The vernier and the slowmotion screw of the vertical arc are shown below the arc, and are attached to the left-hand standard.

Fig. 213.


Attached to the right-hand standard is the "Gradienter" (shown in detail in Fig. 245).
302. A vertical section through the body of the engineer's transit is given in Fig. 212. The lower plate, or "divided limb," B, is supported by the hollow socket C. Through this hollow socket passes the conical spindle which supports the upper plate A.

The upper plate carries the telescope, compass-box, and the verniers. The vernier-scales, $V, V$, are attached to the upper plate, but lie in the same plane as the divisions of the lower plate (so that the two can be viewed together without parallax), and are covered with glass to exclude dust. E is the clamp-screw.
303. The Theodolite. The transit, when furnished with a vertical circle and telescope level, is sometimes called a Theodolite. This name is used almost exclusively in England and on the Continent of Europe. In one form of the theodolite the telescope can

Fig. 214.
 not be revolved on its horizontal axis. This form has been almost entirely superseded in this country by that having a reversible telescope. It is then called a Transit Theodolite, or simply a Transit.
304. Goniasmometre. A very compact instrument, to which this name has been given in France, where it is much used, is shown in the figure. The upper half of the cylinder is movable on its lower half. The observations may be taken through the slits, as in the surveyor's cross, or a telescope may be added to it. Readings may be taken both from the compass and from the divided edge of the lower half of the cylinder, by means of a vernier on the upper half.*

[^36]
## VERNIERS.

305. Definition. A vernier is a contrivance for measuring smaller portions of space than those into which a line is actually divided. It consists of a second line or scale, movable by the side of the first, and divided into equal parts, which are a very little shorter or longer than the parts into which the first line is divided. This small difference is the space which we are thus enabled to measure.*

The vernier scale is usually constructed by taking a length equal to any number of parts on the divided line, and then dividing this length into a number of equal parts, one more or one less than the number into which the same length on the original line is divided.
306. Illustration. The figure represents (to twice the real size) a scale of inches divided into tenths, with a vernier scale beside it, by which hundredths of an inch can be measured. The vernier is

Fig. 215.

made by setting off on it nine tenths of an inch, and dividing that length into ten equal parts. Each space on the vernier is therefore equal to a tenth of nine tenths of an inch, or to nine hundredths of an inch, and is consequently one hundredth of an inch shorter than one of the divisions of the original scale. The first space of the vernier will therefore fall short of, or be overlapped by, the first

[^37]space on the scale by this one hundredth of an inch; the second space of the vernier will fall short by two hundredths of an inch; and so on. If, then, the vernier be moved up by the side of the original scale, so that the line marked 1 coincides, or forms one straight line, with the line of the scale which was just above it, we know that the rernier has been moved one hundredth of an inch. If the line marked 2 comes to coincide with a line of the scale, the

Fig. 216.

vernier has moved up two hundredths of an inch ; and so for other numbers. If the position of the rernier be as in this figure, the line marked 7 on the vernier corresponding with some line on the scale, the zero-line of the vernier is seren hundredths of an inch above the division of the scale next below this zero-line. If this division be, as in the figure, 8 inches and 6 tenths, the reading will be 8.67 inches.*

A vernier like this is used on some leveling-rods, being engraved on the sides of the opening in the part of the target above its middle line. The rod being divided into hundredths of a foot, this vernier reads to thousandths of a foot. It is also used on some French mountain barometers, which are divided to hundredths of a metre, and thus read to thousandths of that unit.
307. General Rules. To find u'hat any vernier reads to-i. e., to determine how small a distance it can measure-obserre how many parts on the original line are equal to the same number increased or diminished by one on the rernier, and divide the length

[^38]of a part on the original line by this last number. It will give the required distance.*

For verniers as usually constructed, the following rule will apply : Divide the value of the smallest division on the original scale by the number of parts on the vernier.

For example, if the limb of a transit be divided into half degrees, and thirty parts on the vernier are equal to twenty-nine on the limb, then the value of the smallest division on the limb (30 minutes), divided by the number of parts on the vernier (30) equals one minute. This is what the vernier reads to.

To read any vernier, first, look at the zero-line of the vernier (which is sometimes marked by an arrow-head), and if it coincides with any division of the scale, that will be the correct reading, and the vernier divisions are not needed. But if, as usually happens, the zero-line of the vernier comes between any two divisions of the scale, note the nearest next less division on the scale, and then look along the vernier till you come to some line on it which exactly coincides, or forms a straight line, with some line (no matter which) on the fixed scale. The number of this line on the vernier (the ryth, in the last figure) tells that so many of the subdivisions which the vernier indicates are to be added to the reading of the entire divisions on the scale.

When several lines on the rernier appear to coincide equally with lines of the scale, take the middle line.

When no line coincides, but one line on the vernier is on one side of a line on the scale, and the next line on the vernier is as far on the other side of it, the true reading is midway between those indicated by these two lines.
308. Retrograde Verniers. The spaces of the vernier in modern instruments are usually each shorter than those on the scale, a certain number of parts on the scale being divided into a larger number

[^39]of parts on the vernier.* In the contrary case, $\dagger$ there is the inconvenience of being obliged to number the lines of the vernier and to count their coincidences with the lines of the scale, in a retrograde or contrary direction to that in which the numbers on the scale run. We will call such arrangements retrograde verniers.
309. Illustration. In this figure, the scale, as before, represents (to twice the real size) inches divided into tenths, but the vernier is made by dividing eleven parts of the scale into ten equal

Fig. 217.

parts, each of which is therefore one tenth of eleren tenths of an inch-i. e., cleven hundredths of an inch, or a tenth and a hundredth. Each space of the vernier therefore overlaps a space on the scale by one hundredth of an inch. The manner of reading this vernier is the same as in the last one, except that the numbers run in a reverse direction. The reading of the figure is 30•16.

This vernier is the one generally applied to the common barometer, the zero-point of the vernier being brought to the level of the top of the mercury, whose height it then measures. It is also employed for leveling-rods which read downward from the middle of the target.
310. Fig. 218 represents (to double size) the usual scale of the English mountain barometer. The scale is first dirided into inches. These are subdivided into tenths by the longer

$$
\text { * i. e., algebraically, } v=\frac{m}{m+1} . \quad \quad \text { i. e., when } v=\frac{m}{m-1} s .
$$

lines, and the shorter lines again divide these into half tenths, or to 5 hundredths; 24 of these smaller parts are set off on the ver-

Fig. 218.

nier, and divided into 25 equal parts, each of which is therefore $=\frac{24 \times \cdot 05}{25}=\cdot 048$ inch, and is shorter than a division of the scale by $\cdot 050-\cdot 048=\cdot 002$, or two thousandths of an inch, a twentyfifth part of a division on the scale, to which minuteness the vernier can therefore read. The reading in the figure is 30.686 ( $30 \cdot 65$ by the scale and $\cdot 036$ by the vernier), the dotted line marked, D showing where the coincidence takes place.
311. Circle divided in $\ddagger 0$ Degrees. The following illustrations apply to the measurements of angles, the circle being variously divided. In this article, the circle is supposed to be divided into degrees.

If 6 spaces on the vernier are found to be equal to 5 on the circle, the vernier can read to one sixth of a space on the circlei. e., to $10^{\prime}$.

If 10 spaces on the vernier are equal to 9 on the circle, the vernier can read to one tenth of a space on the circle-i. e., to $6^{\prime}$.

If 12 spaces on the vernier are equal to 11 on the circle, the vernier can read to one twelfth of a space on the circle-i. e., to $5^{\prime}$.

Fig. 219 shows such an arrangement. The index, or zero, of the vernier is at a point beyond $358^{\circ}$, a certain distance, which the coincidence of the third line of the vernier (as indicated by the dotted and crossed line) shows to be $15^{\prime}$. The whole reading is therefore $358^{\circ} 15^{\prime}$.

If 20 spaces on the vernier are equal to 19 on the circle, the vernier can read to one twentieth of a division on the circle-i. e.,

Fig. 219.

to 3 '. English compasses, or "circumferentors," are sometimes thus arranged.

If 60 spaces on the vernier are equal to 59 on the circle, the vernier can read to one sixtieth of a division on the circle-i. e., to $1^{\prime}$.
312. Circle divided to $30^{\prime}$. Such a graduation is a very common one. The vernier may be rariously constructed.

Fig. 220.


Suppose 30 spaces on the vernier to be equal to 29 on the circle. Each space on the vernier will be $=\frac{29 \times 30^{\prime}}{30}=29^{\prime}$, and will therefore 'be less than a space of the circle by 1 ', to which the vernier will then read.

Fig. 220 shows this arrangement. The reading is $0^{\circ}$, or $360^{\circ}$.
In Fig. 221 the dotted and crossed line shows what divisions coincide, and the reading is $20^{\circ} 10^{\prime}$; the vernier being the same as in the preceding figure, and its zero being at a point of the circle $10^{\prime}$ beyond $20^{\circ}$.

Fig. 221.


In Fig. 222, the reading is $20^{\circ} 40^{\prime}$, the index being at a point beyond $20^{\circ} 3^{\prime}$, and the additional space being shown by the vernier to be $10^{\prime}$.

Sometimes 30 spaces on the vernier are equal to 31 on the circle. Each space on the vernier will therefore be $=\frac{31 \times 30^{\prime}}{30}=31^{\prime}$, and will be longer than a space on the circle by $1^{\prime}$, to which it will therefore read, as in the last case, but the vernier will be "retrograde." This is the vernier of the compass. The peculiar manner in which it is there applied is shown in Fig. 229.

Fig. 222.


If 15 spaces on the vernier are equal to 16 on the circle, each space on the vernier will be $=\frac{16 \times 30^{\prime}}{15}=32^{\prime}$, and the rernier will therefore read to $2^{\prime}$.
313. Circle divided to $20^{\prime}$. If 20 spaces of the rernier are equal to 19 on the circle, each space of the latter will be $=$ $\frac{19 \times 20^{\prime}}{20}=19^{\prime}$, and the vernier will read to $20^{\prime}-19^{\prime}=1^{\prime}$.

If 40 spaces on the vernier are equal to 41 on the circle, each
Fig. 223.

space on the vernier will be $=\frac{41 \times 20^{\prime}}{40}=20 \frac{1}{2}^{\prime}$, and the vernier will therefore read to $20 \frac{1}{2}^{\prime}-20^{\prime}=30^{\prime \prime}$. It will be retrograde. In Fig. 223 the reading is $360^{\circ}$, or $0^{\circ}$; and it will be seen that the

Fig. 224.


40 spaces on the vernier (numbered to whole minutes) are equal to $13^{\circ} 40^{\prime}$ on the limb-i. e., to 41 spaces, each of $20^{\prime}$.

If 60 spaces on the vernier are equal to 59 on the circle, each of the former will be $=\frac{59 \times 20}{60}=19^{\prime} 40^{\prime \prime}$, and the vernier will Fig. 225.

20


10
 $\begin{array}{llllllllll}10 & 9 & 3 & 7 & 6 & 5 & 4 & 3 & 2 & 1\end{array}$

therefore read to $20^{\prime}-19^{\prime} 40^{\prime \prime}=20^{\prime \prime}$. Fig. 224 shows such an arrangement. The reading in that position would be $40^{\circ} 46^{\prime} 20^{\prime \prime}$.
314. Circle divided to 15 . If 60 spaces on the vernier are equal to 59 on the circle, each space on the vernier will be $=$ $\frac{59 \times 15^{\prime}}{60}=14^{\prime} 45^{\prime \prime}$, and the vernier will read to $15^{\prime \prime}$. In Fig. 225 the reading is $10^{\circ} 20^{\prime} 45^{\prime \prime}$, the index pointing to $10^{\circ} 15^{\prime}$, and something more, which the vernier shows to be $5^{\prime} 45^{\prime \prime}$.
315. Circle divided to $\mathbf{1 0}^{\prime}$. If 60 spaces on the vernier be equal to 59 on the limb, the vernier will read to $10^{\prime \prime}$. In Fig. 226 the reading is $y^{\circ} 25^{\prime} 40^{\prime \prime}$, the reading on the circle being $7^{\circ} 20^{\prime}$, and the vernier showing the remaining space to be $5^{\prime} 40^{\prime \prime}$.

Fig. 226.

316. Reading backward. When an index carrying a vernier is moved backward, or in a contrary direction to that in which the numbers on the circle run, if we wish to read the space which it has passed over in this direction from the zero-point, the rernier must be read backward (i. e., the highest number be called 0 ), or its actual reading must be subtracted from the ralue of the smallest space on the circle. The reason is plain ; for, since the rernier
shows how far the index, moving in one direction, has gone past one division-line, the distance which it is from the next divisionline (which it may be supposed to have passed, moving in a contrary direction) will be the difference between the reading and the value of one space.

Thus, in Fig. 219, the reading is $358^{\circ} 15^{\prime}$. But, counting backward from the $360^{\circ}$, or zero-point, it is $1^{\circ} 45^{\prime}$.

Caution on this point is particularly necessary in using small angles of deflection for railroad-curves.

31\%. Arc of Excess. On the sextant and similar instruments, the divisions of the limb are carried onward a short distance beyond the zero-point. This portion of the limb is called the "Arc of Excess." When the index of the vernier points to this arc, the

Fig. 227.

reading must be made as explained in the last article. Thus, in the figure, the reading on the arc from the zero of the limb to the zero of the vernier is $4^{\circ} 20^{\prime}$, and something more, and the reading of the vernier from 10 toward the right, where the lines coincide, is $3^{\prime} 20^{\prime \prime}$ (or it is $10^{\prime}-6^{\prime} 40^{\prime \prime}=3^{\prime} 20^{\prime \prime}$ ), and the entire reading is therefore $4^{\circ} 23^{\prime} 20^{\prime \prime}$.
318. Double Verniers. To avoid the inconveniences of reading backward, double verniers are sometimes used. Fig. 228
shows one applied to a transit. Each of the verniers is like the one described in Art. 312, Figs. 220, 221, and 222. When the degrees are counted to the left, or as the numbers run, as is usual,

Fig. 228.

the left-hand vernier is to be read, as in Art. 312 ; but when the degrees are counted to the right, from the $360^{\circ}$ line, the righthand vernier is to be used.
319. Compass-Vernier. Another form of double rernier, often applied to the compass, is shown in Fig. 229. The limb is

Fig. 229.

divided to half-degrees, and the rernier reads to minutes, 30 parts on it being equal to 31 on the limb. But the rernier is only half as long as in the previous case, going only to $15^{\prime}$, the upper figures on one half being a sort of continuation of the lower figures on the other half. Thus, in moring the index to the right, read the lower figures on the left-hand vernier (it being retrograde)
at any coincidence, when the space passed over is less than $15^{\prime}$; but if it be more, read the upper figures on the right-hand vernier, and vice versa.

## ADJUSTMIENTS.

320. The purposes for which the transit (as well as most surveying and astronomical instruments) is to be used, require and presuppose certain parts and lines of the instrument to be placed in certain directions with respect to others ; these respective directions being usually parallel or perpendicular. Such arrangements of their parts are called their Adjustments. The same word is also applied to placing these lines in these directions. In the following explanations the operations which determine whether these adjustments are correct, will be called their Verifications ; and the making them right, if they are not so, their Rectifications.*
321. In observations of horizontal angles with the transit it is required-
322. That the circular plates shall be horizontal in whatever way they may be turned around.
323. That the telescope, when pointed forward, shall look in precisely the reverse of its direction when pointed backward-i. e., that its two lines of sight (or lines of collimation) forward and backward shall lie in the same plane.
324. That the telescope, in turning upward or downward, shall move in a truly vertical plane, in order that the angle measured between a low object and a high one may be precisely the same as would be the angle measured between the low object and a point exactly under the high object, and in the same horizontal plane as the low one.

We shall see that all these adjustments are finally resolvable into these : 1 . Making the vertical axis of the instrument perpendicular to the plane of the levels ; 2. Making the line of collima-

[^40]tion perpendicular to its axis ; and, 3. Making this axis parallel to the plane of the levels. They are all best tested by the invaluable principle of "reversion."

We have now, first, to examine whether these things are sothat is, to "verify" the adjustments; and, second, if we find that they are not so, to make them so-i. e., to "rectify" or "adjust" them correctly. The above three requirements produce as many corresponding adjustments.
322. First Adjustment. To cause the circle to be horizontal in every position.

Verification. Turn the vernier-plate, which carries the levels, till one of them is parallel to one pair of the parallel plate-screms. The other will then be parallel to the other pair. Bring each bubble to the middle of its tube, by that pair of screms to which it is parallel. Then turn the vernier-plate half-way around-i. e., till the index has passed over $180^{\circ}$. If the bubbles remain in the centers of the tubes, they are in adjustment. If either of them runs to one end of the tube, it requires rectification.

Rectification. The fault which is to be rectified is that the plane of the lerel (i. e., the plane tangent to the highest point of the level tube) is not perpendicular to the vertical axis on which

the plate turns. For, let A B represent this plane, seen edgemise, and CD the center line of the rertical axis, which is here drawn as making an acute angle with this plane on the right-hand side. The first figure represents the bubble brought to the center of the tube. The second figure represents the plate turned half around. The center line of the axis is supposed to remain unmored. The
acute angle will now be on the left-hand side, and the plate will no longer be horizontal ; consequently, the bubble will run to the higher end of the tube. The rectification necessary is evidently to raise one end of the tube and lower the other. The real error has been doubled to the eye by the reversion. Half of the motion of the bubble was caused by the tangent plane not being perpendicular to the axis, and half by this axis not being vertical. Therefore, raise or lower one end of the level by the screws which fasten it to the plate, till the bubble comes about half-way back to the center, and then bring it quite back by turning its pair of parallel platescrews. Then again reverse the vernier-plate $180^{\circ}$. The bubble should now remain in the center. If not, the operation should be repeated. Thre same must be done with the other level, if required. Then the bubbles will remain in the center during a complete revolution. This proves that the axis of the vernier-plate is then vertical ; and, as it has been fixed by the maker perpendicular to the plate, the latter must then be horizontal.

It is also necessary to examine whether the bubbles remain in the center, when the divided circle is turned round on its axis. If not, the axes of the two plates are not parallel to each other. The defect can be remedied only by the maker ; for, if the bubbles be altered so as to be right for this reversal, they will be wrong for the vernier-plate reversal.
323. Second Adjustment. To cause the line of collimation to revolve in a plane.

Verification. Set up the transit in the middle of a level piece of ground, as at A in the figure. Level it carefully. Set a stake,

Fig. 232.

with a nail driven into its head, or a chain-pin, as far from the instrument as it is distinctly visible, as at B. Direct the telescope
to it, and fix the intersection of the cross-hairs very precisely upon it. Clamp the instrument. Measure from A to B. Then turn over the telescope, and set another stake at an equal distance from the transit, and also precisely in the line of sight. If the line of collimation has not continued in the same plane during its halfrevolution, this stake will not be at E , but to one side, as at C . To discover the truth, loosen the clamp and turn the vernier-plate half around without touching the telescope. Sight to B, as at first, and again clamp it. Then turn orer the telescope, and the line of sight will strike, as at D in the figure, as far to the right of the point as it did before to its left.

Rectification. The fault which is to be rectified is that the line of collimation of the telescope is not perpendicular to the horizontal axis on which the telescope revolves. This will be seen by

the figures, which represent the position of the lines in each of the four observations which have been made. In each of the figures the long, thick line represents the telescope, and the short one the axis on which it turns. In Fig. 233 the line of sight is directed to B. In Fig. 234 the telescope has been turned over, and with it the axis, so that the obtuse angle marked 0 in the first figure has taken the place, $0^{\prime}$, of the acute angle, and the telescope points to C instead of to E. In Fig. 235 the rernier-plate has been turned
half around so as to point to B again, and the same obtuse angle has got around to $0^{\prime \prime}$. In Fig. 236 the telescope has been turned over, the obtuse angle is at $0^{\prime \prime \prime}$, and the telescope now points to $D$.

To make the line of collimation perpendicular to the axis, the former must have its direction changed. This is effected by moving the vertical hair the proper distance to one side. By loosening the left-hand screw and tightening the right-hand one, the ring, and with it the cross-hairs, will be drawn to the right, and vice versa. Two holes at right angles to each other pass through the outer heads of the screws. Into these holes a stout steel wire is inserted, and the screws can thus be turned around. Screws so made are called "capstan-headed." One of the other pair of screws may need to be loosened to avoid straining the threads. In some French instruments, one of each pair of screws is replaced by a spring.

To find how much to move this vertical hair, measure from C to D, Fig. 232: Set a stake at the middle point E, and set another at the point F, midway between D and E. Move the vertical hair till the line of sight strikes F . Then the instrument is adjusted ; and, if the line of sight be now directed to E , it will strike B when the telescope is turned over, since the hair is moved half of the doubled error, DE. The operation will generally require to be repeated, not being quite perfected at first.

It should be remembered that, if the telescope used does not invert objects, its eye-piece will do so. Consequently, with such a telescope, if it seems that the vertical hair should be moved to the left, it must be moved to the right, and vice versa. An inverting telescope does not invert the cross-hairs.

If the young surveyor has any doubts as to the perfection of his rectification, he may set another stake exactly under the instrument by means of a plumb-line suspended from its center; and then, in like manner, set his transit over B or E. He will find that the other two stakes, A and the extreme one, are in the same straight line with his instrument.

In some instruments, the horizontal axis of the telescope can be taken out of its supports and turned over, end for end. In such a case, the line of sight may be directed to any well-defined point,
and the axis then taken out and turned over. If the line of sight again strikes the same point, this line is perpendicular to the axis.

Fig. 237.


If not, the apparent error is double the real error, as appears from the figures, the obtuse angle 0 coming to $0^{\prime}$, and the desired perpendicular line falling at C midway between B and $\mathrm{B}^{\prime}$. The rectification may be made as before ; or, in some large instruments, in which the telescope is supported on $Y_{S}$, by moving one of the $Y_{s}$ laterally.
324. Third Adjustment. To cause the line of collimation to revolve in a vertical plane.

Verification. Suspend a long plumb-line from some high point. Set the instrument near this line, and level it carefully. Direct the telescope to the plumb-line, and see if the intersection of the cross-hairs follows and remains upon this line when the telescope is turned up and down. If it does, it mores in 2 , rertical plane.

The angle of a new and well-built house will form an imperfect substitute for the plumb-line.

Otherwise: The instrument being set up and leveled as abore, place a basin of some reflecting liquid (quicksilver being the best, though molasses, or oil, or even water will answer, though less perfectly) so that the top of a steeple, or other point of a high object, can be seen in it through the telescope by reflection. Make the intersection of the cross-hairs cover it. Then turn up the telescope, and, if the intersection of the cross-hairs bisects also the object seen directly, the line of sight has mored in a rertical plane. If a star be taken as the object, the star and its reflection will be
equivalent (if it be nearly overhead) to a plumb-line at least fifty million million miles long.

Otherwise: Set the instrument as close as possible to the base of a steeple or other high object ; level it, and direct it to the top of the steeple, or to some other elevated and welldefined point. Clamp the plates. Turn down the telescope, and set up a pin in the ground precisely "in line." Then loosen the clamp, turn over the telescope, and turn it half-way around, or so far as to again sight to the high point. Clamp the plates, and again turn down the telescope. If the line of sight again strikes the pin, the telescope has moved in a vertical plane. If not, the apparent error is double the real error. For, let S be the top of the steeple (Fig. 239),

Fig. 239.
 and $\mathrm{P}^{\prime}$ the pin ; then the plane in which the telescope moves, seen edgewise, is $\mathrm{SP}^{\prime}$; and, after being turned around, the line of sight moves in the plane $\mathrm{S} \mathrm{P}^{\prime \prime}$, as far to one side of the vertical plane $\mathrm{S} P$ as $\mathrm{S} \mathrm{P}^{\prime}$ was on the other side of it.

Rectification. Since the second adjustment causes the line of sight to move in a plane perpendicular to the axis on which it turns, it will move in a vertical plane if that axis be horizontal. It can be made so by raising or lowering one end of the axis by means of a screw placed in the standard for that purpose.
325. Centering Eye-Piece. In some instruments, such as that of which a longitudinal section is shown in the margin, the inner end of the eye-piece may be moved so that the cross-hairs shall be seen precisely in the center of its field of view. This is done by means of four screws, arranged in pairs, like those of the cross-hair ring-screws, and capable of moving the eye-piece up and down, and to right or left, by loosening one and tightening the opposite one. Two of them are shown at A, A, in the figure, in which B, B, are two of the cross-hair screws.
326. Centering Object-Glass. In some instruments four screws, similarly arranged, two of which are shown at C, C, can move, in any direction, the inner end of the slide which carries the object-
glass. The necessity for such an arrangement arises from the

Fig. 240.
 impossibility of drawing a tube perfectly straight. Consequently, the line of collimation, when the tube is drawn in, will not coincide with the same line when the tube is pushed out. If adjusted for one position, it will therefore be wrong for the other. These screws, however, can make it right in both positions. They are used as follows:

Sight to some well-defined point as far off as it can be distinctly seen. Then, having the plates firmly clamped, move out the ob-ject-glass slide, and fix a point in the line of sight as close to the instrument as can be distinctly seen. Then turn the limb halfway around horizontally, reverse the telescope, and again sight to the near point, by clamping the plates and bringing the rertical cross-hair on the point by means of the tangent-screw. Then draw in the objectglass slide until the distant object is distinctly seen. If the rertical cross-hair bisects it, no adjustment is necessary. If not, correct one half of the apparent error by means of the screws C C in Fig. 240. This may disturb the second adjustment. Try that over again, and again perform the operation of centering the object-glass.

This adjustment is always performed by the maker, and its screws are covered by a short tube.

All the adjustments should be meddled with as little as possible, lest the screws should get loose ; and, when once made right, they should be kept so by careful usage.

32\%. Fourth Adjustment. To cause the line of collimation of the telescope to be horizontal when the bubble of the level attached to it is in the center of its tube.

Drive two pegs several hundred feet apart, and set the instrument midway between them. Level, and sight to the rod held on each peg. The difference of the readings will be the true difference of the heights of the pegs, no matter how much the level may be out of adjustment.

Then set the instrument over one peg, and sight to the rod held at the other. Measure the height of the cross-hairs above the first peg. The difference of this height and the reading on the rod should equal the difference of the heights of the two points, as previously determined. If it does not, set the target to the sum or difference of the height of, the cross-hairs above the first peg and the true difference of height of the points, according as the first point is higher or lower than the second, and hold the rod on the second point. Sight to it, and raise or lower one end of the bubbletube until the horizontal cross-hair does bisect the target when the bubble is in the center.

Instead of setting over one peg, it is generally more convenient to set near to it, and sight to a rod held on it, and use this reading instead of the measured height of the cross-hairs.
328. Fifth Adjustment. To make the vernier of the vertical circle read zero when the bubble of the telescope-level is in the center.

This is verified in various ways:

1. By simple inspection.
2. By reversion. Sight to some point. Note the reading on the vertical circle. Turn the telescope half-way around horizontally. Turn over the telescope and again observe the same point, and note the reading. Half the difference (if any) of the two readings is the error.
3. By reciprocal observations. Observe successively from each of two points to the other. Half the difference of the readings equals the index-error.

When the verification has been made, the error may be rectified
on the instrument by moving the vernier-plate, or the circle, or noted as a correction to each obserration when the instrument is large and delicate.

## THE FIELD-WORK.

329. To measure a Horizontal Angle. Set up the instrument so that its center shall be exactly over the angular point, or in the intersection of the two lines whose difference of direction is to be measured, as at $B$ in the figure. A plumb-line must be suspended from under the center. Dropping a stone is an imperfect substi-

Fig. 241.
 tute for this. Set the instrument so that its lower parallel plate may be as nearly horizontal as possible. The levels will serve as guides if the four parallel-plate screws be first so screwed up or down that equal lengths of them shall be above the upper plate. Then level the instrument carefully. Direct the telescope to a rod, stake, or other object, A in the figure, on one of the lines which form the angle. Tighten the clamps, and by the tangent-screw move the telescope so that the intersection of the cross-hairs shall rery precisely bisect this object. Note the reading of the rernier. Then loosen the clamp of the vernier, and direct the telescope on the other line (as to C) precisely as before, and again read. The difference of the two readings will be the desired angle, A B C. Thus, if the first reading had been $40^{\circ}$ and the last $190^{\circ}$, the angle mould be $150^{\circ}$. If the vernier had passed $360^{\circ}$ in turning to the second object, $360^{\circ}$ should be added to the last reading before subtracting. Thus, if the first reading had been $300^{\circ}$, and the last reading $90^{\circ}$, the angle would be found by calling the last reading, as it really is, $360^{\circ}+90^{\circ}=450^{\circ}$, and then subtracting $300^{\circ}$.

It is best to sight first to the left-hand object and then to the right-hand one, turning "with the sun" or like the hands of a watch, since the numbering of the degrees usually runs in that direction.

It is convenient, though not necessary, to begin by setting the
vernier at zero by the upper movement (that of the vernier-plate on the circle), and then, by means of the lower motion (that of the whole instrument on its axis), to direct the telescope to the first object. Then fasten the lower clamp, and sight to the second object as before. The reading will then be the angle desired. An objection to this is that the two verniers seldorn read alike.*

After one or more angles have been observed from one point, the telescope must be directed back to the first object, and the reading to it noted, so as to make sure that it has not slipped. A watch-telescope renders this unnecessary.

The error arising from the instrument not being set precisely over the center of the station will be greater the nearer the object sighted to. Thus, a difference of one inch would cause an error of only $3^{\prime \prime}$ in the apparent direction of an object a mile distant, but one of nearly $3^{\prime}$ at a distance of a hundred feet.
330. Reduction of High and Low Objects. When one of the objects sighted to is higher than the other, the "plunging telescope" of this instrument causes the angle measured to be the true horizontal angle desired-i. e., the same angle as if a point exactly under the high object and on a level with the low object (or vice versa) had been sighted to. For the telescope has been caused to move in a vertical plane by the third adjustment, and the angle measured is therefore the angle between the vertical planes which pass through the two objects, and which "project" the two lines of sight on the same horizontal plane.

This constitutes the great practical advantage of these instruments over those which are held in the planes of the two objects observed, such as the sextant and the "circle," much used by the French.
331. Notation of Angles. The angles observed may be noted in various ways. Thus, the observation of the angle ABC, in

[^41]Fig. 241, may be noted "At B, from A to C, $150^{\circ}$," or, better, "At B, between A and $\mathrm{C}, 150^{\circ}$." In column form, this becomes

$$
\begin{gathered}
\text { Between } \mathrm{A} 150^{\circ} \text { and } \mathrm{C} . \\
\text { At } \left\lvert\, \begin{array}{c}
\mathrm{B}
\end{array}{ }^{2}\right.
\end{gathered}
$$

When the vernier had been set at zero before sighting to the first object, and other objects were then sighted to, those objects, the readings to which were less than $180^{\circ}$, will be on the left of the first line, and those to which the readings were more than $180^{\circ}$ will be on its right, looking in the direction in which the surrey is proceeding, from $A$ to $B$, and so on.

In surveying a farm, the angle of deflection at station, or the traverse angle, may be noted, together with the lengths of the courses.
332. To repeat an Angle. Begin as in Art. 329, and measure the angle as there directed. Then unclamp below, and turn the circle around till the telescope is again directed to the first object, and made to bisect it precisely by the lower tangent-screw. Then unclamp above and turn the vernier-plate till the telescope again points to the second object, the first reading remaining unchanged. The angle will now have been measured a second time, but on a part of the circle adjoining that on which it was first measured, the second are beginning where the first ended. The difference between the first and last readings will therefore be twice the angle.

This operation may be repeated a third, a fourth, or any number of times, always turning the telescope back to the first object by the lower movement (so as to start with the reading at which the preceding observation left off), and turning it to the second object by the upper morement. Take the difference of the first and last readings and divide by the number of obserrations.

The advantage of this method is that the errors of observation (i. e., sighting sometimes to the right and sometimes to the left of the true point) balance each other in a number of repetitions, while the constant error of graduation is reduced in proportion to this number. This beautiful principle has some imperfections in practice, probably arising from the slipping and straining of the clamps.
333. Angles of Deflection. The angle of deflection of one line from another is the angle which one line makes with the other line produced. Thus, in the figure, the angle of deflection of $B C$ from $A B$ is $\mathrm{B}^{\prime} \mathrm{B} C$. It is evidently the supplement of the angle

Fig. 242.
 A B C.

To measure it with the Transit, set the instrument at B , direct the telescope to $A$, and then turn it over. It will now point in the direction of $A B$ produced, or to $B^{\prime}$, if the second adjustment has been performed. Note the reading. Then direct the telescope to C. Note the new reading, and their difference will be the required angle of deflection, $\mathrm{B}^{\prime} \mathrm{B} \mathrm{C}$.

If the vernier be set at zero before taking the first observation, the readings for objects on the right of the first line will be less than $180^{\circ}$, and more than $180^{\circ}$ for objects on the left, conversely to Art. 331.
334. Line-Surveying. The survey of a line, such as a road, etc., can be made by the transit with great precision, measuring the angle which each line makes with the preceding line, and noting their lengths, and the necessary offsets on each side.

Short lines of sight should be avoided, since a slight inaccuracy in setting the center of the instrument exactly over or under the point previously sighted to would then much affect the angle. Very great accuracy can be obtained by using three tripods. One would be set at the first station and sighted back to from the instrument placed at the second station, and a forward sight be then taken to the third tripod placed at the third station. The instrument would then be set on this third tripod, a back-sight taken to the tripod remaining on the second station, and a foresight taken to the tripod brought from the first station to the fourth station, to which the instrument is next taken, and so on. This is especially valuable in surveys of mines.

The field-notes may be taken as directed in compass-surveying, the angles taking the place of the bearings. The "checks by
intersecting bearings," before explained, should also be employed. The angles made on each side of the stations may both be measured, and the equality of their sum to $360^{\circ}$ would at once prove the accuracy of the work.

If the magnetic bearing of any one of the lines be given, and that of any of the other lines of the series be required, it can be deduced by constructing a diagram, or by modifications of the rules given for the reverse object.
335. Traversing ; or, surveying by the Back-Angle. This is a method of observing and recording the different directions of successive portions of a line

Fig. 243.
 (such as a road, the boundaries of a farm, etc.), so as to read off on the instrument, at each station, the angle which each line makes-not with the preceding line-but with the first line observed, or some other constant line. This line is, therefore, called the meridian of that survey.

The operation consists essentially in taking each back-sight by the lower motion (which turns the circle without changing the reading), and taking each forward sight by the upper motion, which moves the vernier orer the are measuring the new angle; and thus adds it to or subtracts it from the previous reading.

Set up the instrument at some station, as B ; put the rernier at zero, and, by the lower motion, sight back to A. Tighten the lower clamp, reverse the telescope, loosen the upper clamp, sight to C by the upper motion, and clamp the rernier-plate again. Remove the instrument to C , sight back to B by the lower motion. Then clamp below, reverse the telescope, loosen the upper clamp, and sight to $D$ by the upper motion. Then go to $D$ and proceed as at C ; and so on. The reading gires the angles measured to the right or "with the sun," as shown by the ares in the figure.

Care should be taken to keep the same side of the instrument
ahead, and, if only one vernier is read, to read from the same vernier.

The chief advantage of this method is its greater rapidity in the field and in platting, the angles being all laid down from one meridian, as in compass-surveying.
336. Use of the Compass. The chief use of the compass attached to a transit is as a check on the observations; for the difference between the magnetic bearings of any two lines should be the same, approximately, as the angle between them, measured by the more accurate instruments. The bearing also prevents any ambiguity as to whether an angle was taken to the right or to the left.

The instrument may also be used like a simple compass, the telescope taking the place of the sights, and requiring similar tests of accuracy. A more precise way of taking a bearing is to turn the plate to which the compass-box is attached, till the needle points to zero, and note the reading of the vernier ; then sight to the object, and again read the vernier. The bearing will thus be obtained more minutely than the divisions on the compass-box could give it.

33\%. Ranging out Lines. This is the converse of surveyinglines. The instrument is fixed over the first station with great precision, its telescope being very carefully adjusted to move in a vertical plane. A series of stakes, with nails driven in their tops, or otherwise well defined, are then set in the desired line as far as the power of the instrument extends. It is then taken forward to a stake three or four from the last one set, and is fixed over it, first by the plumb and then by sighting backward and forward to the first and last stake. The line is then continued as before. A good object for a long sight is a board painted like a target, with black and white concentric rings, and made to slide in grooves cut in the tops of two stakes set in the ground about in the line. It is moved till the vertical hair bisects the circles (which the eye can determine with great precision), and a plumb-line dropped from their center gives the place of the stake. "Mason and Dixon's Line" was thus ranged.

When the transit is used for ranging, its "Second Adjustment" is most important, to insure the accuracy of the reversal of its telescope.
338. Farm-Surveying, etc. A farm can be much more accurately surveyed with the transit than with the compass. The farm should be kept on the right hand, and then the angles measured will be the supplements of the interior angles. If the angles to the right be called positive, and those to the left negative, their algebraic sum should equal $360^{\circ}$.

If the boundary-lines be surveyed by "Traversing," the reading, on getting back to the last station and looking back to the first line, should be $360^{\circ}$, or $0^{\circ}$.

The content of any surface surveyed by "Trarersing" with the transit can be calculated by the traverse-table, by the following modification : When the angle of deflection of any side from the first side, or meridian, is less than $90^{\circ}$, call this angle the bearing, find its latitude and departure, and call them both plus. When the angle is between $90^{\circ}$ and $180^{\circ}$, call the difference between the

Fig. 244.

angle and $180^{\circ}$ the bearing, and call its latitude minus and its departure plus. When the angle is between $180^{\circ}$ and $270^{\circ}$, call its difference from $180^{\circ}$ the bearing, and call its latitude minus and its departure minus. When the angle is more than $2 \% 0^{\circ}$, call its difference from $360^{\circ}$ the bearing, and call its latitude plus and its departure minus. Then use these as in getting the content of a compass-survey. The signs of the latitudes and departures follow those of the cosines and sines in the successive quadrants.

Fig. 244 is a plat of the survey worked out in Art. 255.
The following table gives the deflection angle at each station, the traverse angle (i. e., the angle which each line makes with the first one), and the reduced bearing, calling the first line (1 to 2 ) the meridian :

| statrons. | deflection anales. | traverse angles. | bearings. |
| :---: | :---: | :---: | :---: |
| 1 | $911^{\circ}$ | $0^{\circ}$ or $360^{\circ}$ | North. |
| ${ }_{3}{ }^{2}$ | $481^{10}$ | $48 \frac{1}{2}^{\circ}$ | N. $481^{\circ}{ }^{\circ} \mathrm{E}$. |
| 3 | ${ }^{3911^{\circ}}$ |  | N. $88^{\circ} \mathrm{E}$. |
| $\stackrel{4}{5}$ |  | $\begin{aligned} & 1799^{\circ} \\ & 2683_{2}^{\circ} \end{aligned}$ |  |

If the deflection angle at station $1\left(91 \frac{1}{2}^{\circ}\right)$ be added to the traverse angle at station 5, the sum will be $360^{\circ}$.

Any side may be taken as the meridian of the survey.
If the true bearing of one side be known, the true bearings of the other sides may be determined by Art. 189.

The content is calculated by latitudes and departures, as in compass-surveying.

The latitudes and departures may be taken from the tables, interpolating for minutes as in Art. 242, or they may be calculated with a table of natural sines and cosines, as in Art. 240.

## Example. <br> FIELD-BOOK.

| stations. | Angles of deflection. | Distances in chains. |
| :---: | :---: | :---: |
| 1 | $62^{\circ} 15^{\prime}$ | 4.64 |
| 2 | $86^{\circ} 38^{\prime}$ | $3 \cdot 60$ |
| 3 | $59^{\circ} 20^{\prime}$ | $4 \cdot 15$ |
| 4 | $80^{\circ} 6^{\prime}$ | $4 \cdot 22$ |
| 5 | $71^{\circ} 41^{\prime}$ | 3.25 |

CALCULATION OF AREAS，CALLING COURSE 1 TO 2 THE MERIDIAN，AND USING SINES AND COSINES INSTEAD OF A TRAVERSE TABLE．

| $\begin{aligned} & \text { 离 } \\ & \text { Z } \\ & \text { 旨 } \\ & \text { 会 } \end{aligned}$ | bearings． |  |  | $\begin{aligned} & \text { 曷 } \\ & Z y_{0}^{2} \\ & 0 \end{aligned}$ | Lati－ |  | DEPART•URES． |  | double <br> Longi－ <br> tudes． | DOUble AREAS． |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  | ＋ | － |  | － |  | ＋ | － |
| 1 | $+00^{\circ}, 00^{\prime}+$ | $4 \cdot 64$ | －00000 | $1 \cdot 00000$ | $4 \cdot 64$ |  | $0 \cdot 00$ |  | $0 \cdot 00$ | $0 \cdot 0000$ |  |
| 2 | $+86^{\circ}, 38^{\prime}+$ | $3 \cdot 60$ | －99827 | －05873 | 0．21 |  | $3 \cdot 59$ |  | $+3 \cdot 59$ | $\cdot 7539$ |  |
|  | $-34^{\circ}, \quad 2^{\prime}+$ |  | －55968 | －82871 |  | $3 \cdot 44$ | $2 \cdot 32$ |  | ＋9•50 |  | $32 \cdot 6800$ |
| 4 | $-46^{\circ}, 4^{\prime}-$ |  | $\cdot 72015$ | －69382 |  | $2 \cdot 93$ |  | $3 \cdot 04$ | ＋8．78 |  | $25^{\prime} 7254$ |
| 5 | $+62^{\circ}, 15^{\prime}-$ |  | －88499 | $\cdot 46561$ | $1 \cdot 52$ |  |  | 2.87 | ＋2．87 | 4．3624 |  |
|  |  |  |  |  | $6 \cdot 37$ | 6．37 | $5 \cdot 91$ | 5．91 |  | $5 \cdot 1163$ | 58.4054 |
|  |  |  |  |  |  |  |  |  |  |  | $5 \cdot 1163$ |
|  |  |  |  |  |  |  |  |  |  |  | 53．2891 |
|  |  |  |  |  |  |  |  |  | quare | ains， | 26.6445 |

339．When the lengths of the sides are measured with an engi－ neer＇s chain，and the distances are determined in feet，the process of calculating the area is the same as for chains and decimals．The area is obtained in square feet instead of square chains，and to reduce it to acres it will be necessary to divide by 43560 ，the num－ ber of square feet in an acre．

340．Platting．Any of these surveys can be platted by any of the methods explained and characterized in Chapter III．A circu－ lar protractor may be regarded as a theodolite placed on the paper． ＂Platting Bearings＂can be employed when the surrey has been made by＂Traversing．＂But the method of＂Latitudes and de－ partures＂is by far the most accurate．

## THE GRADIENTER．

341．This is an attachment to the transit for determining grades and distances．It consists of an arm，attached to the axis of the telescope，and a micrometer－screw，by means of which the more－ ment of the arm，and consequently of the telescope，can be accu－ rately measured．

The arm is placed on the inside of one of the standards，and is attacied to the telescope axis by means of a clamp－screm，so that it may be clamped or loosened at pleasure．

The method of measuring the movement of the arm is shown in Fig. 245.

C is a section of the axis of the telescope. B is the arm, which

is clamped to the axis by the screw D. M is the micrometer-screw. A is a lip projecting from a plate fastened to the standards.

The screw is accurately cut, so that one revolution of the screw will cause the horizontal cross-hair of the telescope to move over a given space (say one foot) on a rod held at a given distance, as 100 feet. The head of the screw is graduated into equal parts, usually 50 or 100. Above the graduated head is a scale so graduated that one revolution of the screw will move the head over one space on the scale. Thus the number of whole revolutions given to the screw may be read on the scale, and the parts of a revolution read on the graduated head.

The point of the screw presses against the lip, A, and is held firmly against it by the opposing spiral spring, S.

When the arm is made fast to the axis by the clamp-screw, D, and the gradienter-screw, M , is turned, it will turn the telescope vertically on its axis, and the distance which the horizontal cross-
hair will pass over on a rod, toward which the telescope is pointed, will vary directly with the distance from the transit to the rod.
342. To establish Grades. Let us suppose that one revolution of the gradienter-screw will move the horizontal cross-hair orer a space of one foot, on a rod held at a distance of 100 feet from the transit. Then, to set grades, we have only to level the telescope, clamp the gradienter-arm, and turn the micrometer-screw through as many divisions of the head (graduated into 100 parts) as there are hundredths of a foot rise or fall per hundred feet of horizontal distance ; raising the cross-hair for an up-grade, and lowering it for a down-grade. The line of sight will then be on the required grade.

If the transit be set over a point of the required grade-line, set the target on the rod at the height of the center of the telescopeaxis above the given point, and then the bottom of the rod, held at any point on the line, will be at a point in the desired gradeline when the horizontal cross-hair bisects the target.

Thus, if the grade is to be 1.64 feet per hundred, turn the screw one entire revolution and 64 of the divisions on the graduated head, and the line of sight will then be on the required grade.
343. To measure Distances. When the ground is level or approximately so, see what space on the rod the horizontal cross-hair moves over for one revolution of the gradienter-screw. Then the distance in feet will be equal to the space on the rod, expressed in feet and decimals, multiplied by 100 .

Thus, if the space on the rod, mored orer by the cross-hair
Fig. 246.
 for one rerolution of the gradienter-screw, was $4 \cdot 2 \%$ feet, the distance at which the rod was held was $42 \%$ feet.

For, in Fig. 246, let
A be the position of the transit; C B, the reading on the rod, held at a distance of 100 feet, for one revolution of the screw ; and DE the space passed orer on the rod for one rerolution of the screw
when the rod is held at the unknown distance AD. It is evident that the triangles ABC and ADE are similar, and that

$$
\begin{array}{r}
\text { C B : A B :: E D : A D, } \\
\text { or, } \quad 1: 100:: 4 \cdot 2 \%: 42 \%
\end{array}
$$

If the rod sighted to is only graduated to feet-as an ordinary transit-rod-find how many revolutions and parts of revolutions will move the horizontal cross-hair over a whole number of feet on the rod. Then, since one revolution of the screw will move the cross-hair over a space of one foot on the rod at a distance of 100 feet, we have the proportion : $A s$ the number of revolutions of the screw (whole numbers and decimals) is to 100 feet, so is the number of feet passed over on the rod by the cross-hair to the required distance. For, from Fig. 246 we have, as before: .

$$
\mathrm{CB}: \mathrm{AB}:: \mathrm{DE}: \mathrm{AD} .
$$

CB now represents what the reading on the rod (in feet and decimals), held at a distance of 100 feet, would be for the given number of revolutions: A B is $100^{\prime}, \mathrm{DE}$ is the reading on the rod in feet, and AD is the required distance.

Suppose, for example, the gradienter-screw be turned 1.25 time, and the space passed over on the rod by the cross-hair be 3 feet. Then we have :

$$
1 \cdot 25: 100:: 3: 240
$$

$\therefore$ The required distance is 240 feet.
Problem.-When no graduated rod is available, to determine a distance by using, in place of a rod, a stick whose length can afterward be measured.

On sloping ground, the methods given will apply, if the rod be held perpendicular to the line of sight. This, however, is not easily done. It will be better to apply methods specially adapted to sloping ground.
344. On Sloping Ground. In Fig. 247 , let A be the position of the tran-

sit ; $G$ the point over which it is set ; $C$ where the rod is held; A B a horizontal line through the axis of the telescope; A C the distance from the horizontal axis of the telescope to the foot of the rod; and CD the distance, on a vertical rod, passed over by the horizontal cross-hair for one revolution of the gradienter-screw. Let CF be perpendicular to A C, and D B to AB.

Represent the angle of elevation, BAC , by $e$, the angle CAD by $s$, and the distance DC by $k$. Then we have :

$$
\mathrm{D} \mathrm{~B}=\mathrm{DC}+\mathrm{CB}
$$

$\therefore \mathrm{AB} \tan .(s+e)=k+\mathrm{AB} \tan . e$, and $\mathrm{AB}=\frac{k}{\tan .(s+e)-\tan . e}$.
For convenience of computation, this may be put in another form. Add and subtract $100 \%$, and we have :

$$
\begin{aligned}
& \mathrm{AB}=100 k-100 k+\frac{k}{\tan \cdot(s+e)-\tan \cdot e} \\
& \quad \text { And, since } \tan . s=\frac{1}{100} \\
& \mathrm{AB}=100 k-k(100 \sin \cdot e+\cos . e) \sin \cdot e
\end{aligned}
$$

TABLE FOR GRADIENTER.

| ANGLE OF Elevation. | $\begin{gathered} (100 \mathrm{sin} . e+\cos . e) \\ \times \text { SIN.e. } \end{gathered}$ |
| :---: | :---: |
| $0^{\circ}$ | $\cdot 0$ |
| $1^{\circ}$ | $\cdot 1$ |
| $2^{\circ}$ | $\cdot 2$ |
| $3^{\circ}$ | $\cdot 3$ |
| $4^{\circ}$ | $\cdot 5$ |
| $5^{\circ}$ | -8 |
| $6^{\circ}$ | $1 \cdot 2$ |
| $7{ }^{\circ}$ | $1 \cdot 6$ |
| $8^{\circ}$ | $2 \cdot 1$ |
| $9^{\circ}$ | $2 \cdot 6$ |
| $10^{\circ}$ | $3 \cdot 2$ |
| $11^{\circ}$ | $3 \cdot 8$ |
| $12^{\circ}$ | $4 \cdot 5$ |
| $13^{\circ}$ | $5 \cdot 3$ |
| $14^{\circ}$ | $6 \cdot 1$ |
| $15^{\circ}$ | $7 \cdot 0$ |
| $16^{\circ}$ | $7 \cdot 9$ |
| $17^{\circ}$ | $8 \cdot 8$ |
| $18^{\circ}$ | $9 \cdot 8$ |
| $19^{\circ}$ | $10 \cdot 9$ |
| $20^{\circ}$ | $12 \cdot 0$ |

The quantity ( $100 \sin . e+\cos . e$ ) $\sin . e$, for angles from $1^{\circ}$ to $20^{\circ}$ will be found in the table for the gradienter. Hence the rule :

Multiply the rod-reading by 100 , and deduct the product of the rodreading by the tabular number corresponding to the angle of elevation, e. The result will be the horizontal distance A B.

Example. Angle of eleration, $4^{\circ}$; rod-reading, $2 \cdot 63$ feet.

$$
\begin{aligned}
& 2 \cdot 63 \times 100=263 \\
& 2.63 \times .5=1 \cdot 3 \\
& \text { Horizontal distance, } \\
& 261 \cdot \%
\end{aligned}
$$

The table for the correction is computed to tenths only, as the unaroid-
able errors in using the instrument would render any more exact computation useless.

For ordinary cases, when the angle of elevation is small, the computation for the distance and correction can be made mentally.
345. The horizontal distance, A B , is the one almost always required, as all measurements of distances in surveying and engineering should be made horizontally.

The distance from the transit to the point at which the rod is held (i. e., A C) is equal to the horizontal distance, A B, divided by cos. $e$.

The distance GC may be found by solving the triangle CAG, of which the sides $A G$ and $A C$, and the included angle $C A G$, are known.

When the angle $e$ is an angle of depression, the top of the rod is taken for the point $c$, and the distance CD is measured downward from the top of the rod.

In using the micrometer-screw, care must be taken, when measuring, to always turn the screw in the same direction, in order to aroid any lost motion in the screw. In determining the space passed over by the cross-hair for one revolution of the screw, set the screw back of the first reading, and bring it up by turning the screw in the same direction in which it is to be turned for making the measurement.

## THE STADIA OR TELEMETER.

346. On the cross-hair ring of the telescope stretch two more horizontal cross-hairs of spider-web or platinum wire, at equal distances above and below the original one. The two additional wires are called Stadia Wires. The stadia wires may be either fixed or adjustable. In the former case they may be attached directly to the cross-hair ring. When they are adjustable, each may be fastened to a separate slide, actuated by a capstan-screw on the outside of the telescope-tube, as shown in Figs. 248 and 249.

The slides to which the stadia wires $b b$ and $c c$ are attached are held apart by the hoop-spring, shown in the figure, and are adjusted by the capstan-screws $d d$.

It is erident that, in looking through the telescope at a graduated rod, a certain portion
 of the rod will be intercepted between the stadia wires, and that the greater the distance at which the rod is held, the longer will be the space on the rod intercepted by the stadia wires.

Referring to Art. 287, Fig. 201, we see that the objective of the telescope forms an image, $B$, of the arrow, $A$. A may represent the part of the rod intercepted by the stadia wires, and B the distance between the wires. The farther the rod is carried from the telescope, the nearer the image is formed to the objective. If the rod were at an infinite distance, the image would be formed at the principal focus of the objective.

Call the distance of the principal focus from the lens, $f$; the distance from the lens to the rod held at any point, $p$; the distance from the lens to the image, $q$; the space intercepted on the rod by the stadia wires, $k$; and the distance apart of the stadia wires, $a$.

As $p$ increases, $k$ increases, $q$ decreases, and $a$ remains constant. From similar triangles, Fig. 201, we hare:

$$
\begin{equation*}
p: q:: k: a, \tag{1.}
\end{equation*}
$$

and from the principles of optics-

$$
\begin{equation*}
\frac{1}{p}+\frac{1}{q}=\frac{1}{f} \tag{2.}
\end{equation*}
$$

From [1] $\frac{p}{q}=\frac{k}{a}$.
From [2] $\frac{p}{q}=\frac{p}{f}-1$.
$\therefore \frac{p}{f}-1=\frac{k}{a}$.
and $p=\frac{f}{a} k+f$.

Formula [3] is not perfectly accurate, as $p$ and $q$ are measured from the surface of the lens instead of its center, and the objective of the telescope is not a simple double-convex lens. It is, however, sufficiently exact for this purpose.

We see by the formula [3] that, as $f$ and $a$ are constants, the distance, $p$, from the objective to the rod is equal to the reading on the rod, multiplied by a constant quantity, plus the principal focal distance of the objective. To obtain the distance from the center of the instrument to the rod, it is also necessary to add the distance from the center of the instrument to the objective. Call this distance $c$. Then, for the distance from the center of the instrument to the rod, we have :

$$
\begin{equation*}
\text { distance }=\frac{f}{a} k+f+c . \tag{4.}
\end{equation*}
$$

The distance from the objective to the center of the instrument is not precisely the same for all lengths of sight. The farther off the object sighted to is, the nearer the image will be formed to the objective, and hence the objective must be drawn in, in order that the image may be formed at the cross-hairs. When the object sighted to is near, the image is formed farther from the objective, and the objective-slide must be moved out in order that the image may be formed at the cross-hairs. Hence, we see that the quantity $c$ is not rigidly constant. The difference in value, however, is not enough to be taken into consideration. A mean value of $c$ can be determined by sighting to some object at a distance of the mean length of sight (say five

Fig. 250.
 hundred feet), and then measuring the distance from the objective to the center of the telescope-axis.

34\%. Formula [4] is for level ground. For sloping ground, this must be modified. In Fig. 250 let A be the center of the telescopeaxis ; $C E$, the reading on the rod; $D$, the point on the rod where the center cross-hair intersects the rod; A B, the horizontal distance ; H, a point in front of the object-glass, and at a distance equal to its focal length; e, the angle of elevation; ML, perpendicular to the line of sight; $f, a, c$, and $k$ as in [4]. Then we have :

$$
\begin{gather*}
\mathrm{ML}=\mathrm{CE} \cos \cdot e=k . \cos \cdot e \text { and } \mathrm{H} \mathrm{D}=\frac{f}{a} k \cos \cdot e \\
\mathrm{HI}=\mathrm{H} \mathrm{D} \cos \cdot e=\frac{f}{a} k \cos .^{2} e \\
\mathrm{~A} \mathrm{~B}=\mathrm{A} \mathrm{~N}+\mathrm{N} \mathrm{~B}(=\mathrm{HI}) \\
\therefore \mathrm{AB}=(c+f) \cos \cdot e+\frac{f}{a} k \cos ^{2} e
\end{gather*}
$$

The height $\mathrm{BD}=\mathrm{AB}$ tan. $e$

$$
\begin{equation*}
\mathrm{B} \mathrm{D}=(c+f) \sin \cdot e+\frac{f}{a} k \frac{\sin \cdot 2 e}{2} \tag{6.}
\end{equation*}
$$

To find the value of $a$ in any case, measure off from the point over which the instrument is set a base-line, B (say one thousand feet), and hold the stadia-rod at the farther end. Let the reading on the rod be $k^{\prime}$.

$$
\begin{aligned}
& \text { Then, by [4] } \mathrm{B}=\frac{f}{a} k^{\prime}+f+c \\
& \text { and } a=\frac{f k^{\prime}}{\mathrm{B}-f-c}
\end{aligned}
$$

Substituting this value of $a$ in equations [5] and [6], we hare:
Horizontal distance $=(c+f) \cos . e+\frac{\hbar}{k^{\prime}}(\mathrm{B}-f-c) \cos { }^{2} e . \quad[\%]$
Difference of level $=(c+f) \sin . e+\frac{k}{2 k^{\prime}}(\mathrm{B}-f-c) \sin .2 e$. [8.]
348. The Stadia-Tables* giren in this volume were calculated from formulas [ 7 ] and [8], using the following ralues :

The measured base, $B=1,000$ feet, and $k^{\prime}=$ the reading on the rod for that distance-i. e., the distance indicated by the stadiareading is 1,000 feet.

[^42]$$
(c+f)=1 \cdot 4 \text { feet }
$$

The quantities in the columns headed $a$ and $b$ are computed respectively from the expressions $(c+f)$ cos. $e$, and $(c+f)$ sin. $e$, in the formulas. They are constant for all readings if the angle $e$ remains the same.

The horizontal distances, and the differences of level, are computed by the tables in a manner similar to that employed in calculating latitudes and departures with a table.

Example. Let $e=4^{\circ} 27^{\prime \prime}$, and $k=$ reading corresponding to 1,384 feet when the ground is horizontal.

Take from the table as follows :

HORIZONTAL DISTANCE.


## DIFFERENCE OF LEVEL.



The difference of level given by formula [8] is the difference in height between the instrument at A and the point where the central cross-hair strikes the rod at C . The difference between the height of the instrument above the ground, and the height of C above the ground, must be applied as a correction to the difference of level, obtained by the formula, to get the true difference of height of the ground at the instrument, and at the rod.
349. The stadia-wires may be adjusted to use with a rod already graduated to feet and decimals, or, if the wires are fixed, a rod may be graduated to suit the wires.

In the first case the wires are adjusted so that one foot is included between the wires at a given distance ( 50 or 100 feet) plus the constant $c$. Suppose the space included between the wires was one foot, at a distance from the center of the instrument of 100 feet $+c$. Then, if the reading on the rod held at some unknown distance was $3 \cdot 46$ feet, the distance would be 346 feet $+c$.

If the wires are fixed, measure off from the center of the instrument 500 feet $+c$, and note the space on the rod, intercepted by the cross-hairs at that distance. Divide this space into five equal
parts, subdivide the parts to tenths and hundredths, and graduate the remainder of the rod with similar divisions. This rod can then be used in the same way as the rod, graduated to feet, was in the first case. Suppose, on holding up this rod at an unknown distance, that the stadia-wires intercepted 3.67 of the parts. Then the distance is 367 feet $+c$.

The rod may be supplied with one or two targets, or may be used as a "speaking-rod"-that is, it may be graduated and marked so as to be read by the observer at the instrument.

For forms of targets, and methods of graduating and marking rods, see subject "Rods," Part II.
350. Several different formulas and methods have been used in stadia-surveying, depending upon the object and extent of the survey, and the degree of accuracy required. Another method is given in the following communication,* together with results in practice :

## 351. Results of Telemeter Traverse between Triangulation-Points on the Shores of Lake George, New York.

Instrument. Engineer's transit of W. \& L. E. Gurley. Focal length = 0.565 feet; distance of cross-wires from center of instrument $=0.13$ feet. One extra cross-wire was added to the diaphragm. At 103 feet from the center of the instrument, the distance included between the wires was found to be 1.0253 feet-

$$
\begin{align*}
\text { by the formula, } t & =0.01005 d-0.01 \text { feet, }  \tag{1.}\\
\text { or, } d & =99.48 t+1 \text { foot, } \tag{2.}
\end{align*}
$$

where $t=$ distance included between the wires at any distance, $d$, from the center of the instrument.

Stadia-Rod or Telemeter. This was graduated especially for the instrument from formula [1], the zero of graduation being displaced 0.01 foot to allow for the constant of the formula. The least reading of the rod was $2 \frac{1}{3}$ feet. Distances were estimated and recorded to single feet.

Circumstances of Measurement. Traverse-lines were run between trian-gulation-points; the distances between the latter were computed from the traverse and compared with the results from triangulation, in nine cases. The aggregate length of these nine lines was about $10_{\frac{1}{10}}^{\frac{4}{0}}$ miles.

Four closed traverses were run around islands, and the errors of closure were obtained.

The lines of sight generally passed orer mater, which circumstance mas favorable to precise reading.

[^43]The results of comparison are given below. They indicate that the constants used in graduating the telemeter-rod were not exactly obtained. The error of measurement averaged $+2 \cdot 2$ feet to 1,000 . If this allowance had been made in graduating the rod, or this constant error had been allowed for, the purely accidental errors would have been only $\pm 1.2$ foot to the 1,000 . The law of propagation of errors of length is favorable to close linear measurements with the telemeter upon traverse-lines, as was found to be actually the case here. In traverse-lines, the larger part of the total error is due to angular errors which overweigh the linear ones, unless exceptional means are taken to avoid this.

| (1) | (2) | (3) | (4) | (5) |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $\stackrel{\text { feet. }}{5183 \cdot 5}$ | $\begin{gathered} \text { feet. } \\ +12.8 \end{gathered}$ | $\begin{gathered} \text { feet. } \\ +2 \cdot 47 \end{gathered}$ | $\begin{gathered} \text { feet. } \\ +0.31 \end{gathered}$ | 9 | angulation-points as computed from traverse. |
| $3988 \cdot 0$ | + 7.5 | +1.88 | -0.28 | 7 | $(2)=$ Distance by traverse |
| $4925 \cdot 7$ | + $7 \cdot 6$ | +1.54 | $-0.62$ | 9 | minus distance by triangulation. |
| $8427 \cdot 8$ | $+11.7$ | $+1 \cdot 39$ | $-0.77$ | 17 |  |
| $2995 \cdot 0$ | $+15 \cdot 0$ | $+5 \cdot 01$ | $+2.85$ | 7 | (3) = Error to 1,000 feet, including constant error. |
| $3104 \cdot 6$ | + $9 \cdot 7$ | $+3 \cdot 12$ | $+0.96$ | 5 |  |
| 9593.2 | $+20 \cdot 2$ | $+2 \cdot 11$ | $-0.05$ | 15 | (4) = Purely accidental error to 1,000 feet |
| $6987 \cdot 9$ | + $6 \cdot 0$ | $+0.86$ | $-1.30$ | 20 |  |
| $9850 \cdot 0$ | $+10 \cdot 0$ | +1.02 | $-1.14$ | 20 | (5) $=$ Number of sides to |
| $55055 \cdot 7$ |  | $+2 \cdot 16$ | $\pm 1 \cdot 21=\frac{\sqrt{s s}}{n}$ |  | adi |

CLOSED TRAVERSES.

| LOCALITY. | (1) | (2) | (3) | (4) | (1) $=$ Sum of distances by |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  | feet. | feet. | feet. | feet. | $(2)=$ Closing error. |
| Mother Bunch Islands. | 4061 | $13 \cdot 9$ | 14 | $3 \cdot 42$ |  |
| Vicar's Island. | 2316 | $7 \cdot 1$ | 10 | $3 \cdot 06$ | (3) $=$ closed traverse. |
| Harbor Islands. | 5722 | $1 \cdot 9$ | 12 | $0 \cdot 33$ | (4) = Error to 1,000 feet, in- |
| Hatcher Island. | 1610 | $3 \cdot 5$ | 6 | $2 \cdot 17$ | cluding constant error. |

352. In 1881 a stadia-survey for a road was made in Mexico,* from Culiacan to Durango. Two different routes were followed, one in going up the mountains to Durango, and the other on the return to Culiacan. The total distance run was 606 miles, and difference of elevation 11,000 feet. When the entire traverse was closed, the error of closure was found to be 1,100 feet.
[^44]
## CHAPTER V.

## OBSTACLES IN ANGULAR SURVEYING.

353. The obstacles, such as trees, houses, hills, valleys, rivers, etc., which prevent the direct alinement or measurement of any desired course, can be overcome much more easily and precisely with any angular instrument than with the chain, methods for using which were explained in Chapter II. They will, however, be taken up in the same order. As before, the given and measured lines are drawn with fine full lines; the visual lines with broken lines ; and the lines of the result with heary full lines. Part of the demonstrations of the problems are given, and part are left as exercises for the student.

## PERPENDICULARS AND PARALLELS.

354. Erecting Perpendiculars. To erect a perpendicular to a line at a given point, set the instrument at the given point, and, if it be a compass, direct its sights on the line, and then turn them till the new bearing differs $90^{\circ}$ from the original one. A convenient approximation is to file notches in the compass-plate, at the $90^{\circ}$ points, and stretch over them a thread, sighting across which will give a perpendicular to the direction of the sights.

The transit being set as above, note the reading of the vernier, and then turn it till the new reading is $90^{\circ}$ more or less than the former one.
355. To erect a perpendicular to an inacessible line, at a given point of $i t$. Let AB be the line and A the point. Calculate the distance from $A$ to any point $C$, and the angle $C A B$, by the
method of Art. 381. Set the instrument at C , sight to A , turn an angle $=\mathrm{CAB}$, and measure in the direction thus obtained a distance $\mathrm{C} \mathrm{P}=\mathrm{CA} . \cos . \mathrm{CAB}$. PA will be the required perpendicular.

Fig. 251.

356. Letting fall Perpendiculars. To let fall a perpendicular to a line from a given point. With the compass, take the bearing of the given line, and then from the

Fig. 252.
 given point run a line, with a bearing differing $90^{\circ}$ from the original bearing, till it reaches the given line.

With the transit, set it at any point of the given line, as A, and observe the angle between this line and a line thence to the given point, P. Then set at P, sight to the former position of the instrument, and turn a number of degrees equal to what the observed angle at A wanted of $90^{\circ}$. The instrument will then point in the direction of the required perpendicular P B.

35\%. To let fall a perpendicular to a line from an inaccessible point. Let AB be the line and P the point. Measure the angles PAB and P B A. Measure A B. The angles A P C and BPC are known, being the complements of the angles measured. Then is $\mathrm{AC}=\mathrm{AB} \cdot \frac{\tan . \mathrm{APC}}{\tan . \mathrm{A} \mathrm{P} \mathrm{C}+\tan . \mathrm{BPC}}$.

Fig. 253.


Proof: $\mathrm{AC}=\mathrm{PC} . \tan . \mathrm{APC}$; and $\mathrm{C} \mathrm{B}=\mathrm{PC} . \tan . \mathrm{B} \mathrm{P} \mathrm{C} \mathrm{[Trigo-}$ nometry, Art. 4].

Hence $\quad \mathrm{AC}: \mathrm{CB}:: \tan . \mathrm{APC}: \tan . \mathrm{BPC}$; and $A C: A C+O B:: \tan . A P C: \tan . A P C+\tan . B P C$. Consequently, since $\mathrm{AC}+\mathrm{CB}=\mathrm{AB}, \mathrm{AC}=\mathrm{AB} \cdot \frac{\tan . \mathrm{APC}}{\tan . \mathrm{APC}+\tan . \mathrm{BPC}}$.
358. To let fall a perpendicular to an inaccessible line from a given point. Let C be the point and A B the line. Calculate the

Fig. 254.

angle C A B by the method of Art. 381. Set the instrument at $C$, sight to $A$, and turn an angle $=90-\mathrm{CAB}$. It will then point in the direction of the required perpendicular, C E.
359. Running Parallels. To trace a line through a given point parallel to a given line. With the compass, take the bearing of the given line, and then, from the given point, run a line with the same bearing.

With the transit or theodolite, set it at any convenient point of the given line, as $A$, direct it on this line, and note the reading. Then turn the vernier till the cross-hairs bisect the given point, P. Take the instrument to this point and sight back to the former

Fig. 255.
 station, by the lower motion, without changing the reading. Then move the vernier till the reading is the same as it was when the telescope was directed on the given line, or $180^{\circ}$ different. It will then be directed on P Q , a parallel to $A B$, since equal angles have been measured at $A$ and $P$. The manner of reading them is similar to the method of "Traversing."
360. To trace a line through a given point parallel to an inac-

Fig. 256.
 cessible line. Let C be the giren point and AB the inaccessible line. Find the angle C A B, as in Art. 381. Set the instrument at $C$, direct it to $A$, and then turn it so as to make an angle with CA equal to the supplement of the angle CAB. It will then point in a direction, CE , parallel to A B.

## OBSTACLES TO ALINEMENT.

## A. To prolong a Line.

361. The instrument being set at the farther end of a line and directed back to its beginning, the sights of the compass, if that be used, will at once give the forward direction of the line. A distant point being thus obtained, the compass is taken to it and the process repeated. The use of the transit for this purpose has been fully explained.
362. By Perpendiculars. When a tree or house obstructing the line is met with, place the instrument at a point B of the line, and set off there a perpendicular to $C$; set off another at C to D , a third at D to

Fig. 257.
 E , making $\mathrm{DE}=\mathrm{BC}$, and a fourth at E , which last will be in the direction of AB prolonged. If perpendiculars can not be conveniently used, let BC and DE make any equal angles with the line AB , so as to make CD parallel to it.
363. By an Equilateral Triangle. At B turn aside from the line at an angle of $60^{\circ}$, and measure
 some convenient distance B C. At C turn $60^{\circ}$ in the contrary direction, and measure a distance CD $=\mathrm{BC}$. Then will D be a point in the line A B prolonged. At D turn $60^{\circ}$ from C D prolonged, and the new direction will be in the line of A B prolonged. This method requires the measurement of one angle less than the preceding.
364. By Triangulation. Let AB be the line to be prolonged. Choose some station C, whence

Fig. 259.
 can be seen A, B, and a point beyond the obstacle. Measure A B
and the angles A and B of the triangle ABC , and thence calculate the side AC. Set the instrument at C , and measure the angle ACD, CD being any line which will clear the obstacle. Let E be the desired point in the lines AB and CD prolonged. Then in the triangle ACE will be known the side AC and its including angles, whence CE can be calculated. Measure the resulting distance on the ground, and its extremity will be the desired point E. Set the instrument at $\mathbf{E}$, sight to $\mathbf{C}$, and turn an angle equal to the supplement of the angle AEC, and you will have the direction, EF , of AB prolonged.
365. When the Line to be prolonged is inaccessible. In this case, before the preceding method can be applied, it will be necessary to determine the lengths of the lines AB and A C, and the angle A, by the method given in Art. 381.
366. To prolong a Line with only an Angular Instrument. This may be done when no means

Fig. 260.
 of measuring any distance can be obtained. Let A B be the line to be prolonged. Set the instrument at $B$ and deflect angles of $45^{\circ}$ in the directions C and D. Set at some point, C, on one of these lines and deflect from C B $45^{\circ}$, and mark the point D where this direction intersects the direction BD. Also, at C, deflect $90^{\circ}$ from B. Then, at D, deflect $90^{\circ}$ from D B. The intersections of these last directions will fix a point E . At E deflect $135^{\circ}$ from EC or ED, and a line EF, in the direction of $A B$, will be obtained and may be continued.*

## B. To interpolate Points in a Line.

367. The instrument being set at one end of a line and directed to the other, intermediate points can be found, etc. If a ralley in-

[^45]tervenes, the sights of the compass (if the compass-plate be very carefully kept level crosswise), or the telescope of the transit, answer as substitutes for the plumb-line.
368. By a Random Line. When a wood, hill, or other obstacle prevents one end of the line, Z , from being seen from the other, A , run a random line A B with the compass or transit, etc., as nearly in the desired direction as can be guessed, till you arrive opposite the point Z. Meas-

Fig. 261.
 ure the error, B Z, at right angles to AB , as an offset. Multiply this error by ${ }_{5}{ }^{18} \frac{3}{10}$, and divide the product by the distance A B. The quotient will be the degrees and decimal parts of a degree contained in the angle BAZ. Add or subtract this angle to or from the bearing or reading with which AB was run, according to the side on which the error was, and start from A , with this corrected bearing or reading, to run another line, which will come out at Z, if no error has been committed.

Example: A random line was run, by compass, with a bearing of $\mathrm{S} .80^{\circ} \mathrm{E}$. At twenty chains distance a point was reached opposite to the desired point, and ten links distant from it on its right. Required the correct bearing.

Ans. By the rule, $\frac{10 \times 57^{\circ} \cdot 3}{2,000}=0^{\circ} \cdot 2865=17^{\prime}$. The correct bearing is therefore $\mathrm{S} .80^{\circ} 17^{\prime} \mathrm{E}$. If the transit had been used, its reading would have been changed for the new line by the same $17^{\prime}$. A simple diagram of the case will at once show whether the correction is to be added to the original bearing or angle, or subtracted from it.

If trigonometrical tables are at hand, the correction will be more precisely obtained from this equation : Tan. $\mathrm{BAZ}=\frac{\mathrm{BZ}}{\mathrm{AB}}$. In this example, $\frac{\mathrm{B} Z}{\mathrm{AB}}=\frac{10}{2,000}=\cdot 005=\tan .1 \%^{\prime \prime}$.

The $57^{\circ} \cdot 3$ rule, as it is sometimes called, may be variously modified. Thus, multiply the error by $86^{\circ}$, and divide by one and a half time the distance ; or, to get the correction in minutes,
multiply by 3,438 and divide by the distance ; or, if the error is given in feet and the distance in four-rod chains, multiply the former by 52 and divide by the distance, to get the correction in minutes.

The correct line may be run with the bearing of the random line by turning the vernier for the correction.
369. By Latitudes and Departures. When a single line, such as AB, can not be run so as to come opposite to the

Fig. 262.
 given point Z, proceed thus with the compass: Run any number of zigzag courses, A B, B C, CD, D Z, in any convenient direction, so as at last to arrive at the desired point. Calculate the latitude and departure of each of these courses and take their algebraic sums. The sum of the latitudes will be equal to AX , and that of the departures to $X Z$. Then is $\tan . Z A X=\frac{X Z}{\overline{X A}}$; i. e., the algebraic sum of the departures divided by the algebraic sum of the latitudes is equal to the tangent of the bearing.*
370. When the transit is used, any line may be taken as a meridian-i. e., as the line to which the following lines are referred; as in "Traversing," Art. 335, all the successive lines were referred to the first line. In Fig. 263 the same lines as in the preceding figure are represented, but they are referred to the first course, AB, instead of to the magnetic meridian as before, and their latitudes are measured along its produced line, and its departures perpendicular to it. As before, a right-angled triangle will be formed, and the angle ZAY will be the angle at A between the first line $A B$ and the desired line $A Z$.

This method of operation has many useful applications, such as in obtaining data for running railroad-curves, etc., and the student should master it thoroughly.

[^46]The desired angle (and at the same time the distance) can be obtained, approximately, in this and the preceding case, by finding in a traverse-table the final latitude and departure of the desired line (or a latitude and departure having the same ratio), and the bearing and distance corresponding to these will be the angle and distance desired.

$$
\text { Fig. } 263 .
$$


371. By Similar Triangles. Through A measure any line C D. Take a point $E$, on the line CB, beyond the obstacle, and from it set off a parallel to CD, to some point, F , in the line D B. Measure EF, CD, and CA. Then this proportion, $\mathrm{CD}: \mathrm{CA}:: \mathrm{EF}$ : $\mathrm{E} G$, will give the distance E G, from $E$ to a point in the line $A B$. So for other points.
372. By Triangulation. When obstacles prevent the preceding methods being used, if a point, C , can be found from which A and $B$ are accessible, measure the distances C A, CB, and the angle A CB, and thence calculate the angle CAB. Then observe any angle ACD beyond the obstacle. In the triangle ACD a side and its including angles are

Fig. 265.
 known to find CD. Measure it, and a point, D , in the desired line will be obtained.

## OBSTACLES TO MEASUREMENT.

## A. When Both Ends of the Line are accessible.

373. The methods given in the preceding articles for prolonging a line and for interpolating points in it will generally give the length of the line by the same operation. The method of latitudes and departures is very generally applicable. So is the following.

Fig. 266.

374. By Triangulation, Let A B be the inaccessible distance. From any point, C , from which both A and B are accessible, measure CA , $C B$, and the angle $A C B$. Then in the triangle A B C two sides and the included angle are known to find the side A B.*
375. By Angles to Known Points. The length of a line, both ends of which are accessible, may also be determined by angles measured at its extremities between it and the directions of two or more known points. But, as the methods of calculation involve subsequent problems, they will be postponed.

## B. When One End of the Line is inaccessible.

376. By Perpendiculars. Many of the methods giren for the chain may be still more advantageously employed with angular instruments, which can so much more easily and precisely set off the perpendiculars.
377. By Equal Angles. Let A B be the inaccessible line. At A set off A C perpendicular to A B, and as nearly equal to it, by estimation, as the ground will permit. At C measure the angle ACB, and turn the sights or vernier till $\mathrm{ACD}=$ ACB. Find the point, D, at the intersection of the lines CD and

Fig. 267.
 B A produced. Then is A D $=\mathrm{AB}$.
378. By Triangulation. Measure a distance A C, about equal to AB. Measure the angles at A and C. Then, in the triangle A B C, two angles and the included side are known, to find another side, A B $=\frac{\mathrm{AC} \sin . \mathrm{A} \mathrm{CB}}{\sin . \mathrm{A} \mathrm{B} \mathrm{C}}$.

[^47]When the compass is used, the angles between the lines will be deduced from their respective bearings.

If the angle at A is $90^{\circ}, \mathrm{AB}=\mathrm{AC}$. tang. ACB.

If the angle $\mathrm{ACB}=45^{\circ}$, then $\mathrm{AC}=\mathrm{AB}$; but this position could not easily be obtained, except by the use of the sextant, a reflecting

## Fig. 268.

 instrument, described in Part V.
379. When One Point can not be seen from the other. Choose two points, C and D , in the line of A ,

Fia. $268^{\prime}$.
 and such that from $\mathrm{C}, \mathrm{A}$, and B can be seen, and from $\mathrm{D}, \mathrm{A}$, and B. Measure $\mathrm{AC}, \mathrm{AD}$, and the angles C and D . Then, in the triangle BCD , are known two angles and the included side, to find CB. Then, in the triangle ABC, are known two sides and the included angle, to find the third side, A B.
380. To find the Distance from a Given Point to an Inaccessible Line. In Fig. 254, Art. 358, the required distance is CE. The operations therein directed give the line CA and the angle CAB, or CAE. The required distance $\mathrm{CE}=\mathrm{CA} \cdot \sin$. CAE.
C. When Both Ends of the Line are inacoessible.
381. General Method. Let A B be the inaccessible line. Measure any convenient distance, CD , and the angles ACD , B C D, A D C, B D C.

Then, in the triangle $\mathrm{CD} A$, two angles and the included side are given, to find CA. In the

$$
\text { Fig. } 269 .
$$


triangle $C D B$, two angles and the included side are giren, to find $C$ B. Then, in the triangle $A B C$, two sides and the included angle are giren, to find $A B$.

The mork may be rerified by taking another set of triangles, and finding $\triangle B$ from the triangle $\triangle B D$ instead of $\perp B C$.

The following formulas will, howerer, give the desired distaoces mith less labor:

Find on angle $K$, such that tan. $K=\frac{\sin \cdot A D C \cdot \sin \cdot C B D}{\sin \cdot C A D \cdot \sin \cdot B D C}$.
Then find the difference of the unknown angles in the triangle $C \perp B$ from the formula-

$$
\text { Finally, } A B=C D \frac{\sin \cdot B D C \cdot \sin \cdot A C B}{\sin C B D \cdot \sin \cdot C A B}
$$

Demonstration: In the triangle A B C. designate the angles as A. B, C ; snd the sides opposite to them $a s a, b, c$. Let $\mathrm{CD}=d$. The triangle BCD gires [Trig., Art. 12, Theorem $I, a=d \frac{\text { sin. } \mathrm{BDC}}{\sin . \mathrm{CBD}}$. The triangle $\triangle \mathrm{CD}$ simi$\operatorname{larl} \mathrm{f}$ gires $b=d$. $\frac{\sin . A D C}{\operatorname{sid} . C A D}$.

In the triangle A B C, we hare [Trig.. Art. 12, Theorem II,

$$
\tan \frac{1}{\Xi}(\perp-\mathrm{B}): \cot \frac{1}{ \pm} \mathrm{C}:: a-b: a+b ;
$$

mbence

$$
\begin{equation*}
\tan \cdot \frac{\ddagger}{z}(A-B)=\frac{a-b}{a+b} \cdot \cot \neq C \text {. } \tag{1.}
\end{equation*}
$$

Let $\bar{K}$ be an suriliary sngle, such that $b=a \cdot \tan . \overline{\mathrm{F}}$; whepce tan. $\mathrm{K}=\frac{b}{a}$ Diriding the second member of equation [1], abore and below, br $a$, sod sab-


Since tan. $45^{\circ}=1$, we msy substitate it for 1 in the preceding equation, and we get tan. $\frac{1}{3}(A-B)=\frac{\tan .45^{\circ}-\tan \cdot \mathrm{K}}{\tan .45^{2}+\tan . \mathrm{K}} \cdot \cot$. $\frac{1}{3} \mathrm{C}$.

From the expression for the tangent of the difference of two arcs [Trig., Art. 8], the preceding fraction redaces to tan. ( $45^{*}-\mathrm{K}$ ); and the equation becomes

$$
\tan \cdot \frac{1}{2}(A-B)=\tan \cdot\left(\frac{1}{2} 5^{2}-B\right) \cdot \cot \frac{1}{3} C .
$$

In the equation tsin. $K=\frac{\hbar}{a}$, substitute the ralues of $b$ sod $a$ from the formalas at the beginning of this inrestigation. This gires

$$
\text { tan. } \kappa=d \cdot \frac{\sin \cdot A D C}{\sin \cdot C A D} \div d \cdot \frac{\sin \cdot B D C}{\sin \cdot C B D}=\frac{\sin A D C \cdot \sin \cdot C B D}{\sin \cdot A D \cdot \sin \cdot B D C}
$$

$(A-B)$ is then obsained by equation [2]; $(1+B)$ is the supplement of $C$; therefore, the angle $A$ is known.

$$
\begin{aligned}
& \text { Tan. } \frac{1}{3}(C A B-A B C)=\tan \left(45^{2}-h\right) \cdot \cot \frac{1}{4} A C B . \\
& \text { Then is } C A B=\frac{1}{3}(C A B- \pm B C)+\frac{1}{2}(C A B+A C B) \text {. }
\end{aligned}
$$

Then

$$
c=\mathrm{A} \mathrm{~B}=\frac{a \sin . \mathrm{C}}{\sin . \mathrm{A}}=\frac{d \cdot \sin . \mathrm{B} \mathrm{D} \mathrm{C} \cdot \sin . \mathrm{A} \mathrm{C} \mathrm{~B}}{\sin . \mathrm{CBD} \cdot \sin . \mathrm{CAB}} .
$$

The use of the auxiliary angle K avoids the calculation of the sides $a$ and $b$.
Example. Let C D $=7,106.25$ feet ; A CD $=95^{\circ} 17^{\prime} 20^{\prime \prime} ; ~ \mathrm{BCD}=61^{\circ}$ $41^{\prime} 50^{\prime \prime} ; \mathrm{ADC}=39^{\circ} 38^{\prime} 40^{\prime \prime} ; \mathrm{BDC}=78^{\circ} 35^{\prime} 10^{\prime \prime}$; required AB .

The figure is constructed with these data on a scale of 5,000 feet to 1 inch $=1: 60000$.

By the above formulas, K is found to be $30^{\circ} 26^{\prime} 5^{\prime \prime} ; \mathrm{C} \mathrm{A} \mathrm{B}=113^{\circ} 55^{\prime}$ $37^{\prime \prime}$; and, lastly, $\mathrm{A} \mathrm{B}=6598 \cdot 32$.

Both the methods may be used as mutual checks in any important case.
If the lines A B and C D crossed each other, as in Fig. 270, instead of being situated as in the preceding figure, the same method of calculation would apply.
382. Problem. To measure an inaccessible distance, A B, when a point, C, in its line can be obtained. Set the instrument at a point, D , from which $\mathrm{A}, \mathrm{B}$, and

Fig. 270.
 $\sigma$ can be seen, and measure the angles CD A and ADB. Measure also the line D C and the angle C. Then in the triangle ACD two angles and

Fig. 271.
 the included side are given to find A D. In the triangle D A B, the angle D A B is known (being equal to A CD + C DA), and A D having been found, we again have two angles and the included side to find AB.
383. Problem, To measure an inaccessible distance, A B, when only one point, C, can be found from which both ends of the line can be seen. Consider C A and C B as distances to be determined, having one end accessible. Determine them as in Art. 378 , by choosing a point D, from which $C$ and $A$ are visible, and a point $E$, from which $C$ and $B$ are visible. At $C$ observe the angles D C A, A C B, and BCE. Measure the distances CD and CE . Observe the angles ADC and BEC. Then in the triangle A D C, two angles and the included side are given, to find CA ; and the same in the triangle CBE , to find CB. Lastly, in the triangle A C B two sides and the included angle are known, to find A B.
384. Problem. To measure an inacessible distance, A B, when no point can be found from which the two ends can be seen. Let C be a point from which A is risible, and D a point from which B is visible, and also C . Measure CD. Find the distances CA

Fig. 273.
 and DB , as in the preceding problem, i. e., choose a point $E$, from which $A$ and C are visible, and another point, F , from which $D$ and $B$ are risible. Measure E C and D F. Observe the angles $\mathrm{AEC}, \mathrm{ECA}, \mathrm{BDF}$, and DFB; and at the same time the angles ACD and $C \cdot D B$, for the subsequent work. Then $C A$ and $D B$ will be found, as were CA and $C B$ in the last problem. Then in the triangle $C D B, t w o$ sides and the included angle are known to find CB and the angle D CB; and, lastly, in the triangle ACB, two sides and the included angle (the difference of $A C D$ and $D C B$ ) to find $A B$.
385. Problem. Given the angles observed, at the ends of a line which can not be measured, between it and the ends of a line of knoun length but inaccessible, required the length of the former line. This problem is the converse of that given in Art. 381. Its figure, 269, may represent the case, if the distance AB be regarded as known and CD as that to be found. Use the first and second formulas as before, and insert the last formula, obtaining $C D=A B \frac{\sin . C B D \cdot \sin . C A B}{\sin . B D C \cdot \sin \cdot A C B}$.

This problem may also be solred, indirectly, by assuming any length for CD, and thence calculating, as in the first part of Art. 381, the length of AB on this hypothesis. The imaginary figure thus calculated is similar to the true one; and the true length of CD will be given by this proportion: Calculated length of AB : true length of $\mathrm{A} B:$ : assumed length of $C D$ : true length of $C D$.

The length of CD can also be obtained graphically. Take a line of any length, as $\mathrm{C}^{\prime} \mathrm{D}^{\prime}$, and from $\mathrm{C}^{\prime}$ and $\mathrm{D}^{\prime}$ lay off angles equal to those observed at C and D , and thus fix points $\mathrm{A}, \mathrm{B}^{\prime}$. Produce $A B^{\prime}$ till it equals the given distance $A B$,

Fig. 274.
 on any desired scale. From $B$ draw a parallel to $\mathrm{B}^{\prime} \mathrm{D}^{\prime}$, meeting $\mathrm{A} \mathrm{D}^{\prime}$ produced in D ; and from D draw a parallel to $\mathrm{D}^{\prime} \mathrm{C}^{\prime}$ meeting $\mathrm{AC}^{\prime}$ produced in C. Then $C D$ will be the required distance to the same scale as $\mathrm{A} B$.
386. Problem. Three points, A B C, being giren by their distances from each other, and two other points, P and Q , being so situated that from each of them troo of the three points can be seen and the angles A PQ, B P Q. C Q P, B Q P , be mexsured, it is required to determine the positions of P and Q .

Constrootion. Begin by describing a circle passing through $A$ and $B$, and having the central angle subtended by $A B$, equal to twice the given angle APB, and thus containing that angle. The point $P$ will lie somewhere in its circumference. Describe another circle passing through B and C , and having a central angle subtended by BC equal to twice the given angle B Q C. The point Q will lie somewhere in its circumference. From A draw a line making with $A B$ an angle $=B P Q$, and meeting at $X$ the circle first
 drawn. From $C$ draw a line making with CB an angle $=\mathrm{BQP}$, and meeting the second circle in Y . Join $\mathrm{X} Y$ and produce it till it cuts the circles in points P and Q , which will be those required; since $B P X=B A X=B P Q$; and $B Q Y=B C Y=$ B Q P.

Calculation. In the triangle A B C, the sides being given, the angle $\mathrm{A} B \mathrm{C}$ is known. In the triangle $\mathrm{A} B X$, a side and all the angles are known, to find B X. In the triangle C B Y, B Y is similarly found. By subtracting the angle ABC from the sum of the angles ABX and CBY, the angle X B Y can be obtained. Then in the triangle X B Y, the sides B X, B Y, and the included angle are given to find the other angles. Then in the triangle BPX are known all the angles and the side BX to find BP. In the triangle $\mathrm{BQ} \mathrm{Y}, \mathrm{B}$ Q is found in like manner. Finally, in the triangle B PQ , $P Q$ can then be found.

If desired, we can also obtain AP in the triangle APB; and C Q in the triangle C B Q.
387. Problem. Four points, A, B, C, D, being given in position, by their mutual distances and directions, and two other points, P and Q , being so situated that from each of them two of the four points can be seen and the angles A P B, A P Q, PQC, and PQD measured, it is required to determine the position of P and Q .

Construction. Begin, as in the last article, by describing on AB the segment of a circle to contain an angle equal to A PB. From B draw a chord BE, making an angle with BA equal to the supplement of the angle APQ. On CD describe another segment to contain an angle equal to $C$ QD. From $C$ draw a chord CF, making an angle with $C D$ equal to the supplement of the angle D QP. Draw the line EF, and it will cut the two circles in the required points P and Q .

For, the angle $A P Q$ in the figure equals the measured angle $A P Q$, because the supplement of the former, EP A, equals the supplement of the latter, since it is measured by the same are as the angle ABE , equal to that supplement by construction. So too with the angle D QP.

Calcclation. To obtain $P Q=E F-E P-Q F$, we proceed to find those three lines thus: In the triangle $A B E$, we know the side $A B$, the angle $A B E$, and the angle $A E B=A P B$; whence to find EB. In the same way, the triangle CFD gives FC. In the triangle EBC are known

Fig. 276.


EB and BC , and the angle $\mathrm{EBC}=\mathrm{ABC}-\mathrm{ABE}$; whence EC and the angle ECB are found. In the triangle ECF are known EC, FC, and the angle $\mathrm{ECF}=\mathrm{BCD}-\mathrm{ECB}-\mathrm{FCD}$; whence we find EF, and the angles CEF and CFE.

In the triangle BEP , we have EB , the angle $\mathrm{BEP}=\mathrm{BEO}+\mathrm{CEP}$, and the angle $\mathrm{BPE}=\mathrm{BPA}+\mathrm{APE}$; to find EP and PB. In the triangle $Q C F$, we have $C F$, and the angles $C Q F$ and $C F Q$, to find $Q C$ and $Q F$. Then we know $\mathrm{PQ}=\mathrm{EF}-\mathrm{EP}-\mathrm{QF}$.

The other distances, if desired, can be easily found from the abore data, some of the calculations, not needed for PQ , being made with reference to them. In the triangle ABP, we know A B, B P, and the angle BAP, to find the angle ABP and AP. In the triangle QDC we know Q C, C D, and the angle $C Q D$, to find the angle $Q C D$ and $Q D$. In the triangle PBC, we know P B, B C, and the angle P BC=ABC-ABP. to find P C. Lastly, in the triangle Q C B, we know Q C, C B, and the angle Q C B $=\mathrm{DCB}-\mathrm{DCQ}$, to find QB .

The solution of this problem includes the tro preceding; for, let the line BC be reduced to a point so that its two ends come together and the three lines become two, and we have the problem of Art. 386 ; and let the line A B be reduced to a point, $B$, and $C D$ to a point, $C$, and we hare but one line, and the problem becomes that of Art. 35 万.

In these three problems, if the two stations lie in a right line with one of the given points, the problem is indeterminate.
389. Problem of the Eight Points. Four points, A, B, C, D, are inaccessible, but visible from four other points, E, F, G, H; it is required to find the relative distances of these eight points; the only data being the observa-
tion, from each of the points of the second system, of the angles under which are seen the points of the first system.

This problem can be solved, but the great length and complication of the investigation and resulting formulas render it more a matter of curiosity than of utility. It may be found in Puissant's "Topographie," page 55 ; Leferre's "Trigonométrie," page 90, and Lefevre's "Arpentage," No. 387.

## TO SUPPLY OMISSIONS.

389. Any two omissions in a closed survey, whether

Fig. 277.
 of the direction or of the length, or of both, of one or more of the sides bounding the area surveyed, can always be supplied by a suitable application of the principle of latitudes and departures, although this means should be resorted to only in cases of absolute necessity, since any omission renders it impossible to "test the survey." In the following articles the survey will be considered

Fig. 278.
 to have been made with the compass. All the rules will, however, apply to a transit survey, the angles being referred to any line as a meridian, as in " traversing."

To save unnecessary labor, the examples in the various cases now to be examined will all be taken from the same survey, a plat of which is given in the margin on the scale of 40 chains to 1 inch $(1: 31,680)$, and the field-notes of which, with the latitudes and departures carried out to five decimal places, are given on page 258.*

[^48]|  | bearivgs. |  | latitudes. |  | departires. |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | N. | S. | E. | W. |
| ABCDEE | North. | 1284 | $1284 \cdot 00000$ |  | 0 | 0 |
|  | N. $32^{\circ} \mathrm{E}$. | 1782 | $1511 \cdot 22171$ |  | $944 \cdot 31619$ |  |
|  | N. $80^{\circ} \mathrm{E}$. | 2400 | 416.75568 |  | $2363 \cdot 53872$ |  |
|  | S. $48^{\circ} \mathrm{E}$. | 2700 |  | $1806 \cdot 65262$ | 2006•49096 |  |
|  | S. $18^{\circ} \mathrm{W}$. | 2860 |  | $2720 \cdot 02159$ |  | $883 \cdot 78862$ |
|  | N. $73^{\circ} 28^{\prime} 21^{\prime \prime} \mathrm{W}$. | $4621 \frac{1}{3}$ | 1314.69682 |  |  | $4430 \cdot 55725$ |
|  |  |  | 4526.67421 | $4526 \cdot 67421$ | 5314:34587 | $5314 \cdot 34587$ |

Case 1. When the length and the bearing of any one side are wanting.
390. Find the latitudes and the departures of the remaining sides. The difference of the north and south latitudes of these lines is the latitude of the omitted line, and the difference of their departures is its departure. This latitude and departure are two sides of a right-angled triangle of which the omitted line is the hypotenuse. Its length is therefore equal to the square root of the sum of their squares, and the quotient of the departure divided by the latitude is the tangent of its bearing.

In the above surrey, suppose the course from F to A to hare been omitted or lost. The difference of the latitudes of the remaining courses will be found to be $1314 \cdot 69682$, and the difference of the departures to be $4430 \cdot 055 \% 25$. The square root of the sum of their squares is 46215 ; and the quotient of the departure divided by the latitude is the tangent of $73^{\circ} 28^{\prime} 21^{\prime \prime}$. The deficiencies were in north latitude and west departure, and the omitted course is therefore N. $73^{\circ} 28^{\prime} 21^{\prime \prime} \mathrm{W} ., 4621^{\circ}$.

Case 2. When the length of one side and the bearing of another are wanting.
391. When the Deficient Sides adjoin Each Other. Find, as in Case 1 , the length and bearing of the line joining the ends of the remaining courses. This line and the deficient lines will form a triangle, in which two sides will be known, and the angle between the calculated side and the side whose bearing is given can be
found. The parts wanting can then be obtained by the common rules of trigonometry.

In the figure, let the length of EF and the bearing of FA be the omitted parts. The difference of the sums of the N. and S. latitudes, and the E. and W. departures of the complete courses from A to E, are respectively $1405 \cdot 324 \% \%$ north latitude, and $5314 \cdot 3458 \%$ east departure. The course, E A, corresponding to this deficiency, we find, by proceeding as in Case 1, to be S. $75^{\circ} 11^{\prime} 15^{\prime \prime}$ W., $549 \%$ 026. The angle AEF is therefore $=75^{\circ} 11^{\prime} 15^{\prime \prime}-18^{\circ}=57^{\circ}$

Fig. 279.
 $11^{\prime} 15^{\prime \prime}$. Then in the triangle AEF are given the sides $\mathrm{AE}, \mathrm{AF}$, and the angle AEF to find the remaining parts, viz., the angle $\mathrm{AFE}=91^{\circ} 28^{\prime} 21^{\prime \prime}$, whence the bearing of $\mathrm{FA}=91^{\circ} 28^{\prime} 21^{\prime \prime}-18^{\circ}=\mathrm{N} .73^{\circ} 28^{\prime} 21^{\prime \prime} \mathrm{W}$.; and the side $\mathrm{EF}=28 \cdot 60$.

## 392. When the Deficient Sides are separated from Each Other.

 A modification of the preceding method will still apply. In this figure let the omissions be the bearing of FA and the length of C D. Imagine the courses to change places without changing bearings or lengths, so as to bring the deficient lines next to each other by transferring $C D$ to $A G, A B$ to $G H$, and $B C$ to HD. This will not affect their latitudes or departures. Join G F. Then in the figure DEFGH the latitudes and departures of all the sides but F G are known, whence its length and bearing can be found as in Case 1. Then the triangle A GF may be treated like the triangle A E F in the last article, to obtain the length of $\mathrm{A} G=\mathrm{CD}$, and the bearing of FA .

Otherwise, by changing the meridian. Imagine the field to turn around till the side of which the distance is unknown be-
comes the meridian-i. e., comes to be due north and south-all the other sides retaining their relative positions, and continuing to make the same angles with each other. Change their bearings accordingly. Find the latitudes and departures of the sides in their new positions. Since the side whose length was unknown has been made the meridian, it has no departure, whatever may be its unknown length ; and the difference of the columns of departure will therefore be the departure of the side whose bearing is unknown. The length of this side is given. It is the hypotenuse of a rightangled triangle, of which the departure is one side. Hence the other side, which is the latitude, can be at once found, and also the unknown bearing.

Put this latitude in the table in the blank where it belongs. Then add up the columns of latitude, and the difference of their sums will be the unknown length of the side which had been made a meridian.*

Let the omitted quantities be, as in the last article, the length of $C D$ and the bearing of FA. Make CD the meridian. The

| stations. | old bearings. | new bearings. |
| :---: | :---: | :---: |
| A | North. | N. $80^{\circ} \mathrm{W}$. |
| B | N. $32^{\circ} \mathrm{E}$. | N. $48^{\circ} \mathrm{W}$. |
| O | N. $80^{\circ} \mathrm{E}$. | North. |
| I) | S. $48^{\circ} \mathrm{E}$. | N. $52^{\circ} \mathrm{E}$. |
| E | S. $18^{\circ} \mathrm{W}$. | S. $62^{\circ} \mathrm{E}$. |
| F |  |  | changed bearings can then be found to be as in the margin. To aid the imagination, turn the book around till CD points up and down, as north lines are usually placed on a map. Then obtain the latitudes of the courses with their new bearings and old distances, and proceed as has been directed.

Case 3. When the lengths of two sides are wanting.
393. When the Deficient Sides adjoin Each Other. Find the latitudes and departures of the other courses, and then, by Case 1, find the length and bearing of the line joining the extremities of the deficient courses. Then, in the triangle thus formed, are

[^49]known one side and all the angles (deduced from the bearings) to find the lengths of the other two sides.

Thus, in Fig. 279, let E F and F A be the sides whose lengths are unknown. E A is then to be calculated, and its length will be found to be $5497 \cdot 026$, and its bearing $\mathrm{S} .75^{\circ} 11^{\prime} 15^{\prime \prime} \mathrm{W}$., whence the angle AEF $=75^{\circ} 11^{\prime} 15^{\prime \prime}-18^{\circ}=57^{\circ} 11^{\prime} 15^{\prime \prime} ;$ A FE $=18^{\circ}$ $+73^{\circ} 28^{\prime} 21^{\prime \prime}=91^{\circ} 28^{\prime} 21^{\prime \prime}$; and EAF $=31^{\circ} 20^{\prime} 24^{\prime \prime}$; whence can be obtained $\mathrm{EF}=28.60$ and $\mathrm{FA}=46 \cdot 215$.
394. When the Deficient Sides are separated from Each Other. Let the lengths of BC and DE be those omitted. Again imagine the courses to change places, so as to bring the deficient lines together, D E being transferred to CG , and CD to G E. Join B G. Then in the figure ABGEFA are known the latitudes and departures of all the courses except B G, whence its length and bearing can be found, as in Case 1. Then in the triangle BCG, the angle C B G

Fig. 281.
 can be found from the bearings of C B and B G, and the angle C G B from the bearings of B G and G C. Then all the angles of the triangle are known and one side, BG , whence to find the required sides, $\mathrm{BC}=1,782$, and $\mathrm{C} \mathrm{G}=$ $\mathrm{DE}=2, \% 00$.

Otherwise, by changing the meridian. Imagine the field to turn around till one of the sides whose length is wanting becomes a meridian or due north and south. Change all the bearings correspondingly. Find the latitudes and departures of the changed courses. The difference of the columns of departure will be the departure of the second course of unknown length, since the course made meridian has now no departure. The new bearing of this second course being given in the right-angled triangle formed by this course as an hypotenuse, and its departure and latitude, we know one side, the departure, and the acute angles, which are the bearing and its complement. The length of the course is then readily calculated, and also its latitude. This latitude being in-
serted in its proper place, the difference of the columns of latitude will be the length of that wanting side which had been made a meridian.

Thus, let the lengths of B C and D E be wanting, as in the preceding example.

| statioss. | old bearings. | new bearings. |
| :---: | :---: | :---: |
| ${ }_{\text {A }}$ | North, | N. $32^{\circ} \mathrm{W}$. |
| C | $\stackrel{\text { N. } .32^{\circ} \mathrm{E}}{\mathrm{E} .} 80^{\circ} \mathrm{E}$ | $\begin{aligned} & \text { North. } \\ & \text { N. } 48^{\circ} \mathrm{E} \end{aligned}$ |
| D | S. $48^{\circ} \mathrm{E}$. | S. $80^{\circ} \mathrm{E}$. |
| E | S. $18^{\circ} \mathrm{W}$. | S. $14^{\circ} \mathrm{E}$. |
| F | N. $73^{\circ} 28^{\prime} 21^{\prime \prime} \mathrm{W}$. | S. $74^{\circ} 31^{\prime} 59^{\prime \prime}$. W. | Make B C a meridian. The other bearings are then changed as in the margin. Calculate new latitudes and departures. The difference of the departures will be the departure of DE , since B C, being a meridian, has no departure. Hence the length and latitude of DE are readily obtained. This latitude being put in the table, and the columns of latitude then added up, their difference will be the length of B C.

## Case 4. When the bearings of two sides are wanting.

395. When the Deficient Sides adjoin Each Other. Find the latitudes and departures of the other sides, and then, as in Case 1, find the length and bearing of the line joining the extremities of the deficient sides. Then, in the triangle thus formed, we have the three sides to find the angles and thence the bearings.
396. When the Deficient Sides are separated from Each Other. Change the places of the sides so as to bring the deficient ones next to each other. Thus, in the figure, supposing the bearings of CD and EF to be wanting, transfer EF to DG, and DE to GF. Then calculate, as in Case 1, the length and bearing of the line joining the extremities of the deficient sides, C G in the figure. This
 line and the deficient sides form a triangle in which the three sides are given to determine the angles and thence the required bearings.

## CHAPTER VI.

LAYING OUT, PARTING OFF, AND DIVIDING UP LAND.

## IAYING OUT IAND.

397. Its Nature. This operation is precisely the reverse of those of surveying properly so called. The latter measures certain lines as they are ; the former marks them out in the ground where they are required to be, in order to satisfy certain conditions. The same instruments, however, are used as in surveying.

Perpendiculars and parallels are the lines most often employed. Part of the demonstrations of the problems are left as exercises for the student.
398. To lay out Squares. Reduce the desired content to square chains, and extract its square root. This will be the length of the required side, which is to be set out by one of the methods indicated in the preceding article.

An acre, laid out in the form of a square, is frequently desired by farmers. Its side must be made $316 \frac{1}{4}$ links of a Gunter's chain ; or $208 \frac{71}{100}$ feet ; or $69 \frac{57}{100}$ yards. It is often taken at $\%$ paces.

The number of plants, hills of corn, loads of manure, etc., which an acre will contain at any uniform distance apart, can be at once found by dividing 209 by this distance in feet, and multiplying the quotient by itself, or by dividing 43,560 by the square of the distance in feet. Thus, at 3 feet apart, an acre would contain 4,840 plants, etc. ; at 10 feet apart, 436 ; at a rod apart, 160 ; and so on. If the distances apart be unequal, divide 43,560 by the product of these distances in feet ; thus, if the plants were in rows 6 feet apart, and the plants in the rows were 3 feet apart, 2,420 of them would grow on one acre.
399. To lay out Rectangles. The content and length being given, both as measured by the same unit, divide the former by the latter, and the quotient will be the required breadth. Thus, 1 acre or 10 square chains, if 5 chains long, must be 2 chains wide.

The content being given and the length to be a certain number of times the breadth. Divide the content in square chains, etc., by the ratio of the length to the breadth, and the square root of the quotient will be the shorter side desired, whence the longer side is also known. Thus, let it be required to lay out 30 acres in the form of a rectangle 3 times as long as broad; 30 acres $=300$ square chains. The desired rectangle will contain 3 squares, each of 100 square chains, having sides of 10 chains. The rectangle will therefore be 10 chains wide and 30 long.

An acre laid out in a rectangle twice as long as broad will be 224 links by 448 links, nearly ; or, $14 \frac{1}{2}$ feet by 295 feet ; or, $49 \frac{1}{3}$ yards by $98 \frac{2}{5}$ yards. Fifty paces by one hundred is often used as an approximation, easy to be remembered.

The content being given, and the difference between the length and breadth. Let $c$ represent this content, and $d$ this difference. Then the longer side $=\frac{1}{2} d+\frac{1}{2} \sqrt{ }\left(d^{2}+4 c\right)$.

Example. Let the content be 6.4 acres, and the difference 12 chains. Then the sides of the rectangle will be respectively 16 chains and 4 chains.

The content being given, and the sum of the length and breadth. Let $c$ represent this content, and $s$ this sum. Then the longer side $=\frac{1}{2} s+\frac{1}{2} \sqrt{ }\left(s^{2}-4 c\right)$.

Example. Let the content be 6.4 acres, and the sum 20 chains. The above formula gives the sides of the rectangle 16 chains and 4 chains as before.
400. To lay out Triangles. The content and the base being given, divide the former by half the latter to get the height. At any point of the base erect a perpendicular of the length thus obtained, and it will be the vertex of the required triangle.

The content being given and the base having to be $m$ times the height, the height will equal the square root of the quotient obtained by dividing twice the given area by $m$.

The content being given and the triangle to be equilateral, take the square root of the content and multiply it by 1520 . The product will be the length of the side required. This rule makes the sides of an equilateral triangle containing one acre to be $480 \frac{1}{2}$ links. A quarter of an acre laid out in the same form would have each side 240 links long. An equilateral triangle is very easily set out on the ground, as directed under "Platting," using a rope or chain for compasses.

The content and base being given, and one side having to make a given angle, as $B$, with the base AB, the length of the side $\mathrm{BC}=\frac{2 \times \mathrm{ABC}}{\mathrm{AB} \cdot \sin \cdot \mathrm{B}}$.

Example. Eighty acres are to be laid out in the form of a triangle, on a base, AB , of sixty chains, bearing $\mathrm{N} .80^{\circ} \mathrm{W}$., the bearing of the side BC being $\mathrm{N} .70^{\circ}$

Fig. 283.

E. Here the angle B is found from the bearings (reversing one of them) to be $30^{\circ}$. Hence $\mathrm{BC}=53 \cdot 33$. The figure is on a scale of 50 chains to 1 inch $=1: 39600$.

Any right-line figure may be laid out by analogous methods.
401. To lay out Circles. Multiply the given content by ${ }^{7}$, divide the product by 22 , and take the square root of the quotient. This will give the radius, with which the circle can be described on the ground with a rope or chain. A circle containing one acre has a radius of $1788_{4}$ links. A circle containing a quarter of an acre will have a radius of 89 links.
402. Town-Lots. House-lots in cities are usually laid off as rectangles of 25 feet front and 100 feet depth, variously combined in blocks. Part of New York is laid out in blocks 200 feet by 800 , each containing 64 lots, and separated by streets, 60 feet wide, running along their long sides, and avenues, 100 feet wide, on their short sides. The eight lots on each short side of the block front on the avenues, and the remaining forty-eight lots front on the streets. Such a block covers almost precisely $3 \frac{2}{3}$ acres, and $17 \frac{1}{2}$ such lots about make an acre. But, allowing for the streets, land
laid out into lots, 25 by 100 , arranged as abore, would contain only $11 \cdot 9$, or not quite 12 lots per acre.

Lots in small towns and villages are laid out of greater size and less uniformity: 50 feet by 100 is a frequent size for new villages, the blocks being 200 feet by 500 , each therefore containing 20 lots.
403. Land sold for Taxes. A case occurring in the State of New York will serve as an application of the modes of laying out squares and rectangles. Land on which taxes are unpaid is sold at auction to the lowest bidder-i. e., to him who will accept the smallest portion of it in return for paying the taxes on the whole. The lot in question was originally the east half of the square lot ABCD, containing 500 acres. At a sale for taxes in 1830, $\% 0$ acres were bid off, and this area was set off to the purchaser in a square lot, from the northeast corner. Required the side of the

Fig. 284.
 square in links. Again, in 1834, 29 acres more were thus sold, to be set off in a strip of equal width around the square pretiously sold. Required the width of this strip. Once more : in 1839, 42 acres more were sold, to be set off around the preceding piece. Required the dimensions of this third portion. The answer can be proved by calculating if the dimensions of the remaining rectangle will gire the content which it should hare, riz., $250-$ $\left({ }^{4} 0+29+42\right)=109$ acres.

The figure is on a scale of 40 chains to 1 inch $=1: 31680$.
404. New Countries. The operations of laring out land for the purposes of settlers are required on a large scale in new countries, in combination with their surver. There is great difficulty in uniting the necessary precision, rapidity, and cheapness. "Triangular surrering" will insure the first of these qualities, but is deficient in the last tro, and leares the laying out of lots to be
subsequently executed. "Compass-surveying" possesses the last two qualities, but not the first. The United States system for surveying and laying out the public lands admirably combines an accurate determination of standard lines (meridians and parallels) with a cheap and rapid subdivision by compass. The subject is so important and extensive that it will be explained by itself.

## PARTING OFF LAND.

405. It is often required to part off from a field, or from an indefinite space, a certain number of acres by a fence or other boundary-line, which is also required to run in a particular direction, to start from a certain point, or to fulfill some other condition. The various cases most likely to occur will be here arranged according to these conditions. Both graphical and numerical methods will generally be given.*

The given content is always supposed to be reduced to square chains and decimal parts, and the lines to be in chains and decimals.

## A. By a Line parallel to a Side.

406. To part off a Rectangle. If the sides of the field adjacent to the given side make right angles with it, the figure parted off by a parallel to the given side will be a rectangle, and its breadth will equal the required content divided by that side, as in Art. 398.

If the field be bounded by a curved or zigzag line outside of the given side, find the content between these irregular lines and the given straight side, by the method of offsets, subtract it from the content required to be parted off, and proceed with the remainder as above. The same directions apply to the subsequent problems.
407. To part off a Parallelogram. If the sides adjacent to the given side be parallel, the figure parted off will be a parallelogram, and its perpendicular width, C E,

Fig. 285.


[^50]will be obtained as above. The length of one of the parallel sides, as $\mathrm{AC}=\frac{\mathrm{CE}}{\sin \cdot \mathrm{A}}=\frac{\mathrm{ABDC}}{\mathrm{AB} \cdot \sin \cdot \mathrm{A}}$.
408. To part off a Trapezoid. When the sides of a field adjacent to the given side are not parallel, the figure parted off will be a trapezoid.

When the field or figure is given on the ground, or on a plat, begin as if the sides were parallel,

Fig. 286.
 dividing the given content by the base AB. The quotient will be an approximate breadth, CE , or DF; too small if the sides converge, as in the figure, and vice versa. Measure CD. Calculate the content of ABDC . Divide the difference of it and the required content by C D. Set off the quotient perpendicular to CD (in this figure, outside of it), and it will give a new line, G H, a still nearer approximation to that desired. The operation may be repeated, if found necessary.
409. When the field is given by bearings, deduce from them the angles at A and B . The required sides will then be given by these formulas:

$$
\begin{gathered}
\mathrm{CD}=\sqrt{ }\left(\mathrm{A} \mathrm{~B}^{2}-\frac{2 \times \mathrm{ABCD} \cdot \sin \cdot(\mathrm{~A}+\mathrm{B})}{\sin \cdot \mathrm{A} \cdot \sin \cdot \mathrm{~B}}\right) . \\
\mathrm{AD}=(\mathrm{AB}-\mathrm{CD}) \frac{\sin \cdot \mathrm{B}}{\sin \cdot(\mathrm{~A}+\mathrm{B})} . \\
\mathrm{BC}=(\mathrm{AB}-\mathrm{CD}) \frac{\sin \cdot \mathrm{A}}{\sin \cdot(\mathrm{~A}+\mathrm{B})} .
\end{gathered}
$$

Demonstration. Produce B C and A D to meet in E. By similar triangles,

ABE:DCE::A B $: \mathrm{DC}^{2}$.

Fig. 287.


$$
\begin{gathered}
\text { ABE-DCE:ABE::A B2}-\mathrm{DC}^{2}: \mathrm{AB}^{2} \\
\text { Now } \mathrm{ABE}-\mathrm{DCE}=\mathrm{ABCD} ; \text { also, by Art. } 61 \text {, note, } \\
\text { ABE }=\mathrm{AB}^{2} \cdot \frac{\sin \cdot \mathrm{~A} \cdot \sin \cdot \mathrm{~B}}{2 \cdot \sin \cdot(\mathrm{~A}+\mathrm{B})}
\end{gathered}
$$

The above proportion, therefore, becomes

$$
A B C D: A B^{2} \cdot \frac{\sin \cdot A \cdot \sin \cdot B}{2 \cdot \sin \cdot(A+B)}:: A B^{2}-C D^{2}: A B^{2} .
$$

Multiplying extremes and means, cancelFig. 288. ing, transposing, and extracting the square root, we get $\mathrm{C} D=\sqrt{ }\left[\mathrm{A} \mathrm{B}^{2}-\right.$ $\left.\frac{2 \cdot \mathrm{~A} \mathrm{~B} \mathrm{CD} \cdot \sin \cdot(\mathrm{A}+\mathrm{B})}{\sin \cdot \mathrm{A} \cdot \sin \cdot \mathrm{B}}\right]$.

When $\mathrm{A}+\mathrm{B}>180^{\circ}$, $\sin .(\mathrm{A}+\mathrm{B})$ is negative, and therefore the fraction in
 which it occurs becomes positive.

C F being drawn parallel to D A, we have

$$
\begin{aligned}
A D & =F C=F B \cdot \frac{\sin \cdot B}{\sin \cdot B C F}=F B \cdot \frac{\sin \cdot B}{\sin \cdot\left(180^{\circ}-A-B\right)} \\
& =(A B-C D) \frac{\sin \cdot B}{\sin \cdot(A+B)} B C=(A B-C D) \frac{\sin \cdot A}{\sin \cdot(A+B)}
\end{aligned}
$$

When the sides A D and B C diverge, instead of converging, as in the figure, the negative term, in the expression for $C D$, becomes positive ; and, in the expressions for both $A D$ and $B C$, the first factor becomes (CD-AB).

The perpendicular breadth of the trapezoid $=A D . \sin . A$; or $=$ B C. sin. B.

Example. Let A B run north, six chains ; A D, N. $80^{\circ}$ E.; $\mathrm{BC}, \mathrm{S} .60^{\circ} \mathrm{E}$. Let it be required to part off one acre by a fence parallel to AB . Here $\mathrm{AB}=6 \cdot 00, \mathrm{ABCD}=10$ square chains, $\mathrm{A}=80^{\circ}, \mathrm{B}=60^{\circ}$. Ans. $\mathrm{CD}=4 \cdot 5 \%, \mathrm{~A} \mathrm{D}=1 \cdot 92, \mathrm{~B} \mathrm{C}=2 \cdot 18$, and the breadth $=1.89$.

The figure is on a scale of 4 chains to an inch $=1: 3168$.

## B. By a Line perpendicular to a Side.

410. To part off a Triangle. Let F G be the required line.

Fig. 289.
 When the field is given on the ground, or on a plat, at any point, as $D$, of the given side $A B$, set out a "guess-line," D E, perpendicular to $A B$, and calculate the content of D E B. Then the required distance BF, from the angular point to the foot of the desired perpendicular $=\mathrm{BD} \sqrt{ }\left(\frac{\mathrm{BFG}}{\mathrm{BDE}}\right)$.

Since similar triangles are as the squares of their homologous sides, $\mathrm{BDE:BFG}:: \mathrm{BD}^{2}: B F^{2}$; whence $\mathrm{BF}=\mathrm{BD} \sqrt{ }\left(\frac{\mathrm{BFG}}{\mathrm{BDE}}\right)$.

Fig. 290.


Example. Let B D $=30$ chains ; E D $=$ 12 chains ; and the desired area $=24.8$ acres. Then B F $=35 \cdot 22$ chains.

The scale of the figure is 30 chains to 1 inch $=1: 23 \% 60$.

When the field is given by bearings, find the angle B from the bearings; then is
$\mathrm{BF}=\sqrt{ }\left(\frac{2 \times \mathrm{BFG}}{\tan . \mathrm{B}}\right)$.
Example. Let B A bear S. $75^{\circ}$ E., and B C N. $60^{\circ}$ E., and let five acres be required to be parted of from the field by a perpendicular to $B A$. Here the angle $B=45^{\circ}$, and $B F=10.00$ chains.

The scale of the figure is 20 chains to 1 inch $=1: 15840$.
411. To part off a Quadrilateral. Produce the converging sides to meet at B. Calculate the content of the triangle H K B, whether on the ground or plat, or from bearings. Add it to the content of the quadrilateral required

Fig. 291.
 to be parted off, and it will give that of the triangle FGB, and the method of the preceding case can then be applied.
412. To part off any Figure. If the field be very irregularly shaped, find by trial any line which will part off a little less than the required area. This trial-line will represent HK in the preceding figure, and the problem is reduced to parting off, according to the required condition, a quadrilateral, comprised between the trial-line, two sides of the field, and the required line, and containing the difference between the required content and that parted off by the trial-line.

## C. By a Line running in any Given Direction.

413. To part off a Triangle. By construction, on the ground or the plat, proceed nearly as in Art. 410, setting out a line in the required direction, calculating the triangle thus formed, and obtaining BF by the same formula as in that article.
414. If the field be given by bearings, find from them the angles CBA and GFB ; then is $\mathrm{BF}=\sqrt{ }\left(\frac{2 \times \mathrm{BFG} \sin .(\mathrm{B}+\mathrm{F})}{\sin . \mathrm{B} \cdot \sin . \mathrm{F}}\right)$.

Example. Let B A bear S. $30^{\circ}$ E. ; B C, N. $80^{\circ} \mathrm{E}$. ; and a fence be required to run from some point in BA, a due north course, and to part off one acre. Required the distance from B to the point F , whence it must start. Ans. The angle $B=80^{\circ}$, and $F=30^{\circ}$. Then B F $=6.4 \%$.

The scale of Fig. 292 is 6 chains to 1 inch $=1: 4752$.
415. To part off a Quadrilateral. Let it be required to part off, by a line running in a given direction, a quadrilateral from a field in which are given the side $A B$, and

Fig. 293.
 the directions of the two other sides running from A and from B.

On the ground or plat produce the two converging sides to meet at some point E. Calculate the content of the triangle A B E. Measure the side A E. From ABE subtract the area to be cut off, and the remainder will be the content of
the triangle CDE. From A set out a line A F parallel to the given direction. Find the content of ABF. Take it from

ABE , and thus obtain AFE. Then this formula, $\mathrm{ED}=\mathrm{AE}$ $\sqrt{\frac{C D E}{F A E}}$, will fix the point $D$, since $A D=A E-E D$.

When the field and the dividing line are given by bearings, produce the sides as in the last article. Find all the angles from the bearings. Calculate the content of the triangle ABE, by the formula for one side and its including angles. Take the desired content from this to obtain CDE. Calculate the side $\mathrm{A} E=\mathrm{AB}$ $\frac{\sin . \mathrm{B}}{\sin . \mathrm{E}}$. Then is $\mathrm{AD}=\mathrm{AE}-\sqrt{ }\left(\frac{2 \times \mathrm{CDE} \cdot \sin . \mathrm{DCE}}{\sin . \mathrm{E} \cdot \sin . \mathrm{CDE}}\right)$.

Demonstration. Since triangles which have an angle in each equal, are as the products of the sides about the equal angles, we have
$A B E: C D E:: A E \times B E: C E \times D E$.

$$
\begin{aligned}
\mathrm{ABE} & =\frac{1}{2} \cdot \mathrm{AB} \cdot \frac{\sin \cdot \mathrm{~A} \cdot \sin \cdot \mathrm{~B}}{\sin \cdot(\mathrm{~A}+\mathrm{B})} \cdot \\
\mathrm{BE} & =\mathrm{AB} \cdot \frac{\mathrm{sin} \cdot \mathrm{~A}}{\sin \cdot \mathrm{E}} .
\end{aligned}
$$

Substituting these values in the preceding proportion, canceling the common factors, observing that sin. $(\mathrm{A}+\mathrm{B})=$ sin. E , multiplying extremes and means, and dividing, we get $\mathrm{DE}=\sqrt{ }\left(\frac{2 \cdot \mathrm{CDE} \cdot \sin \cdot \mathrm{DCE}}{\sin . \mathrm{E} \cdot \sin . \mathrm{CDE}}\right)$.

Example. Let D A bear S. $20^{1^{\circ}} \mathrm{W} . ;$ A B, N. $\check{1 \frac{1}{2}^{\circ}} \mathrm{W} ., 8 \cdot 19$; $\mathrm{BC}, \mathrm{N} .73 \frac{1}{2}^{\circ} \mathrm{E}$; and let it be required to part off two acres by a fence, D C, running N. $45^{\circ} \mathrm{W}$. Ans. A B E $=32 \cdot 56$ square chains ; whence $\mathrm{CDE}=12.56$ square chains. Also, $\mathrm{A} \mathrm{E}=8.37$; and, finally, $\mathrm{AD}=8.37-5.51=2.86$ chains.

The scale of Fig. 293 is 5 chains to 1 inch $=1: 3960$.
If the sum of the angles at A and B were more than two right angles, the point E would lie on the other side of $\mathrm{A} B$. The necessary modifications are apparent.
416. To part off any Figure. Proceed in a similar manner to that described in Art. 412, by getting a suitable trial-line, producing the sides it intersects, and then applying the method just given.

## D. By a Line starting from a Giten Point iv a Side.

417. To part off a Triangle. Let it be required to cut of from a corner of a field a triangular space of giren content, bs a line starting from a giren point on one of the sides, A in the figure,
the base, AB , of the desired triangle being thus given. If the field be given on the ground or on a plat, divide the given content by half the base, and the quotient will be the height of the triangle. Set off this distance from any point of A B, perpendicular to it, as from A to C ; from C set out a parallel to AB , and its inter-

Fig. 294.
 section with the second side, as at D , will be the vertex of the required triangle.

Otherwise : Divide the required content by half of the perpendicular distance from A to BD , and the quotient will be B D .

If the field be given by the bearings of two sides and the length of one of them, deduce the angle B (Fig. 294) from the bearings. Then is $\mathrm{BD}=\frac{2 \times \mathrm{ABD}}{\mathrm{AB} \cdot \sin \cdot \mathrm{B}}$.

If it is more convenient to fix the point D , by the second ${ }^{\circ}$ method, that of rectangular co-ordinates, we shall have $\mathrm{BE}=$ B D. cos. B ; and E D = B D. . sin. B.

The bearing of $A D$ is obtained from the angle $B A D$, which is known, since $\frac{E D}{E A}=\frac{E D}{A B-B E}=\tan . \mathrm{BAD}$.

Example. Eighty acres are to be set off from a corner of a field, the course AB being $\mathrm{N} .80^{\circ} \mathrm{W}$., sixty chains ; and the bearing of BD being N. $70^{\circ} \mathrm{E}$. Ans. B D $=53.33$; B E $=46 \cdot 19$; $\mathrm{ED}=26.67$; and the bearing of $\mathrm{AD}, \mathrm{N} .17^{70} 23^{\prime} \mathrm{W}$.

The scale of Fig. 294 is 40 chains to 1 inch $=1: 31680$.
If the field were right-angled $\alpha \stackrel{A}{\Delta}$, of course $D B=\frac{2 A B D}{A B}$.
418. To part off a Quadrilateral. Imagine the two converging sides of the field produced to meet, as in Art. 415. Calculate the content of the triangle thus formed, and the question will then be reduced to the one explained in the last two articles.
419. To part off any Figure. Proceed as directed in Art. 416. Otherwise, proceed as follows :

The field being given on the ground or on a plat, find on which side of it the required line will

Fig. 295.
 end, by drawing or running "guess-lines" from the given point to various angles, and roughly measuring the content thus parted off. If, as in the figure, A being the given point, the guess-line A D parts off less than the required content, and A E parts off more, then the desired division-line A Z will end in the side DE. Subtract the area parted off by AD from the required content, and the difference will be the content of the triangle AD Z. Divide this by half the perpendicular let fall from the given point A to the side DE , and the quotient will be the base, or distance from D to Z .

Or, find the content of ADE and make this proportion: ADE:ADZ::DE:D Z.

The field being given by bearings and distances, find as before, by approximate trials on the plat, or otherwise, which side the desired line of division will terminate in, as DE in the last figure. Draw A D. Find the latitude and departure of this line, and thence its length and bearing. Then calculate the area of the space this line parts off, ABCD in the figure, by the usual method, explained in Part I, Chapter III. Subtract this area from that required to be cut off, and the remainder will be the area of the triangle A D Z. Then, as in Art. 41ǒ, D Z = $\frac{2 \mathrm{AD} \mathrm{Z}}{\mathrm{AD} \cdot \sin . \mathrm{ADZ}}$.

This problem may be executed without any other table than that of latitudes and departures, thus: Find the latitude and departure of DA , as before, the area of the space ABCD , and thence the content of A D Z. Then find the latitude and departure of EA, and the content of A D E. Lastly, make this proportion : ADE:ADZ: DE:D Z.*

[^51]Example. In the field ABCDE, etc., part of which is shown in Fig. 295 (on a scale of 4 chains to 1 inch $=1: 3168$ ), one acre is to be parted off on the west side, by a line starting from the angle A. Required the distance from D to Z , the other end of this dividing line.*

The only courses needed are these : A B, N. $53^{\circ} \mathrm{W} ., 1 \cdot 55$; B C, N. $20^{\circ} \mathrm{E}, 2 \cdot 00$; C D, N. $53 \frac{1}{2}^{\circ}$ E., $1 \cdot 32$; D E, S. $57^{\circ}$ E., $5 \% 79$. A rough measurement will at once show that $\mathrm{A} B C D$ is less than an acre, and that ABCDE is more ; hence the desired line will fall on DE. The latitudes and departures of $\mathrm{AB}, \mathrm{BC}$, and CD are then found. From them the course AD is found to be N. $8^{\circ} 1^{\prime} 2 z^{\prime \prime}$ E., $3 \cdot 634$. The content of ABCD will be $3 \cdot 19$ square chains. Subtracting this from one acre, the remainder, 6.81 square chains, is the content of A D Z. A P $=3.63 \times \sin .65^{\circ}=3.29$. Dividing $\mathrm{A} \mathrm{D} \mathrm{Z} \mathrm{by} \mathrm{half} \mathrm{of} \mathrm{this} \mathrm{we} \mathrm{obtain} \mathrm{D} Z=,4 \cdot 14$ chains.

By the second method, the latitude and departure of D A , the area of $A B C D$, and of $A D Z$, being found as before, we next find the latitude and departure of E A from those of A D and D E, and thence the area of $\mathrm{ADE}=9 \cdot 53$. Lastly, we have the proportion $9 \cdot 53: 6 \cdot 81:: 5 \cdot 79: \mathrm{D} \mathrm{Z}=4 \cdot 14$, as before.
E. By a Line passing through a Giten Point within the Field.

## 420. To part off a Triangle.

 Let P be a point within a field through which it is required to run a line so as to part off from the field a given area in the form of a triangle.When the field is given on
line is to fall, a meridian, and changing the bearings. The difference of the new departures will be the departure of the division-line. Its position can then be easily determined.

* If the whole field has been surveyed and balanced, the balanced latitudes and departures should be used. We will here suppose the survey to have proved perfectly correct.

Fig. 296.

the ground or on a plat, the division can be made by construction, thus: Divide the given area by half of the perpendicular distance from P to AC, and set off the quotient from $C$ to $G$. Bisect G C in H. From P draw P E, parallel to the side BC. On HE describe a semicircle. On it set off EK = EC. Join K H. Set off HL=HK. The line L M, drawn from $L$ through $P$, will be the division-line required.* If HK be set off in the contrary direction, it will fix another line L' P M', meeting C B produced, and thus parting off another triangle of the required content.

Demonstration. By construction, G PC= the required content. Now, GPC $=G D C$, since they have the same base and equal altitudes. The bave now to prove that $\mathrm{LMC}=\mathrm{GDC}$. These two triangles have a common angle at C. Hence, they are to each other as the rectangles of the adjacent sides-i. e.,

$$
\text { G D C : L M C : : GC } \times \text { C D :: LC } \times \text { CM. }
$$

Here C MI is unknown, and must be eliminated. We obtain an expression for it by means of the similar triangles L MC and LE P, which give
LE:LC:: EP = CD:CM.

Hence, $\mathrm{CM}=\frac{\mathrm{CD} \times \mathrm{LC}}{\mathrm{LE}}$. Substituting this ralue of CM in the first proportion, and canceling CD in the last two terms, we get

GDC:LMC:: GC: $\frac{L^{12}}{\mathrm{LE}}$; or GDC:LMC:: GC $\times \mathrm{LE}: \mathrm{LC}^{2}$.

$$
\mathrm{LC}^{2}=(\mathrm{LH}+\mathrm{HC})^{2}=\mathrm{LH} \mathrm{H}^{2}+2 \mathrm{LH} \times \mathrm{HC}+\mathrm{HC} \mathrm{C}^{2} .
$$

But, by construction,
$\mathrm{LH}^{2}=\mathrm{HK}^{2}=\mathrm{HE}^{2}-\mathrm{EK}^{2}=\mathrm{HE}^{2}-\mathrm{EC}^{2}=(\mathrm{HE}+\mathrm{EC})(\mathrm{HE}-\mathrm{EC})=\mathrm{HC}(\mathrm{HE}-\mathrm{EC})$. Also, $\quad G C=2 H C$; and $L E=L H+H E$.
Substituting these values in the last proportion, it becomes
GDC:LMO: : $2 . \mathrm{HC}(\mathrm{LH}+\mathrm{HE}): \mathrm{HC}(\mathrm{HE}-\mathrm{EC})+2 \mathrm{LH} \times \mathrm{HC}+\mathrm{HC}^{2}$.

$$
\begin{aligned}
:: 2 L H+2 H E & : H E-E C+2 L H+H C . \\
& : H E-E C+2 L H+H E+E C . \\
& : 2 H E+2 L H .
\end{aligned}
$$

The last two terms of this proportion are thus proved to be equal. Therefore, the first two terms are also equal-i. e., $\mathrm{LMC}=\mathrm{GDC}=$ the required content.

Since $H K=V\left(H E^{2}-E K^{2}\right)$, it will hare a negative as well as a positive ralue. It may therefore be set off in the contrary direction from L- . i. e., to $L^{\prime}$. The line drawn from $L^{\prime}$ through $P$, and meeting $C B$ produced beyond B , will part off another triangle of the required content.

Example. Let it be required to part off $31 \cdot 1 \%$ acres br a fence passing through a point $P$, the distance $P D$ of $P$ from the side

[^52]B C, measured parallel to A C, being 6 chains, and DC 18 chains. The angle at C is fixed by a "tie-line" $\mathrm{AB}=48^{\circ} 00, \mathrm{~B} C$ being $42 \cdot 00$, and CA being 30.00 . Ans. $\mathrm{CL}=2 \% \cdot 31$ chains, or $\mathrm{C} \mathrm{L}^{\prime}=7 \cdot 69$ chains.

The figure is on a scale of 20 chains to 1 inch $=1: 15840$.

If the angle of the field and the position of the point P are given by bearings or angles, proceed thus: Find the perpendicular distances, P Q and $P R$, from the given point to

Fig. 297.
 the sides, by the formulas $\mathrm{P} \mathrm{Q}=\mathrm{PC} . \sin . \mathrm{PCQ}$; and $\mathrm{PR}=$ PC. sin. PCR. Let $\mathrm{PQ}=q, \mathrm{PR}=p$, and the required content $=c$. Then $\mathrm{CL}=\frac{c}{p} \pm \sqrt{ }\left(\frac{c^{2}}{p^{2}}-\frac{2 q c}{p \sin . \mathrm{LCM}}\right)$.

Demonstration. Suppose the line L M drawn. Then, by Art. 61, note, the required content, $c=\frac{1}{2} . \mathrm{CL} \times \mathrm{CM} . \sin$. LCM. This content will also equal the sum of the two triangles LCP and $\mathrm{MCP}-\mathrm{i}$. e., $c=\frac{1}{2} . \mathrm{CL} \times p+$ $\frac{1}{2} . \mathrm{CM} \times q$. The first of these equations gives $\mathrm{CM}=\frac{2 c}{\mathrm{CL} \cdot \sin . \mathrm{L} \mathrm{CM}}$. Subetituting this in the second equation, we have

$$
c=\frac{1}{2} . \mathrm{CL} \times p+\frac{c q}{\mathrm{CL} \cdot \sin . \mathrm{LCM}} .
$$

Whence, $\quad \frac{1}{2} p . \mathrm{CL}^{2} . \sin . \mathrm{LCM}+c q=c . \mathrm{CL} . \sin . \mathrm{LCM}$.
Transposing and dividing by the coefficient of $\mathrm{CL}^{2}$, we get

$$
\begin{gathered}
\mathrm{CL}^{2}-\frac{2 c}{p} \cdot \mathrm{CL}=-\frac{2 c q}{p \cdot \sin \cdot \mathrm{CLM}} \\
\mathrm{CL}=\frac{c}{p} \pm \sqrt{\left(\frac{c^{2}}{p^{2}}-\frac{2 c q}{p \cdot \sin \cdot \mathrm{~L} \mathrm{CM}}\right)} .
\end{gathered}
$$

If the given point is outside of the lines CL and CM, conceive the desired line to be drawn from it, and another line to join the given point to the corner of the field. Then, as above, get expressions for the two triangles thus formed, and put their sum equal to the expression for the triangle which comprehends them both, and thence deduce the desired distance, nearly as above.

Example. Let the angle $\mathrm{LCM}=82^{\circ}$. Let it be required to part off the same area as in the preceding example. Let $\mathrm{PC}=$ $19 . \% 5$, $\mathrm{PCQ}=1 \%^{\circ} 30 \frac{1}{2}^{\prime}$, $\mathrm{PCR}=64^{\circ} 29 \frac{1}{2}^{\prime}$. Required OL. Ans. ${ }^{\mathrm{P}} \mathrm{P}=5 \cdot 94, \mathrm{PR}=1 \% \cdot 82$, and therefore, by the formula, $\mathrm{CL}=$
$2 \% \cdot 31$, or $C L^{\prime}=7 \cdot 69$; corresponding to the graphical solution. The figure is on the same scale.

If the given point were without the field, the division-line could be determined in a similar manner.
421. To part off a Quadrilateral. Conceive the two sides of the field which the division-line will intersect, D A and C B, produced till they meet at a point $G$, not shown in the figure. Calculate the triangle thus formed outside of the field. Its area, increased by the required area, will be that of the triangle EFG. Then the problem is identical with that in the last article. The following example is that given in Gummere's "Surveying." The figure represents it on a scale of 20 chains to 1 inch

Fig. 298.
 $=1: 15840$.

Example. A field is bounded thus : N. $14^{\circ} \mathrm{W} ., 15 \cdot 20$; N. $80 \frac{1}{2}^{\circ}$ E., $20 \cdot 43$; S. $6^{\circ}$ E., $22 \cdot 79$; N. $86 \frac{1}{2}^{\circ}$ W., $18 \cdot 00$. A spring within it bears from the second corner S . $75^{\circ} \mathrm{E} ., 7 \cdot 90$. It is required to cut off 10 acres from the west side of the field by a straight fence through the spring. How far will it be from the first corner to the point at which the division-fence meets the fourth side? Ans. $4.635 \%$ chains.
422. To part off any Figure. Let it be required to part off from

Fig. 299.
 a field a certain area by a line passing through a given point P within the field. Run a guess-line A B through $P$. Calculate the area which it parts off. Call the difference between it and the required area $=d$. Let CD be the desired line of dirision, and let P represent the angle, APC or BPD, which it makes with the given line. Obtain the angles $\mathrm{PAC}=\mathrm{A}$, and $\mathrm{PBD}=\mathrm{B}$, either by measurement, or by de-
duction from bearings. Measure PA and P B. Then the desired angle P will be given by the following formula:

$$
\begin{aligned}
& \text { Cot. } \mathrm{P}=-\frac{1}{2}\left(\cot . \mathrm{A}+\cot . \mathrm{B}-\frac{\mathrm{A} \mathrm{P}^{2}-\mathrm{B} \mathrm{P}^{2}}{2 d}\right) \pm \\
& {\left[\sqrt{ } \frac{\mathrm{A} \mathrm{P}^{2} \cdot \cot \cdot \mathrm{~B}-\mathrm{B}^{2} \cdot \cot \cdot \mathrm{~A}}{2 d}-\cot \cdot \mathrm{A} \cdot \cot \cdot \mathrm{~B}+\right.} \\
& \frac{1\left(\cot . \mathrm{A}+\cot . \mathrm{B}-\frac{\mathrm{A} \mathrm{P}^{2}-\mathrm{B} \mathrm{P}^{2}}{2 d}\right)^{2}}{2 d}
\end{aligned}
$$

If the guess-line be run so as to be perpendicular to one of the sides of the field, at A, for example, the preceding expression reduces to the following simpler form :

$$
\begin{aligned}
& \text { Cot. } \mathrm{P}=-\frac{1}{2}\left(\cot . \mathrm{B}-\frac{\mathrm{AP}^{2}-\mathrm{B} \mathrm{P}^{2}}{2 d}\right) \pm \\
& \sqrt{ }\left[\frac{\mathrm{A} \mathrm{P}^{2} \cdot \cot \cdot \mathrm{~B}}{2 d}+\frac{1}{4}\left(\cot \cdot \mathrm{~B}-\frac{\mathrm{A} \mathrm{P}^{2}-\mathrm{B} \mathrm{P}^{2}}{2 d}\right)^{2}\right] .
\end{aligned}
$$

Demonstration. The difference $d$, between the areas parted off by the guess-line AB, and the required line CD, is equal to the difference between the triangles A PC and B P D.

By Art. 61, note, the triangle APC $=\frac{1}{2} \cdot \mathrm{AP}^{2} \cdot \frac{\sin . \mathrm{A} \cdot \sin . \mathrm{P}}{\sin .(\mathrm{A}+\mathrm{P})}$.
Similarly, the triangle B PD $=\frac{1}{2} . \mathrm{BP}^{2} \frac{\sin . \mathrm{B} \cdot \sin . \mathrm{P}}{\sin .(\mathrm{B}+\mathrm{P})}$.

$$
\therefore d=\frac{1}{2} \cdot \mathrm{AP}^{2} \frac{\sin \cdot \mathrm{~A} \sin \cdot \mathrm{P}}{\sin \cdot(\mathrm{~A}+\mathrm{P})}-\frac{1}{2} \mathrm{BP}^{2} \cdot \frac{\sin \cdot \mathrm{~B} \cdot \sin \cdot \mathrm{P}}{\sin \cdot(\mathrm{~B}+\mathrm{P})} .
$$

By the expression for sin. $(a+b)$ [Trigonometry, Art. 8], we have
$d=\frac{1}{y}$ A P $P^{2} \cdot \frac{\sin . \mathrm{A} \cdot \sin . \mathrm{P}}{\sin . \mathrm{A} \cdot \cos . \mathrm{P}+\sin . \mathrm{P} \cdot \cos . \mathrm{A}}-\frac{1}{2} \mathrm{~B} \mathrm{P}^{2} \cdot \frac{\sin . \mathrm{B}}{\sin . \mathrm{B} \cdot \cos \cdot \mathrm{sin} . \mathrm{P}}+\sin \cdot \mathrm{P} \cdot \cos . \mathrm{B}$
Dividing each fraction by its numerator, and remembering that $\frac{\cos . a}{\sin . \alpha}=$ $=\cot . a$, we have

$$
d=\frac{\frac{1}{1} \mathrm{~A} \mathrm{P}^{2}}{\cot . \mathrm{P}+\cot . \mathrm{A}}-\frac{\frac{1}{2} \mathrm{~B} \mathrm{P}^{2}}{\cot . \mathrm{P}+\cot . \mathrm{B}} .
$$

For convenience, let $p=\cot$. P; $a=\cot$. A; and $b=\cot$. B. The above equation will then read, multiplying both sides by 2 ,

$$
2 d=\frac{\mathrm{A} p^{2}}{p+\mathrm{A}}-\frac{\mathrm{B} \mathrm{P}^{2}}{p+b} .
$$

Clearing of fractions, we have
$2 d p^{2}+2 d a p+2 d b p+2 d a b=p . \mathrm{AP}^{2}+b . \mathrm{AP}^{2}-p . \mathrm{BP}^{2}-a . \mathrm{BP}^{2}$.
Transposing, dividing through by $2 d$, and separating into factors, we get

$$
\begin{aligned}
& p^{2}+\left(a+b-\frac{\mathrm{AP}^{2}-\mathrm{BP}^{2}}{2 d}\right) p=\frac{b \cdot \mathrm{AP}^{2}-a \cdot \mathrm{BP}^{2}}{2 d}-a b . \\
& \because p=-\frac{1}{2}\left(a+b-\frac{\mathrm{AP}^{2}-\mathrm{BP}^{2}}{2 d}\right) \pm \cdot \mathrm{V}\left[\frac{b \cdot \mathrm{AP}^{2}-a \cdot \mathrm{BP}^{2}}{2 d}\right.
\end{aligned}
$$

$$
\left.-a b+\frac{1}{4}\left(a+b-\frac{\mathrm{AP}^{2}-\mathrm{BP}^{2}}{2 d}\right)^{2}\right] .
$$

If $\mathrm{A}=90^{\circ}$, cot. $\mathrm{A}=0$; and the expression reduces to the simpler form given in the article.

Example. It was required to cut off from a field twelre acres by a line passing through a spring P . A guess-line, A B , was run making an angle with one side of the field, at A , of $55^{\circ}$, and with the opposite side, at B , of $81^{\circ}$. The area thus cut off was found to be $13 \cdot 10$ acres. From the spring to A was $9 \cdot 30$ chains, and to B 3.30 chains. Required the angle which the required line, C D, must make with the guess-line, A B, at P. Ans. $20^{\circ} 45^{\prime}$; or $-86^{\circ} 25^{\prime}$. The heavy broken line, $\mathrm{C}^{\prime} \mathrm{D}^{\prime}$, shows the latter.

The scale of the figure is 10 chains to 1 inch $=1: \% 920$.
If the given point were outside of the field, the calculations would be similar.

## F. By the Shortest Possible Line.

423. To part off a Triangle. Let it be required to part off a

Fig. 300.
 triangular space, BDE , of given content, from the corner of a field, ABC , by the shortest possible line, D E.

From B set off BD and BE each equal to $\sqrt{ }\left(\frac{2 \mathrm{~B} \mathrm{DE}}{\sin . \mathrm{B}}\right)$. The line D E thus obtained will be perpendicular to the line, B F , which bisects the angle $B$. The length of $D E=$ $\frac{\sqrt{ }(2 . \mathrm{D} \mathrm{B} \mathrm{E.} \mathrm{sin.} \mathrm{B)}}{\cos \cdot \frac{1}{2} \mathrm{~B}}$.

Demonstration Conceive a perpendicular, BF, to be let fall from B to the required line DE . Let B represent the angle $\mathrm{D} \mathrm{B} \mathrm{E} ,\mathrm{and} \beta$ the unknown angle DBF. The angle BDF $=90^{\circ}-\beta$; and the angle BEF $=90^{\circ}-$ $(B-\beta)=90^{\circ}-B+\beta$. By Art. 61, note, the area of the triangle D BE $=\frac{1}{2} \mathrm{DE}^{2} \cdot \frac{\sin \cdot \mathrm{BDE} \cdot \sin \cdot \mathrm{BED}}{\sin \cdot(\mathrm{B} \mathrm{D} \mathrm{E} \mathrm{+} \mathrm{~B} \mathrm{E} \mathrm{D})}=\frac{1}{2} \cdot \mathrm{DE}^{2} \cdot \frac{\sin \cdot\left(90^{\circ}-\beta\right) \sin \cdot\left(90^{\circ}-\mathrm{B}+\beta\right)}{\sin \cdot \mathrm{B}}$.
Hence, $\mathrm{DE}^{2}=\frac{2 \times \mathrm{DBE} \times \sin . \mathrm{B}}{\sin .\left(90^{\circ}-\beta\right) \cdot\left(\sin .\left(90^{\circ}-\mathrm{B}+\beta\right)\right.}=\frac{2 \times \mathrm{D} \mathrm{BE} \times \sin . \mathrm{B}}{\cos \beta \cdot \cos \cdot(\mathrm{B}-\beta)}$.
Now, in order that D E may be the least possible, the denominator of the last fraction must be the greatest possible. It may be transformed, by the formula, cos. $a \cdot \cos . b=\frac{1}{8} \cos .(a+b)+\frac{1}{2} \cdot \cos .(a-b)$ [Trigonometry, Art. 8], into $\frac{1}{2} \cos . B+\frac{1}{2} \cdot \cos .(B-2 \beta)$. Since $B$ is constant, the ralue of
this expression depends on its second term, and that will be the greatest possible when $\mathrm{B}-2 \beta=0$, in which case $\beta=\frac{1}{2} \mathrm{~B}$.

It hence appears that the required line D E is perpendicular to the line, B F, which bisects the given angle B. This gives the direction in which DE is to be run.

Its starting-point, $D$ or $E$, is found thus: The area of the triangle D B E $=\frac{1}{2}$ B D. BE.sin. B. Since the triangle is isosceles, this becomes

$$
\mathrm{D} \mathrm{BE}=\frac{1}{2} \mathrm{~B} \mathrm{D}^{2} . \sin . \mathrm{B} ; \text { whence } \mathrm{B} \mathrm{D}=\sqrt{ }\left(\frac{2 \mathrm{D} \mathrm{~B} \mathrm{E}}{\sin . \mathrm{B}}\right) .
$$

D E is obtained from the expression for $\mathrm{D} \mathrm{E}^{2}$, which becomes, making $\beta=\frac{1}{2} \mathrm{~B}$,
$\mathrm{DE}^{2}=\frac{2 \times \mathrm{D} \mathrm{B} \mathrm{E} \times \sin . \mathrm{B}}{\cos \cdot \frac{1}{2} \mathrm{~B} \cdot \cos \cdot \frac{1}{2} \mathrm{~B}}$, whence, $\mathrm{DE}=\frac{\vee(2 \cdot \mathrm{D} \mathrm{BE} \cdot \sin . \mathrm{P})}{\cos \cdot \frac{1}{2} \mathrm{~B}}$.
Example. Let it be required to part off 1.3 acre from the corner of a field, the angle, B , being $30^{\circ}$. Ans. $\mathrm{BD}=\mathrm{BE}=\% \cdot 21$; and $\mathrm{DE}=3 \% \%$.

The scale of the figure is 10 chains to 1 inch $=1$ : \% 920 .

## G. Land of Variable Valde.

424. Let the figure represent a field in which the land is of two qualities and values, divided by the "quality-line" EF. It is required to part off from it a quantity of land worth a certain sum, by a straight fence parallel to A B.

Multiply the value per acre of each part by its length (in chains) on the line A B, add the products, multiply the value to be set off by 10 , divide by the above sum, and the quotient will be
 the desired breadth, BC or AD , in chains.

Demonstration. Let $\alpha=$ value per acre of one portion of the land, and $b$ that of the other portion. Let $x=$ the width required, BC or AD. Then the value of BCFE $=a \times \frac{x \times \mathrm{BE}}{10}$, and the value of $\mathrm{A} \mathrm{DFE}=b \times$ $\frac{x \times \mathrm{A} \mathrm{E}}{10}$.

Putting the sum of these equal to the value required to be parted off, we obtain $x=\frac{\text { value required } \times 10}{a \times \mathrm{BE}+b \times \mathrm{AE}}$.

Example. Let the land on one side of E F be worth $\$ 200$ per acre, and on the other side $\$ 100$. Let the length of the former, BE, be 10 chains, and EA be 30 chains. It is required to part off
a quantity of land worth $\$ 7,500$. Ans. The width of the desired strip will be 15 chains.

The scale of the figure is 40 chains to 1 inch $=1: 31680$.
If the "quality-line" be not perpendicular to $\mathrm{A} B$, it may be made so by "giving and taking," or as in the article following this one.

The same method may be applied to land of any number of different qualities; and a combination of this method with the preceding problems will solve any case which may occur.

## H. Straightening Crooked Fences.

425. It is often required to substitute a straight fence for a crooked one, so that the former shall part off precisely the same quantity of land as did the latter. This can be done on a plat by

Fig. 302.

the method given in Art. \%6, by which the irregular figure 1...2... $3 . . .4 . .5$ is reduced to the equivalent triangle $1 . . .5 . . .3^{\prime}$, and the straight line $5 . . .3^{\prime}$ therefore parts off the same quantity of land on either side as did the crooked one. The distance from 1 to $3^{\prime}$, as found on the plat, can then be set out on the ground and the straight fence be then ranged from 3 ' to 5 .

The work may be done on the ground more accurately by ran-
Fig. 303.

ning a guess-line, A C, Fig. 303, across the bends of the fence which crooks from A to B , measuring offsets to the bends on each
side of the guess-line, and calculating their content. If the sums of these areas on each side of A C chanced to be equal, that would be the line desired ; but if, as in the figure, it passes too far on one side, divide the difference of the areas by half of A C , and set off the quotient at right angles to AC , from A to D . D C will then be a line parting off the same quantity of land as did the crooked fence. If the fence at A was not perpendicular to A C, but oblique, as $A E$, then from $D$ run a parallel to $A C$, meeting the fence at $E$, and EC will be the required line.

## DIVIDING UP LAND.

426. Most of the problems for "dividing up" land may be brought under the cases in the preceding articles, by regarding one of the portions into which the figure is to be divided as an area to be "parted off" from it. Many of them, however, can be most neatly executed by considering them as independent problems, and this will be here done. They will be arranged, first, according to the simplicity of the figure to be divided up, and then subarranged, according to the manner of the division.

## Division of Triangles.

427. By Lines parallel to a Side. Suppose that the triangle ABC is to be divided into two equivalent parts by a line parallel to AC. The desired point, D , from which this line is to start, will be obtained by measuring $\mathrm{BD}=\mathrm{AB} \sqrt{ }$ $\frac{1}{2}$. So, too, E is fixed by $\mathrm{E}=\mathrm{BC} \sqrt{ } \frac{1}{2}$.

Fig. 304.


Generally, to divide the triangle into two parts, BDE and ACED, which shall have to each other a ratio $=m: n$, we have $\mathrm{BD}=\mathrm{AB} \sqrt{ } \frac{m}{m+n}$.

Fig. 305.


This may be constructed thus: Describe a semicircle on AB as a diameter. From B set off $\mathrm{BF}=\frac{m}{m+n} . \mathrm{BA} . \quad$ At F erect a perpendicular meeting the semicircle at $G$. Set off $B G$ from B to D. D is the starting-point of the division-
line required. In the figure, the two parts are as 2 to 3 , and BF is therefore $=\frac{2}{5} \mathrm{BA}$.

To divide the triangle A B C into five equivalent parts, we

Fig. 306.
 should have, similarly, $\mathrm{B} D=\mathrm{A} \mathrm{B} \sqrt{\frac{1}{5}}$; $\mathrm{BD}^{\prime}=\mathrm{AB} \sqrt{ } \frac{2}{5} ; \mathrm{BD}^{\prime \prime}=\mathrm{AB} \sqrt{\frac{3}{5}} ; \mathrm{BD}^{\prime \prime \prime}$ $=\mathrm{AB} \sqrt{ } \frac{4}{5}$.

The same method will divide the triangle into any desired number of parts having any ratios to each other.

## 428. By Lines perpendicular to a Side.

 Suppose that A B C is to be divided into two parts having a ratio $=m: n$, by a line perpendicular to $\mathrm{A} C$. Let EF be the dividing line whose position is required. Let BD be a perpendicular let fall from $B$ to $A C$. Then is $A E=\sqrt{ }(A C \times A D \times$ $\left.\frac{m}{m+n}\right)$. In this figure, $\mathrm{AFE}: \mathrm{EFBC}$Fig. 307.
 $:: m: n:: 1: 2$.

If the triangle had to be divided into two equiralent parts, the above expression would become $\mathrm{A} E=\sqrt{ }\left(\frac{1}{2} \mathrm{AC} \times \mathrm{AD}\right)$.

Demonstration. By hypothesis, AEF:EFBC::m:n; whence AEF: $\mathrm{ABC}: m: m+n ;$ and $\mathrm{AEF}=\mathrm{ABC} \frac{m}{m+n}=\frac{\mathrm{AC} \times \mathrm{DB}}{2} \cdot \frac{m}{m+n}$. Also, $\mathrm{AEF}=\frac{1}{2} . \mathrm{AE} \times \mathrm{EF}$.

The similar triangles AEF and ABD give $\mathrm{AD}: \mathrm{DB}:: \mathrm{AE}: \mathrm{EF}=$ $\frac{\mathrm{DB} \times \mathrm{AE}}{\mathrm{AD}}$. The second expression for AEF then becomes $A E F=$ $\frac{1}{2} \mathrm{AE} \cdot \frac{\mathrm{DB} \times \mathrm{AE}}{\mathrm{AD}}$. Equating this with the other ralue of AEF , we have $\frac{\mathrm{A} \mathrm{C} \times \mathrm{D} \mathrm{B}}{2} \cdot \frac{m}{m+n}=\frac{\mathrm{A} \mathrm{E}^{2} \times \mathrm{D} \mathrm{B}}{2 \cdot \mathrm{AD}} ;$ whence $\mathrm{A} \mathrm{E}=\sqrt{ }\left(\mathrm{AC} \times \mathrm{AD} \times \frac{m}{m+n}\right)$.
429. By Lines running in any Given Direction. Let a triangle, $\mathrm{A} \mathrm{B} \mathrm{C} ,\mathrm{be} \mathrm{given} \mathrm{to} \mathrm{be} \mathrm{divided} \mathrm{into} \mathrm{two} \mathrm{parts} \mathrm{haring} \mathrm{a} \mathrm{ratio}=m:$,$n ,$ by a line making a giren angle with a side. Part off, as in Art. 413 or 414 , Fig. 292, an area B F G $=\frac{m}{m+n}$. A B C.
430. By Lines starting from an Angle. Divide the side opposite to the given angle into the required number of parts, and draw lines from the angle to the points of division. In the figure the triangle is represented as being thus divided into two equivalent parts.

If the triangle were required to be divided

Fig. 308.
 into two parts, having to each other a ratio $=$ $m: n$, we should have $\mathrm{A}=\mathrm{AC} \frac{m}{m+n}$, and $\mathrm{DC}=\mathrm{AC} \frac{n}{m+n}$.

Fig. 309.
 parts which should be to each other $:: m: n$ $: p$, we should have $\mathrm{AD}=\mathrm{AC} \frac{m}{m+n+p}$, $\mathrm{D} \mathrm{E}=\mathrm{A} C \frac{m}{m+n+p}$, and $\mathrm{EC}=\mathrm{AC}$ $\frac{p}{m+n+p}$.
Suppose that a triangular field, A B C, had to be divided among five men, two of them to have a quarter each, and three of them each a sixth. Divide AC into two equal parts, one of these again into two equal parts, and the other one into three equal parts. Run the lines from the four points thus obtained to the angle B.
431. By Lines starting from a Point in a Side. Suppose that the triangle ABC is to be divided into two equivalent parts by a line starting from a point D in the side AC . Take a point E in the middle of AC. Join B D, and from E draw a parallel to it, meeting AB in F . D F will be the dividing line required.


The point F will be most easily obtained on the ground by the proportion $\mathrm{AD}: \mathrm{AB}:: \mathrm{AE}=\frac{1}{2} \mathrm{AC}: \mathrm{AF}$.

The altitude of AFD of course equals $\frac{1}{2} \mathrm{ABC} \div \frac{1}{2} \mathrm{AD}$.
If the triangle is to be divided into two parts having any other ratio to each other, divide AC in that ratio, and then proceed as before. Let this ratio $=m: n$, then $\mathrm{AF}=\frac{\mathrm{AB} \times \mathrm{AC}}{\mathrm{AD}} \cdot \frac{m}{m+n}$.

Demonstration. In Fig. 310, conceive the line EB to be drawn. The triangle $\mathrm{AEB}=\frac{1}{2} \mathrm{ABC}$, having the same altitude and half the base; and $A F D=A E B$, because of the equivalency of the triangles EFD and EFB. which, with AEF, make up AFD and AEB.

The point F is fixed by the similar triangles A D B and A E F.
The expression for AF, in the last paragraph, is given by the proportion, ABC:ADF: : AB $\times \mathrm{AC}: \mathrm{AD} \times \mathrm{AF}$;
whence,

$$
A F=\frac{A B \times A C}{A D} \cdot \frac{A D F}{A B C}=\frac{A B \times A C}{A D} \cdot \frac{m}{m+n} .
$$

Next suppose that the triangle A B C is to be divided into three equivalent parts, meeting at $D$. The

Fig. 311.
 altitudes, EF and GH, of the parts ADE and DCG, will be obtained by dividing $\frac{1}{3} \mathrm{ABC}$, by half of the respective bases $A D$ and DC.

If one of these quotients gires an altitude greater than that of the triangle A B C, it will show that the two lines D E and D G would both cut the same side, as in Fig. 312, in which E F is obtained as above, and G H $=\frac{2}{3} \mathrm{ABC} \div \frac{1}{2} \mathrm{AD}$.

In practice it is more convenient to determine the points $F$ and $G$, by these proportions:
BK : AK : : EF : AF ; and BK : AK

Fig. 312.

: : GH: A H.

The division of a triangle into a greater number of parts, haring any ratios, may be effected in a similar manner.

This problem admits of a more elegant solution, analogous to that given for the division into two

Fig. 313.
 parts, graphically. Divide A C into three equal parts at $L$ and M. Join $B \mathrm{D}$, and from L and M draw parallels to it, meeting $A B$ and $B C$ in $E$ and G. Draw ED and GD, which will be the desired lines of division. The figure is the same triangle as Fig. 311.

The points $E$ and $G$ can be obtained on the ground by measur-
ing AD and AB , and making the proportion $\mathrm{AD}: \mathrm{AB}:: \frac{1}{3} \mathrm{AC}$ : AE. The point $G$ is similarly obtained.

The same method will divide a triangle into a greater number of parts.

To divide a triangle into four equivalent triangles by lines terminating in the sides, is very easy. From D, the middle point of AB, draw D E parallel to AC, and from $F$, the middle of AC , draw FD and FE . The problem is now solved.

Fig. 314.


## 432. By Lines passing through a Point

 within the Triangle. Let D be a given point (such as a well, etc.) within a triangular field ABC, fromFig. 315.
 which fences are to run so as to divide the triangle into two equivalent parts. Join A D. Take E in the middle of B C, and from it draw a parallel to D A , meeting A C in F. ED F is the fence required.

If it be required to divide a triangle into two equivalent parts by a straight line passing through a point within it, proceed thus: Let $P$ be the given point. From P draw P D parallel to A C, and PE parallel to BC. Bisect A C at F. Join F D. From B draw BG parallel to DF. Then bisect G C in H. On HE describe a semicircle. On it set off EK = EC. Join KH. Set off $\mathrm{HL}=\mathrm{HK}$. The line LM drawn from L , through P , will be the division-line required.

This figure is the same as that of Art. 416. The triangle A B C contains 62.35 acres, and the distance $\mathrm{CL}=2 \% \cdot 31$ chains, as in the example in that article.

433. Next suppose that the triangle $A B C$ is to be divided into

Fig. 317.
 three equivalent parts by lines starting from a point $D$, within the triangle, given by the rectangular co-ordinates A E and ED. Let ED be one of the lines of division, and $F$ and $G$ the other points required. The point $F$ will be determined if AH is known; AH and $H F$ being its rectangular co-ordinates. From $B$ let fall the perpendicular BK on $\mathrm{A} C$.
Then is $A H=\frac{A K\left(\frac{2}{3} A B C-A E \times E D\right)}{A E \times B K-E D \times A K}$. The position of the other point, $G$, is determined in a similar manner.

Demonstration. Let $\mathrm{A} \mathrm{E}=x, \mathrm{E} \mathrm{D}=y, \mathrm{AH}=x^{\prime}, \mathrm{HF}=y^{\prime}, \mathrm{A} \mathrm{K}=a$, $K B=b$.

The quadrilateral AFDE, equivalent to $\frac{1}{3}$ ABC, but which we will represent generally by $m^{2}$, is made up of the triangle AFH and the trapezoid F HED.

$$
\mathrm{A} \mathrm{~F} \mathrm{H}=\frac{1}{2} \cdot x^{\prime} y^{\prime} . \quad \text { F H E D }=\frac{1}{2}\left(x-x^{\prime}\right)\left(y+y^{\prime}\right) .
$$

$\therefore$ AFDE $=m^{2}=\frac{1}{2} \cdot x^{\prime} y^{\prime}+\frac{1}{2}\left(x-x^{\prime}\right)\left(y+y^{\prime}\right)=\frac{1}{2} x(y+y)-\frac{1}{2} x^{\prime} y$ 。 The similar triangles, A H F and AKB, give

$$
a: b:: x^{\prime}: y^{\prime}=\frac{b x^{\prime}}{a}
$$

Substituting this value of $y^{\prime}$ in the expression for $m^{2}$, we hare

$$
m^{2}=\frac{1}{2} x\left(y+\frac{b x^{\prime}}{a}\right)-\frac{1}{2} x^{\prime} y
$$

whence,

$$
x^{\prime}=\frac{a\left(2 m^{2}-x y\right)}{b x-a y}=\frac{\mathrm{AK}\left(\frac{2}{3} \mathrm{~A} \mathrm{BC}-\mathrm{AE} \times \mathrm{ED}\right)}{\mathrm{KB} \times \mathrm{AE}-\mathrm{AK} \times \mathrm{ED}}
$$

The formula is general, whaterer may be the ratio of the area $m^{2}$ to that of the triangle $A B C$.

Let DB , instead of DE , be one of the required lines of division. Divide $\frac{1}{3}$ A B C by half of the perpendicular DH , let fall from $D$ to $A B$, and the quotient will be the distance BF. To find G, if, as in this figure, the triangle $\mathrm{BDC}(=$ $\mathrm{BC} \times \frac{1}{2} \mathrm{DK}$ ) is less than $\frac{1}{3}$ A BC, divide the excess of the

latter (which will be CDG) by $\frac{1}{2} \mathrm{DE}$, and the quotient will be CG.

Example. Let $\mathrm{AB}=30.00 ; \mathrm{B} \mathrm{C}=45.00 ; \mathrm{CA}=50.00$. Let the perpendiculars from D to the sides be these : $\mathrm{DE}=10.00$; $\mathrm{DH}=20.00 ; \mathrm{DK}=5 \cdot 17 \frac{1}{3}$. The content of the triangle ABC will be $666 \cdot 6$ square chains. Each of the small triangles must therefore contain $222 \cdot 2$ square chains, B D being one division-line. We shall therefore have $\mathrm{BF}=222 \cdot 2 \div \frac{1}{2} \mathrm{DH}=22 \cdot 2$ chains. $B D C=45 \times \frac{1}{2} \times 5 \cdot 1 \% \frac{1}{3}=116.4$ square chains, not enough for a second portion, but leaving 105.8 square chains for CD G; whence $C G=21 \cdot 16$ chains. To prove the work, calculate the content of the remaining portion, GDFA. We shall find DGA $=144{ }^{2} 2$ square chains, and $\mathrm{ADF}=78 \cdot 0$ square chains, making together $222 \cdot 2$ square chains, as required.

The scale of Fig. 318 is 30 chains to 1 inch $=1: 23 \% 60$.
434. The preceding case may be also solved graphically, thus: Take C L $=\frac{1}{3}$ A C. Join D L, and from B draw B G parallel to D L. Join D G. It will be a second line of division. Then take a point, M , in the middle of $B G$, and from it draw a line, M F, parallel to DA. D F will be the third line of division.
 This method is neater on paper than the preceding, but less convenient on the ground.

Demonstration. In Fig. 319 D G is a second line of division, because, drawing BL , the triangle $\mathrm{BLC}=\frac{1}{3} \mathrm{ABC}$; and BDGC is equivalent to BL C, because of the common part B C L D, and the equivalency of the triangles D L G and D LB.

To prove that D F is a third line of division, join MD and MA. Then $\mathrm{BMA}=\frac{1}{2} \mathrm{BGA}$. From BMA take MFA and add its equivalent MFD, and we have MDFB $=\frac{1}{2} B G A=\frac{1}{2}(\mathrm{ABDG}-\mathrm{BDG})=\frac{1}{2}\left(\frac{2}{3} \mathrm{ABC}-\right.$ $B D G)=\frac{1}{3} A B C-\frac{1}{2} B D G$. To MDFB add MDB, and add its equivalent, $\frac{1}{2} \mathrm{BDG}$, to the other side of the equation, and we have $M D F B+M D B=\frac{1}{3} A B C-\frac{1}{2} B D G+\frac{1}{2} B D G ;$ or, $B D F=\frac{1}{3} A B C$.
435. Let it be required to divide the triangle ABC into three equivalent triangles, by lines drawn from the three angular points
to some unknown point within the triangle. This point is now to be found. On any side, as A B, take $\mathrm{AD}=\frac{1}{3} \mathrm{~A} B$. From D draw D E parallel to A C. The middle, F, of D E, is the point required.

If the three small triangles are not to be equivalent, but are to have to each other the ratios $: m: n: p$, divide a side, AB , into parts haring these ratios, and through each point of division, $D, E, d r a w ~ a ~$ parallel to the side nearest to it. The intersection of these parallels, in $F$, is the point required. In the figure the parts ACF, ABF, BCF, are as $2: 3: 4$.

Fig. 321.

436. Let it be required to find the position of a point, D , situated within a given triangle, ABC , and equally distant from the points,

Fig. 322.
 $\mathrm{A}, \mathrm{B}, \mathrm{C}$; and to determine the ratios to each other of the three triangles into which the given triangle is divided.

By construction, find the center of the circle passing through $\mathrm{A}, \mathrm{B}, \mathrm{C}$. This will be the required point.
By calculation, the distance $\mathrm{DA}=\mathrm{DB}=\mathrm{DC}=\frac{\mathrm{AB} \times \mathrm{BC} \times \mathrm{CA}}{4 \times \text { area } \mathrm{ABC}}$.
The three small triangles will be to each other as the sines of their angles at $\mathrm{D}-\mathrm{i}$. e., ADB : A D C : BDC : : $\sin$. AD B $: \sin$. A D C $: \sin$. $B D C$. These angles are readily found, since the sine of half of each of them equals the opposite side divided by twice one of the equal distances.
437. By the Shortest Possible Line. Let it be required to divide the triangle ABC by the short-

est possible line, D E, into two parts, which shall be to each other $:: m: n$; or D B E : A B C $:: m: m+n$.

From the smallest angle, B , of the triangle, measure along the sides, B A and BC , a distance $\mathrm{B} D=\mathrm{BE}=\sqrt{ }\left(\frac{m}{m+n} \times \mathrm{AB} \times \mathrm{BC}\right)$.
DE is the line required. It is perpendicular to the line BF which bisects the angle A B C ; and it is

$$
=\frac{\sin \cdot \mathrm{B}}{\cos \cdot \frac{1}{2} \mathrm{~B}} \sqrt{ }\left(\frac{m}{m+n} \times \mathrm{AB} \times \mathrm{BC}\right) .
$$

The formulas are obtained from Art. 419.

## Division of Rectangles.

438. By Lines parallel to a Side. Divide two opposite sides into the required number of parts, either equal or in any given ratio to each other, and the lines joining the points of division will be the lines desired.

The same method is applicable to any parallelogram.
Example. A rectangular field A B C D, measuring 15.00 chains by 8.00 , is bought by three men, who pay respectively $\$ 300, \$ 400$, and $\$ 500$. It is to be divided among them in that proportion. Ans. The portion of the first, A E E'B, is obtained by making the proportion $300+400+500: 300$ $:: 15 \cdot 00: \mathrm{A} \mathrm{E}=3 \% \mathrm{~F} . \mathrm{EF}$ is in

## Fig. 324.

 like manner found to be $5 \cdot 00$; and $\mathrm{FD}=6 \cdot 25 . \mathrm{BE}$ is made equal to $\mathrm{AE} ; \mathrm{E}^{\prime} \mathrm{F}^{\prime}$ to EF ; and $\mathrm{F}^{\prime} \mathrm{C}$ to FD . Fences from E to $\mathrm{E}^{\prime}$, and from F to $\mathrm{F}^{\prime}$, will divide the land as required.

The scale of the figure is 10 chains to 1 inch $=1: 7920$.
The other modes of dividing up rectangles will be given under the head of "Quadrilaterals," Art. 443, etc.

## Division of Trapezoids.

439. By Lines parallel to the Bases. Given the bases and a third side of the trapezoid, A B C D, to be divided into two parts, such that BCFE:EFDA: : $m: n$.

The length of the desired dividing line,

Fig. 325.


$$
\mathrm{EF}=\sqrt{ }\left(\frac{m \times \mathrm{AD}^{2}+n \times \mathrm{B} \mathrm{C}^{2}}{m+n}\right)
$$

The distance $B E=\frac{A B(E F-B C)}{A D-B C}$.
Demonstration. In Fig. 325, conceive the sides $A B$ and D C, produced, to meet in some point $P$. Then, by reason of the similar triangles, ADP: BCP : : A D ${ }^{2}$ : B C ${ }^{2}$, whence, by "division," $\mathrm{ADP}-\mathrm{BCP}=\mathrm{ABCD}: \mathrm{BCP}:: \mathrm{A} \mathrm{D}^{2}-$ $\mathrm{B} \mathrm{C}^{2}: \mathrm{B} \mathrm{C}^{2}$.

In like manner, comparing $E F P$ and $B C P$, we get $\mathrm{EBCF}: \mathrm{BCP}: E \mathrm{~F}^{2}-\mathrm{BC}^{2}: \mathrm{BC}^{2}$. Combining these two proportions, we hare $\mathrm{ABCD}: \mathrm{EBCF}:: A \mathrm{D}^{2}-\mathrm{BC}^{2}: \mathrm{EF}^{2}-\mathrm{BC}^{2}$; or, $\quad m+n: m:: \mathrm{A}^{2}-\mathrm{BC}^{2}: \mathrm{EF}^{2}-\mathrm{BC}^{2}$. Whence, $(m+n) \mathrm{EF}^{2}-m . \mathrm{BC}^{2}-n \mathrm{BC}^{2}=m$. $\mathrm{AD}^{2}-m . \mathrm{BC}^{2}$;
$\therefore \mathrm{EF}=1\left(\frac{m \times \mathrm{AD}^{2}+n \times \mathrm{B} \mathrm{C}^{2}}{m+n}\right)$.
Also, from the similar triangles formed by drawing BL parallel to C D, we bave

$$
A L: E K:: B A: B E=\frac{B A \times E K}{A L}=\frac{A B(E F-B C)}{A D-B C}
$$

Example. Let $\mathrm{A} D=30$ chains ; $\mathrm{B}=20$ chains ; and $\mathrm{AB}=$ $54 \frac{1}{3}$ chains ; and the parts to be as 1 to 2 ; required EF and B E. Ans. $\mathrm{EF}=23 \cdot 80$; and $\mathrm{BE}=20 \cdot 65$.

The figure is on a scale of 30 chains to 1 inch $=1: 23 \% 60$.
440. Given the bases of a trapezoid, and the perpendicular distance, BH , between them; it is required to divide it as before, and to find EF, and the altitude, B G, of one of the parts. Let $B C F E: E F D A:: m: n$. Then $B G=-\frac{B C \times B H}{A D-B C}+$

$$
\begin{gathered}
\sqrt{\left[\frac{m}{m+n} \times \frac{2 \times \mathrm{ABCD} \times \mathrm{BH}}{\mathrm{AD}-\mathrm{BC}}+\left(\frac{\mathrm{BC} \times \mathrm{BH}}{\mathrm{AD}-\mathrm{BC}}\right)^{2}\right]} \\
\mathrm{EF}=\mathrm{BC}+\mathrm{BG} \times \frac{\mathrm{AD}-\mathrm{BC}}{\mathrm{BH}} .
\end{gathered}
$$

Demonstration. Let $\mathrm{BEFC}=\frac{m}{m+n} . \mathrm{ABCD}=a$; let $\mathrm{BC}=b ; \mathrm{B} \mathrm{H}$ $=h ;$ and $\mathrm{AD}-\mathrm{BC}=c$. Also, let $\mathrm{B} \mathrm{G}=x$; and $\mathrm{EF}=y$. Dram B L parallel to CD. By similar triangles, AL:EK::BA:BE::BH:
$\mathrm{BG} ;$ or, $\mathrm{AD}-\mathrm{BC}: \mathrm{EF}-\mathrm{BC}:: \mathrm{BH}: \mathrm{B} \mathrm{G} ;$ i. e., $c: y-b:: h: x ;$ whence $x=\frac{h(y-b)}{c}$.
Also, the area $\mathrm{BEFC}=a=\frac{1}{2} \cdot \mathrm{~B} \mathrm{G}(\mathrm{EF}+\mathrm{BC})=\frac{1}{2} x(y+b) ;$ whence $y=\frac{2 a}{x}-b$.
Substituting this value of $y$ in the expression for $x$, and reducing, we obtain $x^{2}+\frac{2 b h}{c} x=\frac{2 a h}{c} ;$ whence we have $x=-\frac{b h}{c} \pm \sqrt{ }\left(\frac{2 a h}{c}+\frac{b^{2} h^{2}}{c^{2}}\right)$.
The second proportion above gives $y-b=\frac{c x}{\hbar}$; whence $y=b+\frac{c}{\hbar} \cdot x$.
Replacing the symbols by their lines, we get the formulas in the text.
Example. Let A D $=30.00 ; \mathrm{BC}=20 \cdot 00 ; \mathrm{BH}=54.00$; and the two parts to be to each other $:: 46: 89$.

The above data give the content of $\mathrm{ABCD}=1,350$ square chains. Substituting these numbers in the above formula, we ob$\operatorname{tain} \mathrm{B} G=20 \cdot 96$, and $\mathrm{EF}=23 \cdot 88$.
441. By Lines starting from Points in a Side. To divide a trapezoid into parts equivalent, or having any ratios, divide its parallel sides in the same ratios, and join the corresponding points.

If it "be also required that the division-lines shall start from given points on a side, proceed thus: Let it be required to divide the trapezoid A BCD into three equivalent parts by fences starting from P and Q. Divide the trapezoid, as above directed, into three equivalent trapezoids by the lines E F
 and G H. These three trapezoids must now be transformed, thus : Join EP, and from F draw F R parallel to it. Join P R, and it will be one of the divis-ion-lines required.

The other division-line, QS , is obtained similarly.
442. Other Cases. For other cases of dividing trapezoids, apply those for quadrilaterals in general, given in the following articles.*

[^53]
## Division of Quadrilaterals.

443. By Lines parallel to a Side. Let ABCD be a quadrilateral which it is required to divide, by a line E F, parallel to

Fig. 327.
 A D, into two parts, B E F C and EFD A, which shall be to each other as $m: n$. Prolong $A B$ and $C D$ to intersect in G. Let $a$ be the area of the triangle ADG , obtained by any method, graphical or trigonometrical, and $a^{\prime}=$ the area of the triangle $B C G$, obtained by subtracting the area of the given quadrilateral from that of the triangle AD G. Then $\mathrm{GK}=\mathrm{GH} \sqrt{ } /\left(\frac{m a+n}{(m+n)} a^{\prime}\right)$. Having measured this length of $G K$ from $G$ on $G H$, set off at $K$ a perpendicular to $G K$, and it will be the required line of division.

Demonstration. In Fig. 327, since EF is parallel to A D, we have $\mathrm{ADG}: \mathrm{EGF}:: \mathrm{GH}^{2}: \mathrm{GK}^{2}$. EGF is made up of the triangle $\mathrm{BCG}=a^{\prime}$, and the quadrilateral $\mathrm{BEFC}=\frac{m}{m+n} \cdot \mathrm{ABCD}=\frac{m}{m+n} \cdot\left(a-a^{\prime}\right)$. Hence the above proportion becomes

$$
a: a^{\prime}+\frac{m}{m+n}\left(a-a^{\prime}\right):: \mathrm{GH}^{2}: \mathrm{GK}^{2} ; \text { or, }
$$

$(m+n) a: m a+n a^{\prime}:: \mathrm{GH}^{2}: \mathrm{GK}^{2}$; whence $\mathrm{GK}=\mathrm{GH} /\left(\frac{m a+n a^{\prime}}{(m+n) a}\right)$.
GE is given by the proportion $\mathrm{GH}: \mathrm{GK}:: \mathrm{GA}: \mathrm{GE}=\mathrm{GA} \cdot \frac{\mathrm{GK}}{\mathrm{GH}}$.
In Fig. 328 , the division into $p$ parts is founded on the same principle. The triangle EFG=GBC+EFCB= $a^{\prime}+\frac{\mathrm{Q}}{p}$. Now ADG:EFG: : $\mathrm{A} \mathrm{G}^{2}: \mathrm{EG}^{2} ;$ or, $a^{\prime}+\mathrm{Q}: a^{\prime}+\frac{\mathrm{Q}}{p}:: \mathrm{A} \mathrm{G}^{2}: \mathrm{EG}^{2}$;

$$
\text { whence } \left.G E=A G \sqrt{\left(\frac{a^{\prime}+\frac{Q}{p}}{a^{\prime}+Q}\right.}\right)
$$

zoid, any line drawn through the middle of the first line will diride the trapezoid into two equiralent parts.

GL is obtained by taking the triangle $\mathrm{LMG}=a^{\prime}+\frac{2 \mathrm{Q}}{p}$; and so for the rest.

Otherwise, take $\mathrm{G} \mathrm{E}=\mathrm{G} A \sqrt{ }\left(\frac{m a+n a^{\prime}}{(m+n) a}\right)$; and from E run a parallel to A D.

If the two parts of the quadrilateral were to be equivalent, $m=n$, and we have $\mathrm{GK}=\mathrm{GH} \sqrt{ } /\left(\frac{a+a^{\prime}}{2 a}\right)$; and consequently $G E$ to $G A$ in the same ratio.

Example. Let a quadrilateral, A B C D, be required to be thus divided, and let its angles, B and C , be given by rectangular coordinates, viz., $A \mathrm{~B}^{\prime}=6.00 ; \mathrm{B}^{\prime} \mathrm{B}=9.00 ; \mathrm{D} \mathrm{C}^{\prime}=8.00 ; \mathrm{C}^{\prime} \mathrm{C}=$ $13.00 ; \mathrm{B}^{\prime} \mathrm{C}^{\prime}=24.00$. Here G H is readily found to be 29.64 ; $\mathrm{ADG}=563 \cdot 16$ square chains ; and $\mathrm{B} \mathrm{G} \mathrm{C}=220 \cdot 16$ square chains. Hence, by the formula, $G K=24.72$; whence $K H=G H-G K$ $=4.92$; and the abscissas for the points E and F can be obtained by a simple proportion.

The scale of the figure is 20 chains to 1 inch $=1: 15840$.
If the quadrilateral be given by bearings, part off the desired area $=\frac{n}{m+n} \cdot$ A B C D, by the formulas of Art. 403.

Suppose now that a quadrilateral, ABCD , is to be divided into $p$ equivalent parts, by lines parallel to A D. Measure, or calculate by trigonometry, A G. Let Q be the quadrilateral ABCD, and, as before, $a^{\prime}=\mathrm{BCG}$. Then $\mathrm{GE}=\mathrm{AG} \sqrt{ }\left\{\begin{array}{l}a^{\prime}+\frac{\mathrm{Q}}{p} \\ \frac{a^{\prime}+\mathrm{Q}}{}\end{array}\right\} ; \mathrm{GL}=\mathrm{AG} \sqrt{ }\left\{\begin{array}{l}a^{\prime}+\frac{2 \mathrm{Q}}{p} \\ \frac{a^{\prime}+\mathrm{Q}}{}\end{array}\right\} ;$


If the quadrilateral be given by bearings, part off $\frac{1}{p}$. A B C D, then part off $\frac{2}{p}$. A B C D, etc. ; so in any similar case.
444. By Lines perpendicular to a Side. Let ABCD be a quadrilateral which is to be dirided, by

Fig. 329.
 a line perpendicular to $\mathrm{A} D$, into two parts haring a ratio $=m: n$. By hypothesis, $\mathrm{ABEF}=\frac{m}{m+n}, \mathrm{ABCD}$. Taking array the triangle ABG, the remainder, GBEF, will be to the rest of the figure in a known ratio, and the position of EF, paralle] to B G, will be found as in the last article.
445. By Lines running in any Given Direction. To divide a quadrilateral ABCD into two parts $:: m: n$, part off from it an area $=\frac{m}{m+n} \cdot \mathrm{ABCD}$, by the methods of Art. $40 \%$ or 408, if the area parted off is to be a triangle, or Art. 409 if the area parted off is to be a quadrilateral.

## 446. By Lines starting from

 an Angle. A B CD is to be divided, by the line CE , into two parts haring the ratio $m: n$. Since the area of the triangle $\mathrm{CDE}=\frac{m}{m+n} . \mathrm{ABCD}, \quad \mathrm{DE}$ will be obtained by dividing this area by half of the altitude C F.447. By Lines starting from Points in a Side. Let it be required to diride ABCD into tro

Fig. 351.
 parts : : $m$ : $n$, by a line starting from the point $E$. The area $A B F E$ is known (being $=\frac{m}{m+n}$. ABCD ) as also ABE ; $\mathrm{AB}, \mathrm{BE}$, and EA being giren on the ground. BEF will then be known $=$ ABFE-
$A B E$. Then $G F=\frac{B E F}{\frac{1}{2} B E}$, and the point $F$ is obtained by
running a parallel to $\mathrm{B} E$, at a perpendicular distance from it $=G \mathrm{~F}$.

To divide a quadrilateral, A B C D, graphically, into two equivalent parts by a line from a point, E, on a side, proceed thus : Draw the diagonal CA , and from B draw a parallel to it, meeting D A prolonged in F. Mark the middle point, G, of FD. Join GE. From C draw a parallel to EG , meeting DA in H . EH is the required line. The quad-
 rilateral could also be divided in any ratio $=m: n$, by dividing FD in that ratio.

If the quadrilateral be given by bearings, proceed to part off the desired area, as in Art. 412 or 413.
448. Let it be required to divide a quadrilateral, A BCD , into three equivalent parts.

## Fig. 333.

 From any angle, as C, draw C E, parallel to D A. Divide AD and E C, each into three equal parts, at $F, F^{\prime}$, and $G, G^{\prime}$. Draw BF, BF'. From G draw G H, parallel to FB , and from $G^{\prime}$ draw $G^{\prime} H^{\prime}$, parallel to $\mathrm{F}^{\prime} \mathrm{B}$. FH and $\mathrm{F}^{\prime} \mathrm{H}^{\prime}$ are the required lines of division.

Let it be required to make the above division by lines starting from two given points, P and Q . Reduce the quadrilateral to an equivalent triangle C B E. Divide EB into three equal parts at F and G. Join C Q, and, from G, draw GK parallel

Fig. 334.

to it. Join C P, and from F draw F L parallel to it. Join P L and QK , and they will be the division-lines required.
449. By Lines passing through a Point within the Figure. Proceed to part off the desired area as in Art. 416 or 41\%, according to the circumstances of the case.

## Division of Polygons.

450. By Lines running in any Direction. Let A BCDEFG
 be a given polygon, and B H the direction parallel to which is to be drawn a line $\mathrm{P} Q$, dividing the polygon into two parts in any desired ratio $=m: n$. The area $\mathrm{PCDEQ}=\frac{m}{m+n}$. ABC DEFG. Taking it from the area BCDEH , the remainder will be the area BPQH. The quadrilateral BCEH, CE being supposed to be drawn, can then be dirided by the method of Art. 443 into two parts, B PQH and PQEC, having to each other a known relation.

If D K were the given direction, at right angles to the former, the position of a dividing line $R S$ could be similarly obtained.

Fig. 336.

451. By Lines starting from an Angle. Produce one side, A B, of the given polygon, both ways, and reduce the polygon to a single equivalent triangle, X Y Z. Then divide the base, X Y, in the required ratio, as at W , and draw Z W, which will be the divisionline desired. In this figure the polygon is divided into two equivalent parts.

If the division-line should pass outside of the polygon, as does Z P, through P draw a parallel to B Z, meeting the adjacent side of the polygon in Q , and Z Q will be the division-line desired.
452. By Lines starting from a Point on a Side. See Articles 414 and 415.
453. By Lines passing through a Point within the Figure. Part off, as in Art. 416 or 418, if a straight line be required, or by guess-lines and the addition of triangles, as in Art. 433, if the lines have merely to start from the point, such as a spring or well.
454. Other Problems. The following is from Gummere's "Surveying" : Question. A tract of land is bounded thus: N. $35 \frac{1}{4}^{\circ} \mathrm{E}$., 23.00 ; N. $75 \frac{1}{2}^{\circ}$ E., 30.50 ; S. $3 \frac{1}{4}^{\circ}$ E., $46 \cdot 49$; N. $66 \frac{1}{4}^{\circ}$ W., $49 \cdot 64$. It is to be divided into four equivalent parts by two straight lines, one of which is to run parallel to the third side ; required the distance of the

Fig. 337. parallel division-line from the first corner, measured on the fourth side; also the bearing of the other division-line, and its distance from the same corner measured on the first side. Ans. Distance of the parallel division-line from the first corner, 32.50 ; the bearing of the other, S. $88^{\circ}$ $22^{\prime}$ E. ; and its distance from the same
 corner 5.99.

The scale of the figure is 40 chains to 1 inch $=1: 31680$.
An indefinite number of problems on this subject might be proposed, but they would be matters of curiosity rather than of utility, and exercises in geometry and trigonometry rather than in surveying.

Fig. 338.


## CHAPTER VII.

## THE PUBLIC LANDS OF THE UNITED STATES.*

455. General System. The public lands of the United States of America are generally divided and laid out into squares (approximately), the sides of which run truly north and south, or east and west.

This is effected by means of meridian lines and parallels of latitude, established six miles apart. The squares thus formed are called Townships. They contain 36 square miles, or 23,040 acres, " as nearly as may be." A principal meridian, running due north and south, and a base-line, running due east and west, are first established astronomically, and the half-mile, mile, and six-mile corners are permanently marked on them. These two lines form the basis of all the subsequent subdivision into townships and sections. All of the lines on the public surveys, except these two and the standard parallels, are run with compass and chain.

The map, Fig. 338, represents a portion of the State of Oregon thus laid out. The scale is 10 miles to 1 inch $=1$ : 633600. On it will be seen the "Willamette meridian," running truly north and south, and a "base-line," which is a " parallel of latitude," running truly east and west. Parallel to these, and six miles from them, are other lines, forming townships. All the townships, situated north or south of each other, form a RANGE. The ranges are named by their number east or west of the principal meridian. In the figure are seen three ranges east and west of

[^54]the Willamette meridian. They are noted as R. I. E., R. I. W., etc. The townships in each range are named by their number north or south of the base-line. In the figure, along the principal meridian, are seen four north and five south of the base-line. They are noted as T. 1 N., T. 2 N., T 1 S., etc.*

Each township is divided into 36 sections, each one mile square, and therefore containing, "as
 nearly as may be," 640 acres. The sections in each township are numbered, as in the margin, from 1 to 36 , beginning at the northeast angle of the township, and going west from 1 to 6, then east from 7 to 12 , and so on alternately to section 36 , which will be in the southeast angle of the township. The sections are subdivided into quar-ter-sections, half a mile square, and containing 160 acres, and sometimes into half-quarter-sections of 80 acres, and quarter-quartersections of 40 acres.

By this beautiful system, the smallest subdivision of land can be at once designated ; such as the northeast quarter of section 31, in township two south, in range two east of Willamette meridian.
456. Difficulty. "The law requires that the lines of the public surreys shall be governed by the true meridian, and that the townships shall be six miles square-two things involving in connection a mathematical impossibility-for, strictly to conform to the meridian, necessarily throws the township out of square, by reason of the convergency of meridians; hence, adhering to the true meridian renders it necessary to depart from the strict requirements of law as respects the precise area of townships, and the subdirisional parts thereof, the township assuming something of a trapezoidal form, which inequality develops itself, more and more as such, the higher the latitude of the surreys. In riew of these circumstances, the law provides that the sections of a mile square shall contain

[^55]the quantity of 640 acres, as nearly as may be; and, moreover, provides that, 'in all cases where the exterior lines of the townships, thus to be subdivided into sections or half-sections, shall exceed, or shall not exceed, six miles, the excess or deficiency shall be specially noted, and added to or deducted from the western or northern ranges of sections or half-sections in such township, according as the error may be in running the lines from east to west, or from south to north.'".
" "In order to throw the excesses or deficiencies, as the case may be, on the north and on the west sides of a township, according to law, it is necessary to survey the section-lines from south to north on a true meridian, leaving the result in the northern line of the township to be governed by the convexity of the earth and the convergency of meridians."

Thus, suppose the land to be surveyed lies between $46^{\circ}$ and $47^{\circ}$ of north latitude. The length of a degree of longitude in latitude $46^{\circ} \mathrm{N}$. is taken as 48.0705 statute miles, and in latitude $47^{\circ} \mathrm{N}$. as $47 \cdot 1944$. The difference, or convergency per square degree $=$ $0.8761=70.08$ chains. The convergency per range ( 8 per degree of longitude) equals one eighth of this, or $8 . \% 6$ chains; and per township ( $11 \frac{1}{2}$ per degree of latitude) equals the above divided by $11 \frac{1}{2}-\mathrm{i}$. e., $0 . \% 6$ chain. We therefore know that the width of the townships along their northern line is 76 links less than on their southern line. The townships north of the base-line therefore become narrower and narrower than the six-mile width with which they start, by that amount.
"Standard Parallels (usually called correction-lines) are established at stated intervals of 30 miles,* to provide for or counteract the error that otherwise would result from the convergency of meridians ; and, because the public surveys have to be governed by the true meridian, such lines serve also to arrest errors arising from inaccuracies in measurements. Such lines, when lying north of the principal base, themselves constitute a base to the surveys on the north of them."

The convergency or divergency above noticed is taken up on

[^56]these correction-lines, from which the townships start again with their proper widths. On these, therefore, there are found double corners, both for townships and sections, one set being the closing corners of the surveys ending there, and the other set being the standard corners for the surveys starting there.

Auxiliary Meridiats. These are run north and south from the base-line, at intervals of twenty-four miles, or four townships.
457. Running Township-Lines. "The principal meridian, the base-line, and the standard parallels, having been first astronomically run, measured, and marked, according to instructions, on true meridians, and true parallels of latitude, the process of running, measuring, and marking the exterior lines of townships will be as follows :

Townships situated North of the base-line and west of the principal meridian.* Commence at Station No. 1, being the southwest corner of T. 1 N.-R. 1 W. , as established on the baseline; thence run north, on a true meridian line, 480 chains, establishing the mile and half-mile corners thereon, as per instructions, to No. 2 (the northwest corner of the same township), whereat establish the corner of Tps. 1 and $2 \mathrm{~N} .-\mathrm{Rs} .1$ and 2 W . ; thence east, on a random or trial line, setting temporary mile and halfmile stakes to No. 3 (the northeast corner of the same tornship), where measure and note the distance at which the line intersects the eastern boundary, north or south of the true or established corner. Run and measure westward, on the true line (taking care to note all the land and water crossings, etc., as per instructions), to No. 4, which is identical with No. 2, establishing the mile and half-mile permanent corners on said line, the last half-mile of which will fall short of being forty chains, by about the amount of the calculated convergency per township, 76 links in the case abore supposed. Should it ever happen, however, that such random line materially falls short, or overruns in length, or intersects the eastern boundary of the township at any considerable distance from the true corner thereon (either of which would indicate an im-

[^57]portant error in the surveying), the lines must be retraced, even if found necessary to remeasure the meridional boundaries of the township (especially the western boundary), so as to discover and correct the error ; in doing which, the true corners must be established and marked, and the false ones destroyed and obliterated, to prevent confusion in future ; and all the facts must be distinctly set forth in the notes. Thence proceed in a similar manner north, from No. 4 to No. 5 (the N. W. corner of T. 2 N.-R. 1 W.), east from No. 5 to No. 6 (the N. E. corner of the same township), west from No. 6 to No. 7 (the same as No. 5), north from No. 7 to No. 8 (the N. W. corner of T. 3 N., R. 1 W.), east from No. 8 to No. 9 (the N. E. corner of the same township, and thence west to No. 10 (the same as No. 8), or the southwest corner T. 4 N.-R. 1 W. Thence north, still on a true meridian line establishing the mile and half-mile corners, until reaching the standard parallel or correction-line (which is here four townships north of the baseline) ; throwing the excess over, or deficiency under, four hundred and eighty chains, on the last half-mile, according to law, and at the intersection establishing the "closing corner," the distance of which from the standard corner must be measured and noted as required by the instructions. But should it ever so happen that some impassable barrier will have prevented or delayed the extension of the standard parallel along and above the field of present survey, then the surveyor will plant, in place, the corner for the township, subject to correction thereafter, should such parallel be extended.

Townships situated north of the base-line, and East of the principal meridian. Commence at No. 1, being the southeast corner of T. 1 N.-R. 1 E., and proceed as with townships situated " north and west," except that the random or trial lines will be run and measured west, and the trus lines east, throwing the excess over or deficiency under four hundred and eighty chains on the west end of the line, as required by law ; wherefore, the surveyor will commence his measurement with the length of the deficient or excessive half-section boundary on the west of the township, and thus the remaining measurements will all be even miles and half-miles.
458. Running Section-Lines. The interior or sectional lines of all townships, however situated in reference to the BASE and MEridian lines, are laid off and surveyed as below :


In the above diagram, the squares and large figures represent sections, and the small figures at their corners are those referred to in the following directions:
"Commence at No. 1 (see small figures on the diagram), the corner established on the township boundary for sections $1,2,35$, and 36 ; thence run north on a true meridian ; at 40 chains setting the half-mile or quarter-section post, and at 80 chains (No. 2) establishing and marking the corner of sections 25, 26, 35, and 36. Thence east, on a random line, to No. 3, setting the temporary quarter-section post at 40 chains, noting the measurement to No. 3 , and the measured distance of the random's intersection north or south of the true or established corner of sections 25, 30, 30, and 31, on the township boundary. Thence correct, west, on the true line to No. 4, setting the quarter-section post on this line exactly
at the equidistant point, now known, between the section corners indicated by the small figures Nos. 3 and 4. Proceed, in like manner, from No. 4 to No. 5, 5 to 6, 6 to ${ }^{7}$, and so on to No. 16, the corner to sections $1,2,11$, and 12 . Thence north on a random line, to No. 17 , setting a temporary quarter-section post at 40 chains, noting the length of the whole line, and the measured distance of the random's intersection east or west of the true corner of sections $1,2,35$, and 36 , established on the township boundary; thence southwardly from the latter, on a true line, noting the course and distance to No. 18, the established corner to sections $1,2,11$, and 12 , taking care to establish the quarter-section corner on the true line, at the distance of 40 chains from said section corner, so as to throw the excess or deficiency on the northern halfmile, according to law. Proceed in like manner through all the intervening tiers of sections to No. ${ }^{7} 3$, the corner to sections 31, 32,5 , and 6 ; thence north, on a true meridian line, to No. 74 , establishing the quarter-section corner at 40 chains, and at 80 chains the corner to sections $29,30,31$, and 32 ; thence east, on a random line to No. \%5, setting a temporary quarter-section post at 40 chains, noting the measurement to No. 75 , and the distance of the random's intersection north or south of the established corner of sections $28,29,32$, and 33 ; thence west from said corner, on the true line, setting the quarter-section post at the equidistant point, to No. ${ }^{7} 6$, which is identical with 74 ; thence west, on a random line, to No. ${ }^{7} \%$, and setting a temporary quartersection post at 40 chains, noting the measurement to No. 77 , and the distance of the random's intersection with the western boundary, north or south of the established corner of sections 25, 36, 30, and 31 ; and from No. $77 \%$, correct, eastward, on the true line, giving its course, but establishing the quarter-section post, on this line so as to retain the distance of 40 chains from the corner of sections $29,30,31$, and 32 ; thereby throwing the excess or deficiency of measurement on the most western half-mile. Proceed north, in a similar manner, from No. "8 to 79 , 79 to 80,80 to 81 , and so on to 96 , the southeast corner of section 6 , where haring established the corner for sections $5,6,7$, and 8 , run thence, successively, on random line east to 95 , north to 97 , and west to 99 ; and
by reverse courses correct on true lines back to said southeast corner of section 6, establishing the quarter-section corners, and noting the courses, distances, etc., as before described.
"In townships contiguous to standard parallels, the abore method will be varied as follows: In every township south of the principal base-line, which closes on a standard parallel, the surveyor will begin at the southeast corner of the township, and measure west on the standard, establishing thereon the mile and half-mile corners, and noting their distances from the pre-established corners. He then will proceed to subdivide, as directed under the above head.
"In the townships norti of the principal base-line, which close on the standard parallel, the sectional lines must be closed on the standard by true meridians, instead of by course-lines, as directed under the above head for townships otherwise situated ; and the connections of the closing corners with the pre-established standard corners are to be ascertained and noted. Such procedure does away with any necessity for running the randoms. But in case he is unable to close the lines on account of the standard not having been run, from some inevitable necessity, as heretofore mentioned, he will plant a temporary stake, or mound, at the end of the sixth mile, thus learing the lines and their connections to be finished, and the permanent corners to be planted, at such time as the standard shall be extended."
459. Exceptional Methods. Departures from the general system of subdividing public lands have been authorized by law in certain cases, particularly on water-fronts.

Thus, an act of Congress, March 3, 1811, authorized the surreyors of Louisiana, "in surveying and diriding such of the public lands in the said Territory, which are or may be authorized to be surveyed and divided, as are adjacent to any river, lake, creek, bayou, or water-course, to lay out the same into tracts, as far as practicable, of fifty-eight poles in front, and four hundred and sixty-five poles in depth, of such shape and bounded by such lines, as the nature of the country will render practicable and most convenient." Another act, of May 24, 1824, authorizes
lands similarly situated "to be surveyed in tracts of two acres in width, fronting on any river, bayou, lake, or water-course, and running back the depth of forty acres; which tracts of land, so surveyed, shall be offered for sale entire, instead of in half-quarter-sections."

The "Instructions" from which we have quoted say: "In those localities where it would best subserve the interests of the people to have fronts on the navigable streams, and to run back into the uplands for quantity and timber, the principles of the act of May 24, 1824, may be adopted, and you are authorized to enlarge the quantity, so as to embrace four acres front by forty in depth, forming tracts of one hundred and sixty acres. But in so doing it is designed only to survey the lines between every four lots (or 640 acres), but to establish the boundary posts, or mounds, in front and in rear, at the distances requisite to secure the quantity of 160 acres to each lot, either rectangularly, when practicable, or at oblique angles, when otherwise. The angle is not important, so that the principle be maintained, as far as practicable, of making the work to square in the rear with the regular sectioning.
"The numbering of all anomalous lots will commence with No. 37 , to avoid the possibility of conflict with the numbering of the regular sections."

The act of September 27,1850 , authorizes the Department, should it deem expedient, to cause the Oregon surveys to be executed according to the principles of what is called the "Geodetic Method."

The complete adoption of this has not been thought to be expedient; but "it was deemed useful to institute on the principal base and meridian lines of the public surveys in Oregon, ordered to be established by the act referred to, a system of triangulations from the recognized legal stations, to all prominent objects within the range of the theodolite; by means of which the relative distances of such objects, in respect to those main lines, and also to each other, might be observed, calculated, and protracted, with the view of contributing to the knowledge of the topography of the country in advance of the progressing
linear surveys, and to obtain the elements for estimating the areas of valleys intervening between the spurs of the mountains."
"Meandering" is a name given to the usual mode of surreying with the compass, particularly as applied to navigable streams. The "Instructions" for this are, in part, as follows:
"Both banks of navigable rivers are to be meandered by taking the courses and distances of their sinuosities, and the same are to be entered in the 'meander field-book.' At those points where either the township or section lines intersect the banks of a navigable stream, posTs, or, where necessary, mounds of earth or stone (as noted in Art. 461), are to be established at the time of running these lines. These are called 'meander corners'; and in meandering you are to commence at one of those corners on the township-line, coursing the banks, and measuring the distance of each course from your commencing corner to the next 'meander corner,' upon the same or another boundary of the same township; carefully noting your intersection with all intermediate meander corners. By the same method you are to meander the opposite bank of the same river.
"The crossing distance between the meander corners, on the same line, is to be ascertained by triangulation, in order that the river may be protracted with entire accuracy. The particulars to be given in the field-notes.
"The courses and distances on meandered narigable streams govern the calculations wherefrom are ascertained the true areas of the tracts of land (sections, quarter-sections, etc.) known to the law as fractional, and bounding on such streams.
"You are also to meander, in manner aforesaid, all lakes and deep ponds of the area of twenty-five acres and upward; also navigable bayous.
"The precise relative position of islands, in a township made fractional by the river in which the same are situated, is to be determined trigonometrically. Sighting to a flag or other fixed object on the island, from a special and carefully measured baseline, connected with the surreyed lines, on or near the river-bank, you are to form connection between the meander corners on the river to points corresponding thereto, in direct line, on the bank
of the island, and there establish the proper meander corners, and calculate the distance across."
460. Marking-Lines. "All lines on which are to be established the legal corner boundaries are to be marked after this method, viz. : Those trees which may intercept your line must have two chops or notches cut on each side of them, without any other marks whatever. These are called 'sight-trees,' or 'linetrees.'
"A sufficient number of other trees standing nearest to your line, on either side of it, are to be blazed on two sides, diagonally or quartering toward the line, in order to render the line conspicuous, and readily to be traced, the blazes to be opposite each other, coinciding in direction with the line where the trees stand very near it, and to approach nearer each other, the farther the line passes from the blazed trees. Due care must ever be taken to have the lines so well marked as to be readily followed."
461. Marking-Corners. "After a true coursing, and most exact measurements, the corner boundary is the consummation of the work, for which all the previous pains and expenditure have been incurred. A boundary corner, in a timbered country, is to be a tree, if one be found at the precise spot; and if not, a post is to be planted thereat; and the position of the corner post is to be indicated by trees adjacent (called bearing-trees), the angular bearings and distances of which from the corner are facts to be ascertained and registered in your field-book.
"In a region where stone abounds, the corner boundary will be a small monument of stones alongside of a single marked stone, for a township corner-and a single stone for all other corners.
"In a region where timber is not near, nor stone, the corner will be a mound of earth, of prescribed size, varying to suit the case.
"Corners are to be fixed, for township boundaries, at intervals of every six miles; for section boundaries, at intervals of every
mile, or 80 chains ; and, for quarter-section boundaries, at intervals of every half-mile, or 40 chains.
"Meander Corner Posts are to be planted at all those points where the township or section lines intersect the banks of such rivers, lakes, or islands, as are by law directed to be meandered," as explained in Art. 459.
" When posts are used, their length and size must be proportioned to the importance of the corner, whether township, section, or quarter-section, the first being at least twenty-four inches above-ground, and three inches square.
"Where a township post is a corner common to four townships, it is to be set in the earth diagonally, thus: $\mathrm{T}{\underset{\mathrm{S}}{\mathrm{D}}}_{\mathrm{N}} \mathrm{E}$, and the cardinal points of the compass are to be indicated thereon by a crossline, or wedge (one eighth of an inch deep at least), cut or sawed out of its top, as in the figure. On each surface of the post is to be marked the number of the particular township, and its range, which it faces. Thus, if the post be a common boundary to four townships, say one and two, south of the base-line, of range one, west of the meridian ; also to tornships one and two, south of the base-line, of range two, west of the meridian-it is to be marked thus:

The position of the post, which is

"These marks are to be distinctly and neatly chiseled into the wood, at least the eighth of an inch deep ; and to be also marked with red chalk. The number of the sections which they respectively face will also be marked on the township post.
"Section or mile posts, being corners of sections, when they are common to four sections, are to be set diagonally in the earth (in the manner provided for township corner posts), and with a similar cross cut in the top, to indicate the cardinal points of the compass; and on each side of the squared surfaces is to be marked the appropriate number of the particular one of the four sections, respectively, which such side faces; also on one side thereof are to be marked the numbers of its township and range; and, to make such marks yet more conspicuous (in manner aforesaid), a streak of red chalk is to be applied.
"In the case of an isolated township, subdivided into thirtysix sections, there are twenty-five interior sections, the southwest corner boundary of each of which will be common to four sections. On all the extreme sides of an isolated township, the outer tiers of sections have corners common only to two sections then surveyed. The posts, however, must be planted precisely like the former, but presenting two vacant surfaces to receive the appropriate marks when the adjacent survey may be made.
"A quarter-section or half-mile post is to have no other mark on it than $\frac{1}{4} \mathrm{~S}$., to indicate what it stands for.
"Township corner posts are to be NOTCHED with six notches on each of the four angles of the squared part set to the cardinal points.
"All mile-posts on township lines must have as many notches on them, on two opposite angles thereof, as they are miles distant from the township corners, respectively. Each of the posts at the corners of sections in the interior of a township must indicate, by a number of notches on each of its four corners directed to the cardinal points, the corresponding number of miles that it stands from the outlines of the township. The four sides of the post will indicate the number of the section they respectively face. Should a tree be found at the place of any corner, it will be marked and notched, as aforesaid, and answer for the corner in lieu of a post; the kind of tree and its diameter being given in the field-notes.
"The position of all corner posts, or corner trees of whatever description, which may be established, is to be perpetuated
in the following manner, viz. : From such post or tree the courses shall be taken, and the distances measured, to two or more adjacent trees, in opposite directions, as nearly as may be, which are called 'bearing-trees,' and are to be blazed near the ground, with a large blaze facing the post, and having one notch in it, neatly and plainly made with an axe, square across, and a little below the middle of the blaze. . The kind of tree and the diameter of each are facts to be distinctly set forth in the field-book.
"On each bearing-tree the letters B. T. must be distinctly cut into the wood, in the blaze, a little above the notch, or on the bark, with the number of the range, township, and section.
"At all township corners, and at all section corners, on range or township lines, four bearing-trees are to be marked in this manner, one in each of the adjoining sections.
"At interior section corners four trees, one to stand within each of the four sections to which such corner is common, are to be marked in the manner aforesaid, if such be found.
"From quarter-section and meander corners two bearing-trees are to be marked, one within each of the adjoining sections.
"Stones at township corners (a small monument of stones being alongside thereof) must have six notches cut with a pick or chisel on each edge or side toward the cardinal points; and where used as section corners on the range and township lines, or as section corners in the interior of a township, they will also be notched by a pick or chisel, to correspond with the directions given for notching posts similarly situated.
"'Stones, when used as quarter-section corners, will have $\frac{1}{4}$ cut on them ; on the west side on north and south lines, and on the north side on east and west lines.
"Whenever bearing-trees are not found, Mounds of earth, or stone, are to be raised around posts on which the corners are to be marked in the manner aforesaid. Wherever a mound of earth is adopted, the same will present a conical shape; but at its base, on the earth's surface, a quadrangular trench will be dug; a spade-deep of earth being thrown up from the four sides of the line, outside the trench, so as to form a continuous eleration along
its outer edge. In mounds of earth, common to four townships or to four sections, they will present the angles of the quadrangular trench (diagonally) toward the cardinal points. In mounds common only to two townships or two sections, the sides of the quadrangular trench will face the cardinal points.
" Prior to piling up the earth to construct a mound, in a cavity formed at the corner boundary point is to be deposited a stone, or a portion of charcoal, or a charred stake is to be driven twelve inches down into such center point, to be a witness for the future.
"The surveyor is further specially enjoined to plant, midway between each pit and the trench, seeds of some tree, those of fruittrees adapted to the climate being always to be preferred.
"Double corners are to be found nowhere except on the standard parallels or correction-lines, whereon are to appear both the corners which mark the intersection of the lines which close thereon, and those from which the surveys start in the opposite direction.
"The corners which are established on the standard parallel, at the time of running it, are to be known as 'Standard Corners,' and, in addition to all the ordinary marks (as herein prescribed), they will be marked with the letters S. C. The 'closing corners' will be marked C. C."
462. Field-Books. There should be several distinct and separate field-books, viz. :
" 1 . Field-notes of the meridian and base lines, showing the establishment of the township, section, or mile, and quarter-section or half-mile, boundary corners thereon ; with the crossings of streams, ravines, hills, and mountains ; character of soil, timber, minerals, etc. These notes will be arranged, in series, by mile-stations, from number one to number -.
"2. Field-notes of the 'standard parallels, or correctionlines,' showing the establishment of the township, section, and quarter-section corners, besides exhibiting the topography of the country on line, as required on the base and meridian lines.

[^58]the establishment of the corners on line, and the topography, as aforesaid.
"4. Field-notes of the subdivisions of townships into sections and quarter-sections; at the close whereof will follow the notes of the meanders of navigable streams. These notes will also show, by ocular observation, the estimated rise and fall of the land on the line. A description of the timber, undergrowth, surface, soil, and minerals, upon each section-line, is to follow the notes thereof, and not to be mixed up with them."
5. The "Geodetic Field-Book," comprising all triangulations, angles of elevation and depression, leveling, etc.

The examples on the next two pages, taken from the " Instructions" which we have followed throughout, will show what is required.

The ascents and descents are recorded in the right-hand columns.

For full details of public-land surveying, see "System of Rectangular Surveying," by J. H. Hawes.
"Instructions" are issued from the General Land-Office from time to time, giving any changes in methods of work, or of mark-ing-points.

## FIELD-NOTES OF

THE EXTERIOR LINES<br>OF AN ISOLATED TOWNSIIIP.

Field-notes of the Survey of Township 25 north, of Range 2 west, of ihe Willamette meridian, in the Territory of Oregon, by Robert Acres, Deputy-Surveyor, under his contract No. 1, bearing date the 2d day of January, 1851.

|  | Chs. lks. <br> East. |  | Feet. |
| :---: | :---: | :---: | :---: |
|  |  | Township lines commenced January 20, 1851. Southern boundary variation $18^{\circ} 41^{\prime} \mathrm{E}$. |  |
|  |  | On a random line on the south boundaries of sections $31,32,33$, 34,35 , and 36 . Set temporary mile and half-mile posts, and intersected the eastern boundary 2 chains 20 links north of the true corner 5 miles 74 chains 53 links. <br> Therefore the correction will be 5 chains 47 links W., $37 \cdot 1$ links S. per mile. |  |
|  | West. $40 \cdot 00$ <br> 62:50 | True southern boundary variation $18^{\circ} 41^{\prime} \mathrm{E}$. | a 10 |
| 0.0.00.00.0 |  | On the southern boundary of sec. 36, Jan. 24, 1851. Set qr. sec. post from which |  |
|  |  | a beech 24 in . dia. bears N. 11 E. 3 S links dist. a do. 9 do. do. S. 9 E. 17 do. a brook 8 l. wide, course N. W. $\qquad$ |  |
| OU \# E E | $80 \cdot 00$ | Set post cor. of secs. $35 \& 36,1 \& 2$, from which a beech 9 in . dia. bears S. 46 E .8 l . dist. a do. 8 do. do. S. 62 W .7 do. a w. oak 10 do. do. N. 19 W. 14 do. a b. oak 14 do. do. N. 29 E. 16 do. <br> Land level, part wet and swampy; timber, beech, oak, ash, hickory, etc. | a 5 |
| Deficient timbered corners. | West. $40 \cdot 00$ | On the S. boundary of sec. $35-$ Set qr. sec. post, with trench, from which a beech 6 in. dia. bears N. 80 E. 8 l. dist. planted S. W. a yellow-locust secd. | a 10 |
|  | $\begin{aligned} & 65 \cdot 00 \\ & 80 \cdot 00 \end{aligned}$ | To beginning of hill.. | $\begin{array}{rr} \alpha & 5 \\ a & 20 \end{array}$ |
|  |  | Set post, with trench, cor. of secs. $34: \& 35,2 \& 3$, from which a beech 10 in . dia. bears S. 51 E .13 l . dist. <br> a do. 10 do. do. N. 56 W .9 do. <br> Planted S. W. a white-oak acorn, <br> N. E. a beechnut. <br> Land level, rich, and good for farming; timber same. |  |
|  | West. $40 \cdot 00$ | On the S. boundary of sec. 34Set qr. sec. post, with trench, from which a black oak 10 in. dia. bears N. 2 E. 635 l. dist. Planted S. W. a beechnut. | $a \quad 5$ |
|  | $80 \cdot 00$ | To corner of sections $33,34,3$ and 4 , drove charred stakes; raised mound, with trench, as per instructions, and Planted N. E. a white-oak acorn; N. W. a yellow-locust seed; <br> S. E. a butternut ; S. W. a beechnut. <br> Land level, rich, and good for farming ; some scattering oak and walnut. | a 10 |
|  |  | Etc., etc., etc. |  |

# FiELD-NOTES OF THE <br> SUBDIVISIONAL OR SECTIONAL LINES, and meanders. <br> Township 25 N., Range 2 W., Willamette Mer. 



## Meanders of Chicheeles Rifer.

Beginning at a meander post in the northern township boundary, and thence on the
left bank down-stream. Commenced February 11, 1851.

| Courses. | Distances. Chs. lks. | Remares. |
| :---: | :---: | :---: |
| S. 76 W . | 18.46 | In section 4 bearing to corner sec. 4 on right bank N. $70^{\circ} \mathrm{W}$. |
| S. 61 W . | $10 \cdot 00$ | Bearing to cor. sec. 4 and 5, right bank N. $52^{\circ} \mathrm{W}$. |
| S. 61 W . | $8 \cdot 18$ | To post in line between sections 4 and 5 , breadth of river by triangulation 9 chains 51 links. |
| S. 54 W. | 10.69 | In section 5. |
| S. 40 W . | 5•59 |  |
| S. 50 W . | 8.46 |  |
| S. 37 W . | 16.50 | To upper corner of John Smith's claim, course E. |
| S. 44 W . | 21.96 |  |
| S. 36 W . | 27.53 | To post in line betreen sections 5 and $\delta$, breadth of river by triangulation $S$ chains 78 links. |
|  |  | Etc., ete. etc. |

## THE SOLAR COMPASS.

463. Nearly all of the lines required in the public-land surveys are meridians and parallels of latitude. Meridians may be located by the methods given in. Chapter III, but the easiest method is with the Solar Compass.

There are several varieties of this instrument, all of which are constructed on the same principle, and are modifications of the instrument invented by William A. Burt, and patented by him in 1836.

Before describing the solar compass, it will be necessary to define the terms to be used.
464. Definitions. The axis of the exrth is the imaginary line about which it revolves. The points in which the axis meets the surface of the earth are called the poles of the earth.

Meridians are great circles of the earth's surface, passing through the poles. The equator is a great circle of the earth's surface, $90^{\circ}$ from the poles. Parallels of latitude are small circles of the earth's surface parallel to the equator. Latitude is the distance north or south from the equator, and is measured on a meridian circle. Longitude is distance east or west from some established meridian. The meridian of Greenwich, England, is usually taken as the prime meridian, from which longitude is reckoned.

Astronomical Terms. Conceive all of the heavenly bodies projected upon the concave surface of a sphere, of which the earth is the center, and whose radius is infinitely great when compared with that of the earth. This is called the Celestial Sphere.

If the axis of the earth be prolonged, the points in which it meets the celestial sphere are called the north and sonth poles of the heavens, and the line joining them is called the axis of the celestial sphere. The apparent revolution of the heavenly bodies about the axis of the celestial sphere is due to the rotation of the earth on its axis once in twenty-four hours.

A plane passed tangent to the earth at the feet of an observer is the sensible horizon; and a plane passed, parallel to this, through the center of the earth, is the rational horizon. Since the radius of the earth is infinitely small in comparison with that of the celestial sphere, if the planes of the rational horizon and sensible horizon be extended in every direction indefinitely, they will meet the celestial sphere in one great circle, called the celestial horizon. If the plane of the earth's equator be extended indefinitely, it will meet the celestial sphere in a great circle, called the celestial equator, or equinoctial.

If through any place a line be passed, perpendicular to the plane of the horizon, the point in which it meets the celestial sphere above the observer is called the zenith; and the point in which it meets the celestial sphere below the observer, the nadir.

Great circles passing through the zerith and nadir are vertical circles.
The zenith distance of a heavenly body is its angular distance from the zenith, and is measured on a vertical circle. The altitude of a body is its angular distance above the celestial horizon, and is measured on a vertical circle. Altitude and zenith distance are complements of each other.

Great circles passing through the poles of the celestial sphere are called circles of declination, or hour-circles. The declination of a heavenly body is its angular distance north or south from the equinoctial, and is measured on a circle of declination.

The celestial meridian of any place is a great circle passing through the zenith, and through the poles of the celestial sphere. The line in which the plane of the celestial meridian meets the plane of the horizon is the terrestrial meridian, or true north and south line.

The hour-angle of a heavenly body is the angle at the pole between the meridian and the declination circle passing through the body.

The parallactic angle is the angle at the body between the declination circle and vertical passing through the body.

The azimuth of a heavenly body is the angle between the celestial meridian and a vertical circle passing through the body, and is measured on the celestial horizon.

If an observer be at the equator, the celestial horizon will pass through the pules of the heavens, and the celestial equator through the zenith. For each degree which the observer travels northward on the earth, the north pole of the heavens will appear to rise one degree above the horizon, and the celestial equator will appear to more one degree southward from the zenith. The latitude of a place, then, is equal to the altitude of the elevated pole, or to the declination of the zenith. In the nortbern hemisphere the north pole of the heavens is the elevated pole.

The earth revolves around the sun in an elliptical orbit once in a year. This gives the sun an apparent motion around the earth. The path of the earth, or the apparent path of the sun in the hearens, is called the ecliptic. It is a great circle on the celestial sphere, making an angle with the celestial equator of about $23^{\circ} 27^{\prime}$. The two points in which the ecliptic meets the equinoctial are called the equinoxes. The sun is on the equinoctial the 21 st of March. This is the vernal equinox. It then moves north of the equator, increasing constantly in northern declination, until the 21st of June, when its declination is about $23^{\circ} 27^{\prime}$ north. This is the northern summer solstice. It then decreases in declination until September 21st, when it is again on the equinoctial. This is the autumnal equinox. It then moves south of the equator, increasing in southern declination until December 21st, when its declination is about $23^{\circ} 27^{\prime}$ south. This is the northern winter solstice. It then decreases in declination until March 21st, when it again arrives at the vernal equinox. The declination of the sun is given in the "Nautical Almanac " for every day in the year.

The transit of a heavenly body is its passage across the celestial meridian.

A sidereal day is the interval of time between two successive transits of
the vernal equinox. A solar day is the interval of time between two successive transits of the sun. The apparent motion of the sun is not uniform, and hence use is made of a fictitious, or mean sun, moving on the equinoctial with a uniform motion, and keeping mean solar time. This is the time kept by clocks and watches. The time indicated by the true sun is called apparent solar time. This is the time given by sun-dials. The difference between apparent solar time and mean solar time is called the equation of time. The equation of time is zero four times in a year, and its maximum value is about sixteen minutes. It is given in the "Nautical Almanac" for every day in the year.

A ray of light, passing from a rarer to a denser medium, is bent, or refracted, toward a perpendicular to the surface of the second medium at the point where the ray enters. The atmosphere surrounding the earth varies in density, being denser as we approach the surface of the earth. The light coming from a heavenly body, and passing through the atmosphere, will be constantly bent toward a perpendicular to the surface of the earth, and its path will be a curve, and not a straight line. The apparent direction of a heavenly body will be tangent to this curve where it meets the eye of the observer. The difference between the apparent and the true positions of a heavenly body is called refraction. It is zero at the zenith, and about $33^{\prime}$ at the horizon ; $45^{\circ}$ from the zenith it is about $57^{\prime \prime}$.

Refractionincreases the altitude of a heavenly body and decreases the zenith distance.

In Fig. 339, N S represents the axis of the celestial sphere, $N$ the north pole, and $S$ the south pole. ED Q is the equinoctial,

Fig. 339.
 H A O the horizon, and HZOX the meridian. ZAX is a vertical circle, N D S a declination-circle. C (the position of the earth) is the center of the celestial sphere. Z is the zenith and X the nadir. Let P be any point on the celestial sphere. A P is its altitude, P Z its zenith distance, and PD its declination; Z N P its hourangle, Z P N its parallactic angle, and N Z P its azimuth.
465. The solar compass differs from the ordinary compass, Fig. 135, in having a solar apparatus, instead of a măgnetic needle, for determining the meridian.

In the figure, $a$ is the latitude-arc, whose center of motion is in two pivots, one of which is shown at $d$. It is furnished with a clamp, slow-motion screw, $f$, and vernier, $e$.

The declination-arc is shown at $b$. The movable arm, $h$, has its center of motion in a pirot at $g$, and is furnished with a clamp, vernier, $v$, and a slow-motion screw, $k$.


The plane of the hour-arc, $c$, is at right angles to the latitudearc, and its center is in the polar axis $p$.

The declination-arc and latitude-arc are read to minutes by the verniers. The hour-are is graduated to half-degrees, and is figured both for hours and degrees.

Attached to each end of the arm $h$ is a rectangular block of brass, in which is set a convex lens, whose focus is on a silver plate attached to the face of the opposite block. The silver plate is marked by two sets of parallel lines, at right angles to each other, as shown in Fig. $341 ; b b$ are called the hour-lines, and $c c$ the equatorial lines. The distance between the hour-lines and between the equatorial lines is equal to the diameter of the image of the sun, formed by the lens in the opposite block.

Fig. 341.


The needle-box $n$ contains a magnetic needle, and is furnished with an are of about $36^{\circ}$ in extent, graduated to half-degrees. The needle-box can be moved about its center by the slow-motion screw $t$.

The sight and levels are similar to those of the ordinary compass.
The equatorial sights, $u$ and $n$, attached to the upper side of the rectangular lens-blocks, are used in the adjustments.

The adjuster, also used in adjusting the instrument, is kept in the instrument-box, and is not shown in the figure.

The compass-sights are attached to the lower plate, and the solar apparatus, levels, and needle-box to the upper plate. The horizontal limb is read to single minutes by the vernier.

Suppose the instrument to be set up and leveled, with the lati-tude-are toward the south. If, now, the latitude-are be set to the latitude of the place of observation (that is, so that the plane of the hour-arc makes an angle with the vertical equal to the latitude of the place), the plane of the hour-are will then be in the plane of the celestial equator, and the polar axis will be parallel to the axis of the earth, and will point toward the north pole of the heavens. If the sun be on the celestial equator, the declination-arm, $h$, may be set at zero on the declination-are, and it will then lie in the plane in which the sun appears to move. If the declination-are be turned so as to point toward the sun, the lens in the block toward the sun will form an image on the silver plate attached to the opposite block. By means of the polar axis, $p$, the declinationarm may be turned so as to follow the sun all day.

When the sun is not at the equinoxes, set off its declination on the declination-arc, and the declination-arm, when turned about on
the axis, $p$, will still turn in the plane in which the sun appears to move. When the sun is in south declination, turn the declinationare away from the sun ; and when the sun is in north declination, turn the declination-arc toward the sun.

When the instrument is in perfect adjustment, and is properly set up and leveled, the image of the sun can not be brought between the equatorial lines, unless the sights are in the plane of the meridian.

## Adjustments.

466. The adjustments will be given in the order in which they should be made. In describing each adjustment, it will be supposed that the instrument has been properly set up and leveled, and the latitude-are turned toward the south.
467. First Adjustment. To cause the level-bubbles to remain in the center of the tubes when the instrument is turned around on its verticul axis. The verification and rectification are the same as those given for the common compass.
468. Second Adjustment. To adjust the equatorial lines and solar lenses. Detach the declination-arm, $h$, by removing the necessary screws, and attach in its place the adjuster, replacing the screws of the pivot, and also of the clamp.

Place the arm $h$ on the adjuster, with the same side against the declination-arc as before it was detached. Then, by means of the vertical axis of the instrument, the declination and latitude ares, and the leveling-screws, turn the arm in the direction of the sun, and bring the image of the sun between the equatorial lines. Then turn the arm half over, bringing the opposite faces of the blocks in contact with the adjuster.

If the sun's image remains between the equatorial lines, the silver plate is in its proper position. If not, loosen the screws which hold the plate, and more the plate so as to correct half of the ap-. parent error. Verify the work by repeating the above operation, until the image remains between the lines in both positions of the arm.

To adjust the other plate, turn the arm end for end on the adjuster, and then proceed as for the first plate.

When both plates have been properly adjusted, remove the adjuster, and replace the declination-arm and its attachments.
469. Third Adjustment. To adjust the vernier of the declina-tion-arc. Set the vernier of the declination-arc at zero. Turn the declination-arm $h$ so as to point toward the sun. Bring the sun's image between the equatorial lines, by means of the slow-motion screw of the latitude-are and the parallel plate-screws, as in the second adjustment. Then revolve the arm so as to bring the opposite solar lens toward the sun. If the sun's image now comes between the equatorial lines, no adjustment is necessary. If not, correct half of the apparent error by means of the slow-motion screw $k$. Verify the work by repeating the above operation until the image comes between the lines in both positions of the arm. The zero of the vernier will now not coincide with the zero of the arc. Make it do so by loosening the screws which hold the vernier, and moving the vernier.
470. Fourth Adjustment. To adjust the Solar Apparatus to the Compass-Sights. Set the vernier of the horizontal limb at zero. Raise the latitude-arc until the polar axis is horizontal, and set the vernier of the declination-are at zero. Direct the equatorial sights at some distant point. If the same point is seen through the sights, no adjustment is necessary. If not, the sights must be changed, or some equivalent adjustment made, which can only be done by an instrument-maker.

## Field-Work.

471. Before the instrument can be used in the field, it is necessary to determine what angles are to be set off on the declinationare and on the latitude-arc.

On the declination-arc, both the declination of the sun and the correction for refraction must be provided for.
472. Declination. The declination of the sun at noon at Greenwich, England, is given in the "Nautical Almanac" for every day in the year, together with the hourly change in declination.

To determine the declination at any place for any time, a correction will need to be applied for difference of declination due to
the difference of time corresponding to difference of longitude, and also for change of declination for different hours of the day.

For example, suppose we wish to find the declination of the sun at Schenectady, New York, for the different hours of the day on May 1, 1885. The longitude of Schenectady is $73^{\circ} 55^{\prime} 50^{\prime \prime}$ west. This in time is 4 h .55 m .43 sec ., or approximately (and near enough for this purpose) 5 hours. From the "Nautical Almanac " we find that the declination of the sun at Greenwich, noon on May 1st, to be $15^{\circ} 12^{\prime} 37 \cdot 5^{\prime \prime}$ north, and the hourly difference is $45^{\prime \prime}$.

When it is noon at Greenwich, it is 7 o'clock in the morning at Schenectady, and at that time the declination of the sun is $15^{\circ}$ $12^{\prime} 37^{\prime \prime}$.

For the successive hours of the day we have only to add the hourly difference in declination, $55^{\prime \prime \prime}$ (the sun at that time having a motion northward from the equator).
473. Refraction. Tables of refraction have been calculated, giving the amount of refraction for different altitudes from the horizon. These tables, however, give the refraction in a vertical plane, and are not directly applicable for use as a correction in declination. It is evident that, in revolving the declination-are around the polar axis, the declination-arc will not lie in the plane of a vertical circle, except when it is placed in the plane of the meridian. The correction for refraction, to be set off on the declina-tion-are, will not, therefore, be equal to the refraction given in the tables except at noon.

The proper correction for refraction to be set off on the decli-nation-are varies with the latitude, declination of the sum, and hour-angle of the sun.

From Chauvenet's "Astronomy," Art. 120, we have:
Refraction in declination $=k^{\prime} \cdot \tan . z \cdot \cos . q$.
The value of $k$ ' may be taken from Table II, Chaurenet's "Astronomy." Its mean value is about $57^{7 \prime \prime}$, and this may be employed when very precise results are not required.
$z$ is the zenith distance, and $q$ the parallactic angle.
From Art. 15, Chaurenet's "Astronomy." we have : $\tan . z \cdot \cos . q=\cot .(\delta+\mathrm{N})$,
in which $\delta=$ declination of the sun, and N is an auxiliary quantity. Tan. N equals cot. $\phi$. cos. $t$, in which $\phi$ is the latitude of the place, and $t$ the hour-angle of the sun.

The tables of Refraction in Declination* are calculated by the above formulas.

In the tables the hour-angle denotes the distance of the sun from the meridian in hours. Thus, at 7 o'clock A. m. the value of the hour-angle is five hours. The north declinations are indicated by + and the south declinations by - .

When the sun is in north declination, the refraction in declination given by the tables is additive. When the sun is in south declination, it is subtractive.

No tables of refraction can be relied upon for altitudes of less than five degrees.

To use the tables, suppose the declination, corrected for refraction, be required for each hour of the day, May 1, 1885, at Schenectady, New York.

By Art. $4 \% 2$ we found that the declination at $\%$ o'clock in the morning was $15^{\circ} 12^{\prime} 37^{\prime \prime}$. The latitude of Schenectady is $42^{\circ} 49^{\prime}$. (Take tabular values for $42^{\circ} 30^{\prime}$.)

In the tables we find that the refraction in declination for latitude $42^{\circ} 30^{\prime}$, when the sun's declination is $15^{\circ}$, and hour-angle 5 hours, is $1^{\prime} 36^{\prime \prime}$. Adding this to $15^{\circ} 12^{\prime} 37^{\prime \prime \prime}$, we have $15^{\circ} 14^{\prime}$ to be set off on the declination-arc.
474. To determine the Latitude, Set off on the declination-are the declination of the sun at noon on the given day (corrected for refraction).

A few minutes before noon, set up and level the instrument, set the declination-are at 12 o'clock on the hour-are, and turn the instrument horizontally until the declination-arm is directed toward the sun. Move the latitude-are vertically so as to bring the sun's image between the equatorial lines. As the sun moves toward the meridian, turn the instrument horizontally so as to keep the image between the hour-lines, and move the latitude-are so as to keep the

[^59]image between the equatorial lines. So long as the sun is ascending, the image will move downward on the plate. When the sun has passed the meridian, and begins to descend, the image will move upward. When the image begins to move upward, the reading on the latitude-are will give the latitude of the place.
475. To determine the "Meridian," or true North and South Line. Set off on the latitude-are the latitude of the place, and on the declination-are the declination of the sun at the time, corrected for refraction. Level the instrument, clamp the horizontal plates at zero, turn the latitude-arc approximately south, and direct the declination-arm toward the sun. Then with one hand turn the instrument horizontally, and with the other revolve the declinationarm on the polar axis, until the image of the sun is brought between the equatorial lines. The sights will then point north and south.
476. Running Lines. The meridian being given by the solar compass, it can be used for determining the bearing of lines in the same way as an ordinary compass, but with greater precision, as the meridian is more accurately determined, and the angles are read by the vernier to single minutes.
477. Use of the Magnetic Needle. Since the solar compass gives the true meridian, and the magnetic needle the " magnetic meridian," the declination of the magnetic needle can be read off directly from the magnetic needle. If the needle be kept at zero of the compass-box arc, by turning the box with its tangent-screw, the declination of the needle can be read to minutes on the are which shows the movement of the compass-box.

By constantly noting the declination of the needle, or by moring the needle-box so as to keep the needle reading zero, lines may be run by the needle, while the sun is obscured, or at such times as for any reason the solar apparatus is not reliable, as when the sun is near the horizon or the meridian.
478. Solar Attachment.* The solar apparatus may be attached to a transit, as shown in Fig. 342.

[^60]The "polar axis" of the solar apparatus is attached to the horizontal axis of the telescope, and projects upward. The "hourcircle" is the small graduated circle, shown above the telescope.

Fig. 342.


Engineer's Transit, with Solar Attachment.
On the " polar axis" rests the frame, which carries the " declina-tion-arc," and the "arm " with its slow-motion attachments, "solar lenses," and "equatorial lines," as before described.

The vertical circle, or arc, of the transit, is used for a "lati-tude-arc."

## Adjustments.

479. The first, second, and third adjustments are similar to those of the solar compass, already explained.
480. To adjust the Polar Axis. Level the instrument carefully, and then level the telescope by means of the level attached to it. Set the arm of the declination-arc at zero, and bring it parallel to the telescope. Place an adjusting lerel, shown in Fig. 343, on the

rectangular blocks attached to the declination-arm. If the brbble remains in the center, the polar axis needs no adjustment in the plane of the axis of the telescope. If not, bring the bubble to the center by means of the two capstan-head screws under the hourcircle, and in line with the telescope. Then turn the declinationarm on the polar axis until it is parallel to the telescope axis, and at right angles to its former position. If the bubble now remains in the center, no adjustment is necessary. If not, bring the bubble to the center by means of the pair of capstan-head screws under the hour-circle and in line with the telescope axis. Verify, and repeat the above operations until the bubble of the adjusting level will remain in the center while the declination-arm is revolved horizontally on the polar axis.
481. To adjust the Hour-Arc. When the telescope is• in the plane of the meridian, the index of the hour-circle should gire ap: parent solar time-that is, mean solar time $\pm$ the equation of time. If the index does not point to the proper dirision, it can be made to do so by loosening the screws on the top of the hour-circle, and turning it until the correct time is indicated by the index.

482. The method of using the solar apparatus on the transit is so nearly the same as that on the compass, already given, that no separate directions will be necessary.
483. Fig. 344 represents a transit with another form of solar attachment.* It consists essentially of a small telescope and level, the telescope being mounted in standards, in which it can be elevated or depressed. The standard revolves around an axis, called the polar axis, which is fastened to the telescope axis of the transit instrument. The telescope, called the "solar telescope," can thus be moved in altitude and azimuth. It is provided with shadeglasses to subdue the glare of the sun, as well as a prism to observe with greater ease when the declination is far north. Two pointers attached to the telescope to approximately set the instrument are so adjusted that when the shadow of the one is thrown on the other the sun will appear in the field of view.

## Adjustment of the Apparatus.

First. Attach the "polar axis" to the main telescope axis in the center at right angles to the line of collimation. The base of this axis is provided with three adjusting-screws for this purpose ; by means of the level on the solar telescope this condition can be readily and accurately tested.

Second. Point the transit telescope-which instrument we assume to be in adjustment-exactly horizontal, and bisect any distant object. The transit level will then be in the middle of the scale. Point the "solar telescope" also horizontally by observing the same object, and adjust its level to read zero, for which purpose the usual adjusting-screws are provided.

## Directions for ting the Attachment.

First. Take the declination of the sun as given in the "Nautical Almanac " for the given day and hour, and correct it for refraction and hourly change. Incline the transit telescope until this amount is indicated by its rertical arc. If the declination of the

[^61]sun is north, depress it ; if south, elevate it. Without disturbing the position of the transit telescope, bring the solar telescope to a horizontal position by means of its level. The two telescopes will now form an angle which equals the amount of the declination.

Second. Without disturbing the relative positions of the two telescopes, incline them and set the vernier to the latitude of the place.

The vertical axis of the "solar attachment" will then point to the pole, the apparatus being in fact a small equatorial.

By moving the transit and the "solar attachment" around their respective vertical axes, the image of the sun will be brought into the field of the solar telescope, and after actually bisecting it the transit telescope must be in the meridian, and the compassneedle indicates its deviation at that place.

## To locate a Parallel of Latitude.

484. In Fig. 345, let P be the pole of the earth, P A and P B the meridians, and A B the desired parallel.

First Method. If from A a line, A C, be run perpendicular to the meridian A P, it is evident that, owing to the convergence of the meridians, the perpendicular will not coincide with the parallel of latitude through A. In north latitudes, as in the United States, the perpendicular, A C, will run to the south of the parallel, A B.


To find the distance C B, when the latitude of the starting-point A, and the distance A C are known.

In the triangle PAC, right-angled at A:
$\cos . \mathrm{PC}=\cos . \mathrm{A} \mathrm{P} \times \cos . \mathrm{A} \mathrm{C}$.

$$
\begin{equation*}
\mathrm{BC}=\mathrm{PC}-\mathrm{PB} \text {, and } \mathrm{AP}=\mathrm{BP}=\text { co-latitude. } \tag{1.}
\end{equation*}
$$

$\therefore \cos . \mathrm{PC}=\sin$. latitude $\times \cos . \mathrm{AC}$.
A C, being a measured distance on an arc of a great circle, must be reduced to the corresponding angle.
Angle of any arc in minutes $=\frac{\text { length of are } \times 3437 \cdot 7468}{\text { radius }}$.

$$
(3437 \cdot 7468=57 \cdot 29598 \times 60) . \quad \text { Art. } 280
$$

Treating the earth as a sphere, this becomes:
Angle of arc in minutes $=$ length of arc $\frac{343 \% \cdot \% 468}{20912405}$.
Log. arc in minutes $=$ log. length $-3 \cdot \% 941301$
Then use the value obtained by [2] in formula [1].
BC is found as an angle. To reduce it to feet, we have :

$$
\begin{align*}
& \text { Length in feet }=\frac{\text { angle in minutes } \times \text { radius }}{343 \% \cdot \neg 468} \\
& \text { Length in feet }=\frac{\text { angle in seconds } \times \text { radius }}{60 \times 3437 \cdot \tau 468} \tag{3.}
\end{align*}
$$

Log. length in feet $=$ log. angle in seconds $+2 \cdot 0059759$
485. Otherwise. Find the length of an arc subtending one second at the center.

$$
\frac{2 \pi \times 20912405}{360 \times 60 \times 60}=101 \cdot 386 \text { feet }
$$

i. e., $101 \cdot 386$ feet subtends an angle of one second at the center of the earth. Then, angle in seconds $=\frac{\text { distance in feet }}{101: 386}$, and distance $=$ angle in seconds $\times 101.386$
486. Approximately, $B C$ in seconds $=\frac{1}{4} \mathrm{P}^{2}$ (in seconds) $\times \sin .2 \mathrm{PA} \times \sin , 1^{\prime \prime}$.

$$
\text { To find } P . \quad \tan . P=\frac{\tan \cdot A B}{\sin A P}
$$

487. Example. Latitude $45^{\circ}$ north, and distance 6 miles, required the offset BC.

$$
6 \text { miles }=31650 \text { feet }
$$

By [2]

$$
\begin{aligned}
\log \cdot 31680 & =4 \cdot 500: 852 \\
& -3 \cdot 7941301 \\
\text { log. } 5^{\prime} \cdot 089265 & =\cdot \% 066551 \\
5^{\prime} \cdot 089265 & =5^{\prime} 5^{\prime \prime} \cdot 356
\end{aligned}
$$

By [1]

$$
\log \cdot \sin .45^{\circ}=9 \cdot 8494850
$$

$$
\log . \cos .5^{\prime} 5^{\prime \prime} \cdot 356=9 \cdot 9999995
$$

log. cos. $\mathrm{P} C=\log . \cos .45^{\circ} 0^{\prime} 0^{\prime \prime} \cdot 23 i=9 \cdot 8494845$

$$
\therefore \mathrm{BC}=0^{\prime \prime} \cdot 23 \%
$$

T'o reduce to feet by [3], log. $0^{\prime \prime} \cdot 23 \%=\overline{1} \cdot 3747483$
$\log$. B C in feet $=\log .24 \cdot 029$ feet $\frac{+2 \cdot 0059789}{=1 \cdot 3807272}$
Second Method:

$$
\text { Angle }=\frac{31680}{101 \cdot 38 \overline{6}}=312^{\prime \prime}=5^{\prime} 12^{\prime \prime} \cdot 468
$$

Then, as above, we find $\mathrm{BC}=0^{\prime \prime} \cdot 23 \%$ of arc.
B C in feet $=0^{\prime \prime} \cdot 23 \% \times 101 \cdot 386=24 \cdot 0289$ feet.
Approximate Method:
Solving by formula [5], we find $\mathrm{BC}=24 \cdot 3$ feet.
488. Spheroidal Formula. The preceding methods suppose the earth to be a sphere. Treating it as a spheroid, the following formula is without material error for distances within 100 miles:

$$
\mathrm{C} \mathrm{~B}=\frac{1}{2} k^{2} \tan \cdot \mathrm{~L} \frac{\left(1-\left[e^{2} \cdot \sin _{0}{ }^{2} \mathrm{~L}\right]\right)^{\frac{1}{2}}}{a} .
$$

$k=$ distance in feet, $\mathrm{L}=$ latitude of initial point.
$a$ equatorial radius $=20926062$ feet .
$e=\cdot 08169683$.
Example. Latitude $45^{\circ} \mathrm{N}$. Distance 6 miles.

$$
\begin{aligned}
& \text { log. } e^{2}=\overline{3} .8244104 \text {. } \\
& \text { log. } \sin .{ }^{2} 45^{\circ}=\left\{\begin{array}{l}
9 \cdot 8494850 \\
9 \cdot 8494850
\end{array}\right. \\
& \text { log. } \cdot 0033 \% 18=\overline{3} \cdot 5233804 \\
& 1-\cdot 0033 \% 18=\quad \cdot 9966283 \\
& \log \cdot 0 \cdot 9966283=\overline{1} \cdot 9992666=\text { log. numerator. } \\
& \text { log. } \frac{1}{2}=\overline{1} \cdot 6989 \% 00 \\
& \text { log. } k^{2}=\left\{\begin{array}{l}
4 \cdot 500 \% 852 \\
4 \cdot 500 \% 852
\end{array}\right. \\
& \text { log. tan. } 45=10 \text {. } \\
& \text { log. numerator }=\frac{\overline{1} \cdot 9992666}{8 \cdot 6998070} \\
& \text { log. } a=\quad \% \cdot 32068 \% 5 \\
& \log .23 \cdot 939 \text { feet }=\overline{1 \cdot 3791195}
\end{aligned}
$$

489. Length of Parallels. The radius of any parallel of latitude equals the radius at the equator multiplied by the cos. latitude.

Then length in feet of $1^{\circ}=\frac{\pi}{180}$. radius in feet $\times$ cos. latitude.
Then length in feet of $1^{\circ}=\frac{\pi}{180} \times 20912405 \times \cos$. latitude.
log. length in feet of $1^{\circ}=$ log. cos. latitude $+5 \cdot 5622814$.
Example. To find the length of a degree on the $45^{\circ}$ parallel.

$$
\begin{aligned}
\log . \operatorname{cos.} 45 & =9 \cdot 8494855 \\
\text { log. } 25808 \% & =\frac{5 \cdot 5622814}{5 \cdot 411 \% 669}
\end{aligned}
$$

Conversely. The angle, in minutes, subtended by any arc $=$ length of arc $\times 3437 \cdot 7468$
radius $\times$ cos. latitude .
log. angle in minutes $=$ log. arc in feet $-3 \cdot \% 841301-$ cos. latitude.
Example. Latitude $45^{\circ} \mathrm{N}$. and distance 6 miles.

$$
\begin{aligned}
& \text { log. } 31680=4 \cdot 5007852 \\
& \frac{-3 \cdot 7841301}{\cdot 7166551} \\
& \text { co-log. cos. } 45^{\circ}=\frac{1505150}{.8671701} \\
& \text { log. } 7^{\prime} 21^{\prime \prime} \cdot 89 \%
\end{aligned}
$$

490. The difference of lengths of any two parallels is called the convergence of the meridians between those parallels. This may be obtained more easily, since the distances between the meridians are as the cosines of the latitudes.

Example. Two "range-lines" (meridians) are 6 miles (480 chains) apart on the base-line of $46^{\circ}$.

Required their convergence at $4 \%^{\circ}$ north.

$$
\begin{aligned}
& \text { Length at } 4 \%^{\circ}=480 \frac{\cos .47^{\circ}}{\cos .46^{\circ}}=4 \% 1 \cdot 252 . \\
& \qquad 480-4 \% 1 \cdot 252=8 \text { chains } \% 4 \cdot 8 \text { links. }
\end{aligned}
$$

## PART II.

## LEVELING.

## INTRODUCTION.

491. Leveling in General. A level surface is one which is everywhere perpendicular to the direction of gravity, as indicated by a plumb-line, etc., and consequently parallel to the surface of standing water. It is, therefore, spherical (more precisely, spheroidal), but, for a small extent, may be considered as plane. Any line lying in it is a level line.

A vertical line is one which coincides with the direction of gravity.

The height of a point is its distance from a given level surface, measured perpendicularly to that surface, and therefore in a vertical line.

Leveling is the art of determining the difference of the heights of two or more points.

To obtain a level surface or line, usually the latter, is the first thing required in leveling.

When this has been obtained, by any of the methods to be hereafter described, the desired height of a point may be determined directly or indirectly.
492. Direct Leveling. In this method of leveling, a level line is so directed and prolonged, either actually or visually, as to pass exactly over or under the point in question-i. e., so as to be in the same vertical plane with it-and the height (or depth) of the point above (or below) this level line is measured by a vertical rod, or by some similar means. The height of any other point being
determined in the same manner, the difference of the two will be the height of one of the points above the other. So on, for any number of points.

Direct Leveling is the method most commonly employed. It will form Chapter I of this part.
493. Indirect Leveling. In this method of leveling the desired height is obtained by calculation from certain co-ordinate measured lines or angles, which fix the place of the point.

Thus, the horizontal distance from any point to a tree being known, and also the angle with the horizon made by a straight line passing from the point to the top of the tree, its height above the point can be readily calculated. This is the most simple and most usual form of this method, though many others may be employed.

Indirect Leveling will be developed in Chapter II.
494. Barometric Leveling. This determines the difference of the heights of two points by the difference of the weights of the portions of the atmosphere which are abore each of them, as indicated by a barometer. It is explained in Chapter III.

## CHAPTER I.

## DIRECT LEVELIIVG。

GENERAL PRINCIPLES.
495. Leveling Instruments. The instruments employed to obtain a level line may be arranged in three classes, depending on these three principles:

1. That a line perpendicular to a vertical line is a horizontal or level line.
2. That the surface of a liquid in repose is horizontal.
3. That a bubble of air, confined in a ressel otherwise full of a liquid, will rise to the highest point of that liquid.

They will be described in the following pages.
496. Methods of Operation. When a level line has been obtained, by any means, the difference of heights of any two points may be found by either of these two methods:

First Method. Set the leveling instrument over one of the

Fig. 346.

points, as A, in Fig. 346. Measure the height of the level line above the point. Then direct this line to a rod held on the other
point, and note the reading. The difference of the two measurements at A and B will be the difference of their heights.

Second Method. Let A and B, Fig. 34\%, represent the two points. Set the instru-
 ment on any spot from which both the points can be seen, and at such a height that the level line will pass above the highest one. Sight to a rod held at A, and note the reading. Then turn the instrument toward B , and note the height observed on the rod held at that point. The difference of the two readings will be the difference of the heights required. The absolute height of the level line itself is a matter of indifference.
497. Curvature. The level line giren by an instrument is tangent to the surface of the earth. Therefore, the line of true level is always below the line of apparent level. In Fig. 348, A D represents the line of apparent level, and A B the line of true level. D B is the correction for the earth's curvature. By geometry we have: $\mathrm{A}^{2}=\mathrm{DB} \times(\mathrm{DB}+2 \mathrm{~B} 0)$. But $\mathrm{D} B$, being very small, compared with the diameter of the earth, may be dropped from the quantity in the parenthesis, and we have :

$$
\mathrm{DB}=\frac{\mathrm{AD}^{2}}{2 \mathrm{BO}} ;
$$

Fig. 348.

i. e., the correction equals the square of the distance divided by the diameter of the earth.

The difference of height for a distance of

$$
1 \text { mile }=\frac{1}{7916}=\frac{5280 \times 12}{7916}=8 \text { inches } .
$$

This varies as the square of the distance. The effect, if neglected, is to make distant objects appear lower than they really are.

The effect is destroyed by setting the instrument midway between the two points.
498. Refraction. Rays of light coming through the air are curved downward. The effect is, to make objects look higher than they really are. Its amount is about one seventh that of curvature, and it operates in a contrary direction.

## PERPENDICULAR LEVELS.

499. Principle. The principle upon which these are constructed is, that a line perpendicular to the direction of gravity is a level line.
500. Plumb-line Levels. The A level, Fig. 349, is so adjusted that, when the plumb-line coincides with the mark on the cross-piece, the feet of the level shall be at the same height. It is adjusted by reversion thus: Place its feet on any two points. Mark on the cross-bar

Fig. 350.
 the place of the plumb-line. Turn the instrument end for end, resting it on the same points, and mark the new place of the plumb-line. The point midway between the two is the right one.

Another form is shown in Fig. 350. The above forms are not convenient for prolonging a level line. To do this, invert the preceding form, as in Fig. 351.

To test and adjust this, sight to some distant point nearly on a level, and mark where the plumbline comes to on

Fig. 351.
 the bottom of the rod. Turn the instrument around and sight again, and note the place of the plumb-line. The midway point is the right one.

A modification of the last form is to fasten a common carpenter's square in a slit in the top of a staff, by means of a screw, and

Fig. 352.
 then tie a plumb-line at the angle so that it may hang beside one arm. When it has been brought to do so, by turning the square, then the other arm will be level.
501. Reflecting Levels. In these, the perpendicular to the direction of gravity is not an actual line, but an imaginary reflected line.
It depends on the optical principle that a ray of light which meets a reflecting plane at right angles is reflected back in the same line.

When the eye sees itself in a plane mirror, the imaginary line which passes from the eye to its image is perpendicular to the mirror. Therefore, if the mirror be vertical, the line will be horizontal. It may therefore be used like any other line of sight for determining points at the same height as itself.

The first form, Fig. 353 (Colonel Burel's), consists of a rhomb of lead, of about tro inches on a side, and one inch thick.

One side (the shaded part of the figure) is faced with a mirror. The right-hand corner of the rhomb is cut off, as seen in the figure, and a wire, A B, is stretched across the mirror.

To use this, hold up the instrument, with the

Fig. 353.
 mirror opposite the eye, by the string D , so that the eye seems bisected in the mirror by the wire A B. Then glance through the opening at $B$, and any point in the line of the eye and wire will be in the same horizontal plane with them.

The correctness of the instrument may be verified in the following manner : Hold up the instrument before any plane surface, as a wall, and determine the height of some point, as preriously directed. Then, without changing the height of the instrument, turn it, half around, place yourself between it and the wall, and
note the point of the wall which is seen in the mirror to coincide with the image of the eye.

If the two points on the wall coincide, the instrument is correct. If they do not, the mirror does not hang plumb, and the point midway between the two is the true one.

The instrument is rectified, or made to hang plumb, by means of the pear-shaped piece of lead seen attached to the lower corner of the rhomb.

The second form consists of a hollow brass cylinder, with an opening at the upper end, as seen in Fig. 354. At the opening is a small mirror, whose vertical plane makes an angle with the vertical plane of section by which the cylinder was cut in forming the aperture. The edge of the mirror is marked thus ( x ) in the first half of Fig. 354.


Fig. 354.


The mirror is made to hang plumb by means of a one-sided weight within the cylinder.

This is used by setting it on a stake driven into the ground, or by holding it in the hand, making the lower edge of the opening answer the same purpose as the wire in the other case.

The same methods of verifica-

Fig. 355.
 tion and rectification are used as with the first form of the instrument.

The instrument, in its third form, is simply a small steel cylinder, $4^{\prime \prime}$ or $5^{\prime \prime}$ long, and $\frac{1_{2}^{\prime \prime}}{}$ in diameter, highly polished, and suspended from the center of one end by a fine thread.
To use this, hold it up by the thread with one hand, and with the other hand hold a card between the eye and instrument, using
the upper edge of the card, as seen reflected in the mirror, the same as the wire in the first form.

This instrument is the invention of M. Cousinery.

## WATER-LEVELS.

502. Continuous Water-Levels. These may consist of a channel connecting the two points, and filled with water; or of a tube, usually flexible, with the enảs tursed up, and extending from one point to the other.

By measuring up or down, from the surface of the water at each end, the relative heights of the two points may be determined.
503. Visual Water-Levels. The simplest one is a short surface of water prolonged by sights at equal distances above it, as in Fig. 356.

A portable form is a tube bent up at each end, and nearly filled
Fig. 356.

with water. The surface of the water in one end will always be at the same height as that in the other, howerer the position of the tube may vary. It may be
 easily constructed with a tube of tin, lead, copper, etc., by bending up, at right angles, an inch or two of each end, and supporting the tube, if too flexible, on a wooden bar. In these ends, cement (with putty, twine dipped in white-lead, etc.) thin vials, with their bottoms broken off, so as to leave a free communication between them. Fill the tube and the rials, nearly to their top, with colored water. Blue vitriol or cochineal may be used for coloring it. Cork their mouths, and fit the instrument, by a steady but flexible joint, to a tripod.

To use it, set it in the desired spot, place the tube by eye nearly level, remove the corks, and the surfaces of the water in the two vials will come to the same level. Stand about a yard behind the nearest vial, and let one eye, the other being closed, glance along the right-hand side of one vial, and the left-hand side of the other. Raise or lower the head till the two surfaces seem to coincide, and this line of sight, prolonged, will give the level line desired. Sights of equal height, floating on the water, and rising above the tops of the vials, would give a better-defined line.

## AIR-BUBBLE OR SPIRIT LEVELS.

504. The "spirit-level" consists essentially of a curved glass tube nearly filled with alcohol, but with a bubble of air left within, which always seeks the highest spot in the tube, and will therefore, by its movements, indicate any change in the position of the tube. Whenever the bubble, by raising or lowering one end, has been brought to stand between two marks on the tube, or, in case of expansion or contraction, to extend an equal distance on either side of them, the bottom

Fig. 358.
 of the block (if the tube be in one), or sights at each end of the tube, previously properly adjusted, will be on the same level line. It may be placed on a board fixed to the top of a staff or tripod.

When, instead of the sights, a telescope is made parallel to the level, and various contrivances to increase its delicacy and accuracy are added, the instrument becomes the engineer's spirit-level.

The upper surface of the tube is usually the are of a circle, and, when we speak of lines parallel to a " level," we mean parallel to the tangent of this are at its highest point, as indicated by the middle of the bubble.
505. Sensibility. This is estimated by the distance which the bubble moves for any change of inclination. It is directly proportional to the radius of curvature of the tube. To determine the radius, proceed thus :

Let $S=$ length of the are over which the bubble moves for an inclination of 1 second ( $1^{\prime \prime}$ ).

Let $\mathrm{R}=$ its radius of curvature.

$$
\begin{aligned}
& \text { Then } \mathrm{S}: 2 \pi \mathrm{R}:: 1^{\prime \prime}: 360^{\circ}, \\
& \text { whence } \mathrm{R}=206265 \times \mathrm{S}, \\
& \quad \text { or } \mathrm{S}=\frac{\mathrm{R}}{206265} .
\end{aligned}
$$

S may be found by trial, the level being attached to a finely divided vertical circle. The
 radius may also be found without this, thus : Bring the bubble to center, and sight to a divided rod. Raise or lower one end of the level, and again sight to the rod. Call the difference of the readings $h$, the distance of the rod $d$, and the space which the bubble moved S. Then we have two approximately similar triangles ; whence $r=\frac{d \mathrm{~S}}{h}$.

Example. At 100 feet distance, the difference of readings was 0.02 foot, and the bubble moved 0.01 foot. Then the radius was $\frac{100 \times 0.01}{0.02}=50$ feet.

The sensibility of an air-bubble level equals that of a plumbline level having a plumb-line of the same length as the radius of curvature.
506. Block-Level. If this is marked by the maker, and the bubble does not come to the center, when turned end for end, plane or grind off one end of the bottom until it

Fig. 360.
 does.

Otheriwise, if the bubble-tube is capable of morement, raise or lower one end of it until it will verify, bringing the bubble
half-way back to the middle by this means, and the other half by raising or lowering one end of the block, because the reversion has doubled the error.

Repeat this, if necessary.
Circular Level. The upper surface of this is spherical. It will therefore indicate a level in every direction, instead of only one, as does the preceding. It is adjusted like the last one, but in two directions, at right angles to each other.

507. Level with Sights. The line of sight is made parallel to the tangent of the level. It may be tested thus :

Fig. 362.


Bring the bubble to the center of the tube and make a mark, in the line of sight, as far off as can be seen. Then turn the level end for end, and sight again. If the bubble remains in the same place, " all right." If not, rectify it by altering the sights, or by altering the marks for the bubble to come to, bringing the bubble half-way back, and trying it again.
508. Hand-Reflected Level. This consists of a brass tube, about six inches long, and one inch in diameter. To the inside of the

Fig. 363.
 upper portion of the tube is attached a small level. A small mirror is placed at an angle in the lower side of the tube, so that it will reflect the point to which the bubble must come, in order to have the instrument level, to the eye.

A small hole at one end, and a horizontal cross-hair at the other, give the desired level line. It is used by holding it in the hand.

Fig. 363 is an approved form, made by Young, of Philadelphia. The improvement consists in the patent "Locke sight," which enables the near cross-hair to be distinctly seen at the same time as the distant object.
509. Gurley's Telescopic Hand-Level (Fig. 363', a). "This consists of a tube to which are fitted the lenses of a single opera-

glass, and containing in addition thereto a reflecting prism. crosswire, and small spirit-level, the last being shown in the open part of the tube.
"The eye-lens, as indicated in the cut, is made of two separate pieces, the larger one being the usual concare eye-lens of the opera-glass, the smaller one a segment of a plano-convex lens haring its focus in a cross-wire under the level-vial and above the reflecting prism.
"The observer holds the tube horizontal, with the level opening uppermost, and with the same eye sees the object toward which the instrument is directed, and obserres the position of the bubble. When the level is truly horizontal, the cross-wire will
bisect the bubble, and will also determine the level of any object seen through the telescope.
" In the binocular form of this level (Fig. 363', b) the tube on the right incloses the usual lenses of the opera-glass, while that on the left contains only the prism, level-vial, and cross-wire. The binocular hand-level gives a clearer view of an object than is possible with a single tube, there being no light lost by the interference of the prism and level-vial."
510. The Telescope-Level. In this the line of collimation of the telescope corresponds to the sights of Fig. 362, and is made parallel to the level-i. c., this line is so adjusted as to be horizontal when the bubble of its level is in the center.

There are many different forms of the telescope-level, of which the most important ones will now be given.
511. The Y-Level. This is so named from the shape of the supports of the telescope. It is the variety most used by American engineers.

Fig. 364 represents a Y-level of the usual form. The telescope is held in the wyes by the clips, A A, which are fastened to the

wyes by tapering pins, so that the telescope can be clamped in any position. The milled-headed screws at $M$ and $M$ are used to move
the object-glass and eve-piece in and out, so as to adjust them for long and short sights, and for short-sighted and long-sighted per-

Fig. 365.
 sons. L is a spirit-level ; P and P are parallel plates ; C is the clamp-screw, which fastens the spindle on which the level-bar, B, which supports the wyes, turns; $T$ is the tangentscrew, by which the telescope may be slowly turned around horizontally.
512. The Telescope. The arrangement of the parts of the telescope is shown in Fig. 365. 0 is the object-glass, by which an image of any object, toward which the telescope may be directed, is formed within the tube. EE is the eye-piece-a combination of lenses, so arranged as to magnify the small image formed by the object-glass. The cross-hairs are at X . They are moved by means of the screws shown at B B. A A are screws used for centering the eye-piece. CC are screws used for centering the object-glass. At D D are rings, or collars, of exactly the same diameter, turned very truly, by which the telescope revolves in the wres.

The telescope shown in the figure forms the image erect. Other combinations of lenses are used, some of which insert the image ; but the one here shown is generally preferred.
513. The Cross-Hairs. These are made of very fine platinum wire or of spider-threads. They are attached to a short, thick tube, placed within the telescope-tube, through which pass loosely four screws whose threads enter and take hold of the cross-hair ring, as shown in Fig. 366.

In some instruments, one of each pair of opposite screws is replaced by a spring ; and the screws, instead
of being capstan-headed, and moved by an "adjusting-pin," have square heads, and are moved by a "key," like a watch-key.

The line of collimation (or line of aim) is the imaginary line passing through the intersection of the cross-hairs and the optical center of the objectglass.

The image formed by the object-glass should coincide precisely with the cross-hairs. When this is

Fig. 366.
 not the case, there will be an apparent morement of the crosshairs, about the objects sighted to, on moving the eye of the observer. This is called instrumental parallax. To correct it, move the eye-piece out or in, till the cross-hairs are sharply defined against any white object. Then move the object-glass in or out, till the object is also distinctly seen. The image is now formed where the cross-hairs are, and no movement of the eye will cause any apparent motion of the cross-hairs.
514. The Level. This consists of a. thick glass tube, slightly curved upward, and so nearly filled with alcohol that only a small bubble of air remains in the tube. This always rises to the highest part. The brass case, in which this is inclosed, is attached to the under side of the telescope, and is furnished with the means of moving, at one end vertically, and at the other horizontally. Over the aperture, in the case, through which the bubble-vial is seen, is a graduated level-scale, numbered each way from zero at the center.
515. Supports. The wyes in which the telescope rests are supported by the level-bar, B, and fastened to it by two nuts at each end (one above, one below the bar), which may be moved with an adjusting-pin. The use of these nuts will be explained under "Adjustments." Attached to the center of the level-bar is a steel
spindle, made so as to turn smoothly and firmly in a hollow cylinder of bell-metal ; this, again, is fitted to the main socket of the upper parallel plate.
516. Parallel Plates. It is by the aid of these that the instrument is leveled. The plates are united by a ball-and-socket joint,

and are held apart by the four plate-screws, $Q Q Q Q$. which pass through the upper one, and press against the lower one.

To level the instrument, turn the telescope till it is brought over a pair of opposite parallel plate-screws. Then turn the pair of screws, to which the telescope has been made parallel, equally in opposite directions, screwing one in and the other out, till the bubble is brought to the center. Then turn the telescope so as to bring it over the other pair of opposite screws, and bring the bubble to the center, as before.

Repeat the operation, as moving one pair of screws may affect the other.

Sometimes one of each pair of opposite screws is replaced by a strong spring, and in some instruments only three screws are used.

The lower plate is screwed on to the tripodhead.

51\%. Fig. 367 is a twenty-inch Y-level, and Fig. 368 is a longitudinal section of it, showing its construction.


In Fig. 368, B B are
the screws attached to the cross-hair ring. At A are four screws holding a ring through which the inner end of the eye-piece passes. At C are four screws holding a ring, through which the inner
end of the object-glass slide passes. The use of these sets of screws will be explained under "Adjustments."

The interior spindle, $D$, which supports the instrument, and on which it turns, is made of steel, and is carefully fitted to the interior of a hollow socket of bell-metal, which has its exterior surface fitted to the main socket, E , of the tripod-head. The hollow bell-metal socket is held in place by a washer and screw, shown at D.

A screw, passing through the main socket, E, enters a groove in the exterior of the bell-metal socket, and fastens the instrument to the tripod-head.

## ADJUSTMMENTS.

518. The line of collimation of the telescope should be horizontal when the bubble is in the center of the tube ; which will be the case when this line is parallel to the plane of the level. But both this line and this plane are imaginary, and can not be compared together directly. They are therefore compared indirectly. The line of collimation is made parallel to the bottom of the collars, and the plane of the level is then made parallel to them.
519. First Adjustment. To make the line of collimation parallel to the bottoms of the collars.

Sight to some well-defined point, as far off as it can be dis-
Fig. 369.

tinctly seen. Then revolve the telescope half around in its sup-ports-i. e., turn it upside down. If the line of collimation was not in the imaginary axis of the rings, or collars, on which the telescope rests, it will now no longer bisect the object sighted to. Thus, if the horizontal hair was too high, as in Fig. 369, this line of collimation would point at first to A, and, after being turned over, it would point to $B$. The error is doubled by the reversion, and it should point to C, midway between $A$ and B. Make it do
so, by unscrewing the upper capstan-headed screw, and screwing in the lower one, till the horizontal hair is brought half-way back to the point B. Remember that, in an erecting telescope, the crosshairs are reversed, and vice versa. Bring it the rest of the way by means of the parallel plate-screws. Then revolve it in the wyes back to its original position, and see if the intersection of the crosshairs now bisects the point, as it should. If not, again revolve, and repeat the operation till it is perfected. If the vertical hair passes to the right or to the left of the point when the telescope is turned half around, it must be adjusted in the same manner by the other pair of cross-hair screws. One of these adjustments may disturb the other, and they should be repeated alternately. When they are perfected, the intersection of the cross-hairs, when once fixed on a point, will not move from it when the telescope is revolved in its supports. This double operation is called adjusting the line of collimation.

It has now been brought into the center line, or axis, of the collars, and is therefore parallel to their bottoms, or the points on which they rest, if they are of equal diameters. We have to assume this as having been effected by the maker.

In making this adjustment, the level should be clamped, but need not be leveled.
520. Second Adjustment. To make the bottoms of the collars parallel to the plane of the level-i. e., to insure their being horizontal when the bubble is in the center.

Clamp the instrument, and bring the bubble to the center by the parallel plate-screws. Take the telescope out of the wyes, and turn it end for end. If the bubble returns to the center, "all right." If not, rectify it, by bringing the bubble half-way back, by means of the nuts which are above and below one end of the bubble-tube, and which work on a screw. Bring it the rest of the way by the plate-screws, and again turn end for end. Repeat the operation, if necessary.

If, in revolving the telescope (as in the first adjustment), the bubble runs toward either end, it must be adjusted sidewise, by means of two screws which press horizontally against the other end
of the bubble-tube. This part of the adjustment may derange the preceding part, which must, therefore, be tried again.
521. Third Adjustment. To cause the bubble to remain in the center of the tube when the telescope is turned around horizontally.

To verify this, bring the bubble to the center of the tube, and then turn the telescope half-way around horizontally. If the bubble does not remain in the center, adjust it by bringing it half-way back by means of the nuts at the end of the level-bar. Test it by bringing it the rest of the way back by the parallel plate-screws, and again turning half-way around.

The cause of the difficulty is, that the plane of the level is not perpendicular to the axis about which it turns, and that this axis is not vertical. The above operations correct both these faults.

This adjustment is mainly for convenience, and not for accuracy, except in a very small degree.

Some instruments have no means of making the third adjustment. They must be treated thus :

Use the screws at the end of the bubble-tube, to cause the bubble to remain in the center when the level is turned around horizontally. Then make the line of collimation parallel to the level by raising or lowering the cross-hairs.
522. When levels are provided with the means of centering the eye-piece and object-glass, these operations should precede the first three which we have just explained.

Centering the Object-Glass. After adjusting the line of collimation for a distant object (as explained in the "First Adjustment") move out the slide, which carries the object-glass, until a point ten or fifteen feet distant can be distinctly seen. Then turn the telescope half over, as before, and see if the intersection of the crosshairs bisects the point. If not, bring it half-way back by the screws C C, Fig. 365, moving only one pair of screws at a time. Repeat the operation for a distant point, and then again for a near one, if necessary. We have now adjusted the line of collimation for long and short sights, and may assume it to be in adjustment for intermediate ones, since the bearings of the slides are supposed to be true, and their planes parallel to each other.

Centering the Eye-Piece. This is to enable the observer to see the intersection of the cross-hairs precisely in the center of the field of view of the eye-piece. It is adjusted by means of four screws, two of which are shown at A A.

These operations are performed by the maker so permanently as to need no further attention from the engineer, and the heads of the screws, by which these adjustments are made, are covered by a thin ring which protects them from disturbance.
523. Adjustment by setting between two points, or the "PegMethod." Drive two pegs several hundred feet apart, and set the instrument midway between them. Level, and sight to the rod held on each peg. The difference of the readings will be the true difference of the heights of the pegs, no matter how much the level may be out of adjustment.

Then set the level over one peg, and sight to the rod at the other. Measure the height of the cross-hairs above the first peg. The difference of this and the reading on the rod should equal the difference of the heights of the two points, as previously determined. If it does not, set the target to the sum or difference of the height of the cross-hairs above the first peg, and the true difference of height of the points, according as the first point is higher or lower than the second, and hold the rod on the second point. Sight to it, and raise or lower one end of the bubble-tube until the horizontal cross-hair does bisect the target when the bubble is in the center. Then perform the " third adjustment."

Instead of setting over one peg, it is generally more convenient

Fig. 370.

to set near to it, and sight to a rod held on it, and use this reading instead of the measured height of the cross-hairs.
N. B. -This verification should always be used for every level, even after the three usual adjustments have been made; for it is independent of the equality of the collars.

In running a long line of levels, let the last sight at night be taken midway between the last two "turning-point" pegs, and in the morning try their difference by setting close to the last one. This tests the level every day with very little extra labor.
524. Egault's Level, In this level the bubble-tube is not connected with the telescope. It is used thus:

Level and sight as usual. Then turn the telescope upside down, end for end, and half-way around horizontally, and sight again. Half the sum of the two readings is the correct one, no matter how much the instrument is out of adjustment (assuming the collars to

Fig. 3 3is.
 be of equal size) ; for the errors then cancel each other. This is the one used principally in France.

The rod used with it is marked with numbers only half the real heights abore its bottom. Then the sum of the readings is the true one. Thus the rod itself takes the mean of the readings.

Fig. 372.

525. Troughton's Level. In this the bubble-tube is permanently fastened in the top of the telescope-tube. It is adjusted by the "peg method," or some similar one, the cross-hair being mored up or down until the observation gives the trne difference of height of the pegs when the bubble is in the center. Then make the "third adjustment," by means of the screws under the telescope.
526. Gravatt's Level, or the "Dumpy Level." Its diameter is very great, thus giving more light. Its bubble is on the top, and can be seen in a small inclined mirror, by the observer. It also has a cross-level.

52\%. Lenoir's Level. In this, the telescope carries, at each end, a steel block, whose upper and lower faces are made perfectly parallel. They are placed

Fig. 373.
 on a brass circle, which is made level by reversing a level placed upon the upper surface of the steel blocks.
528. Tripods. These consist of three legs, shod with iron, and connected by joints at the top. There are many different forms,

Fig. 374.
 the most common of which is given in Fig. 36\% Other forms are given in Art. 476. Lightness and stiffness are the desired qualities.

Stephenson's tripod has a ball-andsocket joint below the parallel plates, so as to admit of being at once set nearly level on very steep slopes.
"Quick-leveling" tripod-heads, for quickly setting the levelingplates nearly level, are made of various patterns.

Extension tripods are manufactured which provide for lengthening and shortening the legs of the tripod.
529. Rods. These should be made of light, well-seasoned wood. A plumb or level attached to them will show when they are held
vertically. To detect whether the rod leans to or from the instrument, its front may be angular or curred. If angular, when held leaning toward the instrument, the lines of di-

Fig. 375. Fig. 376.
 vision will appear as in Fig. 3\%j. When leaning from the instrument, they will appear as in Fig. 3i6. They are usually divided to feet, tenths, and hundredths.
530. Target. This is a plate of iron or brass, attached to the rod in such a way that it may be moved up and down the rod and clamped in any position. The face of the target should be painted of such a pattern that, when sighting to it, it may be rery precisely bisected by the horizontal cross-hair. Some of the many rarieties are given in Figs. 377-385.

Those represented in Figs. 3\%\%, 3\%8, and 3\%9 are bad, because

Fig. 377.


Fig. 380.


Fig. 383.


Fig. 378.


Fig. 381.


Fig. 384.


Fig. 379.


Fig. 382.


Fig. 385.

the cross-hair may be above or below the middle of the target by its full thickness, as magnified by the eye-piece of the telescope without the error being perceptible. The nest three, Figs. 380,

381, and 382 , depend upon the nicety with which the eye can determine if a line bisects an angle. Fig. 383 depends upon the accuracy with which the eye can bisect a space. Fig. 384 depends upon the accuracy with which the eye can bisect a circle. Figs. 381, 382, and 385 are the best forms for use. Red and white are the best colors.
531. Vernier. The target carries a vernier, by which smaller spaces may be measured than those into which the rod is divided. It may be placed on the side of an aperture, in the face of the target, through which the divisions on the rod can be seen, or carried on the back or side of the rod by the target-clamp.

## 532. The New York

 Rod (Fig. 386). This is usually in two pieces, sliding one upon the other, and connected by a tongue. It is graduated to tenths and hundredths of a foot, and
half the target is used as on other rods. For greater heights, the target is fixed at six and a half feet, and the back part of the rod, which carries the target, is shoved up (Fig. 386) until the target is bisected by the cross-hairs. Its height is then read off on the side of the rod, on which the numbers run downward, and on

Fig. 389.
 which is a second rernier, which gives the precise reading. It is convenient for its portability, but apt to be too tight or too loose, as the weather is moist or dry. Sometimes it is in three pieces, as in Fig. $88 \%$
533. The Boston Rod (Fig. 388). This is usually in two parts, like the New York rod. The target is rectangular, and is fastened to one of the pieces near its extremity. For heights less than six feet, the rod is held with the target-end down, and the target is moved up by sliding up the piece which carries it. For heights abore six feet, the rod is turned end for end, bringing the target-end up, and then sliding up the piece which carries the target.
534. The Philadelphia Rod (Fig. 389). This is in two parts, held together by brass clamps, and is furnished with a target. It is graduated and painted so as to be used as a "speaking-rod," or with a target. When the target is used, the vernier on the target is read for height up to seven feet. For greater heights, the target is clamped at seren feet, and the part to which the target is clamped is slid up, and the vernier on the upper clamp is used.
535. Speaking-Rods. These are rods which are read without targets, the dirisions and subdivisions being painted on the face of the rod. They produce great saring of time and increase of accuracr.

In one form (Fig. 390) the face of the rod is dirided into tenths of feet, and smaller dirisions estimated. In Bourdaloue's rod the dirisions are each four centimetres
( $1 \cdot 6 \mathrm{inch}$ ), and are numbered at half their value. He arranges them as in Fig. 391.

Gravatt's Rod (Fig. 392). This is divided to 0.01 foot. The upper hundredth of each tenth extends across the rod. Each half-tenth is marked by a dot; each halffoot by two dots. Every other tenth is numbered, and the numbers are each $0 \cdot 1$ high. It is in three parts, which slide into each other like a telescope.

Barlow's Rod (Fig.

Fig. 390.


Fig. 391.
 393). In this the divisions are marked by triangles, each 0.02 foot high, so that it

Fig. 393.


Fig. 394.
 reads to hundredths, and less by estimation. This is based on the power the eye has in bisecting angles.

Stephenson's Rod (Fig. 394). This is based upon the principle of the diagonal scale. Each tenth is bisected by a horizontal line, and the diagonals enable the observer to read to hundredths.

Conybeare's Rod (Fig. 395). It reads to hundredths of a foot by means of the cross-hair bisecting the tops and bottoms and angles of hexagons. The
odd tenths are made white and the even ones black. The figures are placed so that their centers are opposite the divisions they refer to.

Fig. 395.


Fig. 396.


Pemberton's Rod (Fig. 396). This is on the principle of nine verniers placed side by side. It reads to hundredths, which are given by counting up from the dot which the hair bisects, to the dot in the same rertical line which is bisected by one of the horizontal lines which mark the tenths. The inventor claims that it can be read nine times as far as Gravatt's.

On all speaking-rods, to avoid confounding numbers, such as 3 and 8 , they may be marked thus :

$$
1 \text {. } 2 . \mathrm{III} .4 \text {. V . } 6 \text {. \% . } 8 \text {. IX . X . } 11 \text {. XII. }
$$

The French, who go by tenths, use the following :

$$
1 \cdot 2 \cdot \mathrm{~T} \cdot 4 \cdot \mathrm{~V} \cdot 6 \cdot \% \cdot 8 \cdot \mathrm{~N} \cdot \mathrm{X} .
$$

The figures are sometimes placed with their tops on a level with the tops of the dimensions they mark-e. g., feet ; and sometimes with their middles on the dividing line.

## THE PRACTICE.

536. Field Routine ; or, how to start and go on :
537. The rodman holds the rod on the starting-point, which may be a peg, a door-sill, or other "bench-mark." He stands square behind his rod, and holds it as nearly rertical as possible.
538. The leveler sets up the instrument, somewhere in the direction in which he is going, but not necessarily, or usually. in the precise line. He then levels the instrument by the parallel platescrews, sights to the rod, and notes the reading, whether of target or speaking-rod, as a "back-sight" (B. S.), or + (plus) sight; entering it in the proper column of one of the tabular forms of field-book, given in the following articles.
539. The rodman is then sent ahead about as far as he was behind, and he there drives a "level-peg" nearly to the surface of the ground, or finds a hard, well-defined point, and holds the rod upon it.
540. The leveler then again sights to the rod, and notes the reading as a "fore-sight" (F. S.), or - (minus) sight. The difference of the two readings is the difference of the heights of the points.
541. He then takes up the instrument, goes beyond the rod, any convenient distance, sets up again, and proceeds as in paragraph 2 ; and so on for any number of points, which will form a series of pairs. The successive observations of each pair give their difference of heights, and the combination of all these gives the difference of heights of the first and last points of the series.
542. If the vertical cross-hair be strictly vertical, it will determine whether the rod leans to the right or left. To know whether the cross-hair is vertical or not, try whether it coincides with a plumbline, or sight to some fixed point, turn the telescope from side to side horizontally, and see if the horizontal cross-hair continues to cover the spot. If it does not, turn the telescope around in the wyes till it does; then it is truly horizontal, and the other hair, being perpendicular to it, is truly vertical. To know whether the rod leans forward or backward, have the rodman move it from and to himself. If the line bisected by the cross-hair descends in both motions, the rod was vertical ; if the line rises, the rod was leaning. The lowest reading is the true one.
\%. When a target is used, signals are made by the leveler with the hand, "up" and "down," to indicate in which direction to move the target. Drawing the hand to the side signifies "stop," and both hands brought together above the head signifies "all right." The rodman should move the target fast at first, and slowly after having passed the right point. When signaled "all right," he should clamp the target and show again. Then call out the reading before moving, and show it to the leveler, as either passes the other.
543. We have thus far supposed that only the difference of heights of the two extreme points is desired. But when a section or profile of the ground is required, the rod must be held and observed, at
each change of slope of the ground, or at regular distances; usually, for railroad-work, at every hundred feet, and also at any change of slope between those points.

Any number of points, within sight, may have their relative heights determined at one setting of the level.

The names back-sight (B. S.) and fore-sight (F. S.) do not necessarily mean sights taken looking forward or backward (though they are generally so for turaing-points), but the first sight taken, after setting up the instrument, is a B. S. or + (plus) sight, and all following ones, taken before remoring the instrument, are F. S.'s, or - (minus) sights. The full meaning of this will appear in considering the forms of field-book.

All but the first and last points sighted to are called intermediate points, or "intermediates." The last point sighted to before moving the instrument is called a turning-point, or changingpoint.

The first and last sights, taken at any one setting of the instrument, require the greatest possible accuracy. The intermediate points may be taken only to the nearest tenth, or hundredth at most ; because any error in them will not affect the final result, but only the height of that single point at which it was taken.

Two rodmen are often used to save the time of the leveler. Then it is well to use a target-rod for the "turning-points," which are often distant and need most precision, and a speaking-rod for the intermediate points. Where one rod is used, the rodman should keep notes of the readings at the turning-points.
537. Field-Notes. The beginner may sketch the heights and distances measured, in a profile or side view, as in Fig. 39\%. But when the observations are numerous, they should be placed in one of the tabular forms given on the following pages.
538. First Form of Field-Book. In this, the names of the points or "stations," whose heights are demanded, are placed in the first column, and their heights, as finally ascertained, in reference to the first point, in the last column. The heights abore the startingpoint are marked + , and those below it are marked -. The back-
sight to any station is placed on the line below the point to which it refers. When a back-sight exceeds a fore-sight, their difference

Fig. 397.

is placed in the column of " Rise" ; when it is less, their difference is a "Fall." The following table represents the same observations as the last figure, and their careful comparison will explain any obscurities in either :

| stations. | distances. | BACKSIGHTS. | FORE- SIGHTS. | Rise. | fall. | $\begin{aligned} & \text { Total } \\ & \text { HEIGHTS. } \end{aligned}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| A |  |  |  | $\begin{aligned} & +1 \cdot 00 \\ & +5 \cdot 00 \end{aligned}$ |  | $0 \cdot 00$ |
| B | 100 | $2 \cdot 00$ | $6 \cdot 00$ |  | $-4 \cdot 00$ | $-4 \cdot 00$ |
| C | 60 | $3 \cdot 00$ | $4 \cdot 00$ |  | $-1 \cdot 00$ | $-5 \cdot 00$ |
| D | 40 | $2 \cdot 00$ | $1 \cdot 00$ |  |  | $-4 \cdot 00$ |
| E | 70 | $6 \cdot 00$ | $1 \cdot 00$ |  |  | $+1 \cdot 00$ |
| F | 50 | $2 \cdot 00$ | 6.00 |  | $-4 \cdot 00$ | $-3 \cdot 00$ |
|  |  | 15.00 | $18 \cdot 00$ |  | $-3 \cdot 00$ |  |

The above table shows that $B$ is 4 feet below $A$; that $C$ is 5 feet below A ; that E is 1 foot above A ; and so on. To test the calculations, add up the back-sights and fore-sights. The difference of the sums should equal the last "total height."

An objection to this form is that the back-sights come on the line below the station to which they are taken, which is embarrassing to a beginner.

When " intermediate" observations are taken, the "foresights" taken to these intermediate points are put down in their proper column, and are also set down in the column of "backsights" ; so that, when the two columns are added up, any error in
these intermediate sights (which are usually not taken rery accurately) will be canceled, and will not affect the final result. The effect is the same as if, after the fore-sight to the intermediate point had been taken, the instrument had been taken up and set down again at precisely the same height as before, and a back-sight had then been taken to the same point. Hence, in this form, the "turning-points" are those stations which have different backsights and fore-sights, while those which have them the same are "intermediates."

The following figure and table represent the same ground as the

- Fig. 398.

preceding one, but with ouly two settings of the instrument. D is the turning-point :

| stations. | distances. | $\begin{aligned} & \text { BACK- } \\ & \text { SIGHTS. } \end{aligned}$ | $\begin{aligned} & \text { FORE- } \\ & \text { SIGHTS. } \end{aligned}$ | Rise. | Fall. | Total HEIGHTs. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| A |  |  |  | $\begin{aligned} & 1 \cdot 00 \\ & 5 \cdot 00 \end{aligned}$ | $\begin{aligned} & 4 \cdot 00 \\ & 1 \cdot 00 \end{aligned}$ | $\begin{array}{r} 0 \cdot 00 \\ -4 \cdot 00 \\ -\cdot \cdot 00 \\ -+\cdot 00 \\ +1 \cdot 00 \\ -3 \cdot 00 \end{array}$ |
| C |  | 6.00 | $\uparrow \cdot 00$ |  |  |  |
| I |  | $7 \cdot 00$ | $6 \cdot 00$ |  |  |  |
| E |  | $9 \cdot 00$ | $4 \cdot 00$ |  |  |  |
| F |  | $4 \cdot 00$ | $8 \cdot 00$ |  | $4 \cdot 00$ |  |
|  |  | $+28.00$ | $-31 \cdot 00$ |  | 300 |  |

In leveling for "sections," the distances between the .points leveled must be recorded. They are usually put down after the stations to which they are measured; although in survering with the compass, etc., they are put down after the stations from which they are measured. In the following notes, which contain inter-
mediate stations, they are put down before the stations to which they are measured. It should be remembered that these distances are measured between the points at which the rod is held, and have no reference to the points at which the instrument is set up:

| distances. | stations. | $\begin{aligned} & \text { BACK- } \\ & \text { SIGHTS. }+ \end{aligned}$ | $\begin{aligned} & \text { Fore- } \\ & \text { SIGHTS. } \end{aligned}$ | RISE. | FALL. | $\begin{aligned} & \text { Total } \\ & \text { HEIGHTS. } \end{aligned}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 100 | 260 261 | $4 \cdot 576$ | $3 \cdot 726$ | 0.850 |  | $91 \cdot 397$ $92 \cdot 247$ |
| 100 | 262 | $5 \cdot 420$ | $4 \cdot 500$ | 0.920 |  | $93 \cdot 167$ |
| 100 | 263 | $4 \cdot 500$ | $3 \cdot 170$ | 1•330 |  | $94 \cdot 497$ |
| 40 | $263 \cdot 40$ | $4 \cdot 910$ | $4 \cdot 938$ |  | $0 \cdot 028$ | $94 \cdot 469$ |
| 60 | 264 | $4 \cdot 938$ | $6 \cdot 386$ |  | $1 \cdot 448$ | $93 \cdot 021$ |
| 100 | 265 | $3 \cdot 380$ | $4 \cdot 640$ |  | $1 \cdot 260$ | $91 \cdot 761$ |
| 100 | 266 | $4 \cdot 640$ | $5 \cdot 400$ |  | $0 \cdot 760$ | $91 \cdot 001$ |
| 70 | $266 \cdot 70$ | $2 \cdot 760$ | $3 \cdot 070$ |  | $0 \cdot 310$ | $90 \cdot 691$ |
| 30 | 267 | $3 \cdot 070$ | $3 \cdot 750$ |  | 0.680 | $90 \cdot 011$ |
| 100 | 268 | 6.750 | 5.925 |  | $3 \cdot 175$ | $86 \cdot 836$ |
|  |  | $41 \cdot 944$ | 46.505 |  | $-4.561$ |  |
|  |  |  | $41 \cdot 944$ |  | + 91.397 |  |
|  |  |  | -4.561 |  | 86.836 |  |

539. Second Form of Field-Book. This is presented below. It refers to the same stations and levels noted in the first table, and shown in Fig. 397 :

| stationg. | distances. | BACK- SIGHTS. | $\begin{gathered} \text { HFIGHT OF } \\ \text { INSTRUMENT } \\ \text { ABOVE DATUM. } \end{gathered}$ | FORE- SIGHTS. | $\begin{gathered} \text { Total } \\ \text { HEIGHTS. } \end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| A |  |  |  |  | $0 \cdot 00$ |
| B | 100 | $2 \cdot 00$ | $+2 \cdot 00$ | $6 \cdot 00$ | $-4.00$ |
| $\bigcirc$ | 60 | $3 \cdot 00$ | $-1.00$ | $4 \cdot 00$ | $-5.00$ |
| D | 40 | $2 \cdot 00$ | $-3.00$ | $1 \cdot 00$ | $-4.00$ |
| E | 70 | 6.00 | $+2.00$ | $1 \cdot 00$ | $+1 \cdot 00$ |
| F | 50 | $2 \cdot 00$ | $+3.00$ | $6 \cdot 00$ | $-3.00$ |
|  |  | 15.00 |  | 18.00 | $-3 \cdot 00$ |

In the preceding form it will be seen that a new column is introduced, containing the height of the instrument-i. e., of its line of sight-not above the ground where it stands, but above the Datum, or starting-point, of the levels. The former columns of " rise" and "fall" are omitted. The preceding notes are taken thus: The height of the starting-point, or "datum," at A, is 0.00 . The instrument being set up and leveled, the rod is held at A .

The back-sight upon it is $2 \cdot 00$; therefore the height of the instrument is also $2 \cdot 00$. The rod is next held at B. The fore-sight to it is 6.00 . That point is therefore 6.00 below the instrument, or $2.00-6.00=-4.00$ below the datum. The instrument is now moved, and again set up, and the back-sight to $B$, being $3 \cdot 00$, the height of the instrument is $-4 \cdot 00+3 \cdot 00=-1 \cdot 00$, and so on ; the height of the instrument being always obtained by adding the back-sight to the height of the peg on which the rod is held, and the height of the next peg being obtained by cubtracting the foresight to the rod held on that peg, from the height of the instrument.

This form is better than the first form, in leveling for a section of the ground to make a profile; or when several observations are to be made at one setting of the level ; or when points of desired heights are to be established, as in "leveling-location."

This form may be modified by putting the back-sights on the same line with the stations to which they are taken. This avoids the defect of the first form, but introduces the new defect of writing them down after the number which they precede, in a backhanded way, which may be a source of error.

This modification is shown in the following table, which corresponds to Fig. 398. In the column of fore-sights, the "turn-ing-points" (T. P.), and "intermediate points" (Int.), are put in separate columns ; so that, to prove the work, the difference of the sum of the back-sights and of the sum of the turning-point foresights, is the number which should equal the difference of the heights of the first and last points :

| stations. | distances. | $\begin{aligned} & \text { Back- } \\ & \text { sIGHTs. }+ \end{aligned}$ | $\begin{aligned} & \text { HEIGHT OF } \\ & \text { HNSTRUMENT. } \end{aligned}$ | fore-sights. - |  | $\begin{aligned} & \text { TOTAL } \\ & \text { HEIGHTS. } \end{aligned}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  | т. p. | гیт. |  |
| $\begin{aligned} & \mathrm{A} \\ & \mathrm{~B} \\ & \mathrm{C} \\ & \mathrm{D} \\ & \mathrm{E} \\ & \mathrm{~F} \end{aligned}$ |  | $2 \cdot 00$ | +2.00 |  | 6.00 | 0.09 -4.00 |
|  |  |  |  |  | $7 \cdot 00$ | $-5.00$ |
|  |  | $9 \cdot 00$ |  | $6 \cdot 00$ |  | $-4.00$ |
|  |  |  | +5.00 | $8 \cdot 00$ | $4 \cdot 00$ | $\begin{array}{r} +1 \cdot 00 \\ -3 \cdot 00 \end{array}$ |
|  |  | $+11.00$ |  | -14.00 +11.00 |  |  |
|  |  |  |  | $-3.00$ |  |  |

When a line is divided up into stations of 100 feet each, as on railroad-work, the number of the station indicates its distance from the starting-point. When an observation is taken at a point between these hundred-feet stations, it is noted as a decimal, thus : Station $4 \cdot 60$ is 460 feet from the starting point. In the field-notes of such work, the column of distances may be omitted, as in the following table. The heights and distances are the same as in the last table under Art. 538 :

| stations. | back-sights. | height of instrument | Fore-sights. |  | $\begin{gathered} \text { TOTAL } \\ \text { HEIGHTS. } \end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | т. $\mathbf{P}$. | int. |  |
| 260 | $4 \cdot 576$ | $95 \cdot 973$ |  |  | $91 \cdot 397$ |
| 261 | $5 \cdot 420$ | $97 \cdot 667$ | $3 \cdot 726$ |  | $92 \cdot 247$ |
| 262 |  |  |  | 4:500 | $93 \cdot 167$ |
| 263 | $4 \cdot 910$ | 99•407 | 3•170 |  | $94 \cdot 497$ |
| $263 \cdot 40$ |  |  |  | $4 \cdot 938$ | $94 \cdot 469$ |
| 264 | $3 \cdot 380$ | $96 \cdot 401$ | $6 \cdot 386$ |  | $93 \cdot 021$ |
| 265 |  |  |  | $4 \cdot 640$ | ${ }^{91} \cdot 761$ |
| ${ }_{266}{ }^{66} 70$ | $2 \cdot 760$ | $93 \cdot 761$ | $5 \cdot 400$ | 3.070 | $9{ }^{91} \cdot 691$ |
| 267 |  |  |  | $3 \cdot 750$ | 90.011 |
| 268 |  |  | 6.925 |  | 86.836 |
|  | $+21 \cdot 046$ |  | $-25 \cdot 607$ $+21 \cdot 046$ |  |  |
|  |  |  | $-4.561$ |  |  |
|  |  |  | +91397 |  |  |
|  |  |  | +86.836 |  |  |

540. Third Form of Field-Book. In this the back-sights are placed directly under the height of the station to which they are taken, which lessens the chance of making mistakes in adding to get the height of instrument. The height of instrument is distinguished by being included between two horizontal lines. The following table refers to the same ground as the preceding one:

| stations. | Fore-sights. | heights. | remarks. |
| :---: | :---: | :---: | :---: |
| 260 | $3 \cdot 726$ | $\begin{array}{r} 91 \cdot 397 \\ 4 \cdot 576 \end{array}$ |  |
|  |  | 95.973 |  |
| 261 |  | $\begin{array}{r} 92 \cdot 247 \\ 5.420 \end{array}$ |  |
|  |  | $97 \cdot 667$ |  |
| $\begin{aligned} & 262 \\ & 263 \end{aligned}$ | $\begin{aligned} & 4 \cdot 500 \\ & 3 \cdot 170 \end{aligned}$ | $\begin{array}{r} 93 \cdot 167 \\ 94 \cdot 497 \\ 4 \cdot 910 \end{array}$ |  |
|  | $\begin{aligned} & 4 \cdot 938 \\ & 6.386 \end{aligned}$ | $99 \cdot 407$ |  |
| $\begin{array}{r} +40 \\ +264 \end{array}$ |  | $\begin{array}{r} 94 \cdot 469 \\ 93.021 \\ 3.380 \end{array}$ |  |
|  | $\begin{aligned} & 4 \cdot 640 \\ & 5 \cdot 400 \end{aligned}$ | $96 \cdot 401$ |  |
| $\begin{aligned} & 265 \\ & 266 \end{aligned}$ |  | $91 \cdot 761$ <br> $91 \cdot 001$ <br> 2.760 |  |
|  |  | $93 \cdot 761$ |  |
| +70 +267 | $\begin{aligned} & 3 \cdot 070 \\ & 3.750 \end{aligned}$ | $90 \cdot 691$ 90.011 |  |
| 268 | 6.925 | 86.836 |  |

541. Best Length of Sights. There are two classes of inaccuracies. With rery long sights, the errors of imperfect adjustment and curvature are greatest ; the former varying as the length, and the latter as the square of the length. With rery short sights, and therefore more numerous, the errors of inaccurate sighting at the target are greatest. The best usual mean is from 200 feet to 300 feet, or more if equal distances for back-sights and fore-sights to turning-points can be obtained.
542. Equal Distances of Sight. They are always rery desirable. They are most easily determined, when no stakes hare been previously set, by "stadia" cross-hairs in the telescope of the level.
543. Datum-Level. This is the plane of reference, from which, above it or below it, usually the former, the heights of all points of the line are reckoned.

It may be taken as the height of the starting-point. If the line descends, it is better to call the starting-point 10 feet or 100 feet above some imaginary plane, so that points below the startingpoint may not have minus-signs.

It is desirable to refer all levels in a country to some one datum. This is usually the surface of the sea, and, for general purposes, mean tide is best. Low-water mark should be the datum when the levelings are connected with harbor-surveys, whose soundings always refer to low water. High-water mark should be used when the levelings relate to the drainage of a country.
544. Bench-Marks (B. M.). These are permanent objects, natural or artificial, whose heights above the datum are determined and recorded for future reference.

Good objects are these: Pointed tops of rocks, tops of milestones, stone door-sills, tops of gate-posts or hinges, and generally any object not easily disturbed, and easily described and found.

A knob made on the spreading root of a tree is good. A nail may be driven in it, and the tree "blazed" and marked, as in Fig. 399. A stake will do till frost.

Bench-marks should be made near the starting-point of a line of levels ; near where the line crosses a road ; on
 each side of a river crossed by it; at the top and bottom of any high hill passed over ; and always at every half-mile or mile.

The precise location and description of every bench-mark should be noted very fully and precisely, and in such a way that an entire stranger could find it, with the aid of the notes.
545. Check-Levels, or Test-Levels. No single set of levels is to be trusted ; but they must be tested by another set, run between the bench-marks (B. M.'s), though not necessarily over the same ground.

A set of levels will verify themselves if they come around to the starting-point again.
546. Limits of Precision. Errors and inaccuracies should be carefully distinguished. For the latter, every leveler must make a standard for himself, so as to be able, in testing his work, to distinguish any real error from his usual inaccuracy.

The result of four sets of levelings, in France, of from 45 to 140 miles, averaged a difference of $\frac{1}{10}$ foot in 43 miles, and the greatest error was $\frac{1}{3}$ foot in 56 miles.

A French leveler, M. Bourdaloue, contracts to level the benchmarks of a railroad survey to within 0.002 foot per mile, or $\frac{1}{10}$ foot per 50 miles.

In Scotland, the difference of two sets of levels of 26 miles was 0.02 foot.
547. Trial-Levels, or Flying-Levels. Their object is to get a general approximate idea of the comparative heights of a portion of the country, as a guide in choosing lines to be leveled more accurately. More rapidity is required, and less precision is necessary. The distances may be measured at the same time by stadia-hairs.
548. Leveling for Sections. The object of this is to measure all the ascents and descents of the line, and the distances between the points at which the slope changes ; so that a section or profile of it can be made from the observations taken.

The line of a railroad is usually set out by a party with compass or transit, who drive at erery hundred feet a large stake with the number of the station on it, and beside it a small level-peg, even with the surface of the ground. On this the rod is held for the observations. The level-peg is set in " line," and the large stake a foot or two to one side.
549. Profiles. A profile is a section of ground by a vertical plane or cylindrical surface,* passing through the line along which a profile is desired. It represents to any desired scale the heights and distances of the various points of a line, its ascents and descents, as seen in a side riew. It is made thus : Any point on the

[^62]paper being assumed for the first station, a horizontal line is drawn through it ; the distance to the next station is measured along it, to the required scale ; at the termination of this distance a vertical line is drawn ; and the given height of the second station above or below the first is set off on this vertical line. The point thus fixed determines the second station, and a line joining it to the first station represents the slope of the ground between the two. The process is repeated for the next station, etc.

But the rises and falls of a line are always very small in proportion to the distances passed over, even mountains being merely as the roughnesses of the rind of an orange. If the distances and the heights were represented on a profile to the same scale, the latter would be hardly visible. To make them more apparent, it is usual to "exaggerate the vertical scale" tenfold, or more-i. e., to make the representation of a foot of height ten times as great as that of a foot of length, as in Fig. 397, in which one inch represents one hundred feet for the distances, and ten feet for the heights.

In practice, engraved profile-paper is generally used, which is ruled in squares or rectangles, to which any arbitrary values may be assigned.

When the line leveled over is not straight, the profile, whose length is that of the line straightened out, will extend beyond the "plan" when both are on the same sheet.
550. Cross-Levels. These show the heights of the ground on a line at right angles to the main line. They give "cross-sections"

Fig. 400.

of it. In the note-book they are put on the right-hand page. They may be taken at the same time with the other levels, or inde-
pendently. In taking cross-levels where the slopes are quite steep, as in mountain districts, frequent settings of the instrument are necessary.

A much more rapid method is by the use of "cross-section rods." These are two rods, one of which is about ten or twelve feet long, provided with a bubble-tube near each end, so as to be held level, and graduated to feet, tenths, and hundredths. The other is simply a graduated rod. The manner of using them is shown in Fig. 400.

A slope-level is sometimes used. (See "Angular Surveying.")

## DIFFICULTIES.

551. Steep Slopes. In descending or ascending a hill, the instrument and the rod should be so placed that the sight should strike as near as possible to the bottom of the rod on the up-hill side, and the top of the rod on the down-hill side.

Try this by leveling over two screws, setting the instrument so that one pair of opposite plate-screws shall point in the direction of the line, but do not be too particular ; it is a waste of time.

Doing this produces sights of unequal length. The rod being about three times as high as the instrument, the down-hill sights will be about double the length of the up-hill ones, as shown in Fig. 401. Then set to one side of the line. This is necessary on

Fig. 401.

slopes so steep that the rod is too near the lerel to be read. If this be impossible, keep notes of the lengths of the sights to the turn-ing-points, backward and forward, and as soon as possible take
sights unequal in the contrary direction till the differences of lengths balance the former ones. When approaching a long ascent or descent, make these compensations in advance.

In leveling over a line of stakes already set, as on a railroad, at every 100 feet, if the line of sight strikes not quite up to one, drive a peg as high as you can see it, and make it a turning-point, noting it "peg" in the field-book.

In leveling across a hill or hollow, instead of setting the instru-
Fig. 402.

ment on the top of the hill or bottom of the hollow, time will be saved by the method represented in Figs. 402 and 403.

Fig. 403.

552. When the rod is a little too low, raise it alongside of a stake, or the body, and put the top of the rod "right"; then measure down from the bottom of the rod, and add it to its length.
553. When the rod is a little too high, so that the line of sight strikes the peg below the bottom of the rod, measure down from the top of the peg, and put down the sight with a contrary sign to what it would have had-i. e., if a back-sight make it minus, and if a fore-sight make it plus.
554. Then the rod is too near. When no figure is risible, raise the rod slowly till a figure comes in sight. If too near to read, and there is no target, use a field-book as target. If the instrument is exactly over the peg, measure up to the height of the crosshairs, as given by the side-screws.
555. Water. A.-A pond too wide to be sighted across. Drive a peg to the level of the water, on the first side, and observe its height, as an F. S. Then drive a peg on the other side of the pond, also to the surface of the water. Hold the rod on it. Set

Fig. 404.

up the level beyond it, and sight to it as a B. S., and put down the obserration as if it had been taken to the first peg.

| Fore-sights. | stations. | heights. | back-sights. | $\ominus$ |
| :---: | :---: | :---: | :---: | :---: |
| 5•0 | 74 74.89 | $\begin{aligned} & 50.00 \\ & 48.00 \end{aligned}$ | $3 \cdot 00$ | 53.00 |
|  | 81.89 \} |  | $6 \cdot 00$ | 54•00 |

There must be no wind in the direction of the line of Jerel.
B.-For leveling across a running stream. Set the two pegs in a line at right angles to the current, although the line to be leveled may cross it obliquely.

If a profile or section of the ground under the water be required, find the height of the surface, and measure the depths below this at a sufficient number of points, measuring the distances also, and put these depths down as fore-sights.
556. A Swamp, or Marsh. This can not be treated like a poud, for the water mar seem nearly staguant while its surface has considerable slope, its flow being retarded by regetation. If only slightly "shakr," hare an obserrer at each end of the level. If
more so, push the legs down as far as they will go, and let both observers lie down on their sides. If still more "shaky," drive three stakes or piles, to support the legs of the tripod, and stand the tripod on them.

A water-level will level itself. Use that for intermediate points on the swamp, and test the result by leveling around the swamp with the spirit-level.
557. Underwood. If it can not be cut away, set the instrument on some eminence, natural or artificial.
558. Board Fence. Run a knife-blade through one of the boards, and hold the rod upon it on each side of the fence, as if it were a peg, keeping the blade in the same horizontal position while the rod and instrument are taken over.
559. A Wall. First Method. Drive a peg at the bottom of the wall, on the first side, and observe on it. Measure the height of the wall above the peg, and put this down as a B. S. Drive another peg on the other side of the wall ; measure down to it from the top of the wall, and put that down as an F. S., just as if the level had been set in the air at the height of the top of the

Fig. 405.

wall, and this B. S. and F. S. had been really taken. Set up the instrument beyond the wall, take a B. S. to this peg, and go on as usual.

| Fore-sights. | stations. | Heights. | back-sights. | $\oplus$ |
| :---: | :---: | :---: | :---: | :---: |
|  | 50 | 74.00 | 5.00 | 79.00 |
|  | Peg. | 76.00 | 18.00 | 89.00 |
| 12.00 | Peg. | 77.00 | 2.00 | 79.00 |
| 1.00 | 51 | 78.00 |  |  |

25

Second Method. Mark where the line of sight strikes the wall ; measure up to the top of the wall, and put this down as an F. S., with a plus-sign, as in 553, where the line of sight struck below the top of the peg.

On the other side of the wall, sight back to it, and mark where the line of sight strikes. Measure to the top of the wall, and put this down as a B. S., with a minus-sign, and then go on as usual.
560. House. First try to find some place for the instrument from which you can see through, by opening doors or windows. Or, find some place in the house where you can set the instrument and see both ways, or hold the rod at some point inside, and look to it from front and back. A straight stick may be used if the rod can not be held upright, and the height measured on the rod.
561. The Sun. It often causes the leveler much difficulty-

1. By shining in the object-glass. If the instrument has a shade on it, draw it out. If not, shade the glass with jour hand or hat, or set the instrument to one side of the line.
2. By heating the level unequally in all its parts. Holding an umbrella over it will remedy this.
3. By causing irregular refraction. Some parts of the ground become heated more than others, and therefore rarefy the air at those places. This can not be avoided nor corrected.
4. Wind. Watch for lulls of rind, and observe then sereral times, and take the mean. The least wind is at daybreak.
5. Idiosyncrasies. Different persons do not see things precisely alike. Each indiridual may hare an inaccuracy peculiar to himself. One may read an observation higher or lower than another equal in skill. Also, a person's right and left eye may differ. This difference in individuals is termed their "personal equation."

To test the accuracy of jour eje, turn the head so as to bring the eyes in the same vertical line, and sight to the rod held horizontally. Nota where the rertical hair strikes. Then turn the
head to the other side, so as to invert the position of the eyes, and then sight again. As before, the mean of the two readings is the correct one.
564. Reciprocal Leveling. This is to be used when it is impossible to set midway between the two points, and the distance can not be readily determined.

Set the instrument over $A$, and sight to a rod at $B$, and note Fig. 406.

reading. The difference of the reading and of the height of the cross-hairs gives $a$ difference of height of A and B. Then set up at B , and observe to A , similarly. A new difference of height is obtained. The mean of these two is the correct one.

Itt. of cross-hairs above peg at $\mathrm{A}=4 \cdot 3^{\prime} \mathrm{Ht}$. of cross-hairs above peg at $\mathrm{B}=4 \cdot 9^{\prime}$ Observation to $B=7 \cdot 0^{\prime} \quad$ Observation to $A=4 \cdot 2^{\prime}$ Diff. of height $=2 \cdot 7^{\prime} \quad$ Diff. of height $=0 \cdot 7^{\prime}$ True difference $=\frac{1}{2}\left(2 \cdot 7^{\prime}+0 \cdot 7^{\prime}\right)=1 \cdot 7^{\prime}$.

Otherwise, set the instrument at an equal distance from each point, as $\mathrm{A}^{\prime}$ and $\mathrm{B}^{\prime}$, and observe to each in turn. The mean of the two differences of height obtained will be the true difference, as before.

## LEVELING LOCATION.

565. Its Nature. It is the converse of the general problem of leveling, which is to find the difference of heights of two given points. This consists in determining the place of a point of any required height above or below any given point.

To do this, hold the rod on some point of known height above
the datum-level ; sight to it, and thus determine the height of the cross-hairs. Subtract from this the desired height of the required point, and set the target at the difference. Hold the rod at the place where the height is desired, and raise or lower it till the cross-hair bisects the target. Then the bottom of the rod is at the desired height. Usually, a peg is driven till its top is at the given height above the datum.
566. Difficulties. If the difference of height be too much to be measured at one setting of the instrument, take a series of levels up or down to the desired point. So, too, if they be far apart; and thus find a place where, the instrument having a known height of cross-hairs, the target can finally be set, as before.

If the ground be so low or so high that a peg can not be set with its top at the required height, drive a peg till its top is just above the surface of the ground. Observe to the rod on it, determine its height above or below the desired point, and note this on a large stake driven beside it ; or, place its top a whole number of feet above or below the required height, and mark the difference on it, or on a stake beside it.
567. Staking out Work. When embankments and excarations are to be made for roads, etc., side-stakes are set at points in their intended outside edges-

Fig. 407.
 i. e., where their slopes will meet the surface of the ground ; and the height which the ground at those points is abore or below the required height or depth of the top or bottom of the finished work, is marked on these stakes with the words "cut," or "fill," or the signs + or - .

The places of the stakes are found by trial. (See Gillespie's "Road-Making," page 145.) These stakes are set to prepare the work for contractors. When the work is nearly finished, other stakes are set at the exact required height.

In staking out foundation-pits, set temporary stakes exactly above the intended bottom angles of the completed pit, thus marking out on the surface of the ground its intended shape. Take the heights of each of these stakes and move them outward such distances that cutting down from them with the proper depth and slope will bring you to the desired bottom angle.

Fig. 408.

568. To locate a Level-Line. This consists in determining on the surface of the ground a series of points which are at the same level-i. e., at the same height above some datum. Set one peg at the desired height, as in Art. 565. Sight to the rod held thereon, and make fast the target when bisected. Then send on the rod in the desired direction, and have it moved up or down along the slope of the ground, until the target is again bisected. This gives a second point. So go on as far as sights can be correctly taken, keeping unchanged the instrument and target. Make the last point sighted to a "turning-point." Carry the instrument beyond it, set up again, take a B. S., and proceed as at first.

The rod should be held and pegs driven at points so near together that the level-line between them will be approximately straight.
569. Applications. One use of this operation is to mark out the line which will be the edge of the water of a pond to be formed by a dam. In that case, a point of a height equal to that of the top of the proposed dam, plus the height which the water will stand on it (to be determined by hydraulic formulas), will be the starting-point. Then proceed to set stakes as directed in the last article.

The line from stake to stake may then be surveyed like the sides of a field, and the area to be overflowed thus determined.

Strictly, the surface of the water behind a dam is not level,
but is curved concavely upward, and is therefore higher and sets back farther than if level. The backing up of the water is called Remous.

Another important application of this problem is to obtain " contour-lines" for topography.
570. To run a Grade-Line. This consists in setting a series of pegs so that their tops shall be points in a line which shall hare any required slope, ascending or descending.

When a grade-line is to be run straight between two given points, set the level over one point, set the target at the height of the cross-hairs, hold the rod on the other point, and raise or lower one end of the instrument till the cross-hair bisects the target. Then send the rod along the line, and drive pegs to such heights that when the rod is held on them the cross-hair will bisect the target. A stake may be driven at the extreme point to the height of the target.

Another Method. Knowing the horizontal distance between the two giren points,

Fig. 403.
 and their difference of level, determine the rise or fall per hundred feet. Then drive stakes at every hundred feet, so that the top of each succeeding one is the given grade per hundred feet higher or lower, according as the grade is ascending or descending.

For example, suppose the horizontal distance from A to B is 1,200 feet, and that $B$ is 16.8 feet higher than A. The rise per hundred feet from A is 1.4 foot. Beginning at A , set stakes at every hundred feet, so that the top of each one is 1.4 foot higher than the preceding one.

A line of uniform grade or slope is not a straight line. Calling the globe spherical, this line, when traced in the plane of a great circle, would be a logarithmic spiral. On a length of sir miles, the difference in the middle between it and its straight chord would be six feet.

## CHAPTER II.

## INDIRECT LEVELING.

## METHODS AND INSTRUMENTS.

571. Vertical Surveying. Leveling may be named Vertical Surveying, or Up-and-down Surveying; Land-Surveying being Horizontal Surveying, or Right-and-left and Fore-and-aft Surveying.

All the methods of determining the position of a point in horizontal surveying may be used in vertical surveying.

The point may be determined by co-ordinates situated in a vertical plane, as in any of the systems employed in a horizontal plane.

Thus, if a balloon be held down by a single rope attached to a point in a level surface, its height above that surface is found by measur-
 ing the length of the rope. This is the direct method. It resembles that of "rectangular co-ordinates," though here only one of the co-ordinates is measured.

Fig. 411.
 The other might be situated anywhere in the surface.

If, however, the balloon be held down by two cords, its height can be determined by measuring the length of the cords and the distance between their lower ends. They correspond to the "focal co-ordinates." The required vertical heightcan be calculated by triconometry. So in the following other indirect methods :

Fig. 412.


The length of the string of a kite, and the angle which this string makes with the horizon, are the "polar co-ordinates" of the kite.

The "angles of elevation" of a meteor, observed by two persons in the same vertical plane with it, and

Fig. 413.
 at known distances apart, are its "angular co-ordinates."

Finally, an aëronaut could determine his own height by observing the angles subtended by three given objects situated on the earth's surface, at known distances, and in the same vertical plane with him. These would be "trilinear co-ordinates."

Many other systems of co-ordinate lines and angles, variously combined, may be employed.

The desired heights may also be determined by various other methods, analogous to those given for "inaccessible distances."

Combinations of measurements not in the same vertical plane may also be used, as will be shown in this chapter.
572. Vertical Angles. The vertical angles measured may be those made-either with a level line,

Fig. 415.


Fig. 414.
 or with a vertical line-by the line passing from one point to the other.

The angle B AC is called an "angle of elevation," and the angle B' A C an "angle of depression." The former angle may be called positive, and the latter negative.

The angle BAZ or $\mathrm{B}^{\prime} \mathrm{AZ}$ is called the zenith-distance of the object. It is the complement of the former an-gle-i. e., $=90^{\circ}-$ that angle taken with its proper algebraic sign. An angle of eleration, $\mathrm{BAC}=10^{\circ}$, would be a zenith-distance of $80^{\circ}$. An angle of depression, $\mathrm{B}^{\prime} \mathrm{AC}=-10^{\circ}$, would be a zenithdistance of $100^{\circ}$. The zenith-distance is preferable in important
and complicated operations, as avoiding the ambiguity of the other mode of notation.
573. Instruments. All contain a divided circle, or arc, placed vertically, and a level or plumb line. By these is measured the desired vertical angle made by the inclined line with either a level line or vertical line.

This inclined line may be an actual line or a visual line. In the former case, it may be a rod, or cord, or wire, as shown in Figs. 416-418.

This last arrangement of a cord or wire (Fig. 418) is used in mine - surveying. A
 light surveyor's chain may be similarly used, with the advantage of giving, at the same time, difference of heights and distance.

Fig. 417.


Fig. 418.


Difference of heights $=$ length of chain $\times$ sin. angle.
Horizontal distance $=$ length of chain $\times$ cos. angle.
These instruments are all "slope-measurers." They are also called Clinometers, Clisimeters, Eclimeters, etc., all meaning the same thing.
574. Slopes. These may be designated by their angles with the horizon, or by the relations of their bases and heights. The French engineers name a slope by the ratio of its height to its base-i. e.,

Fig. 419.

$\frac{\mathrm{BC}}{\mathrm{AC}}$; which is the tangent of the angle BA C. The English and Americans use the ratio of the base to the height-i. e., $\frac{\mathrm{AC}}{\mathrm{BC}}$, and make the height the unit, so that if AC $=2 \mathrm{CB}$, the slope is called 2 to 1 ; and so on.

When the inclined line is a visual line, such as the line of sight of a telescope, whose angular movements are measured on a vertical circle beside it, and when with these is combined a porizontal circle for measuring horizontal angles, the instrument is called a "transit."

## 575. Angular Profiles

Fig. 420.


A section or profile of a tolerably uniform slope is most easily obtained, as shown in the figures, by measuring the heights or depths below an inclined
 line, instead of below a level line.

A cross-section for a road may be taken in the same way.
576. Burnier's Level. It is a pear-shaped instrument, haring two graduated circles : one vertical, haring a weight attached so as

Fig. 421.
 to keep it in the same rertical posidion when in use ; and the other, a horizontal graduated circle, made light and carried around by a mag. metic needle, so that the instrument can be used as a compass as well as a slope or angular level. It has a conrex-glass, or lens, in the smaller end, through which can be seen a hair which corers, on the circle, the number of the degrees of the angle of inclination, or of the horizontal angle.

The sights are on the top or sides, according as it is used as a compass or slope-measurer. It is used by sighting to the object, and at the same time reading off the angle, the hair covering the zero-mark when the instrument is level.
577. German Universal Instrument. Its use is to enable the observer to sight to an object nearly or quite overhead. It consists of a telescope having the part which carries the eyepiece at right angles to the part carrying the object-glass, instead of being in the same straight line, as in an ordinary tele-

Fig. 422.
 scope. The part containing the eye-piece is connected with the other part at the axis, and is in the same line with the axis.

In the telescope is placed a small mirror, or reflector, or (what is still better) a triangular prism of glass, at an angle of $45^{\circ}$ to the line of sight. Thus the observer can keep his eye at the same place at any inclination of the telescope.

## SIMPLE ANGULAR LEVELING.

## A. For Short Distances.

578. Principle. For short distances, curvature and refraction

Fig. 423.
 may be neglected. Thus, if the height of a wall, house, tree, etc., be desired, note the point where the horizontal line strikes the wall, etc., and add its height above the ground to that calculated by the formula :

$$
\mathrm{BC}=\mathrm{AC} . \text { tang. B A C. . . . [1.] }
$$

579. The "best-condition" angle for obserration is $45^{\circ}$. Hence, in setting the instrument, we should, where practicable, have the distance about equal to the height of the point whose height we wish to ascertain.

## B. For Greater Distances.

580. Correction for Curvature. A C is the line of apparent level, as given by the instrument, and $\mathrm{AC}^{\prime}$

Fig. 424.
 is the line of true level. Calling the angle $\mathrm{ACB}=90^{\circ}$ (which it is approximately for moderately great distances), formula [1] gives $B C$ as the height of $B$ above $A$. But $\mathrm{BC}^{\prime}$ is the true difference of heights of A and B .

A correction for the curvature of the earth must therefore be made. It may be done in two ways: either by calculating C C', and adding it to B C, obtained by formula [1], or by calculating the angle $\mathrm{CA} \mathrm{C}^{\prime}$, adding it to BAC , and then applying the formula [1] to the angle B A C'.
581. Correcting the Result. Expressing the distance by $k$, we have, by Art. 497 :
In feet $\mathrm{CC}^{\prime}=\frac{k^{2}}{2 \mathrm{R}}=\frac{k^{2}}{2 \times 20912405}=0.000000023909 k^{2}$.
Then, calling A C B a right angle, we have :
B $\mathrm{C}^{\prime}=k \times$ tang. B A C $+0.000000023909 k^{2}$ in feet.
The arc $\mathrm{AC}^{\prime}$ and the straight lines $\mathrm{AC}^{\prime}$ and AC are all three approximately equal.
582. Correcting the Angle. The angle $\mathrm{CAC}^{\prime}=\frac{1}{2} \mathrm{AOC}^{\prime}$, the central angle, which is measured by the $\operatorname{arc} \mathrm{AC}^{\prime}$, or $k$.

The length of the are subtending one minute

$$
=\frac{2 \pi \times 20912405}{360 \times 60}=6083 \text { feet. }
$$

Then for any arc, $k$, the angle 0 in minutes

$$
=\frac{k}{6083}=0.00016438 k ;
$$

and the angle CA C' (in minutes) $=0.000082193 \%$.
Adding this to the obserred angle, B A C, and calling $\mathrm{A}^{\prime} \mathrm{B}$ a right angle, we have, by [1]:

$$
\begin{equation*}
\mathrm{B} \mathrm{C}^{\prime}=k \text { tang. }(\mathrm{B} \mathrm{~A} \mathrm{C}+0.000082193 k) \tag{3.}
\end{equation*}
$$

583. Correction for Refraction. The effect of refraction causes the angle actually observed to be, not C A B, but C A B', which will be designated by $a^{\circ}$. For small distances, B and $\mathrm{B}^{\prime}$ sensibly coincide. The correction for refraction may be made in two ways, as for curvature.

To correct the result by finding B B'. It varies very irregularly, with wind, barometer, temperature, etc. ; but is usually taken, as an average, $\mathrm{BB}^{\prime}=0.16 \mathrm{C} \mathrm{C}^{\prime}$.

Subtracting this from the value of $\mathrm{BC}^{\prime}$, in formula [2], it becomes $\mathrm{B} \mathrm{C}^{\prime}=k$. tang. $\mathrm{B}^{\prime} \mathrm{AC}+0.000000022 k^{\circ}$.

To correct the observed angle. Subtract from it the angle $\mathrm{BAB}^{\prime}$, which is about 0.16 of the angle $\mathrm{CA} \mathrm{C}^{\prime}$.

This changes formula [3] to

$$
\begin{equation*}
\mathrm{B} \mathrm{C}^{\prime}=k \cdot \tan g .\left(\mathrm{B}^{\prime} \mathrm{A} \mathrm{C}+0.00006844 k\right) . \tag{5.}
\end{equation*}
$$

## C. For Very Great Distanoes.

584. Correction for Curvature. As before, there are two methods of making the correction.

For these distances we can not consider the angle at $\mathrm{C}^{\prime}$ a right angle. The triangle ABC gives

$$
\mathrm{B} \mathrm{C}=k \cdot \frac{\sin \cdot \mathrm{~B} \mathrm{~A} \mathrm{C}}{\sin . \mathrm{B}} .
$$

To find the angle B , we have, in the triangle BA 0 ,

$$
\begin{aligned}
& B=180^{\circ}-(0+\text { B A } 0), \\
& B=180^{\circ}-\left(0+90^{\circ}+B \text { A C }\right), \\
& B=90^{\circ}-(0+\text { A C }) ;
\end{aligned}
$$

$$
\text { Hence, } \sin . B=\cos .(0+B A C) .
$$

$$
\text { Then, B C }=k \cdot \frac{\sin . \mathrm{B} \mathrm{~A} \mathrm{C}}{\cos \cdot(0+\mathrm{B} \mathrm{~A} \mathrm{C)}} \text {, }
$$

and $\mathrm{BC}^{\prime}=\mathrm{BC}+\mathrm{CC}^{\prime}=k \cdot \frac{\sin . \mathrm{BAC}}{\cos \cdot(0+\mathrm{BAC})}+0.000000023909 k^{2}$.
$\mathrm{B} \mathrm{C}^{\prime}=k \frac{\sin . \mathrm{B} \mathrm{A} \mathrm{C}}{\cos (\mathrm{BAC}+0.0001646 k)}+0.000000023909 k^{2} \quad$. [6.]

Correcting the Angle. In the triangle A B C', getting expressions for the angles, and using the sine proportion, as before, in A B C, we have :

$$
\begin{align*}
& \mathrm{B} \mathrm{C}^{\prime}=k \cdot \frac{\sin \cdot\left(\mathrm{BAC}+\frac{1}{2} \mathrm{O}\right)}{\cos \cdot(\mathrm{BAC}+0)} . \\
& \mathrm{BC}^{\prime}=k \cdot \frac{\sin \cdot(\mathrm{BAC}+0 \cdot 000082193 k)}{\cos \cdot(\mathrm{BAC}+0.00016438 \% k)} .
\end{align*}
$$

585. Correction for Refraction. Formula [6] becomes $\mathrm{BC}^{\prime}=k \cdot \frac{\sin .\left(\mathrm{B}^{\prime} \mathrm{AC}-0.00001375 \%\right)}{\cos .\left(\mathrm{B}^{\prime} \mathrm{AC}+0.000150636 k\right)}+0.000000023909 k^{2}$.

Formula [7] becomes, diminishing BAC in both numerator and denominator by 0.08 of 0 ,

$$
\begin{equation*}
\mathrm{BC}^{\prime}=\pi \cdot \frac{\sin \cdot\left(\mathrm{B}^{\prime} \mathrm{A} C+0.000068442 k\right)}{\cos \cdot\left(\mathrm{B}^{\prime} \mathrm{A} C+0.000150636 / k\right)} . \tag{9.}
\end{equation*}
$$

586. Reciprocal Observations for canceling Refraction. Observe the reciprocal zenith-distances from each point to the other. Call these angles $\Delta$ and $\Delta^{\prime}$.

The angle ZAB is the observed zenith-distance $(\Delta)$ of $\beta$, plus the refraction $\rho-$ i. e., Z A B $=\Delta+\rho$, and $Z^{\prime} B A=\Delta^{\prime}+\rho^{\prime}$.

Let $\delta=\Delta+\rho$ and $\delta^{\prime}=\Delta^{\prime}+\rho^{\prime}$, Then $\delta+\delta^{\prime}=\Delta+\Delta^{\prime}+\rho+\rho^{\prime}=180$

$$
+0
$$

The obserrations should be simultaneous as well as reciprocal.
When this is the case, we may take $\rho=\rho^{\prime}$.

$$
\begin{gathered}
\text { Then } \rho=90+\frac{1}{2} 0-\frac{1}{2}\left(\Delta+\Delta^{\prime}\right), \\
\delta^{\prime}=\Delta^{\prime}+\rho=90+\frac{1}{2} 0+\frac{1}{2}\left(\Delta^{\prime}-\Delta\right), \\
Z_{\text {A C }}{ }^{\prime}=90+\frac{1}{2} 0 .
\end{gathered}
$$

In the triangle $\mathrm{BAC}^{\prime}, \mathrm{BC}^{\prime}: \mathrm{A}^{\prime}(=k):: \sin . \mathrm{BAC}^{\prime}: \sin . \mathrm{ABC}$.

$$
\begin{aligned}
\therefore \mathrm{BC}^{\prime} & =k \frac{\sin . \mathrm{BAC}^{\prime}}{\sin \cdot \mathrm{A} \mathrm{C} \mathrm{C}}=k \frac{\sin .\left(\mathrm{ZAB}+\mathrm{C}^{\prime} \mathrm{A} \mathrm{O}\right.}{\sin . \mathrm{Z}^{\prime} \mathrm{BA}}, \\
\mathrm{BC}^{\prime} & =k \frac{\sin \cdot\left[180^{\circ}-\frac{1}{2}\left(\Delta^{\prime}-\Delta\right)\right]}{\sin \cdot\left[90^{\circ}+\frac{1}{2} 0+\frac{1}{2}\left(\Delta^{\prime}-\Delta\right)\right]},
\end{aligned}
$$

$$
\mathrm{BC}^{\prime}=k \frac{\sin \cdot \frac{1}{2}\left(\Delta^{\prime}-\Delta\right)}{\cos \cdot \frac{1}{2}\left(\Delta^{\prime}-\Delta+0\right)}
$$

When the angle 0 is very small compared with the other angles, this becomes : $\mathrm{B}^{\prime}=k$. tan. $\frac{1}{2}\left(\Delta^{\prime}-\Delta\right)$.

Or, using angles of elevation and depression ( $\alpha$ and $\beta$ ) we have :

$$
\begin{equation*}
\mathrm{BC}^{\prime}=k \cdot \frac{\sin \cdot \frac{1}{2}(\alpha+\beta)}{\cos \cdot \frac{1}{2}(\alpha+\beta+0)} \tag{10.}
\end{equation*}
$$

Note.-Angle O, in minutes $=0.000164387 k$.
Log. $0 \cdot 000164387=\overline{4} \cdot 2158699$.
When 0 is very small, compared with the other angles, by neglecting it we have :

$$
\mathrm{BC}^{\prime}=\kappa \cdot \text { tang. } \frac{1}{2}(\alpha+\beta) . \quad . \quad . \quad . \quad[11 .]
$$

The following is from the "New York State Survey Report," 1882 :

The formula employed in deducing differences of height from reciprocal zenith-distance observations is

$$
\mathrm{H}^{\prime}-\mathrm{H}=\mathrm{K} \tan \cdot \frac{\mathrm{Z}^{\prime}-\mathrm{Z}}{2}\left(1+\frac{\mathrm{H}+\mathrm{H}^{\prime}}{2 r}\right),
$$

where $\mathrm{H}^{\prime}$ and H are the heights of the stations above sea-level, K is the distance between the stations in metres, as given by the triangulation, and consequently reduced to sea-level, $Z^{\prime}$ and $Z$ are the observed zenith-distances; $r$ is the mean radius of the earth in metres; its logarithm is 6.80454 for latitude $43^{\circ}$, according to Bessel's determination. This mean value may be safely taken as constant throughout the area of New York State without any practical error in the resulting differences of height.

The factor $\left(1+\frac{\mathrm{H}+\mathrm{H}^{\prime}}{2 r}\right)$ will never in this State affect $\mathrm{H}^{\prime}-\mathrm{H}$ by more than $\frac{1}{4000}$ part of its value; it is usual, therefore, to compute the difference of height from the formula $H^{\prime}-H=K \tan . \frac{Z^{\prime}-Z}{2}$; and if by inspection of a short table of values of the omitted factor it is seen that its effect will be appreciable, it is then introduced.

For computing differences of height from zenith-distances observed at one station only, the formula

$$
\mathrm{H}^{\prime}-\mathrm{H}=\mathrm{K} \cot . \mathrm{Z}\left(1+\frac{\mathrm{H}+\mathrm{H}^{\prime}}{2 r}\right)+\frac{1-2 m}{2 r} \mathrm{~K}^{2}
$$

is employed. The symbols here have the same significance as before, and $2 m$ is the ratio of the radius of the earth to the radius of the curve of light. The value of $m$ may be approximately determined by means of reciprocal zenith-distance observations. From 137 of such observations the State Survey has found $m=0.0730$; its value is liable to considerable fluctuation, but it may be considered constant within the hours to which the observations are confined on the survey without any material error.

The factor $\left(1+\frac{\mathrm{U}+\mathrm{H}^{\prime}}{2 r}\right)$ is treated as before. The logarithm of the coefficient $\frac{1-2 m}{2 r}$ is 2 . 82589 . The quantity $\frac{1-2 m}{2 r} \mathrm{~K}^{2}$ has been tabulated for values of K up to 18,000 metres for office use.
587. Reduction to the Summits of the Signals. Stations $a$ and $b$ can not be seen from each other.

Fig. 427.
 Signals are erected at each point, and from $a$ the angle $\mathrm{B} a \mathrm{C}=\mathrm{A}$ is obserred ; and from $b$ the angle $A b D=B$. The heights of the signals above the instrument at $a$ and $b$ are $h$ and $h^{\prime}$. The distance between the signals is $k$.

Required the reduced angles $a=c a b$ and $\beta=\mathrm{D} b a$.

$$
\left.\begin{array}{l}
\alpha=\mathrm{A}-\frac{h \cdot \cos \cdot \mathrm{~A}}{h \cdot \sin \cdot 1^{\prime \prime}} \\
\beta=\mathrm{B}+\frac{h^{\prime} \cdot \cos \cdot \mathrm{B}}{k \cdot \sin \cdot 1^{\prime \prime}} \tag{12.}
\end{array}\right\}
$$

The difference is in seconds.
Usually, in such cases, zenith-distances are taken, and the observed angles are called $\Delta$ and $\Delta^{\prime}$. The reduced angles are $\delta$ and $\delta^{\prime}$.

Draw a line in the figure from A to B . Then in the triangle $\mathrm{A} B a$ we have :

$$
\begin{gather*}
\sin . \mathrm{AB} a: \sin . \Delta:: h: k . \\
\text { or, } \sin . \mathrm{AB} a=\frac{h \sin \cdot \Delta}{k \sin \cdot 1^{n}}, \\
\operatorname{and} a \mathrm{~B} \mathrm{~A}=\frac{h \sin \cdot \Delta^{\prime}}{k \sin \cdot 1^{\prime \prime}} \\
\therefore \delta=\Delta+\frac{h \cdot \sin . \Delta}{k \cdot \sin \cdot 1^{n}} \text {, and } \delta^{\prime}=\Delta^{\prime}+\frac{h^{\prime} \cdot \sin \cdot \Delta^{\prime}}{k \cdot \sin \cdot 1^{n}} . \tag{13.}
\end{gather*}
$$

The difference is seconds.
Instead of $h$ and $h^{\prime}$, some writers use $d \mathrm{H}$ and $d \mathrm{H}^{\prime}$; or $d \mathrm{~A}$ and $d \mathrm{~A}^{\prime}$, meaning difference of height, and difference of altitude.

For great exactness, instead of using the mean radius of the earth to get 0 , the radius at the point of obserration is used.
588. When the height of the signal above the instrument can not be measured, if the signal be conical, like a spire, etc., to find $\mathrm{B}^{\prime}$ we measure two diameters, 2 R and $2 r$, and the distance apart, $h$.
Then, $\mathrm{B} \mathrm{B}^{\prime}=\frac{\mathrm{R} h}{\mathrm{R}-r} . \quad$ [14.]

Fig. 428.


If the oblique distance
$l$ be measured instead of $h$, then

$$
\mathrm{B} \mathrm{~B}^{\prime}=\frac{\mathrm{R}}{\mathrm{R}-r} \sqrt{ }[l+(\mathrm{R}-r)][l-(\mathrm{R}-r)] .
$$

When a Spire is very
Fig. 429.


$$
B B^{\prime}=\frac{K \cdot \tan \cdot\left(\delta^{\prime \prime}-\delta\right)}{\cos \cdot \frac{1}{2}\left(\Delta^{\prime}-\Delta+0\right)} .
$$

589. Leveling by the Horizon of the Sea. Owing to refraction, the apparent zenith-distance will be Z B A '.

Let $\mathrm{R}=$ radius of the earth ; $\mathrm{H}^{\prime}=$ horizon.

$$
\text { Then } R+B B^{\prime}=\frac{R}{\cos . C} .
$$

$$
\therefore \mathrm{BB}^{\prime}=\mathrm{R} \frac{(1-\cos . \mathrm{C})}{\cos . \mathrm{C}} \cdot[16 .]
$$

Now, $(1-\cos . C)=2 \sin .^{2} \frac{1}{2}$ C. Transposing, we have $\cos$. C

$=\cos ^{2} \frac{1}{2} \mathrm{C}-\sin ^{2} . \frac{1}{2} \mathrm{C}$. Substituting these values in equation (1), we get $B B^{\prime}=\frac{R\left(2 \sin .^{2} \frac{1}{2} C\right)}{\operatorname{cos.}^{2} \frac{1}{2} C-\sin .^{2} \frac{1}{2} C}=2 R \frac{\sin \cdot{ }^{2} \frac{1}{2} C}{\cos ^{2} \frac{1}{2} C-\sin .^{2} \frac{1}{2} C}$.
(Developing by the binomial formula) -

$$
=2 \mathrm{R} \tan ^{2}{ }^{2} \frac{1}{2} \mathrm{C}\left(1+\tan ^{2}{ }^{\frac{1}{2}} \mathrm{C}-\tan ^{4} \frac{1}{2} \mathrm{C}+\text {, etc. }\right)
$$

Using the first two terms of the series, we have

$$
\mathrm{B}^{\prime}=2 \mathrm{R} \tan ^{2} \frac{1}{2} \mathrm{C}\left(1+\tan ^{2} \frac{1}{2} \mathrm{C}\right) .
$$

As the angle C is rery small, we may express the tangent as an are in terms of the radius, without greater error than one foot in an altitude of 45,000 .

Then we have $\mathrm{BB}^{\prime}=\frac{\mathrm{R}}{2} \mathrm{C}^{2}\left(1+\frac{\mathrm{C}^{2}}{4}\right)$. [1\%.]
The angle $\mathrm{C}=\mathrm{HB} \mathrm{A}^{\prime}+\mathrm{ABA}=\delta-90^{\circ}-n \mathrm{C}, n$ being the coefficient of refraction. $\quad \therefore \mathrm{C}=\frac{\delta-90^{\circ}}{1-n}$.

In order to introduce the value of C into equation (2), we multiply it by the sine of $1^{\prime \prime}$, to reduce are to linear measure.

Then we have
$B B^{\prime}=\frac{1}{2} R\left(\frac{\sin .1^{\prime \prime}}{1-n}\right)^{2}\left(\delta-90^{\circ}\right)^{2}\left\{1+\frac{1}{4}\left(\frac{\sin .1^{\prime \prime}}{1-n}\right)^{2}\left(\delta-90^{\circ}\right)\right\} .[18$.

## COMPOUND ANGULAR LEVELING.

590. The following problems may mostly be reduced to a combination of : first, determining the inaccessible distance to a point immediately under (or over) the point whose height is desired, and then using this distance to obtain that height.
591. By Angular Co ordinates in one Plane. Take two

Fig. 431.
 stations, A and D , in the same vertical plane with B . At A observe the angles of elevation of B and D . Measure A D. At D observe the angle ADB . Then, in the triangle A B D we get AB, and in the triangle BAC we get BC .

$$
\begin{equation*}
\mathrm{BC}=\mathrm{A} \mathrm{D} \cdot \frac{\sin \cdot \mathrm{BDA} \cdot \sin . \mathrm{B} \mathrm{~A} \mathrm{C}}{\sin \cdot \mathrm{~A} \mathrm{D} \mathrm{D}} . \tag{19.}
\end{equation*}
$$

For great distances, the corrections for curvature and refraction are to be made as in the preceding articles.

If AD be horizontal, the same formula ap-

Fig. 432.
 plies; but there is one angle less to measure, since $\mathrm{BAC}=\mathrm{BAD}$. Formula [19] gives the height of B above A .

If the height of B above D, in Fig. 432, be desired, find BD in the triangle BAD , observe the angle of elevation of B from D , and then the desired height equals
B D . sin. B D E.

Otherwise, find height of D above A , and subtract it from B C.
592. By Angular Co-ordinates in Several Planes. On irregular ground, when the distance between the two points is unknown, the
operations for finding it by the various methods already given may be combined with the observation of vertical angles, thus :

Fig. 433.


At A measure the vertical angle of elevation, B A C. Also measure the horizontal angle, CAD , to some point, D , and measure horizontally the distance, AD. At D measure the horizontal angle, AD C. Then,

$$
\begin{align*}
& \mathrm{AC}=\mathrm{AD} \frac{\sin \cdot \mathrm{ADC}}{\sin \cdot \mathrm{ACD}} \cdot \quad \mathrm{BC}=\mathrm{AC} \cdot \operatorname{tang} \cdot \mathrm{BAC} . \\
& \mathrm{BC}=\mathrm{AD} \frac{\sin \cdot \mathrm{ADC} \cdot \tan \cdot \mathrm{BAC}}{\sin \cdot \mathrm{ACD}} . . . . . \tag{20.}
\end{align*}
$$

593. Conversely. The distance may be obtained when the height is known.

Let C B be a known height. Then, AC = C B.tan. ABC. BC is a known height, and DE an inaccessible line in the same

Fig. 434.


Fig. 435.

horizontal plane as C. Find CD and C E by the last method, and measure the horizontal angle ECD subtended at C by ED.

Then the two sides and the included angle of a triangle are known, to find the third side.

## CHAPTER III.

## BAROMETRIC LEVELING.

## PRINCIPLES AND FORMULAS.

594. Principles. The difference of the heights of two places may be determined by finding the difference of their depths below the top of the atmosphere in the same way as the comparative heights of ground under water are determined by the difference of the depths below the top of the water. The desired height of the atmosphere above any point, such as the top of a mountain, or the bottom of a valley, is determined by weighing it. This is done by trying how high a column of mercury or other liquid the column of air above it will balance ; or what pressure it will exert against an elastic box containing a vacuum, etc. Such instruments are called Barometers.
595. Applications. Since the column of mercury in the barometer is supported by the column of air above it, the mercury sinks when the barometer is carried higher, and vice versa.

The weight of any portion of air decreases from the surface of the earth to the assumed surface of the atmosphere. It has been found that, as the heights to which the barometer is carried increase in arithmetical progression, the weights of the column of air above the barometer, and consequently its readings, decrease in geometrical progression. Consequently, the difference of the heights of any two not very distant points on the earth's surface is proportional to the difference of the logarithms of the readings of the barometer at those points-i. e., equal to this latter difference multiplied by some constant coefficient. This is found by experiment to be 60159 , at the freezing-point, or temperature of $32^{\circ}$ Fahr., the
readings of the mercury being in inches, and the product, which is the difference of height, being in feet.

Several corrections are necessary.
596. Correction for Temperature of the Mercury. If the temperature of the mercury be different at the two stations, it is expanded at the one, and contracted at the other, to a height different from that which is due to the mere weight of the air above it.

Mercury expands about $\frac{1}{10000}$ of its bulk for each degree of $F$. Therefore, this fraction of the height of the column is to be added to the height of the colder column, or subtracted from the height of the warmer one, in order to reduce them to the same standard. A thermometer is therefore attached to the instrument in such a manner as to give the temperature of the mercury.

If a brass scale is used, the correction is $\frac{9}{100000}$ for each degree F .

59\%. Correction for Temperature of the Air. The warmer the air is, the lighter it is ; so that a column of warm air of any height will weigh less than when it is colder. Consequently, the mercury in warm air falls less in ascending any height, and is higher at the place than it otherwise would be. Hence the height giren by the preceding approximate result will be too small, and must be increased by $\frac{1}{491}$ part for each degree $F$. that the temperature of the air is above $32^{\circ}$. The effect of moisture in the air changes this fraction to $\frac{1}{450}$.
598. Other Corrections. For very great accuracy, we should allow for the variation of grarity, corresponding to the rariation of latitude on either side of the mean. So, too, we should allow for the decrease of gravity corresponding to any increase of height of the place.
599. Rules for calculating Heights by the Mercurial Barometer.

1. At each station read the barometer ; note its temperature by the attached thermometer, and note the temperature of the air by a detached thermometer.
2. Multiply the height of the upper column by the difference
of readings of the attached thermometer, and that by $\frac{99}{100000}$, and add the product to the upper column, if that be the colder, or subtract it, if that be the warmer. This gives the corrected height of the mercury.
3. Multiply the difference of the logarithms of the corrected heights of the mercury-i.e., the corrected upper one and the lower one-by 60159, and the product is the approximate difference of heights of the places in feet for the temperature of $32^{\circ}$.
4. Subtract $32^{\circ}$ from the arithmetical mean of the temperatures of the detached thermometer ; multiply the approximate altitude by this difference; divide the product by 450 ; add the quotient to the approximate altitude, and the sum is the corrected altitude.
5. Formulas. The rules just given are best expressed in formulas, thus :

|  | at lower station. | at upper station. |
| :---: | :---: | :---: |
| Height of mercury | H | $h^{\prime}$ |
| Temperature of mercury | T | T' |
| Temperature of air. . | $t$ | $t^{\prime}$ |

Calling the reduced height of mercury at the upper station $h$, we have, by Rule 2 :

$$
\begin{equation*}
\hbar=\hbar^{\prime}+0.00009\left(\mathrm{~T}-\mathrm{T}^{\prime}\right) h^{\prime} \tag{1.}
\end{equation*}
$$

N. B. -If $T^{\prime}$ is more than $T$, the product will be subtractive.

Then, by Rule 3, we have :
Approximate height $=60159$ (log. $\mathrm{H}-\log . ~ h)$.
By Rule 4, the correction for temperature of air

$$
=\text { approximate height } \times \frac{t+t^{\prime}-64}{900}
$$

Adding this correction to the approximate height, and factoring the sum, we get:
Corrected ht. $=60159$ (log. $\mathrm{H}-\log . h)\left(1+\frac{t+t^{\prime}-64}{900}\right)$.
601. To correct for Latitude. Multiply the preceding result by 0.00265 . cos. 2 L ( L being the latitude), and add (algebraically) the product to the preceding result.

At $45^{\circ}$, correction is zero. At equator it is +0.00265 . At pole it is -0.00265 .

To correct for Elevation of the Place. Call the last corrected height $x^{\prime}$, and the height of the lower place above the level of the sea, S , and add to $x^{\prime}$ this quantity :

$$
\frac{x^{\prime}+52251}{20912405}+\frac{\mathrm{S}}{10456203} .
$$

602. Final English Formula. Combining the previous results into one formula, we get :
$H \mathrm{t} .=60159(\log . \mathrm{H}-\log . h)\left\{\begin{array}{l}\left(1+\frac{t+t^{\prime}-64}{900}\right), \\ (1+0.00265 . \cos .2 \mathrm{~L}), \\ \left(1+\frac{x^{\prime}+52251}{20912405}+\frac{\mathrm{S}}{10456203}\right)\end{array}\right\}[$ [3.]
In this formula, the three quantities under each other are three factors.

Usually, only the first factor is required, and then we have formula [2]. Using the second, also, we correct for latitude ; and, using the third, for the elevation.
603. French Formulas. French barometers are graduated in French millimetres, each $=0.03937$ inch, and the thermometer is centigrade, in which the freezing-point is zero, and boiling-point $100^{\circ}$ :

$$
a^{\circ} \text { cent. }=\left(\frac{9}{5} a+32\right)^{\circ} \mathrm{F} .
$$

Then, the French formula corresponding to [3] is the following ( H and $h^{\prime}$ being in millimetres, and the temperatures centigrade) :

$$
h=h^{\prime}\left(1+\frac{T-T^{\prime}}{6200}\right)
$$

And the difference of heights in metres

$$
=18336(\log . \mathrm{H}-\log \cdot h)\left\{\begin{array}{l}
\left(1+\frac{2\left(t+t^{\prime}\right)}{1000}\right), \\
(1+0 \cdot 00265 \cdot \cos .2 \mathrm{~L}), \\
\left(\frac{1+x^{\prime}+15926}{63 \cdot 2481}\right)+\frac{\mathrm{S}}{3186241}
\end{array}\right\} \text { [4.] }
$$

604. Babinet's Simplified Formula, without Logarithms.
$h^{\prime}$ is reduced to $h$, as before, viz. : $h=h^{\prime}\left(1+\frac{\mathrm{T}-\mathrm{T}}{6200}{ }_{\prime}\right)$.
Then, the difference of heights in metres

$$
\begin{equation*}
=16000 \cdot \frac{\mathrm{H}-\hbar}{\overline{\mathrm{H}}+h}\left(1+\frac{2\left(t+t^{\prime}\right)}{1000}\right) . \tag{5.}
\end{equation*}
$$

The heights are in millimetres and the temperatures centigrade.

$$
\begin{aligned}
& \text { Example. } \mathrm{H}=755 . h=745 \\
& t=15^{\circ} t^{\prime}=10^{\circ} . \\
& \mathrm{Ht} .=16000 \frac{10}{1500}\left(1+\frac{50}{1000}\right)=112 \mathrm{~m} .
\end{aligned}
$$

Correct result is 111.6 m .
This formula is a very near approximation for moderate heights.
Babinet's formula in English measures (the heights being in inches, and temperatures Fahrenheit) is in feet:

$$
\begin{equation*}
52494\left(\frac{\mathrm{H}-h}{\mathrm{H}+h}\right) \quad\left(1+\frac{t+t^{\prime}-64}{900}\right) . \tag{6.}
\end{equation*}
$$

Leslie's formula is :

$$
\text { height in feet }=55000 \frac{\mathrm{~B}-b}{\mathrm{~B}+b} \text {. }
$$

In which $B=$ upper reading, and $b=$ lower reading. This is for a temperature of $55^{\circ}$ Fahr.
605. Tables. These shorten the operations greatly. The most portalle are in "Annuaire du Bureau des Longitudes." The most complete are Professor Guyot's, published by the Smithsonian Institution at Washington.
606. Approximations. One tenth of an inch difference of readings in two places corresponds to about ninety feet difference of elevation. One millimetre difference of readings corresponds to about ten and a half metres difference of height, or about thirtyfour feet.

This is correct near the freezing-point, and near the level of the sea. The height corresponding to any given difference of readings increases, however, with the temperature and with the height of the station. Thus, at $70^{\circ} \mathrm{F}$., $\frac{1}{10}$ of an inch corresponds to an ele-
vation of 95 feet ; and one millimetre at $30^{\circ}$ cent. corresponds to $11 \frac{3}{4}$ metres, or about 40 feet.

## Instruments.

607. Barometers made for leveling are called Mountain Ba-

Fig. 436.
 rometers. They are either cistern barometers or siphon barometers.

Fig. 436 is a cistern barometer.* This consists of a column of mercury, contained in a glass tube, whose lower end is placed in a cistern of mercury. The tube is corered with a brass case, terminating at the top in a ring, A, for suspension, and at the bottom in a flange, B , to which the cistern is attached.

At C is a rernier, by which the height of the mercury is read off. The vernier is mored by means of a rack, worked by the milled head shown at D.

The zero of the scale is a small irory point, shown below the flange B. The mercury in the cistern is raised or lowered, by means of the milled-headed screw 0 , till its surface is just in contact with the irory point. The upper part of the cistern is of glass, so that the surface of the mercury in the cistern, and the irory point, may be readily seen. At E is the attached thermometer which indicates the temperature of the mercury. When it is carried, the mercury is screwed up to prevent breaking the glass.
608. The Aneroid Barometer. This is a thin box of corrugated copper, exhausted of air. When the air grows hearier, the box is compressed ; and when the air grows lighter, it is expanded by a spring inside. This motion is communicated

[^63]by suitable levers to the index-hand, on the face, which indicates the pressure of the atmosphere, the face being graduated to correspond with a common barometer.

There are several varieties of this instrument, differing principally in the method of determining the movement of the corrugated box due to changes in the density of the atmosphere.

They are made in
 sizes varying from two to six inches in diameter. They are much used on account of their portability, but are not as reliable as the mercurial barometer.

Approximately, a difference of reading of $\frac{1}{100}$ of an inch corresponds to a difference of height of nine feet. The following table is more nearly accurate :

| mean temperature. | $30^{\circ}$ | $40^{\circ}$ | $50^{\circ}$ | $60^{\circ}$ | $70^{\circ}$ | $80^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Mean pressure, 27 inches. | $9 \cdot 7$ | $9 \cdot 9$ | $10 \cdot 1$ | $10 \cdot 3$ | 10.5 | 10.8 |
| 28 " | $9 \cdot 3$ | $9 \cdot 5$ | $9 \cdot 8$ | $10 \cdot 0$ | $10 \cdot 2$ | $10 \cdot 4$ |
| 29 " | $9 \cdot 0$ | $9 \cdot 2$ | $9 \cdot 4$ | $9 \cdot 6$ | $9 \cdot 8$ | 10. |
| " 30 " | $8 \cdot 7$ | $8 \cdot 9$ | $9 \cdot 1$ | $9 \cdot 3$ | $9 \cdot 5$ | $9 \cdot 7$ |

609. The Hypsometer. The temperature at which water boils varies with the pressure of the atmosphere, and therefore decreases in ascending heights. Then a thermometer becomes a substitute for a barometer.

Approximately, each degree of difference (Fahr.) corresponds to about 5 5ั0 feet difference of

| TEMPERATURE OF BOILING WATER. | CORRESPONDING BAROMETER READINGS |
| :---: | :---: |
| $213^{\circ}$ | $30^{\prime \prime} .522$ |
| $212{ }^{\circ}$ | $29^{\prime \prime} \cdot 922$ |
| $211{ }^{\circ}$ | $29^{\prime \prime} \cdot 331$ |
| $210^{\circ}$ | $28^{\prime \prime} \cdot 751$ |
| $209^{\circ}$ | $28^{\prime \prime} \cdot 180$ |
| $208^{\circ}$ | $27^{\prime \prime} \cdot 618$ | elevation, subject to the usual barometric corrections for the temperature of the air. (For minute tables, see Guyot's.)

610. Accuracy of Barometric Observations. This increases with the number of repetitions of them, the mean being taken. With great skill and experience they may be depended upon to a very few feet.
611. The observations at the two places, whose difference of heights is to

PROFESSOR GUYOT"S RESULTS.

| HEIGHTS FOUND BY THE BAROMETER. | CORRESPONDING HEIGHTS FOUND BY THE SPIRIT-LEVEL. |
| :---: | :---: |
| 6707 feet. 2752 " 6291 | $\begin{aligned} & 6711 \text { feet. } \\ & 2752 \text { "6 } \\ & \{285 \\ & 6293 \end{aligned}$ | be determined, should be taken simultaneously at a series of intervals preriously agreed upon, the distance apart of the places being as short as possible. Distant places should be connected by a series of intermediate ones.

## PART III.

## TOPOGRAPHY.

## INTRODUCTION.

612. Definition. Topography is the complete determination and representation of any portion of the surface of the earth, embracing the relative position and heights of its inequalities; its hills and hollows, its ridges and valleys, level plains, slopes, etc., telling precisely where any point is, and how high it is.

It therefore determines the three co-ordinates of any point; the horizontal ones by surveying, and the vertical ones by leveling.

The results of these determinations are represented in a conventional manner, which is called "topographical mapping."

The difficulty is, that we see hills and hollows in elevation, while we have to represent them in plan.
613. Systems. Hills are represented by various systems :

1. By level contour-lines, or horizontal sections.
2. By lines of greatest slope, perpendicular to the former.
3. By shades from vertical light.
4. By shades from oblique light.

The most usual method is a combination of the first, second, and third systems.

## CHAPTER I.

## FIRST SYSTEM.

## BY HORIZONTAL CONTOUR-IINES.

614. General Ideas. Imagine a hill to be sliced off by a number of equidistant horizontal planes, and their intersections with it

Fig. 438.
 to be drawn as they would be seen from above, or horizontally projected on the map, as in Fig. 438. These are "contour-lines."

They are the same lines as would be formed by water surrounding the hill, and rising one foot (or any other height) at a time till it reached the top of the hill. The edge of the water, or its shore, at each successive rise, would be one of these horizontal contour-lines. It is plain that their nearness or distance on the map would indicate the steepness or gentleness of the slopes. A

Fig. 439.


Fig. 440.


Fig. 441.

right cone would thus be represented by a series of concentric circles, as in Fig. 439 ; an oblique cone, by circles not concentric, but nearer to each other on the steep side than on the other, as in Fig. 440 ; and by a half-egg, somewhat as in Fig. 441.
615. Plane of Reference. The horizontal plane on which the contour-lines are projected, and to which they are referred, is called the "plane of reference." This plane may be assumed in any position, and the distance of the contour-lines above or below it is noted on them. It is usually best to assume the position of the plane of reference lower than any point to be represented ; so that all the contour-lines will be above it, and none of them have minus signs.
616. Vertical Distances of the Horizontal Sections. These depend on the object of the survey, the population of the country, the irregularity of the surface, and the scale of the map. In mountainous districts they may be 100 feet apart. On the United States Coast Survey they are twenty feet; for engineering purposes, five feet, or less. One rule is to make the distance in feet equal to the denominator of the ratio of the scale of the map, divided by 600 .
617. Methods for determining Contour-Lines. They are of two classes: 1. Determining them on the ground at once ; 2. Determining the highest and lowest points, and thence deducing the contour-lines.

## First Method.

618. General Method. Determine one point at the desired height of one line, and then "locate" a line at that level.

The "reflected hand-level," or "reflecting-level," or "waterlevel," are sufficiently accurate between " bench-marks" not very distant.

One such line having been determined, a point in the next higher or the next lower one is fixed, and the preceding operations repeated.
619. On a Long, Narrow Strip of Ground, such as that required for locating a road: Run a section across it at every quarter or half mile, about in the line of greatest slope. Set stakes on these sections at the heights of the desired contour-lines, and then get inter-

Fig. 442.

mediate points at these heights between the stakes. These sections check the levels.
620. On a Broad Surface. Level around it setting-stakes, at points of the desired height, and then run sections across it, and from them obtain the contour-lines as before.

The external lines here serve as checks to the cross-lines.
621. Surveying the Contour-Lines. The contour-lines thus found may be surresed by any method. If they are long, and not very much curred, the compass and chain and the method of "progression" is best. If they are curred irregularly, the method of radiation is best. When straight lines exist among them, such as fences, etc., or can conreniently be established, then rectangular co-ordinates are most convenient.

## Second Method.

622. General Nature. This method consists in determining the heights and positions of the principal points, where the surface of the ground changes its slope in degree or in direction-i. e., determining all the highest and lowest points and lines, the tops of the
hills and bottoms of the hollows, ridges and valleys, etc., and then, by proportion or interpolation, obtaining the places of the points which are at the same desired level. The heights of the principal points are found by common leveling, and their places fixed as in Art. 621.

The first method is more accurate ; the second is more rapid.
623. Irregular Ground. When the ground has no very marked features, run lines across it in various directions, and level along them, taking heights at each change of slope, just as in taking sections for profiles.

Otherwise, thus: Set stakes on four sides of the field, so as to inclose it in a rectangle, if possible, as in Fig. 443. Place the stakes equidistant, so that the imaginary visual lines connecting them would divide the surface into rectangles. Send the rod along one of these lines till it gets in the range of a cross-one, and observe to it there. Put down the observed heights of these points at the corresponding points on the plat,

Fig. 443.
 on which these lines have been drawn. The contour-lines are determined as in Art. 626.
624. On a Single Hill. Proceed thus : From its top, range lines down the hill, in various directions, and take their bearings. Set stakes on them at each change of slope, and note the heights and distances of these stakes from the starting-point, and plat their places. The contour-lines are then put in as in Art. 626.

With a transit, the heights of the points could be determined by vertical angles ; and also their distances with stadia-hairs, their directions being given by the horizontal circle of the transit. The French use for this purpose a "leveling-compass."
625. For an Extensive Topographical Survey. Proceed thus : Set up and get the height of the cross-hairs from some bench-mark, and get the heights of high and low prominent points all around. Then go beyond these points and set up again. Sight to one of these known points as a "turning-point," and get the heights of all the points now in sight, as before. Then go beyond these again, and so on. The places of these new points are fixed as before.
626. Interpolation. The heights and the places of the principal points being determined, by either of the preceding methods, points of any intermediate height, corresponding to any desired contour-curre, are obtained by proportion.

If, in Fig. 444, the heights of the intersection of the lines being found, as in Art. 623, and their distance apart being
 100 feet, it is required to construct contour-curres whose difference of heights is 5 feet: Taking, for example, the one whose height is 45 feet, we see it must fall between the points $A$ and $B$, whose heights are 50 feet and 35 feet : and its distance from A will be found by the proportion, as 15 is to 5 so is 100 to the required distance. So on for any number of points. To save the labor of continually calculating the fourth proportional, a scale of proportion may be constricted.
627. Interpolating with the Sector. This is one of the easiest ways. The problem is: having given on a plat two points of known height, to interpolate between them a point of any desired intermediate height.

Take in the dividers the distance betreen the giren points on
the plat ; open the sector so that this distance shall just reach between numbers, on the scale marked L, corresponding to the difference of the heights of the two given points-i. e., from 6 to 6 , or ${ }^{7}$ to $\%$, and so on. The sector is then set for all the interpolations between these two points.

Then note the difference of height between the desired point and one of the given points, and extend the dividers between the corresponding numbers on the scale. This opening will be the distance to be set off on the plat from the given point to the desired point.

628. Ridges and Thalwegs. The general character of the surface of a country is given by two sets of lines: the ridge-lines, or water-shed lines; and the "thalweys," or "lowest lines of valleys."

The former are lines which divide the water falling upon them, and from which it passes off on contrary sides. They are the lines of least slope when looking along them from above downward ; and they are the lines of greatest slope when looking from below upward. They can therefore be readily determined by the slope-level, etc. They are the lines of least zenith-distances when viewed from either direction.

On these lines are found all the projecting or protruding bends of the contour-lines, convex toward the lower ground, as shown in Fig. 396.

The second set of lines, or the "thalwegs," are the converse of the former. They are indicated by the water-courses which follow them or occupy them. They are the lines of greatest slope when looked at from above, and of least slope when looked at from below. They are the lines of greatest zenith-distance when viewed from either direction.

On these lines are the receding or re-entering points of the contour-curves, concave toward the lower ground.

The general system of the surface of a country is most easily characterized by putting down these two sets of lines, and marking the changes of slope,

Fig. 446.
 especially the beginning and the end.

The most important points to be determined are :

1. At the top and bottom of slopes.
2. At the changes of slopes in degree.
3. On the water-shed lines, and on the thalwegs.
4. On "cols," or culminating points of passes.
5. Forms of Ground. It will be found, on the inspection of a "contour-map" (which shows ground much more plainly to the eye than does the ground itself), that its infinite variety of form may, for the purposes of the engineer, be reduced to five :
6. Sloping down on all sides-i. e., a hill (Fig. 44\%).

Fig. 447.


Fig. 448.

2. Sloping up on all sides-i. e., a hollow (Fig. 448).
3. Sloping down on three sides and up on one-i. e., a croupe,

Fig. 449.


Fig. 450 Fig. 451.

or shoulder, or promontory, the end of a ridge or water-shed line (Fig. 449).
4. Sloping up on three sides and down on one-i. e., a valley, or thalweg (Fig. 450).
5. Sloping up on two sides and sloping down on two, alter-nately-i. e., a "pas," or "col," or " saddle" (Fig. 451).
[Note.--The arrows in the figures indicate the direction in which water would run.]
630. Sketching Ground by Contours. A valuable guide is, the observation that the contour-lines are perpendicular to the watershed lines and thalwegs. Note especially the contour-lines at the bottoms of hills and ridges, and at the tops of hollows and valleys, putting them down, in their true relative positions and distances, to an estimated scale.

On a long slope or hill, draw first the bottom contour-line, and the top one ; then the middle one; and afterward interpolate others. Remember that two of them can never meet, except on a perpendicular face ; and that, if one of them passes entirely around a hill or hollow, it will come back to its starting-point. Hold the field-book so that the lines on it have their true direction. As far as possible, all of the work should be done in the field with the ground in sight, and not trust to finishing from memory.
631. Ambiguity. In contour-maps of ground, if the heights of the contour-lines are not written upon them, it may be doubtful which are the highest and lowest ; which are ridges and which valleys, etc.

1. Numbers remove this.
2. The water-courses show the slopes. If there are none, put some in, in the thalwegs of a rough sketch.
3. Put hatchings on the lower sides of the contour-lines, as if water were draining off.
4. Tint the valleys and low places.
5. Conventionalities. Sometimes the spaces between contourlines are colored with tints of Indian-ink, sepia, etc., increasing in darkness as the depth increases.

Ground under water is commonly so represented, beginning at the low-water line and covering the space to the six-feet-deep con-tour-line with a dark shade of Indian-ink; then a lighter shade from 6 to 12 ; a still lighter from 12 to 18 ; and the lightest from 18 to 24.

Greater depths are noted in fathoms and fractions.
633. Applications of Contour-Lines. They have many important uses besides their representation of ground:

1. To obtain vertical sections-i. e., profiles.
2. To obtain oblique sections.
3. To locate roads.
4. To calculate excaration and embankment. Consider the contour-lines to represent sections of the mass by horizontal planes. Then each slice between them will have its contents equal, approximately, to half the sum of its upper and lower surfaces multiplied by the vertical distance apart of the sections. This method is used to get the cubic contents of a hill to be cut away ; of a hollow to be filled up ; of a great reservoir in a valley, either only projected, or full of water, etc.
5. Sections by Oblique Planes. This method was much used by the old military topographers. It is picturesque, but not precise. The cutting-planes are parallel, and may make any angle with the horizon.

## CHAPTER IT.

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SECOND SYSTEM.
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## BY LINES OF GREATEST SLOPE.

635. Their Direction. It is that which water would take in running down a slope. They are drawn perpendicularly to the contour-lines, and are the "lines of greatest slope." They are called " hatchings."

Fig. 452 represents an oval hill by this system.
636. Sketching Ground by this System. This is rapid and effective, but not precise. In doing this,

Fig. 452.
 hold the book to correspond with your position on the ground, and always draw toward you. If at the top of a hill, begin by drawing lines from the bottom, and vice versa. The hatchings are guided by contourlines lightly sketched in.
637. Details of Hatchings. They must be drawn very truly perpendicular to the contour-lines. But if the contour-lines are not parallel, the hatchings must curve. In finishing drawings, sketch in the curved hatchings with a pencil at some distance apart as guides. When the contours are very far apart, as on nearly level ground, pencil in intermediate ones.

Hatchings in adjoining rows should not be continuous, but
"break joints," to indicate the places of the contour-lines, which are usually penciled in to guide the hatchings, and then rubbed ont. The rows of hatchings must neither orerlap nor separate, and the lines should be made slightly tremulous. When they are put in without contour-lines to guide them, take care nerer to let two rows run into one ; for the breaks between the rows represent con-tour-lines, and two contour-lines of different heights can never meet except on a rertical surface.

In drawing a hill begin at the top. When hatchings diverge very much, as on hill-tops, put in alternate short ones. When the formation is very convex or concare, short auxiliary contours may be used.

## CHAPTER III.

## THIRD SYSTEM.

## BY SHADES FROM VERTICAL LIGHT.

638. Degree of Shade. The steeper the slope is, the less light it receives, in the inverse ratio of its length-i. e., inversely as the secant of the angle $a$ which it makes with the horizon, or directly as $\cos . a$. Then the ratio of the black to the white is,

$$
:: 1-\cos . a: \cos . a
$$

In practice, the difference of shade is much exaggerated.

Tables have been prepared by various nations, establishing the ratio of black and white.

Fig. 453.


The proper degree of shade may be given to the hills and hollows on the map by various means.
639. Shades by Tints. Indian-ink, or sepia, is used. The shades are put on with proper darkness, according to a previously prepared "diapason of tints." The tints are made light for gentle slopes, and dark for steep slopes, in a constant ratio, a slope of $60^{\circ}$ being quite black, one of $30^{\circ}$ a tint midway between that and white, and so on. The edges at the top and bottom are softened off with a clean brush. This is rapid and effective, but not very definite or precise, except in combination with contour-lines.
640. Shades by Contour-Lines. This is done by making the con-tour-lines more numerous-i. e., interpolating new ones between
those first determined. One objection to this is confusion of these lines with roads.
641. Shades by Lines of Greatest Slope. The lines of steepest slope-i. e., the hatchings between the contours - have their thickness and distance apart made proportional to the steepness of the slope, in some definite ratio. This is the most usual method.

The tints may be produced by varying the thickness of the hatchings, or their distance apart. Both are usually combined.
642. The French Method. In this the degree of inclination is indicated by varying the distances between the centers of the hatchings. The rule is : the distance between the centers of the lines shall equal $\frac{2}{100}$ of an inch, plus $\frac{1}{4}$ of the denominator of the fraction denoting the declivity-i. e., tangent of the angle made by the surface of the ground with the plane of reference-expressed in hun-. dredths of an inch.

The lines are made hearier as the slope is steeper, being fine for the most gentle slopes, and increasing in breadth till the blank space between them equals $\frac{1}{2}$ the breadth of the lines.

Only slopes of from $\frac{1}{1}$ to $\frac{1}{6 \pm}$ inclusive are represented by this method.
643. The German, or Lehmann's Method. He uses nine grades for slopes from $0^{\circ}$ to $45^{\circ}$, the first being white and the last black.

Fig. 454.


For the intermediate slopes, he makes the white to the black in the following proportion:

The white : the black : : 45 ${ }^{\circ}$ - angle of slope : angle of slope For example, for $30^{\circ}$ :

$$
\text { light : shade : : } 45^{\circ}-30^{\circ}: 30^{\circ}:: 1: 2 .
$$

Hence, the space between the strokes is to their thickness, as $45^{\circ}$ minus the angle of the slope is to the angle of the slope. Slopes

Fig. 455.

steeper than $45^{\circ}$ are represented by short, heavy lines, parallel to the contour-lines, as shown in the upper right-hand corner of Fig. 455-a hill drawn by Lehmann's method.
644. Another Diapason of Tints:

| Slope...... | $2 \frac{1^{\circ}}{2}$ | $5^{\circ}$ | $10^{\circ}$ | $15^{\circ}$ | $25^{\circ}$ | $35^{\circ}$ | $45^{\circ}$ | $60^{\circ}$ | $75^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Black $\ldots .$. | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 |
| White $\ldots .$. | 10 | 9 | 8 | 7 | 6 | 5 | 4 | 3 | 2 |

This distinguishes gentle slopes better. It makes them darker, and the steeper slopes lighter, and provides for slopes beyond $45^{\circ}$. To use this standard, make it on the edge of a strip of paper, and apply that to the map in various parts, and draw a few lines corresponding to the slope of those parts ; then fill up the intervening portions with suitable gradations. The angle of the slope is known from the map, since its tangent equals the rertical distance between the contours, divided by the horizontal distance. A scale can be made for any given vertical distance.

## FOURTH SYSTEM.

## BY SHADES PRODUCED BY OBLIQUE LIGHT.

645. Light is supposed to fall from the upper left-hand corner, as in drawing an "elevation," although the map is in plan. Then slopes facing the light will have a light tint, and those on the opposite side a dark tint.

This is picturesque, but not precise. It gives apparent "relief " to the ground, but does not show the degree of steepness.

The shades may be produced, as in the last method, by any means-tints, contours, or hatchings.

By making a map with contour-lines, and shaded obliquely, it will be both effective and precise.

## CHAPTER IV.

## CONVENTIONAL SIGNS。

646. Signs for Natural Surface. Sand is represented by fine dots made with the point of the pen ; gravel, by coarser dots. Rocks are drawn in their proper places, in irregular angular forms, imitating their true appearance as seen from above. The nature of the rocks, or the geology of the country, may be shown by applying the proper colors, as agreed on by geologists, to the back of the map, so that they may be seen by holding it up against the light, while they will thus not confuse the usual details.
647. Signs for Vegetation. Woods are represented by scalloped circles, irregularly disposed, imitating trees seen ""in plan," and closer or farther apart according to the thickness of the forest (Fig. 456). It is usual to shade their lower and right-hand sides, and to represent their shadows, as in the figure, though, in strictness, this is inconsistent with the hypothesis of vertical light,

Fig. 456.


Fig. 45\%.

usually adopted for "hill-drawing." For pine and similar forests, the signs may have a star-like form, as in the lower part of Fig. 45\%. When it is desired to distinguish deciduous trees, they are represented as in the upper part of Fig. 45\%. Trees are sometimes drawn "in elevation," or sidewise, as usually seen (Fig. 458). This makes them more easily recognized, but is in

Fig. 45 8.


Fig. 459.

utter violation of the principles of mapping in horizontal projection, though it may be defended as a pure convention. Orchards are represented by trees arranged in rows (Fig. 459). Bushes may be drawn like trees, but smaller. Fig. 460 represents trees and bushes intermingled.

Grass-land is drawn with irregularly scattered groups of short lines, as in Fig. 4.61, the lines being arranged in odd numbers, and

Fig. 460.


Fig. 461.

so that the top of each group is convex, and its bottom horizontal or parallel to the base of the drawing. Meadows are sometimes represented by pairs of diverging lines which may be regarded as tall blades of grass. Uncultivated land is indicated by appropriately intermingling the signs for grass-land, bushes, sand, and rocks. Cultivated land is shown by parallel rows of broken and dotted lines, as in the figure, representing furrows. In Fig. 462 is represented on the right cultivated land with fences, and on the

Fig. 462.


Fig. 463.

left, uncultivated land or "common." Crops are so temporary that signs for them are unnecessary, though often used. They are usually imitative, as for cotton, sugar, tobacco, rice, vines, hops,

Fig. 464.


Fig. 465.


Fig. 466.


Fig. 467.

etc. Gardens are drawn with circular and other beds and walks. Fig. 463 represents a house with grounds.
648. Signs for Water. The sea-coast is represented by drawing a line parallel to the shore, following all its windings and indentations, and as close to it as possible ; then another parallel line a little more distant ; then a third still more distant, and so on, as in Fig. 468. If these lines are drawn from the low-tide mark, a

Fig. 468.

similar set may be drawn between that and the high-tide mark, and dots, for sand, be made over the included space. Fig. 464 represents a sea-coast with rocks and reefs.

Rivers have each shore treated like the sea-shore, as in Fig. 469. Brooks would be shown by only tro lines, or one, according to
their magnitude. Ponds may be drawn like sea-shores, or represented̉ by parallel horizontal lines ruled across them. Marshes and swamps are represented by an irregular intermingling of the preceding sign with that for grass and bushes. Fig. 465 represents a fresh-water marsh. Fig. 466 represents a salt marsh on the right and mud on the left. Fig. $46 \%$ represents osier-beds on the right, and mangrove on the left.
649. Colored Topography. The conventional signs which have been described, as made with the pen, require much time and
 labor. Colors are generally used by the French as substitutes for them, and combine the advantages of great rapidity and effectiveness. Only three colors (besides In-dian-ink) are required, viz., gamboge (yellow), indigo (blue), and lake (scarlet) ; sepia, burnt sienna, yellow ochre, red-lead, and vermilion, are also sometimes used. The last three are difficult to work with. To use these paints, moisten the end of a cake and rub it up with a drop of water, afterward diluting this to the proper tint, which should always be light and delicate. To cover any surface with a uniform flat tint, use a large camel's-hair or sable brush, keep it always moderately full, incline the board toward you, previously moisten the paper with clean water if the outline is very irregular, begin at the top of the surface, apply a tint across the upper part, and continue it downward, never letting the edge dry. This last is the secret of a smooth tint. It requires rapidity in returning to the beginning of a tint to continue it, and dexterity in following the outline. Marbling, or variegation, is produced by having a brush at each end of a stick, one for each color, and applying first one, and then the other, beside it before it dries, so that they may blend, but not mix, and produce an irregularly clouded appearance. Scratched parts of the paper may be painted over by first applying strong alum-water to the place.

The conventions for colored topography, adopted by the French
military engineers，are as follows：Woods，yellow；using gamboge and a very little indigo．Grass－LAND，green；made of gamboge and indigo．Cultivated Land，brown；lake，gamboge，and a little Indian－ink；＂burnt sienna＂will answer．Adjoining fields should be slightly varied in tint．Sometimes furrows are indicated by strips of various colors．GARDENS are represented by small rectangular patches of brighter green and brown．Uncultivated LaND，marbled green and light brown．Bresh，Brambles，etc．， marbled green and yellow．Heath，Furze，etc：，marbled green and pink．Vineyards，purple；lake and indigo．Sands，a light brown；gamboge and lake ；＂yellow ochre＂will do．LaKes and rivers，light blue，with a darker tint on their upper and left－hand sides．Seas，dark blue，with a little yellow added．Marshes，the blue of water，with spots of grass，green，the touches all lying hori－ zontally．Roads，brown ；between the tints for sand and cultivated ground，with more Indian－ink．Hills，greenish－brown；gamboge， indigo，lake，and Indian－ink．Woods may be finished up by draw－ ing the trees and coloring them green，with touches of gamboge toward the light（the upper and left－hand side），and of indigo on the opposite side．

650．Signs for Miscellaneous Objects．Too great a number of these will cause confusion．A few leading ones will be giren ：

| Signal of survey， |  |  |  | Saw－mill， | ¢ ${ }^{\text {² }}$ |  | $4 i 9$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Telegraph， | nmed | ، | $4 \% 1$ | Wind－mill， | CX | ${ }^{6}$ | 450 |
| Court－house， | 方 | 66 | $4 \% 2$ | Steam－mill， | 0 | ＇6 | 481 |
| Post－office， | 星 | ‘ | 473 | Furnace， | 9 | ، | 482 |
| Tavern， | ［四 | ، | $4 \% 4$ | Woolen－factory， | 湥 | ＇6 | 483 |
| Blacksmith＇s shop， | 凧 | c | 475 | Cotton－factory， |  | 6 | 484 |
| Guide－board， | I | ＇6 | $4 \% 6$ | Glass－works， | ） | 6 | 455 |
| Quarry， | $\chi$ | ＇6 | $4 \% \%$ | Church， | す | 6 | 486 |
| Grist－mill， | $\bigcirc$ | 66 | $4 \% 8$ | Graveyard， | － | ، | $48 \%$ |

An ordinary house is drawn in its true position and size, and the ridge of its roof shown, if the scale of the map is large enough. On a very small scale, a small shaded rectangle represents it. If colors are used, buildings of masonry are tinted a deep crimson (with lake), and those of wood with Indian-ink. Their lower and right-hand sides are drawn with heavier lines. Fences of stone or wood, and hedges, may be drawn in imitation of the realities ; and, if desired, colored appropriately.

Mines may be represented by the signs of the planets, which were anciently associated with the various metals. The signs here given represent respectively :

Gold. Silver. Iron. Copper. Tin. Lead Quicksilver.



A large black circle, (3) may be used for coal.
Boundary-lines, of private properties, of townships, of counties, and of States, may be indicated by lines formed of various combinations of short lines, dots, and crosses, as below :
-.-.-.-.-.-.-.-.-.-.-. .-. .-. .-. .-. .-.
$+++++++++++++++++++$
651. Scales. The scale to which a topographical map should be drawn depends on several considerations. The principal ones are these : It should be large enough to express all necessary details, and yet not so large as to be unwieldy. The scale should be so chosen that the dimensions measured on the ground can be easily
Fig. 489.


[^64]
converted, without calculation, into the corresponding dimensions on the map. (See " Scales," Part I.)

For specimens of topographical drawing, see Enthoffer's "Topography," and " United States Coast and Geodetic Survey Reports."

## THE PLANE-TABLE.

652. The Plane-Table is in substance merely a drawing-board fixed on a tripod, so that lines may be drawn on it by a ruler placed so as to point to any object in sight. All its parts are mere additions to render this operation more convenient and precise.*

Such an arrangement may be applied to any kind of "Angular Surveying," such as the Third Method, "Polar Surveying," inits two modifications of Radiation and Progression, and the Fourth Method, by Intersections. Each of these will be successively explained. The instrument is very convenient for filling in the details of a survey, when the principal points have been determined by the more

[^65]precise method of "Triangular Surveying," and can then be platted on the paper in advance. It has the great adrantage of dispensing with all notes and records of the measurements, since they are platted as they are made. It thus saves time and lessens mistakes, but is wanting in precision.
653. The Table. It is usually a rectangular board of well-seasoned pine, about twenty inches wide and thirty long. The paper to be drawn upon may be attached to it by drawing-pins, or by clamp-ing-plates fixed on its sides for that purpose, or by springs pressed upon it, or it may be held between rollers at opposite sides of the table. Tinted paper is less dazzling in the sun. Cugnot's joint, or a pair of parallel plates, like those of the transit, may be used for connecting it with its tripod. A detached level is placed on the board to test its horizontality ; though a smooth ball, as a marble, will answer the same purpose approximately.

A pair of sights, like those of the compass, are sometimes placed under the board, serving, like a " watch-telescope," to detect any movement of the instrument. To find what point on the lower side of the board is exactly under a point on the upper side, so that by suspending a plumb-line from the former the latter may be exactly over any desired point of ground, a large pair of "callipers," or dividers with curved legs, may be used, one of their points being placed on the upper point of the board, and their other point then determining the corresponding under point; or a frame forming three sides of a rectangle, like a slate-frame, may be placed so that one end of one side of it touches the upper point, and the end of the corresponding side is under the table precisely below the giren point, so that from this end a plumb-line can be dropped. A compass is sometimes attached to the table, or a detached compass. consisting of a needle in a narrow box (called a Declinator), is placed upon it, as desired. The edges of the table are sometimes divided into degrees, like the "Drawing-board Protractor." It then becomes a sort of goniometer.
654. The Alidade. The ruler has a fiducial or feather edge, which may be divided into inches, tenths, etc. At each end it
carries a sight like those of the compass. Two needles would be tolerable substitutes. The sights project beyond its edge so that their center lines shall be precisely in the same vertical plane as this edge, in order that the lines drawn by it may correspond to the lines sighted on by them. To test this, fix a needle in the board, place the alidade against it, sight to some near point, draw a line by the ruler, turn it end for end, again place it against the needle, again sight to the same point, and draw a new line. If it coincides with the former line, the above condition is satisfied. The ruler and sights together take the name of Alidade. If a point should be too high or too low to be seen with the alidade, a plumb-line, held between the eye and the object, will remove the difficulty.

A telescope is sometimes substituted for the sights, being supported above the ruler by a standard, and capable of pointing upward or downward. It admits of adjustments similar in principle to the second and third adjustments of the transit.

But even without these adjustments, whether of the sights or of the telescope, a survey could be made which would be perfectly correct as to the relative position of its parts, however far the line of sight might be from lying in the same vertical plane as the edge of the ruler, or even from being parallel to it ; just as in the transit or theodolite the index or vernier need not to be exactly under the vertical hair of the telescope, since the angular deviation affects all the observed directions equally.
655. The plane-table shown in Fig. 491 is one of the standard forms.* The table is leveled by means of three leveling-screws, and tested by a spirit-level' on the alidade. The telescope of the alidade is " transit-mounted "-that is, has both ends of the axis supported.

Distances may be determined by means of stadia-wires placed in the telescope, and heights by means of the vertical arc.
656. Method of Radiation. This is the simplest, though not the best, method of surveying with the plane-table. It is especially

[^66]Fig. 491.

applicable to surveying a field, as in the figure. In it and the following figures, the size of the table is much exaggerated. Set the instrument at any convenient point, as O ; level it, and fix a needle (having a head of sealing-wax) in the board to represent the station. Direct the alidade to any corner of the field, as A, the fiducial edge of the ruler touching the needle, and draw an indefinite line
 by it. Measure 0 A , and set off the distance, to any desired scale, from the needle-point, along the line just drawn, to $\alpha$. The line 0 A is thus platted on the paper of the table as soon as determined in the field. Determine and plat in the same way, $\mathrm{O} B, \mathrm{OC}$, etc., to $b, c$, etc. Join $a b, b c$, etc., and a complete plat of the field is obtained. Trees, houses, hills, bends of rivers, etc., may be determined in the same manner. The corresponding method with the compass or transit has been described. The table may be set at one of the angles of the field, if more convenient. If the alidade has a telescope, the method of measuring distances with a stadia may be here applied with great advantage.
657. Method of Progression. Let A B C D, etc., be the line to be surveyed. Fix a needle at a convenient point of the plane-table, near a corner so as to leave room for the plat, and set up the table at $B$, the second angle of the line, so that the needle, whose point represents $B$, and which should be named $b$, shall be exactly orer that station. Sight to A, pressing the fiducial edge of the ruler against the needle, and draw a line by it. Measure B A, and set off its length, to the desired scale, on the line just drawn, from $b$ to a point $\alpha$, representing A. Then sight to C, draw an indefinite line by the ruler, and on it set off the length of BC from $b$ to $c$. Fix the needle at $c$. Set up at $C$, the point $c$ being over this station, and make the line $c b$ of the plat coincide in direction with
$C B$ on the ground, by placing the edge of the ruler on $c b$, and turning the table till the sights point to B . The compass, if the

Fig. 493.

table have one, will facilitate this. Then sight forward from C to D, and fix CD, $c d$ on the plat, as $b c$ was fixed. Set up at D, make $d c$ coincide with DC , and proceed as before. The figure shows the lines drawn at each successire station. The table drawn at A shows how the survey might be commenced there.

In going around a field, the work would be-proved by the last line "closing". at the starting-point; and, during the progress of the survey, by any direction, as from C to A on the ground, coinciding with the corresponding line, $c a$, on the plat.

This method is substantially the same as the method of surreying a line with the transit. It requires all the points to be accessible. It is especially suited to the surver of a road, a brook, a winding path through roods, etc. The offsets required may often be sketched in by the eye with sufficient precision.

When the paper is filled, put on a new sheet, and begin by fixing on it two points, such as C and D, which were on the former sheet, and from them proceed as before. The sheets can then be afterward united, so that all the points on both shall be in theil true relative positions.
658. Method of Intersection. This is the most usual and the most rapid method of using the plane-table. Set up the instru-
ment at any conrenient point, as X in the figure, and sight to all the desired points, A, B, C, etc., which are visible, and draw in-

Fig. 494.

definite lines in their directions. Measure any line X Y, Y being one of the points sighted to, and set off this line on the paper to any scale. Set up at Y, and turn the table till the line X Y on the paper lies in the direction of X Y , on the ground, as at C in the last method. Sight to all the former points and draw lines in their directions, and the intersections of the two lines of sight to each point will determine them, by the Fourth Method. Points on the other side of the line X Y could be determined at the same time. In surveying a field, one side of it may be taken for the base X Y. Very acute or obtuse intersections should be avoided$30^{\circ}$ and $150^{\circ}$ should be the extreme limits. The impossibility of always doing this renders this method often deficient in precision. When the paper is filled, put on a new sheet, by fixing on it two known points, as in the preceding method.
659. Method of Resection. This method (called by the French Recoupement) is a modification of the preceding method of intersection. It requires the measurement of only one distance, but all the points must be accessible. Let A B be the measured distance. Lay it off on the paper as $a b$. Set the table up at B, and turn it till the line $b a$ on the paper coincides with BA on the ground, as in the Method of Progression. Then sight to C, and draw an indefinite line by the ruler. Set up at C , and turn the line last
drawn so as to point to B. Fix a needle at $a$ on the table, place the alidade against the needle and turn it till it sights to A . Then

the point in which the edge of the ruler cuts the line drawn from B will be the point $c$ on the table. Next sight to D, and draw an indefinite line. Set up at $D$, and make the line last drawn point to C. Then fix the needle at $a$ or $b$, and by the alidade, as at the last station, get a new line back from either of them, to cut the last-drawn line at a point which will be $d$. So proceed as far as desired.
660. To orient the Table.* The operation of orientation consists in placing the table at any point so that its lines shall have the same directions as when it was at previous stations in the same survey.

With a compass this is rery easily effected by turning the table till the needle of the attached compass, or that of the declinator, placed in a fixed position, points to the same degree as when at the previous station.

Without a compass the table is oriented, when set at one end of a line previously determined, by sighting back on this line, as at C in the Method of Progression.

[^67]To orient the table, when at a station unconnected with others, is more difficult. It may be effected thus: Let $a b$ on the table represent a line A B on the ground. Set up at A, make $a b$ coincide with A B , and draw a line from $a$ directed toward a steeple, or other conspicuous object, as $S$. Do the same at $B$. Draw a line $c d$, parallel to $a b$,

Fig, 496.
 and intercepted between $a \mathrm{~S}$, and $b \mathrm{~S}$. Divide $a b$ and $c d$ into the same number of equal parts. The table is then prepared. Now let there be a station, $\mathrm{P}, p$ on the table, at which the table is to be oriented. Set the table, so that $p$ is over P , apply the edge of the ruler to $p$, and turn it till this edge cuts $c d$ in the division corresponding to that in which it cuts $a b$. Then turn the table till the sights point to S , and the table will be oriented.
661. To Find One's Place on the Ground. This problem may be otherwise expressed as interpolating a point in a plat. It is most easily performed by revers-

Fig. 497.
 ing the Method of Intersection. Set up the table over the station, $O$ in the figure, whose place on the plat already on the table is desired, and orient it, by one of the means described in the last article. Make the edge of the ruler pass through some point, $a$ on the table, and turn it till the sights point to the corresponding station, A on the ground. Draw a line by the ruler. The desired point is somewhere in this line. Make the ruler pass through another point, $b$ on the table, and make the sights point to $B$ on the ground. Draw a second line, and its intersection with the first will be the point desired. Using C in the same way would give a third line to prove the work. This operation may be used as a
new method of surveying with the plane-table, since any number of points can have their places fixed in the same manner.

This problem may also be executed without orientation on the principle of trilinear surveying. Three points being given on the table, lay on it a piece of transparent paper, fix a needle anywhere on this, and with the alidade sight and draw lines.toward each of these three points on the ground. Then use this paper to find the desired point, precisely as directed in the last sentence of Art. 720 , page $48 \%$.

When it is desired to set up the plane-table at some undetermined point, not connected by known lines with any other point in the survey, and the table can be readily only approximately oriented, the table may be accurately oriented and. the point be determined by means of the "three-point problem." For the solution of this problem, and for treatise on the plane-tabie, see "United States Coast and Geodetic Survey Report," 1880, Appendix XIII.
662. Inaccessible Distances. Many of the problems in Part I, Chapter V, can be at once solved on the ground by the plane-table, since it is at the same time a goniometer and a protractor. Thus, the Problem of Art. 385 may be solved as follows, on the principle of the construction in the last paragraph of that article : Set the table at C. Mark on it a point, $c^{\prime}$, to represent C, placing $c^{\prime}$ vertically over C . Sight to $\mathrm{A}, \mathrm{B}$, and D , and draw corresponding lines from $c^{\prime}$. Set up at D , mark any point on the line drawn from $c^{\prime}$ toward D , and call it $d^{\prime}$. Let $d^{\prime}$ be exactly over D , and direct $d^{\prime} c^{\prime}$ toward U. Then sight to A and B, and draw corresponding lines, and their intersections with the lines before drawn toward A and B will fix points $a^{\prime}$ and $b^{\prime}$. Then on the line joining $a$ and $b$, given on the paper to represent A and $\mathrm{B}, a b$ being equal to A B on any scale, construct a figure, $a b c d$, similar to $a^{\prime} b^{\prime} c^{\prime} d^{\prime}$, and the line $c d$ thus determined will represent CD on the same scale as A B.
663. Contouring with the Plane-Table. It is used to map the points on the contour-lines as soon as obtained, thus: Range out an approximately level line, and on it set equidistant stakes. At
these stakes range out perpendiculars to the line, and set up several stakes on them for the alignment of the rodman. Draw these lines on the plane-table. Set up and " orient" the table on the ground. Send the rod along one of the perpendiculars till it comes to a point of the right height. Then sight to it with the alidade, and its edge will cut the corresponding line on the table at the correct place on the plat. So for the other perpendiculars.

# PART IV. <br> <br> TRIANGULAR SURVEYING: 

 <br> <br> TRIANGULAR SURVEYING:}
or
By the Fourth Method.

## CHAPTER I.

PLANE SURFACES.
664. Triangular Surveying is founded on the method of determining the position of a point by the intersection of two known lines. Thus, the point P is determined by knowing the length of the line A B, and the angles P B A and P A B, which the lines PA and P B make with A B. By an extension of the principle, a field,

Fig. 498.
 a farm, or a country, can be surveyed by measuring only one line, and calculating all the other desired distances, which are made sides of a connected series of imaginary triangles, whose angles are carefully measured. The district surveyed is covered with a sort of network of such triangles, whence the name given to this kind of surrering. It is more commonly called "Trigonometrical Surrering," and sometimes "Geodesic Surveying," but improperly, since it does not necessarily take into account the curvature of the earth, though always adopted in the great survers in which that is considered.
665. Outline of Operations. A base-line, as long as possible (five or ten miles in surveys of countries), is measured with extreme accuracy.

From its extremities, angles are taken to the most distant objects visible, such as steeples, signals on mountain-tops, etc.

The distances to these and between these are then calculated by the rules of trigonometry.

The instrument is then placed at each of these new stations, and angles are taken from them to still more distant stations, the calculated lines being used as new base-lines.

This process is repeated and extended till the whole district is embraced by these " primary triangles" of as large sides as possible.

One side of the last triangle is so located that its length can be obtained by measurement as well as by calculation, and the agreement of the two proves the accuracy of the whole work.

Within these primary triangles, secondary or smaller triangles are formed, to fix the position of the minor local details, and to serve as starting-points for common surveys with chain and compass, etc. Tertiary triangles may also be required.

The larger triangles are first formed, and the smaller ones based on them, in accordance with the important principle in all surveying operations, always to work from the whole to the parts, and from greater to less.
666. Measuring a Base. Extreme accuracy in this is necessary, because any error in it will be multiplied in the subsequent work. The ground on which it is located must be smooth and nearly level, and its extremities must be in sight of the chief points in the neighborhood. Its point of beginning must be marked by a stone set in the ground with a bolt let into it. Over this a theodolite or transit is to be set, and the line "ranged out." The measurement may be made with chains, steel tapes, etc., or with rods.
667. Measuring a Base with Rods. We will notice, in turn, their materials, supports, alignment, leveling, and contact.

As to materials, iron, brass, and other metals, have been used, but are greatly lengthened and shortened by changes of temperature. Wood is affected by moisture. Glass rods and tubes are preferable on both these accounts ; but wood is the most convenient. Wooden rods should be straight-grained white pine, etc., well sea-
soned, baked, soaked in boiling oil, painted, and varnished. They may be trussed, or framed like a mason's plumb-line level, to prevent their bending. Ten or fifteen feet is a converient length. Three are required, which may be of different colors, to prevent mistakes in recording. They must be very carefully compared with a standard measure.

Supports must be provided for the rods, in accurate work. Posts, set in line at distances equal to the length of the rods, may be driven or sawed to a uniform line, and the rods laid on them, either directly or on beams a little shorter. Tripods or trestles, with screws in their tops to raise or lower the ends of the rods resting on them, or blocks with three long screws passing through them and serving as legs, may also be used. Staves, or legs, for the rods have been used, these legs bearing pieces which can slide up and down them, and on which the rods themselves rest.

The alignment of the rods can be effected if they are laid on the ground, by strings, two or three hundred feet long, stretched between the stakes set in the line, a notched peg being driven when the measurement has reached the end of one string, which is then taken on to the next pair of stakes; or, if the rods rest on supports, by projecting points on the rods being aligned by the instrument.

The leveling of the rods can be performed with a common mason's level ; or their angle measured, if not horizontal, by a "slope-level."

The contacts of the rods may be effected by bringing them end to end. The third rod must be applied to the second before the first has been removed, to detect any movement. The ends must be protected by metal, and should be rounded (with radius equal to length of rod), so as to touch in only one point. Round-headed nails will answer tolerably. Better are small steel cylinders, horizontal on one end and rertical on the other. Sliding ends, with verniers, have been used. If one rod be higher than the nest one, one must be brought to touch a plumb-line which touches the other, and its thickness be added. To prevent a shock from contact, the rods may be brought not quite in contact, and a wedge be let down between them till it touches both at known points on
its graduated edges. The rods may be laid side by side, and lines drawn across the end of each be made to coincide or form one line. This is more accurate. Still better is a "visual contact," a double microscope with cross-hairs being used, so placed that one tube bisects a dot at the end of one rod, and the other tube bisects a dot at the end of the next rod. The rods thus never touch. The distance between the two sets of cross-hairs is of course to be added.

A base could be measured over very uneven ground, or even water, by suspending a series of rods from a stretched rope by rings in which they can move, and leveling them and bringing them into contact as above.
668. Measuring a Base with a Steel Tape. The tape should be from two hundred to five hundred feet long, furnished at one end with a spring-balance for determining the pull on the tape when measuring. It should be tested under the same conditions in which it is to be used-that is, supported at points from ten to twentyfive feet apart, and subjected to a pull of from ten to twenty pounds. The temperature at which the test is made should be noted.

Let us suppose that the tape was tested, resting on supports twenty feet apart, and under a pull of fifteen pounds.

To measure the base, drive stakes along the base-line twenty feet apart, and with one face in line. Drive nails in the lined face of the stakes at the same level, or on an even grade if the ground is not level.

Set a post solidly in the ground at each tape-length along the line, so that the top of the post shall be at the height at which the tape is to be held.

Place the tape on the nails in the stakes, or, better still, on hooks swinging from the nails, and apply a pull of fifteen pounds, bringing the ends of the tape over the posts. Hold the first graduation of the tape over the starting-point on the first post, and mark where the last graduation comes on the second post, by making a line on the head of a copper tack driven into the post, or on a piece of metal fastened on the top of the post. Bring the first graduation on the tape to the mark on the second post, and mark the place
of the last graduation on the third post. So proceed for the whole length of the line.

A steel tape will expand $00000 \%$ of its length for each degree (Fahr.) of rise in temperature. The temperature should be carefully noted when the measurement is made, and the proper correction applied.

The measurement is best made on a still, cloudy day.
If the measured line be on a slope, its measured length must be multiplied by the cosine of the angle of inclination, to reduce it to the horizontal distance between its extremities.
669. Corrections of Base. If the rods were not level, their length must be reduced to its horizontal projection. This would be the square root of the difference of the squares of the length of the rod (or of the base) and of the height of one end above the other ; or the product of the same length by the cosine of the angle which it makes with the horizon.*

If the rods were metallic, they would need to be corrected for temperature. Thus, if an iron bar expands $\frac{1000000}{}$ of its length for $1^{\circ}$ Fahrenheit, and had been tested at $32^{\circ}$, and a base had been measured at $72^{\circ}$ with such a bar ten feet long, and found to contain 3,000 of them, its apparent length would be 30,000 feet, but its real length would be 8.4 feet more. An iron and a brass bar can be so combined that the difference of their expansion causes two points attached to their ends to remain at the same distance at all temperatures. Such a combination is used on the United States Coast Survey.

$$
\begin{aligned}
& \text { Expansion for } 1^{\circ} \text { Fahrenheit. } \\
& \text { Brass bar }=0.00001050903 ; \\
& \text { Iron bar }=0.000006963535 ; \\
& \text { Platinum }=0.0000051344 ; \\
& \text { Glass }=0.0000043119 ; \\
& \text { White-pine }=0.0000022685
\end{aligned}
$$

670. Reducing the Base to the Level of the Sea. Let $\mathrm{A} B=a$

[^68]be the measured base, and $\mathrm{A}^{\prime} \mathrm{B}^{\prime}=x$, the base reduced to the level of the sea, $h$ the height of the measured base above the level of the sea, and $r$ the radius of the earth to the level of the sea. Then we have :
\[

$$
\begin{aligned}
& r+h: r:: a: x . \\
& \therefore x=a \frac{r}{r+h} .
\end{aligned}
$$
\]



$$
a-x=\frac{a h}{r+h}=\frac{\frac{a h}{r}}{1+\frac{h}{r}}=\frac{a h}{r}\left(1+\frac{h}{r}\right)^{-1}
$$

Developing by the binomial formula, we get:

$$
a-x=a \frac{\hbar}{r}-a \frac{h^{2}}{r^{2}}+a \frac{h^{3}}{r^{3}}-, \text { etc. }
$$

As $h$ is very small in comparison with $r$, the first term of the correction is generally sufficient.
671. A Broken Base. When the angle C is very obtuse, the lines AC and CB being measured, and forming nearly a straight

Fig. 500.

line, the length of the line $\mathrm{A} B$ is found thus: Naming the lines, as is usual in trigonometry, by small letters corresponding to the capital letters at the angles to which they are opposite, and letting $\mathrm{K}=$ the number of minutes in the supplement of the angle C , we shall have :

$$
\mathrm{AB}=c=a+b-0.000000042308 \times \frac{a b \mathrm{~K}^{2}}{a+b}
$$

Log. $0 \cdot 000000042308=2 \cdot 6264222-10$.
Proof. Art. 12, Theorem III [Trigonometry, Appendix A], gives, cos. $\mathrm{C}=\frac{a^{2}+b^{2}-c^{2}}{2 a b}$; or $c^{2}=a^{2}+b^{2}-2 a b . \cos$. C. This becomes [Trig., Art. 6], K being the supplement of $\mathrm{C}, c^{2}=a^{2}+b^{2}+2 a b . \cos$. K. The
series [Trig., Art. 5] for the length of a cosine gives, taking only its first two terms, since K is very small, cos. $\mathrm{K}=1-\frac{1}{2} \mathrm{~K}^{2}$. Hence,
$c^{2}=a^{2}+b^{2}+2 a b-a b \mathrm{~K}^{2}=(a+b)^{2}-a b \mathrm{~K}^{2}=(a+b)^{2}\left(1-\frac{a b \mathrm{~K}^{2}}{(a+b)^{2}}\right) ;$
whence,

$$
c=(a+b) /\left(1-\frac{a b K^{2}}{(a+b)^{2}}\right) .
$$

Developing the quantity under the radical sign by the binomial theorem, and neglecting the terms after the second, it becomes

$$
1-\frac{1}{2} \cdot \frac{a b \mathrm{~K}^{2}}{(a+b)^{2}}+\text {, etc. }
$$

Substituting for K minutes, K. siu. $1^{\prime}$ [Trig., Art. 5], and performing the multiplication by $a+b$, we obtain

$$
c=a+b-\frac{a b \mathrm{~K}^{2} \cdot\left(\sin \cdot 1^{\prime}\right)^{2}}{2(a+b)} . \text { Now, } \frac{\left(\sin .1^{\prime}\right)^{2}}{2}=0.0000000423079 ;
$$

whence the formula, $c=a+b-0.000000042308 \times \frac{a b \mathrm{~K}^{2}}{a+b}$.
672. Problem to interpolate a Base. Four inaccessible objects,

Fig. 501.
 $\mathrm{A}, \mathrm{B}, \mathrm{C}, \mathrm{D}$, being in a right line, and visible from only one point, E , it is required to determine the distance between the middle points, B and C , the exterior distances, A B and CD , being known.

Let $\mathrm{A} \mathrm{B}=a, \mathrm{C} \mathrm{D}=b, \mathrm{~B} \mathrm{C}=x$; $A E B=P, A E C=Q, A E D=R$.
Calculate an auxiliary angle, K , such that

$$
\operatorname{tang}{ }^{2} \mathrm{~K}=\frac{4 a b}{(a-b)^{2}} \cdot \frac{\sin \cdot \mathrm{Q} \cdot \sin \cdot(\mathrm{R}-\mathrm{P})}{\sin \cdot \mathrm{P} \cdot \sin \cdot(\mathrm{R}-\mathrm{Q})}
$$

Then is $x=-\frac{a+b}{2} \pm \frac{a-b}{2 \cdot \cos \mathrm{~K}}$.
Of the two values of $x$, the positive one is alone to be taken.
This problem is used when a portion of a base-line passes over water, etc.

Proof. In Fig. 501, produce AD to some point F. The exterior angles, $\mathrm{EBC}=\mathrm{A}+\mathrm{P} ; \quad \mathrm{ECD}=\mathrm{A}+\mathrm{Q} ; \quad \mathrm{EDF}=\mathrm{A}+\mathrm{R} . \quad$ The triangle ABE gives $\frac{\mathrm{BE}}{a}=\frac{\sin . \mathrm{A}}{\sin . \mathrm{P}}$. The triangle A C E gives $\frac{\mathrm{CE}}{a+x}=\frac{\sin . \mathrm{A}}{\sin . \mathrm{Q}}$.
Dividing member by member, we get $\frac{\mathrm{B}}{\mathrm{C}} \frac{\mathrm{E}}{\mathrm{E}}=\frac{a \cdot \sin \cdot \mathrm{Q}}{(a+x) \sin . \mathrm{P}}$.

In the same way the triangle $B E D$ and $C E D$ give $\frac{B E}{b+x}=\frac{\sin .(A+R)}{\sin .(R-P)}$; and $\frac{\mathrm{CE}}{b}=\frac{\sin \cdot(\mathrm{A}+\mathrm{R})}{\sin \cdot(\mathrm{R}-\mathrm{Q})}$. Whence as before, $\frac{\mathrm{BE}}{\mathrm{CE}}=\frac{(b+x) \sin \cdot(\mathrm{R}-\mathrm{Q})}{b \cdot \sin \cdot(\mathrm{R}-\mathrm{P})}$.

Equating these two values of the same ratio, we get

$$
\begin{gathered}
\frac{a \cdot \sin \cdot \mathrm{Q}}{(a+x) \sin \cdot \mathrm{P}}=\frac{(b+x) \sin \cdot(\mathrm{R}-\mathrm{Q})}{b \cdot \sin \cdot(\mathrm{R}-\mathrm{P})} ; \text { and thence } \\
\frac{a b \cdot \sin \mathrm{Q} \cdot \sin \cdot(\mathrm{R}-\mathrm{P})}{\sin \cdot \mathrm{P} \cdot \sin \cdot(\mathrm{R}-\mathrm{Q})}=(a+x)(b+x)=a b+(a+b) x+x^{2} .
\end{gathered}
$$

To solve this equation of the second degree, with reference to $x$, make

$$
\tan .{ }^{2} \mathrm{~K}=\frac{4 a b}{(a-b)^{2}} \cdot \frac{\sin . \mathrm{Q}(\sin . \mathrm{R}-\mathrm{P})}{\sin . \mathrm{P}(\sin . \mathrm{R}-\mathrm{Q})}
$$

Then the first member of the preceding equation $=\frac{1}{4} \cdot(a-b)^{2} \times \tan ^{2} \mathrm{~K}$, and we get $\quad x^{2}+(a+b) x=\frac{1}{4}(a-b)^{2} \cdot \tan ^{2} \mathrm{~K}-a b$, . and $\quad x=-\frac{1}{2}(a+b) \pm \sqrt{ }\left[\frac{1}{4}(a-b)^{2} \cdot \tan ^{2} \mathrm{~K}-a b+\frac{1}{4}(a+b)^{2}\right]$,
 $=-\frac{1}{2}(a+b) \pm \frac{1}{\frac{2}{2}}(a-b) \sqrt{ }\left(\tan ^{2} \mathrm{~K}+1\right)$.
Or, since $\sqrt{ }\left(\tan .{ }^{2} \mathrm{~K}+1\right)=\operatorname{secant} \mathrm{K}=\frac{1}{\cos . \mathrm{K}}$, we have $x=-\frac{a+b}{2} \pm$ $a-b$
$2 \cdot \cos . K^{\circ}$
When $a=b$, or when the two known parts are equal to each other, the above solution is indeterminate. For this case put

$$
\tan .{ }^{2} \mathrm{~K}^{\prime}=\frac{a b \sin . \mathrm{Q} \sin .(\mathrm{R}-\mathrm{P})}{\sin . \mathrm{P} \cdot \sin \cdot(\mathrm{R}-\mathrm{Q})}
$$

and the solution gives:

$$
x=-\frac{1}{2}(a+b) \pm \sqrt{\tan ^{2} \mathrm{~K}^{\prime}+\frac{(a-b)^{2}}{4}}
$$

If $a=b$, this becomes:

$$
x=-\frac{1}{2}(a+b) \pm \tan . \mathrm{K}^{\prime} .
$$

673. Base of Verification. As mentioned in Art. 665, a side of the last triangle is so located that it can be measured, as was the first base. If the measured and calculated lengths agree, this proves the accuracy of all the previous work of measurement and calculation, since the whole is a chain of which this is the last link, and any error in any previous part would affect the very last line, except by some improbable compensation. How near the agreement should be, will depend on the nicety desired and attained in the previous operations. Two bases, 60 miles distant, differed on one great English survey 28 inches ; on another, 1 inch; and on a French triangulation extending over 500 miles, the difference was less than 2 feet. Results of equal or greater accuracy are obtained
on the United States Coast Survey. "The Fire Islaud base, on the south side of Long Island, and the Kent Island base in Chesapeake Bay, are connected by a primary triangulation. This Kent Island base is 5 miles and 4 tenths long, and the original Fire Island base is 8 miles and 7 tenths. The shortest distance between them is 208 miles, but the distance through the triangulation is 320. The number of intervening triangles is 32 , yet the computed and measured lengths of the Kent Island base exhibit a discrepancy no greater than 4 inches."
674. Choice of Stations. The stations, or " trigonometrical points," which are to form the vertices of the triangles, and to be observed to and from, must be so selected that the resulting triangles may be "well-conditioned"-i. e., may have such sides and angles that a small error in any of the measured quantities will cause the least possible errors in the quantities calculated from them. The higher calculus shows that the triangles should be as nearly equilateral as possible. This is seldom attainable, but no angle should be admitted less than $30^{\circ}$, or more than $120^{\circ}$. When two angles only are observed, as is often the case in the secondary

triangulation, the unobserved angle ought to be nearly a right angle.


To extend the triangulation, by continually increasing the sides of the triangles, without introducing "ill-conditioned " triangles, may be effected as in Fig. 502. A B is the measured base, C and D are the nearest stations. In the triangles ABC and ABD , all the angles being observed, and the side A B known, the other sides can be readily calculated. Then, in each of the triangles D AC and DBC, two sides and the contained angles are given to find D C, one calculation checking the other. DC then becomes a base to calculate EF, which is then used to find GH, and so on.

The fewer primary stations used the better, both to prevent confusion and because the smaller number of triangles makes the correctness of the results more " probable."

The United States Coast and Geodetic Surrey displays some fine illustrations of these principles, and of the modifications they may undergo to suit various localities. Fig. 乞ّ03 represents part of the scheme of the primary triangulation resting on the Massachusetts base, and including some remarkably well-conditioned triangles, as well as the system of quadrilaterals, which is a raluable feature of the scheme when the sides of the triangles are extended to considerable lengths, and quadrilaterals, with both diagonals determined, take the place of simple triangles.

The engraving is on a scale of $1: 1,200,000$.
675. Signals. They must be high, conspicuous, and so made that the instrument can be placed precisely under them.

Three or four timbers framed into a
 pyramid, as in Fig. 504, with a long mast projecting abore, fulfill the first and last conditions. The mast may be made rertical by directing two theodolites to it, and adjusting it so that their telescopes follow it up and down, their lines of sight being at right angles to each other. Guy ropes may be used to keep it rertical.

Another form of signal is represented in the three following figures. It consists merely of three stout sticks, which form a tripod, framed with the
signal-staff, by a bolt passing through their ends and its middle. Fig. 505 represents the signal as framed on the ground ; Fig. 506 shows it erected and ready for observation, its base being steadied

Fig. 505.


Fig. 507.

with stones ; and Fig. $50 \%$ shows it with the staff turned aside, to make room for the theodolite and its protecting tent. The heights of these signals varied between fifteen and eighty feet.

Another good signal consists of a stout post let into the ground, with a mast fastened to it by a bolt below and a collar above. By opening the collar, the mast can be turned down and the theodolite set exactly under the former summit of the signal, i. e., in its vertical axis.

A tripod of gas-pipe has been used to support the signal in positions exposed to the sea, as on shoals. It is taken to the desired spot in pieces, and there screwed together and set up.

Signals should have a height equal to at least $\frac{1}{8000}$

Fig. 508.
 of their distance, so as to subtend an angle of half a minute, which experience has shown to be the least allowable.

To make the tops of the signal-masts conspicuous, flags may be attached to them : white and red, if to be seen against the ground ; and red and green, if to be seen against the sky.* The motion of

Fig. 509.


* To determine at a station A, whether its signal can be seen from B, projected against the sky or not, measure the vertical angles BAZ and ZAC . If their sum equals or exceeds $180^{\circ}$, A will be thus seen from B. If not, the signal at $A$ must be raised till this sum equals $180^{\circ}$.
flags renders them visible, when much larger motionless objects are not; but they are useless in calm weather. A disk of sheet-iron, with a hole in it, is very conspicuous. It should be arranged so as to be turned to face each station. A barrel, formed of muslin sewed together, four or five feet long, with two hoops in it two feet apart, and its loose ends sewed to the signal-staff, which passes through it, is a cheap and good arrangement. A tuft of pineboughs fastened to the top of the staff will be well seen against the sky.

In sunshine a number of pieces of tin, nailed to the staff at different angles, will be very conspicuous. A truncated cone of burnished tin will reflect the sun's rays to the eye in almost every situation.

The most perfect arrangement is the "heliotrope." This consists of a mirror a few inches in diameter, so mounted on a tele-

scope, near the eye-end, that the reflection of the sun may be thrown in any desired direction. They hare been observed on at a distance of nearly two hundred miles, when the outlines of the mountains on which they were placed were invisible. A man, called a "heliotroper," is stationed at the instrument. He directs the telescope toward the station at which the transit is placed for observation, and keeps the mirror turned so as to reflect the sun in a direction parallel to the axis of the instrument. This he accomplishes by causing the reflection to pass through two perforated
disks, mounted on the telescope, one near the object-end, and the other near the mirror.

For night-signals, an Argand lamp has been used ; or, better still, a Drummond light, or a magnesium-light. The distinctness of the light is exceedingly increased by a parabolic reflector behind it, or a lens in front of it.
676. Observations of the Angles. These should be repeated as often as possible. In extended surveys, three sets, of ten each, are recommended. They should be taken on different parts of the circle. In ordinary surveys, it is well to employ the method of "traversing." In long sights, the state of the atmosphere has a very remarkable effect on both the visibility of the signals and on the correctness of the observations.

When many angles are taken from one station, it is important to record them by some uniform system. The form given below is convenient. It will be noticed that only the minutes and seconds of the second vernier are employed, the degrees being all taken from the first :

Observations at

| Stations observed то. | Readings. |  | MEAN READINGS. | RIGHT OR LEFT of PRECEDING OBJECT. | REmarks. |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  | vernier a. | VERNIER b . |  |  |  |
|  | - ' " | ' "' | - '" |  |  |
| A | $7019 \quad 0$ | 1840 | 701850 |  |  |
| B | 1033220 | 3240 | 1033230 | R. |  |
| C | 1151420 | 1450 | 1151435 | R. |  |

When the angles are "repeated," the multiple ares will be registered under each other, and the mean of the seconds shown by all the verniers at the first and last readings be adopted.

When the country over which the triangulation extends is flat, it has been found necessary to elevate the transit some distance from the surface of the ground, the stratum of air near the surface being so disturbed by exhalations and inequalities of temperature and density as to render accurate observation impossible. The plan adopted on the Coast Survey is as follows: On the top of a signal-tripod, forty-three feet high, is placed a cap-block, into which is mortised a square hole to receive the signal-pole. Around
the tripod, but not touching it, is erected a rectangular scaffold, forty feet high. On the top of it is a platform, from which the observations are taken, the signal-pole being removed from the capblock, and the transit placed so that its center shall be precisely over the station-point.
677. Reduction to the Center. It is often impossible to set the instrument precisely at or under the signal which has been observed. In such cases proceed

Fig. $\mathfrak{\text { ofl }}$.
 thus: Let C be the center of the signal, and RCL the desired angle, R being the righthand object and L the left-hand one. Set the instrument at D, as near as possible to C , and measure the angle RDL. It may be less than R C L, or greater than it, or equal to it, according as D lies without the circle passing through C , L , and R , or within it, or in its circumference. The instrument should be set as nearly as possible in this last position. To find the proper correction for the obserred angle, obserre also the angle LD C (called the angle of direction), counting it from $0^{\circ}$ to $360^{\circ}$, going from the left-hand object toward the left, and measure the distance DC. Calculate the distances $C R$ and $C L$ with the angle $R D L$, instead of RCL, since they are sufficiently nearly equal. Then, $R C L=R D L+\frac{C D \cdot \sin \cdot(R D L+L D C)}{C R \cdot \sin \cdot 1^{\prime \prime}}-\frac{C D \cdot \sin \cdot L D C}{C L \cdot \sin \cdot 1^{\pi}}$

The last two terms will be the number of seconds to be added or subtracted. The trigonometrical signs of the sines must be attended to. The $\log$. $\sin .1^{\prime \prime}=4 \cdot 6855 \% 49$. Instead of diriding by $\sin .1^{\prime \prime}$, the correction without it, which will be a very small fraction, may be reduced to seconds by multiplying it by 206265.

Example. Let R D L $=32^{\circ} 20^{\prime} 18.06^{\prime \prime}$; L D C $101^{\circ} 15^{\prime} 32 \cdot 4^{\prime \prime}$; C D $=0 \cdot 9 ; ~ C R=35845 \cdot 12 ; C L=29 * 83 \cdot 1$.

The first term of the correction will be $+3 \cdot \% 50^{\prime \prime}$, and the second term $-6 \cdot 113^{\prime \prime}$. Therefore, the obserred angle R D L
must be diminished by $2 \cdot 363^{\prime \prime}$, to reduce it to the desired angle R C L.

Much calculation may be saved by taking the station D so that all the signals to be observed can be seen from it. Then only a single distance and angle of direction need be measured.

It may also happen that the center, C , of the signal can not be seen from D. Thus, if the signal be a solid circular tower, set the theodolite at D , and turn its telescope so that its line of sight becomes tangent to the tower at $\mathrm{T}, \mathrm{T}^{\prime}$; measure on these tangents equal distances, D E, D F, and

Fig. 512.
 direct the telescope to the middle, $G$, of the line E F. It will then point to the center, C ; and the distance D C will equal the distance from D to the tower plus the radius obtained by measuring the circumference.

If the signal be rectangular, measure $\mathrm{D} E, \mathrm{DF}$.
Fic. 513. Take any point G on D E, and on D F set off D H
 $=D G \frac{D}{D} \overline{\mathrm{E}}$. Then is $G H$ parallel to $E F$ (since D G: DH: : DE:DF), and the telescope directed to its middle, K , will point to the middle of the diagonal EF. We shall also have D C $=\mathrm{DK}$ $\frac{D E}{D G}$.
Any such case may be solved by similar methods.
The "phase" of objects is the effect produced by the sun shining on only one side of them, so that the telescope will be directed from a distant station to the middle of that bright side instead of to the true center. It is a source of error to be guarded against. Its effect may, however, be calculated.

When the signal is a tin cone :
Let $r=$ radius of the signal ;
$\mathrm{Z}=$ angle at the point of observation between the sun and the signal ;
$\mathrm{D}=$ the distance.
Then, the correction $= \pm \frac{r \cos { }^{2}{ }^{2} \frac{1}{2} Z}{D \sin 1^{\prime \prime}}$
678. Correction of the Angles. When all the angles of any triangle can be observed, their sum should equal $180^{\circ}$.* If not, they must be corrected. If all the observations are considered equally accurate, one third of the difference of their sum from $180^{\circ}$ is to be added to, or subtracted from, each of them. But if the angles are the means of unequal numbers of observations, their errors may be considered to be inversely as those numbers, and they may be corrected by this proportion ; As the sum of the reciprocals of each of the three numbers of observations is to the whole error, so is the reciprocal of the number of observations of one of the angles to its correction. Thus, if one angle was the mean of three observations, another of four, and the third of ten, and the sum of all the angles was $180^{\circ} 3^{\prime}$, the first-named angle must be diminjshed by the fourth term of this proportion $; \frac{1}{3}+\frac{1}{4}+\frac{1}{10}: 3^{\prime}:: \frac{1}{3}: 1^{\prime} 27 \cdot 8^{\prime \prime}$. The second angle must in like manner be diminished by $1^{\prime} 5 \cdot 9^{\prime \prime}$; and the third by $26 \cdot 3^{\prime \prime}$. Their corrected sum will then be $180^{\circ}$.

It is still more accurate, but laborious, to apportion the total error, or difference from $180^{\circ}$, among the angles inversely as the "weights." On the United States Coast Surrey, in six triangles measured in 1844 by Professor Bache, the greatest error was six tenths of a second.

678 ${ }^{1}$. Calculation and Platting. The lengths of the sides of the triangles should be calculated with extreme accuracy, in two ways if possible, and by at least two persons. Plane trigonometry may be used for even large surveys ; for, though these sides are really arcs and not straight lines, the difference will be only one twentieth of a foot in a distance of $11 \frac{1}{2}$ miles ; half a foot in 23 miles ; a foot in $34 \frac{1}{2}$ miles, etc.

The platting is most correctly done by constructing the triangles, by means of the calculated lengths of their sides. If the measured angles are platted, the best method is that of chords. If many triangles are successively based on one another, they will be platted most accurately by referring all their sides to some one

[^69]meridian line by means of "Rectangular Co-ordinates." In the survey of a country, this meridian would be the true north and south line passing through some well-determined point.
679. Interior Filling up. The stations whose positions have been determined by the triangulation are so many fixed points, from which more minute surveys may start and interpolate any other points. The trigonometrical points are like the observed latitudes and longitudes which the mariner obtains at every opportunity, so as to take a new departure from them and determine his course in the intervals by the less precise methods of his compass and log. The chief interior points may be obtained by "Secondary Triangulation," and the minor details be then filled in by any of the methods of surveying, with chain, compass, or transit, already explained, or by the plane-table. With the transit, "Traversing" is the best mode of surveving, the instrument being set at zero, and being then directed from one of the trigonometrical points to another, which line therefore becomes the "meridian" of that survey. On reaching this second point, in the course of the survey, and sighting back to the first, the reading• should of course be $0^{\circ}$.
680. Radiating Triangulation. This name may be given to a method shown in the figure. Choose a conspicuous point, 0 , nearly in the center of the field or farm to be surveyed. Find other points, A, B, C, D, etc., such that the signal at $O$ can be seen from all of them, and that the triangles A B O, B C O, etc., shall be as nearly equilateral as possible. Measure one side, A B for example. At A measure the angles 0 AB and 0 AG ; at B measure the angles OBA and OBC ; and so on, around the polygon. The correctness
 of these measurements may be tested by the sum of the angles. It may also be tested by the trigonometrical principle that the product of the sines of every alternate angle,
or the odd numbers in the figure, should equal the product of the sines of the remaining angles, the even numbers in the figure.

The triangles AOB, BOC, COD, etc., give the following proportions [Trigonometry, Art. 12, Theorem I] : A O : O B :: sin. (2) : sin. (1); O B : O C : : $\sin$. (4) : $\sin$. (3); O C : O D : : $\sin .(6): \sin$. 5 ; and so on around the polygon. Multiplying together the corresponding terms of all the proportions, the sides will all be canceled, and there will result
$1: 1:: \sin$. (2) $\times \sin$. (4) $\times \sin .(6) \times \sin$. (8) $\times \sin$. (10) $\times \sin .(12) \times \sin$. (14); $\sin$. (1) $\times \sin$. (3) $\times \sin$. (5) $\times \sin$. (7) $\times \sin$. (9) $\times \sin$. (11) $\times \sin$. (13). Hence the equality of the last two terms of the proportion.

The calculations of the unknown sides are readily made. In the triangle ABO , one side and all the angles are given to find $\mathrm{A} O$ and BO . In the triangle $\mathrm{BCO}, \mathrm{BO}$ and all the angles are given to find BC and CO ; and so with the rest. Another proof of the accuracy of the work will be given by the calculation of the length of the side A O in the last triangle, agreeing with its length as obtained in the first triangle.
681. Farm Triangulation. A farm or field may be surveyed by the previous methods, but the following plan will often be more convenient: Choose a base, as X Y ,

Fig. 515.
 within the field, and from its ends measure the angles between it and the direction of each corner of the field, if the theodolite or transit be used, or take the bearing of each, if the compass be used. Consider first the triangles which have XY for a base, and the corners of the field, A, B, C, etc., for vertices. In each of them one side and the angles will be known to find the other sides, X A, X B, etc. Then consider the field as made up of triangles which have their rertices at X . In each of them two sides and the included angle will be given to find its content. If Y be then taken for the common vertex, a test of the former work will be obtained.

The operation will be somewhat simplified by taking for the base-line a diagonal of the field, or one of its sides.
682. Inaccessible Areas. A field or farm mar be surrered, by this "Fourth Method," without entering it. Chocse a base-line

X Y, from which all the corners of the field can be seen. Take their bearings, or the angles between the base-line and their directions. The distances from X and Y to each of them can be calculated as in the last article. The figure will then show in what manner the content of the field is the difference between the contents of the triangles, having X (or Y) for a vertex, which lie outside of it, and those which lie partly within the field and partly outside of it. Their contents can be calculated as in the last

Fig. 516.
 article, and their difference will be the desired content. If the figure be regarded as generated by the revolution of a line one end of which is at $X$, while its other end passes along the boundaries of the field, shortening and lengthening accordingly, and if those triangles generated by its movement in one direction be called plus and those generated by the contrary movement be called minus, their algebraic sum will be the content.
683. Inversion of the Fourth Method. In all the operations which have been explained, the position of a point has been determined, as in Art. 6, by taking the angles, or bearings, of two lines passing from the two ends of a base-line to the unknown point. But the same determination may be effected inversely, by taking from the point the bearings, by compass, of the two ends of the base-line, or of any two known points. The unknown point will then be fixed by platting from the two known points the opposite bearings, for it will be at the intersection of the lines thus determined.
684. Defects of the Method of Intersection. The determination of a point by the Fourth Method, founded on the intersection of lines, has the serious defect that the point sighted to will be very indefinitely determined if the lines which fix it meet at a very acute or a very obtuse angle, which the relative positions of the points observed from and to often render unavoidable. Intersections at right angles should therefore be sought for, so far as other considerations will permit.

## CHAPTER II.

SPHERICAL SURVEYING, OR GEODESY.
685. Nature. It comprises the methods of surveying areas of such extent that the curvature of the earth can not be neglected.

The general method is the same as that given in Chapter I, but more precise methods of measurement and of computation are required, since the triangles into which the surface is divided are spherical triangles.

The United States Coast and Geodetic Survey, the Lake Surrey, and the State Surreys organized by several of the States, are works of this character.

The subject is too extensive to be properly treated within the

Fig. 517.

limits of this work. Only a general sketch of it will be giren, with references to such authorities as mill enable the student to further investigate the subject.

## Field-Work.

686. Reconnaissance. The first step in making a geodetic surrer is the selection of a series of points. A, B, C, etc. (Fig. $51 \%$ ), as the basis of a system of triangulation. In case the country is broken or open, but little difficulty will be experienced in locating these points, and often lines of great length may be secured. Thus, in the triangulation of California,* the line Mount Helena-Mount Shasta

[^70]is 192 miles in length. It is in general advisable to choose the points so that the resulting triangle sides are as nearly equal as possible. To do this, it may be necessary to build towers or scaffolds at the stations A, B, etc., on which to place the instrument. Signals must also be placed at the stations sighted at, their general character depending on the length of the lines of sight.
687. The Base. In order to compute a triangulation, we must have at least one side measured. This measured side is called the base-line, or simply the base. In geodetic work the base must be measured with great accuracy, though it is more important that many bases occur in a system, and these be measured with moderate precision, than that only a few occur, and these be measured with great precision. The reason is, that a check can be more frequently had of the character of the work.

Several different forms of base-measuring apparatus * have been designed, of which probably the simplest and best consists of a steel bar packed in melting ice. The bar will remain of the same length throughout the measurement, as its temperature is always $32^{\circ}$ Fahr.
688. The Angles. Suppose the observer at any station, as D for example. The angles to be measured would be A D C, C D E, EDF. Each of these angles should be measured independently a number of times, depending on the quality of the instrument used, and the mean of the results taken. As a check against mistakes and accidental errors of various kinds, combinations of these angles should be measured, as A D E, A D F, C D F. On the method of measuring an angle with a theodolite, see Wright's "Adjustment of Observations," pp. 253, 254.

## Office-Work.

689. Computation of the Sides of the Triangles. The triangles observed are supposed to have sides of such length that the sum of

[^71]the three angles exceeds $180^{\circ}$ by a certain sensible quantity called the spherical excess. This is usually only a few seconds. For a triangle containing about 76 square miles, which, if equilateral, would have sides 13 miles long, the spherical excess is only one second. For a triangle with sides of 102 miles it is one minute. It must be determined before we can know how much the error of closure is, and therefore what the correction to each angle should be.
690. Spherical Excess. Calling the earth a sphere, the spherical excess $e$ (in seconds) of a triangle is found from the relation
$$
e=\frac{\text { area of triangle }}{R^{2} \sin \cdot 1^{\prime \prime}}
$$
when $\mathrm{R}=$ the radius of the earth.
The triangle surface being small, compared with $\mathrm{R}^{2}$, may be obtained with sufficient accuracy by treating it as if it were plane. Thus, when two sides and the contained angle are given, we have :
\[

$$
\begin{array}{ll} 
& \text { area }=\frac{1}{2} a b \sin . \mathrm{C} ; \\
\text { and therefore } \quad & e=\frac{a b \sin . \mathrm{C}}{2 \mathrm{R}^{2} \sin \cdot 1^{\prime \prime}} .
\end{array}
$$
\]

The earth, howerer, instead of being spherical, is spheroidal in form ; and since a spheroidal triangle may be computed as a spherical triangle on a sphere whose radius is $\sqrt{\overline{R N}}$, when $R$ and $N$ are the radii of curvature of the meridian and of the section normal to the meridian at the mean of the latitudes of the triangle rertices, we replace $R^{2}$ in the above value of $e$ by $R N$. We have then :

$$
\text { excess in seconds }=\frac{a b \sin C}{2 \mathrm{RN} \operatorname{arc} 1^{\prime \prime}}
$$

Writing this in the form

$$
e=m a b \sin . \mathrm{C}
$$

the values of $m$ may be taken from the following table, the argument being the mean latitude of the triangle rertices. The metre is the unit of length:

| Lat. | LOG. $m$. | Lat. | LOG. $m$. | Lat. | LOG. $m$. | Lat. | LOG. $m$. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| - |  | - |  | - |  | - |  |
| 10 | 1.40675 | 25 | $1 \cdot 40589$ | 40 | 1.40451 | 55 | $1 \cdot 40299$ |
| 11 | $1 \cdot 40672$ | 26 | $1 \cdot 40581$ | 41 | 1.40441 | 56 | $1 \cdot 40289$ |
| 12 | $1 \cdot 40668$ | 27 | $1 \cdot 40573$ | 42 | 1.40431 | 57 | $1 \cdot 40280$ |
| 13 | $1 \cdot 40663$ | 28 | $1 \cdot 40564$ | 43 | $1 \cdot 40420$ | 58 | 1-40271 |
| 14 | $1 \cdot 40659$ | 29 | $1 \cdot 40555$ | 44 | 1.40410 | 59 | $1 \cdot 40262$ |
| 15 | $1 \cdot 40654$ | 30 | $1 \cdot 40547$ | 45 | $1 \cdot 40400$ | 60 | $1 \cdot 40253$ |
| 16 | $1 \cdot 40649$ | 31 | $1 \cdot 40537$ | 46 | $1 \cdot 40390$ | 61 | $1 \cdot 40244$ |
| 17 | 140643 | 32 | $1 \cdot 40528$ | 47 | $1 \cdot 40380$ | 62 | $1 \cdot 40235$ |
| 18 | $1 \cdot 40637$. | 33 | $1 \cdot 40519$ | 48 | 1.40369 | 63 | $1 \cdot 40226$ |
| 19 | $1 \cdot 40631$ | 34 | $1 \cdot 40509$ | 49 | $1 \cdot 40359$ | 64 | $1 \cdot 40218$ |
| 20 | $1 \cdot 40625$ | 35 | $1 \cdot 40500$ | 50 | $1 \cdot 40349$ | 65 | $1 \cdot 40210$ |
| 21 | $1 \cdot 40618$ | 36 | $1 \cdot 40491$ | 51 | $1 \cdot 40339$ | 66 | $1 \cdot 40202$ |
| 22 | $1 \cdot 40611$ | 37 | $1 \cdot 40481$ | 52 | $1 \cdot 4.0329$ | 67 | $1 \cdot 40195$ |
| 23 | $1 \cdot 40604$ | 38 | $1 \cdot 40471$ | 53 | 1.40319 | 68 | $1 \cdot 40188$ |
| 24 | $1 \cdot 40597$ | 39 | $1 \cdot 40461$ | 54 | 1.40309 | ¢9 | $1 \cdot 40181$ |
|  |  |  |  |  |  | 70 | $1 \cdot 40174$ |

Example. In a spherical triangle, given $a=122 \% 55, b=$ $94616^{\mathrm{m}}$, angle $\mathrm{C}=50^{\circ} 10^{\prime} 20^{\prime \prime}$, mean latitude of rertices, $\mathrm{A}, \mathrm{B}, \mathrm{C}=$ $45^{\circ} 15^{\prime}$; required the spherical excess.

$$
\begin{aligned}
& \text { log. } a, \check{0} \cdot 08904 \\
& \text { log. b, 4.97596 } \\
& \text { log. sin. C, } 9 \cdot 88535 \\
& \text { log. } m, 1.40398 \\
& \text { log. } 22 \cdot 61,1 \cdot 35433 \\
& \text { whence excess } e=22^{\prime \prime} \cdot 61 \text {. }
\end{aligned}
$$

691. Having found the spherical excess, if the sum of the angles of the triangle is not equal to $180^{\circ}$ plus this excess, the difference is distributed among them, and each angle is corrected by one third of this difference. The angles are then said to be "adjusted."

|  | stations. | observed | angles. | adjusted angles. |
| :---: | :---: | :---: | :---: | :---: |
|  | Prince. | ${ }^{\circ} 11^{\prime} 7$ | $41 \cdot 79$ | $41 \cdot 19$ |
|  | Buck. | 8113 | $13 \cdot 78$ | $13 \cdot 18$ |
|  | Hill | 5659 | $07 \cdot 39$ | 06.79 |
|  |  | $18000$ | 02.96 | $01 \cdot 16$ check. |
|  |  |  | 01•16 |  |
|  |  | $1.80 \div 3=0.60$ |  |  |  |

The difference between the sum of the observed angles and $180^{\circ}$ plus the spherical excess $\left(1^{\prime \prime} \cdot 16\right)$ is $1^{\prime \prime} \cdot 80$, which will make a correction for each angle of $0^{\prime \prime} \cdot 60$. Subtracting this from the obserred angles, we get the corrected or adjusted spherical angles as in the table.
692. Having now the length of one side (or base), and the adjusted values of the three angles of a triangle, the other sides might be computed by the principles of spherical trigonometry. This would be very laborious, but by the help of Legendre's theorem the triangle may be computed as if it were a plane one, and the work be greatly shortened. The theorem is as follows :

Legendre's Theorem. "In any spherical triangle, the sides of which are small compared with the radius of the sphere, if each of the angles be diminished by one third of the spherical excess, the sines of these angles will be proportional to the lengths of the opposite sides."

Example.

| stations. | spherical angles. | PLaNE ANGLES AND distavces. | Logartthms. |
| :---: | :---: | :---: | :---: |
| Prince. Buck . Hill | Buck to Hill. | $\stackrel{\mathrm{m} .}{19189 \cdot 80}$ | 4•2830705 |
|  | $414741 \cdot 19$ | $40 \cdot 80$ | 0•1762239 |
|  | $811313 \cdot 18$ | $12 \cdot 79$ | 9.9948811 |
|  | $\begin{array}{llll}56 & 59 & 6 \cdot 79\end{array}$ | 06.41 | 99235180 |
|  | $1 \cdot 16$ |  |  |
|  | Prince to Hill. Prince to Buck. | $\begin{array}{r} \mathrm{m} \cdot \mathrm{f} \cdot 10 \\ 28456 \\ 24144 \cdot 18 \end{array}$ | $\begin{aligned} & 4 \cdot 4541755 \\ & 4 \cdot 3828124 \end{aligned}$ |

One third of the spherical excess is subtracted from the spherical angles to reduce them to plane angles, which are placed in the third column. Using these plane angles, and the given side, and applying the sine proportion, we have:

| To find b. |  | To find c. |  |
| :--- | :--- | :--- | :--- |
| Log. $a$ | $=4 \cdot 2830 \% 05$ | Log. $a$ | $=4 \cdot 2830 \% 05$ |
| Log. sin. B | $=9 \cdot 9948811$ | Log. sin. C | $=9 \cdot 9235180$ |
| Co-log. sin. A | $=0.1 .62239$ | Co-log. sin. A | $=0 \cdot 1.62239$ |
| Log. $b$ | $=\overline{4.4541 \% 55}$ |  | Log. $c$ |

The logarithms of the sides and of the sines of the plane angles are placed in the last column. For convenience in calculation, the co-log. of angle opposite the given side is taken.
693. In this manner, starting from the base A B (Fig. 51\%), a single chain of spherical triangles may be computed. If another base were measured at EF, a check of the accuracy of the work would be afforded by comparing the computed and measured values of E F. In the Lake-Survey triangulation of Lake Erie, the measured value of the Sandusky base differed from the value computed from the Buffalo base through a chain of thirty-six triangles intervening, by about one inch and a half.
694. Adjustment of a Triangulation. We have considered the measurement and computation of a single chain of triangles proceeding from a single measured base A B. Suppose now that the observer while at station $B$ had sighted over the line $B D$, measuring the angles A B D, C B D, and while at D had measured the angles A D B, C D B. We should then have been able to compute C D from AB, by using any one of the pairs of triangles ABC , B C D : ABC, ACD:ABD, B CD : ABD, ACD. A contradiction is to be expected, as the measurements are not perfect, and therefore before beginning the computation of the sides, an "adjustment" of the angles must be made, so that their most probable values alone enter, and no contradiction will appear in the computed lengths.

The question becomes more complicated when bases are measured at intervals. Thus, suppose the triangulation adjusted from A B as base and E F computed. Another adjustment is needed to harmonize this value with the measured value of E F .

Still further contradictions arise from the introduction of the astronomical determination of the direction of a line (or azimuth), which must be adjusted for before the work is ready for mapping.

Consult " Report of the United States Coast and Geodetic Survey," 1854, 1864 ; Wright, "Adjustment of Observations," chaps. v to ix. On mapping, see "Report United States Coast and Geodetic Survey," 1880.
695. Co-ordinates of the Points. The polar spherical co-ordinates of a point with respect to another point are these : the length of the arc of the great circle passing through the points, and its azimuth, i. e., the angle it makes with the meridian passing through one of its points.

The rectangular spherical co-ordinates of a point have for axes the meridian passing through the origin, and a perpendicular to it. For short distances these may be regarded as in one plane. For greater distances new meridians must be taken-say, not farther apart than fifty miles.

Within that limit the successive triangles may be conceived to be turned down into the same plane.

The astronomical co-ordinates of a point are its latitude and longitude. These are determined by practical astronomy.

See "Report of the United States Coast and Geodetic Survey," 1866, 1868, 1872, 1876, 1880 ; Chauvenet's "Astronomy," rol. ii ; Brunnow's "Astronomy " ; Doolittle's "Astronomy."

The methods of transformation from one system of co-ordinates to another are of great importance in practice. Two problems of common occurrence are the following :
696. Problem. Given the latitude and longitude of $A$, and the azimuth and distance from A to B. Required the latitude and

Fig. 518.
 longitude of $B$, and the azimuth from B to A.

When the triangle sides do not exceed fifteen miles, the geodetic latitudes, longitudes, and azimuths required are computed as follows :

Let $\mathrm{K}=$ distance in metres between two stations, the latitude and longitude of one of which are known.
$\mathrm{L}=$ latitude of first station.
$\mathrm{M}=$ longitude of first.
$\mathrm{Z}=$ azimuth of second station from first, counted from the south around by the rest, from $0^{\circ}$ to $360^{\circ}$. The
algebraic signs of the sine and cosine of this angle must be carefully attended to.
$\mathrm{L}^{\prime}, \mathrm{M}^{\prime}, \mathrm{Z}^{\prime}$, the same things at second station, or quantities required.
$e=$ the eccentricity.
$\mathrm{R}=$ the radius of curvature of the meridian, in metres.
$\mathrm{N}=$ the radius of curvature of a section perpendicular to the meridian, in metres.

Then we have

$$
\begin{aligned}
\mathrm{L}^{\prime} & =\mathrm{L}-\frac{\mathrm{K} \cos . \mathrm{Z}}{\mathrm{R} \operatorname{arc} 1^{\prime \prime}}-\mathrm{K}^{2} \sin .{ }^{2} \mathrm{Z} \frac{\tan . \mathrm{L}}{2 \mathrm{RNarc} 1^{\prime \prime}}-\frac{\mathrm{K}^{2} \mathrm{e}^{2} \sin .2 \mathrm{~L} \cos { }^{2} \mathrm{Z}}{\left.\mathrm{R}^{2}\left(1-\mathrm{e}^{2} \sin .{ }^{2} \mathrm{~L}\right)\right)^{\frac{3}{2}} \operatorname{arcc} 1^{\prime \prime}} \\
& =\mathrm{L}-\mathrm{K} \mathrm{~B} \mathrm{\cos . Z-K}^{2} \mathrm{C} \sin .{ }^{2} \mathrm{Z}-\mathrm{K}^{2} \mathrm{~B}^{2} \mathrm{D} \cos .^{2} \mathrm{Z} \\
\mathrm{M}^{\prime} & =\mathrm{M}+\frac{\mathrm{K} \sin . \mathrm{Z}}{\mathrm{~N}^{\prime} \cos . \mathrm{L}^{\prime} \operatorname{arc} 1^{\prime \prime}} \\
& =\mathrm{M}+\frac{\mathrm{A}^{\prime} \mathrm{K} \sin . \mathrm{Z}}{\cos \cdot \mathrm{~L}^{\prime}}
\end{aligned}
$$

$$
\mathrm{Z}^{\prime}=\mathrm{Z}+180-\left(\mathrm{M}^{\prime}-\mathrm{M}\right) \frac{\sin \cdot \frac{1}{2}\left(\mathrm{~L}^{\prime}+\mathrm{L}\right)}{\cos \cdot \frac{1}{2}\left(\mathrm{~L}^{\prime}-\mathrm{L}\right)}
$$

when the quantities $\mathrm{B}, \mathrm{C}, \mathrm{D}, \mathrm{A}^{\prime}$ may be tabulated for given values of the latitude entering. Tables for this purpose will be found in "Report of the United States Coast and Geodetic Survey," 1884.

Example. Given latitude and longitude of station Victory and length and azimuth of line Victory-Oswego, to find latitude and longitude of Oswego and azimuth of line Oswego-Victory.

The computation may be conveniently arranged in the following tabular form :

| $\begin{gathered} \mathrm{Z} \\ \mathrm{Z}^{\prime}-\mathrm{Z} \\ 180^{\circ} \\ \mathrm{Z}^{\prime} \end{gathered}$ | Victory to Oswego |  |  |  | $196$ |  | 3 | $\begin{gathered} \prime \prime \\ 39 \cdot 23 \\ 48 \cdot 46 \end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | Oswego to Victory ............. |  |  |  | . $\begin{array}{r}180 \\ \hline 16\end{array}$ |  |  | $27 \cdot 69$ |
| $\begin{gathered} L^{\prime}-L \\ L^{\prime} \end{gathered}$ | - | , | " | Victory . . <br> Oswego . . | $\begin{gathered} \mathrm{M} \\ \mathrm{M}^{\prime}-\mathrm{M} \\ \mathrm{M}^{\prime} \end{gathered}$ | 。 | , | " |
|  | 43 | 13 | 06.82 |  |  | 76 | 36 | $22 \cdot 13$ |
|  |  | 13 | $30 \cdot 48$ |  |  |  | 5 | $32 \cdot 92$ |
|  | 43 | 26 | $37 \cdot 30$ |  |  | 76 | 30 | 49-21 |


| $\left\|\begin{array}{c} \mathrm{K} \\ \mathrm{Cos.} \mathrm{Z} \\ \mathrm{~B} \\ \mathrm{~K} \\ \mathrm{~B} \cos . \mathrm{Z} \end{array}\right\|$ | $\begin{aligned} & 4 \cdot 4168423 \\ & 9 \cdot 9813739_{\mathrm{n}} \\ & 85106052 \\ & \hline 2 \cdot 9088214_{\mathrm{n}} \end{aligned}$ | $\begin{gathered} \mathrm{K}^{2} \\ \operatorname{Sin}_{\mathrm{C}}{ }^{2} \mathrm{Z} \end{gathered}$ | $\begin{aligned} & 8 \cdot 83368 \\ & 8.91488 \\ & 1 \cdot 37716 \end{aligned}$ | $\left(\begin{array}{c} (\mathrm{K} \mathrm{~B} \operatorname{cos.} \mathrm{Z})^{2} \\ \mathrm{D} \end{array}\right.$ |  | $\begin{aligned} & 5 \cdot 8176 \\ & 2 \cdot 3924 \end{aligned}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | 9•12572 |  |  | $8 \cdot 2100$ |
| 1st term. 2d term. 3d term. | $\begin{array}{r} -810 \cdot 63 \\ 0.13 \\ 0.02 \end{array}$ |  |  |  |  |  |
| $\begin{aligned} & -\left(L^{\prime}-L\right) \\ & \frac{1}{2}\left(L^{\prime}+L\right) \\ & \frac{1}{2}\left(L^{\prime}-L\right) \end{aligned}$ | $\left[\begin{array}{ccc} \circ & 1 & \prime \prime \\ -13 & 30 \cdot 48 \\ 43 & 19 & 52.06 \\ & 6 & 45 \end{array}\right.$ | $\begin{gathered} \mathrm{N}^{\prime}-\mathrm{M} \\ \sin . \frac{1}{2}\left(\mathrm{~L}^{\prime}+\mathrm{L}\right) \\ \operatorname{Cos.} \frac{1}{2}\left(\mathrm{~L}^{\prime}-\mathrm{L}\right) \\ \text { ar. co. } \\ \mathrm{Z}^{\prime}-\mathrm{Z} \end{gathered}$ | $\begin{aligned} & 2 \cdot 5223459_{\mathrm{n}} \\ & 9 \cdot 836459 \\ & 0 \end{aligned}$ |  |  | $\begin{aligned} & 90305 \\ & 68429 \\ & 743999_{n} \\ & 790332 \end{aligned}$ |
|  |  |  | $\begin{aligned} & 2 \cdot 3588052_{\mathrm{n}} \\ & -228.46^{\prime} \end{aligned}$ |  |  | $\begin{aligned} & 23459_{\mathrm{x}} \\ & 32^{\prime \prime} .94 \end{aligned}$ |

697. Problem. Given latitude and longitude of two stations, to find the distance between them and the azimuth from each to the other.

This is the inverse problem of the preceding. It is solved by dividing

$$
M^{\prime}-M=A^{\prime} K \sin . Z \text { sec. } L^{\prime}
$$

by the first term for $L^{\prime}-L$, namely,

$$
L^{\prime}-L=B K \cos . Z,
$$

whence

$$
\tan . \mathrm{Z}=\frac{\left(\mathrm{M}^{\prime}-\mathrm{M}\right) \mathrm{B}}{\left(\mathrm{~L}^{\prime}-\mathrm{L}\right) \mathrm{A}^{\prime}} \cos . \mathrm{L}^{\prime},
$$

which would give us the azimuth at once if we knew $L^{\prime}-\mathrm{L}$. We therefore seek to compute the smaller terms for the difference of latitude in order to obtain $\mathrm{K} \mathrm{B} \cos$. Z. by subtracting them from the known difference of latitude.
698. In addition to the authorities already quoted, and which give the methods in use in the United States, the following list may be of service: "Ordnance Surreer of Great Britain " ; "Great Trigonometrical Surrey of India" ; "Die Preussische Landestriangulation"; Bessel, "Gradmessung in Ostpreussen" ; Jordan, "Handbuch der Vermessungskunde"; Helmert, "Geodäsie"; Puissant, "Géodésie."

## PART V.

## MARITIME OR HYDROGRAPHICAL SURVEYING.

## INTRODUCTION.

699. The object of this is to fix the positions of the deep and shallow points in harbors, rivers, etc., and thus to discover and record the shoals, rocks, channels, and other important features of the locality.

The relative positions of prominent points on the shore are first very precisely determined by "Trigonometrical Surveying," Part IV. These form the basis of operations, and afford the means of correcting the results obtained by the less accurate methods employed for filling in the details.

In addition to the surveying-instruments already described, the sextant is much used in liydrographical surveying. When the sextant is used for determining the position of a point, the angles are measured between three lines, passing from the required point to three known points. The required point is thus determined by trilinear co-ordinates, or by the fifth method, as explained in Art. 8.

## CHAPTER I.

THE SEXTANT.
700. Principle. The angle subtended at the eye br lines passing from it to two distant objects, may be measured by so arranging two mirrors that one object is looked at directly, and the other object is seen by its image, reflected from one mirror to the second, and from the second mirror to the eye. If the first mirror be moved so that the doubly reflected image of the second object be made to cover or coincide with the object seen directly, then is the

Fig. 519.
 desired angle equal to twice the angle which the mirrors make with each other.

Proof. In Fig. 519, let D and E be two mirrors, perpendicular to the plane of the paper. Let a ray of light from the object A be reflected from the mirrors $D$ and $E$ to the eye at $C$, and B be the other object, looked at directly. Erect perpendiculars to the mirrors, and prolong them until ther meet at F. Prolong the line AD until it meets the line BEat C. The angle D F E is equal to the angle which the tro mirrors make with each other.

Since the angle of incidence equals the angle of reflection, A $D G$ $=\mathrm{GDE}$, and $\mathrm{DEF}=\mathrm{FEC}$,
then we have : D C E = A D E - D E C

$$
\begin{aligned}
& \text { DCE }=2(\mathrm{GDE}-\mathrm{DEF}) \\
& \text { DCE }=2 \mathrm{DFE}
\end{aligned}
$$

701. Description of the Sextant (Fig. 520). The frame is usually of brass, constructed so as to combine strength with lightness. The


Sextant.
handle by which it is held is of wood. The index-arm is movable about a pivot in the center of the graduated arc. The index-glass is a small mirror, attached to the index-arm at the pivot, so as to be perpendicular to the plane of the graduated arc. The horizonglass on the left in the figure is attached perpendicularly to the plane of the instrument, and parallel to the index-glass when the index is at zero. The lower half of this glass is silvered, to make it a reflector, and the upper half is transparent. The telescope is attached so as to point toward the horizon-glass. Sets of colored glasses are used to moderate the light of the sun, when that body is observed.

The sextant has an arc of one sixth of a circle, and measures angles up to $120^{\circ}$, the divisions of the graduated are being num-
bered with twice their real value, so that the true desired angle, subtended by the two objects, is read off at once. The arc is usually graduated to $10^{\prime}$ and read by a vernier to $10^{\prime \prime}$.
702. The box or pocket sextant has the same glasses as the larger sextant, inclosed in a circular box, about three inches in diameter. The lower part, which answers for a handle when in use, screws off and is used for a cover.

The octant has an arc of one eighth of a circumference, and measures angles to $90^{\circ}$.
703. The Reflecting Circle. This is an instrument constructed on the same principle, and used for the same purposes, as the sextant. In it the graduated arc extends to the whole circumference, and more than one vernier may be used by producing the indexarm to meet the circumference in one or two more points.
704. Adjustments of the Sextant. 1. To make the index-glass perpendicular to the plane of the are:

Bring the index near the center of the are and place the eye near the index-glass, and nearly in the plane of the arc. See if the part of the are reflected in the mirror appears to be a continuation of the part seen directly. If so, the glass is perpendicular to the plane of the arc. If not, adjust it by the screws behind it.
2. To make the horizon-glass perpendicular to the plane of the arc:

The index-glass having been adjusted, sight to some well-defined object, as a star, and if, in moring the index-arm, one image seems to separate from or overlap the other, then the horizon-glass is not perpendicular to the plane of the arc. It must be made so by the screws attached to it.

Another method of testing the perpendicularity of the horizonglass is as follows: Hold the instrument vertically, and bring the direct and reflected images of a smooth portion of the distant horizon into coincidence. Then turn the instrument until it makes an angle with the vertical. If the two images still coincide, the glasses are parallel ; and, as the index-glass has been made perpendicular to the plane of the are, the horizon-glass is in adjustment.
3. T'o make the line of collimation of the telescope parallel to the plane of the arc:

The line of collimation of the telescope is an imaginary line, passing through the optical center of the object-lens, and a point midway between the two parallel wires. These wires are made parallel to the plane of the sextant by revolving the tube in which they are placed.

To see whether the line of collimation of the telescope is in adjustment, bring the images of two objects, such as the sun and moon, into contact at the wire nearest the instrument, and then, by moving the instrument, bring them to the other wire. If the contact remains perfect, the line of collimation is parallel to the plane of the arc ; if it does not, the adjustment must be made by the screws in the collar of the telescope.
4. To see if the two mirrors are parallel when the index is at zero:

Bring the direct and reflected images of a star into coincidence. If the index is at zero, then no correction is necessary ; if not, the reading is the "index-error," and is positive or negative, according as the index is to the right or left of the zero.

The "index-error" may be rectified by moving the horizonglass until the images do coincide when the index is at zero, but it is usually merely noted, and used as a correction, being added to each reading if the error is positive, or subtracted from each reading if the error is negative.
705. How to observe. Hold the instrument so that its plane is in the plane of the two objects to be observed, and hold it loosely. Look through the eye-hole, or plain tube, or telescope, at the lefthand or lower object, by direct vision, through the unsilvered part of the horizon-glass. Then move the index-arm till the other object is seen in the silvered part of the horizon-glass, and the two are brought to apparently coincide. Then the reading of the vernier is the angle desired.

If one object be brighter than the other, look at the former by reflection. If the brighter object be to the left or below, hold the instrument upside down.

If the angular distance of the object be more than the range of the sextant (about $120^{\circ}$ ), observe from one of them to some intermediate object, and thence to the other.

A good rest for a sextant is an ordinary telescope-clamp, through which is passed a stick, one end of which is fitted into a hole made in the sextant-handle, and the other end of which is weighted for a counterpoise.

## THE PRACTICE.

706. To set out Perpendiculars. Set the index at $90^{\circ}$. Hold the instrument over the given point by a plumb-line, and look along the line by direct vision. Send a rod in about the desired direction, and when it is seen by reflection to coincide with the point on the line looked at directly, it will be in a line perpendicular to the given line at the desired point.

Conversely, to find where a perpendicular from a given point would strike a line :

Set the index at $90^{\circ}$, and walk along the line, looking directly at a point on it, until the given point is seen by reflection to coincide with the point on the line. A plumb-line let fall from the eye will give the desired point.

Fig. 521.

707. The Optical Square (Fig. 521). This is a box containing two mirrors, fixed at an angle of $45^{\circ}$ to each other, and therefore
giving an angle of $90^{\circ}$, as does the sextant with its glasses fixed at that angle. It is used to set out perpendiculars.
708. To measure a Line, One End being inaccessible. Let A B be the required line, and B the inaccessible point.

At A set off a perpendicular, A C, by Art. 706 . Then set the index at $45^{\circ}$, and walk backward from $A$ in the line of CA pro-

Fig. 522.

longed, looking by direct vision at $C$, until you arrive at some point, D , from which B is seen by reflection to coincide with C . Then is $\mathrm{AD}=\mathrm{AB}$.

If more convenient, after setting off the right angle, set the index at $63^{\circ} 26^{\prime}$, and then proceed as before. The objects will be seen to coincide when at some point, $\mathrm{D}^{\prime}$. Then $A \mathrm{D}^{\prime}=\frac{1}{2} \mathrm{~A} B$. If Fig. 523.

the index be set at $71^{\circ} 34^{\prime}$, then the measured distance will be $\frac{1}{3} \mathrm{~A}$ $B$, and so on.

If the index be set at the complements of the above angles, the
distance measured will be, in the first case, twice, and in the second case three times the desired one.

When the distance A D can not be measured, as in Fig. 523, fix D as before. Set the index at $26^{\circ} 34^{\prime}$, and go along the line to E , where the objects are seen to coincide with each other ; then is A E twice AB , and hence $\mathrm{ED}=\mathrm{AB}$.
709. Otherwise. At A set off an angle, as C A D (A D being a prolongation of AB). Then walk along the line A C with the index

set to half that angle, looking at A directly, and B by reflection, till you come to some point, C , at which they coincide. Then is $\mathrm{CA}=\mathrm{AB}$.
710. To measure a Line when Both Ends are inaccessible. Let AB be the required line. At any point, C , measure the angle ACB. Set the sextant to half that angle, and walk back in the line B C prolonged till at some point, D, A, and B are seen to coincide, as in last problem ; thus making $\mathrm{AC}=\mathrm{C} D$. Do the same on A C produced to some point, E. Then is D E = A B.

Fig. 525.

711. All the methods for overcoming obstacles to measurement, determining inaccessible distances, etc. (Part I, Chapter V), with the transit or theodolite, can be executed with the sextant.
712. To measure Heights. Measure the vertical angle between the top of the object and a mark at the height of the eye, as with a theodolite or transit, and then calculate the height as in Part II, Art. $5 \% 8$.

Otherwise. Set the index at $45^{\circ}$, and walk backward till the mark and the top of the object are brought to coincide. Then the horizontal distance equals the height.

So, too, if the index is set at $63^{\circ} 26^{\prime}$, the height equals twice the distance, and so on. The ground is supposed to be level.

Fig. 526.


When the base is inaccessible: Make $\mathrm{C}=45^{\circ}$, and $\mathrm{D}=26^{\circ} 34^{\prime}$. Then C D $=\mathrm{A}$ B. So, too, if $\mathrm{C}=26^{\circ} 34^{\prime}$, and $\mathrm{D}=18^{\circ} 26^{\prime}$.

This may be used when a river flows along the base of a hill whose height is desired, or in any other like circumstance.
713. To observe Altitudes in an Artificial Horizon. In this method we measure the angle subtended at the eye between the object and its image reflected from an artificial horizon of mercury, molasses, oil, or water. The image of the object in the mercury is looked at directly, and the object itself is viewed by reflection. The object observed is supposed to be so distant that the rays from it,

which strike respectively the index-glass and the artificial horizon, are parallel ; i. e., $S$ and $S^{\prime}$, Fig. $52 \%$, are the same point.

Then will the observed angle $\mathrm{SES} \mathrm{S}^{\prime \prime}$ be double the required angle SEH.

Demonstration.

$$
\begin{gathered}
a=a^{\prime}, a^{\prime}=a^{\prime \prime}, \text { and } a^{\prime \prime}=a^{\prime \prime \prime} . \quad \text { Hence } a^{\prime \prime \prime}=a . \\
\text { SE S" }=a+a^{\prime \prime \prime}=2 a=2 \text { S E H. }
\end{gathered}
$$

714. When the sun is the object observed, to determine whether it is his upper or lower limb whose altitude has been observed, proceed thus:

Having brought two limbs to touch, push the indes-arm from you. If one image passes over the other, so that the other two limbs come together, then you had the lower limb at first. If they separate, you had the upper limb.

In the forenoon, with an inverting telescope, the lower limbs are parting, and the upper limbs are approaching; and vice versa in the afternoon.

Fig. 528.

715. To observe very small altitudes and depressions with the artificial horizon :

Stretch a string over the artificial horizon. Place your head so that you see the string cover its image in the mercury. Then the eye and string determine a vertical plane.

Then observe, looking at the string by direct vision, and seeing the object-by reflection, and you have the angle S E N, in Fig. 528, the supplement of the zenith-distance.

Otherwise. Fix behind the horizon-glass a piece of white paper with a small hole in it, and with a black line on it perpendicular to the plane of the arc.

Then look into the mercury, so as to see in it the image of the line. Your line of sight is then vertical, and the angle to the object seen by reflection is measured as before.
716. To measure Slopes with the Sextant and Artificial Horizon. Let AB be the surface of the ground, and AF a horizontal

Fig. 529.

line. Mark two points equally distant from the eye. Measure, by the preceding method, the angles $\beta$ and $\beta^{\prime}$, which $C A$ and $C B$ make with the vertical CD. Then will half the difference of these angles equal the angle which the slope makes with the horizon.

Demonstration. Continue the vertical line $C D$ to meet the horizontal line in F , and draw C E perpendicular to A B. Then the triangles $C D E$ and $A D F$ are similar, being right-angled and having the acute angles at D equal. Consequently, the angle $\mathrm{DCE}=\mathrm{DAF}$, which is the angle of the slope with the horizon. But $\mathrm{DCE}=\frac{1}{2}\left(\beta^{\prime}-\beta\right)$, hence $\frac{1}{2}\left(\beta^{\prime}-\beta\right)=$ the angle which the slope of the ground makes with the horizon.

If the points $A$ and $B$ be not equally distant from $C$, but jet far apart, this method will still give a very near approximation, the error, which is additive, being $\frac{1}{2}\left(a^{\prime}-a\right)$.

## Demonstration.

$$
\begin{aligned}
\mathrm{DCE} & =\beta^{\prime}+a^{\prime}-90^{\circ} \\
\mathrm{DCE} & =-\beta-a+90^{\circ} \\
2 \mathrm{DCE} & =\beta^{\prime}-\beta+a^{\prime}-a \\
\mathrm{DCE} & =\frac{1}{2}\left(\beta^{\prime}-\beta\right)+\frac{1}{2}\left(a^{\prime}-a\right)
\end{aligned}
$$

717. Oblique Angles. When the plane of two objects, observed by the sextant, is rery oblique to the horizon, the observed angle will differ much from the horizontal angle which is its horizontal projection, and which is the angle needed for platting. The projected angle may be larger or smaller than the obserred angle.

This difficulty may be obviated in various ways:

1. Observe the angular distance of each object from some third object, very far to the right or left of both. The difference of these angles will be nearly equal to the desired angle.
2. Note, if possible, some point abore or below one of the objects, and on the same level with the other, and observe to it and the other object.
3. Suspend two plumb-lines, and place the eye so that these lines cover the two objects. Then observe the horizontal angle between the plumb-lines.
4. For perfect precision, observe the oblique angle itself, and
also the angle of elevation or depression of each of the objects. With these data the oblique angle can be reduced to its horizontal projection, either by descriptive geometry or more precisely by calculation, thus:

Let A H B be the observed angle, and $\mathrm{A}^{\prime} \mathrm{H} \mathrm{B}^{\prime}$ the required horizontal angle.

Conceive a vertical HZ, and a spherical surface, of which H, the vertex of the angle, is the center. Then will the vertical

Fig. 530.

planes, $\mathrm{A} H \mathrm{~A}^{\prime}$ and $\mathrm{BH} \mathrm{B}^{\prime}$, and the oblique plane AHB , cut this sphere in arcs of great circles, $\mathrm{Z}^{\prime \prime}$, $\mathrm{Z} \mathrm{B}^{\prime \prime}$, and $\mathrm{A}^{\prime \prime} \mathrm{B}^{\prime \prime}$, thus forming a spherical triangle, $\mathrm{A}^{\prime \prime} \mathrm{Z} \mathrm{B}^{\prime \prime}$, in which $\mathrm{A}^{\prime \prime} \mathrm{B}^{\prime \prime}=h$ measures the observed angle ; $\mathrm{Z} \mathrm{A}^{\prime \prime}=\mathrm{Z}$ measures the zenith-distance of the point A ; and $\mathrm{Z}^{\prime \prime}=\mathrm{Z}^{\prime}$ measures the zenith-distance of the point B .

These zenith-distances are observed directly, or given by the observed angles of elevation or depression. Then we have the three sides of the triangle to find the angle $\mathrm{B}=\mathrm{A}^{\prime} \mathrm{H}^{\prime}$.

Calling P the half sum of the three sides, we have :

$$
\operatorname{Sin} . \frac{1}{2} B=\sqrt{\frac{\sin \cdot(P-Z) \sin \cdot\left(P-Z^{\prime}\right)_{1}}{\sin \cdot Z \cdot \sin \cdot Z^{\prime}}} .
$$

An approximate correction, when the zenith-distances do not differ from $90^{\circ}$ by more than $2^{\circ}$ or $3^{\circ}$, is this: $\left(90^{\circ}-\frac{Z+Z^{\prime}}{2}\right)^{2}$ tang. $\frac{1}{2} h \cdot \sin .1^{\prime \prime}-\left(\frac{Z-Z^{\prime}}{2}\right)^{2} \cot \cdot \frac{1}{2} h \cdot \sin .1^{\prime \prime}$.

The quantities in the parentheses are to be taken in seconds.
The answer is in seconds, and additive.
717 ${ }^{1}$. The advantages of the sextant over the theodolite are these :
$48 \pm$ MARITIME OR HYDROGRAPHICAL SURVEYING.

1. It does not require a fixed support, but can be used while the observer is on horseback, or on a surface in motion, as at sea.
2. It can take simultaneous observations on two moving bodies, as the moon and a star.

It can also do all that the theodolite can. Its only defect is in observing oblique angles in some cases. By these properties it determines distances, heights, time, latitude, longitude, and true meridian, and thus is a portable observatory.

## CHAPTER II.

TRILINEAR SURVEYING.
718. Trilinear Surveying is founded on the fifth method of determining the position of a point, by measuring the angles between three lines conceived to pass from the required point to three known points, as illustrated in Art. 8.

To fix the place of the point from these data is much more difficult than in the preceding methods, and is known as the "Problem of the three points." It will be here solved geometrically, instrumentally, and analytically.
719. Geometrical Solution. Let A, B, and C be the known objects observed from S, the angles ASB and BSC being there Fig. 531.

measured. To fix this point, S , on the plat containing $\mathrm{A}, \mathrm{B}$, and C , draw lines from A and B , making angles with $\mathrm{A} B$ each equal
to $90^{\circ}-\mathrm{ASB}$. The intersection of these lines at 0 will be the center of a circle passing through A and B , in the circumference of which the point S will be situated.* Describe this circle. Also draw lines from $B$ and $C$, making angles with $B C$, each equal to $90^{\circ}-\mathrm{BSC}$. Their intersection, $\mathrm{O}^{\prime}$, will be the center of a circle passing through B and C . The point S will lie somewhere in its circumference, and therefore in its intersection with the former circumference. The point is thus determined.

In the figure the observed angles, A S B and B S C, are supposed to have been respectively $40^{\circ}$ and $60^{\circ}$. The angles set off are therefore $50^{\circ}$ and $30^{\circ}$. The central angles are consequently $80^{\circ}$ and $120^{\circ}$, twice the observed angles.

The dotted lines refer to the checks explained in the latter part of this article.

When one of the angles is obtuse, set off its difference from $90^{\circ}$ on the opposite side of the line joining the two objects to that on which the point of observation lies.

When the angle ABC is equal to the supplement of the sum of the observed angles, the position of the point will be indeterminate, for the two centers obtained will coincide, and the circle described from this common center will pass through the three points, and any point of the circumference will fulfill the conditions of the problem.

A third angle, between one of the three points and a fourth point, should always be observed, if possible, and used like the others, to serve as a check.

Many tests of the correctness of the position of the point determined may be employed. The simplest one is that the centers of the circles, 0 and $0^{\prime}$, should lie in the perpendiculars drawn through the middle points of the lines A B and BC .

Another is that the line BS should be bisected perpendicularly by the line $00^{\prime}$.

A third check is obtained by drawing at A and C perpendiculars to AB and CB , and producing them to meet BO and BO ',

[^72]produced, in D and E. The line D E should pass through S ; for, the angles B S D and B S E being right angles, the lines D S and S E form one straight line.

The figure shows these three checks by its dotted lines.
720. Instrumental Solution. The preceding process is tedious where many stations are to be determined. They can be more readily found by an instrument called a Station-pointer, or Chorograph. It consists of three arms, or straight-edges, turning about a common center, and capable of being set so as to make with each other any angles desired. This is effected by means of graduated arcs carried on their ends, or by taking off with their points (as with a pair of dividers) the proper distance from a scale of chords constructed to a radius of their length. Being thus set so as to make the two observed angles, the instrument is laid on a map containing the three given points, and is turned about till the three edges pass through these points. Then their center is at the place of the station, for the three points there subtend on the paper the angles observed in the field.

A simple and useful substitute is a piece of transparent paper, or ground glass, on which three lines may be drawn at the proper angles and moved about on the paper as before.
721. Analytical Solution. The distances of the required point from each of the known points may be obtained analytically. Let $\mathrm{AB}=c ; \mathrm{BC}=a ; \mathrm{ABC}=\mathrm{B} ; \mathrm{ASB}=\mathrm{S} ; \mathrm{BSC}=\mathrm{S}^{\prime}$. Also, make $T=360^{\circ}-\mathrm{S}-\mathrm{S}^{\prime}-\mathrm{B}$. Let $\mathrm{BAS}=\mathrm{U} ; \mathrm{BCS}=\mathrm{V}$. Then we shall have :

$$
\begin{aligned}
\text { Cot. } \mathrm{U} & =\cot . \mathrm{T}\left(\frac{c \cdot \sin \cdot \mathrm{~S}^{\prime}}{a \cdot \sin \cdot \mathrm{~S} \cdot \cos \cdot \mathrm{~T}}+1\right), \\
\mathrm{V} & =\mathrm{T}-\mathrm{U}, \\
\mathrm{~S} \mathrm{~B} & =\frac{c \cdot \sin . \mathrm{U}}{\sin \mathrm{~S}} ; \text { or, }=\frac{a \cdot \sin . \mathrm{V}}{\sin \mathrm{~S}^{\prime}}, \\
\mathrm{S} \mathrm{~A} & =\frac{c \cdot \sin . \mathrm{A} \mathrm{~B} \mathrm{~S}}{\sin . \mathrm{S}}, \quad \mathrm{~S} \mathrm{C}=\frac{a \cdot \sin . \mathrm{CB} \mathrm{~S}}{\sin . \mathrm{S}^{\prime}} .
\end{aligned}
$$

Proof. In the triangle A B S, we have
$\sin . \mathrm{ASB}: \sin . \mathrm{BAS}:: \mathrm{AB}: \mathrm{S} \mathrm{B}=\frac{\mathrm{AB} \cdot \sin . \mathrm{BAS}}{\sin \cdot \mathrm{AS} \mathrm{B}}=\frac{c \cdot \sin \cdot \mathrm{U}}{\sin . \mathrm{S}}$.

In the triangle CB , we have
$\sin$. B S C : $\sin$. B C S :: B C : S B $=\frac{\mathrm{BC} \cdot \sin \text {. B C S }}{\sin . \mathrm{B} \mathrm{S} \mathrm{C}}=\frac{a \cdot \sin V}{\sin . \mathrm{S}^{\prime}}$.
Hence, $\frac{c \cdot \sin . \mathrm{U}}{\sin . \mathrm{S}}=\frac{a \cdot \sin . \mathrm{V}}{\sin . \mathrm{S}^{\prime}}$; whence, $c \cdot \sin . \mathrm{S}^{\prime} . \sin . \mathrm{U}-a \cdot \sin . \mathrm{S} . \sin$. $\mathrm{V}=0$ 。

In the quadrilateral ABCS , we have
$\mathrm{BCS}=360^{\circ}-\mathrm{ASB}-\mathrm{BSC}-\mathrm{ABC}-\mathrm{BAS} ;$ or $\mathrm{V}=360^{\circ}-\mathrm{S}-\mathrm{S}^{\prime}$ $-\mathrm{B}-\mathrm{O}$.

Let $\mathrm{T}=360^{\circ}-\mathrm{S}-\mathrm{S}^{\prime}-\mathrm{B}$, and we have $\mathrm{V}=\mathrm{T}-\mathrm{U}$.
Substituting this value of V, in equation [3], we get [Trig., Art. 8],
$c . \sin . \mathrm{S}^{\prime} \sin . \mathrm{U}-a . \sin . \mathrm{S}(\sin . \mathrm{T} . \cos . \mathrm{U}-\cos . \mathrm{T} . \sin . \mathrm{U})=0$.
Dividing by sin. U, we get

$$
c . \sin . \mathrm{S}^{\prime}-a \cdot \sin . \mathrm{S}\left(\sin . \mathrm{T} \cdot \frac{\cos . \mathrm{U}}{\sin . \mathrm{U}}-\cos . \mathrm{T}\right)=0
$$

Whence we have

$$
\frac{\cos \cdot \mathrm{U}}{\sin \cdot \mathrm{U}}=\cot \cdot \mathrm{U}=\frac{c \cdot \sin \cdot \mathrm{~S}^{\prime}+a \cdot \sin \cdot \mathrm{~S} \cdot \cos \cdot \mathrm{~T}}{a \cdot \sin \cdot \mathrm{~S} \cdot \sin \cdot \mathrm{~T}} .
$$

Separating this expression into two parts, and canceling, we get

$$
\cot . \mathrm{U}=\frac{c \cdot \sin \cdot \mathrm{~S}^{\prime}}{a \cdot \sin \cdot \mathrm{~S} \cdot \sin \cdot \mathrm{~T}}+\frac{\cos . \mathrm{T}}{\sin \cdot \mathrm{~T}} .
$$

Separating the second member into factors, we get

$$
\begin{aligned}
& \cot \cdot \mathrm{U}=\frac{\cos \cdot \mathrm{T}}{\sin \cdot \mathrm{~T}}\left(\frac{c \cdot \sin \cdot \mathrm{~S}^{\prime}}{a \cdot \sin \cdot \mathrm{~S} \cdot \cos \cdot \mathrm{~T}}+1\right) \text {; or } \\
& \cot \cdot \mathrm{U}=\cot \cdot \mathrm{T}\left(\frac{c \cdot \sin \cdot \mathrm{~S}^{\prime}}{a \cdot \sin \cdot \mathrm{~S} \cdot \cos \cdot \mathrm{~T}}+1\right)
\end{aligned}
$$

Having found U , equation [4] gives V ; and either [1] or [2] gives S B ; and S A and SC are then given by the familiar "Sine proportion "[Trig., Art. 12].

Attention must be given to the algebraic signs of the trigonometrical functions.

Example. A S B $=33^{\circ} 45^{\prime} ; \mathrm{BSC}=22^{\circ} 30^{\prime} ; \mathrm{A} \mathrm{B}=600$ feet ; $\mathrm{BC}=400$ feet ; $\mathrm{AC}=800$ feet. Required the distances and directions of the point $S$ from each of the stations.

In the triangle A B C, the three sides being known, the angle ABC is found to be $104^{\circ} 28^{\prime} 39^{\prime \prime}$. The formula then gires the angle $\mathrm{BAS}=\mathrm{U}=105^{\circ} 8^{\prime} 10^{\prime \prime}$; whence BCS is found to be $94^{\circ} 8^{\prime}$ $11^{\prime \prime}$; and $\mathrm{S} \mathrm{B}=1042 \cdot 51$; $\mathrm{S} \mathrm{A}={ }^{7} 10 \cdot 193$; and $\mathrm{S} \mathrm{C}=934 \cdot 291$.

## CHAPTER III.

## SURVEYING THE SHORE-LINE.

722. The High-water Line. The principal points on the highwater line are determined by triangulating. The sections between these points are surveyed with the compass and chain, by running a series of straight lines so as to follow, approximately, the shoreline, and taking offsets from the straight lines of the survey to the bends in the shore-line. The straight lines can be more accurately determined by " traversing" with the transit.
723. The Low-water Line. In "tidal-waters" this is more difficult, because low and bare for only a short time. The survey is best made with the sextant, observing from prominent points to three signals, by the trilinear method, and sketching, by the eye, bends of the shore between the stations observed from.

There should be one to observe and one to record. Let 1 and 2, Fig. 532, be two points on the lowwater line, whose position it is desired to determine. The observations taken will be as follows :

$$
\begin{array}{lllll}
\text { (1.) } \begin{array}{l}
\text { A and } \\
\\
\mathrm{B} \text { and }
\end{array} . & . & . & 18^{\circ} \\
\text { (2.) } & \mathrm{B} \text { and C } & . & . & 20^{\circ} \\
& \mathrm{C} \text { and D } & . & . & 15^{\circ} \\
\hline
\end{array}
$$

Fig. 532.


When the shore is inaccessible, a base-line must be measured on the water, and points on the shore fixed by angles from its ends, as in Art. \%29.
724. Measuring a Base on the Water. 1. By sound. Sound travels at the rate of 1,090 feet per second, with the temperature at $30^{\circ}$ Fahr. For higher or lower temperatures, add or subtract $1 \frac{1}{7}$ foot for each degree. If the wind blows with or against the movement of the sound, its velocity must' be added or subtracted. If it blows obliquely, the correction will be its velocity multiplied by the cosine of the angle which the direction of the wind makes with the direction of the sound.
2. By measuring with the sextant the angular height of the mast of a vessel, then we have:

Distance $=$ height of mast $\div$ tan. of the angle

## CHAPTER IV.

## SOUNDINGS.

725. Is sounding, the object is to determine the contour of the bottom of any river, lake, bay, etc., so that a chart of it may be drawn, showing the depth of water at all points covered by the survey. The heights of the points on the bottom are referred to the surface of the water as a "datum-plane," and contour-lines may be determined in the manner described in "Topography."

For the same extent of surface, however, if the same degree of accuracy is required, it will be necessary to measure the height of more points in sounding than in topographical surveying, as the surface between the points, whose heights are measured, can not be seen and sketched.
726. For depths up to eighteen feet a sounding-rod, graduated to feet and tenths, may be used. For greater depths, a lead-line marked to fathoms and half-fathoms will be necessary. The size of the line and the weight of the lead will depend upon the depth of the water. A lead weighing ten pounds will be sufficient for depths up to twenty fathoms. Before using a lead-line it should be thoroughly wet and stretched, and the length of the line should be frequently tested.
727. Before commencing the soundings, stations should be erected on all of the principal points on the shore, such as headlands, bights of bays, etc.

A good station-mark is a post, set in the ground about three feet, leaving about one foot above the surface. The flag-pole is
placed in an auger-hole made in the top of the post. The flag-pole can readily be lifted out, and the transit set over the center of the station. The number of the station should be marked on each post, and it should be distinguished by the combination of colors on the flag, or by the number and arrangement of cross-pieces on the staff.

A permanent "bench-mark" must be established, and the height of the water, when the soundings are made, noted and recorded.

Stations on the water are marked by buoys. A buoy may be made of a light wood float, in which is a hole for the flag-pole. The float is anchored with a stone, or by some other means.
728. The position of the station-buoys, and of the boat when sounding, is determined in various ways.
729. From the Shore. A point on the water may be determined by observing to it with a transit from two stations on the shore, at a given signal or fixed time. In Fig. 533, the length of the line


A B, and the angles which the lines of sight make rith it would then be known, and its place would be fixed by angular co-ordinates. Two observers are necessary.
730. From the Boat with a Compass. Observe from the boat with a prismatic compass, or a Burnier's compass, to two signals on
shore. The place of the boat is then determined, and may be fixed on the map by drawing, from the two known points, lines having the opposite bearings, and their intersection will be the required point. This is rapid and easy, but not precise.
731. From the Boat with the Sextant. Observe with the sextant to three signals on shore, noting the two angles. Two observers, or one observer with two sextants, are necessary. This is the trilinear method, given in Chapter II of this part.
732. Between Stations. Positions of the boat are thus determined only at considerable distances apart, and the boat is rowed.

Fig. 534.

from one of these points to a second one, and soundings taken at regular intervals of time between them.

The distance apart of the soundings depends on the regularity of the bottom, the depth of the water, and the object of the survey. Care should be taken to leave no spot unexplored.

For great accuracy, anchor at some point, and determine its place as above, and then proceed to another point, paying out a line, fastened to the anchor, and sounding at regular distances. Cast anchor at the second point, go back to the first, take up the anchor, go on to the second, and then proceed as before.
733. In a river or narrow water, the soundings may be taken in zigzag lines, from shore to shore, at equal intervals of time, as in Fig. 535.

Where soundings can be made through the ice, the position of
all the points can be determined by any of the methods of surveying. This is the most

Fig. à3ธ.
 accurate method of sounding.
734. On the seacoast the soundings must all be reduced to mean low spring-tides.
735. Tide-Gauges. Tidal observations consist in recording the heights of the water at stated times. In order to determine this, tide-gauges are necessary. The simplest form is a stick of timber, graduated to feet and inches, or tenths, and either set up in the water, or: fastened to the face of a dock, or pier, so that the rise of the tide may be noted upon it. The zero-point of each gauge is taken at or below the lowest tide, and is referred to a permanent "bench-mark" on the shore. On account of the difficulty of sustaining a timber of considerable height against the force of the wind and waves, several successive gauges are sometimes used-the bottom mark on each gauge higher up being on a level with the top line of the next lower. Such an arrangement is required on gentle slopes.

On the sea-coast, where the wares make the reading of the staff difficult, the staff may be attached to a float, inclosed in an upright tube, pierced with holes. The holes in the tube should be of such a size as to allow the water to find the mean height inside, and yet reduce the oscillations to very small limits. Permanent tide-gauges should be self-registering. For a description of a self-registering tide-gauge, see "United States Coast Surrey Report," 1853.
736. "Establishment of the Port." Owing to the obstructions which the tidal wave meets with from the formation of the sea-bed as it approaches the shore, and the character and direction of the channels, the time of high water will differ for different ports in the same ricinity. In order that narigators, entering a port, mar be able to find the time of high water, a standard tide-time is
established-i. e., the number of hours at which high water occurs after the moon's transit over the meridian. This is called the "Establishment of the Port." This time varies with the age of the moon. When observed on the days of full or change, it is the "Vulgar Establishment of the Port." The " Corrected Establishment of the Port" is the mean of the intervals between the times of the transit of the moon and the times of high tide for half a month. This is used for finding the time of high water on any given day, and tables are constructed, from observations at the principal ports, for finding the correction for semi-monthly inequality.
737. In rivers, a number of tide-gauges are necessary, at moderate distances apart, especially at the bends, because the tidal lines of high and low water are not parallel to one another.

The soundings are to be reduced by the nearest gauge, or by the mean of the two between which they may be taken.
738. Beacons and Buoys. Beacons are permanent objects, such as piles of stones with signals on them, usually on shoals and dangerous rocks.

Buoys are floating objects, such as barrels, or hollow iron spheres or cylinders, anchored by a chain, and variously painted, to indicate either dangers or channels.

Those placed by the United States Coast Survey are so colored and numbered that, in entering a bay, harbor, or channel, red buoys with even numbers shall be passed on the starboard or right hand, black buoys with odd numbers on the port hand or left hand, and buoys with red and black stripes on either hand. Buoys in channel-ways are colored with alternate white and black vertical stripes.

## CHAPTER V.

## THE CHART.

739. Hating determined the lines of high and low water, the position of the channels, rocks, shoals, etc., Fig. 536. and the soundings, a chart must be made, on which all these are laid down in their proper places. For scales, see Arts. 43-45.

The high-water line is platted like the bounding lines of a farm. The points determined in the low-water line, and the positions of the boat, determined by the method given in Arts. $728-i 31$, are fixed on the chart by one of the methods given in Arts. $719-i 21$. Contour curves are drawn as in land topography (Part III), for the first four fathoms. These may be indicated by dotted lines, as

in Fig. 536, or they may be shaded with Indian-ink, as in Fig. $53 \%$.

Beyond four fathoms, the depths are noted in fathoms and vulgar fractions.
740. Various conventional signs are used ; some of the principal ones are given in Figs. 538-558.

Fig. 538.


Rocky shore.
Fig. 541.


Fig. 539.


Rocks always bare.

Fig. 540.


Low, swampy shore.

Fig. 545.


Fig. 550.


Fig. 543.


Sandy shore, with hillocks.

Reef of rocks.
Fig. 542.


Rocks sometimes bare.

Fig. 544.


Fig. 546.


Fig. 548.


Fig. 547.


Fig. 549.


Fig. 551.


Buoys.

Fig. 552.


Light-bouse.

Fig. 553.
\& Anchorage for coasters.

Fig. 555.


Signal-house.
Fig. 556.

Rocks always covered.

Fig. 557.
必
Harbors.

Fig. 558.
cos
Channel-marks.

## PART VI. <br> UNDERGROUND OR MINING SURVEYING.

741. It has three objects :
742. To determine the directions and extent of the present workings of a mine.
743. To find a point on the surface of the ground from which to sink a shaft, to meet a desired spot of the underground workings.
744. To direct the underground workings to meet a shaft or any other desired point.

It attains these objects by a combination of surreying and lereling.

## CHAPTER I.

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SURV゙EFINGG AN゙D LEVELING OLD LIVES.
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742. First Object. To determine the direction and estent of the present workings of a mine.

We have to measure:

1. Azimuths, or directions right and left.
2. Lengths or distances.
3. Heights, or distances up and down, either by perpendicular or angular lereling ; usually the latter.

This being done, the relative positions of all the points are known by their three rectangular co-ordinates.

They are referred, first, to a rertical plane (which may be either north and south, or pass through the first line of the sur-
vey) ; second, to another vertical plane, perpendicular to the preceding one ; and, third, to a horizontal datum-plane.
743. In making an underground survey, the same rules and principles apply as to work on the surface. Some differences in methods and detail are necessary, on account of the entire dependence upon artificial light, and the circumscribed limits within which the surveyor is obliged to work.

As the headings and air-ways of a mine are generally driven far in advance of the other workings, it is essential that they should be surveyed with great accuracy, in order to give an intelligent idea of the territory about to be mined. It is also essential, in order that they may serve as a base from which to continue and check the surveys of the interior portions of the mine.
744. Stations. The work may often be much simplified by a careful selection of the stations. See that the average distance between them is as long as possible ; that they are convenient for future use ; and are so chosen that the instrument can be easily set over them. It is also important to locate them where they can be easily and permanently marked. Frequently a station may be so chosen that several different sights can be taken from it-thus economizing much time.
745. Marking the Stations. Whenever possible, all stations should be plainly marked with white paint, and given some distinguishing number or letter. This is necessary for use in extending the surveys at some future time, and also to make the map of use when wishing to identify some particular locality in the mine. The precise point may be indicated by an iron spud like a horseshoe nail, with a hole through the head large enough to take the line of a plumb-bob or plummet-lamp. The spud is driven in a crack in the roof, or in a wooden plug which is driven in a hole that has been previously drilled. The objections to this method are, the length of time it takes to get the spuds in the roof, and also the difficulty in using them when the roof is high. Another objection is that mischievous workmen will drive the spuds up in
the plugs out of sight with the ends of their drills. Probably, as satisfactory a way as any to mark the point is to drill a shallow

Fig. 5 อั9.
 hole, about one eighth of an inch in diameter, in the center of a painted + , or a circle about six inches in diameter. Fig. 559 shows a very convenient device for marking the stations, and plumbing down from them when the roof is high. It is made of light gas-pipe, about half an inch in diameter, and of any convenient length. At one end is a drill ; the other end is bent about three inches out of line, and tapered at the end to fit into the hole made with the drill. There is also a notch in the end large enough to hold the line of a plumb-bob. Attached to the pipe are two rings with shanks about an inch in length. The lower one is fixed, the other is adjustable with a clamp-screw. The upper ring is split in the back wide enough to take a plumb-line easily. To use this device in marking the stations, first strike the drill against the roof, then twist it around a few times. This will generally make a mark large enough to be easily identified. Then reverse the instrument, put the handle of the paint-brush in the upper ring, adjust to the proper height, and clamp it fast. Put the claw, or notch, in the drill-hole and descrive a circle, and also paint the number or letter. To plumb down from the point in the roof, remore the brush, put the plamb-line in the small notch, and through the upper ring, which can be easily done through the split. Hold the claw with the plumb-line in it against the roof at the proper point, then pay out the plumb-line until the plumb-bob reaches the bottom, when the point can be fixed. When not in use, bring the tro rings together, gripping the plumb-bob between them, and clamp fast. Wrap the cord around the shanks of the rings, and fasten with a half-hitch.
746. Points for setting the Transit over. These may be made in a variety of ways, as a nail in a tie, a chalk $\times$ on a rail or stone, a $X$ scratched with a measuring-pin, a speck of paint, or a spot of white paint with a speck of coal in the center. If the chalked $X$ is too coarse, rub away a portion of it with the finger. Special cases may arise where it would be advisable to carry along weights of lead with a short piece of brass wire projecting above the surface, to give a precise point. A center-mark on the top of the telescope will afford the means of placing the transit in position under a plumb-bob suspended from the roof.
747. Giving the Sights. A measuring-pin, if held plumb, with a lamp in front, and a little to one side, makes a very good sight. The pin should be whitened with chalk to make a background for the cross-hair. The cord of a plumb-bob can be seen distinctly up to three or four hundred feet, if a piece of white paper is held behind it and a light is held in front. Care must be taken not to mistake the shadow of the line for the line itself. It is difficult to hold the plumb-bob steady unless it can be hung in the iron spuds mentioned in Art. ${ }^{7} 45$, or the device shown in Fig. 559 is used. Where the mine is smoky, or the sights are very long, sight to the center of the blaze of the lamp, which must be carefully plumbed over the point. To meet cases of this kind, the plummet-lamp has been devised (Fig. 560). It consists of a brass lamp hung in gimbals and supported by two chains. The lamp terminates below in a conical plummet. A shield at the top prevents the flame from burning the string. The sight is taken to the center of the flame. These lamps are generally used in pairs, for back-and-forward sights. They are inconvenient to use, as they require the iron spuds with a hole through the head to support them from the top. Where the roof is high, it is difficult to get up to, the station to put the string through the hole.

Fig. 560.


If care is taken not to make them too heary, they can be supported with the device mentioned in Art. 74.5. Another objection is the additional load they impose upon the party to carry.
748. The Transit. The essential features of a transit to be used for surveys in mines are that the verniers should be so placed as to be easily read by lamp-light,
 and that the marking should be very distinct, on account of the imperfect light arailable. Again, the instrument should not be too heary, as there is often difficult climbing to be done over fallen rock and other mine débris. If the instrument be easily detached from its tripod, it will often be found a conrenience, as thereby the load may be lightened and the instrument itself more carefully carried and more fully protected.
Graduations on solid silver are apt to be tarnished by the pow-der-smoke of the mines. Some makers claim to obviate this by making the graduations on platinum.

If the telescope has a level attached, see that the lamp is not held under it for any length of time, as the heat may explode it. Accidents of this kind have occurred, producing serious results.

In one form of mining

Fig. 562.

transit an extra telescope is attached on one side, as shown in Fig. 561 , and is balanced by a weight on the opposite side. The advantage of this form is, that sights may be taken vertically up or down, as is sometimes necessary in connecting the underground surveys with those on the surface.

In another form, the extra telescope is attached to the transittelescope, as shown in Fig. 562.

The diagonal prism, shown in Fig. 211, may be used with advantage on the extra telescope.
749. Taking the Sights. The beginner will at first have some trouble in catching the light through the telescope. A little practice will overcome this. Hold a lamp a little above the instrument, sight over the top of the telescope, and turn it until it points to the light which it is desired to observe. Now sight through the telescope, and turn it a little each way, until the eye catches the light. Clamp the instrument, and move the object-glass until the light looks like a large round blur. This will form a background on which the cross-hairs can be plainly seen. " Bisect" the blur, then focus the object-glass, and the cross-hairs will be so near the right place that there will be no trouble to find them in bisecting a plumb-line, or whatever else is sighted to. Some instruments have a reflector for illuminating the cross-hairs by throwing a light into the telescope (Fig. 210). The same result can be accomplished by holding a lamp two or three feet in front of the object-glass, and a little to one side, so as to be out of the line of sight.
750. Measuring the Angles. Proceed as in making a traverse on the surface, noting whether the angles are to the right or left. It is generally more satisfactory to put the vernier at zero every time rather than to survey or traverse by the back-angle. The instrument gets some hard usage, and when the surveyor reviews the angle, after having moved to the next station preparatory to measuring a new angle, he has the unsatisfied feeling of not knowing whether the upper motion has slipped, or that he read the angle wrong before. It is also more troublesome to set the vernier at odd degrees and minutes than at 0 , in case there should be a slip of
the upper motion. The surveyor should never omit to check the reading of his angles, either by noting whether the sum of the two readings on each side of the 0 of the vernier is equal to $180^{\circ}$ or by repeating the angle. The latter method is the most satisfactory. If the graduated circle has a double row of figures reading $180^{\circ}$ each way, and the deflection should be greater than $90^{\circ}$, it is only necessary to read the supplement or smaller angle, noting at the same time whether it reads to the right or left on the limb.

The needle-readings, which should always be taken, will prevent the gross error of getting into the wrong quadrant.

Thus, \begin{tabular}{|c|c|c|}
\hline back-sights. \& angles. \& Fore-sights. <br>
\hline S. $30^{\circ} 00^{\prime} \mathrm{W}$. \& $165^{\circ} 00^{\prime} \mathrm{L}$ \& N. $45^{\circ} 00^{\prime} \mathrm{E}$. <br>
\hline S. $30^{\circ} 00^{\prime} \mathrm{W}$. \& $15^{\circ} 00^{\prime} \mathrm{R}$ \& N. $45^{\circ} 00^{\prime} \mathrm{E}$. <br>
\hline

 the the same 

thedle,
\end{tabular}

showing that the last course should be N. E. instead of S. W., as the angle would seem to indicate.

The advantage of this method is that it is a little more convenient to use in working out the courses. It also relieves the surveyor of the inquiry as to whether his rernier has passed the $90^{\circ}$, and he should use the larger or smaller angle. He reads the vernier as it stands, and lets the needle determine the quadrant. It is almost impossible to set up an instrument so solidly that when the cross-hairs are put on a given point they will remain there for any length of time. For this reason it is best not to begin to measure the angle until everything is all ready; then measure and check by doubling it as quickly as can be done mith accuracy. Occasions sometimes arise in which a surveyor has but a few hours in which to make an extended surrey. For a necessity of this kind the use of three transits will be found to expedite the work rery greatly. This prevents loss of time in setting the instrument orer a given point, the work being carried on from the plumb-line of one instrument to that of the next.
751. Plumbing the Shaft. In order that the lines underground may be worked from the same meridian as those on the surface,
they must be deflected from some line whose azimuth is known. Should it not be considered justifiable to depend upon the needle to determine the azimuth, and should it be impossible to enter the mine by a slope or a tunnel, the surveyor will be obliged to resort to plumbing the shaft. Two plumb-lines are carefully put into some known line on the surface, and their direction, which will be in the same line, is again taken at the foot of the shaft, as a meridian from which all the lines underground are deflected. As the two plumb-lines are necessarily but a few feet apart, and as the integrity of all the subsequent work depends upon the accuracy with which the direction of the line on the surface is reproduced by the plumb-lines at the foot of the shaft, it is necessary that extreme care should be exercised in doing the work. Much time will be saved by studying the local conditions of the shaft, and making thorough preparations before beginning the work. In the selection of wires, iron and steel are excellent, when new, as their strength enables a fine wire to support a heavy weight. The objection is that they rust and become treacherous, breaking at most inopportune times. Hard-rolled brass wire, though free from this objection, has to be very carefully used, as it is liable to kink, and then break. If it slips out of the hands while attaching the weights at the bottom, it will fly up the shaft in an almost inextricable tangle. Copper stretches and the weights have to be carefully watched to see that they do not touch the bottom of the vessel in which they are suspended. On the whole, however, it seems to give the best satisfaction. Have the wire wound on two strong reels, set in frames which can be securely anchored. The reels should have stops, so that the weights can be held at any point that may be desired.
752. Suspending the Wires. Nail two boards on the sides of the head-frame, at right angles to the line of sight, and about four feet from the ground. Place on each of these boards a scantling about twelve feet long, letting one end rest on the ground a little out of the line of sight. The upper end should project over the shaft far enough to clear the sides. Put the reels in position, about twenty feet back from the shaft, and also a little out of the
line of sight, and anchor them securely. Fasten weights of about five pounds each to the ends of the wires, and pass them over the ends of the scantlings. Then pay out the wires until the bottom of the shaft is reached. Bring the wires approximately into line by tapping the scantlings with a hammer. In the mean time the assistants at the foot of the shaft will attach the large weights and place them in pails of water. When the signal is given that all is right below, the wires are brought precisely into line, putting in the wire farthest from the instrument first, then bringing the other to it. This can be very easily and accurately done by tapping the scantling gently with a hammer. Examine the wires from the top to the bottom of the shaft to be sure they touch no projecting points. Make all secure at the surface, and, before taking up the instrument to go below, review the work, to be sure that all is correct. Be very careful that no work is done over the head of the shaft while men are at work in the shaft at the foot, lest accidents should occur. At the bottom of the shaft nail two boards across the foot-frame, the same as at the surface. On these place two other boards, about ten inches wide and one quarter of an inch apart, and reaching across the shaft so that the wires will swing freely in the crack between them. These boards serve as a rest for the hand in steadying the vibrations of the wires. They also prevent drops of water from falling into the pails and producing currents which will move the weights. Take a small piece of board and bevel one edge slightly with a knife. Then lay it across the crack between the boards, and bring the beveled edge slowly up to one of the wires until it almost touches. Make a mark on the edge where it bisects the wire, then watch to see if the wire is perfectly still. In deep shafts the oscillations of the wire are rery slow, and it is trying to the eye to watch them through the telescope until they are perfectly still.

Sometimes wires may be steadied by uniting them with a thread or string slightly shorter than the distance between them. The weights are also sometimes placed in oil or mercury. Molasses has also been suggested. If it is impossible to perfectly steady the wires, fasten them at the mean of the oscillations.
753. Getting the instrument into line is not an easy task for the beginner, owing to the difficulty in distinguishing between the lines when looking through the telescope. This is overcome by an assistant holding a white paper with a light alternately in front of and behind the wire farthest away. Another method is to put a couple of round rings in the first wire, and then the second wire can be seen through the openings in the rings. Another very good way is to tack a piece of sheet-iron, of about eight by ten inches, to a piece of board of the same size. Make a hole about one sixteenth of an inch in diameter in the center of the sheet-iron, and at the height of the center of the blaze of a mine-lamp above the board. Bend the sheet-iron so that it will be slightly convex with the bend at the hole. Place this contrivance behind and as close as possible to the rear wire, with the small hole bisecting it. Place a lighted lamp behind the sheet-iron so that the blaze will cover the hole. Put a small piece of board with white paper tacked on it behind the first wire ; also a lighted lamp in front. The instrument can now very quickly be brought into line with the first wire, and the point of light at the second. Verify by bolding white paper, with a light, behind the second wire, and noting whether it is entirely concealed by the other wire.

If possible, use two transits, placed on opposite sides of the shaft, then verify by seeing if they bisect each other's plumb-lines. Do not try to set up the instrument too far away, as it increases the difficulty of getting a clear sight of the wires. Watch, also, that the shadow of the wire is not mistaken for the wire itself. When all is completed, mark the line permanently for future use. Where great accuracy is required, plumb the shaft several times, and take the mean, depending also upon which of the several plumbings has been done 'with the least probability of error.
754. Second Method. When there are two shafts convenient to each other, let a plumb-line down each shaft ; then connect them by a careful survey, both on the surface and underground. Calculate the course between the lines on the surface. Calculate also the course between the wires underground from an assumed meridian. The difference between the two courses will be the correction to be
applied to the underground courses to make them correspond with the azimuth assumed on the surface.
755. Third Method. Use a transit with a telescope outside the standards (Fig. 561). Place the instrument in line directly over the shaft, then produce the line to the foot of the shaft by revolving the telescope so as to sight directly down the shaft. Get two points as far apart as possible at the foot of the shaft, then stretch a fine wire carefully over them, producing the line far enough to make a convenient station over which the transit can be set. In shallow shafts, where communication between the top and bottom is easy, the wire may be lined in directly with the instrument.
756. Fourth Method. If no local attraction exists, and extreme accuracy is not required, use the needle. The needle can be read to within five minutes, and the errors have the probability of correcting each other in the different courses taken. If there is only time and means to do ordinary work, it is better to depend exclusively upon the needle than upon plumbing and deflections poorly done.

The beginner should remember that the greatest care is necessary, and that, when his best has been done, there are possibilities of error. A surveyor who appreciates these errors will not fail to verify his work by repetitions at a later date; as, by making a connection with other openings to the surface, such as a drill-hole, an opening for air, or a connection through a neighboring mine, should such an opportunity present itself.
757. Keeping the Notes. These will depend very much upon the character of the work to be done. Some surveyors prefer to use two note-books. In one are recorded all the instrumental work done with the transit, together with the stations, and whatever explanatory remarks may be necessary. In another, made especially for the purpose, are kept all measurements and references, accompanied with a sketch showing where they were taken. Where the party is large enough, it may be divided so that both of these kinds of work may be kept going at the same time. Another
method, much used, is to keep all the work in one book, where everything will be all together when it is wanted. By having the figures represent certain things when in particular places, and the use of a few symbols and small sketches in special localities, a note-book kept in this manner can generally be made to convey all needed information. Below will be found the right- and left-hand pages of a note-book kept in this manner ; also a map showing the portion of the mine included in the survey of which the notes are a part.

In the first column are the numbers of the stations; also $P$ $\times$, indicating that the station is marked, and in what manner. In the second column are the needle-courses of the back-sights. The third column shows the angles, with R. and L. for right and left. Fourth column, the needle-courses of the fore-sights; the corrected courses can afterward be placed above them in red ink. Fifth column, distances. Sixth column, slopes, and whether $\pm$. Seventh column, height to roof. On the right-hand page, station 1 would be called out by the chairman as follows: Produce 1 and 2 back. At 12, 4 right ; at 20, pillar 7 right ; at 25, 2 left ; at 50, leave point for future reference; at 0,5 right and 9 left ; at 25,3 right and 8 left ; at 58, 1 right and 10 left ; at 58, entrance right, 8 wide and walled ; at 100, 9 right and 3 left ; at 119, entrance right, 8 wide and walled; at distance, 8 right and 2 left, etc.

There will occur to the surveyor, in practice, various symbols and abbreviations which he can use to lessen the labor of recording.

March 4，1886．－Near Foot of Shaft 14.
Set up at point on line of $\times 52$ and $\times 51$ ，produced $39 \cdot 6$ from $\times 51$ ．B．S．on $\times 52$ ．

|  | васк－sights． | angles． | fore－sights． | dis－ | slofe $\pm$. | $\begin{array}{\|l} \text { Heiget } \\ \text { To } \\ \text { Roof. } \end{array}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| P．$\times 70.0$ |  | 40－1ヶ L． | 。 | $39 \cdot 2$ | －0－45 | $\underset{\sim}{\text { Rail．}}$ |
| $\times 7 i$ ． |  |  | $\begin{aligned} & \mathrm{S} .84-15 \mathrm{~W} . \\ & \mathrm{S} .83-00 \end{aligned}$ | $121 \cdot 0$ | 2－05 | ${ }_{10}^{\text {Rail．}}$ |
|  |  |  | N． $73-89 \mathrm{~W}$ ． |  |  | Pave． |
| P．$\times 72.2$ | S． $85-30 \cdot \mathrm{~W}$ ． | 22－06．R． | N． $72-10^{\circ} \mathrm{W}$ ． | $126 \cdot 0$ | ＋0－55 | $9 \cdot 73$ |
| From 2. | N． $73-00 \cdot \mathrm{~W}$ ． | 84－25．R． | S．10－46 w． <br> S． $12-15 \cdot \mathrm{~W}$ ． | $93 \cdot 0$ | ＋7－35 | $\begin{gathered} \text { Pave. } \\ 5 \because 21 \end{gathered}$ |
| 3 | N． $73-00 \cdot \mathrm{~W}$. | 12－29• L． | $\begin{gathered} \text { N. } 86-\sim \text { W. } \\ \text { N. } 86-05^{\circ} \cdot \mathrm{W} . \end{gathered}$ | 104：3 | ＋1－02 | $\begin{aligned} & \text { Pave. } \\ & 11 \cdot 43 \end{aligned}$ |
| P．$\times$ V．1．A． | N． $86-00 \cdot \mathrm{~W}$ ． | 74－27．L． | s．19－25 W． <br> S． $19-30 \cdot{ }^{-1}$ ． | 84－5 | ＋10－02 | $\begin{aligned} & \text { Tie. } \\ & 4 \cdot 23 \end{aligned}$ |
| P．$\times 74.4$ | N． $86-00 \cdot$ W． | 89－55．L． | $\begin{aligned} & \mathrm{N} \cdot \frac{3-57 \cdot}{\mathrm{E} \cdot \mathrm{E}} . \\ & \mathrm{N} .00 \cdot \mathrm{E} \end{aligned}$ | $41 \cdot 8$ | －3－01 | $\begin{aligned} & \text { Tie. } \\ & 6.75 \end{aligned}$ |
| P．$\times 75$. | N．4－00．E． | 88－39 L． | $\begin{aligned} & \text { S. } 84-42 \cdot \mathrm{E} . \\ & \text { S. } 8440 \mathrm{E} . \end{aligned}$ | 78.3 | －0－32 | ${ }_{7}^{\text {Tie．}}$ ， |
| P．$\times 76.6$ | S． $84-45 \cdot \mathrm{E}$ ． | $10 \cdot 09 \cdot \mathrm{R}$ ． | S．${ }_{\text {S．}}$ ． $74-23.3 \cdot \mathrm{E}$ E． | 125.7 | －0－22 | Ruil. |
|  |  |  | N. 8e-98.E. |  |  | Pave． |
| P．$\times 77$. | S． $74-30 \cdot \mathrm{E}$ ． | 16－59• L． | N．89－00 E． | 144.9 | －0－08 | $14 \cdot 12$ |
| From 77. | S．86－35 E． | 3－07．L． |  | $217 \cdot 0$ | －0－15 | Pare． $7 \cdot 52$ |
| P．$\times$ H．From 77 | S． $86-35 \cdot$ E． | 43－17．R． | N．${ }_{\text {N．} 50-150} 50 . \mathrm{W}$ ． | $\uparrow 3 \cdot 6$ | －4－1 | Pave． 6． 25 20． |
|  |  |  | S． $57-58 \cdot \mathrm{E}$ ． |  |  | Rock． |
| 8 | S．86－35．E． | $33-34 \cdot \mathrm{R}$ ． | S． $53-35 \cdot \mathrm{E}$ ． | $99 \cdot 3$ | －0－80 | $7 \cdot 15$ |
| 9 | S．53－35• E． | $3716 . \mathrm{L}$ ． | $\begin{aligned} & \text { N. } 85-46 . \text { E. } \\ & \text { N. } 89-10 \cdot \mathrm{E} . \end{aligned}$ | N． $85^{\circ}$ | E．Err | $0^{\circ}-01^{\prime}$ 。 |

Begin at P．$\times$ V． 1 above to run short chambers．

| P．$\times$ V． 2. | 1 | S． $80-45^{\circ} \mathrm{E}$ ． | 81－44．R． | $36 \cdot 40$ | －2．01 | Pave． 4.92 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| P．$\times$ V． 3. | 2 | S． $1-00 \cdot \mathrm{~W}$ ． |  | $20 \cdot 00$ | ＋1508 | $\begin{gathered} \text { Pave. } \\ 5^{\prime} \cdot 23 \end{gathered}$ |
| P．$\times$ V． 5. | 3 | N． 8515 E． | 81－24• L | 79.80 | $-6 \cdot 30$ | $\begin{aligned} & \text { Pare. } \\ & 5 \because 21 \end{aligned}$ |
|  | 4 | S．3－50．W． |  | 43.00 | ＋25．00 | $\begin{aligned} & \text { Rail. } \\ & 8: 20 \end{aligned}$ |

Set up at $\oplus 53$ on line between $\times 74$ and $\times 75$ ．B．S．on $\times 74$ ．

|  | $\begin{gathered} \text { S. } 84-42 \cdot \mathrm{E} . \\ \text { S. } 84-40 \cdot \mathrm{E} . \end{gathered}$ | 8003 L ． |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| P．N． |  |  | N． $15-15$ E． | 77•20 | $-10-12$ | 8．7 |





$$
\begin{array}{lllll}
\frac{10 \mathrm{rb}}{5 \cdot 4} & \frac{20}{7 \cdot 4} & 30 & 7 \cdot 3 & 50 \\
8 \cdot 3 & 8 \cdot 3 & \text { dist. } \\
\hline
\end{array}
$$

| F. R. $=$ far rib. | $\square=$ blind $\mathrm{n}^{\text {ntrance } .}$ | b. | hdg. = heading. |
| :---: | :---: | :---: | :---: |
| N. R. $=$ near | dist. $=$ distance . | ch. $=$ chamb | ave. |

Fig. 563.

758. Tabling the Survey. On pages 514 and 515 will be found a form and the tabling of the above field-notes for office use and record. It is best to have a specially prepared book already ruled to the reqaired form. All the work of tabling can then be done in this book. Should there ever be an occasion to review the work, it can easily be found.

The two double columns headed 1 and 2 are for convenience in taking down the numbers as they are called off from Gurden's "Traverse Tables," which are to single minutes, and distances to one hundred feet. For convenience in description, we will suppose two persons, A and B , to be tabling the above survey. A will take the sheet or book on which have been recorded the stations, corrected courses, distances, and slopes, and call out the angle, which in the present case we will suppose to be N. $55^{\circ} 30^{\prime}$ W., distance $39 \cdot 19$. B finds this in the book of tables, and on the edge of a sheet of blank paper checks the heary line on the center of the page ; also, the two minute columns. A then calls out the distance, $39 \cdot 19$, which B sets down on his sheet of paper, and then, using his paper as a straight-edge, slides it down the page until he comes to 39, taking care to keep the check on the center line. He will then call out the numbers under the checks for the minute columns, always reading the left-hand one first, to A , who will record them as he receives them in columns 1 and 2 . The same operation is repeated for the 19. A will then call out the next angle, and, while B is searching for it, he will add the numbers given, and, if he has time, carry the results out to the proper columns of N., S., E., and W. A glance at the course, noting whether it is greater or less than $45^{\circ}$, will tell him whether the larger number should be put in the column of Latitude or Departure. The same operation is repeated for all the courses.

For convenience in plotting and calculations, the latitudes and departures should all be referred to a common origin of co-ordinates. In this survey the origin is taken at the west plumb-line of the shaft. Station 51 has been found by previous work to have latitude north +112 , and departure west 159. In like manner, 51 has been found to have a + elevation of $18 \% \cdot \%$. The slopes and distances should be reduced first, then the

| stations. | course. | (tis- $\begin{gathered}\text { dis- } \\ \text { tance. }\end{gathered}$ | $\begin{gathered} \text { sLOPE } \\ + \text { OR } \end{gathered}$ | $\begin{gathered} \text { SLOPE DISTANCEs } \\ \text { REDUCED. } \\ \hline \end{gathered}$ |  | COURSE AND DIS TANCE REDUCED |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  | 1st. | 2d. | 1st. | 2d. |
|  | N. 55.30 W . | $39 \cdot 19$ | $\begin{array}{r} \circ \\ -0 \cdot 45 \end{array}$ | $\begin{array}{r}38.99 \\ \hline 20\end{array}$ | $0.51$ | $\begin{aligned} & 32 \cdot 14 \\ & \cdot 16 \end{aligned}$ | 22.09 <br> 11 |
|  |  | $39 \cdot 20$ |  | 39.19 | $\overline{-0.51}$ | 3230 | 22-20 |
|  | S. $84 \cdot 15 \mathrm{~W}$. | $120 \cdot 96$ | +1.31 | 999.97 20.99 | 2.66 | $\begin{array}{r}119.40 \\ \hline 96\end{array}$ | ${ }^{12 \cdot 02}$ |
| 71............. |  | $121 \cdot 00$ |  | - $120 \cdot 99$ | $\frac{56}{+3 \% 22}$ |  | 12.11 |
|  |  |  |  | 99.99 | 119 | 95.96 23.99 | ${ }_{\text {2 }}^{28} 15$ |
| 72............ | N. 7339 W. | 125.99 | + $0 \cdot 41$ | ${ }_{26 \cdot 00}^{99.99}$ | ${ }^{1.31}$ | ${ }^{23.95}$ | $\stackrel{\text { r-04 }}{ }$ |
|  |  | $126 \cdot 00$ |  | 125.99 | +1.50 | $\overline{120.90}$ | $\overline{35 \cdot 47}$ |
| 73 is 20 beyond station...... |  |  |  | 99.95 | 1•30 | $99 \cdot 77$ 3.99 | 6.74 .27 |
|  | N. 8608 W. | 104.28 | +1-02 | $4 \cdot 30$ | . 77 | -28 | . 02 |
|  |  | $104 \cdot 30$ |  | 104.25 | +2.57 | 104.04 | 7.03 |
|  | N. 357 E . | 41.74 |  | ${ }^{40 \cdot 94}$ | 2.15 | ${ }^{40} \cdot 90$ | 2.82 .05 |
| 74.... ....... |  | $41 \cdot 80$ | -3 01 | $41 \cdot 74$ | $\overline{-2 \cdot 19}$ | $\stackrel{41 \cdot 63}{ }$ | 2.7 |
| 75............ |  | 78.30 | -0 52 | 78*00 | 1118 | 77.67 | 7.20 .03 |
|  | S. 8442 E . | $78 \cdot 30$ |  | 7530 | $-1 \cdot 18$ | $\frac{87}{77}$ | $7 \cdot 23$ |
|  |  | 125•\% | -0 45 | $120 \cdot 00$ $5 \% \%$ | 1.57 | 115.65 5.49 | ${ }^{31 \cdot 97}$ |
| 76............ | S. 7433 E . | $125 \cdot 70$ |  | 125\%\% | -164 | 121•14 | $3{ }^{3} \cdot 18$ |
|  |  | 144.90 | -0 20 | 140.00 4.90 | 0.82 | $139 \cdot 95$ 4.90 | $\stackrel{3}{3} \cdot 12$ |
| 77. | N. 8828 E. | 144.90 |  | 141:90 | -0.84 | $144 \cdot 55$ | $3 \cdot 5$ |
| Point on line between $77 \& 50$ | S. $57-58 \mathrm{E}$. | 56.89 | -1 00 | $\frac{5 \check{5} \cdot 99}{50 \cdot 99}$ | 0.93 .00 | 47.47 | 29:70 |
|  |  | 56.90 |  |  | $-0.98$ | $45 \cdot 22$ | 3017 |
| Close on 70.... | S. 8415 W . | $50 \cdot 00$ 50.00 | -0 32 | 50 | $-0.46$ | 49:75 | 5.01 |
|  | S. 1046 W . | $\begin{aligned} & 91 \cdot 65 \\ & 92 \cdot 46 \end{aligned}$ | +7.35 | ${ }^{91 \cdot 20}$ | 12.14 | S9.40 | $17 \cdot 00$ |
| From 72 to V. 5. |  |  |  | 91.65 | +1220 | 90.04 | ${ }^{17} \cdot 13$ |
| From 20 back of 73 to V. 1... | S. 1925 W . | 82.21 | +10 02 | 82:72 | 14:63 | 75-28 | 27.59 |
|  |  | 84.50 |  | $\frac{49}{83 \div 1}$ | + $+\frac{\cdot 9}{1+\cdot \overbrace{}^{2}}$ | \% 3.45 | $\frac{.07}{2-66}$ |
| From 77 to $40 \ldots$. | N. 8521 E . | 216.99 |  | $209 \cdot 99$ <br> $7 \cdot 00$ | -92 | $\underset{6.97}{209.31}$ | ${ }^{17} \cdot{ }^{17}$ |
|  |  | 217.00 | -0 15 | $216 \cdot 99$ | -0.95 | $216 \cdot 25$ | 1759 |
| From 77 to H... | N. 4815 T . | 73.49 | -4 12 | 72.81 | 5.85 | $\begin{array}{r}54 \cdot 46 \\ 3 \\ \hline 37\end{array}$ | ${ }^{48 \cdot 6.61}$ |
|  |  | $73 \cdot 68$ |  | 78.49 | $-5.40$ | $\overline{54 \times 3}$ | 4-94 |
| From $77 \begin{gathered}\text { old sta. } \\ \text { to } \\ 50 . . .\end{gathered}$ | S. 5758 E . | 99.30 | -0 30 | 9900 | -6 | $\begin{array}{r}83 \\ \hline 85 \\ \hline 29\end{array}$ | 52.51 .16 |
|  |  | $99 \cdot 30$ |  | 99:30 | -'s6 | 84.18 | 52.67 |
| $\begin{aligned} & \text { From V. } 1 \text { to } \\ & \text { V. } 2 \ldots \ldots . \end{aligned}$ | S. 8045 E . | 36.38 | -2 01 | $\begin{array}{r}35.98 \\ \hline 40\end{array}$ | ${ }^{1 \cdot 27}$ | 35.53 37 | 5.79 .116 |
|  |  | $36 \cdot 40$ |  | 36.38 | $-1 \cdot 25$ | $35 \cdot 90$ | 5*:5 |
| V. 3..... | S. 100 W. | $19 \cdot 30$ | +15.08 |  |  | ${ }^{19.00 .}{ }^{\circ}{ }^{\circ}$ | $\begin{array}{r}0.83 \\ \hline 101\end{array}$ |
|  |  | 20.00 |  | $19 \cdot 39$ | +5 22 | $\overline{19.30}$ | 0.44 |
| V. 5...... ${ }_{\text {雄 }}$ | N. 8515 E. | 79.28 | $-630$ | \% 79$79 \% 9$ | 8.94 | -8:78 | 6.54 $\cdot 12$ |
|  |  | $79 \cdot 80$ |  |  | $-9 \cdot 03$ | 79.01 | 6.56 |
|  |  | 43.00 |  | ${ }^{42} \cdot 7.88$ | 11.96 19 |  |  |
| Close to D. 1. | S. 350 W . | $44 \cdot 70$ | +15 46 | $43 \cdot 00$ | $+12 \cdot 15$ | $42 \cdot 9$ | $2 \cdot 87$ |
| From 74 to $\oplus 53$. | S. 84-42 E. | 53.00 | -0.oั2 | 53.00 | -0.50 | 50:7 | $4 \cdot 90$ |
|  |  | \%5.97 |  | 75.78 | -13:63 | 2\% 286 | $19 \cdot 78$ -25 |
| From $\oplus 53$ to N . | N. 15-15 E. | $77 \cdot 20$ | $-10 \cdot 12$ | 7597 | $-18.67$ | 7830 | \% |


corrected horizontal distances placed over the others in red ink.

Problem. It is desired to drive the heading from H so that it will intersect the slope at N . Required, the course and distance. From the columns of total latitudes and departures in the sheet of calculations take :

$$
\begin{array}{cc}
\text { Latitude. } & \text { Departure } \\
\mathrm{N}=+2.74 .60 & -460 \cdot 98 \\
\mathrm{H}=+218 \cdot 42 & -244.60 \\
+56 \cdot 18 & -216.38
\end{array}
$$

Tangent, of course, equal departure divided by latitude.

$$
\begin{aligned}
& \text { log. } 216 \cdot 38=2 \cdot 33521 \% 1 \\
& \log \text {. } 56 \cdot 18=1 * \div 49581 \% \\
& \text { tem. } \mathfrak{\%}-2 \%=10 \cdot 5556354=\imath 5^{\circ}-2 \tilde{\imath}^{\prime}=\text { course. } \\
& \log .56 \cdot 18=1 \cdot ヶ 49581 \% \\
& \cos .75^{\circ} 2 \gamma^{\prime}=\underline{9 \cdot 4000625} \\
& 2 \cdot 3 \pm 95192=223 \cdot 62=\text { distance } .
\end{aligned}
$$

N , being north and west of H , shows the course to be N . T ., or N. $15^{\circ}-26 \mathrm{~T}$. $223 \cdot 36$.

Unless in special cases where great accuracy is required, the more common method of solving this and similar problems is to take the course and distance from the map with a protractor and scale, this being sufficiently accurate for all practical purposes.
759. Making the Map. If the map is to be much handled, use the best quality of cloth-backed paper. The edges should be bound with linen tape, which, if semed, should be double-stitched, with about three stitches to the inch. If the stitches are made closer than this, the binding will break off in the line of the needle-holes. Ascertain from existing maps, or whatever data mar be at hand, the most adrantageous direction for the meridian of the surrer to assume on the map. Fix also upon a point for the origin of coordinates. Begin at the origin and rule the paper into fire- or teninch squares, parallel with the meridian of the surres. Vers great care is required in doing this work, in order to make all the squares check precisely with the scale and be rectangular. Owing to the
expansion and contraction of the paper, the work of laying out the squares should be concluded on the same day it is started. In addition to the underground workings, the map should show all land-lines, dwellings, roads, streams, ponds of water, and any other features of the surface that may have a bearing on an intelligent working of the mine. Both surveys should be referred to the same origin of co-ordinates. In plotting an underground traverse, it is generally more convenient to locate only every fifth or tenth station by its co-ordinates, and use a protractor for filling in the balance.

Take a paper protractor, and letter it N. S. E. W., and fix it at any convenient place on the paper, so its N. and S. points will correspond with the meridian of the survey. Fasten with weights; then transfer the courses from the protractor to where they are wanted on the map, scaling off the distances as required. The stations that have been located by ordinates will check the slight errors in the plotting from the protractor. Having plotted all the courses, proceed to fill in the interior work from the references and sketches shown on the right-hand page of the note-book.

In inking the map, use only colors that will wash. A diluted solution of bichromate of potash mixed with India-ink will prevent spreading of the lines when touched with a wet tinting-brush.

The map should show all the survey-stations, stoppings of entrances, inclination of strata, and elevation of the stations above tide or other datum.

When different "levels" are to be represented, with their connecting shafts, etc., " isometrical projection" has been used, but " military or cavalier projection" is best.

## CHAPTER II.

## LOCATING NEW LINES.

760. Second Object. To determine, on the surface of the ground, where to sink a shaft to meet a desired point in the underground workings.

To do this, repeat on the surface of the ground the surrey made under it-i. e., trace on it the courses and distances of the galleries, or their equivalents (Art. 764).

The chief difficulty is to get a starting-point, and to determine the direction of the first line.
761. When the Mine is entered by an Adit (Fig. 564). Set the transit at the entrance, and get the direction of the adit, and prolong it up the

Fig. 564.
 hill-i. e., in the same vertical plane. The third adjustment is here important.

If the line has to be prolonged by setting the instrument farther on, the second adjustment is important.
762. When the Mine is entered by a Shaft. Get the magnetic bearing of the first underground line, at the bottom of the shaft, with great care. Bring up the end of the line through the shaft by a plumb-line, and set the compass over this point. Set out a
line with the same bearing and length as the first underground line, and repeat the succeeding courses.

When the compass can not be set over the point, proceed thus:

1. Find, by trial, a spot, as B (Fig. 565), which is in the correct course, and measure off a distance equal to the length of the first underground course, and then proceed as before.
2. Otherwise. Set up anywhere, as at $\mathrm{A}^{\prime}$ (Fig. 566), take the bearing and distance of $A$ from $\mathrm{A}^{\prime}$; run a line corresponding with the one underground, from $\mathrm{A}^{\prime}$ to $\mathrm{B}^{\prime}$. Repeat the
 course $A^{\prime} A$ from $B^{\prime} B$; then $A B$ is the desired line.
3. To dispense with the Magnetic Needle. First Method. Let down two plumb-lines on opposite sides of the shaft, so that their lower ends shall be

## Fig. 566.

 very precisely in the underground line (see Art. ${ }^{\text {7 }}$ 751).

Second Method. Set, by repeated trials, two transits on opposite sides of the shaft, so that they shall at the same time point to one another, and each, also, to one of two points in the underground line. They will then give the direction of the line above-ground.

Third Method. If the telescope of the transit be eccentric, as in Fig. 561, set the instrument on a platform over the mouth of the shaft, so that the line of collimation of the telescope shall be in the same vertical plane with two points in the underground line, on opposite sides of the shaft. When the instrument is so placed that, in turning the telescope, the intersection of the crosshairs strikes the two points in the underground line, the line of sight, when directed along the surface, will gire the required line.
764. Having determined the first line, the courses of the underground survey may be repeated on the surface ; or the bearing and length of a single line be calculated, which shall arrive at the desired point.

Let the zigzag line, A B, B C, C D, D Z (Fig. 56\%), be the courses surveyed underground, A being an adit, or at the bottom of a shaft, and Z the point to which it is desired to

Fig. 567.
 sink a shaft. It is required to find the direction and length of the straight line A Z.

When the compass is used, calculate the latitude and departure of each of the courses, A B, B C, etc. The algebraic sum of their latitudes will be equal to $\Lambda \mathrm{X}$, and the algebraic sum of their departures will be equal to $\mathrm{X} Z$. Then is $\tan . \mathrm{ZAX}=\frac{\mathrm{XZ}}{\mathrm{XA}}$; that is, the algebraic sum of the departures divided by the algebraic sum of the latitudes is equal to the tangent of the bearing. The length of the line $\mathrm{A} Z$ equals the square root of the sum of the squares of A X and XZ ; or equals the latitude divided by the cosine of the bearing.

When the transit is used, instead of referring all of the lines to the magnetic meridian, as in the preceding case, any line of the survey may now be taken as the meridian, as in "traversing."

In Fig. 568 all of the courses are referred to the first line of the survey. As before, a
 right-angled triangle will be formed.
Tan. $\mathrm{Z} \mathrm{A} \mathrm{X}=\frac{\mathrm{XZ}}{\overline{\mathrm{XA}}}$, and the length of $\mathrm{AZ}=\sqrt{\overline{\overline{\mathrm{AX}^{2}}+\overline{\mathrm{X}}^{2}}}$; or A $\mathrm{X} \div \cos . \mathrm{XAZ}$.

Two or more lines may be substituted for the single line in the two preceding cases; the condition being, that the algebraic sums of their latitudes and of their departures shall be equal to those of the underground survey.
765. Third Object. Tio direct the workings of a mine to any desired point.

This is the converse of the second object. We repeat under the ground the courses run above-ground ; or their equivalents, as in Art. 764.

In Fig. 569, let A B, B C, C D, D Y, be the present workings of a mine, and Z the shaft to which the workings are to be directed.

Find the latitude and departure of A Z. Then the difference between the algebraic sum of the latitudes of the underground courses already run, and the latitude of A Z, is the latitude of the required course ; and the difference between the algebraic sum of the departures of the underground lines, and the departure of $\mathrm{A} Z$, is the departure of the required course.

The length of $\mathrm{Y} Z$ equals the square root of the sum of the squares of its latitude and departure.

Fig. 569.

766. Problems. Most of the problems which arise in miningsurveying can be solved by an application of the familiar principles of geometry and trigonometry:

1. Given the angle which a vein makes with the horizon, and the place where it meets the sur-

Fig. 570.
 face, to find how deep a shaft at D will be required to strike the vein :

$$
\mathrm{DC}=\mathrm{AD} \cdot \tan \cdot \mathrm{DAC}
$$

2. Given the depth of the shaft D C, and the "dip" of the vein, to find where it crops out:

$$
\mathrm{AD}=\mathrm{DC} \cdot \cot . \mathrm{DAC}
$$

3. Given the depth of a shaft when the vein "crops out," and the "dip" of the vein, to find the distance from the bottom of the shaft to the vein :

$$
\mathrm{BC}=\mathrm{A} \mathrm{~B} \cdot \cot . \mathrm{ACB}
$$

If the ground makes an angle with the horizon, then the problems involve oblique-angled triangles instead of right-angled tri-
angles, as in the preceding cases. Their solution, however, is quite as simple.

In the more difficult problems, the measurement of lines is required, one or both ends of which are inaccessible. (For a full investigation of this subject, see Part I, Chapter V.)

## APPENDIX.

## APPENDIX A.

## SYNOPSIS OF PLANE TRIGONOMETRY.*

1. Definition. Plane Trigonometry is that branch of mathematical science which treats of the relations between the sides and angles of plane triangles. It teaches how to find any three of these six parts, when the other three are given, and one of them, at least, is a side.
2. Angles and Arcs. The angles of a triangle are measured by the arcs described, with any radius, from the angular points as centers, and intercepted between the legs of the angles. These ares are measured by comparing them with an entire circumference, described with the same radius. Every circumference is regarded as being divided into 360 equal parts, called degrees. Each degree is divided into 60 equal parts, called minutes, and each minute into 60 seconds. These divisions are indicated by the marks ${ }^{\circ}$ ' $\prime$. Thus 28 degrees, 17 minutes, and 49 seconds, are written $28^{\circ} 17^{\prime} 49^{\prime \prime}$ Fractions of a second are best expressed decimally. An arc, including a quarter of a circumference and measuring a right angle, is therefore $90^{\circ}$. A semicircumference comprises $180^{\circ}$. It is often represented by $\pi$, which equals $3 \cdot 14159$, etc., or $3 \frac{1}{7}$ approximately, the radius being unity.

The length of $1^{\circ}$ in parts of radius $=0.01745329 ;$ that of $1^{\prime}=0.00029089$; and that of $1^{\prime \prime}=0.00000485$.

The length of the radius of a circle in degrees, or 360ths of the circumference $=57 \cdot 29578^{\circ}=57^{\circ} 17^{\prime}$ $24 \cdot 8^{\prime \prime}=3437 \cdot 747^{\prime}=206264 \cdot 8^{\prime \prime} . \dagger$

An arc may be regarded as generated by a point, M , moving from an origin, A , around a circle, in the direction of the arrow. The point may thus describe

Fig. 571.
 arcs of any lengths, such as $\mathrm{AM} ; \mathrm{AB}=90^{\circ}=\frac{1}{3}$ $\pi ; \mathrm{A} \mathrm{B} \mathrm{C}=180^{\circ}=\pi ; \mathrm{ABCD}=270^{\circ}=\frac{3}{2} \pi ; \mathrm{ABCDA}=360^{\circ}=2 \pi$.

The point may still continue its motion, and generate ares greater than a

[^73]circumference, or than two circumferences, or than three; or even infinite in length.

While the point, M, describes these arcs, the radius, OM, indefinitely produced, generates corresponding angles.

If the point, $M$, should move from the origin, $A$, in the contrary direction to its former morement, the arcs generated by it are regarded as negatice, or minus ; and so too, of necessity, the angles measured by the arcs.

Arcs and angles may therefore vary in length from 0 to $+\infty$ in one direction, and from 0 to $-\infty$ in the contrary direction.

The Complement of an arc is the arc which would remain after subtracting the arc from a quarter of the circumference, or from $90^{\circ}$. If the are be more than $90^{\circ}$, its complement is necessarily negative.

The Suppiement of an arc is what would remain after subtracting it from half the circumference, or from $180^{\circ}$. If the are be more than $180^{\circ}$, its supplement is necessarily negative.
3. Trigonometrical Lines. The relations of the sides of a triangle to its angles are what is required ; but it is more convenient to replace the angles by arcs; and, once more, to replace the arcs by certain straight lines depending upon them, and increasing and decreasing with them, or, conrersely, in such a way that the length of the lines can be found from that of the arcs, and vice versa. It is with these lines that the sides of a triangle are compared.* These lines are called Trigonometrical Lines, or Circular Functions, because their length is a function of that of the circular arcs. The principal trigonometrical lines are Sines, Tangents, and Secants. Chords and rersed sines are also used.

The SINE of an are, AM, is the perpendicular, M P, let fall, from one extremity of the arc, upon the diameter which

Fig. 572.
 passes through the other extremity.

The TANGENT of an are, A M, is the distance, A T, intercepted, on the tangent dramn at one extremity of the arc, between that extremity and the prolongation of the radius which passes through the other extremity.

The SECANT of an arc, AM, is the part, O T , of the prolonged radius, comprised between the center and the tangent.

The sine, tangent, and secant of the complement of an are are called the Co-sine, Co-tangent, and Co-secant of that arc. Thus, $M Q$ is the cosine of $A M, B S$ its cotangent, and $O S$ its cosecant. The cosine $M Q$ is equal to $O P$, the part of the radius comprised between the center and the foot of the sine.

The chord of an are is equal to twice the sine of half that arc.
The versed-sine of an arc, AM, is the distance, A P, comprised betreen the origin of the arc and the foot of the sine. It is consequently equal to the difference between the radius and the sine.

[^74]The trigonometrical lines are usually written in an abbreviated form. Calling the arc $\mathrm{A} \mathrm{M}=a$, we write,

$$
\begin{array}{lll}
\mathrm{M} \mathrm{P}=\sin . a . & \mathrm{A} \mathrm{~T}=\tan . a . & \text { O T }=\text { sec. } a . \\
\mathrm{MQ}=\cos . a . & \mathrm{B} \mathrm{~S}=\cot . a . & \mathrm{OS}=\operatorname{cosec} . a .
\end{array}
$$

The period after sin., tan., etc., indicating abbreviation, is frequently omitted.

The ares whose sines, tangents, etc., are equal to a line $=\alpha$, are written,

$$
\begin{aligned}
& \sin . a, \text { or } \operatorname{arc}(\sin .=a) ; \\
& \tan . a, \text { or } \operatorname{arc}(\tan .=a) ; \text { etc. }
\end{aligned}
$$

4. The Lines as Ratios. The ratios between the trigonometrical lines and the radius are the same for the same angles, or number of degrees in an are, whatever the length of the radius or arc. Consequently, radius being unity, these lines may be expressed as simple ratios. Thus, in the right-angled triangle A B C, we would have

$$
\begin{array}{ll}
\sin . A=\frac{B C}{A} \frac{\text { opposite side }}{\text { hypotenuse }}, & \text { cos. } A=\frac{A C}{A B}=\frac{\text { adjacent side }}{\text { hypotenuse }}, \\
\tan A=\frac{B C}{A C}=\frac{\text { opposite side }}{\text { adjacent side }}, & \text { cot. } A=\frac{A C}{B} \frac{\text { adjacent side }}{C}=\frac{\text { adposite side }}{\text { oppotende }} \\
\text { sec. } A=\frac{A B}{A C}=\frac{\text { hypotenuse }}{\text { adjacent side }}, & \text { cosec. } A=\frac{A B}{B C}=\frac{\text { hypotenuse }}{\text { opposite side }}
\end{array}
$$

When the radius of the ares which measure the angles is unity, these ratios may be used for the lines. If the radius be any other length, the results which have been obtained by the above supposition must be modified by dividing each of the trigonometrical lines in the result by radius, and thus rendering the equations of the results "homogeneous." The same effect would be produced by multiplying each term in the expression by such a power of radius as would make it contain a number of linear factors equal to the greatest number in any term.

Fig. 574.
 The radius is usually represented by $r$, or R .

## 5. Their Variations in Length.

As the point $M$ moves around the circle, and the are thus increases, the sines, tangents, and secants, starting from zero, also increase; till, when the point $M$ has arrived at $B$, and the are has become $90^{\circ}$, the sine has become equal to radius, or unity, and the tangent and secant have become infinite. The complementary lines have decreased, the cosine being equal to radius or unity at starting and becoming zero, and the cotangent and cosecant passing from infinity to zero.

When the point $M$ has passed the first quadrant at $B$, and is proceeding toward C , the sines, tangents, and secants begin to decrease, till, when the point has reached C, they have the same values as at A. They then begin to increase again, and so on. The table on page 527 indicates these variations.

The sines and tangents of very small ares may be regarded as sensibly proportional to the arcs themselves; so that for sin. $a^{\prime \prime}$, we may write $a$. sin. $1^{\prime \prime}$; and similarly, though less accurately, for $\sin . a^{\prime}$, we may write $a$. sin. $1^{\prime}$.

The sines and tangents of very small arcs may similarly be regarded as sensibly of the same length as the ares themselves.*
$a$ being the length of any arc expressed in parts of radius, the lengths of its sine and cosine may be obtained by the following series:

$$
\begin{aligned}
\sin . a & =a-\frac{a^{3}}{2 \cdot 3}+\frac{a^{5}}{2 \cdot 3 \cdot 4 \cdot 5}-\frac{a^{7}}{2 \cdot 3 \ldots \cdot 7}+, \text { etc. } \\
\operatorname{cos.} a & =1-\frac{a^{2}}{2}+\frac{a^{4}}{2 \cdot 3 \cdot 4}-\frac{a^{6}}{2 \ldots .6}+, \text { etc. }
\end{aligned}
$$

Let it be required to find cos. $30^{\circ}$, by the above series.

$$
30^{\circ}=\frac{30}{180} \pi=\frac{1}{6} \times 3.1416=5236 .
$$

Substituting this number for $a$, the series becomes, taking only three terms of it,

$$
1-\frac{(\cdot 5236)^{2}}{2}+\frac{(\cdot 5236)^{4}}{24}-, \text { etc. }=1-0.137078+0.003130=.866052 ;
$$

which is 'the correct value of cos. $30^{\circ}$ for the first four places of decimals.
The lengths of the other lines can be obtained from the mutual relations given in Art. 7. Some particular values are given belor:

$$
\begin{array}{lll}
\sin .30^{\circ}=\frac{1}{2} & \sin .45^{\circ}=\frac{1}{2} \vee 2 . & \sin .60^{\circ}=\frac{1}{2} \sqrt{ } .3 \\
\tan .30^{\circ}=\frac{1}{3} \sqrt{ } 3 . & \tan .45^{\circ}=1 . & \tan .60^{\circ}=\sqrt{ } . \\
\text { sec. } 30^{\circ}=\frac{2}{3} \sqrt{ } 3 . & \sec .45^{\circ}=\sqrt{ } . & \text { sec. } 60^{\circ}=2
\end{array}
$$

6. Their Changes of Sign. Lines measured in contrary directions from a common origin usually receive contrary algebraic signs. If, then, all the lines in the first quadrant are called positire, their signs will change in some of the other quadrants. Thus the sines in the first quadrant being all measured upward, when they are measured downward, as they are in the third and fourth quadrants, they will be negative. The cosines in the first quadrant are measured from left to right, and when they are measured from right to left, as in the second and third quadrants, they will be negative. The tangents and secants follow similar rules.

The variations in length and the changes of sign are all indicated in the following table, radius being unity. The terms "increasing" and "decreasing " apply to the lengths of the lines without any reference to their signs:

[^75]Lengths and Signs of the Trigonometrical Lines for Arcs from $0^{\circ}$ to $360^{\circ}$.

| Arcs. | $0^{\circ}$ | Between $0^{\circ}$ and $90^{\circ}$. | $90^{\circ}$ | Between $90^{\circ}$ and $180^{\circ}$. | $180^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| Sine. . | 0 | +, and increasing, | +1 | +, and decreasing, | 0 |
| Tangent. | 0 | +, and increasing, | $\pm \infty$ | -, and decreasing, | 0 |
| Secant. | +1 | +, and increasing, | $\pm \infty$ | -, and decreasing, | -1 |
| Cosine. | +1 | +, and decreasing, | 0 | -, and increasing, | -1 |
| Cotangent. | $\pm \infty$ | +, and decreasing, | 0 | -, and increasing, | $\mp \infty$ |
| Cosecant. | $\pm \infty$ | +, and decreasing, | +1 | +, and increasing, | $\pm \infty$ |


| Arcs. | $180^{\circ}$ | Between $180^{\circ} \mathrm{and} 270^{\circ}$. | $270{ }^{\circ}$ | Between $2700^{\circ} \mathrm{AND} 360^{\circ}$. | $360^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| Sine. | 0 | -, and increasing, | -1 | -, and decreasing, | 0 |
| Tangent.. | 0 | +, and increasing, | $\pm \infty$ | -, and decreasing, | 0 |
| Secant | -1 | -, and increasing, | $\mp \infty$ | +, and decreasing, | +1 |
| Cosine | -1 | -, and decreasing, | 0 | +, and increasing, | +1 |
| Cotangent | 干 $\infty$ | +, and decreasing, | 0 | -, and increasing, | $\mp \infty$ |
| Cosecant. | $\pm \infty$ | -, and decreasing, | -1 | -, and increasing, | $\mp \infty$ |

From this table, and Fig. 574, we see that an arc and its supplement have the same sine; and that their tangents, secants, cosines, and cotangents are of equal length but of contrary signs; while the cosecants are the same in both length and sign.

We also deduce from the figure the following consequences:

$$
\begin{array}{lc}
\sin .\left(a^{\circ}+180^{\circ}\right)=-\sin . a^{\circ} . & \cos \left(a^{\circ}+180^{\circ}\right)=-\cos . a^{\circ} . \\
\tan .\left(a^{\circ}+180^{\circ}\right)=\tan . a^{\circ} . & \cot \left(a^{\circ}+180^{\circ}\right)=\cot . a^{\circ} . \\
\text { sec. }\left(a^{\circ}+180^{\circ}\right)=-\sec a^{\circ} . & \operatorname{cosec}\left(a^{\circ}+180^{\circ}\right)=-\operatorname{cosec} a^{\circ} . \\
\sin .\left(-a^{\circ}\right)=-\sin . a^{\circ} & \cos \left(-a^{\circ}\right)=\cos \cdot a^{\circ} . \\
\tan .\left(-a^{\circ}\right)=-\tan . a^{\circ} & \cot \left(-a^{\circ}\right)=-\cot a^{\circ} . \\
\sec .\left(-a^{\circ}\right)=\sec . a^{\circ} . & \operatorname{cosec} .\left(-a^{\circ}\right)=-\operatorname{cosec} . a^{\circ} .
\end{array}
$$

An infinite number of arcs have the same trigonometrical lines; for, an arc $a$, the same arc plus a circumference, the same arc plus two circumferences, and so on, would have the same sine, etc.
"To bring back to the first quadrant" the trigonometrical lines of any large arc, proceed thus: Let $1029^{\circ}$ be an arc the sine of which is desired. Take from it as many times $360^{\circ}$ as possible. The remainder will be $309^{\circ}$. Then we shall have sin. $309^{\circ}=\sin .\left(180^{\circ}-309^{\circ}\right)=\sin .-129^{\circ}=-\sin$. $129^{\circ}=-\sin .\left(180^{\circ}-129^{\circ}\right)=-\sin .51^{\circ}$.
7. Their Mutual Relations. Radius being unity,

$$
\begin{aligned}
\tan . a^{\circ}=\frac{\sin . a^{\circ}}{\cos \cdot a^{\circ}} & \cot a^{\circ}=\frac{\cos . a^{\circ}}{\sin \cdot a^{\circ}} . \\
\text { sec. } a^{\circ}=\frac{1}{\cos . a^{\circ}} & \operatorname{cosec} . a^{\circ}=\frac{1}{\sin . a^{\circ}} . \\
\tan . a^{\circ} \times \cot . a^{\circ}=1 . & \left(\sin . a^{\circ}\right)^{2}+\left(\cos a^{\circ}\right)^{2}=1 . * \\
1+\left(\tan . a^{\circ}\right)^{2}=\left(\sec a^{\circ}\right)^{2} . & 1+\left(\cot a^{\circ}\right)^{2}=\left(\operatorname{cosec} a^{\circ}\right) .
\end{aligned}
$$

$$
\tan a^{\circ} \times \cot a^{\circ}=1
$$

[^76]Hence, any one of the trigonometrical lines being giren, the rest can be found from some of these equations.
8. Two Arcs. Let $a$ and $b$ represent any two arcs, $a$ being the greater. Then the following formulas apply:

$$
\begin{aligned}
& \sin .(a+b)=\sin \cdot a \cdot \cos . b+\cos \cdot a \cdot \sin . b . \\
& \sin .(a-b)=\sin . a \cdot \cos . b-\cos . a \cdot \sin . b \text {. } \\
& \cos .(a+b)=\cos \cdot a \cdot \cos . b-\sin . a \cdot \sin . b \text {. } \\
& \cos .(a-b)=\cos . a \cdot \cos . b+\sin . a \cdot \sin . b . \\
& \tan .(a+b)=\frac{\tan \cdot a+\tan \cdot b}{1-\tan \cdot a \cdot \tan . b} . \\
& \tan .(a-b)=\frac{\tan . a-\tan . b}{1+\tan \cdot a \cdot \tan \cdot b} . \\
& \text { cot. }(a+b)=\frac{\text { cot. } a \cdot \cot \cdot b-1}{\cot . b+\cot \cdot a} . \\
& \cot .(a-b)=\frac{\cot \cdot a \cdot \cot \cdot b+1}{\cot . b-\cot \cdot a} . \\
& \sin . a \cdot \sin . b=\frac{1}{2} \cdot \cos \cdot(a-b)-\frac{1}{2} \dot{\cos }(a+b) \text {. } \\
& \cos \cdot a \cdot \cos \cdot b=\frac{1}{2} \cdot \cos \cdot(a+b)+\frac{1}{2} \cos \cdot(a-b) . \\
& \sin . a \cdot \cos \cdot b=\frac{1}{2} \cdot \sin \cdot(a+b)+\frac{1}{2} \sin \cdot(a-b) \text {. } \\
& \cos \cdot a \cdot \sin \cdot b=\frac{1}{2} \cdot \sin \cdot(a+b)-\frac{1}{2} \sin \cdot(a-b) . \\
& \sin . a+\sin . b=2 \sin . \frac{1}{2}(a+b) \cos \cdot \frac{1}{2}(a-b) \text {. } \\
& \cos . a+\cos . b=2 \cos \cdot \frac{1}{2}(a+b) \cos \cdot \frac{1}{2}(a-b) \text {. } \\
& \sin . a-\sin . b=2 \sin \cdot \frac{1}{2}(a-b) \cos . \frac{1}{2}(a+b) \text {. } \\
& \cos . b-\cos . a=2 \sin . \frac{1}{2}(a-b) \sin . \frac{1}{2}(a+b) \text {. } \\
& \tan . a+\tan . b=\frac{\sin \cdot(a+b)}{\cos \cdot a \cdot \cos \cdot b} \text {. } \\
& \tan . a-\tan . b=\frac{\sin .(a-b)}{\cos . a \cdot \cos . \bar{b}} . \\
& \cot . b+\cot . a=\frac{\sin .(a+b)}{\sin \cdot a \cdot \sin . b} \text {. } \\
& \cot . b-\cot . a=\frac{\sin .}{\sin . a-b)} \text {. }
\end{aligned}
$$

9. Double and Half Arcs. Letting $a$ represent any arc, as before, we have the following formulas:

$$
\begin{aligned}
& \sin .2 a=2 \sin \cdot a \cdot \cos \cdot a . \\
& \cos .2 a=(\cos \cdot a)^{2}-(\sin \cdot a)^{2}=2(\cos \cdot a)^{2}-1=1-2(\sin \cdot a)^{2} . \\
& \tan .2 a=\frac{2 \tan \cdot a}{1-(\tan \cdot a)^{2}}=\frac{2 \cot \cdot a}{(\cot . a)^{2}-1}=\frac{2}{\cot \cdot a-\tan \cdot a} . \\
& \cot .2 a=\frac{(\cot \cdot \alpha)^{2}-1}{2 \cot \cdot a}=\frac{1}{2}(\cot \cdot a-\tan \cdot a) .
\end{aligned}
$$

number of the degrees in the arc, thus: $\operatorname{Sin} .^{2} a^{\circ}$, $\tan .^{2} a^{\circ}$, etc. But the notation giren above places the index as used by Gauss, Delambre, Arbogast, ete., though the first two omit the parentheses.
$\sin \cdot \frac{1}{2} a=V\left[\frac{1}{2}(1-\cos . a)^{\prime}\right.$.
$\cos . \frac{1}{2} a=V\left[\frac{1}{2}(1+\cos . u)\right]$.
$\tan . \frac{1}{2} a=\frac{\sin . a}{1+\cos . a}=\frac{1-\cos \cdot a}{\sin . a}=\vee\left(\frac{1-\cos \cdot \alpha}{1+\cos \cdot a}\right)$.
cot. $\frac{1}{2} a=\frac{1+\cos \cdot a}{\sin . a}=\frac{\sin . a}{1-\cos . a}=V\left(\frac{1+\cos \cdot a}{1-\cos . a}\right)$.
10. Trigonometrical Tables. In the usual tables of the natural trigonometrical lines, the degrees from $0^{\circ}$ to $45^{\circ}$ are found at the top of the table, and those from $45^{\circ}$ to $90^{\circ}$ at the bottom; the latter being complements of the former. Consequently, the columns which have Sine and Tangent at top have Cosine and Cotangent at bottom, since the cosine or cotangent of any arc is the same thing as the sine or tangent of its complement. The minutes to be added to the degrees are found in the left-hand column, when the number of degrees at the top of the page are used, and in the right-hand column for the degrees when at the bottom of the page. The lines for arcs intermediate between those in the tables are found by proportion. The lines are calculated for a radius equal unity. Hence, the valnes of the sines and cosines are decimal fractions, though the point is usually omitted. So too are the tangents from $0^{\circ}$ to $45^{\circ}$, and the cotangents from $90^{\circ}$ to $45^{\circ}$. Beyond those points they are integers and decimals.

The calculations, like all others involving large numbers, are shortened by the use of logarithms, which substitute addition and subtraction for multiplication and division; but the young student should avoid the frequent error of regarding logarithms as a necessary part of trigonometry.

## SOLUTION OF TRIANGLES.

## 11. Right-angled Triangles. Let A B C

 be any right-angled triangle. Denote the sides opposite the angles by the corresponding small letters. Then any one side and one acute angle, or any two sides being given, the other parts can be obtained by one of the following equations:Fig. 575.


| GIVEN. | REqUIRED. | formulas. |
| :---: | :---: | :---: |
| $a, b$ | $c, \mathrm{~A}, \mathrm{~B}$ | $c=\sqrt{ }\left(a^{2}+b^{2}\right) ; \tan . \mathrm{A}=\frac{a}{b} ; \cot . \mathrm{B}=\frac{a}{b}$. |
| $a, c$ | $b, \mathrm{~A}, \mathrm{~B}$ | $b=\sqrt{ }\left(c^{2}-a^{2}\right) ; \sin . \mathrm{A}=\frac{a}{c} ; \cos . \mathrm{B}=\frac{a}{c}$. |
| $a, \mathrm{~A}$ | $b, c, \mathrm{~B}$ | $b=a . \cot . \mathrm{A} ; c=\frac{a}{\sin . \mathrm{A}} ; \mathrm{B}=90^{\circ}-\mathrm{A}$. |
| $b, \mathrm{~A}$ | $a, c, \mathrm{~B}$ | $a=b . \tan . \mathrm{A} ; c=\frac{b}{\cos . \mathrm{A}} ; \mathrm{B}=90^{\circ}-\mathrm{A}$. |
| $c, \mathrm{~A}$ | $a, b, \mathrm{~B}$ | $a=c \cdot \sin . \mathrm{A} ; b=c \cos . \mathrm{A} ; \mathrm{B}=90^{\circ}-\mathrm{A}$. |

12. Oblique-angled Triangles. Let ABC be any oblique-angled triangle, the angles and sides being noted as in the figure. Then any three of its six parts being given, and one of them
 being a side, the other parts can be obtained by one of the following methods, which are founded on these three theorems:

Theorem I.- In every plane triangle, the sines of the angles are to each other as the opposite sides.
Theorem II.-In every plane triangle, the sum of two sides is to their difference as the tangent of half the sum of the angles opposite those sides is to the tangent of half their difference.

Theorem III.-In every plane triangle, the cosine of any angle is equal to a fraction whose numerator is the sum of the squares of the sides adjacent to the angle, minus the square of the side opposite to the angle, and whose denominator is twice the product of the sides adjacent to the angle.

All the cases for solution which can occur may be reduced to four:
Case 1.-Given a side and two angles. The third angle is obtained by subtracting the sum of the two given angles from $180^{\circ}$. Then either unknown side can be obtained by Theorem I.

Calling the given side $a$, we have $b=a \cdot \frac{\sin . \mathrm{B}}{\sin . \mathrm{A}}$; and $c=a \frac{\sin . \mathrm{C}}{\sin . \mathrm{A}}$.
Case 2.-Given two sides and an angle opposite one of them. The angle opposite the other given side is found by Theorem I. The third angle is obtained by subtracting the sum of the other two from $180^{\circ}$. The remaining side is then obtained by Theorem I.

Calling the given sides $a$ and $b$, and the given angle A , we hare $\sin . \mathrm{B}=$ $\sin . \mathrm{A} \cdot \frac{b}{a}$.

Since an angle and its supplement have the same sine, the result is ambiguous; for the angle B may have either of the two supplementary values indicated by the sine, if $b>a$, and A is an acute angle.

$$
\mathrm{C}=180^{\circ}-(\mathrm{A}+\mathrm{B}) . \quad c=\sin . \mathrm{C} \frac{a}{\sin . \mathrm{A}}
$$

Case 3.-Given two sides and their included angle. Applying Theorem II (obtaining the sum of the angles opposite the giren sides by subtracting the given included angle from $180^{\circ}$ ), we obtain the difference of the unknown angles. Adding this to their sum we obtain the greater angle, and subtracting it from their sum we get the less. Then Theorem I will give the remaining side.

Calling the given sides $a$ and $b$, and the included angle $C$, we hare $\mathrm{A}+\mathrm{B}=180^{\circ}-\mathrm{C}$. Then

$$
\tan \cdot \frac{1}{2}(\mathrm{~A}-\mathrm{B})=\tan \cdot \frac{1}{2}(\mathrm{~A}+\mathrm{B}) \cdot \frac{a-b}{a+b}
$$

$\frac{1}{2}(\mathrm{~A}+\mathrm{B})+\frac{1}{2}(\mathrm{~A}-\mathrm{B})=\mathrm{A} \cdot \quad \frac{1}{2}(\mathrm{~A}+\mathrm{B})-\frac{1}{2}(\mathrm{~A}-\mathrm{B})=\mathrm{B} . \quad c=a \frac{\sin \mathrm{C}}{\sin \cdot \mathrm{A}}$.

In the first equation cot. $\frac{1}{2} \mathrm{C}$ may be used in the place of $\tan . \frac{1}{2}(\mathrm{~A}+\mathrm{B})$. Case 4.-Given the three sides. Let $s$ represent half the sum of the three sides $=\frac{1}{2}(a+b+c)$. Then any angle, as A, may be obtained from either of the following formulas, founded on Theorem III:

$$
\begin{aligned}
\sin . \frac{1}{2} \mathrm{~A} & =\sqrt{ }\left[\frac{(s-b)(s-c)}{b c}\right] \\
\cos \cdot \frac{1}{2} \mathrm{~A} & =\sqrt{ } /\left[\frac{s(s-a)}{b c}\right] \\
\tan \cdot \frac{1}{2} \mathrm{~A} & =\sqrt{c}\left[\frac{(s-b)(s-c)}{s(s-a)}\right] \\
\sin . \mathrm{A} & =\frac{2 \mathcal{V}^{\prime}[s(s-a)(s-b)(s-c)]}{b c} \\
\cos \mathrm{~A} & =\frac{b^{2}+c^{2}-a^{2}}{2 b c}
\end{aligned}
$$

The first formula should be used when $\mathrm{A}<90^{\circ}$, and the second when A $>90^{\circ}$. The third should not be used when A is nearly $180^{\circ}$; nor the fourth when A is nearly 90 ; nor the fifth when A is very small. The third is the most convenient, when all the angles are required.

## APPENDIX B.

## TRANSIERSALS.

Theorem I.-If a straight line be drawn so as to cut any two sides of a triangle, and the third side prolonged, thus dividing them into six parts (the prolonged side and its prolongation being two of the parts), then will the product of any three of those parts, whose extremities are not contiguous, equal the product Fig. 577. of the other three parts.

That is, in Fig. ort, $^{7}$ A B C being the triangle, and DF the transrersal, $\mathrm{BE} \times \mathrm{AD} \times \mathrm{CF}=$ $\mathrm{EA} \times \mathrm{DC} \times \mathrm{BF}$.

To prove this, from $B$ draw $B G$, parallel to C A. From the similar triangles BE G and AED, we have BG:BE::AD:AE. From the similar triangles $B F G$ and $C F D$, we have
 CD : CF:: B G: BF. Multiplying these proportions tngether, we hare $B G \times C D: B E \times C F:: A D \times B G: A E \times B F$. Multiplying estremes and means, and suppressing the common factor B G, we have $\mathrm{BE} \times \mathrm{AD} \times$ $\mathrm{CF}=\mathrm{EA} \times \mathrm{DU} \times \mathrm{BF}$.

These six parts are sometimes said to be in incolution.
If the transrersal passes entirely outside of the triangle and cuts the prolongations of all three sides, as in Fig. 578, the theorem still holds good. The same demonstration applies.without any change.*

Theorem II.-Conversely : If three
 points be taken on two sides of a triangle, and on the third side prolonged, or on the prolongations of the three sides, dividing them into six parts, such that the product of three non-consecutice parts equals the product of the other three parts, then will these three points lie in the same straight line.

This theorem is proved by a reductio ad absurdum.

Theorem III.-If, from the summits

* This theorem may be extended to polygons.
of a triangle, lines be drawn, to a point situated either within or without the triangle, and prolonged to meet the sides of the triangle, or their prolongations, thus dividing them into six parts, then will the product of any three non-consecutive parts be equal to the product of the other three parts.

That is, in Fig. 579, or Fig. 580, $A E \times B F \times C D=E B \times F C \times D A$.
For, the triangle ABF, being cut by the transver-
 sal E C, gives the relation (Theorem I).

Fig. 5 s0.

$A E \times B C \times F P=E B \times F C \times P A$.
The triangle ACF, being cut by the transversal D B, gives $\mathrm{DC} \times \mathrm{FB} \times \mathrm{PA}=\mathrm{AD} \times \mathrm{CB} \times$ FP.
Multiplying these equations together, and suppressing the common factors P A, C B, and F P, we have $A E \times B F \times C D=E B \times F C \times$ D A.

Theorem IV. - Conversely: If three points are situated on the three sides of a triangle, or on their prolongations (either one, or three, of these points being on the sides), so that they divide these lines in such a way that the product of any three non-consecutive parts equals the product of the other three parts, then will lines drawn from these points to the opposite angles meet in the same point.

This theorem can be demonstrated by a reductio ad absurdum.

## corollaries of the preoeding theorems.

Corollary 1.-The MEDIANS of a triangle (i. e., the lines drawn from its summits to the middles of the opposite sides) meet in the same point.

For, supposing, in Fig. 579, the points D, E, and F to be the middles of the sides, the products of the non-consecutive parts will be equal-i. e., $\mathrm{AE} \times \mathrm{BF} \times \mathrm{CD}=\mathrm{DA} \times \mathrm{EB} \times \mathrm{FC}$; since $\mathrm{AE}=\mathrm{EB}, \mathrm{BF}=\mathrm{FC}, \mathrm{CD}$ $=\mathrm{D} A$. Then Theorem IV applies.

Cor. 2.-The BISSECTRICES of a triangle (i. e., the lines bisecting its angles) meet in the same point.

For, in Fig. 579, supposing the lines A F, B D, C E to be bissectrices, we have (Legendre, IV, 17) :

Multiplying these equations together, and omitting the common factors, we have $\mathrm{BF} \times \mathrm{CD} \times \mathrm{A} \mathrm{E}=\mathrm{FC} \times \mathrm{DA} \times \mathrm{EB}$. Then Theorem IV applies.

Cor. 3.-The ALTITUDES of a triangle (i. e., the lines drawn from its summits perpendicular to the opposite sides) meet in the same point.

For, in Fig. 579, supposing the lines AF, B D, and CE to be altitudes, we have three pairs of similar triangles, B C D and FCA, CAE and D A B, $A B F$ and EBC, by comparing which we obtain relations from which it is easy to deduce $\mathrm{BF} \times \mathrm{CD} \times \mathrm{AE}=\mathrm{EB} \times \mathrm{FC} \times \mathrm{DA}$; and then Theorem IV again applies.

Cor. 4.-If, in Fig. 579, or Fig. 580, the point F be taken in the middle of B C , then will the line E D be parallel to BC .

For, since $\mathrm{BF}=\mathrm{FC}$, the equation of Theorem III reduces to $\mathrm{A} \mathrm{E} \times \mathrm{CD}$ $=\mathrm{EB} \times \mathrm{DA}$; whence $\mathrm{AE}: \mathrm{EB}:: \mathrm{AD}: \mathrm{DC}$; cousequently E D is parallel to BC.

Cor. 5.-Conversely: If E D be parallel to B C , then is $\mathrm{B} \mathrm{F}=\mathrm{FC}$.
For, since $A E: E B:: A D: D C$, we have $A E \times D C=E B \times A D$; whence, in the equation of Theorem III, we must have $\mathrm{BF}=\mathrm{FC}$.

Cor. 6.-From the preceding corollary, we derive the following:
If two sides of a triangle are divided proportion-

Fig. 581.
 ally, starting from the same summit, as A , and lines are drawn from the extremities of the third side to the points of division, the intersections of the corresponding lines will all lie in the same straight line joining the summit A , and the middle of the base.

Cor. 7.-A particular case of the preceding corollary is this:

In any trapezoid, the straight line which joins the intersection of the diagonals and the point of meeting of the non-parallel sides produced, passes through the middle of the two parallel bases.

Cor. 8.-If the three lines drawn through the corresponding summits of two triangles cut each other in the same point, then the three points in which the corresponding sides, produced if necessary, will meet, are situated in the same straight line.

This corollary may be otherwise enunciated, thus:
If two triangles have their summits situated, two and two, on three lines which meet in the same point, then, etc.

This is proved by obtaining by Theorem I three equations, which, being multiplied together, and the six common factors canceled, gire an equation to which Theorem II applies.

Triangles thus situated are called homologic ; the common point of meeting of the lines passing through their summits is called the center of homology; and the one on which the sides meet, the axis of homology.

## HARMONIC DIVISION.

Fig. 582.


Defintions.-A straight line, A B, is said to be harmonically divided at the points C and D , when these points determine tro additive segments, A C, B C, and two sub-
tractive segments, $\mathrm{AD}, \mathrm{BD}$, proportional to one another; so that $\Lambda \mathrm{C}: \mathrm{BC}$ $:: A D: B D$. It will be seen that A C must be more than BC, since A D is more than B D.*

This relation may be otherwise expressed, thus: The product of the whole line by the middle part equals the product of the extreme parts.

Reciprocally, the line D C is harmonically divided at the points B and A , since the preceding proportion may be written $\mathrm{DB}: \mathrm{CB}:: \mathrm{DA}: \mathrm{CA}$.

The four points, A, B, C, D, are called harmonics. The points C and D are called harmonic conjugates. So are the points A and B .

When a straight line, as A B, is divided harmonically, its half is a mean proportional between the distance from the middle of the line to the two points, C and D , which divide it harmonically.

If, from any point, $O$, lines be drawn so as to divide a line harmonically,

Fig. 583.
 these lines are called an harmonic pencil. The four lines which compose it, O A, O C, $\mathrm{OB}, \mathrm{O}$, in the figure, are called its radii, and the pairs which pass through the conjugate points are called conjugate radii.

Theorem V.-In any harmonic pencil, a line drawn parallel to any one of the radii is divided by the three other radii into two equal parts.
Let EF be the line, drawn parallel to O A. Through B draw G H, also parallel to OA. We have,
$\mathrm{GB}: \mathrm{OA}:$ : BD:AD; and
Fig. 584.
BH:OA: : BC:AC.
But, by hypothesis, AC:BC::A乌゙: B D.

Hence, the first two proportions reduce to. $\mathrm{GB}=\mathrm{BH}$; and, consequent$\mathrm{ly}, \mathrm{EK}=\mathrm{K} \mathrm{F}$.

The reciprocal is also true-i. e.,
If four lines radiating from a point are such that a line drawn parallel
 to one of them is divided into two equal parts by the other three, the four lines form an harmonic pencil.

Theorem VI.-If any transversal to an harmonic pencil be drawn, it will be divided harmonically.

Let LM be the transversal. Through K , where L M intersects O B, draw E F parallel to OA . It is bisected at K by the preceding theorem; and the

[^77]similar triangles, FMK and LMO, EKN and LNO, give the proportions

LM: KM: OL: FK, and LN:NK: OL:EK; whence, since $\mathrm{F} \mathrm{K}=\mathrm{E} \mathrm{K}$, we have $\mathrm{L} \mathrm{N}: \mathrm{N} \mathrm{K}:: \mathrm{LM}: \mathrm{KM}$.
Corollary.-The two sides of any angle, together with the bissectrices of the angle and of its supplement, form an harmonic pencil.

Theorem VII.-If, from the summits of any triangle, A B C, through any point, P, there be drawn the transversals $\mathrm{A}, \mathrm{BE}, \mathrm{C} \mathrm{F}$, and the transversal E D be drawon to meet A B prolonged in $\mathrm{F}^{\prime}$, the points F and $\mathrm{F}^{\prime}$ will divide the base A B harmonically.

Fig. $\check{6}$ 8̌.


This may be otherwise expressed, thus:
The line, C P, which joins the intersection of the diagonals of any quadrilateral, A B D E, with the point of meeting, C, of two opposite sides prolonged, cuts the side A B in a point F , which is the harmonic conjugate of the point of meeting, $\mathrm{F}^{\prime}$ of the other twoo sides, E D and A B , prolonged.

For, by Theorem $I, \mathrm{AF}^{\prime} \times \mathrm{BD} \times \mathrm{CE}=\mathrm{F}^{\prime} \mathrm{B} \times \mathrm{DC} \times \mathrm{EA}$; and
by Theorem III, $\mathrm{AF} \times \mathrm{BD} \times \mathrm{CE}=\mathrm{FB} \times \mathrm{DC} \times \mathrm{EA}$;
whence $A F: F B:: F^{\prime}: F^{\prime} B$.

## THE COMPLETE QUADRILATERAL.

A Complete Quadrilateral is formed by drawing any four straight lines, so that each of them shall cut each of the other three, so as to give six differ-

Fig. 586.
 ent points of intersection. It is so called because in the figure thus formed are found three quadrilaterals; riz., in Fig. 586, ABCD, a common convex quadrilateral; EAFC, a uni-concave quadrilateral; and EBAFD, a bi-concave quadrilateral, composed of two opposite triangles.

The complete quadrilateral, A E B C D F, has three diagonals; viz., two interior, A C, B D ; and one exterior, EF.

Theorem VIII. - In every complete quadrilateral the middle points of its three diagonals lie in the same straight line.
AEBCDF is the quadrilateral, and LMN the middle points of its three diagonals. From A and D draw parallels to B C , and from B and C draw parallels to A D. The triangle E D C being cut by the transrersal B F, we have (Theorem D), $\mathrm{DF} \times \mathrm{CB} \times \mathrm{EA}=\mathrm{CF} \times \mathrm{EB} \times \mathrm{D} A$. From the equality of parallels between parallels, we hare $\mathrm{CB}=\mathrm{E}^{\prime} \mathrm{B}^{\prime}, \mathrm{EA}=\mathrm{CA}^{\prime}, \mathrm{E} \mathrm{B}=$ $D B^{\prime}, D A=E^{\prime} A^{\prime}$. Hence, the above equation becomes $D F \times E^{\prime} B^{\prime} \times \mathrm{CA}^{\prime}$
$=\mathrm{CF} \times \mathrm{DB}^{\prime} \times \mathrm{E}^{\prime} \mathrm{A}^{\prime}$; therefore, by Theorem II, the points, $\mathrm{F}, \mathrm{B}^{\prime}, \mathrm{A}^{\prime}$, lie in the same straight line. Now, since the diagonals of the parallelogram $\mathrm{ECA}^{\prime} \mathrm{A}$ bisect each other at N , and those of the parallelogram EBB'D at M , we have $\mathrm{EN}: \mathrm{N} \mathrm{A}^{\prime}:: \mathrm{EM}: \mathrm{MB}^{\prime}$. Then $\mathrm{M} N$ is parallel to $\mathrm{FA}^{\prime}$, and we have $\mathrm{EN}: \mathrm{NA}^{\prime}:: \mathrm{EL}: \mathrm{LF}$, or $\mathrm{EL}=\mathrm{L} \mathrm{F}$, so that L is the middle of E F, and the same straight line passes through L, M, and N.

Theorem IX.-In every complete quadrilateral each of the three diagonals is divided harmonically by the two

Fig. 587.
 others.

CEBADF is the complete quadrilateral. The diagonal EF is divided harmonically at G and H by D B and A C produced ; since A H, D E, and FB are three transversals drawn from the summits of the triangle AEF through the same point C; and therefore, by Theorem VII, D B G and A C H divide EF harmonically.
So too, in the triangle ABD, C B , C A , CD, are the three transversals passing through $C$; and $G$ and $K$ therefore divide the diagoual $B D$ harmonically.

So, too, in the triangle, $\mathrm{ABC}, \mathrm{D} \mathrm{A} ,\mathrm{D} \mathrm{B} ,\mathrm{D} \mathrm{C} \mathrm{are} \mathrm{the} \mathrm{transversals}$, and K the points which divide the diagonal A C harmonically.

Theorem X.-If from a point, A, any number of lines be drawn, cutting the sides of an angle POQ , the intersections of the diagonals of the quadrilaterals thus formed will all lie in the same straight line passing through the summit of the angle.

By the preceding theorem, the diagonal $\mathrm{B} \mathrm{C}^{\prime}$ of the complete quadrilateral, $\mathrm{BAB} \mathrm{B}^{\prime} \mathrm{C}^{\prime} \mathrm{CO}$, is divided harmonically

Fig. 588.
 at D and E. Hence, O A, O P, O D, and O Q, form an harmonic pencil. So do OA, OP, O D', and O Q. Therefore, the lines OD, O D', coincide. So for the other intersections.

If the point A moves on OA , the line OD is not displaced. If, on the contrary, O A is displaced, O D turns around the point O. Hence, the point A is said to be a pole with respect to the line O D, which is itself called the polar of the point A. Similarly, D is a pole of O A, which is the polar of D. OD is likewise the polar of any other point on the line OA ; and this property is necessarily reciprocal for the two conjugate radii $\mathrm{OA}, \mathrm{OD}$, with respect to the lines O P, O Q which are also conjugate radii. Hence : in every harmonic pencil, each of the radii is a polar with respect to each point of its conjugate; and each point of this latter line is a pole with respect to the former.

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oR,
LaTITUDES AND DEPARTURES OF COURSES

CALCULATED TO

THREE DECIMAL PLACES:

FOR

EACH QUARTER DEGREE OF BEARING.

LATITUDES AND DEPARTURES．

| $\begin{gathered} 0.0 \\ \text { 品 } \\ 0.0 \end{gathered}$ | I |  | 2 |  | 3 |  | 4 |  | 5 | － |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | Lat． | Dep． | Lat | Dep． | Lat． | Dep． | Lat． | Dep． | Lat． |  |
| $0^{\text {c }}$ | 1.000 | o． | 2.000 | 0.000 | 3.000 | 0.000 | $4 \cdot 000$ | n．00n | 5.000 | $90^{\circ}$ |
| $0 \frac{1}{4}$ | 1.000 | o． | 2.0010 | riong | 3.000 | 0.013 | 4.000 | 0.017 | 5.000 | $9{ }^{3}$ |
| $\bigcirc \frac{1}{2}$ | 1.000 | 0．rncy | 2. | 0.017 | 3．000 | 0.026 | 4.000 | 0.035 | 5.000 | 8 |
| $0 \frac{3}{4}$ | 1.000 | 0．013 | 2.000 | 0．02t | 3．000 | 0.039 | 4.000 | 0.052 | 5.000 | $89 \frac{1}{4}$ |
| 1 | 1. | 0.01 | 2. | 0.035 | 3.000 | 0.052 | 3.999 | 0.070 | $4 \cdot 999$ | $89^{\circ}$ |
| $1 \frac{1}{4}$ | 1.00 | $0 \cdot 022$ | 2.000 | 0.044 | 2.999 | 0.065 | 3.999 | 0.087 | 4.999 | 883 |
| I 1 | 1.000 | 0.026 | 1.999 | 0.052 | 2.999 | 0.079 | 3.999 | －． 105 | $4 \cdot 998$ | $88 \frac{1}{2}$ |
| $1 \frac{3}{4}$ | 1.000 | 0.031 | 1.999 | 0．061 | 2.999 | 0.092 | 3.998 | 0.122 | $4 \cdot 998$ | $88 \frac{1}{4}$ |
| $2^{2}$ | 0.999 | 0．035 | 1.999 | 0.070 | 2.998 | 0．105 | 3.998 | 0.140 | $4 \cdot 997$ | $88^{\circ}$ |
| 2.1 | － 0.999 | 0．039 | 1.998 | 0.079 | 2.998 | 0.118 | 3.997 | 0.157 | 4.996 | $87 \frac{3}{4}$ |
| $2 \frac{1}{2}$ | 0.999 | 0．044 | 1．998 | 0.087 | 2.997 | 0．J31 | 3.996 | $0 \cdot 174$ | $4 \cdot 995$ | $87 \frac{1}{2}$ |
| $2 \frac{3}{4}$ | 0.999 | 0.048 | 1．998 | 0.096 | 2.997 | 0．1 14 | 3.995 | $0 \cdot 192$ | 4.994 | $87 \frac{1}{4}$ |
| $3^{\circ}$ | 0.999 | 0.052 | 1．997 | 0． 105 | 2.996 | 0.157 | 3.995 | 0.209 | 4.993 | $8 \%^{\circ}$ |
| 31 | 0.998 | 0.057 | 1．997 | 0．113 | 2.995 | 0.170 | 3.994 | 0.227 | $4 \cdot 992$ | $86 \frac{3}{4}$ |
| $3 \frac{1}{2}$ | 0.998 | 0.061 | 1．996 | 0．122 | 2.994 | 0．183 | 3.993 | 0． 244 | $4 \cdot 991$ | $86 \frac{1}{2}$ |
| $3 \frac{3}{5}$ | 0.998 | 0.065 | 1．996 | －．131 | 2.994 | －． 196 | 3.991 | 0.262 | 4.989 | $86 \frac{1}{4}$ |
| $4^{\circ}$ | $0 \cdot 998$ | 0.070 | 1．995 | 0.140 | $2 \cdot 993$ | 0.209 | 3.990 | 0.279 | $4 \cdot 988$ | $86^{\circ}$ |
| 44 | $0 \cdot 997$ | 0.074 | 1－995 | 0． 148 | 2.992 | 0.222 | 3.989 | 0． 296 | $4 \cdot 986$ | $85 \frac{3}{4}$ |
| $4 \frac{1}{2}$ | $0 \cdot 997$ | 0．078 | I．994 | 0． 157 | $2 \cdot 991$ | 0.235 | 3.988 | 0.314 | $4 \cdot 985$ | $85 \frac{1}{2}$ |
| $4 \frac{3}{4}$ | 0.957 | 0.083 | I 9993 | o． 166 | 2.990 | 0． 248 | $3 \cdot 986$ | 0.33 I | 4.983 | 85 $\frac{1}{4}$ |
| $5^{\circ}$ | 0.996 | 0．08－ | $1 \cdot 992$ | － | $2 \cdot 989$ | 0.261 | 3.985 | 0.349 | 4.981 | $85^{\circ}$ |
| 5 | $0 \cdot 996$ | 0.092 | 1.992 | 0.183 | $2 \cdot 987$ | 0.275 | $3 \cdot 983$ | 0．366 | $4 \cdot 979$ | 844 |
| $5 \frac{1}{2}$ | $0 \cdot 995$ | 0.096 | 1．991 | 0．192 | $2 \cdot 986$ | 0． 288 | $3 \cdot 982$ | 0.383 | $4 \cdot 977$ | $84 \frac{1}{2}$ |
| 5 | － 0.995 | 0．10） | 1．990 | 0.200 | $2 \cdot 985$ | 0．301 | $3 \cdot 980$ | $0 \cdot 401$ | $4 \cdot 975$ | 84 |
| $6^{\circ}$ | － 0.995 | 0．105 | 1.989 | 0.209 | $2 \cdot 984$ | －0．314 | 3．978 | 0.418 | $4 \cdot 973$ | $84^{\circ}$ |
| $6 \frac{1}{4}$ | － 0.994 | $0 \cdot 109$ | 1．988 | 0．218 | $2 \cdot 982$ | 0． 327 | $3 \cdot 976$ | 0.435 | 4.970 | 833 |
|  | －0．994 | －．113 | 1．987 | 0.226 | $2 \cdot 981$ | －． 340 | 3.974 | 0.453 | $4 \cdot 968$ | 83 ${ }^{2}$ |
| $6 \frac{3}{4}$ | －－993 | 0．118 | I． 986 | 0． 235 | $2 \cdot 979$ | 0.353 | $3 \cdot 972$ | 0.470 | $4 \cdot 965$ | $83 \ddagger$ |
| $7^{3}$ | $0 \cdot 993$ | $0 \cdot 122$ | I． 985 | 0.244 | $2 \cdot 978$ | 0．366 | 3.970 | 0.487 | $4 \cdot 963$ | $83^{2}$ |
| $7 \frac{1}{4}$ | $0 \cdot 992$ | $0 \cdot 126$ | 1．984 | 0.252 | $2 \cdot 976$ | 0.379 | 3．968 | 0.505 | $4 \cdot 960$ | $82 \frac{3}{4}$ |
| 7 $\frac{1}{2}$ | $0 \cdot 991$ | 0．131 | I． 983 | 0．26I | $2 \cdot 974$ | $0 \cdot 392$ | 3．966 | 0.522 | $4 \cdot 957$ | 32 $\frac{1}{2}$ |
| 73 | $0 \cdot 991$ | 0．135 | 1．932 | 0.270 | $2 \cdot 973$ | $0 \cdot 405$ | 3．963 | 0.539 | $4 \cdot 954$ | $82 \frac{1}{4}$ |
| $8^{\circ}$ | 0.990 | 0．139 | 1．981 | 0.278 | $2 \cdot 971$ | 0.418 | 3．961 | 0.557 | $4 \cdot 951$ | $82^{\circ}$ |
| $8 \frac{1}{4}$ | $0 \cdot 990$ | 0.143 | 1.979 | 0． 287 | 2.969 | $0 \cdot 430$ | $3 \cdot 959$ | 0.574 | 4．948 | 913 |
| $8 \frac{1}{2}$ | 0.989 | 0.148 | 1．978 | 0． 296 | $2 \cdot 967$ | 0.443 | 3．956 | 0.591 | $4 \cdot 945$ | $31 \frac{1}{2}$ |
| $8 \frac{3}{4}$ | 0．988 | 0．152 | 1.977 | 0．304 | $2 \cdot 965$ | 0.456 | $3 \cdot 953$ | 0.608 | 4．942 | 814 |
| $9^{\circ}$ | －0．988 | 0．156 | 1．975 | －0．313 | $2 \cdot 963$ | 0.469 | $3 \cdot 951$ | 0.626 | $4 \cdot 938$ | $81^{\circ}$ |
| $9 \frac{1}{4}$ | 0.98 .7 | 0．161 | 1．974 | 0.321 | $2 \cdot 961$ | 0.482 | 3．948 | 0.643 | $4 \cdot 935$ | $80 \frac{3}{4}$ |
| $9 \frac{1}{2}$ | $0 \cdot 986$ | －． 165 | 1．973 | 0．33o | $2 \cdot 959$ | 0.495 | 3．945 | 0．660 | $4 \cdot 931$ | $80 \frac{1}{2}$ |
| 94 | $0 \cdot 986$ | c． 169 | 1.971 | 0．339 | $2 \cdot 957$ | 0.508 | $3 \cdot 942$ | 0.677 | $4 \cdot 928$ | 80¢ |
| $10^{\circ}$ | 0.985 | 0．174 | 1.970 | 0.347 | 2.954 | 0.521 | 3.939 | $0 \cdot 695$ | 4.924 | $80^{\circ}$ |
| $10 \frac{1}{4}$ | 0.984 | 0．178 | I．968 | 0.356 | $2 \cdot 952$ | 0.534 | 3.936 | 0.712 | 4.920 | $79{ }^{\frac{3}{4}}$ |
| $10 \frac{1}{4}$ | 0.983 | 0． 182 | 1．967 | －． 364 | 2.950 | 0.547 | 3.933 | 0.729 | $4 \cdot 916$ | $79 \frac{1}{2}$ |
| $\mathrm{ic}^{3}$ | 0.982 | 0． 187 | 1．965 | 0.373 | 2.947 | 0.560 | 3.930 | $0 \cdot 746$ | 4.912 | 799 |
| $11^{\circ}$ | 0.982 | 0．191 | 1．963 | 0.382 | 2.945 | 0.572 | 3.927 | $0 \cdot 763$ | 4.908 | $79^{\circ}$ |
| $11 \frac{1}{4}$ | 0.981 | 0．195 | 1.962 | 0．390 | 2.942 | 0.585 | 2.923 | 0.780 | $4 \cdot 904$ | 78 |
| $11 \frac{1}{2}$ | 0．980 | － 199 | 1.960 | c． 399 | $2 \cdot 9.40$ | 0．598 | 3.920 | 0.747 | 4.900 | $78 \frac{1}{2}$ |
| 113 | 0.979 | 0． 204 | 1．958 | 0.407 | 2.937 | 0.611 | 3.916 | 0．815 | $4 \cdot 895$ | 78 |
| $12^{\circ}$ | 0.978 | 0． 208 | 1.956 | 0.416 | 2.934 | 0.624 | $3 \cdot 913$ | 0．832 | 4.891 | $78^{\circ}$ |
| $12 \frac{1}{4}$ | $0 \cdot 977$ | 0.212 | 1.954 | 0.424 | 2.932 | 0．637 | $3 \cdot 909$ | 0.849 | 4.856 | 773 |
| $12 \frac{1}{2}$ | $0 \cdot 976$ | 0．216 | 1.953 | 0.433 | 2.929 | 0.649 | 3.905 | $0.866^{\circ}$ | 4.881 | 776 |
| 123 | 0.975 | 0.221 | 1．951 | 0.441 | 2.926 | 0.662 | 3.901 | 0.883 | 4.877 | $77 \ddagger$ |
| $13^{\circ}$ | $0 \cdot 974$ | 0． 225 | 1．949 | 0．450 | 2.923 | 0.675 | 3.897 | c． 900 | 4.872 | $78^{\circ}$ |
| $13 \frac{1}{4}$ | 0.973 | 0.229 | 1.947 | 0． 458 | 2.920 | 0.688 | 3.894 | $0 \cdot 917$ | 4.867 | 763 |
| $13 \frac{1}{2}$ | 0.972 | －． 233 | 1.945 | 0.467 | 2.917 | $0 \cdot 700$ | 3.859 | 0.934 | 4.862 | $76 \frac{1}{2}$ |
| 133 | 0.971 | 0． 238 | 1．943 | 0.475 | 2.914 | 0.713 | 3.885 | 0.951 | 4.857 | 70¢ |
| 140 | 0.970 | 0.242 | I．94I | 0.484 | $2 \cdot 911$ | 0.726 | 3.881 | 0.968 | 4.851 | $76^{\circ}$ |
| $14 \frac{1}{4}$ | 0.969 | 0.246 | 1．938 | 0.492 | $2 \cdot 908$ | 0.738 | 3.877 3.873 | $0 \cdot 985$ | 4.8 .46 | 75 |
| $14 \frac{1}{2}$ | $0 \cdot 968$ | 0.250 | 1．936 | 0.501 | 2.904 | 0.751 | 3.873 | 1.002 | 4.845 | $75 \frac{1}{2}$ |
| $14 \frac{3}{4}$ | －$\cdot 967$ | 0． 255 | 1.934 | 0.509 | 2.901 | 0．764 | 3.868 | 1．018 | 4.835 | 75才 |
| $15^{\circ}$ | － 966 | 0.259 | $1 \cdot 932$ | 0.518 | $2 \cdot 898$ | 0.776 | 3.864 | I $\cdot 035$ | 4.830 | $75^{\circ}$ |
| 둥 | Dep． | Lat． | Dep． | Lat． | Dep． | Lat． | Dep． | Lat． | Dep． | $00^{0}$ |
| 岛 |  |  |  |  |  |  |  |  | 5 |  |

LATITUDES AND DEPARTURES．

| $\begin{aligned} & \text { 品 } \\ & \text { 品. } \end{aligned}$ | 5 | 6 |  | 7 |  | 8 |  | 0 |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | Dep． | Lat． | Dep． | Lat． | Dep． | Lat． | Dep． | Lat． | Dep． |  |
| $0^{\circ}$ | 0.000 | $6 \cdot 000$ | 0.000 | 7. | 0.000 | 8.000 | 0.000 | $9 \cdot 000$ | 0.000 | $90^{\circ}$ |
| $0 \frac{1}{4}$ | 0.022 | $6 \cdot 000$ | 0.026 | 7－000 | 0．031 | 8.000 | 0．035 | $9 \cdot 000$ | 0．039 | $89{ }^{\frac{3}{4}}$ |
| O $\frac{1}{2}$ | 0.044 | 6．000 | 3．052 | 7．0no | 0.061 | 8.000 | 0.070 | $9 \cdot 000$ | 0.079 | $89 \frac{1}{2}$ |
| $0 \frac{3}{4}$ | 0.065 | 5.999 | 0.079 | 6.999 | 0.092 | 7.999 | o．105 | 8.999 | 0.118 | $89 \frac{1}{4}$ |
| $1{ }^{\circ}$ | 0.087 | $5 \cdot 999$ | －． 105 | 6.999 | 0．122 | $7 \cdot 999$ | 0．140 | 8.999 | o． 157 | $89^{\circ}$ |
| $1 \frac{1}{4}$ | － 109 | 5.999 | －．13I | $6 \cdot 998$ | 0． 153 | $7 \cdot 998$ | 0．175 | 8.998 | c． 196 | $88 \frac{3}{4}$ |
| $1 \frac{1}{2}$ | $0 \cdot 131$ | $5 \cdot 998$ | －． 157 | $6 \cdot 998$ | 0．183 | $7 \cdot 997$ | $0 \cdot 209$ | $8 \cdot 997$ | 0． 236 | $88 \frac{1}{2}$ |
| $1 \frac{3}{4}$ | 0．153 | 5.997 | 0．183 | $6 \cdot 997$ | 0.214 | $7 \cdot 996$ | 0． 244 | $8 \cdot 996$ | 0.275 | $88 \frac{1}{4}$ |
| $2{ }^{\circ}$ | $0 \cdot 174$ | $5 \cdot 996$ | 0.209 | $6 \cdot 996$ | 0．244 | $7 \cdot 995$ | 0.279 | $8 \cdot 995$ | 0．314 | $88^{\circ}$ |
| 21 | －．19\％ | $5 \cdot 995$ | 0．236 | $6 \cdot 995$ | 0． 275 | 7.994 | 0．314 | $8 \cdot 993$ | 0.353 | $87 \frac{3}{4}$ |
| $2 \frac{1}{2}$ | 0.218 | $5 \cdot 994$ | 0． 262 | $6 \cdot 993$ | 0．305 | $7 \cdot 992$ | o． 349 | $8 \cdot 991$ | 0．393 | $87 \frac{1}{2}$ |
| $2 \frac{3}{4}$ | 0.240 | $5 \cdot 993$ | 0.288 | $6 \cdot 992$ | 0.336 | $7 \cdot 991$ | 0．384 | $8 \cdot 990$ | 0.432 | $87 \frac{1}{4}$ |
| $3^{\circ}$ | 0．262 | $5 \cdot 992$ | 0.314 | 6.990 | o． 366 | $7 \cdot 989$ | 0．419 | $8 \cdot 988$ | 0.471 | 880 |
| $3 \frac{1}{4}$ | － 283 | $5 \cdot 990$ | 0.340 | 6.989 | 0.397 | $7 \cdot 987$ | 0．454 | 8－986 | 0.510 | 863 |
| $3 \frac{1}{2}$ | 0.305 | $5 \cdot 989$ | 0.366 | $6 \cdot 987$ | 0.427 | $7 \cdot 985$ | －$\cdot 488$ | 8－983 | 0.549 | $86 \frac{1}{2}$ |
| 3 | 0.327 | $5 \cdot 987$ | 0．392 | $6 \cdot 985$ | 0.458 | $7 \cdot 983$ | 0.523 | $8 \cdot 981$ | 0.589 | 864 |
| $4^{\circ}$ | 0．349 | $5 \cdot 985$ | 0.419 | $6 \cdot 983$ | 0.488 | $7 \cdot 981$ | 0．558 | $8 \cdot 978$ | 0． 628 | $86^{\circ}$ |
| $4 \frac{1}{4}$ | 0.371 | $5 \cdot 984$ | 0.445 | $6 \cdot 981$ | 0.519 | $7 \cdot 978$ | 0．593 | $8 \cdot 975$ | 0.667 | $85 \frac{3}{4}$ |
| 41 ${ }^{2}$ | $0 \cdot 392$ | $5 \cdot 982$ | 0.471 | $6 \cdot 978$ | 0.549 | $7 \cdot 975$ | $0 \cdot 623$ | $8 \cdot 972$ | － 706 | $85 \frac{1}{2}$ |
| $4 \frac{3}{4}$ | 0.414 | 5.979 | 0.497 | $6 \cdot 976$ | 0.580 | $7 \cdot 973$ | 0.662 | 8－969 | $0 \cdot 745$ | $85 \frac{1}{4}$ |
| $5{ }^{\circ}$ | 0.436 | $5 \cdot 977$ | 0.523 | $6 \cdot 973$ | 0.610 | $7 \cdot 970$ | 0.697 | $8 \cdot 966$ | 0.784 | $85^{\circ}$ |
| $5 \frac{1}{4}$ | 0.458 | $5 \cdot 975$ | 0.549 | $6 \cdot 971$ | 0.641 | $7 \cdot 966$ | 0.732 | $8 \cdot 962$ | 0.824 | 84 |
| $5 \frac{1}{3}$ | 0.479 | $5 \cdot 972$ | 0.575 | $6 \cdot 968$ | 0.671 | $7 \cdot 963$ | 0.767 | $8 \cdot 959$ | 0.863 | $84 \frac{1}{2}$ |
| $5{ }^{\frac{3}{4}}$ | 0.501 | 5.970 | 0.601 | $6 \cdot 965$ | 0.701 | $7 \cdot 960$ | 0．802 | $8 \cdot 955$ | $0 \cdot 902$ | $84 \frac{1}{4}$ |
| $6^{\circ}$ | 0.523 | $5 \cdot 967$ | 0.627 | $6 \cdot 962$ | 0.732 | $7 \cdot 956$ | 0．836 | $8 \cdot 951$ | －0．941 | 81 ${ }^{\circ}$ |
| 61 | 0.544 | $5 \cdot 964$ | 0.653 | $6 \cdot 958$ | 0.762 | $7 \cdot 952$ | 0．871 | $8 \cdot 947$ | －0．980 | 833 |
| $6 \frac{1}{2}$ | 0.566 | $5 \cdot 961$ | 0.679 | 6.955 | 0.792 | $7 \cdot 949$ | $0 \cdot 906$ | 8．942 | 1．019 | $83 \frac{1}{2}$ |
| 63 | 0.588 | $5 \cdot 958$ | o． 705 | 6.951 | 0．823 | $7 \cdot 945$ | －． 940 | 8－938 | I． 058 | $83 \frac{1}{4}$ |
| $8{ }^{\circ}$ | 0.609 | $5 \cdot 955$ | $0 \cdot 731$ | 6.948 | 0.853 | 7－940 | －． 975 | $8 \cdot 933$ | $1 \cdot 097$ | $83^{\text {c }}$ |
| $7 \frac{1}{4}$ | － 0.631 | $5 \cdot 952$ | 0.757 | 6.944 | 0.883 | 7－936 | I－ OIO | $8 \cdot 928$ | I•136 | $82 \frac{3}{4}$ |
| $7 \frac{1}{2}$ | 0.653 | $5 \cdot 949$ | 0.783 | 6.940 | 0.914 | $7 \cdot 9^{32}$ | I－ 044 | 8－923 | I－175 | $82 \frac{1}{2}$ |
| 73 ${ }^{4}$ | 0.674 | $5 \cdot 945$ | 0.809 | $6 \cdot 936$ | 0.944 | $7 \cdot 927$ | I－079 | 8．918 | I． 214 | $82 \frac{1}{4}$ |
| $8^{\circ}$ | 0.696 | $5 \cdot 942$ | 0．835 | 6.932 | 0.974 | $7 \cdot 922$ | I－113 | $8 \cdot 912$ | I－253 | $82^{\circ}$ |
| 84 | 0.717 | $5 \cdot 938$ | － 0.861 | $6 \cdot 928$ | I． 004 | $7 \cdot 917$ | I $\cdot 148$ | $8 \cdot 907$ | 1．291 | $8 \mathrm{I} \frac{3}{4}$ |
| $8 \frac{1}{2}$ | 0.739 | $5 \cdot 934$ | $0 \cdot 887$ | $6 \cdot 923$ | I． 035 | $7 \cdot 912$ | I 1． 182 | $8 \cdot 901$ | 1．33o | $81 \frac{1}{2}$ |
| $8{ }^{8}$ | 0．761 | $5 \cdot 930$ | $0 \cdot 913$ | $6 \cdot 919$ | I． 065 | 7907 | 1－217 | 8．895 | I． 369 | $81 \frac{1}{4}$ |
| $9^{\circ}$ | $0 \cdot 782$ | $5 \cdot 926$ | － $9^{3} 9$ | $6 \cdot 914$ | 1．095 | $7 \cdot 902$ | I－25ı | $8 \cdot 889$ | I． 408 | $81^{\circ}$ |
| $9 \frac{1}{4}$ | 0.804 | $5 \cdot 922$ | －$\cdot 964$ | $6 \cdot 909$ | I．125 | $7 \cdot 896$ | I $\cdot 286$ | $8 \cdot 883$ | 1．447 | $80 \frac{3}{4}$ |
| $9 \frac{1}{2}$ | 0.825 0.847 | $5 \cdot 918$ 5 | － 999 | $6 \cdot 904$ | I．I 55 | $7 \cdot 890$ | I－ 320 | $8 \cdot 877$ | I－485 | $80 \frac{1}{2}$ |
| $9{ }^{\frac{3}{4}}$ | 0.847 | $5 \cdot 913$ | 1．016 | 6.899 | I 185 | 7.884 | I $\cdot 355$ | $8 \cdot 870$ | I． 524 | $80 \frac{1}{4}$ |
| $10^{\circ}$ | 0.868 | $5 \cdot 909$ | 1．042 | 6.894 | I． 216 | $7 \cdot 878$ | 1．389 | 8.863 | I． 563 | $80^{\circ}$ |
| $10 \frac{1}{4}$ | 0.890 | $5 \cdot 904$ | I－ 068 | 6.888 | I－246 | $7 \cdot 872$ | 1.424 | 8.856 | 1.601 | $79 \frac{3}{4}$ |
| $10 \frac{1}{2}$ | 0.911 | $5 \cdot 900$ | 1.093 | 6.883 | 1．276 | $7 \cdot 866$ | 1． 458 | 8.849 | 1.640 | $79 \frac{1}{2}$ |
| 104 | 0.933 | 5.895 | 1．119 | 6.877 | I． 306 | $7 \cdot 860$ | 1．492 | 8.842 | 1．679 | $79^{\frac{1}{4}}$ |
| $11^{\circ}$ | 0.954 | 5.890 5.885 | 1． 145 | 6.871 | 1．336 | $7 \cdot 853$ | 1． 526 | 8.835 | 1.717 | $79^{\circ}$ |
| 119 | 0.975 | 5.885 | 1．171 | 6.866 | I $\cdot 366$ | $7 \cdot 846$ | I $\cdot 561$ | 8.827 | 1．756 | $78 \frac{3}{4}$ |
| $11 \frac{1}{3}$ | 0.997 | 5．880 | I－ 196 | 6.859 | 1．396 | $7 \cdot 839$ | 1．595 | $8 \cdot 819$ | 1.794 | $78 \frac{1}{2}$ |
| 113 | 1.018 | $5 \cdot 874$ | I． 222 | 6.853 | I． 425 | $7 \cdot 832$ | I． 629 | 8．811 | 1.833 | $78 \pm$ |
| $12^{\circ}$ | I．040 | 5.869 | 1． 247 | 6.847 | I． 455 | $7 \cdot 8.25$ | I． 663 | $8 \cdot 8 \mathrm{o} 3$ | 1．871 | $78^{\circ}$ |
| $12 \frac{1}{4}$ | 1．061 | 5.863 5.858 | 1． 273 | 6.841 | I． 485 | 7.818 | 1.697 | $8 \cdot 795$ | 1.910 | 77 |
| $12 \frac{1}{3}$ | I $\cdot 082$ | 5.858 | I． 299 | 6.834 | I． 515 | $7 \cdot 810$ | $1 \cdot 732$ | $8 \cdot 787$ | I． 948 | $77 \frac{1}{2}$ |
| 123 | I $\cdot 103$ | $5 \cdot 852$ | 1．324 | 6.827 | I． 545 | $7 \cdot 803$ | 1．766 | $8 \cdot 778$ | $1 \cdot 986$ | $77 \frac{1}{4}$ |
| $13^{\circ}$ | I． 125 | $5 \cdot 846$ | 1．350 | 6．821 | 1．575 | $7 \cdot 795$ | 1.800 | $8 \cdot 769$ | 2.025 | 780 |
| $13 \frac{1}{4}$ | I． 146 | 5.840 | I． 375 | 6.814 | 1．604 | 7－787 | I .834 | $8 \cdot 760$ | 2.063 | $76 \frac{3}{4}$ |
| $13 \frac{1}{2}$ $13 \frac{3}{4}$ 18 | ı $\cdot 167$ － 188 | 5.834 5.828 | 1．401 | 6.807 | I． 634 | 7－779 | I． 868 | $8 \cdot 751$ | $2 \cdot 101$ | $76 \frac{1}{2}$ |
| $13 \frac{3}{4}$ | I 188 | $5 \cdot 828$ | 1．426 | 6.799 | 1．664 | $7 \cdot 771$ | 1．902 | $8 \cdot 742$ 8.733 | $2 \cdot 139$ | 764 |
| $11^{\circ}$ | I．210 | $5 \cdot 822$ | 1．452 | 6.792 | I $\cdot 693$ | 7•762 | 1．935 | $8 \cdot 733$ | $2 \cdot 177$ | $86^{\circ}$ |
| $14 \frac{1}{4}$ | 1．23I | 5．815 | 1．477 | $6 \cdot 785$ | $1 \cdot 723$ | 7－754 | 1.969 | $8 \cdot 723$ | 2.215 | $75 \frac{3}{4}$ |
| $14 \frac{1}{2}$ | 1．252 | 5.809 | 1．502 | 6.777 | I $\cdot 753$ | 7－745 | $2 \cdot 003$ | 8．713 | $2 \cdot 253$ | $75 \frac{1}{2}$ |
| 143 | 1． 273 | $5 \cdot 8 \mathrm{uz}$ | I． 528 | $6 \cdot 769$ | I $\cdot 782$ | 7－736 | 2.037 | $8 \cdot 703$ | $2 \cdot 291$ | $75 \frac{1}{4}$ |
| $15^{\circ}$ | I－ 294 | $5 \cdot 746$ | 1．553 | $6 \cdot 761$ | I．812 | $7 \cdot 727$ | 2.071 | 8－693 | 2.329 | 75 ${ }^{\circ}$ |
| 区 | Lat． | Dep． | Lat． | ep | Lat． | Dep． | Lat． | Dep． | Lat． |  |
|  | 5 |  |  |  |  |  |  |  |  |  |


| LATITUDES AND DEPARTURES． |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\begin{aligned} & \text { 땅 } \\ & \text { 军. } \\ & 0 \end{aligned}$ | 1 |  | 2 |  | 3 |  | 4 |  |  | $\begin{aligned} & \text { •号 } \\ & \text { 品 } \\ & \hline \end{aligned}$ |
|  | Lat． | Dep． | Lat． | Dep． | Lat． | Dep． | Lat． | Dep． |  |  |
| $15^{\circ}$ | c． 966 | 0.259 | 1.932 | 0.518 | 2.898 | 0．776 | 3.864 | I． 035 | 4.830 | $75^{\circ}$ |
| $15 \frac{1}{4}$ | 0.965 | 0.263 | 1．930 | 0.526 | 2.894 | 0.789 | 3.859 | 1．052 | 4.824 | 74 |
| $15 \frac{1}{3}$ | 0.964 | 0.267 | 1.927 | 0.534 | 2.891 | 0.802 | 3.855 | I． 069 | 4.818 | 743 |
| $15 \frac{3}{4}$ | 0.962 | 0.271 | 1．925 | 0.543 | 2.887 | 0.814 | 3.850 | I 1.086 | 4.812 | $74 \frac{4}{4}$ |
| $16^{\circ}$ | $0.96{ }^{\text {c }}$ | 0．275 | 1．923 | 0.551 | 2.884 | 0.827 | 3.845 | I 1 103 | 4.806 | $74^{\circ}$ |
| $16 \frac{1}{4}$ | －．9600 | 0.280 | I．920 | c 560 | 2.880 | 0．839 | 3.840 | I•119 | 4.800 | 733 |
| $16 \frac{1}{2}$ | 0.959 | 0．284 | 1．918 | 0.568 | 2.876 | 0.852 | 3.835 | I．I36 | 4.794 | $73 \frac{1}{2}$ |
| $16 \frac{3}{4}$ | 0.958 | 0.283 | 1．915 | 0.576 | 2.873 | 0.865 | 3.830 | I•153 | $4 \cdot 788$ | $73 \frac{1}{4}$ |
| $17^{\circ}$ | 0.956 | 0.292 | 1．913 | 0.585 | 2.869 | 0.877 | 3.825 | I． 169 | $4 \cdot 782$ | $73^{\circ}$ |
| $17 \frac{1}{4}$ | 0.955 | 0.297 | $1 \cdot 910$ | $0 \cdot 59^{3}$ | 2.865 | 0.890 | 3.820 | I $\cdot 186$ | $4 \cdot 775$. | 723 |
| $17 \frac{1}{2}$ | 0.954 | 0.301 | $1 \cdot 907$ | 0.601 | 2.861 | 0.902 | $3 \cdot 815$ | 1． 203 | $4 \cdot 769$ | $72 \frac{1}{3}$ |
| $178 \frac{3}{4}$ | 0.952 | 0.305 | 1．905 | c．610 | 2.857 | 0.915 | 3.810 | I． 220 | $4 \cdot 762$ | $72 \frac{1}{4}$ |
| $18^{\circ}$ | 0.95 I | 0．309 | $1 \cdot 902$ | 0.618 | 2.853 | 0.927 | 3.804 | I $\cdot 236$ | $4 \cdot 755$ | $72^{\circ}$ |
| $18 \frac{1}{\ddagger}$ | 0.950 | $0 \cdot 313$ | 1．899 | 0.626 | 2.849 | $0 \cdot 939$ | 3.799 | I． 253 | $4 \cdot 748$ | $71 \frac{3}{4}$ |
| $18 \frac{1}{2}$ | －0．948 | 0.317 | I． 897 | 0.635 | 2.845 | 0.952 | 3.793 | I． 269 | $4 \cdot 742$ | $71 \frac{1}{2}$ |
| $18 \frac{3}{4}$ | 0.947 | 0.321 | 1．894 | 0.643 | 2.841 | －．964 | 3．788 | I $\cdot 286$ | $4 \cdot 735$ | $71 \frac{1}{4}$ |
| $19^{\circ}$ | 0.946 | 0．326 | 1.891 | － 0.651 | 2.837 | $0 \cdot 977$ | 3.782 | I 3 302 | $4 \cdot 728$ | $71{ }^{\circ}$ |
| $19 \frac{1}{4}$ | 0.944 | 0．330 | I． 888 | 0.659 | 2.832 | 0.989 | 3.776 | $1 \cdot 319$ 1.335 | $4 \cdot 720$ | $70 \frac{3}{4}$ |
| $19 \frac{1}{2}$ | 0.943 | － 0.334 | 1．885 | $0 \cdot 668$ | 2.828 | 1．001 | 3.771 | 1．335 | $4 \cdot 713$ | $70 \frac{1}{2}$ |
| $19 \frac{3}{4}$ | 0.94 T | 0.338 | I .882 | $0 \cdot 676$ | 2.824 | I．014 | 3．765 | I． 352 | 4．706 | $70 \frac{1}{4}$ |
| $20^{\circ}$ | 0.940 | 0.3 .42 |  | 0.684 |  | 1.026 | 3.759 | 68 | 4.698 | $70^{\circ}$ |
| $20 \frac{1}{4}$ | 0.938 | 0.346 | 1．876 | 0.692 | 2.815 | I． 038 | 3.753 | I． 384 | 4.691 | 609 |
| $20 \frac{1}{2}$ | 0.937 | 0.350 | 1．873 | 0.700 | 2.810 | I． 05 I | 3.747 | I． 401 | 4.683 | $69 \frac{1}{2}$ |
| $20 \frac{3}{4}$ | 0.935 | 0.354 | 1.870 | 0.709 | 2.805 | 1．063 | 3.741 | I． 417 | $4 \cdot 676$ | $69 \ddagger$ |
| $21^{\circ}$ | 0.934 | 0.358 | I． 867 | 0.717 | 2.801 | 1．075 | 3.734 | I． 433 | 4.668 | $69^{\circ}$ |
| 21 ${ }^{1}$ | 0.932 | 0.362 | I． 864 | 0.725 | $2 \cdot 796$ | I．087 | 3.728 | I－450 | 4.660 | 58 |
| $2 \mathrm{~S} \frac{1}{2}$ | 0.930 | $0 \cdot 367$ | I． 861 | 0.733 | $2 \cdot 791$ | 00 | 3.722 | I． 466 | 4.652 | $68 . \frac{1}{2}$ |
| $21 \frac{3}{4}$ | 0.929 | 0.371 | 1.858 | 0.741 | $2 \cdot 786$ | I．II 2 | 3.715 | 1．482 | 4.644 | $68 \ddagger$ |
| $22^{\circ}$ | 0.927 | 0.375 | 1．854 | 0.749 | $2 \cdot 782$ | I－124 | 3.709 | 1． 498 | 4.636 | $68^{\circ}$ |
| $22 \frac{1}{7}$ | $0 \cdot 926$ | 0.379 | 1.851 | 0.757 | 2.777 | I－136 | $3 \cdot 702$ | I． 515 | 4.628 | 6－3 |
| $22 \frac{1}{2}$ | 0.924 | 0.383 | I． 848 | 0.765 | $2 \cdot 772$ | I－148 | $3 \cdot 696$ | I． 531 | $4 \cdot 619$ | $67 \frac{1}{2}$ |
| 223 | 0.922 | 0.387 | 1.844 | $0 \cdot 773$ | $2 \cdot 767$ | 1－160 | 3.689 | I． 547 | $4 \cdot 611$ | 67 $\frac{1}{4}$ |
| $23^{\circ}$ | 0.921 | $0 \cdot 391$ | I．841 | 0．781 | $2 \cdot 762$ | I•172 | 3.682 | I 563 | 4.603 | $67^{\circ}$ |
| $23 \frac{1}{4}$ | 0.919 | 0.395 | I． 838 | 0.789 | $2 \cdot 756$ | 1－184 | 3.675 | 1．579 | 4.594 | $66 \frac{3}{4}$ |
| $33 \frac{1}{2}$ | 0.917 | 0.399 | I． 834 | 0.797 | $2 \cdot 751$ | 1－196 | 3.668 | 1．595 | 4.585 | $66 \frac{1}{2}$ |
| 233 | 0.915 | 0.403 | I．83I | 0.805 | $2 \cdot 746$ | I－208 | 3.661 | 1．611 | 4.577 | $66 \frac{1}{4}$ |
| $21^{\circ}$ | 0.914 | 0.407 | I． 827 | 0．813 | $2 \cdot 741$ | I－220 | $3 \cdot 654$ | 1． 027 | 4.568 | $66^{\circ}$ |
| $24 \frac{1}{4}$ | $0 \cdot 912$ | 0.411 | I． 824 | 0.821 | $2 \cdot 735$ | 1． 232 | 3.647 | I． 643 | 4.559 | $65 \frac{3}{4}$ |
| $24 \frac{1}{4}$ | $0 \cdot 91$ | 0.415 | I． 820 | 0.829 | $2 \cdot 730$ | I． 244 | 3.640 | I．659 | 4.550 | $65 \frac{1}{2}$ |
| $24 \frac{3}{4}$ | $0 \cdot 908$ | 0.419 | I．816 | 0.837 | $2 \cdot 724$ | I． 256 | 3.633 | $1 \cdot 675$ | 4.54 I | 65 |
| $25^{\circ}$ | 0.90 | 0.423 | 1.813 | 0.845 | 2719 | I－268 | 3.625 | 1．690 | 4.532 | $65^{\circ}$ |
| $25 \frac{1}{4}$ | $0 \cdot 904$ | 0.427 | 1．809 | 0．853 | 2.713 | 1．280 | 3.618 | I． 706 | 4.522 | 64 |
| 25 $\frac{1}{2}$ | $0 \cdot 903$ | 0.431 | 1．803 | 0.861 | 2.708 | 1．292 | 3.610 | I． 722 | 4.513 | $64 \frac{1}{2}$ |
| $25 \frac{3}{4}$ | 0.901 | 0.434 | 1.801 | 0.869 | 2.702 | 1．303 | 3.603 | I． 738 | 4.503 | 64才 |
| $26^{\circ}$ | 0.899 | 0.438 | 1.798 | $0.87 \%$ | 2.696 | 1．315 | 3.595 | 1．753 | 4． 294 | $61^{\circ}$ |
| $26 \frac{1}{2}$ | 0.897 | 0.442 | 1．794 | 0.885 | 2.691 | 1．327 | 3.557 | 1.769 | $4 \cdot 48.4$ | $63 \frac{3}{4}$ |
| $26 \frac{1}{2}$ | 0.895 | 0.446 | 1．790 | 0.892 | 2.685 | 1．339 | 3.580 | 1．785 | 4.475 | $63 \frac{1}{2}$ |
| 263 ${ }^{2}$ | 0.893 | 0.450 | 1．786 | 0.900 | 2.679 | 1．350 | 3.572 3.564 | 1．800 | 4.465 | $63 \ddagger$ |
| $27^{\circ}$ | 0.891 | 0.454 | 1.782 | 0.908 | $2 \cdot 673$ | I $\cdot 362$ | 3．564 | I．816 | 4.455 | $63^{\circ}$ |
| $27 \frac{1}{4}$ | 0.889 0.857 | 0．458 | 1.778 | 0.916 | 2.667 | I． 374 I 385 | 3.556 3.548 | I．83I | 4.445 | $62 \frac{3}{4}$ |
| $27 \frac{1}{1}$ | 0.887 0.885 | 0．462 | 1.774 | 0.923 | $2 \cdot 661$ | 1385 | 3.548 | 1．847 | 4.435 | 62. |
| 278 | 0.885 | 0.466 | 1．770 | 0.931 | 2.655 | 1.397 | 3.540 | 1． 562 | 4.425 | $62 \frac{1}{4}$ |
| $28^{\circ}$ | 0.883 | $0 \cdot 459$ | I． 766 | 0.939 | 2.649 | 1.408 | 3.532 | 1．8－5 | 4.415 | $62^{\circ}$ |
| $28 \ddagger$ | 0.881 | 0.473 | 1．762 | 0.947 | 2．643 | 1.420 | 3.524 | 1.893 | 4.404 | $61 \frac{3}{4}$ |
| $28 \frac{1}{2}$ | 0.875 | 0.477 | 1．758 | 0．954 | 2.636 | I． 43 I | 3．515 | 1.909 | 4．394 | 612 |
| $283{ }^{2}$ | 0.877 0.875 | 0．48I | 1．753 | 0.962 | 2.030 | 1．443 | 3.507 | 1．924 | $4 \cdot 384$ | $61 \ddagger$ |
| $29^{\circ}$ | 0.875 | 0.485 | $1 \cdot 749$ | c．970 | 2.624 | $1 \cdot 454$ | $3.49^{8}$ | 1．939 | 4.373 | $61^{\circ}$ |
| $29 \frac{1}{\ddagger}$ | 0.872 0.870 | 0.495 0.492 | 1.745 | 0.977 | 2.617 -611 | 1．466 | 3.490 | 1.954 | 4.362 | $6 \mathrm{n}{ }^{3}$ |
| $29 \frac{1}{2}$ 208 | 0.870 0.868 | 0.492 | $1 \cdot 741$ | 0.985 | －611 | $1 \cdot 477$ | 3．681 | ： 9.970 | 4.35 | 601 |
| $30^{\circ}$ | 0.866 | 0.496 0.500 | 1.730 1.732 | 0.992 1．000 | 2.605 2.598 | 1.459 $\pm .500$ | 3.40 .4 | I 985 2.000 | 4.330 | $60^{\circ}$ |
| ¢ | D | Lat． | Dep． | Lat． | ep． | Lat | Dep． | Lat． | Dep． | － |
|  |  |  |  |  |  |  |  |  | 5 | 馬 |

LATITUDES AND DEPARTURES．

| $\begin{aligned} & \text { 마 } \\ & \text { 籴. } \end{aligned}$ | $\frac{5}{\text { Dep. }}$ | 6 |  | 7 |  | 8 |  | 0 |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | Lat． | Dep． | Lat． | Dep． | Lat． | Dep． | Lat． | Dep． |  |
| $15^{\circ}$ | I $\cdot 294$ | $5 \cdot 796$ | I． 553 | $6 \cdot 761$ | I．8I 2 | $7 \cdot 727$ | 2.071 | 8.693 | 2.329 | $75^{\circ}$ |
| $15 \frac{1}{4}$ | I．315 | $5 \cdot 789$ | 1.578 | $6 \cdot 754$ | 1.84 I | $7 \cdot 718$ | $2 \cdot 104$ | 8.683 | $2 \cdot 367$ | 74 |
| 15 $\frac{1}{2}$ | I． 336 | 5－782 | 1．603 | $6 \cdot 745$ | 1.871 | $7 \cdot 709$ | 2．138 | 8.673 | 2.405 | $74 \frac{1}{8}$ |
| $15 \frac{3}{4}$ | I $\cdot 357$ | $5 \cdot 775$ | 1． 629 | $6 \cdot 737$ | $1 \cdot 900$ | $7 \cdot 700$ | 2．172 | 8.662 | 2.443 | 744 |
| $16^{\circ}$ | I． 378 | $5 \cdot 768$ | I． 654 | 6．72ら | 1．929 | 7.690 | $2 \cdot 205$ | 8．65I | 2.48 I | $74^{\circ}$ |
| 164 | 1．399 | 5．76o | 1． 679 | $6 \cdot 720$ | 1.959 | 7.680 | $2 \cdot 239$ | 8640 | 2．518 | 733 |
| $16 \frac{1}{2}$ | 1.420 | $5 \cdot 753$ | $1 \cdot 704$ | $6 \cdot 712$ | 1．988 | 7.671 | 2.272 | 8.629 | 2.556 | $73 \frac{1}{2}$ |
| $16 \frac{3}{4}$ | I． 441 | $5 \cdot 745$ | $1 \cdot 729$ | $6 \cdot 703$ | 2.017 | 7.661 | $2 \cdot 306$ | 8.618 | 2.594 | $73 \frac{1}{4}$ |
| $18^{\circ}$ | 1． 462 | $5 \cdot 738$ | I． 754 | 6.694 | 2.047 | 7.650 | $2 \cdot 339$ | 8.607 | $2 \cdot 63$ I | $73^{\circ}$ |
| $17 \frac{1}{4}$ | 1.483 | $5 \cdot 730$ | I $\cdot 779$ | 6.685 | 2.076 | 7.640 | 2.372 | $8 \cdot 595$ | $2 \cdot 669$ | $72 \frac{3}{4}$ |
| $17 \frac{1}{2}$ | 1.504 | $5 \cdot 722$ | 1.804 | $6 \cdot 676$ | $2 \cdot 105$ | 7.630 | $2 \cdot 406$ | 8.583 | $2 \cdot 706$ | $72 \frac{1}{2}$ |
| $17 \frac{3}{4}$ | 1．524 | $5 \cdot 714$ | 1．829 | 6.667 | 2.134 | $7 \cdot 619$ | 2.439 | 8.572 | 2.744 | $72 \frac{1}{4}$ |
| $18^{\circ}$ | I． 545 | 5．706 | I． 854 | 6.657 | $2 \cdot 163$ | $7 \cdot 608$ | 2.472 | 8．56．0 | $2 \cdot 781$ | $72^{\circ}$ |
| $18 \frac{1}{4}$ | I． 566 | $5 \cdot 698$ | I $\cdot 879$ | 6.648 | $2 \cdot 192$ | $7 \cdot 598$ | $2 \cdot 505$ | 8.547 | 2.818 | 713 |
| $18 \frac{1}{2}$ | I． 587 | $5 \cdot 690$ | 1－904 | 6.638 | $2 \cdot 22.1$ | $7 \cdot 587$ | 2.538 | 8.535 | 2.856 | $-1 \frac{1}{2}$ |
| $18 \frac{3}{4}$ | 1.607 | $5 \cdot 682$ | 1．929 | 6.629 | 2.250 | $7 \cdot 575$ | 2.572 | $8 \cdot 522$ | $2 \cdot 893$ | $71 \frac{1}{4}$ |
| $19^{\circ}$ | I． 628 | $5 \cdot 673$ | I $\cdot 953$ | $6 \cdot 619$ | 2.279 | $7 \cdot 564$ | $2 \cdot 605$ | $8 \cdot 5 \mathrm{I}$ o | $2 \cdot 930$ | $71{ }^{\circ}$ |
| $19{ }^{\frac{1}{4}}$ | I． 648 | 5.665 | I $\cdot 978$ | $6 \cdot 609$ | $2 \cdot 308$ | $7 \cdot 553$ | $2 \cdot 638$ | 8．497 | $2 \cdot 967$ | $70 \frac{3}{4}$ |
| 1919 | I． 669 | 5.656 | $2 \cdot 003$ | 6.598 | 2.337 | $7 \cdot 541$ | $2 \cdot 670$ | $8 \cdot 484$ | $3 \cdot 004$ | 701 |
| $19 \frac{3}{4}$ | 1.690 | $5 \cdot 647$ | $2 \cdot 028$ | 6.588 | $2 \cdot 365$ | $7 \cdot 529$ | $2 \cdot 703$ | 8．471 | 3．041 | $70 \frac{1}{4}$ |
| $20^{\circ}$ | 1.710 | 5.638 | $2 \cdot$ | 6.578 | 2.394 | 7－518 | $2 \cdot 736$ | 8.457 | 3.078 | $70^{\circ}$ |
| $20 \frac{1}{4}$ | $1 \cdot 731$ | 5.629 | $2 \cdot 077$ | 6.567 | 2.423 | $7 \cdot 506$ | $2 \cdot 769$ | 8.444 | 3．115 | $69 \frac{3}{4}$ |
| $20 \frac{1}{2}$ | I－751 | 5.620 | 2．101 | 6.557 | 2.45 I | $7 \cdot 493$ | 2.802 | 8．43o | 3．152 | 69t |
| $20 \frac{3}{4}$ | 1－771 | 5．611 | $2 \cdot 126$ | $6 \cdot 546$ | 2.480 | 7－481 | 2.834 | $8 \cdot 416$ | 3．189 | $69 \frac{1}{4}$ |
| $21^{\circ}$ | 1．792 | $5 \cdot 60$ I | $2 \cdot 150$ | 6.535 | 2.509 | 7－469 | $2 \cdot 867$ | 8．402 | 3.225 | $69^{\circ}$ |
| $21 \frac{1}{4}$ | 1．812 | $5 \cdot 592$ | $2 \cdot 175$ | 6.524 | 2.537 | 7－456 | $2 \cdot 900$ | 8.388 | $3 \cdot 262$ | 683 |
| $21 \frac{1}{2}$ | 1.833 | 5.582 | 2．199 | 6.513 | 2.566 | 7－443 | $2 \cdot 932$ | $8 \cdot 374$ | 3.299 | $68 \frac{1}{2}$ |
| $21 \frac{3}{4}$ | 1．853 | 5.573 | $2 \cdot 223$ | 6.502 | 2.594 | 7．430 | $2 \cdot 964$ | $8 \cdot 359$ | 3.335 | $68 \frac{1}{4}$ |
| $22^{\circ}$ | 1.873 | 5.563 | $2 \cdot 248$ | 6.490 | 2.622 | $7 \cdot 417$ | $2 \cdot 997$ | $8 \cdot 345$ | 3.371 | $68^{\circ}$ |
| $22 \frac{1}{4}$ | 1.893 | 5.553 | $2 \cdot 272$ | 6.479 | 2.65 I | 7－404 | $3 \cdot 029$ | 8．330 | 3.408 | 673 |
| 22 $\frac{1}{2}$ | 1．913 | 5.543 5 | $2 \cdot 296$ | 6.467 | 2.679 | $7 \cdot 391$ | 3．06I | $8 \cdot 3 \mathrm{I}$ | 3.444 | $67 \frac{1}{2}$ |
| $22 \frac{3}{4}$ | I－934 | $5 \cdot 533$ | $2 \cdot 320$ | 6.455 | 2.707 | 7－378 | 3．094 | $8 \cdot 300$ | 3．480 | $67 \frac{1}{4}$ |
| $23^{\circ}$ | I． 954 | $5 \cdot 523$ | $2 \cdot 344$ | 6.444 | $2 \cdot 735$ | $7 \cdot 364$ | 3－126 | 8－285 | 3.517 | $67^{\circ}$ |
| $23 \frac{1}{4}$ | 1．974 | $5 \cdot 513$ | $2 \cdot 368$ | 6.432 | $2 \cdot 763$ | $7 \cdot 350$ | 3－158 | $8 \cdot 269$ | 3.553 | $66 \frac{3}{4}$ |
| $23 \frac{1}{2}$ | 1－994 | $5 \cdot 502$ | $2 \cdot 392$ | 6.419 | $2 \cdot 791$ | $7 \cdot 336$ | 3－190 | $8 \cdot 254$ | j．589 | $66 \frac{1}{2}$ |
| 233 | 2.014 | 5.492 | $2 \cdot 416$ | $6 \cdot 407$ | 2.819 | $7 \cdot 322$ | $3 \cdot 222$ | $8 \cdot 238$ | 3.625 | $66 \frac{1}{4}$ |
| $24^{\circ}$ | $2 \cdot 034$ | 5．481 | 2.440 | 6.395 | 2.847 | 7－308 | 3－254 | $8 \cdot 222$ | $3 \cdot 661$ | $66^{\circ}$ |
| 244 | $2 \cdot 054$ | 5．471 | $2 \cdot 464$ | 6.382 | 2.875 | 7－294 | 3－286 | $8 \cdot 206$ | $3 \cdot 696$ | 653 |
| $24 \frac{1}{2}$ | $2 \cdot 073$ | 5．460 | $2 \cdot 488$ | 6.370 | $2 \cdot 903$ | 7280 | 3．3i8 | $8 \cdot 190$ | $3 \cdot 732$ | 65 $\frac{1}{2}$ |
| 24妥 | $2 \cdot 09^{3}$ | $5 \cdot 449$ | 2．5I 2 | 6.357 | 2．93I | 7－265 | $3 \cdot 349$ | 8－173 | 3－768 | $65 \frac{1}{4}$ |
| $25^{\circ}$ | 2．113 | 5.438 | 2.536 | 6．344 | 2.958 | 7－250 | $3 \cdot 381$ | $8 \cdot 157$ | 3.804 | $65^{\circ}$ |
| $25 \frac{1}{4}$ | 2．133 | 5．427 | 2.559 | $6 \cdot 33 \mathrm{I}$ | $2 \cdot 986$ | $7 \cdot 236$ | 3.413 | $8 \cdot 140$ | 3．839 | $64 \frac{3}{4}$ |
| 25⿺𠃊 | $2 \cdot 153$ | 5．416 | 2.583 | 6．318 | 3.014 | 7－221 | 3.444 | 8－123 | 3.875 | 64t |
| $25 \frac{3}{4}$ | 2．172 | $5 \cdot 404$ | 2.607 | $6 \cdot 305$ | 3．041 | 7－206 | 3.476 | 8－106 | 3.910 | 64t |
| $26^{\circ}$ | 2．192 | $5 \cdot 39^{3}$ | 2.630 | $6 \cdot 292$ | 3.069 | 7．190 | 3.507 | 8.089 | 3.945 | $64^{\circ}$ |
| $26 \frac{1}{4}$ | 2.211 | $5 \cdot 381$ | 2.654 | $6 \cdot 278$ | 3.096 | 7－175 | 3.538 | 8.072 | 3.981 | 633 |
| $26 \frac{1}{2}$ | $2 \cdot 231$ | $5 \cdot 370$ | $2 \cdot 677$ | $6 \cdot 265$ | 3．123 | 7－160 | 3.570 | 8.054 | 4.016 | $63 \frac{1}{2}$ |
| 263 | 2.250 | $5 \cdot 358$ | $2 \cdot 701$ | $6 \cdot 25 \mathrm{I}$ | 3．151 | 7－144 | 3.601 | 8.037 | 4.051 | $63 \frac{1}{4}$ |
| $28^{\circ}$ | 2.270 | $5 \cdot 346$ | 2.724 | $6 \cdot 237$ | 3．178 | 7－128 | 3.632 | $8 \cdot 019$ | 4.086 | $63^{\circ}$ |
| 274 | 2.289 | 5．334 | 2.747 | 6.223 | 3．205 | 7－112 | 3.663 | $8 \cdot 001$ | 4.121 | 623 |
| 27 $\frac{1}{2}$ | 2.309 | $5 \cdot 322$ | 2.770 | $6 \cdot 209$ | 3.232 | $7 \cdot \mathrm{cg} 6$ | 3.694 | $7 \cdot 983$ | 4．156 | $62 \frac{1}{2}$ |
| $27 \frac{3}{4}$ | 2.328 | $5 \cdot 310$ | 2.794 | $6 \cdot 195$ | $3 \cdot 259$ | $7 \cdot 080$ | $3 \cdot 725$ | $7 \cdot 965$ | 4.190 | $62 \frac{1}{4}$ |
| $28^{\circ}$ | 2.347 | $5 \cdot 298$ | 2.817 | $6 \cdot 181$ | 3．286 | $7 \cdot 064$ | $3 \cdot 756$ | $7 \cdot 947$ | 4.225 | $6{ }^{\circ}$ |
| $28 \frac{1}{4}$ | 2367 | $5 \cdot 285$ | 2.840 | $6 \cdot 166$ | 3．313 | $7 \cdot 047$ | 3.787 | 7－928 | $4 \cdot 26 r$ | 613 |
| $28 \frac{1}{2}$ | 2.386 | $5 \cdot 273$ | 2.863 | 6．152 | 3.340 | 7－031 | 3.817 | 7.909 | $4 \cdot 294$ | $6 \mathrm{I} \frac{1}{2}$ |
| 283 | $2 \cdot 405$ | $5 \cdot 260$ | 2.886 | $6 \cdot 137$ | $3 \cdot 367$ | 7－014 | 3.848 | $7 \cdot 891$ | $4.32 y$ | $61 \frac{1}{4}$ |
| 99 ${ }^{\circ}$ | 2.424 | 5 | 2.909 | $6 \cdot 122$ | $3 \cdot 394$ | 6.997 | 3.878 | 7.872 | $4 \cdot 363$ | $61^{\circ}$ |
| 291 | 2.443 | $5 \cdot 2 \mathrm{j}$ | $2 \cdot 932$ | $6 \cdot 107$ | $3 \cdot 420$ | $6 \cdot 980$ | 3.909 | $7 \cdot 852$ | $4 \cdot 398$ | $60 \frac{3}{4}$ |
| $29 \frac{1}{2}$ | 2.462 | $5 \cdot 222$ | $2 \cdot 955$ | $6 \cdot 093$ | $3 \cdot 447$ | $6 \cdot 963$ | $3 \cdot 939$ | 7.833 | 4.432 | $60 \frac{1}{2}$ |
| $\begin{array}{r}298 \\ 30 \\ \hline 0\end{array}$ | 2.481 2.500 | $5 \cdot 209$ $5 \cdot 196$ | $2 \cdot 977$ $3 \cdot 000$ | 6.077 6.062 | 3.474 3.509 | 6.946 6.928 | 3.970 4.000 | 7.814 | 4.466 | $60 \frac{1}{4}$ |
| $30^{\circ}$ | $2 \cdot 500$ | 5－196 | 3．000 | $6 \cdot 002$ | 3.50 | 6.928 | $4 \cdot 000$ | $7 \cdot 794$ | $4 \cdot 500$ | $60^{\circ}$ |
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|  | Lat. | Dep. | Lat. | Dep. | Lat. | Dep. | Lat. | Dep. | Lat. |  |
| $30^{\circ}$ | 0.866 | 0.500 | 1.732 | I $\cdot 000$ | 2.598 | 1.500 | 3464 | $2 \cdot 000$ | 4.330 | $60^{\circ}$ |
| $30 \frac{1}{4}$ | 0.864 | 0.504 | 1.728 | I .008 | 2.592 | 1.511 | 3.455 | $2 \cdot 015$ | 4.319 | 593 |
| 30 $\frac{1}{3}$ | 0. 862 | 0.508 | 1.723 | I. 015 | 2.585 | I. 523 | 3.447 | 2.030 | $4 \cdot 308$ | $59 \frac{1}{2}$ |
| $30 \frac{3}{4}$ | 0.859 | 0.511 | 1.719 | t.023 | 2.578 | I. 534 | 3.438 | 2.045 | 4.297 | $59 \frac{1}{4}$ |
| $31^{\circ}$ | 0.857 | 0.515 | 1.714 | I.030 | 2.572 | I. 545 | 3.429 | 2.060 | $4 \cdot 286$ | $59^{\circ}$ |
| 311 | 0.855 | 0.519 | 1.710 | I . 038 | 2.565 | 1. 556 | 3.420 | $2 \cdot 075$ | $4 \cdot 275$ | $58 \frac{3}{4}$ |
| $3 \mathrm{I} \frac{1}{2}$ | 0.853 | 0.522 | $1 \cdot 705$ | I $\cdot 045$ | 2.558 | I. 567 | $3 \cdot 411$ | 2.090 | $4 \cdot 263$ | $58 \frac{1}{2}$ |
| $31 \frac{3}{4}$ | 0.850 | 0. 526 | 1.701 | $1 \cdot 052$ | 2.55 I | I. 579 | $3 \cdot 401$ | 2.105 | 4.252 | $58 \frac{1}{4}$ |
| $32^{\circ}$ | 0.848 | 0.530 | 1.696 | I $\cdot 060$ | 2.544 | I. 590 | $3 \cdot 392$ | $2 \cdot 120$ | $4 \cdot 240$ | $58^{\circ}$ |
| $32 \frac{1}{4}$ | 0.846 | 0.534 | I. 691 | I.067 | 2.537 | 1.601 | 3.383 | 2.134 | 4.229 | $57 \frac{3}{4}$ |
| $32 \frac{1}{2}$ | 0.843 | 0.537 | 1.687 | I.075 | 2.530 | I.612 | 3.374 | 2.149 | 4.217 | $57 \frac{1}{2}$ |
| $32 \frac{3}{4}$ | 0.84i | 0.541 | 1.682 | I. 082 | 2.523 | 1.623 | 3.364 | $2 \cdot 164$ | $4 \cdot 205$ | $57 \frac{1}{4}$ |
| $33^{\circ}$ | 0.839 | 0.545 | 1. 677 | I.089 | 2.516 | I. 634 | 3.355 | 2.179 | $4 \cdot 193$ | $57^{\circ}$ |
| $33 \frac{1}{4}$ | 0.836 | 0.548 | I. 673 | I. 097 | 2.509 | I. 645 | 3.345 | $2 \cdot 193$ | $4 \cdot 181$ | $56 \frac{3}{4}$ |
| $33 \frac{1}{2}$ | 0.834 | 0.552 | I. 668 | 1.104 | 2.502 | 1.656 | 3.336 | $2 \cdot 208$ | $4 \cdot 169$ | $56 \frac{1}{2}$ |
| 333 | 0.831 | 0.556 | I. 663 | I.111 | 2.494 | I. 667 | 3.326 | $2 \cdot 222$ | $4 \cdot 157$ | $56 \frac{1}{4}$ |
| $31^{\circ}$ | 0.829 | 0. 559 | I. 658 | 1-118 | 2.487 | I $\cdot 678$ | 3.3I6 | $2 \cdot 237$ | $4 \cdot 145$ | $56^{\circ}$ |
| $34 \frac{1}{4}$ | 0.827 | 0.563 | 1.653 | I-I26 | 2.480 | I. 688 | 3.306 | $2 \cdot 25 \mathrm{I}$ | $4 \cdot 133$ | $55 \frac{3}{4}$ |
| $34 \frac{1}{2}$ | 0.824 | 0.566 | I. 648 | I•I33 | 2.472 | I. 699 | 3.297 | $2 \cdot 266$ | $4 \cdot 121$ | $55 \frac{1}{2}$ |
| 343 | 0.822 | 0.570 | I. 643 | I $\cdot 140$ | $2 \cdot 465$ | I.710 | $3 \cdot 287$ | $2 \cdot 280$ | 4.108 | 551 |
| $35^{\circ}$ | 0.819 | 0.574 | 1. 638 | I 147 | 2.457 | 1.721 | 3.277 | $2 \cdot 294$ | $4 \cdot 096$ | $55^{\circ}$ |
| $35 \frac{1}{4}$ | 0.817 | 0.577 | I. 633 | I. 154 | 2.450 | 1-731 | 3.267 | 2.309 | 4.083 | 54 |
| $35 \frac{1}{2}$ | 0.814 | 0.581 | I. 628 | I-I61 | 2.442 | 1.742 | $3 \cdot 257$ | 2.323 | $4 \cdot 071$ | $54 \frac{1}{2}$ |
| 353 | 0.812 | 0.584 | 1.623 | I. 168 | 2.435 | I $\cdot 753$ | $3 \cdot 246$ | 2.337 | 4.058 | $54 \frac{1}{4}$ |
| $36^{\circ}$ | 0.809 | 0.583 | 1.618 | I $\cdot 176$ | 2.427 | I $\cdot 763$ | 3. 236 | 2.35 I | 4.045 | $54^{\circ}$ |
| $36 \frac{1}{4}$ | 0.806 | 0.591 | 1.613 | 1.183 | 2.419 | I.774 | $3 \cdot 226$ | $2 \cdot 365$ | 4.032 | 533 |
| $36 \frac{1}{2}$ | 0.804 | 0.595 | 1. 608 | I. 190 | 2.412 | 1-784 | 3.215 | 2.379 | $4 \cdot 019$ | $53 . \frac{1}{2}$ |
| $36 \frac{3}{4}$ | 0.801 | 0.598 | 1.603 | 1-197 | 2.404 | 1.795 | $3 \cdot 205$ | $2 \cdot 393$ | 4.006 | $53 \ddagger$ |
| $3 \%^{\circ}$ | 0.799 | 0.602 | 1. 597 | I-204 | 2.396 | I $\cdot 805$ | $3 \cdot 195$ | 2.407 | 3.993 | $53^{\circ}$ |
| $37 \frac{1}{4}$ | $0 \cdot 796$ | 0.605 | 1.592 | I-21I | $2 \cdot 388$ | I.816 | $\begin{array}{llll}3 & 184\end{array}$ | $2 \cdot 421$ | 3.980 | $52 \frac{3}{4}$ |
| $37 \frac{1}{2}$ | 0.793 | 0.609 | 1.587 | 1.218 | 2.380 | $\mathrm{I} \cdot 826$ | 3.173 | 2.435 | 3.967 | $52 \frac{1}{2}$ |
| $37 \frac{3}{4}$ 38 | $0 \cdot 791$ | $0 \cdot 612$ | 1.581 | I $\cdot 224$ | $2 \cdot 372$ | I.837 | 3.163 | 2.449 | 3.953 | $52 \frac{1}{4}$ |
| $38^{\circ}$ | $0 \cdot 788$ | -0.616 | 1.576 | I $\cdot 231$ | 2.364 | 1.847 | 3.152 | 2.463 | 3.940 | $52^{\circ}$ |
| $38 \pm$ | 0.785 | -0.619 | 1.571 | I-238 | $2 \cdot 356$ | 1.857 | 3.141 | 2.476 | $3 \cdot 927$ | $51 \frac{3}{4}$ |
| $38 \frac{1}{2}$ | $0 \cdot 783$ | 0.623 | 1. 565 | I $\cdot 245$ | 2.348 | I.868 | 3.130 | 2.490 | 3.913 | $51 \frac{1}{2}$ |
| $38 \frac{3}{4}$ | $0 \cdot 780$ | 0.626 | I. 560 | I. 25.2 | 2.340 | 1.878 | 3.120 | 2.504 | 3.899 | $51 \frac{1}{4}$ |
| $39^{\circ}$ | $0 \cdot 777$ | -0.629 | 1.554 | I $\cdot 259$ | $2 \cdot 33 \mathrm{I}$ | 1-888 | 3.109 | 2.517 | 3.886 | $51^{\circ}$ |
| $39 \frac{1}{4}$ | 0.774 | 0.633 | I . 549 | I. 265 | $2 \cdot 323$ | I-898 | 3.098 | 2.531 | 3.872 | $50 \frac{3}{4}$ |
| $39 \frac{1}{3}$ | 0.772 | 0.636 | I. 543 | I $\cdot 272$ | $2 \cdot 3 \mathrm{I} 5$ | 1-908 | 3.086 | 2.544 | 3.858 | $50 \frac{1}{2}$ |
| 394 | 0.769 | 0.639 | I. 538 | I-279 | $2 \cdot 307$ | $1 \cdot 918$ | 3.075 | 2.558 | 3.844 | $50 \frac{1}{4}$ |
| $40^{\circ}$ | $0 \cdot 766$ | 0.643 | I. 532 | I $\cdot 286$ | 2.298 | 1.928 | 3.064 | 2.571 | 3.830 | $50^{\circ}$ |
| $40 \frac{1}{4}$ | 0.763 | 0. 646 | I. 526 | I-292 | 2.290 | 1.938 | 3.053 | 2.584 | 3.816 | $49{ }^{\frac{3}{4}}$ |
| $40 \frac{1}{2}$ | 0.760 | 0.649 | I. 521 | I. 299 | $2 \cdot 281$ | 1.948 | 3.042 | 2.598 | 3.802 | $49 \frac{1}{2}$ |
| $40 \frac{3}{4}$ | $0 \cdot 758$ | 0.653 | 1.515 | I 306 | 2273 | 1.958 | 3.030 | 2.611 | 3.788 | 49 |
| $41^{\circ}$ | 0.755 | 0. 656 | 1.509 | 1.312 | 2264 | 1.968 | 3.019 | 2.624 | 3.774 | $49^{\circ}$ |
| $41 \frac{1}{4}$ | $0 \cdot 752$ | - 0.659 | 1.504 | I.319 | 2.256 | 1.978 | 3.007 | 2.637 | 3.759 | $45^{3}$ |
| $41 \frac{1}{3}$ | $0 \cdot 749$ | 0. 663 | 1.498 | I. 325 | 2.247 | I.988 | 2.996 | 2.650 | 3.745 | $48 \frac{1}{2}$ |
| 413 | 0.746 | 0.666 | 1.492 | I.332 | $2 \cdot 238$ | 1.998 | $2 \cdot 954$ | 2.664 | 3.730 | 48 |
| $42^{\circ}$ | 0.743 | 0.669 | 1.486 | I $\cdot 338$ | $2 \cdot 229$ | 2.007 | $2 \cdot 973$ | 2.677 | 3.716 | $48^{\circ}$ |
| 42 $\frac{1}{4}$ | 0.740 | 0.672 | 1.480 | I. 345 | $2 \cdot 221$ | 2.017 | 2.961 | 2.689 | 3.701 3.686 | 473, |
| $42 \frac{1}{2}$ | 0.737 0.734 | 0.676 | I. 475 | I $\cdot 351$ <br> I 358 | 2.212 | 2.027 | 2.949 | 2.702 | 3.686 | 476 |
| 42 年 | 0.734 | 0.679 | I. 469 | I $\cdot 358$ | $2 \cdot 203$ | 2.036 | 2.937 | 2.715 | 3.672 | $47 \ddagger$ |
| $43^{\circ}$ | $0 \cdot 7^{3} 1$ | 0.682 | I. 463 | +.364 | $2 \cdot 194$ | 2.046 | 2.925 | $2 \cdot 728$ | 3.657 | $47^{\circ}$ |
| 431 435 | 0.728 | 0.685 | 1.457 | $1 \cdot 370$ | 2.185 | 2.056 | $2 \cdot 913$ | $2 \cdot 741$ | 3.642 | 463 |
| $4{ }^{4} 3 \frac{1}{3}$ | 2.725 | 0.688 | 1.451 | 1.377 1.383 | 2.176 | 2.065 | $2 \cdot 901$ | $2 \cdot 753$ | 3.627 | $46 \frac{1}{2}$ |
| $43 \frac{3}{4}$ | $0 \cdot 722$ | 0.692 | I. 445 | 1.383 | 2.167 | 2.075 | 2.889 | $2 \cdot 766$ | 3.612 | $46 \ddagger$ |
| $44^{\circ}$ | $0 \cdot 719$ | 0.695 | I. 439 | I. 384 | $2 \cdot 158$ | 2.084 | 2.877 | $2 \cdot 779$ | 3.597 | $46^{\circ}$ |
| $44 \frac{1}{1}$ | 0.716 | 0.698 | 1.433 | 1.396 | 2.149 | $2 \cdot 093$ | 2.865 | $2 \cdot 791$ | 3.582 | $45 \frac{3}{4}$ |
| $44 \frac{1}{2}$ 443 | 0.713 | 0.701 | I. 427 | $1 \cdot 402$ | 2.140 | 2.103 | 2.85 ? | 2.804 | $3 \cdot 566$ | $45 \frac{1}{2}$ |
| $\begin{array}{r}443 \\ 45^{\circ} \\ \hline\end{array}$ | $0 \cdot 710$ | 0.704 0.707 | 1.420 1.414 | 1.408 | $2 \cdot 131$ | $2 \cdot 112$ | 2.841 2.828 | 2.816 2.828 | 3.551 3.536 | 45t |
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## LATITUDES AND DEPARTURES．

| 망 | 5 | 6 |  | 7 |  | 8 |  | 9 |  | $\begin{aligned} & \text { 品 } \\ & \text { 荡 } \\ & \text { p } \end{aligned}$ |
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| 旨 | Dep． | Lat． | Dep． | Lat． | Dep． | Lat． | Dep． | Lat． | Dep． |  |
| $30^{\circ}$ | $2 \cdot 500$ | 5．196 | 3．000 | 6.062 | 3.500 | 6.928 | 4．000 | 7－794 | 4.500 | $60^{\circ}$ |
| 307 | 2.519 | 5．183 | 3．023 | 6.047 | 3.526 | 6.911 | 4．030 | 7．775 | 4.534 | 593 |
| 30，$\frac{1}{2}$ | 2.538 | 5．170 | 3．045 | 6.031 | 3.553 | 6.893 | 4．060 | 7．755 | $4 \cdot 568$ | 591 |
| $30 \frac{3}{4}$ | 2.556 | 5．156 | 3．068 | 6.016 | 3.579 | 6.875 | $4 \cdot 090$ | $7 \cdot 735$ | $4 \cdot 602$ | 591 |
| $31^{\circ}$ | $2 \cdot 575$ | 5．143 | 3．090 | 6.000 | $3 \cdot 605$ | 6.857 | $4 \cdot 120$ | 7－715 | $4 \cdot 635$ | $59^{\circ}$ |
| $31 \frac{1}{4}$ | 2.594 | 5．199 | 3－113 | 5．984 | $3.63{ }_{1}$ | 6.839 | $4 \cdot 150$ | 7.694 | $4 \cdot 669$ | 58 |
| 31 3 | 2.612 | 5．116 | 3－135 | 5．968 | 3.657 | 6.821 I | 480 | 7.674 | $4 \cdot 702$ | $58 \frac{1}{2}$ |
| $31 \frac{3}{4}$ | 2.631 | 5．102 | $3 \cdot 157$ | $5 \cdot 952$ | 3683 | 6.803 | $4 \cdot 210$ | 7.653 | $4 \cdot 736$ | $58 \frac{1}{4}$ |
| $32^{\circ}$ | 2.650 | 5．088 | 3．180 | 5．936 | 3．709 | $6 \cdot 784$ | 4．239 | 7.632 | $4 \cdot 769$ | $58^{\circ}$ |
| $32 \frac{1}{4}$ | 2.668 | $5 \cdot 074$ | $3 \cdot 202$ | $5 \cdot 920$ | $3 \cdot 735$ | $6 \cdot 766$ | $4 \cdot 269$ | 7－612 | $4 \cdot 802$ | 573 |
| $32 \frac{1}{2}$ | $2 \cdot 686$ | $5 \cdot 060$ | $3 \cdot 224$ | $5 \cdot 904$ | 3－761 | $6 \cdot 747$ | $4 \cdot 298$ | 7．591 | 4.836 | $57 \frac{1}{2}$ |
| 323 | $2 \cdot 705$ | 5．046 | 3．246 | 5.887 | $3 \cdot 787$ | $6 \cdot 728$ | $4 \cdot 328$ | 7.569 | $4 \cdot 869$ | 57 |
| $33^{\circ}$ | $2 \cdot 723$ | $5 \cdot 032$ | 3．268 | 5.871 | 3．812 | 6.709 | $4 \cdot 357$ | $7 \cdot 548$ | $4 \cdot 902$ | $57^{\circ}$ |
| 334 | $2 \cdot 741$ | j－018 | $3 \cdot 290$ | $5 \cdot 854$ | 3．838 | 6.690 | $4 \cdot 386$ | 7.527 | $4 \cdot 935$ | 563 |
| 33. | $2 \cdot 760$ | $5 \cdot 003$ | $3 \cdot 312$ | 5.837 | 3.864 | 6.671 | $4 \cdot 416$ | $7 \cdot 505$ | $4 \cdot 967$ | $56 \frac{1}{2}$ |
| 333 | $2 \cdot 778$ | $4 \cdot 989$ | $3 \cdot 333$ | 5.820 | $3 \cdot 889$ | $6 \cdot 652$ | $4 \cdot 445$ | $7 \cdot 483$ | $5 \cdot 000$ | $56 \frac{1}{4}$ |
| $34^{\circ}$ | $2 \cdot 796$ | $4 \cdot 974$ | $3 \cdot 355$ | $5 \cdot 803$ | 3－914 | 6.632 | $4 \cdot 474$ | 7．461 | 5．033 | $56^{\circ}$ |
| 344 | 2.814 | $4 \cdot 960$ | $3 \cdot 377$ | 5．786 | $3 \cdot 940$ | 6．6ı3 | $4 \cdot 502$ | 7.439 | 5－065 | $55 \frac{3}{4}$ |
| $34 \frac{1}{2}$ | 2.832 | $4 \cdot 945$ | $3 \cdot 398$ | 5－769 | $3 \cdot 965$ | 6.593 | $4 \cdot 53 \mathrm{I}$ | 7.417 | 5－0，9 | $55 \frac{1}{2}$ |
| 343 | 2.850 | $4 \cdot 930$ | $3 \cdot 420$ | 5．752 | $3 \cdot 990$ | 6.573 | $4 \cdot 560$ | 7．395 | 5．130 | $55 \frac{1}{4}$ |
| $35^{\circ}$ | 2.868 | $4 \cdot 915$ | $3 \cdot 441$ | 5．734 | 4．015 | 6.553 | 4.589 | $7 \cdot 372$ | 5．162 | $55^{\circ}$ |
| $35 \frac{1}{4}$ | 2.886 | $4 \cdot 900$ | $3 \cdot 463$ | $5 \cdot 716$ | $4 \cdot 040$ | 6.533 | $4 \cdot 617$ | 7350 | 5．194 | 54 |
| 35 $\frac{1}{2}$ | 2.904 | 4.885 | 3.484 | 5.699 | $4 \cdot 065$ | 6.513 | $4 \cdot 646$ | $7 \cdot 327$ | 5．226 | $54 \frac{1}{2}$ |
| $35 \frac{3}{4}$ | $2 \cdot 921$ | 4.869 | $3 \cdot 505$ | 5.681 | $4 \cdot 090$ | 6.493 | $4 \cdot 674$ | $7 \cdot 304$ | 5．258 | $54 \frac{1}{4}$ |
| $36^{\circ}$ | 2.939 | 4.854 | 3.527 | 5.663 | 4．115 | 6.472 | $4 \cdot 702$ | 7．281 | 5.290 | $54^{\circ}$ |
| 364 | $2 \cdot .957$ | $4 \cdot 839$ | 3.548 | 5.645 | $4 \cdot 139$ | 6.452 | $4 \cdot 730$ | $7 \cdot 258$ | $5 \cdot 322$ | 533 |
| $36 \frac{1}{2}$ | 2.974 | $4 \cdot 823$ | 3.569 | 5.627 | 4．164 | 6.43 r | $4 \cdot 759$ | 7－235 | 5.353 | $53 \frac{1}{2}$ |
| 363 | 2．992 | $4 \cdot 808$ | 3.590 | 5．609 | 4．188 | 6.410 | $4 \cdot 787$ | 7－211 | 5.385 | 531 |
| $38^{\circ}$ | 3.009 | $4 \cdot 792$ | 3.611 | $5 \cdot 590$ | $4 \cdot 213$ | 6.389 | $4 \cdot 815$ | 7－188 | 5．416 | $53^{\circ}$ |
| 374 | 3.026 | $4 \cdot 776$ | 3.632 | 5.572 | $4 \cdot 237$ | 6.368 | $4 \cdot 842$ | 7－164 | 5．448 | $52 \frac{3}{4}$ |
| $37 \frac{1}{2}$ | 3.044 | $4 \cdot 760$ | 3.653 | 5.554 | $4 \cdot 261$ | 6.347 | $4 \cdot 870$ | 7．140 | 5．479 | $52 \frac{1}{2}$ |
| 373 | 3.061 | $4 \cdot 744$ | 3.673 | $5 \cdot 535$ | $4 \cdot 286$ | 6.326 | $4 \cdot 898$ | 7－116 | 5．510 | 52．$\frac{1}{4}$ |
| $38^{\circ}$ | $3 \cdot 078$ | $4 \cdot 728$ | 3.694 | 5．516 | $4 \cdot 310$ | 6.304 | $4 \cdot 925$ | $7 \cdot 092$ | 5．541 | $52^{\circ}$ |
| $38 \frac{1}{4}$ | $3 \cdot 095$ | $4 \cdot 712$ | 3.715 | $5 \cdot 497$ | $4 \cdot 334$ | 6.283 | $4 \cdot 953$ | 7．068 | 5.572 | $5 \mathrm{I} \frac{3}{4}$ |
| $38 \frac{1}{2}$ | 3．113 | $4 \cdot 696$ | $3 \cdot 735$ | $5 \cdot 478$ | $4 \cdot 358$ | $6 \cdot 2.61$ | $4 \cdot 980$ | $7 \cdot 043$ | $5 \cdot 603$ | 51 $\frac{1}{2}$ |
| $38 \frac{3}{4}$ | 3．130 | $4 \cdot 679$ | $3 \cdot 756$ | $5 \cdot 459$ | $4 \cdot 38$ I | 6.239 | $5 \cdot 007$ | $7 \cdot 019$ | 5.633 | 51 $\frac{1}{4}$ |
| $39^{\circ}$ | 3．147 | $4 \cdot 663$ | $3 \cdot 776$ | $5 \cdot 440$ | $4 \cdot 405$ | $6 \cdot 217$ | 5．035 | $6 \cdot 994$ | $5 \cdot 664$ | $51^{\circ}$ |
| 394 | 3．164 | 4.646 | 3.796 | 5.421 | $4 \cdot 429$ | $6 \cdot 195$ | $5 \cdot 062$ | $6 \cdot 970$ | 5．694 | $50 \frac{3}{4}$ |
| $39 \frac{1}{2}$ | 3．180 | 4.630 | $3 \cdot 816$ | $5 \cdot 401$ | $4 \cdot 453$ | $6 \cdot 173$ | $5 \cdot 089$ | $6 \cdot 945$ | 5．725 | $50 \frac{1}{2}$ |
| $39^{\frac{3}{4}}$ | 3－197 | $4 \cdot 613$ | 3.837 | 5.382 | $4 \cdot 476$ | $6 \cdot 151$ | 5．116 | $6 \cdot 920$ | $5 \cdot 755$ | $50 \frac{1}{4}$ |
| $40^{\circ}$ | 3.214 | 4.596 | 3.857 | 5.362 | $4 \cdot 500$ | $6 \cdot 128$ | 5．142 | 6.894 | $5 \cdot 785$ | $50^{\circ}$ |
| 40，${ }^{4}$ | 3．23I | 4.579 | 3.877 | $5 \cdot 343$ | $4 \cdot 5 \cdot 23$ | $6 \cdot$ ı06 | 5－169 | 6.869 | 5.815 | $49 \frac{3}{4}$ |
| $40 \frac{1}{2}$ | 3.247 | 4.562 | 3.897 | 5.323 | $4 \cdot 546$ | $6 \cdot 083$ | 5．196 | $6 \cdot 844$ | 5.845 | $49 \frac{1}{2}$ |
| 403 | $3 \cdot 264$ | 4.545 | 3.917 | 5.303 | 4.569 | 6.061 | 5．222 | $6 \cdot 818$ | 5.875 | $49 \frac{1}{4}$ |
| $41^{\circ}$ | 3．280 | 4.528 | 3．936 | 5．283 | $4 \cdot 592$ | 6.038 | 5．248 | $6 \cdot 792$ | $5 \cdot 905$ | $49^{\circ}$ |
| $41 \frac{1}{4}$ | $3 \cdot 297$ | $4 \cdot 511$ | $3 \cdot 956$ | $5 \cdot 263$ | $4 \cdot 615$ | 6.015 | $5 \cdot 275$ | $6 \cdot 767$ | 5．934 | 483 |
| $41 \frac{1}{2}$ | 3.313 | 4.494 | 3.976 | 5．243 | $4 \cdot 638$ | 5.992 | 5.301 | $6 \cdot 741$ | 5．964 | 481 |
| 413 490 | 3.329 3.346 | 4.476 | 3.995 | 5．222 | 4.661 | $5 \cdot 968$ | 5.327 5 | $6 \cdot 715$ | $5 \cdot 993$ | $48 \frac{1}{4}$ |
| $42^{\circ}$ | 3.346 | $4 \cdot 459$ | $4 \cdot 015$ | $5 \cdot 202$ | $4 \cdot 684$ | $5 \cdot 945$ | 5.353 | $6 \cdot 688$ | 6．022 | $48^{\circ}$ |
| $42 \frac{1}{4}$ | 3.362 3.378 | $4 \cdot 441$ | $4 \cdot 034$ | 5－182 | $4 \cdot 707$ | $5 \cdot 922$ | 5.379 | $6 \cdot 662$ | $6 \cdot 05 \mathrm{I}$ | 473 |
| $42 \frac{1}{2}$ | 3.378 | 4.424 | $4 \cdot 054$ | 5．161 | $4 \cdot 729$ | 5.898 | $5 \cdot 405$ | $6 \cdot 635$ | $6 \cdot 080$ | 472 |
| 423 | 3.394 | $4 \cdot 406$ | $4 \cdot 073$ | 5．140 | $4 \cdot 752$ | 5．875 | 5.430 | $6 \cdot 6$ | $6 \cdot 109$ | $47 \frac{1}{4}$ |
| $43^{\circ}$ | 3.410 | $4 \cdot 388$ | $4 \cdot 092$ | 5．119 | $4 \cdot 774$ | 5.85 I | $5 \cdot 456$ | $6 \cdot 582$ | 6．138 | $48^{\circ}$ |
| $43 \frac{1}{4}$ | 3.426 | 4.370 | 4．111 | 5.099 | $4 \cdot 796$ | 5.827 | 5．481 | $6 \cdot 555$ | $6 \cdot 167$ | $46 \frac{3}{4}$ |
| 43 $\frac{1}{2}$ | 3.442 | $4 \cdot 352$ | 4．130 | 5.078 | $4 \cdot 818$ | 5．8o3 | $5 \cdot 507$ | $6 \cdot 528$ | $6 \cdot 195$ | $46 \frac{1}{2}$ |
| 433 | 3.458 | $4 \cdot 334$ | $4 \cdot 149$ | 5.057 | 4.841 | $5 \cdot 779$ | 5.532 | $6 \cdot 501$ | 6． 224 | 461 |
| $44^{\circ}$ | 3.473 | $4 \cdot 316$ | 4．168 | 5.035 | 4.863 | $5 \cdot 755$ | 5.557 | $6 \cdot 474$ | $6 \cdot 252$ | $46^{\circ}$ |
| $44 \frac{1}{4}$ | 3.489 3.505 | $4 \cdot 298$ | $4 \cdot 187$ | 5.014 | 4.885 | 5．730 | $5 \cdot 582$ | $6 \cdot 447$ | $6 \cdot 280$ | $45 \frac{3}{4}$ |
| $44 \frac{1}{2}$ | 3.505 3.5 | $4 \cdot 280$ | $4 \cdot 206$ | $4 \cdot 993$ | $4 \cdot 906$ | $5 \cdot 706$ | $5 \cdot 607$ | $6 \cdot 419$ | 6.308 | $45 \frac{1}{2}$ |
| 443 | 3.520 3.536 | $4 \cdot 261$ | 4．22．4 | $4 \cdot 971$ | $4 \cdot 928$ | 5.68 t | 5.632 | $6 \cdot 392$ | 6.336 | $45 \frac{1}{4}$ |
| $45^{\circ}$ | 3.536 | $4 \cdot 243$ | $4 \cdot 243$ | $4 \cdot 950$ | $4 \cdot 950$ | 5.657 | $5 \cdot 657$ | $6 \cdot 364$ | $6 \cdot 364$ | $45^{\circ}$ |
| ¢ | Lat． | Dep | Lat． | Dep | Lat． | Dep． | Lat． | Dep． | Lat． | ロ |
|  | 5 |  |  |  |  |  |  |  |  | － |

TABLE OF CHORDS: $\{$ Radius $=1.0000]$.

| M. | $0^{\circ}$ | $1{ }^{\circ}$ | $2^{\circ}$ | $3^{\circ}$ | $4{ }^{\circ}$ | $5{ }^{\circ}$ | $6^{\circ}$ | $7{ }^{\circ}$ | $8^{\circ}$ | $9^{\circ}$ | $10^{\circ}$ | m. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 0 | . 0 | . 0175 |  |  | .0690 |  | -1047 | -1221 |  |  |  |  |
| 1 | . 0003 | - 0177 | .o352 | . 0526 | . 0701 | . 0875 | - 1050 | - 1224 | - I398 | . 1572 | . 1746 |  |
| 2 | -000 | - or 80 | . 0355 | . 0529 | . 0704 | . 0878 | - 1053 | . 1227 | -1401 | . 1575 | . 1749 | 2 |
| 3 | -0009 | . -183 | . 0358 | . 0532 | . 0707 | . 0881 | - 1055 | . 1230 | . 1404 | -1578 | . 1752 | 3 |
| 4 | - 0 OI 2 | . 0186 | .o36r | . 0535 | . 0710 | . 0884 | - 1058 | . 1233 | . 1407 | . 158 I | . 1755 | 4 |
| 5 | - 001 | - 0189 | . 0364 | . 0538 | . 0713 | . 0887 | -106I | . 1235 | . 1410 | . 1584 | -1758 | 5 |
| 6 | -0017 | - 0192 | . 0366 | .0541 | . 0715 | . 0890 | - 1064 | . 1238 | . 1413 | . 1587 | - 1761 | 6 |
| 7 | - 0 | - org5 | . 0369 | . 0544 | . 0718 | . 0893 | - 1067 | . 1241 | -1415 | . 1589 | - 1763 | 7 |
| 8 | -0023 | - or 98 | .0372 | . 0547 | . 0721 | . 0896 | - 1070 | . 1244 | -1418 | . 1592 | - 1766 | 8 |
|  | - 0026 | - 0201 | .0375 | - 0 | . 07 | . 0899 | - IC73 | - 1247 | . 1421 |  | -1769 | 9 |
| 10 | - 0029 | . 0204 | . 0378 | . 0553 | -0727 | .0901 | -1076 | -1250 | . 1424 | . 1598 | -1772 | 0 |
| 11 | -0032 | - 0 |  |  | . 0730 |  | 9 | 3 | . 1427 |  |  |  |
| 12 | . 0035 | . 0209 | . 0384 | . 0558 | . 0733 | . 0907 | . 1082 | . 1256 | . 1430 | . 1604 | -1778 | 12 |
| 13 | . 0038 | .0212 | . 0387 | . 0561 | . 0736 | -0910 | - 1 | . 1259 | . 1433 | . 1607 | -1781 | 3 |
| 14 | . 004 I | . 0215 | . 0390 | . 0564 | .0739 | -0913 | . 1087 | - 1262 | . 1436 | -1610 | - 1784 | 4 |
| 15 | - 0044 | - 02 | -0393 | . 0567 | . 0742 | . 0916 | - 1090 | - 1265 | -1439 | - 1613 | - 1787 | 5 |
| 16 | . 0047 | - | -0396 | . 0570 | . 0745 | .0919 | -1093 | - 1267 | -1442 | -1616 | -1789 | 6 |
| 17 | -0049 | - | - 0398 | . 0573 | . 0747 | .0922 | - 1096 | - 1270 | . 1444 | -1618 | -1792 | 7 |
| 18 | - 0052 | . 0227 | -040I | . 0576 | . 0750 | -0925 | - 1099 | -1273 | . 1447 | -1621 | -1795 | 8 |
| 19 | - 0055 | .0230 | - 0404 | . 0579 | . 0753 | .0928 | 2 | - 1276 | - 1450 | - 1624 | - I798 | 19 |
| 20 | - 0058 | . 0233 | - 0407 | . 0582 | . 0756 | .093 I | - I | -1279 | . 1453 | -1627 | -1801 | 2.0 |
| 21 | -00 | . 0 |  |  |  |  |  | 2 | 56 | O |  | 21 |
| 22 | - | .0239 | .0413 | - 0 | . 0762 | -0936 | -1111 | - 1285 | . 1459 | 33 | . 1807 | 22 |
| 23 | - 00 | . 024 I | .0416 | . 0 | . 0765 | . 0939 | - I | . 1288 | - 1462 | - 1636 | . 1810 | 23 |
| 24 | - 0070 | . 0244 | . 0419 | . 0593 | -0768 | . 0942 | . 1 | . 1291 | - I 465 | . 1639 | . 1813 | 4 |
| 25 | -0073 | . 0247 | . 0422 | . 0596 | .0771 | . 0945 | - II | . 1294 | - 1468 | -1642 | . 1816 |  |
| 26 | -0076 | . 025 | . 0425 |  | . 0774 | -0948 | -1122 | - 1296 | -1471 | -1645 | . 1818 | 6 |
| 27 | - 0079 | . 0253 | . 0428 | . 0602 | -0776 | -095 | - I I | - 299 | . 1473 | -1647 | -1821 | 7 |
| 28 | -008 | . 0256 | .0430 | . 0605 | . 0779 | . 0954 | -1128 | - 1302 | . 1476 | - 1650 | -1824 | 28 |
| 29 | - 0 | . 0259 | . 0433 | . 0608 | .0782 | -0957 | . 113 | 1305 | . 1479 | - 1653 | 27 | 29 |
| 30 | -0087 | . 0262 | - 0436 | .0611 | -0785 | .0960 | . 1134 | . 1308 | . 1482 | - 1656 | -1830 |  |
| 31 |  | . 0265 | . 0439 |  |  |  |  |  |  |  | 3 | 31 |
| 32 | . 0093 | . 0268 | . 0442 | .0617 | . 0791 | . 0965 | . 1140 | . 1314 | . 1488 | . 1662 | . 1836 |  |
| 33 | -0096 | . 0 | . 0445 | . 0619 | . 0794 | -0968 | . 1143 | .1317 | . 1491 | - 1665 | -1839 | 33 |
| 34 | - 0099 | . 0273 | . 0448 | -0622 | . 0797 | . 0971 | - 1145 | . 1320 | . 1494 | - 1668 | -1842 |  |
| 35 | - oro | . 0276 | - 045 I | -0625 | -0800 | . 0974 | -1 I 48 | . 1323 | - 1497 | -1671 | - 1845 | 35 |
| 36 | - 0 | . 0279 | - 0454 | .0628 | -0803 | . 0977 | -1151 | . 1325 | . 1500 | . 1674 | . 1847 | 36 |
| 37 | . 0 | . 028 | - 0457 | .063I | - c806 | . 0980 | . 1154 | . 1328 | . 1502 | . 1676 | - 1850 |  |
| 38 | - 0 | . 0285 | -0.460 | . 0634 | . 0808 | . 0983 | . 1157 | . 133 r |  | 1679 | - I 853 | 38 |
| 39 | - 011 | . 0288 | - 0462 | . 0637 | .0811 | . 0988 | - 1160 | . 1334 | . 1508 | 1682 | I 856 | 39 |
| 40 | . 0116 | .0291 | - 0465 | - 0640 | -0814 | . 0989 | - I I 63 | . 1337 | -15II | . 1685 | 9 | 40 |
| 41 |  | . 0294 | . 0468 | . 0643 | .0817 | -0992 | 66 | . 1340 | . 1514 | . 1688 | - 1862 | 41 |
| 42 | - 0122 | - 0297 | . 0471 | . 0646 | . 0820 | -0994 | - I 169 | . 1343 | .1517 | -1691 | - 1865 | 42 |
| 43 | . 0125 | . 0300 | . 0474 | . 0649 | . 0823 | - 0997 | -1172 | . 1346 | . 1520 | -1694 | - 1868 | 43 |
| 44 | - 0128 | - 0303 | . 0477 | .065 I | . 0826 | - 1000 | -117 | . 1349 | . 1523 | -1697 | -1871 | 44 |
| 45 | -or3i | - 0305 | -0480 | . 0654 | . 0829 | - 1003 | -1177 | . 1352 | - 1526 | . 1700 | . 1873 | 45 |
| 46 | -or34 | -0308 | . 0483 | . 0657 | . 0832 | -1006 | -118 | . 1355 | . 1529 | . 1703 | . 1876 | 46 |
| 47 | - or37 | - 03 l I | . 0486 | .0660 | . 0835 | -1009 | . 1183 | . 1357 | . 153 I | -1705 | - 1879 | 47 |
| 48 | - 0140 | .0314 | -0489 | . 0663 | . 0838 | - 1012 | . 1186 | . 1360 | . 1534 | -1708 | - 1882 | 48 |
| 49 | . 0143 | .0317 | . 0492 | . 0666 | . 0840 | -1015 | -1189 | . 1363 | . 1537 | 711 | . 1855 | 49 |
| 50 | . 0145 | . 6320 | . 0494 | . 0669 | . 0843 | -1018 | -1192 | - 1366 | . 1540 | -1714 | - i 888 | 5 |
| 5ı | -0148 | . 0323 | . 0497 | . 0672 | . 0846 | -102I | -1195 | -1369 | . 1543 | -1717 | - IS91 | 51 |
| 52 | .0151 | . 0326 | . 0500 | .0675 | . 0849 | -1023 | - II98 | -1372 | . 1546 | - 1720 | . 1894 | 52 |
| 53 | -0154 | .0329 | -05n3 | .0678 | . 0852 | - 1026 | -1201 | -1375 | . 1549 | -1.723 | . 1897 | 53 |
| 54 | - 0157 | .0332 | . 0.506 | .0681 | . 0855 | -1029 | -1204 | -1378 | . 1552 | -1726 | - iquo | 54 |
| 55 | - 0160 | . 0335 | . 0509 | . 0683 | . 0858 | - 1032 | - 1206 | . 1381 | . 1555 | -1729 | -1902 | 55 |
| 56 | - or 63 | -0337 | .0512 | - 0686 | -0861 | -1035 | -1209 | 1384 | . 1558 | -1732 | - Iqu5 | 56 |
| 57 | - or 66 | - 3440 | . 0515 | . 0689 | -0864 | - 1038 | -1212 | - I 386 | - 1560 | -1734 | - 1908 | 57 |
| 58 | -0169 | -0343 | .0518 | .0692 | -0867 | -104I | . 1215 | - 1389 | - 1563 | -1737 | . 1911 | 58 |
| 59 | -0172 | o346 | -052I | .0695 | -0569 | -1044 | 1218 | -1392 | -1560 | 17 | -1914 | 59 |
| 60 | -0175 | . 0349 | . 0524 | .0698 | .0872 | -1047 | 122 | . 1395 | . 1569 | . 1743 | 191 | 6o |

TABLE OF CHORDS: [Radius $=1.0000$ ].

| m. | $11^{\circ}$ | $12^{\circ}$ | $13^{\circ}$ | $14^{\circ}$ | $15^{\circ}$ | $16^{\circ}$ | $17^{\circ}$ | $18^{\circ}$ | $19^{\circ}$ | $20^{\circ}$ | $21^{\circ}$ | 4. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| ${ }^{\prime}$ | -1917 | - 2091 | - 2264 | - 2437 | . 261 | . 2783 | - 2956 |  | . 3301 | . 3473 | - 3645 |  |
| 1 | -102 | - 2093 | . 2267 | - 2440 | -2613 | . 2786 | - 2959 | .3132 | . $33 \times 4$ | 3.4-6 | - 3n748 |  |
| 2 | -1923 | - 2096 | . 2270 | - 2443 | - 2616 | . 2789 | - 2962 | .3134 | . 3307 | 34-9 | 3nto |  |
| 3 | - 1926 | - 2099 | . 2273 | - 2446 | -2619 | . 2792 | - 2965 | .3137 | . 3310 | . 3482 | . 3655 | 3 |
| 4 | -1928 | - 2102 | . 2276 | . 2449 | - 2622 | . 2795 | - 2968 | . 3140 | . 3312 | . 3484 | . 3n56 | 4 |
| 5 | -1931 | - 2105 | . 2279 | . 2452 | - 2625 | - 2798 | - 2971 | -3143 | .33:5 | . 3487 | - 3659 | 5 |
| 6 | -1934 | - 2108 | - 2281 | - 2455 | - 2628 | - 2801 | - 2973 | -3ı 46 | -33ı8 | . 3490 | - 3662 | 6 |
|  | -1937 | - 2111 | . 2284 | - 2458 | - 2631 | -2804 | - 2976 | .3149 | -332I | -3493 | - 3665 | 7 |
| 8 | - 1940 | - 2 | . 2287 | - 2460 | - 2634 | - 2807 | - 2979 | -3!52 | . 3324 | - 3496 | - 3668 |  |
| 9 | -1943 | - 2117 | - 2290 | - 2463 | - 2636 | - 2809 | - 2982 | -3155 | . 3327 | . 3499 | - 3670 | 9 |
| 10 | -1946 | - 2119 | - 2293 | - 2466 | - 2639 | - 28 I 2 | - 2985 | -3157 | . 3330 | . 3502 | - 3673 | 0 |
| 1 |  | - 2 | - 2 | - 2 | - 2642 |  | 8 | - 3160 | . 3333 | . 3504 | - 3676 | II |
| 12 | -1952 | - 2125 | . 2299 | - 2472 | - 2645 | -2818 | 2991 | -3ı63 | . 3335 | . 3507 | - 3679 | 12 |
| 13 | - 1955 | - 2128 | - 2302 | - 2475 | - 2648 | - 2821 | . 2994 | - 3 : 66 | - 3338 | . 3510 | - 3682 | 13 |
| 14 | -1957 | -2131 | - 2305 | - 2478 | - 265 I | - 2824 | - 2996 | -3169 | -334 | . 35 r 3 | . 3685 | 4 |
| 15 | -1960 | - 2134 | - 2307 | - 2481 | - 2654 | - 2827 | - 2999 | -3172 | - 3344 | -35ı6 | - 3688 | 15 |
| 16 | - 1963 | - 2137 | - 2310 | - 2484 | - 2657 | - 2830 | - 3002 | -3ı75 | - 3347 | -3519 | - 3690 | 16 |
| 17 | - 1966 | - 2140 | . 2313 | - 2486 | - 2660 | - 2832 | -3005 | -3178 | - 3350 | . 3522 | -3693 | 17 |
| 18 | -1969 | - 2143 | - 2316 | - 2489 | - 2662 | - 2835 | -3008 | -3180 | . 3353 | - 3525 | - 3696 | 8 |
| 19 | -1972 | - 2146 | -2319 | - 2.492 | - 2665 | - 2838 | -301 1 | 3183 | . 3355 | . 3527 | - 3699 | 19 |
| 20 | -1975 | - 2148 | - 2322 | - 2495 | - 2668 | - 284 I | -3014 | - 3186 | - 3358 | . 3530 | -3702 | 20 |
| 21 | -19 |  |  |  |  | - 2844 |  |  | . 336 I | - 3533 |  | 21 |
| 22 | -1981 | - 2154 | - 2328 | - 2501 | - 2674 | - 2847 | -3019 | -3192 | - 3364 | - 3536 | - 3708 | 22 |
| 23 | - 1983 | - 2157 | - 2331 | - 2504 | - 2677 | - 2850 | -3022 | -3195 | - 3367 | - 3539 | . 3710 | 2 |
| 24 | - 1986 | - 2160 | - 2333 | - 2507 | - 2680 | - 2853 | -3025 | -3198 | - 3370 | - 3542 | -3713 | 24 |
| 25 | -1989 | - 2163 | . 2336 | . 2510 | - 2683 | - 2855 | - 3028 | - 3200 | . 3373 | . 3545 | . 3716 | 25 |
| 26 | -1992 | - 2166 | - 2339 | -2512 | - 2685 | - 2858 | -3o3I | - 3203 | . 3376 | . 3547 | -3719 | 26 |
| 27 | - 1995 | - 2169 | . 2342 | - 2515 | - 2688 | - 2801 | - 3034 | - 3206 | . 3378 | . 3550 | - 3722 | 27 |
| 28 | - 1998 | - 2172 | - 2345 | - 25 I 8 | - 2691 | - 2864 | - 3037 | - 3209 | . 3381 | - 3553 | - 3725 | 8 |
| 29 | - 2001 | - 2174 | - 2348 | - 2521 | - 2694 | - 2867 | - 3040 | -3212 | . 3384 | - 3556 | - 3728 | 29 |
| 3c | - 2054 | - 2177 | - 235 I | - 2524 | - 2697 | - 2870 | -3042 | - 3215 | - 3387 | - 3559 | -3730 | 30 |
| 3 I | - 2 | - 2 | - 2 | - 2527 |  | - 2873 | . 3045 | . 3218 | . 3390 | . 3562 | . 3733 | 31 |
| 32 | - 2 | . 2183 | . 2357 | - 2530 | - 2703 | - 2876 | - 3048 | - 3221 | . 3393 | . 3565 | . 3736 | 32 |
| 33 | -2 | - 2186 | - 2359 | . 2533 | - 2706 | - 2878 | -305I | - 3223 | . 3396 | - 3567 | - 3739 | 33 |
| 34 | -2015 | . 2189 | - 2362 | - 2536 | - 2709 | . 2881 | - 3054 | - 3226 | . 3398 | . 3570 | - 3742 | 34 |
| 35 | - 20 | - 2192 | - 2365 | - 2538 | -2711 | . 2884 | - 3057 | - 3229 | -3401 | . 3573 | - 3745 | 35 |
| 36 | - 2 | . 2195 | - 2368 | . 2541 | - 2714 | - 2887 | - 3060 | - 3232 | -3404 | . 3576 | . 3748 | 36 |
| 37 | - 2 | - 2198 | -2371 | - 2544 | - 2717 | - 289 | -3063 | - 3235 | - 3407 | - 3579 | -3750 | 37 |
| 38 | - 2 | - 2200 | - 2374 | - 2547 | - 2720 | . 2893 | - 3065 | - 3238 | . 3410 | . 3582 | - 3753 | 38 |
| 39 | - 2030 | - 22203 | - 2377 | - 2550 | - 2723 | - 2896 | - 3068 | -3241 | -3413 | . 3585 | - 3756 | 39 |
| 4 c | - 2033 | - 2206 | - 2380 | - 2553 | - 2726 | - 2899 | - 3071 | - 3244 | -3416 | . 3587 | -3759 | 40 |
| 4 I | - 2 | - 2204 | - 2383 | . 2556 | - 2729 | - 2902 | - 3074 | - 3246 | .3419 | . 3590 |  | 41 |
| 42 | - 2038 | - 2212 | - 2385 | - 2559 | - 2732 | - 2904 | - 3077 | - 3249 | . 342 I | -3593 | . 3765 | 42 |
| 43 | - 2041 | . 2215 | - 2388 | - 256 t | - 2734 | - 2907 | - 3080 | - 3252 | . 3424 | - 3596 | - 3768 | 43 |
| 44 | - 2044 | . 22 | - 2391 | - 2564 | -2.737 | - 2910 | - 3083 | - 3255 | - 3427 | - 3599 | - 3770 | 44 |
| 45 | - 2047 | - 2221 | - 2394 | - 2567 | - 2740 | - 2913 | - 3086 | - 3258 | . 3430 | . 3602 | - 3773 | 45 |
| 46 | - 2050 | . 22 | - 2397 | - 2570 | - 2743 | - 2916 | - 3088 | - 3261 | - 3433 | - 3605 | - 3776 | 46 |
| 4. | - 2053 | - 2226 | - 2400 | - 2573 | -27-16 | - 2919 | -3091 | - 3264 | - 3436 | - 3608 | - 3779 | 47 |
| 48 | - 2056 | - 2229 | - 2.403 | - 2576 | - 2749 | - 2922 | -3094 | - 3267 | -3439 | -36ı0 | . 3782 | 48 |
| 49 | - 2059 | . 2232 | - 2406 | . 2579 | - 2752 | - 2925 | -3097 | - 3269 | -344I | -36ı3 | . 3785 | 49 |
| 50 | - 2062 | - 2235 | - 2409 | - 2582 | . 2755 | - 2927 | - 3roo | - 3272 | . 3444 | -36ı6 | . 3788 | 5 |
| 51 | - 2065 | - 2238 | - 2411 | - 2585 | - 2758 | - 2930 | . 3 ro3 | - 3275 | . 3447 | -3619 | . 3790 | 51 |
| 52 | - 2067 | - 2241 | - 2414 | - 2587 | 2760 | - 2933 | - 3 io6 | - 3278 | . 3450 | - 3622 | - 3793 | 52 |
| 53 | - 2070 | - 2244 | - 2417 | - 2590 | . 2763 | - 2936 | -3ro9 | -328I | . 3453 | - 3625 | - 3796 | 53 |
| 54 | - 2073 | - 2247 | - 2420 | - 2593 | . 2766 | - 2939 | -3ıII | - 3284 | - 3456 | . 3628 | - 3799 | 54 |
| 55 | - $21-6$ | . 2250 | - 2423 | - 2596 | - 2769 | - 2942 | -3ıI4 | - 3287 | - 3459 | . 3630 | . 3802 | 55 |
| 56 | - 2079 | - 2253 | - 2426 | - 2599 | - 2772 | - 2945 | -3117 | -3289 | - 3462 | - 3633 | . 3805 | 56 |
| 57 | - 2082 | - 2255 | . 2429 | - 2602 | - 2775 | - 29 尔 8 | - 31 20 | - 3292 | - 3464 | - 3636 | . 3808 | 57 |
| 58 | - 2085 | - 2258 | - 2432 | - 2605 | - 2778 | - 2950 | -3123 | - 3295 | - 3467 | -3639 | - 38 ıo | 58 |
| 59 | - 2088 | -2261 | - 2434 | - 2608 | - 2781 | - 2953 | -3126 | - 3298 | -3470 | . 3642 | .3813 | 59 |
| 20 | 209 | - 226 | . 2437 | - 261 | - 2783 | -2956 | 3129 | . 330 |  | 13645 | -38ı6 |  |

TABLE OF CHORDS: [RADIUS $=1.0000$ ].

| M. | $22^{\circ}$ | $23^{\circ}$ | $24^{\circ}$ | $25^{\circ}$ | $26^{\circ}$ | $27^{\circ}$ | $28^{\circ}$ | $29^{\circ}$ | $30^{\circ}$ | $31{ }^{\circ}$ | 3 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| - | . 3816 | . 3987 | . 4158 | . 4329 | - 4499 | - 4669 | - 4838 | . 5008 | . 5176 | . 5345 | .55r3 |  |
|  | - 3819 | . 3990 | -4161 | . 4332 | . 4502 | -4672 | . 4841 | . 5010 | . 5179 | . 5348 | . 5516 |  |
| 2 | . 3822 | . 3993 | - 4164 | - 4334 | - 4505 | -4675 | -4844 | . 5013 | .5182 | . 5350 | . 5518 |  |
| 3 | . 38 | . 3996 | 4167 | - 4337 | - 4508 | - 4677 | -4847 | . 512 | . 5185 | . 5353 | . 5 | 3 |
| 4 | . 3828 | - 3999 | -4170 | - 4340 | -4510 | -4680 | -4850 | . 5019 | . 5188 | -5356 | . 5524 | 4 |
| 5 | - 3838 | -4002 | -4172 | - 4343 | - 4513 | . 4683 | - 4853 | . 5022 | . 5190 | - 5359 | . 5527 | 5 |
| 6 | . 3833 | - 4004 | -4175 | - 4346 | - 4516 | . 4686 | -4855 | . 5024 | . 5193 | . 5362 | . 5530 | 6 |
| 7 | . 3836 | -4007 | . 4178 | . 43439 | -4519 | - 4689 | . 4858 | . 5027 | -5196 | . 5364 | . 55532 | 7 |
| ) | - 384 | -4013 | . 418 | - 4354 | - 4525 | - 4694 | - 4864 | . 5033 | . 5 | . 537 | . 5538 | 9 |
| 10 | - 3845 | - ¢016 | -4187 | - 4357 | -4527 | - 4697 | - 4867 | . 5036 | . 5204 | . 5373 | . 554 I |  |
| 11 | . 3848 | - 4019 | -4190 | - 4360 | -453 | . 47 | -4869 | . 5039 | . 5207 | . 5376 | . 554 |  |
| 12 | . 3850 | -4022 | -4192 | - 4363 | - 4533 | -4703 | - 4872 | . 5041 | . 5210 | . 5378 | . 5546 | 12 |
| 13 | . 3853 | - 4024 | -4195 | - 4366 | -4536 | -4706 | . 4875 | . 5044 | . 5213 | . 538 I | . 5549 | 13 |
| 14 | - 3856 | - 4027 | -4198 | . 4369 | -4539 | -4708 | . 4878 | . 5047 | . 5216 | . 5384 | . 5552 | 14 |
| 15 | . 3859 | . 403 | - 4201 | -4371 | - 4542 | -4711 | -4881 | . 5050 | . 5219 | . 5387 | . 5555 | 15 |
| 16 | -3862 | - 4033 | - 4204 | - 4374 | -4544 | -4714 | -4884 | - 5053 | . 5221 | . 5390 | . 5557 | 6 |
|  | - 3865 | - 4036 | - 4202 | - 4377 | - 4547 | -4717 | - 4886 | . 5055 | . 2224 | . 5392 | . 5560 | 17 |
| 18 | - 3868 | - 4039 | - 4209 | - 4380 | - 4550 | - 4720 | - 4889 | - 5058 | - 5227 | . 5395 | . 5563 | 18 |
| 19 | - 3870 | - 4042 | -4212 | - 4383 | - 4553 | -4723 | -4892 | . 5061 | . 5230 | -5.398 | . 5566 | 19 |
| 20 | . 3873 | - 4044 | - 4215 | - 4386 | -4556 | -4725 | -4895 | . 5064 | . 5233 | . 5401 |  | 20 |
| 21 | .3876 |  |  |  |  | -4728 |  |  |  |  |  | 21 |
| 22 | . 3879 | - 4050 | -4221 | -4391 | -456I | -473I | - 4901 | . 5070 | . 5238 | . 5406 | . 5574 | 22 |
| 23 | . 3882 | - 4053 | - 4224 | . 4394 | - 4564 | -4734 | -4903 | . 5072 | . 5241 | . 5409 | . 5577 | 23 |
| 24 | . 3885 | - 4056 | - 4226 | - 4397 | -4567 | -4737 | - 4906 | . 5075 | . 5244 | . 5412 |  | 24 |
| 25 | - 3888 | - 4059 | - 422 | -4400 | -4570 | -4740 | - 4909 | -5078 | . 5247 | . 5415 | . 5583 | 5 |
| 26 | . 3890 | - 4061 | - 4232 | -4403 | -4573 | -4742 | -4912 | . 5081 | . 5249 | - 5418 | . 5585 | 26 |
| 27 | . 3893 | - 4064 | - 4235 | . 4405 | - 4576 | -4745 | -4915 | . 5084 | . 5252 | . 5420 | . 5588 | 7 |
| 28 | . 3896 | -4067 | - 4238 | - 4408 | - 4578 | - 4748 | - 4917 | - 5086 | . 5255 | . 5423 | . 559 y 1 | 28 |
| 29 | . 3899 | -4170 | - 4241 | -4411 | - 4581 | - 475 | -4920 | . 5089 | . 5258 | . 5426 |  | 29 |
| 30 | - 39 | - 4073 | - 4244 | -4414 | -4584 | - 47 | -4923 | -5092 | . 5261 | . 5429 | 97 | 30 |
| 31 | - 3905 |  | . 4246 |  |  |  |  |  | . 5263 | .5432 |  | 31 |
| 32 | - 3908 | - 4079 | - 4249 | -4420 | . 4590 | - 4759 |  | . 5098 | . 5266 | . 5434 | . 5602 | 32 |
| 33 | . 3910 | - 4081 | -4252 | -4422 | - 4593 | -4762 | - 4932 | . 5100 | . 5269 | . 5437 | . 5605 | 33 |
| 34 | . 3913 | - 4084 | - 4255 | -4425 | -4595 | - 4765 | - 4934 | . 5103 | . 5272 | . 5440 | . 5608 | 34 |
| 35 | -3916 | - 408 | - 4258 | -4428 | - 4598 | - 4768 | -4937 | . 5106 | . 5275 | . 5443 | . 5611 | 35 |
| 36 | -3919 | -4090 | -4261 | -4431 | -4601 | -4771 |  | . 5109 | . 5277 | - 5446 | . 56 r 3 | 30 |
| 38 | -3922 | -4093 | - 4263 | . 4434 | -4604 | - 4773 | - 4943 |  |  | . 5448 |  | 37 |
| 38 | . 3925 | -4096 | -4266 | - 4437 | - 4607 | - 4776 | -4946 | . 5115 | . 5283 | . 545 I |  | 38 |
| 39 | - 3727 | - 4098 | - 4269 | -4439 | - 4609 | - 4779 | -4948 | . 5117 | . 5286 | . 5454 |  | 39 |
| 40 | -39 | -4101 | - 4 | - 4442 | - 46 | - 47 | 495 | . 51 | . 5289 | . 5457 | . 5025 | 40 |
| 41 | . 3933 |  | . 4275 |  | -4615 | -4785 | -4954 | . 5123 |  | 460 |  | 41 |
| 42 | 3936 | -4107 | -42.78 | . 4448 | -4618 | -4788 | - 4957 | . 5126 | . 5294 | . 5462 | . 5630 | 42 |
| 43 | 3939 | -4110 | -4280 | . 4451 | -4621 | -4790 | - 4960 | . 5129 | . 5297 | -5465 | . 5633 | 43 |
| 44 | . 3942 | . 4113 | - 4283 | . 4454 | -462.4 | -4793 | - 4963 | -5131 | . 5300 | - 5468 | . 5636 | 44 |
| 结 | . 3945 | -4116 | - 4286 | . 4456 | -4626 | -4796 | - 496 | . 5134 | . 5303 | . 5471 | . 5638 | 45 |
| 46 | . 3947 | -4118 | -4289 | . 4459 | -4629 | -4799 | - 4968 | - 5137 |  | - 5474 | 5641 | 46 |
| 47 | -3950 | - 4121 | -4292 | - 4462 | -4632 | -4802 | - 4971 | - 5140 | . 5308 | . 5476 | . 5644 | 47 |
| 48 | -3953 | -4124 | -4295 | - 4465 | - 4635 | -4805 | - 4974 | . 5143 | .5311 | . 5479 | . 5647 | 48 |
| 49 | 3956 | .4127 | - 4298 | -4468 | -4638 | -4807 | -4977 |  |  | . 54885 | . 5650 | 49 |
| 50 | -3959 | -413 | -43no | -4.771 | -4641 | - 48 | - 4979 |  | $.53$ | . 5485 | $\cdot 56$ | 50 |
| 5I | . 3962 | .4133 | . 4303 | -4474 | . 4643 | -4813 | . 4982 | . 5151 | . 5320 | . 5488 | . 5655 | 51 |
| 52 | - 3965 | -4135 | -4306 | . 4476 | - 4646 | -4816 | - 4955 | . 5154 | . 5322 | . 5490 | . 5658 | 52 |
| 53 | -3967 | -4138 | -4309 | - 4479 | - 4649 | -4819 | - 4988 | . 5157 | . 5325 | . 5493 | . 5661 | 53 |
|  | -3970 | -4141 | -4312 | -4482 | - 4652 | -4822 | - 4991 | . 5160 | . 5328 | -5496 | . 5664 | 5 |
| 5 | -3973 | 4:44 | -4315 | . 4485 | - 4655 | -4824 | - 4994 |  |  | . 5499 | . 5666 | 5 |
| 56 | - 3976 | - 4147 | - 4317 | -4488 | - 4658 |  |  | . 5165 | . 5334 | - 5502 | . 5669 | 56 |
| 58 | -3979 | .4150 .453 | -4320 | . 44491 | - 4660 | . 4830 | . 4999 | - 5168 | . 5336 | . 5504 |  | 57 |
| 58 |  | . 4153 | . 43323 | . 44493 | . 46663 | . 48383 | . 50002 | .5171 .5174 . | . 5339 | . 55 | . 5675 |  |
| $\begin{aligned} & 59 \\ & 60 \end{aligned}$ | $\cdot$ $\cdot$ $\cdot 3985$ | -415 | -423 | . 44499 | - 466 | - 4838 | 5 |  | . 53 | 55 r 3 | . 568 。 | 9 |

TABLE OF CHORDS: [RADIUS $=1.0000]$.

| M. | $33^{\circ}$ | $34^{\circ}$ | $35^{\circ}$ | $\mathbf{3 6}^{\circ}$ | $37^{\circ}$ | $38^{\circ}$ | $39^{\circ}$ | $40^{\circ}$ | $41^{\circ}$ | $42^{\circ}$ | $43^{\circ}$ | . |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $0^{\prime}$ | . 568 | - 584 | . 6014 | . 61 | . 6346 | . 651 I | . 6676 |  | - 7004 | -7167 |  |  |
| 1 | . 5683 | . 5850 | . 6017 | -6183 | . 6349 | . 65.4 | . 6679 | . 6843 | 7007 | - 7170 | - 7333 |  |
| 2 | . 5686 | . 5853 | . 6020 | .6186 | . 6352 | . 6517 | . 6682 | . 6846 | 7010 | 7173 | 7335 |  |
| 3 | . 5689 | . 5856 | . 6022 | . 6189 | . 6354 | . 6520 | . 6684 | . 6849 | - 70 | -7176 | -7338 |  |
| 4 | . 5691 | . 5859 | . 6025 | -6191 | . 6357 | . 6522 | . 6687 | . 6851 | - 7015 | 7178 | -734 1 | 4 |
| 5 | . 5694 | . 5861 | - 6028 | -6194 | . 6360 | . 6525 | . 6690 | . 6354 | - 7018 | -7181 | - 7344 | 5 |
| 6 | . 5697 | . 5864 | . 60.31 | -6197 | - 6363 | . 6528 | . 6693 | . 6857 | - 70 | - 7184 | - 7346 | 6 |
| 7 | . 5700 | . 5867 | -6o34 | - 6200 | . 6365 | . 6531 | - 6695 | . 6860 | - 7023 | - 7186 | - 734 |  |
| 8 | - 5703 | . 5870 | -6o36 | - 6202 | - 6368 | . 6533 | - 6698 | . 6862 | - 7 | -7189 | -7352 | $\bigcirc$ |
| 9 | . 5705 | . 5872 | . 6039 | . 6205 | . 6371 | . 6536 | . 6701 | . 6865 | -7029 | -7192 | - 7354 | 9 |
| 10 | . 5708 | . 5875 | -6042 | . 6208 | . 6374 | . 6539 | . 6704 | . 6868 | -7031 | -7195 | 357 | 0 |
| 11 |  |  |  |  |  |  |  |  |  |  |  |  |
| 12 |  | . 5881 | - 6047 | -6214 | . 6379 | . 6544 | . 6709 | . 6873 | -7037 | - 7200 | - 7362 | 12 |
| 13 | . 5717 | . 5884 | -6050 | . 6216 | . 6382 | . 6547 | . 6712 | . 6876 | -7040 | - 7203 | - 7365 | 13 |
| 14 | -5719 | - 5886 | - 6053 | - 6219 | . 6385 | . 6550 | . 6715 | . 6879 | -7042 | -7205 | - 7368 | 14 |
| 15 | . 5722 | . 5889 | -6o56 | - 6222 | . 6387 | . 6553 | . 6717 | . 6881 | - 7045 | - 7208 | -7371 | 5 |
| 16 | -5725 | . 5892 | -6o58 | - 6225 | .6390 | . 6555 | . 6720 | . 6884 | -7048 | -7211 | - 7373 | 16 |
| 17 | - 5728 | . 5895 | . 6061 | - 6227 | . 6393 | . 6558 | . 6723 | . 6887 | -7050 | -7214 | - 7376 | 17 |
| 18 | . 5730 | - 5897 | - 6064 | . 6230 | . 6396 | . 656 r | . 6725 | . 6890 | -7053 | - 7 | - 7379 | 8 |
| 19 | . 5733 | - 5900 | - 6067 | -6233 | . 6398 | . 6564 | -6728 | -6892 | - 7056 | 19 | -7381 | 9 |
| 20 | - 5736 | - 5903 | - 6070 | - 6236 | . 6401 | - 6566 | .673I | . 6895 | -7059 | - 7222 | - 7384 | 20 |
| 21 |  |  | -6072 | - 6238 | . 6404 | . 6569 |  |  | -7061 | - 7224 | - 7387 | 21 |
| 22 | - 5742 | - 5909 | . 6075 | -624I | - 6407 | . 6572 | -6736 | . 6901 | - 7 | - 7227 | - 7390 | 2 |
| 23 | - 5744 | -5911 | -6078 | - 6244 | -6410 | . 6575 | -6739 | . 6903 | - 7067 | -7230 | - $73{ }^{\prime} 9^{2}$ | 23 |
| 24 | . 5747 | . 5914 | . 6081 | - 6247 | -6412 | . 6577 | - 6742 | - 6906 | - 7069 | - 7232 | -7395 |  |
| 25 | . 5750 | . 5917 | . 6083 | -6249 | -6415 | . 6530 | - 6745 | -6919 | -7072 | - 7235 | -7398 | 25 |
| 26 | . 5753 | - 5920 | . 6086 | . 6252 | -64ı8 | . 6583 | - 6747 | -6911 | - 7075 | - 7238 | -740) | 26 |
| 27 | - 5756 | - 5922 | - 6089 | - 6255 | -642I | . 6586 | -6750 | . 6914 | -7078 | - 724 I | -7403 | 27 |
| 28 | - 5758 | . 5925 | -60g2 | - 6258 | -6423 | . 6588 | -6753 | . 6917 | -7080 | - 7243 | - 7406 | 28 |
| 29 | -5761 | - 5928 | -6095 | -626o | -6426 | . 6591 | -6756 | . 6920 | - 7083 | - 7246 | - 7408 | 29 |
| 30 | . 5764 | . 593 I | - 6097 | - 6263 | . 6429 | . 6594 | -6758 | 6922 | - 7086 | - 7249 | -7411 | 30 |
| 3 I |  | - 5934 | - 6 | - 6266 | . 6432 | . 6597 | . 6761 | -6925 | 089 | -725I | -7414 | 31 |
| 32 | - 5769 | . 5936 | . 6103 | - 6269 | . 6434 | . 6599 | . 6764 | - 6928 | -7091 | - 7254 | -7417 | 32 |
| 33 | . 5772 | . 59.39 | -6ı6 | - 6272 | . 6437 | -6602 | . 6767 | -693i | -7094 | - 7257 | -7419 | 33 |
| 34 | -5775 | - 5942 | -6io8 | - 6274 | -6440 | . 6605 | . 6769 | - 6933 | -7097 | - 7260 | - 7422 | 34 |
| 35 | - 5778 | . 5945 | -6III | - 6277 | - 6443 | . 6608 | - 6772 | -6936 | - 7099 | - 7262 | - 7425 | 35 |
| 36 | -578 I | . 5947 | -6ir 4 | - 6280 | - 6445 | . 6610 | . 6775 | . 6939 | - 7102 | - 7265 | - 7427 | 36 |
| 37 | - 5783 | . 5950 | -6117 | - 6283 | - 6448 | .66ı3 | . 6777 | -6941 | -7105 | - 7268 | -7430 | 37 |
| 38 | - 5786 | . 5953 | -6119 | - 6285 | -645I | -66ı6 | . 5780 | -6944. | 8 | - 7270 | - 7433 | 38 |
| 39 | . 5789 | . 5956 | .6122 | - 6288 | . 6454 | -6619 | . 6783 | - 6947 | - 7110 | - 7273 | - 7435 | 39 |
| 40 | . 5792 | . 5959 |  | . 6291 | . 6456 | . 662 I | . 6786 | . 6950 | -711 | - 7276 | . 7438 | 40 |
| 4 |  |  | -6128 |  |  |  | - 6788 |  |  | - 7279 |  | 41 |
| 42 | . 5797 | . 5964 | . 6130 | - 6296 | . 6462 | - 6627 | -6791 | . 6955 | 7118 | -728I | - 7443 | 42 |
| 43 | - 580 | . 5967 | -6133 | - 6299 | . 6465 | . 6630 | - 6794 | - 6958 | -7121 | - 7284 | - 7446 | 43 |
| 44 | . 5803 | . 5970 | -6136 | -6302 | - 6467 | - 6632 | - 6797 | -6961 | 7 | - 7287 | - 7449 | 44 |
| 45 | . 5806 | . 5972 | -6139 | -6305 | -6470 | - 6635 | . 6799 | - 6963 | -7127 | -7289 | - 7452 | 45 |
| 46 | -5808 | . 5975 | .6142 | -6307 | . 6473 | . 6638 | . 6802 | - 6966 | -7129 | - 7292 | - 7454 | 46 |
| 47 | . 5811 | - 5978 | -6144 | -6310 | . 6476 | -6640 | . 6805 | -69ti9 | - 7132 | - 7295 | - 7457 | 47 |
| 48 | - 58 | . 598 I | -6147 | -63ı3 | . 6478 | -6643 | - 6808 | -6971 | -7135 | - 7298 | - 7460 | 48 |
| 49 |  | . 5984 | -6150 | -63ı6 | -6481 | -6646 | . 6810 | . 6974 | -7137 | - 7300 | - 7462 | 49 |
| 50 |  | 86 | -6153 | -63 | - 6484 | -6649 | - 6 | . 6977 | - 71 | - 7303 | - 7465 | 50 |
| 51 | - 5822 | . 5989 | . 6155 | . 632 I | . 6487 | - 665 I | . 6816 | -6980 | , | 306 | - 7468 | 51 |
| 52 | . 5825 | . 5992 | . 6158 | . 6324 | . 6489 | - 6654 | .6819 | . 6982 | -7146 | - 7308 | - 7471 | 52 |
| 53 | - 5828 | . 5995 | -6161 | - 6327 | -6492 | - 6657 | -682 1 | - 6985 | -7148 | -731 I | -7473 | 53 |
| 54 | -5831 | - 5997 | -6164 | -6330 | . 6495 | -6660 | . 6824 | - 6988 | -7151 | -73*4 | - 7476 | 54 |
| 55 | . 5834 | -6000 | -6166 | . 6332 | - 6498 | - 6662 | - 6827 | -6991 | -7154 | 73ı6 | - 7479 | 55 |
| 56 | - 5836 | -6003 | -6169 | . 6335 | . 6500 | - 6665 | . 6829 | - 6993 | -7156 | -7319 | - 7481 | 56 |
| 57 | . 5839 | -6006 | -6172 | - 6338 | -65o3 | -6668 | . 6832 | - 6996 | - 7159 | - 7322 | - 7484 | 57 |
| 58 | - 5842 | -6009 | -6175 | -6341 | . 6506 | -6671 | - 6835 | - 6999 | - 7162 | -7325 | - 7487 | 58 |
| 59 | - 5845 | -6nit | -6178 | - 6343 | . 6509 | - 6673 | . 6838 | - 7001 | -7165 | 732 | - 7489 | 59 |
| 60 | - 5847 | . 601 | -6ı80 | - 63 | . 65 I 1 | . 6676 | . 684 | - 7004 | -716 | -733o | - $749^{2}$ | 60 |


| A |  |  |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\underline{4}$ | $44^{\circ}$ | $45^{\circ}$ | $46^{\circ}$ | $47^{\circ}$ | $48^{\circ}$ | $49^{\circ}$ | 50 | $51^{\circ}$ | $52^{\circ}$ | $53^{\circ}$ | ${ }^{\circ}$ | $\underline{1}$ |
| $0^{\prime}$ | - 7 | - 7654 | -781 | - 7975 | . 8135 | . $8 \times 94$ |  | . 8610 | . 8767 | . 8924 | O | o' |
| 1 | - 7495 | - 7656 | -7817 | - 7978 | .8ı37 | . 8297 | . 8455 | .86ı 3 | . 8770 | . 8927 | - 9082 |  |
|  | - 7498 | - 7659 | - 7820 | - 7980 | . 8140 | . 8299 | . 8458 | . 8615 | -8773 | . 8929 | - 9085 | 2 |
| 3 | - 7500 | - 7662 | - 7823 | - 7983 | .8143 | . 8302 | .846o | -8618 | . 8775 | .893:2 | - 9088 | 3 |
| 4 | -7503 | - 7664 | - 7825 | - 7986 | .8145 | .8304 | . 8463 | . 8621 | . 8778 | . 8934 | - 9090 | 4 |
| 5 | -7506 | - 7667 | - 7828 | - 7988 | .8148 | .8307 | . 8466 | . 8623 | . 8780 | . 8937 | - 9093 | 5 |
| 6 | -7508 | - 7670 | -783 I | . 7991 | .8151 | .83io | -8468 | . 8626 | . 8783 | . 8940 | - 9095 | 6 |
| 7 | -7511 | - 7672 | - 7833 | - 7994 | .8153 | .8312 | .8471 | . 8629 | . 8786 | . 8942 | -9098 | 7 |
| 8 | . 7514 | - 7675 | - 7836 | - 7996 | -8ı56 | .83.5 | . 8473 | . 863 I | . 8788 | . 8945 | -9IUI | 8 |
| 9 | -75ı6 | - 7678 | - 7839 | - 7999 | -8159 | .8318 | . 8476 | . 8634 | . 8791 | . 8947 | -9103 |  |
| 10 | -7519 | -768ı | - 784 t | . 8002 | .8161 | . 8320 | . 8479 | . 8636 | . 8794 | . 8950 | -9106 | 0 |
| I I |  | -7683 |  | . 8004 |  | . 8323 |  |  |  |  |  | II |
| 12 | - 7524 | - 7686 | - 7847 | . 8007 | -8ı67 | . 8326 | . 8484 | . 8642 | . 8799 | . 8955 | . 9111 | I 2 |
| I 3 | - 7527 | $\cdot 7689$ | - 7849 | . 8010 | .8169 | . 8328 | . 8487 | . 8644 | . 8801 | . 8958 | 9113 | 13 |
| 14 | - 7530 | $\cdot 7691$ | - 7852 | . 8012 | -8172 | -833I | -8489 | . 8647 | . 8804 | -8960 | -9116 | 14 |
| 15 | - 7533 | - 7694 | - 7855 | . 8015 | -8ı75 | . 8334 | . 8492 | . 8650 | -8807 | . 8963 | -9119 | 5 |
| 16 | - 7535 | - 7697 | - 7857 | - 8018 | -8177 | . 8336 | . 8495 | . 8652 | .8809 | - 8966 | -9121 | 6 |
| 17 | - 7538 | - 7699 | - 7860 | . 8020 | -8180 | .8339 | . 8497 | . 8655 | .88I2 | . 8968 | -9124 | 17 |
| 18 | -7541 | -7702 | - 7863 | . 8023 | -8183 | -834I | -8500 | . 8657 | -88ı4 | -8971 | -9126 | 8 |
| 19 | - 7543 | -7705 | - 7865 | . 8026 | -8185 | . 8344 | . 8502 | . 8660 | -8817 | -8973 | -9129 | 19 |
| 20 | - 7546 | - 7707 | - 7868 | -8028 | -8188 | . 8347 | . 8505 | . 8663 | - 8820 | - 8976 | -9132 | 20 |
| 21 |  | - 7 |  | . 803 I |  | . 8349 | . 8508 | - 8665 | . 8822 |  |  | 21 |
| 22 | -755I | -7713 | - 7873 | . 8034 | -8193 | . 8352 | . 8510 | - 8668 | . 8825 | . 8981 | -9137 | 22 |
| 23 | - 7554 | -7715 | - 7876 | . 8036 | -8196 | . 8355 | .85ı3 | . 8671 | . 8828 | . 8984 | -9139 | 2 |
| 24 | - 7557 | -7718 | - 7879 | .8039 | -8198 | . 8357 | . 85 ı6 | -8673 | . 8830 | . 8986 | -9142 | 24 |
| 25 | - 7560 | -7721 | - 7882 | - 8042 | - 8201 | . 8360 | . 8518 | . 8676 | . 8833 | . 8989 | -9145 | 25 |
| 26 | - 7562 | -7723 | - 7884 | -8044 | . 8204 | . 8363 | .8521 | - 8678 | . 8835 | . 8992 | -9147 | 6 |
| 27 | - 7565 | - 7726 | - 7887 | . 8047 | - 8206 | . 8365 | .8523 | -8681 | . 8838 | . 8994 | -9150 | 27 |
| 28 | - 7568 | - 7729 | - 7890 | -8050 | - 8209 | . 8368 | . 8526 | - 8684 | . 884 I | . 8997 | -9152 | 8 |
| 29 | - 7570 | -7731 | - 7892 | . 8052 | .8212 | .8371 | . 8529 | - 8686 | . 8843 | . 8999 | -9155 | $2{ }^{2}$ |
| 30 | $\cdot{ }^{7} 573$ | - 7734 | - 7895 | . 8055 | . 8214 | .8373 | . 8531 | -8689 | . 8846 | - 9002 | -9157 | 30 |
| 31 |  | - 7737 | - 7898 | . 8058 | . 8217 | . 8376 | . 8534 | . 8692 | . 8848 | - 9005 |  | 3 I |
| 32 | - 75 | -7740 | - 7900 | . 8060 | . 8220 | . 8378 | . 8537 | . 86944 | . 885 I | - 9007 | - 9163 | 32 |
| 33 | - 758 | - 7742 | - 7903 | . 8063 | . 8222 | . 838 I | .8539 | . 8697 | . 8855 | -9010 | -9165 | 33 |
| 34 | - 7584 | - 7745 | - 7906 | - 8066 | . 8225 | . 8384 | . 8542 | -8699 | . 8850 6́ | -9012 | -9168 | 34 |
| 35 | - 75 | - 7748 | - 7908 | . 8068 | . 8228 | . 8386 | . 8545 | . 8702 | . 8859 | -9015 | -9170 | 35 |
| 36 | -7589 | - 7750 | -7911 | -8071 | . 8230 | . 8389 | . 8547 | - 8705 | .886I | -9018 | -9173 | 36 |
| 37 | - 7592 | - 7753 | - 7914 | . 8074 | . 8233 | .8392 | . 8550 | . 8707 | . 8864 | -9020 | - 9176 | 37 |
| 38 | -7595 | - 7756 | -7916 | -8076 | . 8236 | . 8394 | . 8552 | . 8710 | . 8867 | - 9023 | -9178 | 38 |
| 39 | - 7597 | - 7758 | -7919 | - 8079 | . 8238 | . 8397 | . 8555 | .8712 | . 8869 | -9025 | -918I | 39 |
| 40 | -7600 | -7761 | - 7922 | . 8082 | . 824 I | . 8400 | . 8555 | . 8715 | . 8872 | -9028 | - 9183 | 10 |
| 41 |  |  |  | . 8084 | . 8244 | . 8402 | . 8560 | -8718 | . 8874 |  | -9186 | 4 |
| 42 | - 7605 | - 7706 | - 7927 | . 8087 | . 8246 | . 8405 | . 8563 | . 8720 | . 8877 | -9033 | - 9188 | 42 |
| 43 | - 760 | -7769 | - 7930 | - 8090 | . 8249 | -8408 | . 8566 | . 8723 | . 8880 | -9036 | -9191 | 43 |
| 44 | -7611 | - 7772 | - 7932 | -8092 | . 825 I | -8410 | . 8568 | . 8726 | . 8882 | -9038 | -9194 | 44 |
| 45 | -76ı3 | - 7774 | - 7935 | -8095 | . 8254 | . 84 I 3 | .8571 | -8728 | . 8885 | -904I | -9196 | 45 |
| 46 | -7616 | - 7777 | - 7938 | -8098 | . 8257 | -8415 | . 8573 | -873I | . 8887 | -9044 | -9199 | 46 |
| 47 | -7619 | 7780 | - 7940 | -8100 | . 8259 | -8418 | . 8576 | -8734 | . 8890 | - 9046 | -9201 | 47 |
| 48 | -762I | -7782 | - 7943 | -8103 | . 8262 | -842I | . 8579 | . 8736 | . 8893 | -9049 | 9204 | 48 |
| 49 | , 624 | - 7785 | - 7946 | -8105 | . 8265 | . 8423 | . 8581 | -8739 | . 8895 | -905 I | - 9207 | 49 |
| 50 | $\cdot 7627$ | - 7788 | - 7948 | - 8108 | . 8267 | - 8426 | . 8584 | . 8741 | . 8898 | -9054 | - 9209 | 5 |
| 51 |  | -7791 | -7951 | -8III | . 8270 | . 8429 | . 8587 | . 8744 | .8900 | . 9056 | 9212 | $5 i$ |
| 52 | - 7632 | - 7793 | - 7954 | -8ı13 | . 8273 | . 8431 | . 8589 | 8747 | . 8903 | - 9059 | - 9214 | 52 |
| 53 | - 7635 | - 7796 | - 7956 | -8ır6 | . 8275 | -8434 | . $859^{2}$ | -8749 | . 8906 | -9062 | 9217 | 53 |
| 54 | - 7638 | - 7799 | - 7959 | 8il9 | . 8278 | -843? | . 8594 | .8752 | . 8908 | - 9064 | -9219 | 54 |
| 55 | - 7640 | -7801 | 79 Cz | .8121 | . 8281 | - 8434 | . 8597 | -8754 | . 8911 | -9067 | - 9222 | 55 |
| 56 | - 7643 | -780-́f | - 7964 | -8124 | . 8283 | . 8442 | - 8600 | . 8757 | . 8914 | - 9069 | 9225 | 56 |
| 57 | - 7646 | - 7807 | - 7967 | -8127 | 8286 | - 8444 | - 8602 | .8760 | . 8916 | -9072 | -9227 | 57 |
| 58 | - 7648 | -7809 | - 7970 | -8129 | . 8289 | . 8447 | -8605 | -8702 | . 8919 | - 9075 | 9230 | 58 |
| 59 | 7651 | -7812 | - 7972 | -8ı32 | . 8291 | -8450 | - 8608 | - 8765 | . 8921 | -9077 | 9232 |  |
| 60 | - |  | - 7975 | . 8 | . 8294 | . $845{ }^{2}$ | .86ı0 | . 8 | 892 |  | -9235 | 60 |

TABLE OF CHORDS: $[\operatorname{RadUS}=1.0000]$.

| m. | $55^{\circ}$ | $56^{\circ}$ | $57^{\circ}$ | $08^{\circ}$ | $59^{\circ}$ | $60^{\circ}$ | $61^{\circ}$ | $62^{\circ}$ | $63^{\circ}$ | $64^{\circ}$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| ' |  |  |  |  |  |  |  |  |  |  | ${ }^{\prime}$ |
|  | -9238 | - $3^{3} 92$ | -9546 | -9699 | -985I | 1.000 | 1.0153 | 1.0303 | I -0452 | 1.0601 |  |
| 2 | -924 | - 9395 | -9548 | -9701 | - 98 | 1. | 6 | 1.0306 | I.0455 | I . 0663 | 2 |
| 3 | - 9243 | -9397 | -9551 | -9704 | -9856 | 1.0 | 1.015 | I -0308 | I.0457 | 1.0606 | 3 |
| 4 | -9245 | -9400 | - 9553 | -9706 | -9859 | I . 0010 | -.oı6ı | I . 0311 | $1 \cdot 0460$ | 1. 0608 | 4 |
|  | - 9248 | -9402 | -9556 | -9709 |  | 1.0013 | 1. | 1.0313 | I 0462 | I.06ı I | 5 |
| 6 | -9250 | - 9405 | -9559 | -9711 |  | 1. | 1.0ı66 | I - 0316 | I 0465 | 1.06ı3 | 6 |
| 7 | -925 | -9407 | -9561 | -9714 | 866 | 1.0018 | I -0168 | $1 \cdot 0318$ | I $\cdot 0467$ | 1.0616 |  |
| 8 | -9256 | -9410 | -9564 | -9717 | -9869 | 1.0020 $1 \cdot 0023$ | i.0ı71 I.or 73 | $1 \cdot 0321$ 1.0323 | 1.0470 1.0472 | 1.0618 | 8 |
| 9 10 | -9258 | -9445 | -9566 | .9719 | -9871 |  | I-0176 | $1 \cdot 0326$ | I 1.0472 <br> 1.0475 | 1-0623 | - |
| 11 |  |  |  |  |  |  |  |  |  |  |  |
| 12 | - 9266 |  |  | 9727 |  | 1. | 1.0 | $1 \cdot 033 \mathrm{I}$ |  |  | 12 |
| 13 | -9268 | -9123 |  | -9729 | -988t | 1.0033 | 1.018 | .033 | 1.0482 | 1.06 | 13 |
| 14 | - 9271 | -9425 | -9579 | -9732 | -9884 | 1.0035 | 1.018 | I 0336 | 1.0485 | 1.0633 | 14 |
| 15 | - 9274 | -9428 | -9581 | -9734 | -9886 | I .oo3 | I. | I - 0338 | 1.0487 | 1.0635 | 15 |
| 16 | -9276 | -9430 | -9584 | -9737 | -9889 | I . 0040 | $1 \cdot 0191$ | 1. | I 0490 |  | 16 |
| 17 | -9279 | -9433 | -9587 | -9739 | -9891 | 1.0043 | $1 \cdot 0193$ | $1 \cdot 0343$ | 1.0492 | I-0640 | 17 |
|  | -9281 | -9436 | -9589 | -9742 | -9894 | 1.0045 | I - 0196 | 1.0346 | 1.0495 | 1-0643 | 18 |
| 19 | - 9288 | -9438 | -9592 | -9744 | -9897 | $1 \cdot 0048$ | $1 \cdot 0198$ | $1 \cdot 0348$ | $1 \cdot 0497$ |  | 19 |
| 20 | -9287 | -9441 | -9594 | -9747 | -9899 | $1 \cdot 0050$ | $1 \cdot 0201$ |  | I . 0500 |  | 20 |
| 21 | -9289 | -9443 |  | -9750 | -9902 | I. | 1.0203 | 3 | 1.0502 |  |  |
| 22 | -9292 | -9446 | -9599 | -9752 | -9904 | 1.0055 | 1. | I $\cdot 0356$ | 1.0504 | $1 \cdot$ | 22 |
| 23 | - 9294 | -9448 |  | -9755 | -9907 | I . 0058 | I $\cdot 0208$ | I . 0358 | 1.0507 | I.0655 | 23 |
| 24 | -9297 | -9451 | -9604 | -9757 | -9909 | Oo | $1 \cdot 0211$ | $\mathrm{I} \cdot 036 \mathrm{I}$ | 1.0509 | I . 0658 | 24 |
| 25 | -9299 | -9454 | -9607 | -9760 | -9912 | I $\cdot 0063$ | $1 \cdot 021$ | I $\cdot 03$ | I 05512 | 1-0660 | 25 |
| 26 | -930 | -9456 | -9610 | -9762 | -9914 | I $\cdot 0065$ | 1. | 1.0366 | 1.0514 | 1-0662 | 26 |
| 27 | -930 | - 9459 | -9612 | -9765 | -9917 | I -0068 | I.0218 | I 0368 | 1.0517 | 1. | 27 |
| 28 | -930 | . 946464 | -9615 | -9767 | -9919 | 1. | 1. | I $\cdot 0370$ - 037 1 | 1.0519 | I -0607 | 28 |
| 30 | -9312 | -9466 | -9620 | -9772 | -9924 | 1•0075 | 1.0226 | 1.0375 | $1 \cdot 0524$ |  | 29 30 |
| 31 | -9315 | -9469 | -9622 |  | -9927 | 1.0078 | 1. | I.0378 | 1.0527 | 1.0675 | 31 |
| 32 | -9317 | -9472 | -9625 | -9778 | -9929 | I -0080 | 1.0231 | 1.0380 | 1.0529 | 1-20077 | 32 |
| 33 |  | -9474 | -9627 | -9780 | -9932 | 1. 0083 | 1.0233 | - 0383 | 1.0532 |  | 33 |
| 34 | -932 | -9477 |  | -9783 |  | 1.008 | 50 | 1.0385 | I -0534 | I | 34 |
| 35 | - 9325 | -9479 | -9633 | $7^{-85}$ | -9937 | 1.0088 | 1.0238 | I - 0388 | $1 \cdot 0537$ | I $\cdot 0688$ | 35 |
| 36 | - 9328 | -9482 | -9635 | . 9788 | -9939 | 1.0091 | I.024 $\mathrm{I} \cdot 024$ | $\mathrm{I} \cdot 0390$ $\mathrm{I} \cdot 039$ | 1.0539 | $1 \cdot 0687$ | 36 |
| 37 | - 3330 | -9484 | -9638 | -9790 | -9942 | 1.0093 | I.0243 | $1 \cdot 0393$ | 1.0542 | I - 0690 | 37 |
| 38 | - 9333 |  |  | -9793 | -9945 | 1-0096 | I-0246 | $1 \cdot 0395$ | I -0544 |  | 38 |
| 38 40 | -9335 |  | -964 | -9795 | -9947 | 1-0098 | I.0248 | I $\cdot 0398$ <br> $1 \cdot 0400$ | 1-0547 | I -0694 | 39 |
| 40 | -9338 | $\cdot 9492$ | -9645 | -9798 | $\cdot 9950$ |  |  | 1-0400 | $1 \cdot 0549$ | ${ }^{1 \cdot 0697}$ | 40 |
| 41 | -9341 |  | -9648 | -9800 | -9952 | 1.0 | 1.0253 | 1.0403 | 1.0551 | 1.0699 | 41 |
| 42 | - 9343 | - 9497 | -9650 | -9803 | -9955 | 1-0106 | 1.0256 | 1.0405 | 1.0554 | 1.0702 | 42 |
| 43 | - 9346 | -9500 | -9653 | -9805 | -9957 | 1.0ios | 1.025 | 1.0408 | I . 0556 | 1-0704 | 43 |
| 44 | - 9348 | -9502 | - 9655 | -9808 | -9960 | 1-01II | 1.0 | 1.0410 | 1.0559 | 1-0707 | 44 |
| 45 | -9351 | -9505 | - 9658 | -0810 | -9962 | I.0113 | I. 0263 | 1.0413 | I -0561 | 1-0709 | 45 |
| 46 | -9353 | - 9507 | -966I | -9813 | -9965 | 1.016 | I. 02 | 1.0415 |  | 1-0712 | 46 |
| 47 | -9356 | -9510 | - 9663 | -9816 | -9967 | I-0118 | 1.0 | 1.0418 | I -0566 | 1-0714 | 47 |
| 48 | -9359 | -951 | - 9666 | -9818 | -9970 | I.0121 | I. 0271 | 1.04 | I -0569 | i $\cdot 0717$ | 48 |
| 49 50 | -9361 | -9515 | -9668 | -9821 | -9972 | $1 \cdot 01$ | $1 \cdot 0273$ | 1.04 | 1.0571 | 1.0719 | 49 |
| 50 | -9364 | -9518 | . 9671 | -9823 | -9975 | 1. | $1 \cdot 02$ | 1.04 | $1 \cdot 0574$ | I 07 | 50 |
| 51 | -9366 | -9520 | -9673 | -9826 | -9977 | 1.0128 | 1.0278 | 1.0428 | I.0576 |  | 51 |
| 52 | - 9369 | -9523 | -9676 | -9828 | -9980 | $1 \cdot 013{ }^{\text {a }}$ | I. 0288 | 1.0430 | 1.0579 | 1:0726 | 52 |
| 53 | -9371 | -9525 | - 9678 | -983I | -9982 | $1 \cdot 0133$ | I. 0283 | 1.0433 | I -0581 | 1-0729 | 53 |
| 54 | -9374 | -9528 | -9681 | -9833 | -9985 | 1-0136 | 1.0286 | I . 0435 | 1.0584 | 1.0731 | 54 |
| 55 | -9377 | -9530 | -9683 | -9836 | -9987 | I.or38 | 1.028 | 1.043 | I. 0586 | 1.0734 | 5 |
| $5 ¢$ | -9379 | -9533 | -9686 | -9838 | -9990 | $1 \cdot 0141$ | I.0291 | 1.0440 | 1.0589 | 1.0736 | 56 |
| 57 58 58 | -9382 | -9538 | -9689 | -9841 | -9992 | I.oil I 0146 | 1.0293 i. 0296 | 1.0443 | I $1 \cdot 059 \mathrm{I}$ | 1.0739 | 57 |
| 59 | - 9387 | -9541 | -9694 | - 9846 | -9998 | $1 \cdot 0148$ | 1.0298 | 1.0447 | 1.0596 | 1.0744 | 5 |
| 60 | -9389 | 5 | -969 | -98 | \%oo | I.015I | - | I .0450 | 05 |  | 60 |

TABLE OF CHORDS: [Raĩils $=1.0000$ ].

| M. | $65^{\circ}$ | $66^{\circ}$ | $67^{\circ}$ | $68^{\circ}$ | $69^{\circ}$ | $70^{\circ}$ | $71^{\circ}$ | $72^{\circ}$ | $73^{\circ}$ | 4. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $0^{\prime}$ |  |  |  |  |  |  |  |  |  |  |
| 1 | I. 0748 |  | I $\cdot 1041$ | I $\cdot 1186$ | I. |  | 6 | 758 |  |  |
| 2 | 1.075 1 | I . 0898 | I 1044 | I. 1189 | I . 1333 | I. 1476 |  | I . 1760 |  |  |
|  | 1.0753 | I.0900 | I . 1046 | I-1191 | I. 1335 | I. 1479 | I $\cdot 162 \mathrm{I}$ | 1.1763 |  |  |
| 4 | I.0756 | I -0903 | I $\cdot 10.48$ | I-1I94 | I. 1338 | I. 148I | I. 1 | I 1765 | I P I 906 |  |
| 5 | I.0758 | I -0905 | 1. | I. II $\mathrm{y}^{6}$ | I - 1340 | I 1.1883 | I. 1626 | 767 |  |  |
| 6 | I $\cdot 0761$ | I $\cdot 0$ |  | 1 I $19^{8}$ | I . 1342 | 1486 | I $\cdot 1628$ | 770 | I-1910 |  |
|  | I $\cdot 0763$ | I 0 | I $\cdot 1056$ | 1-1201 | I. 1345 | I $\cdot 1488$ | I. 163 I | I 1.1772 |  |  |
| 8 | I 0.0766 | $1 \cdot 0912$ |  | 1 | I $\cdot 1347$ | 1.1491 | 33 | 75 |  |  |
| 9 | I. 0 | I.0915 |  | I • 1206 | 1. 1350 |  | I. 1635 | 777 | 1.1917 |  |
| 10 | I. 0 | 1.0917 |  | 1. | I $\cdot 1352$ |  | 8 | 1-1779 | 20 | 0 |
| II |  |  |  |  |  |  |  |  |  |  |
| 12 | I. | I. 0922 |  | 1.1213 | I $\cdot 1357$ | I. | I. 1642 | 4 | 24 | 2 |
| I 3 | 1.07 | 1.09 | 1. 1070 | I. 1215 |  | I. | I. 1645 |  | I. 1927 | + |
| 14 | I. 07 | 1.09 | I 10 | 8 | I . I 362 | I $\cdot 15$ | I. 1647 | 1.1789 | I 1 1929 | 14 |
| 15 | I. O | 1.0929 | I | I 1220 | I | I $\cdot 1507$ | 50 |  |  | 15 |
| 16 | I. 0 | I .0932 | 1. | I . 1222 | I - I 366 | I. 15 I | I. 1 | I $\cdot 1793$ | I. 1934 | 16 |
| 17 | I. 0 | 1.0934 | I 1 | I. 1225 | I. 1369 | I 1515 | I. 1654 |  |  | 17 |
| 18 | I. 0 | 1.0937 | I. | I 1227 | 1.1371 | I $\cdot 1514$ | I . 1657 |  |  | 18 |
| 19 | I $\cdot 0793$ | I -0939 |  |  |  | 1. ${ }^{\text {a }}$ | I. 1659 | Oo | I 1.1941 | 19 |
| 20 | 1.0795 | I $\cdot 0942$ | I. 10 | I. 1232 | I . I 376 | 1.1519 | I. 1661 | I $\cdot 1803$ |  | 20 |
| 21 |  |  |  |  |  |  |  |  |  |  |
| 22 |  | I . 0946 | I. | I. 1237 | I - I 381 | I - 1524 | 6 | I 1807 |  | 22 |
| 23 | I. | I.0949 | I. I | I 1239 | I. 1383 | I. 1526 | I. | I. 1810 |  | 23 |
| 24 | I. 0 | I.095 I | I . IO | I 12 ¢ 2 | I - I 386 | I $\cdot 1529$ | I $\cdot 167 \mathrm{I}$ | 12 |  | 24 |
| 25 | 1. | 1.0954 | I. | I. 1244 | I. I 388 | I $\cdot 153 \mathrm{I}$ | I $\cdot 1673$ | 14 |  | 25 |
| 26 | I. 0 | 1.0956 | I. | I. 1246 | I $\cdot 1390$ | I. I | I $\cdot 1676$ | 1.1817 |  | 26 |
| 27 |  | I -0959 |  |  |  | I - I 536 |  | 19 |  | 27 |
| 28 |  | I.0961 |  |  |  | I $\cdot 1538$ |  |  |  | 28 |
| 29 | 1.0 | I .0963 |  | I $\cdot 1254$ | I. 1398 | I $\cdot 154 \mathrm{r}$ | 1.1683 | I-1824 | I - 1964 | 29 |
| 30 | I. 0 | I -0966 | I-IIII |  | I 1400 | I $\cdot 1543$ | I $\cdot 1685$ |  | I - I 966 | 30 |
| 31 |  |  |  |  |  |  |  |  |  | 1 |
| 32 | 1 | I.0971 | I.IIf6 | I $\cdot 1261$ | I $\cdot 1405$ | I $\cdot 1548$ |  |  |  | 32 |
| 3 | 1.0 | I. | I. I | 1. | I. 1407 |  |  |  | 1-1973 | 33 |
| 3 |  | I. | I. I | I •1266 | I . 1409 | I. 1 |  | $0^{\circ}$ |  |  |
| 35 |  | I . 0978 | I. II 23 |  | I. 1412 | I $\cdot 1555$ | I. 1697 |  | I $\cdot 1978$ | 35 |
| 36 |  | I.0980 | I-1I 26 | I. 1271 | I. 1 | I 155 |  | ¢0 | 1-1980 | 36 |
| 3 |  | I - 0 | I-II28 | I. 1273 | I. 1417 | I $\cdot 1560$ |  | 43 |  |  |
| 38 |  | I . 0985 | I-II3t | I. 1 | I. 1419 | I $\cdot 1562$ |  | . 1845 | I 1-1985 | 38 |
| 39 |  | I. 0 | I. | I. 1 | I. 1421 |  |  |  | I - 1987 | 9 |
| 40 |  | I.00 | 1. |  | I. 1424 |  | I. 1709 |  | I 1990 | 40 |
| 41 |  |  |  |  |  |  |  |  |  | 4 I |
| 42 |  | I 009 | I. 1 |  |  |  |  |  | 994 | 42 |
| 43 | I. | I. 0 | I. I | I. I | Fir 43 I |  |  |  | -1997 | 43 |
| 44 |  | I | I. 1 | I-12 | I'1433 |  |  |  | 999 | 44 |
| 45 | 1 | I $\cdot 1002$ | I. I | 1.1292 | 1-4 43 |  | 1.1720 |  | , | 45 |
| 46 |  |  |  | I - I 295 | I-1438 | I $\cdot 15 \mathrm{~S}$ I | 1.1 | - 1064 | - 200 | 46 |
|  | I.086I | 1.1 | 1. | I 1297 | I. 144 T |  | 1 |  | I. | 47 |
| 48 | I. |  | I | I 1299 | I. 1443 | I. 1586 | I. 1727 |  | I $\cdot 2008$ | 48 |
| 49 | I. 0866 |  | I. I | I - 1302 | I. 1445 | 1.1 |  |  | - 201 | 9 |
| 50 |  | $1 \cdot 1014$ |  | I. | I. 14.48 |  |  |  | I 2013 | 50 |
| 51 |  |  |  | I . 1307 |  | 1.159 |  |  |  | 51 |
| 52 | 1.0873 | I $\cdot 1019$ | 1.1165 | I. 1309 | I $\cdot 1452$ | 1.1595 | 1. |  | 1.2018 | 52 |
| 53 | 1.0876 | I . 1022 | I. 1167 | I. 13 II | I. 1455 | 1-1598 | I $\cdot 1739$ | 1880 | . 2020 | 53 |
| 5 | 1.0878 | 1. | I. I 169 | I. 1314 | I. 1457 | I 1 I 000 | 1.1742 | S 2 | I $\cdot 2022$ | 54 |
| 5 | 1.0 | I $\cdot 1027$ | I-II $7^{2}$ | 1-1316 | I 1460 | 5 | -174 | -1005 | 25 | 55 |
| 56 | 1.0883 | I. 1029 | I.1174 | i. 1319 | I $\cdot 1462$ | I - 1605 | 1. 1746 | 1857 | $1 \cdot 2027$ | 56 |
|  | I. 0885 | I $\cdot 103 \mathrm{I}$ | 1.1177 | I-132I | I - I. 664 | I $\cdot 1607$ | 1.1749 | 1.1889 | I $\cdot 2029$ | 57 |
| 58 | I. 0888 | I $\cdot 1034$ | I. It79 | I -1323 |  | -1609 | t. 175 I | I. IS92; | . 2032 | 58 |
| 59 | 1.0890 | 36 | I. 1181 | I - I 326 | I. 146 | , | 1.1753 | 1.189 |  | 59 |
| 60 | I. 089 | I | I. 1184 | I. 1328 | I. 1472 | I.1614 | $1 \cdot 1750$ | I . ISgós |  | 60 |

TABLE OF CHORDS: [Radius $=1.0000$ ].

| м. |  | $5^{\circ}$ | $76^{\circ}$ | $77^{\circ}$ | $78^{\circ}$ | 89 | $80^{\circ}$ | 81 | $82^{\circ}$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 1 |  |  |  |  | 1. 2722 |  |  |  |  |
|  | 1-2039 |  | 1 | I $\cdot 2453$ | 1.2589 | I $\cdot 2724$ |  |  |  |  |
|  | 1.20 | 1.2180 | 1.2318 |  |  | I $\cdot 2726$ |  | I $\cdot 2993$ |  |  |
|  | $1 \cdot 20$ |  | 1.2320 |  |  | I $\cdot 2728$ |  |  |  |  |
|  | 1.2046 | 1.2 | $1 \cdot 2$ |  |  |  | 1-2865 |  | 1.3130 |  |
|  | , | I. 2 | I. 2 | I $\cdot 2462$ |  | 1.2733 | 1. 2867 | $1 \cdot 3000$ |  |  |
| 6 | I $\cdot 2050$ |  | I. 2327 | I $\cdot 2464$ |  | 1.2735 | 1-2869 | $1 \cdot 3002$ |  |  |
|  |  | I-219 |  |  |  | 1. 2737 | 1.2871 |  |  |  |
| 8 |  | $1 \cdot$ | 1.2332 | I $\cdot 2468$ |  | 1.27 |  |  |  |  |
|  |  | I 2 |  |  |  |  |  |  |  |  |
| 10 | $1 \cdot 2$ | 1.2198 |  |  |  | 1. |  | 1.3011 |  |  |
|  |  |  |  |  |  |  |  |  |  |  |
| I |  |  |  |  |  | , |  |  |  |  |
| 1 | I. 20 | I. 2 | I. 2 | I $\cdot 2480$ | 1. |  | I $\cdot 2885$ | 1.3018 |  |  |
| 14 | I $\cdot 2$ |  |  |  |  | I. 2753 | I . 2887 |  |  |  |
| 15 | I $\cdot 20$ | I. 22 | 1.2 | 1 | 1 | I $\cdot 2755$ |  | I.3022 |  |  |
| 16 | 1.20 | I. 22 |  |  |  |  |  |  | 5 | 6 |
|  | I. 2 | 1.221 |  |  | 1. | 1.2 |  | 1.3027 |  |  |
| 18 |  | 1. |  |  | 1. |  |  |  | I.3i | 8 |
| 19 |  | 1 |  | 1-2493 | 1. | I $\cdot 2$ |  |  |  | 9 |
| 20 | 1. | 1 |  | 1. 2496 | 1. | 1.2766 | 1. | I.3033 |  | 20 |
|  |  |  |  |  |  |  |  |  |  |  |
|  | I $\cdot 20$ | I $\cdot 2$ | I $\cdot 2$ |  | 1 |  |  | 8 |  |  |
| 23 | I $\cdot 2$ | 1 |  |  | 1. | 1. | 1-2907 |  |  | 3 |
| 24 |  | I. 2 |  |  | 1. | 1. |  |  |  |  |
| 25 | I 20 | I $\cdot 2233$ |  |  |  |  | 1.2 |  |  |  |
| 26 |  | 1. |  |  | I - 2645 | 1. |  | 1.3046 |  | 26 |
| 27 |  |  |  |  |  | 1. |  |  |  | 27 |
| 28 |  | I $\cdot 2240$ |  |  | 1-2650 | I $\cdot 2$ | $1 \cdot 2$ |  |  | 88 |
| 29 |  | $1 \cdot 2$ |  |  | I $\cdot 2652$ |  |  | 53 |  | 29 |
| 3 | 1.2106 | I-22 |  |  | 1. | 1-2789 | I 2922 |  |  | 30 |
| 31 |  |  |  |  |  |  |  |  |  |  |
|  |  | 1. |  |  |  |  | I $\cdot 2927$ |  |  | 32 |
|  | 1.2113 |  |  |  |  |  |  |  |  | 3 |
| 34 |  | I 2 |  |  |  |  | I $\cdot 2$ |  |  | 4 |
|  |  | I. 2 |  |  |  | I. | I $\cdot 2934$ | 1. |  |  |
| 36 |  | 1. |  |  | I - 2668 | I |  | I. 3068 | 1.3200 | 36 |
|  |  | 1. |  |  | 1. | I | 1. 29 |  |  |  |
| 38 | I. 21 | I $\cdot 2$ |  |  |  | I $\cdot$ | I $\cdot 29$ |  |  | 38 |
|  |  | 1. |  |  |  |  | 1-2942 |  |  | ) |
| 40 | I $\cdot 2$ |  | I. 2 |  |  |  | I-2945 |  |  |  |
|  |  |  |  |  |  |  |  |  |  |  |
| 42 | I. 2 | 12 | 1. |  |  |  |  | 1.3082 |  | 42 |
| 43 | 1. | 1.22 |  |  |  |  |  | I. 3084 | 1.3215 |  |
|  | 1. | I. 22 |  |  | 1. | $1 \cdot 2$ | 1.29 | 1. | 1. | 44 |
|  |  |  | I 2 |  | 88 | 1-2822 | 1-2956 | I. 3088 | 1. |  |
|  | 1.2143 | 1. | I. 24 |  | I - 2690 |  | 1. 2958 | 1.3090 | I - 3222 | 46 |
| 47 | 2 | I. 2 | 1 |  |  | 1. |  | 3 |  |  |
| 48 |  | 1. |  |  |  |  | I $\cdot 29$ |  |  | 48 |
| 49 |  | 1.2 | I |  |  |  | I $\cdot 2965$ | I.3097 |  | ) |
| 5 | 1. |  | I $\cdot 2428$ |  |  |  | 1-2967 | I 3099 |  | ) |
|  |  |  | I. 2430 |  |  |  |  |  |  |  |
|  | 1. | I $\cdot 22$ | I. | I $\cdot 2568$ |  | I 2838 | 1-297 | 1. | 1.3235 |  |
|  | 1.2159 | I $\cdot 22$ | I. 2 | 1 |  |  | 1. 297 | I. 3 | $1 \cdot 3237$ |  |
|  | 1. |  | I. 2 | 1. | $1 \cdot 2$ | I $\cdot 2$ | 1-2976 | I. 3 |  |  |
|  | 1. | I 23 | 1. 2439 | I $\cdot 2575$ | $1 \cdot 2$ | 1.2 | I 29 |  |  |  |
|  | I. 2 | I - 2304 | I. 2 | 1. | 1. | I $\cdot 28$ | 1-2980 | 1.3112 | 1.3244 | 56 |
|  | I. 2168 | I $\cdot 23$ | 1-2443 | 1 | $1 \cdot 2$ | I $\cdot 2849$ | $1 \cdot 2982$ | 1. | 46 |  |
|  | I. 2171 | $1 \cdot 2$ | 1.2446 |  | 1.2717 | 1.285 I |  |  |  |  |
|  | 1-2173 | I 23 | I $\cdot 2448$ | I - 2584 | I - 2719 | $1 \cdot 2$ | I. 2987 |  |  |  |
|  | 1.2175 | 1.23ı3 | I. |  | 1-2722 | I 2856 | I-2989 |  |  |  |


| TABLE OF CHORDS: [Radius $=1.0000$ ]. |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| M. | $83^{\circ}$ | $84^{\circ}$ | $85^{\circ}$ | $86^{\circ}$ | $88^{\circ}$ | $88^{\circ}$ | $89^{\circ}$ | M. |
| $0^{\prime}$ | 1.3252 | I. 3383 | I. 3512 | I. 3640 | 1.3767 | 1.3893 | 1.4018 | $\mathrm{o}^{\prime}$ |
| 1 | I. 3255 | I. 3385 | 1.3514 | I. 3642 | I. 3769 | 1.3895 | 1.4020 | 1 |
| 2 | I. 3257 | I. 3387 | I.3516 | I. 3644 | I.377 I | I. 3897 | 1.4022 | 2 |
| 3 | I. 3259 | I. 3389 | I.3518 | I. 3646 | I. 3773 | I. 3899 | 1. 4024 | 3 |
| 4 | I.326I | 1.3391 | 1.3520 | I. 3648 | I. 3776 | 1.3902 | 1.4026 | 4 |
| 5 | I. 3263 | 1.3393 | 1.3523 | I.365I | I. 3778 | I.39n4 | 1.4029 | 5 |
| 6 | I. 3265 | 1.3396 | I. 3525 | I. 3653 | I. 3780 | I. 3906 | 1.4031 | 6 |
| 7 | I. 3268 | I.3398 | I. 3527 | I. $\mathrm{E}^{655}$ | 1.3782 | I. 3908 | 1.4033 | 7 |
| 8 | I. 3270 | 1.3400 | I. 3529 | I. 3657 | I. 3784 | I.3910 | I. 4035 | 8 |
| 9 | I. 3272 | 1.3402 | I.353 | I. 3659 | I. 3786 | I.3912 | I. 4037 | 9 |
| 10 | I. 3274 | 1.3404 | I. 3533 | I . 366 I | I. 3788 | I.3914 | 1.4039 | 10 |
| 11 | 1.3276 | I. 3406 | 1.3535 | I. 3663 | 1.3790 | I.3916 | I $\cdot 4041$ | 11 |
| 12 | I. 3279 | I. 3409 | 1.3538 | I. 3665 | 1.3792 | 1.3918 | 1. 4043 | 12 |
| 13 | I.3281 | I. 341 I | 1.3540 | I $\cdot 3668$ | I. 3794 | 1.3920 | 1. 4045 | 13 |
| 14 | 1.3283 | 1.3413 | 1.3542 | I. 3670 | I. 3797 | 1.3922 | I. 4047 | 14 |
| 15 | I. 3285 | I. 3415 | I. 3544 | 1.3672 | I. 3799 | I.3925 | 1.4049 | 15 |
| 16 | I. 3287 | I. 3417 | I. 3546 | 1.3674 | I.3801 | 1.3927 | 1.4051 | 16 |
| 17 | 1.3289 | 1.3419 | I. 3548 | I. 3676 | 1.3803 | 1.3929 | I. 4053 | 17 |
| 18 | I. 3292 | 1.3421 | I. 3550 | I. 3678 | 1.3805 | 1.3931 | I - 4055 | 18 |
| 19 | :. 3294 | I. 3424 | I. 3552 | I. 3680 | I. 3807 | I. 3933 | I. 4058 | 19 |
| 20 | 1.3296 | I. 3426 | I. 3555 | 1.3682 | 1.3809 | I.3935 | 1-4060 | 20 |
| 21 | 1.3298 | 1. 3428 | I. 3557 | I. 3685 | I.381 I | I. 3937 | 1.4062 | 21 |
| 22 | 1.3300 | I. 3430 | 1.3559 | 1.3687 | I.3813 | 1.3939 | I. 4064 | 22 |
| 23 | 1.3302 | 1. 3432 | 1.3561 | I.3689 | 1.3816 | I. 394 I | I $\cdot 4066$ | 23 |
| 24 | 1.3305 | I. 3434 | I. 3563 | I.3691 | I. 3818 | I. 3943 | 1.4068 | 24 |
| 25 | I. 3307 | I. 3437 | I. 3565 | 1.3693 | I. 3820 | I. 3945 | I. 4070 | 25 |
| 26 | I.3309 | I. 3439 | I. 3567 | I. 3695 | I. 3822 | I.3947 | 1.4072 | 26 |
| 27 | I.33ı | I. 3441 | I. 3570 | 1.3697 | I. 3824 | I.3950 | I $\cdot 4074$ | 27 |
| 28 | 1.33ı3 | I. 3443 | 1.3572 | I. 3699 | I. 3826 | I. 395.2 | I. 4076 | 28 |
| 29 | I.33ı5 | I. 3445 | 1.3574 | I.3702 | I. 3828 | I.3954 | 1.4078 | 29 |
| 30 | I.3318 | I. 3447 | 1.3576 | 1.3704 | I.3830 | I. 3956 | 1.4080 | 30 |
| 31 | 1.3320 | I. 3449 | 1.3578 | 1.3706 | I. 3832 | I. 3958 | 1.4082 | 3I |
| 32 | I. 3322 | I. 3452 | 1.3580 | 1.3708 | I. 3834 | I. 3960 | I. 4084 | 32 |
| 33 | I. 3324 | I. 3454 | 1.3582 | 1.3710 | I. 3837 | I. 3962 | 1.4086 | 33 |
| 34 | I. 3326 | I. 3456 | 1.3585 | 1.3712 | 1.3839 | I $\cdot 3964$ | 1.4089 | 34 |
| 35 | I. 33.28 | I. 3458 | 1. 3587 | I. 3714 | I. 3841 | I. 3966 | 1.4091 | 35 |
| 36 | 1.3.331 | I. 3460 | 1.3589 | 1.3716 | I. 3843 | 1.3968 | 1.4093 | 36 |
| 3- | 1. 3333 | I. 3462 | I.3591 | 1.3718 | I. 3845 | I. 3970 | 1.4095 | 37 |
| 38 | 1.3335 | I. 3465 | 1.3593 | 1.3721 | I. 38.47 | 1.3972 | 1.4097 | 38 |
| 39 | 1.3337 | I. 3467 | 1.3595 | 1.3723 | I.38年 9 | I. 3975 | 1.4099 | 39 |
| 40 | 1.3339 | 1.3469 | I.3597 | 1.3725 | I. 3851 | $\underline{\text { I } 3977}$ | 1.4iol | 40. |
| 41 | 1.3341 | I. 3471 | 1.3599 | 1.3727 | 1.3853 | 1.3979 | 1.4103 | 41 |
| 42 | 1.3344 | 1.3473 | 1.3602 | 1.3729 | I. 3855 | I.398I | 1.4105 | 42 |
| 43 | I. 3346 | 1.3475 | 1.3604 | 1.3731 | 1.3858 | I. 3 y 53 | 1.4107 | 43 |
| 44 | 1.3348 | 1.3477 | 1.3606 | 1.3733 | I. 3860 | 1.3985 | 1.4109 | 44 |
| 45 | I.3350 | I.3480 | I. 3608 | 1.3735 | 1.3562 | 1.3957 | 1.4111 | 45 |
| 46 | I.3352 | I.348.2 | I. 3610 | 1.3738 | I. 3864 | 1.3989 | 1.4113 | 46 |
| 47 | I.3354 | I.3484 | I. 3612 | I. 3740 | 1.3866 | 1.3991 | 1.4115 | 4- |
| 48 | I.3357 | I.3486 | I. 3614 | 1.3742 | 1.3868 | I. 3993 | 1.4117 | 48 |
| 49 | 1.3359 | I.3488 | 1.3617 | I.3744 | 1.3870 | 1.3995 | 1.4119 | 49 |
| 50 | I.336I | I $\cdot 3490$ | 1.3619 | I. 3746 | I. 3872 | I. 3997 | 1.4122 | 50 |
| 51 | I. 3363 | 1.3492 | 1.3621 | 1.3748 | I. 3874 | 1. 3999 | 1.4124 | 5 I |
| 52 | I. 3365 | 1.3495 | 1.3623 | 1.3750 | I. $38-6$ | 1.4002 | 1.4126 | 52 |
| 53 | I. 3367 | 1.3497 | 1.3625 | 1.3752 | 1.3879 | - 1.4004 | 1.4128 | 53 |
| 54 | I. 3370 | 1.3499 | I. 3627 | 1.3754 | 1.3881 | I. 4006 | 1.4130 | 54 |
| 55 | 1.3372 | 1.3501 | I. 3629 | I.3757 | I. 3883 | I $\cdot 4008$ | I. 4132 | 55 |
| 56 | I.3374 | I.3503 I. 3505 | I. 363 I | 1.3759 | I.3885 | 1.4010 | 1.4134 | 56 |
| 57 58 | I. 3376 I. $33-8$ | I.3505 I.3508 I | I. 3634 I. 3636 | I. 3761 I. $3-63$ | 1.3887 1.3859 | I. 4012 | I. 4136 I. 4138 | 57 58 |
| 58 | I. 3378 I. 3380 | I.3508 I.35io | I. 3636 I. 3638 | I. 3763 I. $3-65$ | 1.3889 | I. 4014 | I.4138 | 58 |
| 59 60 | I $\cdot 3380$ I .3383 | I.3510 I.35i2 | I. 3638 I. 3640 | I $\cdot 3-65$ I.3-67 | I. 3891 I. 3893 | I•4016 | 1.414 C I. 4142 | 59 60 |
| 60 | I.3383 | 1.3512 | I. 3640 | 1.3-67 | $1 \cdot 3893$ | 1.4010 | 1.4142 | 60 |

## T A B L E 1.

OF

## LOGARITHMS OF NUNBERS

FROM
1 то 10000 .

| N. | Log. | N. | Log. | N. | Log. | N. | Log. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 0.000000 | 26 | 1.414973 | 51 | $1 \cdot 707570$ | 76 | 1.880814 |
|  | 0.30:030 | 27 | 1.431364 | 52 | 1.716003 | 77 | I.886491 |
| 3 | 0.477121 | 28 | $1 \cdot 447158$ | 53 | 1.724276 | 78 | $1 \cdot 892095$ |
| 4 | 0.602060 | 29 | 1.462398 | 54 | $1 \cdot 732394$ | 79 | 1.897627 |
| 5 | 0.698970 | 30 | 1-477121 | 55 | I-740363 | 80 | $1 \cdot 903090$ |
| 6 | 0.778151 | 31 | 1.491362 | 56 | $1 \cdot 748188$ | 81 | $1 \cdot 908485$ |
| 7 | 0.845008 | 32 | 1.505150 | 57 | I. 755875 | 82 | 1.913814 |
| 8 | 0.903090 | 33 | 1.518514 | 58 | $1 \cdot 763428$ | 83 | 1.919078 |
| 9 | -0.954243 | 34 | 1.531479 | 59 | $1 \cdot 770852$ | 84 | 1-924279 |
| 10 | 1-000000 | 35 | 1.544068 | 60 | $1 \cdot 778151$ | 85 | I-929419 |
| 11 | 1.041393 | 36 | 1.556303 | 61 | 1.785330 | 80 | 1-934498 |
| 12 | $1 \cdot 079181$ | 37 | 1.568202 | 62 | 1-792392 | 87 | $1 \cdot 939519$ |
| 13 | $1 \cdot 113943$ | 38 | 1.579784 | 63 | 1.799341 | 88 | I $\cdot 944483$ |
| 14 | 1.146128 | 39 | 1.591065 | 64 | 1.806180 | 89 | I. 949390 |
| 15 | $1 \cdot 176091$ | 40 | 1.602060 | 65 | I-812913 | 90 | I-954243 |
| 16 | 1.204120 | 41 | 1.612784 | 66 | 1.819544 | 91 | I.959041 |
| 17 | $1 \cdot 230449$ | 42 | 1.623249 | 67 | 1.826075 | 92 | 1.963788 |
| 18 | 1-255273 | 43 | 1.633468 | 68 | 1.832509 | 93 | I. 968483 |
| 19 | I. 278754 | 44 | 1.643453 | 69 | I. 838849 | 94 | 1-973128 |
| 20 | $1 \cdot 301030$ | 45 | 1.653213 | 70 | I. 845098 | 95 | 1-977724 |
| 21 | 1.322219 | 46 | 1.662758 | 71 | 1.851258 | 96 | 1.982271 |
| 22 | I. 342423 | 47 | 1.672098 | 72 | I. 857333 | 97 | 1.986772 |
| 23 | 13361728 | 48 | 1.681241 | 73 | 1.863323 | 98 | 1.991226 |
| 24 | 1.380211 | 49 | 1.690196 | 74 | 1.869232 | 99 | $1 \cdot 995635$ |
| 25 | I $\cdot 397940$ | 50 | 1.698070 | 75 | 1.875061 | 100 | $2 \cdot 000000$ |

N. B. In the following table, in the last nine columns of each page, where the first ar loading figures change from 9's to 0 's, the character $\bullet$ is introduced instead of the 0 's, to catch the eye, and to indicate that from thence the annexed firs! two figures of the Logarithm in the second solumn stand in the next lower linn directly under the asterisk.

| 2 |  | LOGARITHMS |  |  |  | OF NUMBERS. |  |  |  | Table I. |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |
| 10 | 000000 | 0434 | 0858 | :301 | 173.4 | 2166 | 2598 | 3029 | 3461 | 3891 | 432 |
| 101 | 4321 | 4751 | 5181 | 3609 | 6038 | 6466 | 6894 | 7321 | 7748 | 8174 | 428 |
| 102 | * 8600 | 9026 | 9451 | 9876 | -300 | C 924 | 1147 | 1570 | 1993 | 2415 | 424 |
| 103 | OI 2837 | 3259 | 3680 | 4:00 | 4521 | 4940 | 5360 | 5779 | 5197 | 6616 | 419 |
| 104 | * 7033 | 7451 | 7868 | 8284 | 8700 | 9116 | 9532 | 9947 | -361 | 0775 | 416 |
| 105 | 021189 | 1603 | 20:6 | 2428 | 2841 | 3252 | 3664 | 4075 | 4486 | 4896 | 412 |
| 106 | 5306 | 5715 | 6125 | 6533 | 6942 | 7350 | 7757 | 8164 | 8571 | 8978 | 408 |
| 107 | * 9384 | 9789 | +195 | 0600 | 1004 | 1408 | 1812 | 2216 | 2619 | 3021 | 404 |
| 108 | 03 3424 | 3826 | 4227 | 4628 | 5029 | 5430 | 583o | 6230 | 6629 | 7028 | 400 |
| 109 | * 7426 | 7825 | 8223 | 8620 | 9017 | 9414 | 981 1 | +207 | 0602 | 0998 | 396 |
| 110 | 041303 | 1787 | 2182 | 2576 | 2959 | 3362 | 3755 | 4148 | 4540 | 4932 | $3 g^{3}$ |
| 111 | 5323 | 5714 | 6105 | $649^{5}$ | 6885 | 7275 | 7664 | 8053 | 8442 | 8830 | 389 |
| 112 | * ${ }^{218}$ | 9,06 | $9 \% 93$ | +380 | 0766 | 1153 | 1538 | 1924 | 2309 | 2694 | 386 |
| 113 | 053078 | 3463 | 3846 | 4230 | 4613 | 4996 | 5378 | 5760 | 6142 | 6524 | 382 |
| 114 | * 6905 | 7286 | 7666 | 8046 | 8426 | 8805 | 9185 | 9563 | 9942 | +320 | 379 |
| 115 | 060698 | 1075 | 1452 | 1829 | 2206 | 2582 | 2958 | 3333 | 3709 | 4083 | 376 |
| 116 | 4458 | 4833 | 5206 | 5580 | 5953 | 6326 | 6699 | 7071 | 7443 | 7815 | 372 |
| 117 | * 8186 | 855.7 | 8928 | 9298 | 9668 | +038 | 0407 | 0776 | 1145 | 1514 | 369 |
| 118 | 071882 | 2250 | 2617 | 2985 | 3352 | 3718 | 4085 | 4451 | 4816 | 5182 | 366 |
| 119 | 5547 | 5912 | 6276 | 6640 | 7004 | 7368 | 7731 | 8094 | 8457 | 8819 | 363 |
| 120 | *9181 | $9^{5.4} 3$ | 9904 | + 266 | 0626 | 0987 | 1347 | 1707 | 2067 | 2426 | 360 |
| 121 | 082785 | 3144 | 3503 | 3861 | 4219 | 4576 | 4934 | 5291 | 5647 | 6004 | 357 |
| 122 | 6360 | 6716 | 7071 | 7426 | 7781 | 8136 | 8490 | 8845 | 9198 | $2^{552}$ | 355 |
| 123 | * 9905 | - 258 | 06II | -963 | 1315 | 1667 | 2018 | 2370 | 2721 | 3071 | 351 |
| 124 | c9 3422 | 3.772 | 4122 | 4471 | 4820 | 5169 | 5518 | 5866 | 6215 | 6562 | 349 |
| 125 | *6910 | 7257 | 7604 | 7951 | 8298 | 8644 | 8990 | 9335 | $\mathrm{q}^{681}$ | -026 | 346 |
| 126 | 100371 | 0715 | 1059 | 1403 | 1747 | 2091 | 2434 | 2777 | 3119 | 3462 | 343 |
| 127 | 3804 | 4146 | 4487 | 4828 | 5169 | 5510 | 5851 | 6191 | 6531 | 6871 | 340 |
| 128 | * 7210 | 7549 | 7888 | 8227 | 8565 | 8903 | 9241 | 95079 | 9916 | -253 | 338 |
| 129 | 1110590 | 0926 | 1263 | 1599 | 1934 | 2270 | 2605 | 2940 | 3275 | 3609 | 335 |
| 130 | 3943 | 4277 | 4611 | 4944 | 5278 | 5611 | 5943 | 6276 | 6608 | 6940 | 333 |
| 131 | * 7271 | 7603 | 7934 | 8265 | 8595 | 8926 | 9256 | 9586 | 9915 | - 245 | 330 |
| 132 | 120574 | 0903 | 1231 | 1560 | 1888 | 2216 | 2544 | 2871 | 3198 | 3525 | 328 |
| 133 | 3852 | 4178 | 4504 | 4830 | 5156 | 5481 | 5806 | 6131 | 6456 | 6781 | 325 |
| 134 | * 7105 | 7429 | 7753 | 8076 | 8399 | 8722 | 9045 | 9368 | 9690 | -012 | 323 |
| 135 | 130334 | 0655 | 0977 | 1298 | 1619 | ${ }^{19} 93$ | 2260 | 2580 |  |  | 321 |
| 136 | 3539 | 3853 | 4177 | 4496 | 4814 | 5133 | 5451 | 5769 | 6086 | 6403 | 318 |
| 137 138 1 | 6721 $\times 8$ | 7037 | 7354 | 7671 | 7987 | 8303 | 8618 | 8934 | 9249 | 9564 | 315 |
| 138 | * 9879 | +194 | 0508 | 0822 | 1136 | 1450 | 1763 | 2076 | 2389 | 2702 | 314 |
| 139 | 143015 | $332 \%$ | 3639 | $39^{51}$ | 4263 | 4574 | 4885 | 5196 | 5507 | 5818 | 311 |
| 140 | 6128 | 6438 | 6748 | 7058 | 7367 | 7676 | 7985 | 8294 | 8603 | 8911 | 309 |
| 141 | * 9219 | 9527 | 9835 | +142 | 0449 | -756 | 1063 | 1370 | 1676 | 1982 | 307 |
| 142 | 152288 | 2594 | 2900 | 3205 | 3510 | 3815 | 4120 | 4424 | 4728 | 50.32 | 305 |
| 143 | 5336 | 5640 | 5943 | 6246 | 6549 | 6852 | 7154 | 7457 | 7759 | 8061 | 303 |
| 144 | * 8362 | 8664 | 8965 | 9266 | 9567 | 9868 | +168 | 0469 | 0769 | 1068 | 301 |
| 145 | 161368 | 1667 | 1967 | 2266 | 2564 | 2863 | 3161 | 3460 | 3758 | 4055 | 299 |
| 146 | 4353 | 4650 | 4947 | 5244 | 5541 | 5838 | 6134 | 6430 | 6726 | 7022 | 297 |
| 147 | ${ }^{7317}$ | 7613 | 7908 | 8203 | 8497 | 8792 | 9086 | 9380 | 9674 | 9968 | 295 |
| 148 | 170262 | 0555 | 0848 | 1141 | 1434 | 1726 | 2019 | 2311 | 2603 | 2895 | 293 |
| 149 | 3186 | 3478 | 3769 | 4060 | 4351 | 4641 | 4932 | 5222 | 5512 | 5802 | 291 |
| 150 | 6091 | 6381 | 6670 | 6259 | 7248 | 7536 | 7825 | 8113 | 8401 |  | 289 |
| 151 | * 8977 | 9264 | 9552 | 9839 | +126 | 0413 | 0699 | 0085 | 1272 | 1558 | :87 |
| 152 | 181844 | 2129 | 2415 | 2700 | 2985 | 3270 | 3555 | 3839 | 4123 | 4407 | 285 |
| 153 | 4691 | 4975 | 5259 | 5542 | 5825 | 6108 | 6391 | 6674 | 6956 | 7239 | 283 |
| 154 | - 7521 | 7803 | 8084 | 8366 | 8647 | 8928 | 9209 | 9490 | 9771 | -051 | 281 |
| 155 | 190332 | 0612 | $\mathrm{org}^{2}$ | 1171 | 1451 | 1730 | 2010 | ${ }_{5}^{289} 9$ | 2567 | 2846 | 279 |
| 156 | 3125 | 3403 | 3681 | 3959 | 4237 | 4514 | 4792 | 5069 | 5346 | 5623 | 27-8 |
| 157 <br> 158 | 5900 +8657 | 6176 | 6453 | 6729 | 7005 | 7281 | 7506 | 7832 | 8107 | 8382 | 276 |
| 158 | * 8657 | 8932 | 9206 | 9481 | 9755 | -029 | O303 | -5̄77 | 0850 | 1124 | 274 |
| 159 | 201397 | 1670 | 1943 | 2216 | 2488 | 2761 | 3033 | 3305 | 3577 | 3848 | 272 |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |

Table I.
LOGARITHMS OF NUMBERS.
3

| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 160 | 204120 | 4391 | 4663 | 4934 | 5204 | 5475 | 5746 | 6016 | 6286 | 6556 | 271 |
| 161 | 6826 | 7096 | 7365 | 7634 | 7304 | 8173 | 8441 | 8710 | 8979 | 9247 | 269 |
| 162 | - 9515 | 9783 | -051 | 0319 | $\bigcirc 586$ | 0853 | 1121 | 1388 | 1654 | 1921 | 267 |
| 163 | 212188 | 2454 | 2720 | 2986 | 3252 | 3518 | 3783 | 4049 | 4314 | 4579 | 266 |
| 164 | 4844 | 5109 | 5373 | 5638 | 5902 | 6166 | 6430 | 6694 | 6,57 | 7221 | 264 |
| 165 | 7484 | 7747 | 8010 | 8273 | 8536 | 8798 | 9060 | 9323 | 9585 | 9846 | 262 |
| 166 | 220108 | 0370 | 0631 | 0892 | 1153 | 1414 | 1675 | 1936 | 2196 | 2456 | 261 |
| 167 | 2716 | 2976 | 3236 | 3496 | 3755 | 4015 | 4274 | 4533 | 4792 | 5051 | 259 |
| 168 | 5309 | 5568 | 5826 | 6084 | 5342 | 6600 | 6858 | 7115 | 7372 | 7630 | 258 |
| 169 | * 7887 | 8144 | 8400 | 8657 | 8913 | 9170 | 9426 | 9682 | 9938 | +193 | 256 |
| 170 | 230449 | 0704 | og60 | 1215 | 1470 | 1724 | 1979 | 2234 | 2488 | 2742 | 254 |
| 171 | 2996 | 3250 | 3504 | 3757 | 4011 | 4264 | 4517 | 4770 | 5023 | 5276 | 253 |
| 172 | 5528 | 5781 | 6033 | 6285 | 6537 | 6789 | 7041 | 7292 | 7544 | 7795 | 252 |
| $17^{3}$ | * 9046 | 8297 | 8548 | 8799 | 9049 | 9299 | 9550 | 9800 | +050 | 0300 | 250 |
| 174 | 240549 | 0799 | 1048 | 1297 | 1546 | 1795 | 2044 | 2293 | 2541 | 2790 | 249 |
| 175 | 3038 | 3286 | 3534 | 3782 | 4030 | 4277 | 4525 | 4772 | 5019 | 5256 | 248 |
| 176 | 5513 | 5759 | 6006 | 6252 | 6499 | 6745 | 6991 | 7237 | 7482 | 7728 | 246 |
| 177 | * 7973 | 8219 | 8464 | 8709 | 8954 | 9198 | 9443 | 9687 | 9932 | +176 | 245 |
| 178 | 250420 | 0664 | - 088 | 1151 | ${ }^{13} 95$ | 1638 | 1881 | 2125 | 2368 | 2610 | 243 |
| 179 | 2853 | 3096 | 3338 | 3580 | 3822 | 4064 | 4306 | 4548 | 4790 | 50 | 242 |
| 180 | 5273 | 5514 | 5755 | 5996 | 6237 | 6477 | 6718 | 6958 | 7198 | 7439 | 241 |
| 181 | 7679 | 7918 | 8158 | 8398 | 8637 | 8877 | 9116 | 9355 | 9594 | 9833 | 239 |
| 182 | 260071 | -310 | 0548 | 0787 | 1025 | 1263 | 1501 | $17^{3} 9$ | 1976 | 2214 | 238 |
| 183 | 2451 | 2688 | 2925 | ${ }_{3} \mathbf{1 t} 2$ | 3399 | 3636 | 3873 | 4109 | 4346 | 4582 | 23- |
| 184 | 4818 | 5054 | 5290 | 5525 | 5761 | 5996 | 6232 | 6467 | 6702 | 6937 | 235 |
| 185 | 7172 | 7406 | 7641 | 7875 | Silo | 8344 | 8578 | 8812 | 9046 | 9279 | 234 |
| 185 | *9513 | 9746 | 9980 | -213 | 3446 | 0679 | 0912 | 1144 | 1377 | 1609 | 233 |
| 187 | 271842 | 2074 | 2306 | 2538 | ${ }_{5} 770$ | 3001 | 3233 | 3464 | 3696 | 3927 | 232 |
| 188 | 4158 | 4389 | 4620 | 4850 | 5081 | 5311 | 5542 | 5772 | 6002 | 6232 | 230 |
| 189 | 6462 | 6692 | 6.91 | 7151 | 7380 | 7609 | 7838 | 8067 | 8296 | 8525 | 229 |
| 190 | * 8754 | 8982 | 9211 | 9439 | 7667 | 9895 | +123 | 0351 | 0578 | 0806 | 228 |
| 195 | 281033 | 1261 | 1488 | 1715 | 1942 | 2169 | 2396 | 2622 | 2849 | 3075 | 227 |
| 192 | 3301 | 3527 | 3753 | 3979 | 4205 | 4431 | 4656 | 4882 | 5107 | 5332 | 226 |
| 193 | 5557 | 5782 | 6007 | 6232 | 6456 | 6681 | 6905 | 7130 | 7354 | 7578 | 225 |
| 194 | 7802 | 8026 | 8249 | $847^{3}$ | 8696 | 8920 | 9143 | 9366 | 9589 | 9812 | 223 |
| 195 | 290035 | 0257 | 0480 | 0702 | 0925 | 1147 | 1369 | 1591 | 1813 | 2034 | 222 |
| 196 | 2256 | 2478 | 2699 | 2920 | 3141 | 3363 | 3584 | 3804 | 4025 | 4246 | 221 |
| 197 | 4466 | 4687 | 4907 | 5127 | 5347 | 5567 | 5787 | 6007 | 6226 | 6446 | 220 |
| 198 | 6665 | 6884 | 7104 | 7323 | 7542 | 7761 | 7979 | 8198 | 8416 | 8635 | 219 |
| 199 | * 8853 | 9071 | 9289 | 9507 | 9725 | 9943 | -161 | 0378 | -5.g5 | 0813 | 218 |
| 200 | 301030 | 1247 | 1464 | 1681 | 1898 | 2114 | 233I | 2547 | 2764 | 2980 | 217 |
| 201 | 3196 | 3412 | 3628 | 3844 | 4059 | 4275 | 4491 | 4706 | 4921 | 5136 | 216 |
| 202 | 5351 | 5566 | 5781 | 5996 | 6211 | 6425 | 6639 | 6854 | 7068 | 7282 | 215 |
| 203 | 7496 | 7710 | 7924 | 8137 | 8351 | 8564 | 8778 | 8991 | 9204 | 9417 | 213 |
| 204 | * 9630 | 9843 | +056 | 0268 | 0481 | 0693 | O906 | 1118 | 1330 | 1542 | 212 |
| 205 | 31 1754 | 1966 | 217 | 2389 | 2600 | 2812 | 3023 | 3234 | 3445 | 3656 | 211 |
| 206 | 3867 | 4078 | 4289 | 4499 |  | 4920 | 5130 | 5340 | 5551 | 5760 | 210 |
| 207 | 5970 | 6180 | 6300 | 6599 | 6809 | 7018 | 7227 | 7436 | 7646 | 7854 | 209 |
| 208 | 8063 | 8272 | 8481 | 8689 | 8898 | 9106 | 9314 | 9522 | 9730 | 9938 | 208 |
| 209 | 320146 | 0354 | 0562 | 0769 | 0977 | 18 | 1391 | 1598 | 1805 | 2012 | 207 |
| 210 | 2219 | 2426 | 2633 | 2839 | 3046 | 3252 | 3458 | 3665 | 3871 | 4077 | 200́ |
| 211 | 4282 | 4488 | 4694 | 4899 | 5105 | 5310 | 5516 | 5721 | 5926 | 6131 | 205 |
| 212 | 6336 | 6541 | 6745 | 6950 | 7155 | 7359 | 7563 | 7767 | 7972 | 8176 | 204 |
| 213 | * 8380 | 8583 | 8787 | 8991 | 9194 | 9398 | 9601 | 9805 | +008 | 0211 | 203 |
| 214 | 330414 | 0617 | 0819 | 1022 | 1225 | 1427 | 1630 | 1832 | 203 | 22 | 202 |
| 215 | 2438 | 2640 | 2842 | 3044 | 3246 | 3447 | 3649 | 3850 | 405I | 4253 | 202 |
| 216 | 4454 | 4655 | 4856 | 5057 | 5257 | 5458 | 5658 | 5859 | 6059 | 6260 | 201 |
| 217 | 6460 | 6660 | 6860 | 7060 | 7260 | 7459 | 7659 | 7858 | 8058 | 8257 | 200 |
| 218 | * 8456 | 8656 | 8855 | 9054 | 9253 | 9451 | 9650 | 9849 | -047 | 0246 | 199 |
| 219 | 340444 | 0642 | 0841 | 1039 | 1237 | 1435 | 1632 | 1830 | 2028 | 2225 | 198 |
| N. | 0 | 1 |  | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |


| 4 |  | LOGARITHMS OF NUMBERS. |  |  |  |  |  |  | Table 1. |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |
| 220 | 342423 | 2050 | 28:7 | 3014 | 3212 | 3409 | 3606 | 3802 | 3999 | 4196 | 197 |
| 221 | 4392 | 4589 | 4785 | 4981 | $517^{8}$ | 5374 | 550 | 5766 | 5962 | 6157 | 196 |
| 222 | 6353 | 6549 | 6744 | $6{ }_{6} 39$ | 7135 | 9330 | 7525 | 7720 | $7{ }^{1} 15$ | 8110 | 195 |
| 223 | * 8305 | 8500 | 8604 | 8889 | 9083 | 92-8 | 9472 | 9666 | 9860 | -054 | 194 |
| 224 | 350248 | 0442 | 0636 | 0829 | 1023 | 1216 | 1410 | 1603 | 1796 | $19^{3} 9$ | $19^{3}$ |
| 225 | 2183 | 2375 | 2568 | 2761 | 2054 | 3147 | 3339 | 3532 | 3724 | 3916 | $19^{3}$ |
| 226 | 4108 | 4301 | 4493 | 4685 | 4876 | 5068 | 5260 | 5452 | 5643 | 5834 | 192 |
| 227 | 6026 | 6217 | 6408 | 6599 | 6790 | 6981 | 7172 | $7^{363}$ | 7554 | 7744 | 191 |
| 228 | 7935 | 8125 | 8316 | 8506 | 8676 | 8886 | 9076 | 9266 | 9456 | 9646 | $1{ }^{1} 0$ |
| 229 | * 9835 | -025 | 0215 | 0404 | -593 | 0783 | 0972 | 1161 | 135c | 1539 | 189 |
| 230 | 361728 | 1917 | 2105 | 2294 | 2482 | 2671 | 2859 | 3048 | 3236 | 3424 | 188 |
| 231 | 3612 | 3800 | 3988 | 4176 | 4363 | 4501 | $47^{39}$ | 4926 | 5113 | 5301 | 188 |
| 232 | 5488 | 5675 | 5862 | 6049 | 6235 | 6423 | 6610 | 6796 | 6983 | 7169 | 187 |
| 233 | 7356 | 7542 | 7729 | 7915 | 8101 | 8287 | 8473 | 8659 | 8845 | 9030 | 186 |
| 234 | * 9216 | 9401 | 9587 | 9772 | 9958 | -143 | 0328 | 0513 | 0698 | 0883 | 185 |
| 235 | 371068 | 1253 | 1437 | 1622 | 1806 | 1991 | 2175 | 2360 | 2544 | 2728 | 18.4 |
| 236 | 2912 | 3096 | 3280 | 3464 | 3647 | 3831 | 4015 | 4198 | 4382 | 4565 | 184 |
| 237 | 4748 | $49^{32}$ | 5115 | 5298 | 5481 | 5664 | 5846 | 6029 | 6212 | 6394 | 183 |
| 238 | 6577 | 6759 | 6942 | 7124 | 7306 | 7488 | 7670 | 7852 | 8034 | 8216 | 182 |
| 239 | * 8398 | 8580 | 8761 | 8943 | 9124 | 9306 | 9487 | 9668 | 9849 | -030 | 181 |
| 240 | 380211 | 0392 | ${ }^{05} 73$ | 0754 | 0934 | 1115 | 1296 | 1476 | 1656 | 1837 | 181 |
| 241 | 2017 | 2197 | 2377 | 2557 | 2737 | 2917 | 3097 | 3277 | 3456 | 3636 | 180 |
| 242 | 3815 | 3995 | 4174 | 4353 | 4533 | 4712 | 4891 | 5070 | 5249 | 5428 | 179 |
| 243 | 5606 | 5785 | 5964 | 6142 | 6321 | 6499 | 6677 | 6856 | 7034 | 7212 | 178 |
| 244 | 7390 | 7568 | 7746 | 7923 | 8101 | 8279 | 8456 | 8634 | 8811 | 8989 | 178 |
| 245 | *9166 | 9343 | 9520 | 9698 | 9875 | -05I | 0228 | 0405 | 0582 | 0759 | 177 |
| 240 | 390935 | 1112 | 1288 | 1464 | 1641 | 1817 | 1993 | 2169 | 2345 | 2521 | 176 |
| 247 | 2697 | 2873 | 3048 | 3224 | 3400 | 3575 | 3751 | 3926 | 4101 | 4277 | 176 |
| 248 | 4452 | 4627 | 4802 | 4977 | 5152 | 5326 | 5501 | 5676 | 5850 | 6025 | 175 |
| 249 | 6199 | 6374 | 6548 | 6.722 | 6896 | 7071 | 7245 | 7419 | 7592 | 7766 | 174 |
| 250 | 7940 | 8114 | 8287 | 846I | 8634 | 8808 | 898I | 9154 | 9328 | 9501 | 173 |
| 251 | * 9674 | 9847 | -020 | -192 | -365 | 0538 | 0711 | 0883 | 1056 | 1228 | $17^{3}$ |
| 252 | 401401 | 1573 | 1745 | 1917 | 2089 | 2261 | 2433 | 2605 | 2777 | 2949 | 172 |
| 253 | 3121 | 3292 | 3464 | 3635 | 3807 | 3978 | 4149 | 4320 | 4492 | 4663 | 171 |
| 254 | 4834 | 5005 | 5176 | 5346 | 5517 | 5688 | 5858 | 6029 | 6199 | 6370 | 171 |
| 255 | 6540 | 6710 | 6881 | 7051 | 7221 | 7391 | 7561 | 7731 | 7201 | 8070 | 170 |
| 256 | 8240 | 8410 | 8579 | 8749 | 8918 | 9087 | 9257 | 9426 | $9^{5} 0^{5}$ | 9764 | 169 |
| 257 | * 9933 | -102 | 0271 | 0440 | -609 | 0777 | 0946 | 1114 | 1283 | 1451 | 169 |
| 258 | 411620 | 1788 | 1956 | 2124 | 2293 | 2461 | 2629 | 2796 | 2964 | 3132 | 168 |
| 259 | 3300 | 3467 | 3635 | 3803 | 3970 | 4137 | 4305 | 4472 | 4639 | 4806 | 157 |
| 260 | 4973 | 5140 | 5307 | 5474 | 5641 | 5808 | 5974 | 6141 | 6308 | 6474 | 167 |
| 261 | 6641 | 6807 | 6973 | 7139 | 7306 | 7472 | 7638 | 7804 | 7970 | 8135 | 166 |
| 262 | 8301 | 8467 | 8633 | 8798 | 8964 | 9129 | 9295 | 9460 | 9625 | 9791 | 165 |
| 263 | * 9956 | +121 | 0286 | 0451 | 0616 | 0781 | 0945 | 1110 | 1275 | 1439 | 165 |
| 264 | 421604 | 1768 | 1933 | 2097 | 2261 | 2426 | 2590 | 2754 | 2918 | 3082 | 164 |
| 265 | 3246 | 3410 | 3574 | 3737 | 3901 | 4065 | 4228 | 4392 | 4555 | 4718 | 164 |
| 266 | 4882 | 5045 | 5208 | 5371 | 5534 | 5697 | 5860 | 6023 | 6186 | 6349 | 163 |
| 267 | 6511 | 6674 | 6836 | 6999 | 7161 | 7324 | 7486 | 7648 | 7811 | 7973 | 162 |
| 268 | 8135 | 8297 | 8459 | 8621 | 8783 | 8944 | 9106 | 9268 | 9429 | $9^{5} 91$ | 162 |
| 269 | * 9752 | 9914 | -075 | 0236 | -398 | -559 | 0720 | 0881 | 1042 | 1203 | 161 |
| 270 | 431364 | 1525 | 1685 | 1846 | 2007 |  | 2328 | 2488 | 2649 |  | 161 |
| 271 | 2969 | 3130 | 3290 | 3450 | 3610 | 3770 | $30^{30}$ | 4090 | 4249 | 4409 | 160 |
| 272 | 4569 | 4729 | 4888 | 5048 | 5207 | 5367 | 5526 | 5685 | 5844 | 6004 | 159 |
| 273 | 6163 | 6322 | 6481 | 6640 | 6799 | 6957 | 7116 | 7275 | 7433 | 7502 | 15 |
| 274 | 7751 | 7909 | 8067 | 8226 | 8384 | 8542 | 8701 | 8859 | 9017 | 9175 | 155 |
| 275 | * 9333 | 9491 | 9648 | $\varsigma^{\text {Ro6 }}$ | 9264 | -122 | 0279 | 0437 | 0594 | 0752 | 158 |
| 276 | 440909 | 1066 | 1224 | 1381 | 1538 | 1695 | 1852 | 2009 | 2166 | 2323 | 157 |
| 277 | 2480 | 2637 | 2793 | 2950 | 3106 | 3263 | 3419 | 3576 | ${ }_{5} 732$ | 3889 | 157 |
| 278 | 4045 | 4201 | 4357 | 4513 | 4669 | 4825 | $49^{81}$ | 5137 | 5293 | 5449 | 156 |
| 279 | 5604 | 5750 | 5915 | 6071 | 6226 | 6382 | 6537 | 6692 | 6848 | 7003 | 155 |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |


| Table I. |  |  | LOGARITHMS OF NUMBERS. |  |  |  |  |  |  |  | 5 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |
| 280 | 447158 | 7313 | 7468 | 7623 | 7778 | 7933 | 8088 | 8242 | 8397 | 8552 | 155 |
| 281 | * 8706 | 8861 | 9015 | 9170 | 9324 | 9478 | 7633 | 9787 | 9941 | -005 | 154 |
| 282 | 450249 | 0403 | 0557 | 071 1 | 0865 | 1018 | 1172 | 1326 | 1479 | 1633 | 154 |
| 283 | 1786 | 1940 | 2093 | 2247 | 2400 | 2553 | 2706 | 2859 | 3012 | 3165 | 153 |
| 284 | 3318 | 3471 | 3624 | 3777 | 3930 | 4082 | 4235 | 4387 | 4540 | 4692 | 153 |
| 285 | 4845 | 4997 | 5150 | 5302 | 5454 | 5606 | 5758 | 5910 | 6062 | 62:4 | 152 |
| 286 | 6366 | 6518 | 6670 | 6821 | 6973 | 7125 | 7276 | 7428 | 7579 | 7731 | 152 |
| 287 | 7882 | 8033 | 8184 | 8336 | 8487 | 8638 | 8789 | 8940 | 9091 | 9242 | 151 |
| 288 | *9392 | 9543 | 9694 | 9845 | 9995 | +146 | 0296 | 0447 | 0597 | 0748 | 151 |
| 289 | 460898 | 1048 | 1198 | 1348 | 1499 | 1649 | 1799 | 1948 | 2098 | 2248 | 150 |
| 290 | 2398 | 2548 | 2697 | 2847 | 2997 | 3146 | 3296 | 3445 | 3504 | 3744 | 150 |
| 291 | 3893 | 4042 | 4191 | 4340 | 4490 | 4639 | 4788 | 4936 | 5085 | 5234 | 149 |
| 292 | 5383 | 5532 | 5680 | 5829 | 5977 | 6126 | 6274 | 6423 | 6571 | 6719 | 149 |
| 293 | 6868 | 7016 | 7164 | 7312 | 7460 | 7608 | $77^{56}$ | 7904 | 8052 | 8200 | 148 |
| 294 | 8347 | 8495 | 8643 | 8790 | 8938 | 9085 | 9233 | 9380 | 9527 | 9675 | 148 |
| 295 | * 9822 | 9969 | -116 | 0263 | 0410 | 0557 | 0704 | 0851 | 0998 | 1145 | 147 |
| 296 | 471292 | 1438 | 1585 | ${ }_{1} 7^{3} 2$ | 1878 | 2025 | 2171 | 2318 | 2464 | 2610 | 146 |
| 297 | 2756 | 2903 | 3049 | ${ }^{31} 195$ | 3341 | 3487 | 3633 | 3779 | 3925 | 4071 | 146 |
| 298 | 4216 | 4362 | 4508 | 4653 | 4799 | 4944 | 5090 | 5235 | 5381 | 5526 | 146 |
| 299 | 5671 | 5816 | 5962 | 6107 | 6252 | 6397 | 6542 | 6687 | 6832 | 6976 | 145 |
| 300 | 7121 | 7266 | 7411 | 7555 | 7700 | 7844 | 7989 | 8133 | 8278 | 8422 | 145 |
| 301 | 8566 | 8711 | 8855 | 8999 | 9143 | 9287 | 9431 | 9575 | 9719 | 9863 | 144 |
| 302 | 480007 | O151 | 0294 | 0438 | 0582 | 0725 | 0869 | 1012 | 1156 | 1299 | 144 |
| 303 | 1443 | 1586 | 1729 | 1872 | 2016 | 2159 | 2302 | 2445 | 2588 | $27^{31}$ | 143 |
| 304 | 2874 | 3016 | 3159 | 3302 | 3445 | 3587 | 3730 | 3872 | 4015 | 4157 | 143 |
| 305 | 4300 | 4442 | 4585 | 4727 | 4869 | 5011 | 5153 | 5295 | 5437 | 5579 | 142 |
| 306 | 5721 | 5863 | 6005 | 6147 | 6289 | 6430 | 6572 | 6714 | 6855 | 6997 | 142 |
| 307 | 7138 | 7280 | 7421 | 7563 | 7704 | 7845 | 7986 | 8127 | 8269 | 8410 | 141 |
| 308 | 8551 | 8692 | 8833 | 8974 | 9114 | 9255 | 9396 | 9537 | 9677 | 9818 | 141 |
| 309 | * 9958 | -099 | 0239 | 0380 | -5̄20 | 0661 | 0801 | 0941 | 1081 | 1222 | 140 |
| 310 | 491362 | 1502 | 1642 | 1782 | 1922 | 2062 | 2201 | 2341 | 2481 | 2621 | 140 |
| 311 | 2760 | 2900 | 3040 | 3179 | 3319 | 3458 | 3597 | 3737 | 3876 | 4015 | 139 |
| 312 312 | 455 | 4294 | 4433 | 4572 | 4711 | 4850 | 4989 | 5128 | 5267 | 5406 | 139 |
| 313 | 5544 | 5683 | 5822 | 5660 | 6099 | 6238 | 6376 | 6515 | 6653 | 6791 | 139 |
| 314 | 6930 | 7068 | 7206 | 7344 | 7483 | 7621 | $77^{5} 9$ | 7897 | 8035 | 8173 | ェ38 |
| 315 | 8311 | 8448 | 8586 | 8724 | 8862 | 8999 | 9137 | 9275 | 9412 | 9550 | 138 |
| 316 | * 9687 | 9824 | 9962 | -099 | 0236 | 0374 | 0511 | 0648 | 0785 | 0922 | 137 |
| 317 318 | 501059 | 1196 | 1333 | 1470 | 1607 | 1744 | 1880 | 2017 | 2154 | 2291 | 137 |
| 318 | 2427 | 2564 | 2700 | 2837 | 2973 | 3109 | 3246 | 3382 | 3518 | 3655 | 136 |
| 319 | 3791 | 3927 | 4063 | 4199 | 4335 | 4471 | 4607 | 4743 | 4878 | 5014 | 136 |
| 320 | 5150 | 5286 | 5421 | 5557 | 5693 | 5828 | 5964 | 6099 | 6234 | 6370 | 136 |
| 321 | 6505 | 6640 | 6776 | 6911 | 7046 | 7181 | 7316 | 7451 | 7586 | 7721 | 135 |
| 322 32 | 7856 | 7991 | 8126 | 8260 | 8395 | 8530 | 8664 | 8799 | 8934 | 9068 | 135 |
| 323 | * ${ }^{9} 9203$ | 9337 | 9471 | 9606 | 9740 | 9874 | -009 | 0143 | 0277 | 0411 | 134 |
| 324 | 510545 | 0679 | 0813 | 0947 | 1081 | 1215 | 1349 | 1482 | 1616 | 1750 | 134 |
| 325 | 1883 | 2017 | 2151 | 2284 | 2418 | 2551 | 2684 | 2818 | 2951 | 3084 | 133 |
| 326 | 3218 | 3351 | 3484 | 3617 | 3750 | 3883 | 4016 | 4149 | 4282 | 4414 | 133 |
| 327 328 | 4548 | 4681 | 4813 | 4946 | 5079 | 5211 | 5344 | 5476 | 5609 | 5741 | 133 |
| 328 | 5874 | 6006 | 6139 | 6271 | 6403 | 6535 | 6668 | 6800 | 6932 | 7064 | 132 |
| 329 | 7196 | 7328 | 7460 | 7592 | 7724 | 7855 | 7987 | 8119 | 8251 | 8382 | 132 |
| 330 | 8514 | 8646 | 8777 | 8909 | 9040 | 9171 | 9303 | 9434 | 9566 | 9697 | 131 |
| 331 | **9828 | 9959 | -090 | 0221 | o353 | 0484 | 0615 | 0745 | 0876 | 1007 | 131 |
| 332 333 | 521138 | 1269 | 1400 | 1530 | 1661 | 1792 | 1922 | 2053 | 2183 | 2314 | 131 |
| 333 334 | 2444 | 2575 | 2705 | 2835 | 2966 | 3096 | 3226 | 3356 | 3486 | $3{ }^{\text {abir }}$ | 130 |
| 334 | 3746 | 3876 | 4006 | 4136 | 4266 | 4396 | 4526 | 4656 | 4785 | 4915 | 130 |
| 335 |  |  | 5304 | 5434 | 5563 | 5693 | 5822 | 5951 | 6081 | 6210 | 129 |
| 336 | 6339 | 6469 | 6598 | 6727 | 6856 | 6985 | 7114 | 7243 | 7372 | 7501 | 129 |
| 337 338 | 7630 .8 | 7759 | 7888 | 8016 | 8145 | 8274 | 8402 | 8531 | 8660 | 8788 | 129 |
| 338 339 | $* 8917$ 530200 | 9045 0328 | 9174 0.56 | 9302 0584 | 9430 0712 | 9550 | 9687 0968 | 9815 1096 | 9943 1223 | -072 | 128 <br> 128 |
|  |  |  |  |  |  |  |  |  |  |  |  |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |


| 6 |  | LOGARITHMS |  |  |  | NUM | BERS |  |  | Table I. |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |
| 340 | 531479 | 1607 | 1734 | 1862 | 1990 | 2117 | 2245 | 2372 | 2500 | 2627 | 128 |
| 341 | 2754 | 2882 | 3009 | 3136 | 3264 | 3391 | 3018 | 3645 | 3772 | 3899 | 127 |
| 342 | 4026 | 4153 | 4280 | 4407 | 4534 | 4661 | 4787 | 4914 | 5041 | 5167 | 127 |
| 343 | 5294 | 5421 | 5547 | 5674 | 5800 | 5927 | 6053 | 6180 | 6306 | 6432 | 126 |
| 344 | 6558 | 6685 | 6811 | 5937 | 7063 | 7189 | 7315 | 7441 | 7567 | 7693 | 126 |
| 345 | 7819 | 7945 | 8071 | 8197 | 8322 | 8448 | 8574 | 8699 | 8825 | 8951 | 126 |
| 346 | - 9076 | 9202 | 9327 | 9452 | 9578 | 9703 | 9829 | 9954 | -079 | 0204 | 125 |
| 347 | 540329 | 0455 | 0580 | 0705 | 0830 | 0955 | 1080 | 1205 | 1330 | 1454 | 125 |
| 348 | 1579 | 1704 | 1829 | 1953 | 2078 | 2203 | 2327 | 2452 | 2576 | 2701 | 125 |
| 349 | 2825 | 29 º | 3074 | 3199 | 3323 | 3447 | 3571 | 3696 | 3820 | 3944 | 124 |
| 350 | 4068 | 4192 | 4316 | 4440 | 4564 | 4688 | 4812 | 4936 | 5060 | 5183 | 124 |
| 351 | 5307 | 5431 | 5555 | 5678 | 5802 | 5925 | 6049 | 6172 | 6296 | 6419 | 124 |
| 352 | 6543 | 6666 | 6789 | 6913 | 7036 | 715 | 7282 | 7405 | 7529 | 7652 | 123 |
| 353 | 7775 | 7898 | 8021 | 8144 | 8267 | 8389 | 8512 | 8635 | 8758 | 8881 | 123 |
| 354 | * 9003 | 9126 | 9249 | 9371 | 9494 | 9616 | $97^{3} 9$ | 9861 | 9984 | -106 | 123 |
| 355 | 550228 | 0351 | 0473 | 0595 | 0717 | 0840 | 0,62 | 1084 | 1206 | 1328 | 122 |
| 356 | 1450 | 1572 | 1694 | 1816 | 1938 | 2060 | 2181 | 2303 | 2425 | 2547 | 122 |
| 357 35 | 2668 | 2790 | 2911 | 3033 | 3155 | 3276 | 3398 | 3519 | 3640 | 3762 | 121 |
| 358 | 3883 | 4004 | 4126 | 4247 | 4368 | 4489 | 4610 | 4731 | 4852 | 4973 | 121 |
| 359 | 5094 | 5215 | 5336 | 5457 | 5578 | 5699 | 5820 | 5940 | 6061 | 6182 | 121 |
| 360 | 6303 | 6423 | 6544 | 6664 | 6785 | 6905 | 7026 | 7146 | 7267 | 7387 | 120 |
| 361 | 7507 | 7627 | 7748 | 7868 | 7988 | 8108 | 8228 | 8349 | 8469 | 8589 | 120 |
| 362 | 8709 | 8829 | 8948 | 9068 | 9188 | 9308 | 9428 | 9548 | 9667 | 9787 | 120 |
| 363 | * 9907 | -026 | 0146 | 0265 | o385 | 0504 | 0624 | 0743 | 0863 | 0982 | 119 |
| 364 | 561101 | 1221 | 1340 | 1459 | 1578 | 1698 | 1817 | 1936 | 2055 | 2174 | 119 |
| 365 | 2293 | 2412 | 2531 | 2650 | 2769 | 2887 | 3006 | 3125 | 3244 | 3362 | 9 |
| 366 | 3481 | 3600 | 3718 | 3837 | 3955 | 4074 | 4192 | 4311 | 4429 | 4548 | 119 |
| 367 | 4666 | 4784 | 4903 | 5021 | 5139 | 5257 | 5376 | 5494 | 5612 | 5730 | 118 |
| 368 | 5848 | 5966 | 6084 | 6202 | 6320 | 6437 | 6555 | 6673 | 6791 | 6909 | 118 |
| 369 | 7026 | 7144 | 7262 | 7379 | 7497 | 7614 | $77^{32}$ | 7849 | 7967 | 8084 | 118 |
| 370 | 8202 | 8319 | 8436 | 8554 | 8671 | 8788 | 8905 | 9023 | 9140 | 9257 | 117 |
| 371 | * 9374 | 9491 | 9608 | 9725 | 9842 | 9959 | -076 | -193 | 0309 | 0426 | 117 |
| 372 3 | 570543 | 0660 | 0776 | 0893 | 1010 | 1126 | 1243 | 1359 | 1476 | 1592 | 117 |
| 373 | 1709 | 1825 | 1942 | 2058 | 2174 | 2291 | 2407 | 2523 | 2639 | 2755 | 116 |
| 374 | 2872 | 2988 | 3104 | 3220 | 3336 | 3452 | 3568 | 3684 | 380 | 3915 | 116 |
| 375 | 4031 | 4147 | 4263 | 4379 | 4494 | 4610 | 4726 | 4841 | 4957 | 5072 | 116 |
| 376 | 5188 | 5303 | 5419 | 5534 | 5650 | 5765 | 5880 | 5996 | 6111 | 6226 | 115 |
| 377 | 6341 | 6457 | 6572 | 6687 | 6802 | 6917 | 7032 | 7147 | 7262 | 7377 | 115 |
| 378 | 7492 | 7607 | 7722 | 7836 | 7951 | 8066 | 8181 | 8295 | 8410 | 8525 | 115 |
| 379 | 8639 | 8754 | 8868 | 8983 | 9097 | 9212 | 9326 | 9441 | 9555 | 9669 | 114 |
| 380 | * 9784 | 9898 | +012 | 0126 | 0241 | 0355 | 0469 | 0583 | 0697 | 0811 | 114 |
| 381 | 580925 | 1039 | 1153 | 1267 | 1381 | 1405 | 1608 | 1722 | 1836 | 1950 | 114 |
| 382 | 2063 | 2177 | 2291 | 2404 | 2518 | 2631 | 2745 | 2858 | 2972 | 3085 | 114 |
| 383 | 3199 | 3312 | 3426 | 3539 | 3652 | 3765 | 3879 | 3992 | 4105 | 4218 | 113 |
| 384 | 4331 | 4444 | 4557 | 4670 | 4783 | 4896 | 5009 | 5122 | 5235 | 5348 | 113 |
| 385 | 5461 | 5574 | 5686 |  | 5912 | 6024 | 6137 | 6250 | 6362 | 6475 | 113 |
| 386 | 6587 | 6700 | 6812 | 6925 | 7037 | 7149 | 7262 | 7374 | 7486 | 7509 | 112 |
| 387 | 771 | 7823 | 7935 | 8047 | 8160 | 8272 | 8384 | 8496 | 8608 | 8720 | 112 |
| 388 | 8832 | S944 | 9056 | 9167 | 9279 | 9391 | 9503 | 9615 | 9726 | 9838 | 112 |
| 389 | * 9950 | -061 | 0173 | 0284 | -396 | 0507 | 0619 | 0730 | 0842 | 0,53 | 112 |
| 390 | 591065 | 1176 | 1287 | 1399 | 1510 | 1621 | 1732 | 1843 | 1955 | 2066 | 111 |
| 391 | 2177 | 2288 | 2399 | 2510 | 2621 | 2732 | 2843 | 2954 | 3064 | 3175 | 111 |
| 392 | 3286 | 3397 | 3508 | 3618 | 3729 | 3840 | 3,50 | 4061 | 4171 | 4282 | 111 |
| 393 | 4393 | 4503 | 4614 | 4724 | 4834 | 4945 | 5055 | 5165 | 5276 | 5386 | 110 |
| 394 | 5496 | 5606 | 5717 | 5827 | 5937 | 6047 | 6157 | 6267 | 6377 | 6487 | 110 |
| 395 | 6597 | 6707 | 6817 | 6927 | 7037 | 7146 | 7256 | 7366 | 7476 | 7586 | 110 |
| 396 | 7695 | 7805 | 7914 | 8024 | 8134 | 8243 | 8353 | 8462 | 8572 | 8681 | 110 |
| 397 | 8791 | 8900 | 9009 | 9119 | 9228 | 9337 | 9446 | 9556 | 9665 | 9774 | 109 |
| 398 3 | $* 9883$ 60.0973 | 9992 | -101 | 0210 | ${ }^{0} 1219$ | 0428 | 0537 | 0646 | 0755 | 0864 | 109 |
| 399 | 600973 | 1082 | 1191 | 1299 | 1408 | 1517 | 1625 | 1734 | 1843 | 1951 | 19. |
| N | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |


| Table I. |  |  | LOGARITHMS OF NUMBERS. |  |  |  |  |  |  |  | 7 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |
| 400 | 602060 | 2169 | 2277 | 2386 | 2494 | 2603 | 2711 | 2819 | 2928 | 3036 | 108 |
| 401 | 3144 | 3253 | 3361 | 3469 | 3577 | 3686 | 3794 | 3 O 2 | 4010 | 4118 | 108 |
| 402 | 4226 | 4334 | 4442 | 4550 | 4658 | 4766 | 4874 | $49^{82}$ | 5089 | 5197 | 108 |
| 403 | 5305 | 5413 | 5521 | 5628 | 5736 | 5844 | 5951 | 6059 | 6166 | 6274 | 108 |
| 404 | 6381 | 6489 | 6596 | 6704 | 6811 | 6919 | 7026 | 7133 | 7241 | 7348 | 107 |
| 405 | 7455 | 7562 | 7669 | 7777 | 7884 | 7991 | 8098 | 8205 | 8312 | 8419 | 107 |
| 406 | 8525 | 9633 | 8740 | 8847 | 8954 | 9061 | 9167 | 9274 | 9381 | 9488 | 107 |
| 407 | * 9594 | 9701 | 9808 | 9914 | $\rightarrow 021$ | O128 | 0234 | . 0341 | 0447 | 0554 | 107 |
| 408 | 610660 | 0767 | ob73 | 0979 | 1086 | 1192 | 1298 | 1405 | 1511 | 1617 | 106 |
| 409 | 1723 | 1829 | 1936 | 2042 | 2148 | 2254 | 2360 | 2466 | 2572 | 2678 | 106 |
| 410 | 2784 | 2890 | 2996 | 3102 | 3207 | 3313 | 3419 | 3525 | 3630 | 3736 | 106 |
| 411 | 3842 | 3947 | 4053 | 4159 | 4264 | 4370 | 4475 | 4581 | 4686 | 4792 | 106 |
| 412 | 4897 | 5003 | 5108 | 5213 | 5319 | 5424 | 5529 | 5634 | 5740 | 5845 | 105 |
| 4 I 3 | 5,50 | 6055 | 6160 | 6265 | 6370 | 6476 | 6581 | 6686 | 6790 | 6895 | 105 |
| 414 | 7000 | 7105 | 7210 | 7315 | 7420 | 7525 | 7629 | 7734 | 7839 | 7943 | 105 |
| 415 | 8048 | 8153 | 8257 | 8362 | 8466 | 8571 | 8676 | 8780 | 8884 | 8989 | 105 |
| 416 | * 9093 | 9198 | 9302 | 9406 | 9511 | 9615 | 9719 | 9824 | 9928 | -032 | 104 |
| 417 | 620136 | 0240 | o344 | 0448 | 0552 | 0656 | 0760 | 0864 | 0968 | 1072 | 104 |
| 418 | 1176 | 1280 | 1384 | 1488 | 1592 | 1695 | 1799 | 1903 | 2007 | 2110 | 104 |
| 419 | 2214 | 2318 | 2421 | 2525 | 2628 | 2732 | 2835 | 2939 | 3042 | 3146 | 104 |
| 420 | 3249 | 3353 | 3456 | 3559 | 3663 | 3766 | 3869 | 3973 | 4076 | 4179 | 103 |
| 421 | 4282 | 4385 | 4488 | 4591 | 4695 | 4798 | 4901 | 5004 | 5107 | 5210 | 103 |
| 422 | 5312 | 5415 | 5518 | 5621 | 5724 | 5827 | 5929 | 6032 | 6135 | 6238 | 103 |
| 423 | 6340 | 6443 | 6546 | 6648 | 6751 | 6853 | 6956 | 7058 | 7161 | 7263 | 103 |
| 424 | 7366 | 7468 | 7571 | 7673 | 7775 | 7878 | 7980 | 8082 | 8185 | 8287 | 102 |
| 425 | 8389 | 8491 | 8593 | 8695 | 8797 | 8900 | 9002 | 9104 | 9206 | 9308 | 102 |
| 426 | * 9410 | 9512 | 9613 | 9715 | 9817 | 9919 | +021 | 0123 | C224 | O326 | 102 |
| 427 | 630428 | -530 | 0631 | 0733 | 0835 | 0936 | 1038 | 1139 | 1241 | 1342 | 102 |
| 428 | 1444 | 1545 | 1647 | 1748 | 1849 | 1951 | 2052 | 2153 | 2255 | 2356 | 101 |
| 429 | 2457 | 2559 | 2660 | 2761 | 2862 | 2963 | 3064 | 3165 | 3266 | 3367 | 101 |
| 430 | 3468 | 3569 | 3670 | 3771 | 3872 | 3973 | 4074 | 4175 | 4276 | 4376 | 100 |
| 431 | 4477 | 4578 | 4679 | 4779 | 4880 | 4981 | 5081 | 5182 | 5283 | 5383 | 100 |
| 432 | 5484 | 5584 | 5685 | 5785 | 5886 | 5986 | 6087 | 6187 | 6287 | 6388 | 100 |
| 433 | 6488 | 6588 | 6688 | 6789 | 6889 | 6989 | 7089 | 7189 | 7290 | 7390 | 100 |
| 434 | 7490 | 7590 | 7690 | 7790 | 7890 | 7990 | 8090 | 8190 | 8290 | 8389 | 99 |
| 435 | 8489 | 8589 | 8689 | 8789 | 8888 | 8988 | 9088 | 9188 | 9287 | 9387 | 99 |
| 436 | * 9485 | 9586 | 9686 | 9785 | 9885 | 9984 | ¢084 | O183 | 0283 | o382 | 99 |
| 437 | 640481 | 0581 | 0680 | 0779 | 0879 | 0978 | 1077 | 1177 | 1276 | 1375 | 99 |
| 438 | 1474 | 1573 | 1672 | 1771 | 1871 | 1970 | 2069 | 2168 | 2267 | 2366 | 99 |
| 439 | 2465 | 2563 | 2662 | 2761 | 2860 | 2959 | 3058 | 3156 | 3255 | 3354 | 99 |
| 440 | 3453 | 3551 | 3650 | 3749 | 3847 | 3946 |  | 4143 | 4242 | 4340 | 98 |
| 441 | 4439 | 4537 | 4636 | 4734 | 4832 | 4931 | 5029 | 5127 | 5226 | 5324 | 98 |
| 442 | 5422 | 5521 | 5619 | 5717 | 5815 | 5 g 13 | 6011 | 6110 | 6208 | 6306 | 98 |
| 443 | 6404 | 6502 | 6600 | 6698 | 6796 | 6894 | 6992 | 7089 | 7187 | 7285 | 98 |
| 444 | 7383 | 7481 | 7579 | 7676 | 7774 | 7872 | 7969 | 8067 | 8165 | 8262 | 98 |
| 445 | 8360 | 8458 | 8555 | 8653 | 8750 | 8849 | 8945 | 9043 | 9140 | 9237 | 97 |
| 446 | * ${ }_{6} 9335$ | 9432 | 9530 | 9627 | 9724 | 9821 | 9919 | -016 | 0113 | 0210 | 97 |
| 447 | 65 o308 | 0405 | 0502 | -599 | 0696 | 0793 | 0890 | 0987 | 1084 | 1181 | 97 |
| 448 | 1278 | 1375 | 1472 | 1569 | 1666 | 1762 | 1859 | 1956 | 2053 | 2150 | 97 |
| 449 | 2246 | 2343 | 2440 | 2536 | 2633 | 2730 | 2826 | 2923 | 3019 | 3116 | 97 |
| 450 | 3213 | 3309 | 3405 |  |  |  | 3791 |  | 3984 | 4080 | 96 |
| 451 | 4177 | 4273 | 4369 | 4465 | 4562 | 4658 | 4754 | 4850 | 4946 | 5042 | \%6 |
| 452 | 5138 | 5235 | 5331 | 5427 | 5523 | 5619 | 5715 | 58ı0 | 5006 | 6002 | 96 |
| 453 | 6098 | 6194 | 6290 | 6386 | 6482 | 6577 | 6673 | 6769 | 6864 | 6960 | 96 |
| 454 | 7056 | 7152 | 7247 | 7343 | 7438 | 7534 | 7629 | 7725 | 7820 | 7916 | 96 |
| 455 | 8011 | 8107 | 8202 | 8298 | 8393 | 8488 | 8584 | 8679 | 8774 | 8870 | 95 |
| 456 | 8965 $* 9016$ | 9060 | 9155 | 9250 | 9346 | 9441 | 9536 | 9631 | 9726 | 9821 | 95 |
|  |  | -011 | 0106 | O201 | 0296 | 0391 | 0486 | 0581 | 0676 | 0771 | 95 |
| 458 459 | $660865$ | 0960 | 1055 | 1150 | 1245 | 1339 | 1434 | 1529 | 1623 | 1718 | 95 |
| 459 | 1813 | 1907 | 2002 | 2096 | 2191 | 2286 | 2380 | 2475 | 2569 | 2663 | 95 |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |


| 8 |  | LOGARITHMS OF NUMBERS. |  |  |  |  |  |  |  | Table L |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |
| 460 | 662758 | 2852 | 2947 | 3041 | 3135 | 3230 | 3324 | 3418 | 3512 | 3607 | 94 |
| 461 | 3701 | 3795 | 3889 | 3983 | 4078 | 4172 | 4266 | 4360 | 4454 | 4548 | 94 |
| 402 | 46.42 | 4736 | 4830 | 4924 | 5018 | 5112 | 5206 | 5299 | 5393 | 5487 | 94 |
| 463 | 5581 | 5675 | 5762 | 5862 | 5.56 | 6050 | 6143 | 6237 | 6331 | 6424 | 94 |
| 464 | 6518 | 6612 | 6705 | 6799 | 6892 | 6986 | 7079 | 7173 | 7266 | 7360 | 94 |
| 465 | 7453 | 7546 | 7640 | 7733 | 7826 | 7920 | 8013 | 8106 | 8199 | 8293 | 93 |
| 466 | 8386 | 8479 | 8572 | 8665 | 8759 | 8852 | 8945 | 9038 | 9131 | 9224 | 93 |
| 467 | *9317 | 9410 | 9503 | 9596 | 9689 | 9782 | $9^{8} 75$ | 9967 | -060 | 0153 | 93 |
| 468 | 670246 | -339 | 0431 | 0524 | 0617 | 0710 | 0802 | 0895 | 0988 | 1080 | 93 |
| 469 | 1173 | 1265 | 1358 | 1451 | 1543 | 1636 | 1728 | 1821 | 1913 | 2005 | $9^{3}$ |
| 470 | 2098 | 2190 | 2:83 | 2375 | 2467 | 2560 | 2652 | 2744 | 2836 | 2729 | 92 |
| 471 | 3021 | 3113 | 3205 | 3297 | 3390 | 3482 | 3574 | 3666 | 3758 | 3850 | 92 |
| 472 | 3942 | 4034 | 4126 | 4218 | 4310 | 4402 | 4494 | 4586 | 4677 | 4769 | 92 |
| 473 | 4861 | 4953 | 5045 | 5137 | 5228 | 5320 | 5412 | 5503 | 5595 | 5687 | 92 |
| 474 | 5778 | 5870 | 5962 | 6053 | 6145 | 6236 | 6328 | 6419 | 6511 | 6602 | 92 |
| $47^{5}$ | 6694 | 6785 | 6876 | 6968 | 7059 | 7151 | 7242 | 7333 | 7424 | 7516 | 91 |
| 476 | 7607 | 7698 | 7789 | 7881 | 7972 | 8063 | 8154 | 8245 | 8336 | 8427 | 91 |
| 477 | 8518 | 8009 | 8700 | 8791 | 8882 | 8973 | 9064 | 9155 | 9246 | 9337 | 91 |
| 478 | * 9428 | 9519 | 9610 | 9700 | 9791 | 9882 | 9973 | +053 | 0154 | 0245 | 91 |
| 479 | 68 o336 | 0426 | ${ }^{0} 517$ | 0607 | 0698 | 0789 | 0879 | 0970 | 1060 | 1151 | 91 |
| 480 | 1241 | 1332 | 1422 | 1513 | 1603 | 1693 | 1784 | 1874 | 1964 | 2055 | 90 |
| 481 | 2145 | 2235 | 2326 | 2416 | 2506 | 2396 | 2686 | 2777 | 2867 | 2957 | 90 |
| 482 | 3047 | 3137 | 3227 | 3317 | 3407 | 3497 | 3587 | 3677 | 3767 | 3857 | 90 |
| 483 | 3947 | 4037 | 4127 | 4217 | 4307 | 4396 | 4486 | 4576 | 4666 | 4756 | 90 |
| 484 | 4845 | 4935 | 5025 | 5114 | 5204 | 5294 | 5383 | 5473 | 5563 | 5652 | 90 |
| 485 | 5742 | 5831 | 5921 | 6010 | 6100 | 6189 | 6279 | 6368 | 6458 | 6547 | 89 |
| 486 | 6636 | 5726 | 6815 | 6904 | 6094 | 7083 | 7172 | 7261 | 7351 | 7440 | 89 |
| 487 | 7529 | 7618 | 7707 | 7796 | 7886 | 7975 | 8064 | 8153 | 8242 | 8331 | 89 |
| 488 | 8420 | 8509 | 8508 | 8687 | 8776 | 8865 | 8 g 53 | 9042 | 9131 | 9220 | 89 |
| 489 | * 9309 | 9398 | 9486 | 9575 | 9664 | 9753 | 9841 | 9930 | +019 | 0107 | 89 |
| 490 | 690196 | 0285 | 0373 | 0462 | 0550 | 0639 | 0728 | 0816 | 0905 | 0993 | 89 |
| 491 | 1081 | 1170 | 1258 | 1347 | 1435 | 1524 | 1612 | 1700 | 1789 | 1877 | 88 |
| 492 | 1965 | 2053 | 2142 | 2230 | 2318 | 2406 | 2494 | 2583 | 2671 | 2759 | 88 |
| 493 | 2847 | 2935 | 3023 | 3111 | 3199 | 3287 | 3375 | 3463 | 3501 | 3639 | 88 |
| 494 | 3727 | 3815 | 3903 | 3991 | 4078 | 4166 | 4254 | 4342 | 4430 | 4517 | 88 |
| 495 | 4605 | 4693 | 4781 | 4868 | 4956 | 5044 | 5131 | 5219 | 5307 | 5394 | 88 |
| 496 | 5482 | 5569 | 5657 | 5744 | 5832 | 5919 | 6007 | 6094 | 6182 | 6269 | 87 |
| 497 | 6356 | 6444 | 6531 | 6618 | 6706 | 6793 | 6880 | 6968 | 7055 | 7142 | 87 |
| 498 | 7229 | 7317 | 7404 | 7491 | 7578 | 7665 | 7752 | 7839 | 7926 | 8014 | 87 |
| 499 | 8101 | 81.88 | 8275 | 8362 | 8449 | 8535 | 8622 | 8709 | 8796 | 8883 | 87 |
| 500 | 8970 | 9057 | 9144 | 923I | 9317 | 9404 | 9491 | 9578 | 9664 | 9751 | 87 |
| 501 | - 9838 | 9924 | -011 | 0098 | 0184 | 0271 | 0358 | 0444 | 0531 | 0617 | 87 |
| 502 | 700704 | 0790 | 0877 | -963 | 1050 | 1136 | 1222 | 1309 | 1395 | 1482 | 86 |
| 503 | 1568 | 1654 | 1741 | 1827 | 1913 | 1999 | 2086 | 2172 | 2258 | 2344 | 86 |
| 504 | 2431 | 2517 | 2603 | 2689 | 2775 | 2861 | 2947 | 3033 | 3119 | 3205 | 86 |
| 505 | 3291 |  | 3463 |  | 3635 | 3721 |  | 3805 |  | 4065 |  |
| 506 | 4151 | 4236 | 4322 | 4408 | 4494 | 4579 | 4665 | 4751 | 4837 | 4922 | 86 |
| 507 | 5008 | 5094 | 5179 | 5265 | 5350 | 5436 | 5522 | 5607 | 5693 | 5778 | 86 |
| 508 | 5864 | 5949 | 6035 | 6120 | 6206 | 6291 | 6376 | 6462 | 6547 | 6632 | 85 |
| 509 | 6718 | 6803 | 6888 | 6974 | 70 9 9 | 7144 | 7229 | 7315 | 7400 | 7485 | 85 |
| 510 | 7570 | 7655 | 7740 | 7826 | 7911 | 7996 | 8081 | 8166 | 8251 | 8336 | 85 |
| 511 | 8421 | 8506 | 8591 | 8676 | 8761 | 8846 | 893i | 9015 | 9100 | 9185 | 85 |
| 512 | * 9270 | 9355 | 9440 | 9524 | 9609 | 9694 | 9779 | 9863 | 9948 | -033 | 85 |
| 513 | 710117 | 0202 | 0287 | 0371 | 0456 | ${ }^{0} 510$ | 0625 | 0710 | 0794 | 0879 | 85 |
| 514 | 0963 | 1048 | 1132 | 1217 | 1301 | 1385 | 1470 | 1554 | 1639 | 1723 | 84 |
| 515 | 1807 | 1892 | 1976 | 2060 | 2144 | 2229 | 2313 | 2397 | 2481 | 2506 | 84 |
| 516 | 2650 | 2734 | 2818 | 2902 | 2986 | 3070 | 3154 | 3238 | 3323 | 3407 | 84 |
| 517 | 3491 | 3575 | 3650 | 3742 | 3826 | 3910 | 3994 | 4078 | 4162 | 4246 | 84 |
| 518 | 4330 | 4414 | 4497 | 4581 | 4665 | 4749 | 4833 | 4916 | 5000 | 5084 | 84 |
| 519 | 5167 | 525 I | 5335 | 5418 | 5502 | 5586 | 5669 | 5753 | 5836 | 5920 | 84 |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |


| Table I. |  |  | LOGARITHMS OF NUMBERS. |  |  |  |  |  |  |  | 0 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |
| 520 | 716003 | 6087 | 6170 | 6254 | 6337 | 6421 | 6504 | 6588 | 6671 | 6754 | 83 |
| 521 | 6838 | 6921 | 7004 | 7088 | 7171 | 7254 | 7338 | 7421 | 7504 | 7587 | 83 |
| 522 | 7671 | 7754 | 7837 | 7920 | 8003 | 8086 | 8169 | 8253 | 8336 | 8419 | 83 |
| 523 | 8502 | 8585 | 8668 | 8751 | 8834 | 8917 | 9000 | 9083 | 9165 | 9248 | 83 |
| 524 | * 9331 | 9414 | 9497 | 9580 | 9663 | 9745 | 9828 | 9911 | ¢994 | +077 | 83 |
| 525 | 720159 | 0242 | -325 | 0407 | 0490 | -573 | 0655 | 0738 | 0821 | s903 | 83 |
| 526 | 0986 | 1068 | 1151 | 1233 | 1316 | 1398 | 1481 | 1563 | 1646 | 1728 | 82 |
| 527 | 1811 | 1893 | 1975 | 2058 | 2140 | 2222 | 2305 | 2387 | 2469 | 2552 | 82 |
| 528 | 2634 | 2716 | 2798 | 2881 | 2963 | 3045 | 3127 | 3209 | 3291 | 3374 | 82 |
| 529 | 3456 | 3538 | 3620 | 3702 | 3784 | 3866 | 3948 | 4030 | 4112 | 4194 | 82 |
| 530 | 4276 | 4358 | 4440 | 4522 | 4604 | 4685 | 4767 | 4849 | 4931 | 5013 | 82 |
| 531 | 5095 | 5176 | 5258 | 5340 | 5422 | 5503 | 5585 | 5667 | 5748 | 5830 | 82 |
| 532 | 5912 | 5993 | 6075 | 6156 | 6238 | 6320 | 6401 | 6483 | 6564 | 6646 | 82 |
| 533 | 6727 | 6809 | 6890 | 6972 | 7053 | 7134 | 7216 | 7297 | 7379 | 7460 | 81 |
| 534 | 7541 | 7623 | 7704 | 7785 | 7866 | 7948 | 8029 | 8110 | 8191 | 8273 | 81 |
| 535 | 8354 | 8435 | 8516 | 8597 | 8678 | 8759 | 8841 | 8922 | 9003 | 9084 | 81 |
| 536 | 9165 | 9246 | 9327 | 9408 | 9489 | 9570 | 9651 | 9732 | 9813 | 9893 | 81 |
| 537 | * 9974 | +055 | 0136 | 0217 | 0298 | 0378 | 0459 | 0540 | 0621 | 0702 | 81 |
| 538 | 730782 | 0863 | 0944 | 1024 | 1105 | 1186 | 1266 | 1347 | 1428 | 1508 | 81 |
| 539 | 1589 | 1669 | 1750 | 1830 | 1911 | 1991 | 2072 | 2152 | 2233 | 2313 | 81 |
| 540 | 2394 | 2474 | 2555 | 2635 | 2715 | 2796 | 2876 | 2956 | 3037 | 3117 | 80 |
| 541 | 3197 | 3278 | 3358 | 3438 | 3518 | 3598 | 3679 | 3759 | 3839 | 3919 | 80 |
| 542 | 3999 | 4079 | 4160 | 4240 | 4320 | 4400 | 4480 | 4560 | 4640 | $47^{20}$ | 80 |
| 543 | 4800 | 4880 | 4960 | 5040 | 5120 | 5200 | 5279 | 5359 | 5439 | 5519 | 80 |
| 544 | 5599 | 5679 | 5759 | 5838 | 5918 | 5998 | 6078 | 6157 | 6237 | 6317 | 80 |
| 545 | 6397 | 6476 | 6556 | 6635 | 6715 | 6795 | 6874 | 6954 | 7034 | 7113 | 80 |
| 546 | 7193 | 7272 | 7352 | 7431 | 7511 | 7500 | 7670 | 7749 | 7829 | 7908 | 79 |
| 547 | 7987 | 8067 | 8146 | 8225 | 8305 | 8384 | 8463 | 8543 | 8622 | 8701 | 79 |
| 548 | 8781 | 8860 | $89^{39}$ | 9018 | 9097 | 9177 | 9256 | 93.35 | 9414 | 9493 | 79 |
| 549 | * 9572 | 9651 | 9731 | 9810 | 9889 | 9968 | +047 | 0126 | 0205 | 0284 | 79 |
| 550 | 740363 | 0442 | 0521 | 0600 | 0678 | 0757 | 0836 | 0915 | 0994 | 1073 | 79 |
| 551 | 1152 | 1230 | 1309 | 1388 | 1467 | 1546 | 1624 | 1703 | 1782 | 1860 | 79 |
| 552 | 1939 | 2018 | 2096 | 2175 | 2254 | 2332 | 2411 | 2489 | 2568 | 2646 | 79 |
| 553 | 2725 | 2804 | 2882 | 2961 | 3039 | 3118 | 3196 | 3275 | 3353 | 3431 | 78 |
| 554 | 3510 | 3588 | 3667 | 3745 | 3823 | 3902 | 3980 | 4058 | 4136 | 4215 | 78 |
| 555 | 4293 | 4371 | 4449 | 4528 | 4606 | 4684 | 4762 | 4840 | 4919 | 4997 | 78 |
| 556 | 5075 | 5153 | 5231 | 5309 | 5387 | 5465 | 5543 | 5621 | 5699 | 5777 | 78 |
| 557 | 5855 | 5933 | 6011 | 6089 | 6167 | 6245 | 6323 | 6401 | 6479 | 6556 | $7^{8}$ |
| 558 | 6634 | 6712 | 6790 | 6868 | 6945 | 7023 | 7101 | 7179 | 7256 | 7334 | 78 |
| 559 | 7412 | 7489 | 7567 | 7645 | 7722 | 7800 | 7878 | 7955 | 8033 | 8110 | 78 |
| 560 | 8188 | 8266 | 8343 | 8421 | 8498 | 8576 | 8653 | 8731 | 8808 | 8885 | 77 |
| 561 | 8963 | 9040 | 9118 | 9195 | 9272 | 9350 | 9427 | 9504 | 9582 | 9659 | 77 |
| 562 | * 9736 | 9814 | 9891 | 9968 | +045 | 0123 | 0200 | 0277 | 0354 | 0431 | 77 |
| 563 | 750508 | 0586 | 0663 | 0740 | 0817 | 0894 | 0971 | 1048 | 1125 | 1202 | 77 |
| 564 | 1279 | 1356 | 1433 | 1510 | 1587 | 1664 | 1741 | 1818 | 1895 | 1972 | 77 |
| 565 | 2048 | 2125 | 2202 | 2279 | 2356 | 2433 | 2509 | 2586 | 2663 | 2740 | 77 |
| 566 | 2816 | 2893 | 2970 | 3047 | 3123 | 3200 | 3277 | 3353 | 3430 | 3506 | 77 |
| 567 | 3583 | 3660 | 3736 | 3813 | 3889 | 3966 | 4042 | 4119 | 4195 | 4272 | 77 |
| 568 | 4348 | 4425 | 4501 | 4578 | 4654 | 4730 | 4807 | 4883 | 4960 | 5036 | 76 |
| 569 | 5112 | 5189 | 5265 | 5341 | 5417 | 5494 | 5570 | 5646 | 5722 | 5799 | 75 |
| 570 | 5875 | 5951 |  | 6ı03 | 6180 | 6256 | 6332 | 6408 | 6484 | 6560 | 76 |
| 571 | 6636 | 6712 | 6788 | 6864 | 6940 | 7016 | 7092 | 7168 | 7244 | 7320 | 76 |
| 572 | 7396 | 7472 | 7548 | 7624 | 7700 | 7775 | 7851 | 7927 | 8003 | 8079 | 76 |
| 573 | 8155 | 8230 | 8306 | 8382 | 8458 | 8533 | 8609 | 8685 | 8761 | 8836 | 76 |
| 574 | 8912 | 8988 | 9063 | 9139 | 9214 | 9290 | 9366 | 9441 | 9517 | 9592 | 76 |
| 575 | * 9668 | 9743 | 9819 | 9894 | 9970 | -045 | 0121 | 0196 | 0272 | 0347 | 75 |
| 576 | 760422 | 0498 | ${ }^{0} 573$ | 0649 | 0724 | 0799 | 0875 | 0,50 | 1025 | 1101 | 75 |
| 577 | 1176 | 1251 | 1326 | 1402 | 1477 | 1552 | 1627 | 1702 | 1778 | 1853 | 75 |
| 578 579 | 1928 2679 | 2003 2754 | 2078 2829 | 2153 | 2228 | 2303 3053 | 2378 3128 | 2453 3203 | 2529 3278 | 2604 3353 | 75 |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |

LOGARITHMS OF NUMBERS.
Table I.

| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 580 | -6 3428 | 3503 | 3578 | 3653 | 3727 | 3802 | 3877 | 3052 | 4027 | 4101 | 75 |
| 581 | 4176 | 4251 | 4326 | 4400 | 4475 | 4550 | 4624 | 4699 | 4774 | 4848 | 95 |
| 582 | 4923 | 4998 | 5072 | 5147 | 5221 | 5296 | 5370 | 5445 | 5520 | 5594 | 75 |
| 533 | 5669 | 5743 | 5818 | 5892 | 5966 | ó041 | 6115 | 6190 | 6264 | 6338 | 74 |
| 584 | 6413 | 6487 | 6562 | 6636 | 6710 | 6785 | 6859 | 6933 | 7007 | 7082 | 74 |
| 585 | 7156 | 7230 | 7304 | 7379 | 7453 | 7527 | 7601 | 7675 | 7749 | 7823 | 74 |
| 586 | 7898 | 7972 | 8046 | 8120 | . 8194 | 8268 | 8342 | 8416 | 8490 | 8564 | 74 |
| 587 | 8638 | 8712 | 8786 | 8860 | 8934 | 9008 | 9082 | 9156 | 9230 | 9303 | 74 |
| 538 | * 9377 | 945 I | 9525 | 9599 | 9673 | 9746 | 9820 | 9894 | 9968 | -042 | 74 |
| 589 | 770115 | -189 | 0263 | -336 | 0410 | 0484 | 0557 | 0631 | 0705 | 0778 | 74 |
| 590 | 0852 | 0926 | 0999 | 1073 | 1146 | 1220 | 1293 | 1367 | 1440 | 1514 | 74 |
| 591 | 1587 | 1661 | 1734 | 1808 | 1881 | 1955 | 2028 | 2102 | 2175 | 2248 | 73 |
| 592 | 2322 | 2395 | 2468 | 25.2 | 2615 | 2688 | 2762 | 2835 | 2908 | 2981 | 73 |
| 593 | 3055 | 3128 | 3201 | 3274 | 3348 | 3421 | 3494 | 3567 | 3640 | 3713 | 73 |
| 594 | 3786 | 3860 | 3933 | 4006 | 4079 | 4152 | 4225 | 4298 | 4371 | 4444 | 73 |
| 595 | 4517 | 4590 | 4663 | 4736 | 4809 | 4882 | 4955 | 5028 | 5100 | 5173 | 73 |
| 596 | 5246 | 5319 | 5392 | 5465 | 5533 | 5610 | 5683 | 5756 | 5829 | 5002 | 73 |
| 597 | 5974 | 6047 | 6120 | 6193 | 6265 | 6338 | 6411 | 6483 | 6556 | 6629 | 73 |
| 598 | 6701 | 6774 | 6846 | 6919 | 6992 | 7064 | 7137 | 7209 | 7282 | 7354 | 73 |
| 599 | 7427 | 7499 | 7572 | 7644 | 7717 | 7789 | 7862 | 7934 | 8006 | 8079 | 72 |
| 600 | 8151 | 8224 | 8296. | 8368 | 8441 | 8513 | 8585 | 8658 | 8730 | 8802 | 72 |
| 601 | 8874 | 8947 | 9019 | 9091 | 9163 | 9236 | 9308 | 9380 | 9452 | 9524 | 72 |
| 602 | * 95956 | 9669 | 9741 | 9813 | 9885 | 9957 | -029 | 0101 | 0173 | 0245 | 72 |
| 603 | 780317 | o389 | 0.461 | 0533 | 0605 | 0677 | 0749 | 0821 | $089^{3}$ | 0965 | 72 |
| 604 | 1037 | 1109 | 1181 | 1253 | 1324 | 1396 | 1468 | 1540 | 1612 | 1684 | 72 |
| 60 | 1755 | 1827 | 1899 | 1971 | 2042 | 2114 | 2186 | 2258 | 2329 | 2401 | 72 |
| 606 | 2473 | 2544 | 2616 | 2688 | 2759 | 2831 | 2902 | 2974 | 3046 | 3117 | 13 |
| 607 | 3189 | 3260 | 3332 | 3403 | 3475 | 3546 | 3618 | 3689 | 3761 | 3832 | 71 |
| 608 | 3004 | 3975 | 4046 | 4118 | 4189 | 4261 | 4332 | 4403 | 4475 | 4546 | 71 |
| 609 | 4617 | 4689 | 4760 | 4831 | 4902 | 4974 | 5045 | 5116 | 5187 | 5259 | 71 |
| 610 | 5330 | 5401 | 5472 | 5543 | 5615 | 5686 | 5757 | 5828 | 5899 | 5970 | 71 |
| 611 | 6041 | 6112 | 6183 | 6254 | 6325 | 6396 | 6467 | 5538 | 6609 | 6680 | 71 |
| 612 | 6751 | 6822 | 6893 | 6964 | 7035 | 7106 | 7177 | 7248 | 7319 | 7390 | 71 |
| 6.3 | 7460 | 7531 | 7602 | 7673 | 7744 | 7815 | 7885 | 7956 | 8027 | 8098 | 71 |
| 614 | 8168 | 8239 | 8310 | 8381 | 8451 | 8522 | 8593 | 8663 | 8734 | 8854 | 71 |
| 615 | 8875 | 8946 | 9016 | 9087 | 915 | 9228 | 9299 | 9369 | 9440 | 9510 | 71 |
| 616 | * 9581 | 9651 | 9722 | 9792 | 9863 | 9933 | -004 | 0074 | 0144 | 0215 | 70 |
| 617 | 790285 | 0356 | 0426 | 0496 | 0567 | 0637 | 0707 | 0778 | 0848 | 0918 | 70 |
| 618 | 0983 | 1059 | 1129 | 1199 | 1269 | 1340 | 1410 | 1480 | 1550 | 1620 | 70 |
| 619 | 1691 | 1761 | 1831 | 1901 | 1971 | 2041 | 2111 | 2181 | 2252 | 2322 | 70 |
| 620 | 2392 | 2462 | 2532 | 2602 | 2672 | 2742 | 2812 | 2882 | 2952 | 3022 | 70 |
| 621 | 3092 | 3162 | 3231 | 3301 | 3371 | 3441 | 3511 | 3581 | 365 I | 3721 | 70 |
| 622 | 3790 | 3860 | 3930 | 4000 | 4070 | 4139 | 4209 | 4279 | 4349 | 4418 | 70 |
| 623 | 4488 | 4558 | 4627 | 4697 | 4767 | 4836 | 4906 | 4976 | 5045 | 5115 | 70 |
| 624 | 5185 | 5254 | 5324 | 5393 | 5463 | 5532 | 5602 | 5672 | 5741 | 5811 | 70 |
| 625 | 5880 | 5949 | 6019 | 6088 | 6158 | 6227 | 6297 | 6366 | 6436 | 6505 | 69 |
| 626 | 6574 | 6644 | 6713 | 6782 | 6852 | 6921 | 6990 | 7060 | 7129 | 7198 | 69 |
| 627 | 7268 | 7337 | 7406 | 7475 | 7545 | 7614 | 7683 | 7752 | 7821 | 7800 | 69 |
| 628 | 7960 | 8029 | 8098 | 8167 | 8236 | 8305 | 8374 | 8443 | 8513 | 8582 | 69 |
| 629 | 8651 | 8720 | 8789 | 8858 | 8927 | 8996 | 9065 | 9134 | 9203 | 9272 | 69 |
| 630 | 9341 | 9409 | 9478 | 9547 | 9616 | 9685 | 9754 | 9823 | $98{ }^{8} 2$ | 9961 | 69 |
| 631 | 800029 | 0098 | 0167 | 0236 | 0305 | 0373 | 0442 | -511 | 0580 | 0648 | 69 |
| 632 | 0717 | 0786 | 0854 | 0923 | 0992 | 1061 | 1129 | 1198 | 1266 | 1335 | 69 |
| 633 | 1404 | 1472 | 1541 | 1609 | 1678 | 1747 | 1815 | 1884 | 1952 | 2021 | 69 |
| 634 | 2089 | 2158 | 2226 | 2295 | 2363 | 2432 | 2500 | 2568 | 2637 | 2705 | 69 |
| 635 | 2774 | 2842 | 2910 | 2979 | 3047 | 3116 | 3184 | 3252 | 3321 | 3389 | 68 |
| 636 | 3457 | 3525 | 3594 | 3662 | 3730 | 3798 | 3867 | 3935 | 4003 | 4071 | 68 |
| 637 | 4139 | 4208 | 4276 | 4344 | 4412 | 4480 | 4548 | 4616 | 4685 | 4753 | 68 |
| 638 | 4821 | 4889 | 4957 | 5025 | 5093 | 5161 | 5229 | 5297 | 5355 | 5433 | 68 |
| 639 | 5501 | 5569 | 5637 | 5705 | 5773 | 5841 | 5908 | 5976 | 6044 | 6112 | 68 |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |


| Table I. |  | LOGARITHMS |  |  |  | NUM | IBERS |  |  |  | 11 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |
| 640 | 80 6180 | 6248 | 6316 | 638 | 6451 | 6519 | 6587 | 6655 | 6723 | 6790 | 68 |
| 641 | 6858 | 6926 | 6994 | 7061 | 7129 | 7197 | 7264 | 7332 | 7400 | 7467 | 68 |
| 642 | 7535 | 7603 | 7670 | 7738 | 7806 | 7873 | 7941 | 8008 | 8076 | 8143 | 68 |
| 643 | 8211 | 8279 | 8346 | 8414 | 8481 | 8549 | 8616 | 8684 | 8751 | 88.8 | 67 |
| 644 | 8846 | 8953 | 9021 | 9088 | 9156 | 9223 | 9290 | 9358 | 9425 | 9492 | 67 |
| 645 | * 9560 | 9627 | 9694 | 9762 | 9829 | 9896 | 9964 | -031 | 0098 | 0165 | 67 |
| 646 | 81 0233 | 0300 | 0367 | 0434 | 0501 | 0569 | 0636 | -703 | 0770 | 0837 | 67 |
| 647 | 0204 | 0971 | 1039 | 1106 | 1173 | 1240 | 1307 | 1374 | 1441 | 1508 | 67 |
| 648 | 1575 | 1642 | 1709 | 1776 | 1843 | 1910 | 1977 | 2044 | 2111 | 2178 | 67 |
| 649 | 2245 | 2312 | 2379 | 2445 | 2512 | 2579 | 2646 | 2713 | 2780 | 2847 | 67 |
| 650 | 2913 | 2980 | 3047 | 3114 | 3181 | 3247 | 3314 | 3381 | 3448 | 3514 | 67 |
| 65 I | 3581 | 3648 | 3714 | 3781 | 3848 | 3914 | 3981 | 4048 | 4114 | 4181 | 67 |
| 652 | 4248 | 4314 | 4381 | 4447 | 4514 | 4581 | 4647 | 4714 | 4780 | 4847 | 67 |
| 653 | 4913 | 4980 | 5046 | 5113 | 5179 | 5246 | 5312 | 5378 | 5445 | 5511 | 66 |
| 654 | 5578 | 5644 | 5711 | 5777 | 5843 | 5910 | 5976 | 6042 | 6109 | 6175 | 66 |
| . 655 | 6241 | 6308 | 6374 | 6440 | 6506 | 6573 | 6639 | 6705 | 6771 | 6838 | 66 |
| 656 | 6904 | 6970 | 7036 | 7102 | 7169 | 7235 | 7301 | 7367 | 7433 | 7499 | 66 |
| 657 | 7565 | 7631 | 7698 | 7764 | 7830 | 7896 | 7962 | 8028 | 8094 | 8160 | 66 |
| 658 | 8226 | 8292 | 8358 | 8424 | 8490 | 8556 | 8622 | 8688 | 8754 | 8820 | $66$ |
| 659 | 8885 | $89^{51}$ | 9017 | 9083 | 9149 | 9215 | 9281 | 9346 | 9412 | 9478 | 66 |
| 660 | * 9544 | 9610 | 9676 | 9741 | 9807 | 9873 | $99^{39}$ | -004 | 0070 | 0.36 | 66 |
| 661 | 820201 | 0267 | -333 | -309 | 0464 | 0530 | 0595 | 0661 | 0727 | 0792 | 66 |
| 662 | 0858 | 0924 | -989 | 1055 | 1120 | 1186 | 1251 | 1317 | 1382 | 1448 | 66 |
| 663 | 1514 | 1579 | 1645 | 1710 | 1775 | 1841 | 1906 | 1972 | 2037 | 2103 | 65 |
| 664 | 2168 | 2233 | 2299 | 2364 | 2430 | 2495 | 2560 | 2626 | 2691 | 2756 | 65 |
| 665 | 2822 | 2887 | 2952 | 3018 | 3083 | 3148 | 3213 | 3279 | 3344 | 3409 | 65 |
| 666 | 3474 | 3539 | 3605 | 3670 | 3735 | 3800 | 3865 | 3 q 30 | 3996 | 4061 | 65 |
| 667 | 4126 | 4191 | 4256 | 4321 | 4386 | 4451 | 45.6 | 4581 | 4646 | 4711 | 65 |
| 668 | 4776 | 4841 | 4906 | 4971 | 5036 | 5101 | 5166 | 5231 | 5296 | 5361 | 65 |
| 869 | 5426 | 5491 | 5556 | 5621 | 5686 | 5751 | 5815 | 5880 | 5945 | 6010 | 65 |
| 670 | 6075 | 6140 | 6204 | 6269 | 6334 | 6399 | 6464 | 6528 | 6593 | 6658 | 65 |
| 671 | 6723 | 6787 | 6852 | 6917 | 6981 | 7046 | 7111 | 7175 | 7240 | 7305 | 65 |
| 672 | 7359 | 7434 | 7499 | 7563 | 7628 | 7692 | 7757 | 7821 | 7886 | 7951 | 65 |
| 673 | 8015 | 8080 | 8144 | 8209 | 8273 | 8338 | 8402 | 8467 | 8531 | $85{ }^{5}$ | 64 |
| 674 | 8660 | 8724 | 8789 | 8853 | 8918 | 8982 | 9046 | 9111 | 9175 | 9239 | 64 |
| 675 | 9304 | 9368 | 9432 | 9497 | 9561 | 9625 | 9690 | 9754 | 9818 | 9882 | 64 |
| 676 | * ${ }^{*} 9947$ | +011 | 0075 | O139 | 0204 | 0268 | o332 | o396 | 0460 | 0525 | 64 |
| 677 | 830589 | 0653 | 0717 | 0781 | 0845 | 0909 | 0973 | 1037 | 1102 | 1166 | 64 |
| 678 | 1230 | 1294 | 1358 | 1422 | 1486 | 1550 | 1614 | 1678 | 1742 | 1806 | 64 |
| 679 | 1870 | 1934 | 1998. | 2062 | 2126 | 2189 | 2253 | 2317 | 238I | 2445 | 64 |
| 680 | 2509 | 2573 | 2637 | 2700 | 2764 | 2828 | $289^{2}$ | 2956 | 3020 | 3083 | 64 |
| 681 | 3147 | 3211 | 3275 | 3338 | 3402 | 3466 | 3530 | 3593 | 3657 | 3721 | 64 |
| 682 | 3784 | 3848 | 3912 | 3975 | 4039 | 4103 | 4166 | 4230 | 4294 | 4357 | 64 |
| 683 | 4421 | 4484 | 4548 | 4611 | 4675 | 4739 | 4802 | 4866 | 4929 | 4993 | 64 |
| 684 | 5056 | 5120 | 5183 | 5247 | 5310 | 5373 | 5437 | 5500 | 5564 | 5627 | 63 |
| 685 | 5691 | 5754 | 5817 | 5881 | 5944 | 6007 | 6071 | 6134 | 6197 | 6261 | 63 |
| 686 | 6324 | 6387 | 6451 | 6514 | 6577 | 6641 | 6704 | 6767 | 6830 | 6894 | 63 |
| 687 688 | 6957 | 7020 | 7083 | 7146 | 7210 | 7273 | 7336 | 7399 | 7462 | 7525 | 63 |
| 688 | 7588 | 7652 | 7715 | 7778 | 7841 | 7904 | 7967 | 8030 | 8093 | 8156 | 63 |
| 689 | 8219 | 8282 | 8345 | 8408 | 8471 | 8534 | 8597 | 8660 | 8723 | 8786 | 63 |
| $6 \mathrm{6m}$ | 8849 | 8912 | 8975 | 9038 | 9101 | 9164 | 9227 | 9289 | 9352 | 9415 | 63 |
| 691 | * 9478 | 9541 | 9604 | 9667 | 9729 | 9792 | 9855 | 9918 | 9981 | +043 | 63 |
| 692 693 | 840106 | 0169 | 0232 | 0294 | 0357 | 0420 | 0482 | 0545 | 0608 | 0671 | 63 |
| 693 | 0733 | -796 | 0859 | $0{ }^{21} 1$ | -984 | 1046 | 1109 | 1172 | 1234 | 1297 | 63 |
| 694 | 1359 | 1422 | 1485 | 1547 | 1610 | 1672 | 1735 | 1797 | 1860 | 1922 | 63 |
| 695 | 1985 | 2047 | 2110 | 2172 | 2235 | 2297 | 2360 | 2422 | 2484 | 2547 | 62 |
| 696 | 2609 |  | 2734 | 2796 | 2859 | 2221 | 2983 | 3046 | 3108 | 3170 | 62 |
| 697 | 3233 3855 | 3295 | 3357 | 3420 | 3482 | 3544 | 3606 | 3669 | 3731 | 3793 | 62 |
| 698 699 | + 4857 | 3918 4539 | 3980 4601 | 4042 4664 | 4104 4726 | 4166 | 4229 4850 | 4291 4912 | 4353 | 4415 5036 | 62 62 |
|  |  |  |  |  |  |  |  |  |  |  |  |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | 1. |


| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 700 | 845098 | 5160 | 5222 | 5284 | 5346 | 5408 | 5470 | 5532 | 5594 | 5656 | 62 |
| 701 | 5718 | 5780 | 5842 | 5904 | 5966 | 6028 | 6090 | 6151 | 6213 | 6275 | 62 |
| 702 | 6337 | 3399 | 6.461 | 6523 | 6585 | 6646 | 6708 | 6770 | 6832 | 6894 | 62 |
| 703 | 6255 | 7017 | 7079 | 7141 | 7202 | 7264 | 7326 | 7388 | 7449 | 7511 | 62 |
| 704 | 7573 | 7634 | 7696 | 7758 | ${ }_{7} 819$ | 7881 | 7943 | 8004 | 8066 | 8128 | 62 |
| 705 | 8189 | 8251 | 8312 | 8374 | 84.35 | 8497 | 8559 | 8620 | 8682 | 8743 | 62 |
| 706 | 8805 | 8866 | 8928 | 8989 | 9051 | 9112 | 9174 | 9235 | 9397 | 9358 | 61 |
| 707 | 95419 | 9481 | 9542 | 9604 | 9665 | 9726 | 9788 | 9849 | 9911 | 9972 | 61 |
| 708 | 850033 | 0095 | 0156 | 0217 | 0279 | 0340 | 0401 | 0462 | 0524 | 0585 | 61 |
| 709 | 0646 | 0707 | 0769 | 0830 | 0891 | 0, 52 | 1014 | 1075 | 1136 | 1197 | 61 |
| 710 | 1258 | 1320 | I38ı | 1442 | :503 | 1564 | 1625 | 1686 | 1747 | 1809 | 61 |
| 711 | 1870 | $12^{31}$ | 1992 | 2053 | 2114 | 2175 | 2236 | -297 | 2358 | 2419 | 61 |
| 712 | 2480 | 2541 | 2602 | 2663 | 2724 | 2785 | 2846 | 2907 | 2968 | 3029 | 61 |
| 713 | 3090 | 3150 | 3211 | 3272 | 3333 | 3394 | 3455 | 3516 | 3577 | 3637 | 61 |
| 714 | 3698 | 3759 | 3820 | 3881 | 3941 | 4002 | 4063 | 4124 | 4185 | 4245 | 61 |
| 715 | 4306 | 4367 | 4428 | 4488 | 4549 | 4610 | 4670 | 4731 | 4792 | 4852 | 61 |
| 716 | 4913 | 4974 | 5034 | 5095 | 5156 | 5216 | 5277 | 5337 | 5398 | 5459 | 61 |
| 717 | 5519 | 5580 | 5640 | 5701 | 5761 | 5822 | 5882 | 5943 | 6003 | 6064 | 61 |
| 718 | 6124 | 6185 | 6245 | 6306 | 6366 | 6427 | 6487 | 6548 | 6608 | 6668 | 60 |
| 719 | 6729 | 6789 | 6850 | 6910 | 6970 | 7031 | 7091 | 7152 | 7212 | 7272 | 60 |
| 720 | 7332 | 7393 | 7453 | 7513 | 7574 | 7634 | 7694 | 7755 | 78.5 | 7875 | 60 |
| 721 | 7935 | 7995 | 8056 | 8116 | 8176 | 8236 | 8297 | 8357 | 8417 | 8477 | 60 |
| 722 | 8537 | 8597 | 8657 | 8718 | 8778 | 8838 | 8898 | 8958 | 9018 | 9078 | 60 |
| 723 | 9138 | 9198 | 9258 | 9318 | 9379 | 9439 | 9499 | 9559 | 9618 | 9679 | 60 |
| 724 | * 9739 | 9799 | 9859 | 9918 | 9978 | -038 | 0098 | 0158 | 0218 | 0278 | 60 |
| 725 | 860338 | -398 | 0458 | -5. 18 | 0578 | 0637 | 0697 | 0757 | 0817 | 0877 | 60 |
| 726 | $09^{37}$ | 0996 | 1056 | 1116 | 1176 | 1236 | 1295 | 1355 | 1415 | 1475 | 60 |
| 727 | 1534 | 1594 | 1654 | 1714 | 1773 | 1833 | $189^{3}$ | 1959 | 2012 | 2072 | 60 |
| 728 | 2131 | 2191 | 2251 | 2310 | 2370 | 2430 | 2489 | 2549 | 2608 | 2668 | 60 |
| 729 | 2728 | 2787 | 2847 | 2906 | 2966 | 3025 | 3085 | 3144 | 3204 | 3263 | 60 |
| 730 | 3323 | 3382 | 3442 | 3501 | 3561 | 3620 | 3680 | 3739 | 3799 | 3858 | 59 |
| 731 | 3 q 17 | 3977 | 4036 | 4096 | 4155 | 4214 | 4274 | 4333 | 4392 | 4452 | 59 |
| 732 | 4511 | 4570 | 4630 | 4689 | 4748 | 4808 | 4867 | 4226 | 4985 | 5045 | 50 |
| 733 | 5104 | 5163 | 5222 | 5282 | 5341 | 5400 | 5459 | 5519 | 5578 | 5637 | 59 |
| 734 | 5696 | 5755 | 5814 | 5874 | 5933 | 5992 | 6051 | 6110 | 6169 | 6228 | 59 |
| 735 | 6287 | 6346 | 6405 | 6465 | 6524 | 6583 | 6642 | 6701 | 6760 | 6819 | 59 |
| 736 | 6878 | $69^{37}$ | 6996 | 7055 | 7114 | 7173 | 7232 | 7291 | 7350 | 7409 | 59 |
| 737 | 7467 | 7526 | 7585 | 7644 | 7703 | 7762 | 7821 | 7880 | 7939 | 7998 | 59 |
| 738 | 8056 | 8115 | 8174 | 8233 | 8292 | 8350 | 8409 | 8468 | 8527 | 8586 | 59 |
| 739 | 8644 | 8703 | 8762 | 8821 | 8879 | 8938 | 8997 | 9056 | 9114 | 9173 | 59 |
| 740 | 9232 | 9290 | 9349 | 9408 | 9466 | 9525 | 9584 | 9642 | 9701 | 9760 | 59 |
| 741 | *9818 | 9877 | 9935 | 9994 | -053 | 0111 | 0170 | 0228 | 0287 | 0345 | 5 |
| 742 | 870404 | 0462 | 0521 | 0579 | 0638 | 0696 | 0755 | 08ı3 | 0872 | $\bigcirc$ | 58 |
| 743 | 0989 | 1047 | 1106 | 1164 | 1223 | 1281 | 1339 | 1308 | 1456 | 15.5 | 58 |
| 744 | 1573 | 1631 | 1690 | 1748 | 1806 | 1865 | 1923 | 1981 | 2040 | 2098 | 58 |
| 745 | 2156 | 2215 | 2273 | 2331 | 2389 | 2448 | 2506 | 2564 | 2622 | 2681 | 58 |
| 746 | 2739 | 2797 | 2855 | 2913 | 2972 | 3030 | 3088 | 3146 | 3204 | 3262 | 58 |
| 747 | 3321 | 3379 | 3437 | 3495 | 3553 | 3611 | 3669 | 3727 | 3785 | 3844 | 58 |
| 748 | 3902 | 3960 | 4018 | 4076 | 4134 | 4192 | 4250 | 4308 | 4366 | 4424 | 58 |
| 749 | 4482 | 4540 | 4598 | 4656 | 4714 | 4772 | 4830 | 4888 | 4945 | 5003 | 58 |
| 750 | 5061 | 5119 | 5177 | 5235 | 5293 | 5351 | 5409 | 5466 | 5524 | 5582 | 58 |
| 751 | 5640 | 5698 | 5756 | 5813 | 5871 | 5929 | 5987 | 6045 | 6102 | 6160 | 58 |
| 752 | 6218 | 6276 | 6333 | 6391 | 6449 | 6507 | 6564 | 6622 | 6680 | 6737 | 58 |
| 753 | 6795 | 6853 | 6910 | 6968 | 7026 | 7083 | 7141 | 7199 | 7256 | 7314 | 58 |
| 754 | 7371 | 7429 | 7487 | 7544 | 7602 | 7659 | 7717 | 7774 | 7832 | 7889 | 58 |
| 755 | 7947 | 8004 | 8062 | 8119 | 8177 | 8234 | 8292 | 8349 | 8407 | 8464 | 57 |
| 756 | 8522 | 8579 | 8637 | 8694 | 8752 | 8809 | 8866 | 8924 | 8081 | 9039 | 57 |
| 757 | 9096 | 9153 | 9211 | 9268 | 9325 | 9383 | 9440 | 9497 | 9555 | 9612 | 57 |
| 758 | * 9669 | 9726 | 9784 | 9841 | 9898 | 9956 | +013 | 0070 | 0127 | 0185 | 57 |
| 759 | 880242 | 0299 | 0356 | 0413 | 0471 | 0528 | 0585 | 0642 | 0699 | 0756 | 57 |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |


| Table I. |  |  | LOGARITHMS OF NUMBERS. |  |  |  |  |  |  |  | 18 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |
| 760 | 880814 | 0871 | -928 | 0985 | 1042 | 1099 | 1156 | 1213 | 1271 | 1328 | 57 |
| 761 | 1385 | 1442 | 1499 | 1556 | 1613 | 1670 | 1727 | 1784 | 1841 | 1898 | 57 |
| 762 | 1955 | 2012 | 2069 | 2126 | 2183 | 2240 | 2297 | 2354 | 2411 | 2468 | 57 |
| 763 | 2525 | 2581 | 2638 | 2695 | 2752 | 2809 | 2866 | 2923 | 2980 | 3037 | 57 |
| 764 | 3093 | 3150 | 3207 | 3264 | 3321 | 3377 | 3434 | 3491 | 3548 | 3605 | 57 |
| 765 | 366I | 3718 | 3775 | 3832 | 3888 | 3945 | 4002 | 4059 | 4115 | 4172 | 57 |
| 766 | 4229 | 4285 | 4342 | 4392 | 4455 | 4512 | 4569 | 4625 | 4682 | 4739 | 57 |
| 767 | 4795 | 4852 | 4909 | 4965 | 5022 | 5078 | 5.35 | 5192 | 52.8 | 5305 | 57 |
| 768 | 5361 | 5418 | 5474 | 5531 | 5587 | 5644 | 5700 | 5757 | 58.3 | 5870 | 57 |
| 769 | 5926 | 5983 | 6039 | 6096 | 6152 | 6209 | 6265 | 6321 | 6378 | 6434 | 50 |
| 770 | 6491 | 6547 | 6604 | 6660 | 6716 | 6773 | 6829 | 6885 | 6942 | 6998 | 56 |
| 771 | 7034 | 7111 | 7167 | 7223 | 7280 | 7336 | 7392 | 7449 | 7505 | 7561 | 56 |
| 772 | 7617 | 7674 | 7730 | 7786 | 7842 | 7898 | 7955 | 8011 | 8067 | 8123 | 56 |
| $77^{3}$ | 8179 | 8236 | 8292 | 8348 | 8404 | 8460 | 8516 | 8573 | 8629 | 8685 | 56 |
| 774 | 8741 | 8797 | 8853 | 8909 | 8965 | 9021 | 9077 | 9134 | 9190 | 9246 | 56 |
| 775 | 9302 | 9358 | 9414 | 9470 | 9526 | 9582 | 9638 | 9694 | 9750 | 9806 | 56 |
| 776 | * 9862 | 9918 | 9974 | -030 | 0086 | 0141 | 0197 | 0253 | -309 | o365 | 56 |
| 777 | 890421 | 0477 | 0533 | -589 | 0645 | 0700 | -756 | 0812 | 0868 | 0924 | 56 |
| 778 | og80 | 1035 | 1091 | 1147 | 1203 | 1259 | 1314 | 1370 | 1426 | 1482 | 56 |
| 779 | 1537 | 1593 | 1649 | 1705 | 1760 | 1816 | 1872 | 1928 | 1983 | 2039 | 56 |
| 780 | 2095 | 2150 | 2206 | 2262 | 2317 | 2373 | 2429 | 2484 | 2540 | 2595 | 56 |
| 781 | 2651 | 2707 | 2762 | 2818 | 2873 | 2929 | 2985 | 3040 | 3096 | 3151 | 56 |
| 782 | 3207 | 3262 | 3318 | 3373 | 3429 | 3484 | 3540 | $352^{5}$ | 3651 | 3706 | 56 |
| 783 | 3762 | 3817 | 3873 | 3928 | 3984 | 4039 | 4094 | 4150 | 4205 | 4261 | 55 |
| 784 | 4316 | 4371 | 4427 | 4482 | 4538 | 4593 | 4648 | 4704 | $47^{5} 9$ | 4814 | 55 |
| 785 | 4870 | 4925 | 4980 | 5036 | 5091 | 5146 | 5201 | 5257 | 5312 | 5367 | 55 |
| 786 | 5423 | 5478 | 5533 | 5588 | 5644 | 5699 | 5754 | 5809 | 5864 | 5920 | 55 |
| 787 | 5975 | 6030 | 6085 | 6140 | 6195 | 6251 | 6306 | 6361 | 6416 | 6471 | 55 |
| 783 | 6526 | 6581 | 6636 | 6692 | 6747 | 6802 | 6857 | 6912 | 6967 | 7022 | 55 |
| 789 | 7077 | 7132 | 7187 | 7242 | 7297 | 7352 | 7407 | 7452 | 7517 | 7572 | 55 |
| 790 | 7627 | 7682 | 7737 | 7792 | 7847 | 7902 | 7957 | 8012 | 8067 | 8122 | 55 |
| 791 | 8176 | 8231 | 8286 | 8341 | 8396 | 8451 | 8506 | 8561 | 86.5 | 8670 | 55 |
| 792 | 8725 | 8780 | 8835 | 8890 | 8944 | 8999 | 9054 | 9109 | 9164 | 9218 | 55 |
| 793 | 9273 | 9328 | 9383 | 9437 | 9492 | 9547 | 9602 | 9656 | 9711 | 9766 | 55 |
| 794 | * 982I | 9875 | 9930 | 9985 | -039 | 0094 | 0149 | 0203 | 0258 | 0312 | 55 |
| 795 | 900367 | 0422 | 0476 | 0531 | 0586 | 0640 | 0695 | 0742 | 0804 | 0859 | 55 |
| 796 | 0913 | 0968 | 1022 | 1077 | 1131 | 1186 | 1240 | 1295 | 1349 | 1404 | 55 |
| 797 | 1458 | 1513 | 1567 | 1622 | 1676 | 1731 | 1785 | 1840 | 1894 | 1948 | 54 |
| 798 | 2003 | 2057 | 2112 | 2166 | 2221 | 2275 | 2329 | 2384 | 2438 | 2492 | 54 |
| 799 | 2547 | 2601 | 2655 | 2710 | 2764 | 2818 | 2873 | 2927 | 2981 | 3036 | 54 |
| 800 | 3090 | 3144 | 3199 | 3253 | 3307 | 3361 | 3416 | 3470 | 3524 | 3578 | 54 |
| 801 | 3633 | 3687 | 3741 | 3795 | 3849 | 3904 | 3 g 58 | 4012 | 4066 | 4120 | 54 |
| 802 | 4174 | 4229 | 4283 | 4337 | 4391 | 4445 | 4499 | 4553 | 4607 | 4661 | 54 |
| 803 | 4716 | 4770 | 4824 | 4878 | 4932 | 4986 | 5040 | 5094 | 5148 | 5202 | 54 |
| 804 | 5256 | 5310 | 5364 | 5418 | 5472 | 5526 | 5580 | 5634 | 5688 | 5742 | 54 |
| 805 | 5796 | 5850 | 5904 | 5, 58 | 6012 | 6066 | 6119 | 6173 | 6227 | 6281 | 54 |
| 806 | 6335 | 6389 | 6443 | 6497 | 6551 | 6604 | 6658 | 6712 | 6766 | 6820 | 54 |
| 807 | 6874 | 6927 | 6981 | 7035 | 7089 | 7143 | 7196 | 7250 | 7304 | 7358 | 54 |
| 808 | 7411 | 7465 | 7 719 | 7573 | 7626 | 7680 | 7734 | $\bigcirc 78$ | 7841 | $78{ }^{5}$ | 54 |
| 809 | 7949 | 8002 | 8056 | 8110 | 8163 | 8217 | 8270 | 8324 | 8378 | 8431 | 54 |
| 810 | 8485 | 8539 | 8592 | 8646 | 8699 | 8753 | 8807 | 8860 | 8914 | 8967 | 54 |
| 811 | 9021 | 9074 | 9128 | 9181 | 9235 | 9289 | 9342 | 9396 | 9449 | 9503 | 54 |
| 812 | * 9556 | 9610 | 9663 | 9716 | 9770 | 9823 | 9877 | 9930 | 9284 | +037 | 53 |
| 813 | 910091 | 0144 | 0197 | 0251 | 0304 | 0358 | 0411 | 0464 | 0518 | 0571 | 53 |
| 814 | 0624 | 0678 | 0731 | 0784 | 0838 | 0891 | 0944 | 0998 | 1051 | 1104 | 53 |
| 815 | 1158 | 1211 | 1264 | 1317 | ${ }^{1371}$ | 1424 | 1477 | 1530 | 1584 | 1637 | 53 |
| 816 | 1070 | 1743 |  | 1850 | 1903 | 1956 | 2009 | 2063 | 2116 | 2169 | 53 |
| 817 | 2222 | 2275 | 2328 | 2381 | 2435 | 2488 | 2541 | 2594 | 2647 | 2700 | 53 |
| 818 | 2753 | 2806 | 2859 | 2913 | 2966 | 3019 | 3072 | 3125 | 3178 | 3231 | 53 |
| 819 | 3284 | 3337 | 3390 | 3443 | 3496 | 3549 | 3602 | 3655 | 3708 | 3761 | 53 |
| N . | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |


| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 820 | 913814 | 3867 | 3920 | 3973 | 4026 | 4079 | 4132 | 4184 | 4237 | 4290 | 53 |
| 821 | 4343 | 4396 | 4449 | 4502 | 4555 | 4608 | 4660 | 4713 | 4766 | 4819 | 53 |
| 822 | 4872 | 4925 | 4977 | 5030 | 5083 | 5ı36 | 5189 | 5241 | 5294 | 5347 | 53 |
| 823 | 5400 | 5453 | 5505 | 5558 | 5611 | 5664 | 5716 | 5769 | 5822 | 5875 | 53 |
| 824 | 5927 | 5980 | 6033 | 6085 | 6, 38 | 6191 | 6243 | 6296 | 6349 | 6401 | 53 |
| 825 | 6454 | 6507 | 6559 | 6612 | 6664 | 6717 | 6770 | 6822 | 6875 | 6927 | 53 |
| 826 | 6980 | 7033 | 7085 | 7138 | 7190 | 7243 | 7295 | 7348 | 7400 | 7453 | 53 |
| 827 | 7506 | 7558 | 7611 | 7663 | 7716 | 7768 | 7820 | 7873 | 7925 | 7978 | 52 |
| 828 | 8030 | 8083 | 8135 | 8188 | 8240 | 8293 | 8345 | 8397 | 8450 | 8502 | 52 |
| 829 | 8555 | 8607 | 8659 | 8712 | 8764 | 8816 | 8869 | 8921 | 8973 | 9026 | 52 |
| 830 | 9078 | 9130 | 9183 | 9235 | 9287 | 9340 | 9392 | 9444 | 9496 | 9549 | 52 |
| 831 | * 9601 | 9653 | 9706 | 9758 | 9810 | 9862 | 9914 | 9967 | +019 | 0071 | 52 |
| 832 | 920123 | 0176 | 0228 | 0280 | 0332 | -384 | 0436 | 0489 | 0541 | -593 | 52 |
| 833 | 0645 | 0697 | 0749 | 0801 | 0853 | 0906 | -958 | 1010 | 1062 | 1114 | 52 |
| 834 | 1166 | 1218 | 1270 | 1322 | 1374 | 1426 | 1478 | 1530 | 1582 | 1634 | 52 |
| 835 | 1686 | 1738 | 1790 | 1842 | 1894 | 1946 | 1298 | 2050 | 2102 | 2154 | 52 |
| 836 | 2206 | 2258 | 2310 | 2362 | 2414 | 2466 | 2518 | 2570 | 2622 | 2674 | 52 |
| 837 | 2725 | 2777 | 2829 | 2881 | 2933 | 2985 | 3037 | 3089 | 3140 | 3192 | 52 |
| 838 | 3244 | 3296 | 3348 | 3399 | 3451 | 3503 | 3555 | 3607 | 3658 | 3710 | 52 |
| 839 | 3762 | 3814 | 3865 | 3917 | 3969 | 4021 | 4072 | 4124 | 4176 | 4228 | 52 |
| 840 | 4279 | 433I | 4383 | 4434 | 4486 | 4538 | 4589 | 4641 | 4693 | 4744 | 52 |
| 841 | 4796 | 4848 | 4899 | 4951 | 5003 | 5054 | 5106 | 5157 | 5209 | 5201 | 52 |
| 842 | 5312 | 5364 | 5415 | 5467 | 5518 | 5570 | 5621 | 5673 | 5725 | 5776 | 52 |
| 843 | 5828 | 5879 | 593I | 5982 | 6034 | 6085 | 6137 | 6188 | 6240 | 6291 | 5I |
| 844 | 6342 | 6394 | 6445 | 6497 | 6548 | 6600 | 6651 | 6702 | 6754 | 6805 | 5I |
| 845 | 6857 | 6908 | 6959 | 7011 | 7062 | 7114 | 7165 | 7216 | 7268 | 7319 | 51 |
| 846 | 7370 | 7422 | 7473 | 7524 | 7576 | 7627 | 7678 | 7730 | 7781 | 7832 | 51 |
| 847 | 7883 | 7935 | 7986 | 8037 | 8088 | 8140 | ${ }^{8191}$ | 8242 | 8293 | 8345 | 51 |
| 848 | 8396 | 8447 | 8498 | 8549 | 8601 | 8652 | 8703 | 8754 | 8805 | 8857 | 51 |
| 849 | 8908 | $89^{59}$ | 9010 | 9061 | 9112 | 9163 | 9215 | 9266 | $9^{317}$ | 9368 | 51 |
| 850 | 9419 | 9470 | 9521 | 9572 | $9^{662}$ | 9674 | 9725 | 9776 | 9827 | 9879 | 51 |
| 851 | * 9930 | 9981 | +032 | 0083 | 0134 | 0185 | 0236 | 0287 | 0338 | -389 | 51 |
| 852 | 930440 | 0491 | 0542 | -5.92 | 0643 | 0694 | 0745 | 0796 | 0847 | -898 | 51 |
| 853 | 0949 | 1000 | 1051 | 1102 | 1153 | 1204 | 1254 | 1305 | 1356 | 1407 | 51 |
| 854 | 1458 | 1509 | 1560 | 1610 | 1661 | 1712 | 1763 | 1814 | 1865 | 1915 | 5I |
| 855 | 1966 | 2017 | 2068 | 2118 | 2169 | 2220 | 2271 | 2322 | 2372 | 2423 | 51 |
| 856 | 2474 | 2524 | 2575 | 2626 | 2677 | 2727 | 2778 | 2829 | 2879 | 2930 | 5 I |
| 857 | 2981 | 3031 | 3082 | 3133 | 3183 | 3234 | 3285 | 3335 | 3386 | 3437 | 51 |
| 858 | 3487 | 3538 | 3589 | 3639 | 3690 | 3740 | 3791 | 3841 | 3892 | 3943 | ${ }_{5}^{5}$ |
| 859 | 3993 | 4044 | 4094 | 4145 | 4195 | 4246 | 4296 | 4347 | 4397 | 4448 | 51 |
| 860 | 4498 | 4549 | 4599 | 4650 | 4700 | 4751 | 4801 | 4852 | 4902 | 4953 | 50 |
| 861 | 5003 | 5054 | 5104 | 5154 | 5205 | 5255 | 5306 | 5356 | 5406 | 5457 | 50 |
| 862 | 5507 | 5558 | 5608 | 5658 | 5709 | 5759 | 5809 | 5860 | 5910 | 5960 | 50 |
| 853 | 6011 | 6061 | 6111 | 6162 | 6212 | 6262 | 6313 | 6363 | 6413 | 6463 | 50 |
| 864 | 6514 | 6564 | 6614 | 6665 | 6715 | 6765 | 6815 | 686 | 6916 | 6,66 | 50 |
| 865 | 7016 | 7066 | 7117 | 7167 | 7217 | 7267 | 7317 | 7367 | 7418 | 7468 | 50 |
| 866 | 7518 | 7568 | 7618 | 7668 | 7718 | 7769 | 7819 | 7869 | 7919 | 7969 | 50 |
| 867 | 8019 | 8069 | 8119 | 8169 | 8219 | 8269 | 8320 | 8370 | 8420 | 8470 | 50 |
| 868 | 8520 | 8570 | 8620 | 8670 | 8720 | 8770 | 8820 | 8870 | 8920 | 8970 | 50 |
| 869 | 9020 | 9070 | 9120 | 9170 | 9220 | 9270 | 9320 | 9369 | 9419 | 9469 | 50 |
| 870 | ${ }^{9519}$ | $\bigcirc{ }^{9} 569$ |  | 9669 |  |  |  |  | 9918 |  | 50 |
| 871 | 940018 | 0068 | 0118 | 0168 | 0218 | 0267 | 0317 | 0367 | 0417 | 0467 | 50 |
| 872 | 0516 | 0566 | 0616 | 0666 | 0716 | 0765 | 0815 | 0865 | 0915 | 0964 | 50 |
| 873 | 1014 | 1064 | 1114 | 1163 | 1213 | 1263 | 1313 | 1362 | 1412 | 1462 | 50 |
| 874 | 1511 | 1561 | 1611 | 1660 | 1710 | 1760 | 1809 | 1859 | 1909 | 1938 | 50 |
| 875 | 2008 | 2058 | 2107 | 215 | 2207 | 2256 | 2306 | 2355 | 2405 | 2455 | 50 |
| 876 | 2504 | 2554 | 2603 | 2653 | 2702 | 2752 | 2801 | 2851 | 2901 | 2950 | 50 |
| 877 | 3000 | 3049 | 3099 | 3148 | 3198 | 3247 | 3297 | 3346 | 3396 | 3445 | 49 |
| 878 | 3495 | 3544 | 3503 | 3643 | 3692 | 3742 | 3791 | $38^{4} 41$ | 3890 | 3939 | 49 |
| 870 | 3989 | 4038 | 4088 | 4137 | 4186 | 4236 | 4285 | 4335 | 4384 | 4433 | 49 |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |


| Table I. |  |  | LOGARITHMS OF NUMBERS. |  |  |  |  |  |  |  | 15 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |
| 880 | 944483 | 4532 | 4581 | 4631 | 4680 | 4729 | 4779 | 4828 | 4877 | 4927 | 49 |
| 881 | 4976 | 5025 | 5074 | 5124 | 5173 | 5222 | 5272 | 5321 | 5370 | 5419 | 49 |
| 882 | 5469 | 5518 | 5567 | 5616 | 5665 | 5715 | 5764 | 5813 | 5862 | 5912 | 49 |
| 883 | 5 c 51 | 6010 | 6059 | 6108 | 6157 | 6207 | 6256 | 6305 | 6354 | 6403 | 49 |
| 884 | 6452 | 6501 | 6551 | 6600 | 6649 | 6698 | 6747 | 6796 | 6845 | 6894 | 49 |
| 885 | 6943 | 6992 | 7041 | 7090 | 7140 | 7189 | 7238 | 7287 | 7336 | 7385 | 49 |
| 886 | 7434 | 7483 | 7532 | 7581 | 7630 | 7679 | 7728 | 7777 | 7826 | 7875 | 49 |
| 887 | 7924 | 7973 | 8022 | 8070 | 8119 | 8168 | 8217 | 8266 | 8315 | 8364 | 49 |
| 888 | 8413 | 8462 | 8511 | 8560 | 8609 | 8657 | 8706 | 8755 | 8804 | 8853 | 49 |
| 889 | 8902 | 8951 | 8999 | 9048 | 9097 | 9146 | 9195 | 9244 | 9292 | 9341 | 49 |
| 8 8о | 9390 | 9439 | 9488 | ${ }_{9} 536$ | 9585 | 9634 | 9683 | 9731 | 9780 | 9829 | 49 |
| 891 | * 9878 | 9926 | 9975 | +024 | 0073 | 0121 | 0170 | 0219 | 0267 | 0316 | 49 |
| 892 | 950365 | 0414 | 0462 | -511 | 0560 | o608 | 0657 | 0706 | 0754 | 0803 | 49 |
| 893 | 0851 | -900 | 0949 | 0997 | 1046 | 1095 | 1143 | 1192 | 1240 | 1289 | 49 |
| 894 | 1338 | 1386 | 1435 | 1483 | 1532 | 1580 | 1629 | 1677 | 1726 | 1775 | 49 |
| 895 | 1823 | 1872 | 1920 | 1969 | 2017 | 2066 | 2114 | 2163 | 2211 | 2260 | 48 |
| 896 | 2308 | 2356 | 2405 | 2453 | 2502 | 2550 | 2599 | 2647 | 2696 | 2744 | 48 |
| 897 | 2792 | 2841 | 2889 | 2938 | 2986 | 3034 | 3083 | 3 I 3 I | 3180 | 3228 | 48 |
| 898 | 3276 | 3325 | 3373 | 3421 | 3470 | 3518 | 3566 | 3615 | 3663 | 3711 | 48 |
| 899 | 3760 | 3808 | 3856 | 3905 | 3953 | 4001 | 4049 | 4098 | 4146 | 4194 | 48 |
| 900 | 4243 | 4291 | 4339 | 4387 | 4435 | 4484 | 4532 | 4580 | 4628 | 4677 | 48 |
| 901 | 4725 | 4773 | 4821 | 4869 | 4918 | 4966 | 5014 | 5062 | 5110 | 5158 | 48 |
| 902 | 5207 | 5255 | 5303 | 5351 | 5399 | 5447 | 5495 | 5543 | 5592 | 5640 | 48 |
| 903 | 5688 | 5736 | 5784 | 5832 | 5880 | 5928 | 5976 | 6024 | 6072 | 6120 | 48 |
| 904 | 6168 | 6216 | 6265 | 6313 | 6361 | 6409 | 6457 | 6505 | 6553 | 6601 | 48 |
| 905 | 6649 | 6697 | 6745 | 6793 | 6840 | 6888 | 6936 | 6984 | 7032 | 7080 | 48 |
| 906 | 7128 | 7176 | 7224 | 7272 | 7320 | 7368 | 7416 | 7464 | 7512 | 7559 | 48 |
| 907 | 7607 | 7655 | 7703 | 7751 | 7799 | 7847 | 7894 | 7942 | 7990 | 8038 | 48 |
| 908 | 8086 | 8134 | 8181 | 8229 | 8277 | 8325 | 8373 | 8421 | 8468 | 8516 | 48 |
| 909 | 8564 | 8612 | 8659 | 8707 | 8755 | 8803 | 8850 | 8898 | 8946 | 8994 | 48 |
| 910 | 9041 | 9089 | 9137 | 9185 | 9232 | 9280 | 9328 | 9375 | 9423 | 9471 | 48 |
| 911 | 9518 | 9566 | 9614 | 9661 | 9709 | 9757 | 9804 | 9852 | 9900 | 9947 | 48 |
| 912 | * 9995 | +042 | 0090 | -138 | - 185 | 0233 | 0280 | o328 | o376 | 0423 | 48 |
| 913 | 960471 | -518 | o566 | 0613 | 0661 | 0709 | 0750 | 0804 | 0851 | -899 | 48 |
| 914 | 0946 | 0994 | 1041 | 1089 | 1136 | 1184 | 1231 | 1279 | 1326 | 1374 | 47 |
| 915 | 1421 | 1469 | 1516 | 1563 | 1611 | 1658 | 1706 | 1753 | 1801 | 1848 | 47 |
| 916 | 1895 | 1943 | 1990 | 2038 | 2085 | 2132 | 2180 | 2227 | 2275 | 2322 | 47 |
| 917 | 2369 | 2417 | 2464 | 2511 | 2559 | 2606 | 2653 | 2701 | 2748 | ${ }^{2} 795$ | 47 |
| 918 | 2843 | 2890 | 2937 | 2985 | 3032 | 3079 | 3126 | 3174 | 3221 | 3268 | 47 |
| 919 | 3316 | 3363 | 3410 | 3457 | 3504 | 3552 | 3599 | 3646 | 3693 | 3741 | 47 |
| 920 | 3788 | 3835 | 3882 | 3929 |  | 4024 | 4071 | 4118 | 4165 | 4212 | 47 |
| 921 | 4260 | 4307 | 4354 | 4401 | 4448 | 4495 | 4542 | 4590 | 4637 | 4684 | 47 |
| 922 | 4731 | 4778 | 4825 | 4872 | 4919 | 4966 | 5013 | 5061 | 5108 | 5155 | 47 |
| 923 | 5202 | 5249 | 5296 | 5343 | 5390 | 5437 | 5484 | 5531 | 5578 | 562.5 | 47 |
| 924 | 5672 | 5719 | 5766 | 5813 | 5860 | 5907 | 5954 | 6001 | 6048 | 6095 | 47 |
| 925 | 6142 | 6189 | 6236 | 6283 | 6329 | 6376 | 6423 | 6470 | 6517 | 6564 | 47 |
| 926 | 6611 | 6658 | 6705 | 6752 | 6799 | 6845 | 6892 | 6939 | 6986 | 7033 | 47 |
| 927 | 7080 | 7127 | 7173 | 7220 | 7267 | 7314 | 7361 | 7408 | 7454 | 7501 | 47 |
| 928 | 7548 | 7595 | 7642 | 7688 | 7735 | 7782 | 7829 | 7875 | 7922 | 7969 | 47 |
| 929 | 8016 | 8062 | 8109 | 8156 | 8203 | 8249 | 8296 | 8343 | 8390 | 8436 | 47 |
| 930 | 8483 | 8530 | 8576 | 8623 | 8670 | 8716 | 8763 | 88ı0 | 8856 | 8903 | 47 |
| 931 | 8950 | 8996 | 9043 |  | 9136 | 9183 |  | 9276 | 9323 | 9369 | 47 |
| 932 | + 9416 | 9463 | 9509 | 9556 | 9602 | 9649 | 9695 | 9742 | 9789 | 9835 | 47 |
| 933 | * 9882 | 9928 | 9975 | +021 | 0068 | 0114 | 0161 | 0207 | 0254 | 0300 | 47 |
| 934 | 970347 | -393 | 0440 | 0486 | $\bigcirc 533$ | -5̇79 | 0626 | 0672 | 0719 | 0765 | 46 |
| 935 | 0812 | ט858 | 0904 | 0, 51 | 0997 | 1044 | 1090 | 1137 | 1183 | 1229 | 46 |
| 936 | 1276 | 1322 | 1369 | 1415 | 1461 | 1508 | 1554 | 1601 | 1647 | $162^{3}$ | 46 |
| 437 | 1740 | 1786 | 1832 | 1879 | 1225 | 1971 | 2018 | 2064 | 2110 | 2157 | 46 |
| 938 939 | 2203 | 2249 2712 | 2295 2758 | 2342 2804 | 2388 2851 | 2434 2897 | 2481 2943 | 2527 2989 | 2573 3035 | 2619 3082 | 46 46 |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |

LOGARITHMS OF NUMBERS.
Table I.

| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | D. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 940 | 973128 | 3174 | 3220 | 3266 | 3313 | 3359 | 3405 | 3451 | 3497 | 3543 | 46 |
| 941 | 35190 | 3636 | 3682 | 3728 | 3774 | 3820 | 3866 | 3و13 | 3959 | 4005 | 46 |
| 942 | 4051 | 4097 | 4143 | 4189 | 4235 | 4281 | 4327 | 4374 | 4420 | 4466 | 46 |
| 943 | 4512 | 4508 | 4604 | 4650 | 4696 | 4742 | 4788 | 4834 | 4880 | 4926 | 46 |
| 944 | 4972 | 5018 | 5064 | 5110 | 5156 | 5202 | 5248 | 5294 | 5340 | 5386 | 46 |
| 945 | 5432 | 5478 | 5524 | 5570 | 5616 | 5662 | 5707 | 5753 | 5799 | 5845 | 46 |
| 946 | 5891 | 5937 | 5983 | 6029 | 6075 | 6121 | 6167 | 6212 | 6258 | 6304 | 46 |
| 947 | 6350 | 6396 | 6442 | 6488 | 6533 | 6579 | 6625 | 6671 | 6717 | 6763 | 46 |
| 948 | 6808 | 6854 | 6900 | 6946 | 6992 | 7037 | 7083 | 7129 | 7175 | 7220 | 46 |
| 949 | 7266 | 7312 | 7358 | 7403 | 7449 | $749^{5}$ | 7501 | 7586 | 7632 | 7678 | 46 |
| 950 | 7724 | 7769 | 7815 | 7851 | 7906 | 7952 | 7998 | 8043 | 8089 | 8135 | 46 |
| 951 | 8181 | 8226 | 8272 | 8317 | 8363 | 8409 | 8454 | 8500 | 8546 | 85 g 1 | 46 |
| $9{ }^{5} 2$ | 8637 | 8683 | 8728 | 8774 | 88ı9 | 8865 | 8911 | 8956 | 9002 | 9047 | 46 |
| 950 | 9093 | 9138 | 9184 | 9230 | 9275 | 9321 | 9366 | 9412 | 9457 | 9503 | 46 |
| 954 | 950 | 9594 | 9639 | 9685 | 9730 | 9776 | 9821 | 9867 | 9912 | 9958 | 46 |
| 955 | 980003 | 0049 | 0094 | 0140 | 0185 | 0231 | 0276 | 0322 | 0367 | 0412 | 45 |
| 956 | 0458 | -5́o3 | -5549 | -0594 | 0640 | 0685 | 0730 | 0776 | 0821 | 0867 | 45 |
| 950 | 0912 | 0, ${ }^{\text {¢ }}$ | 1003 | 1048 | 1093 | 1139 | 1184 | 1229 | 1275 | 1320 | 45 |
| 938 | 1366 | 1411 | 1456 | 1501 | 1547 | 1592 | 1637 | 1683 | 1728 | 1773 | 45 |
| 9プ9 | 1819 | 1864 | 1909 | 1954 | 2000 | 2045 | 2090 | 2135 | 2181 | 2226 | 45 |
| 960 | 2271 | 2316 | 2362 | 2407 | 2452 | 2497 | 2543 | 2588 | 2633 | 2678 | 45 |
| 961 | 2723 | 2769 | 2814 | 2859 | 2904 | 2949 | 2994 | 3040 | 3085 | 3130 | 45 |
| 962 | 3175 | 3220 | 3265 | 3310 | 3356 | 3401 | 3446 | 3491 | 3536 | 3581 | 45 |
| 963 | 3626 | 3671 | 3716 | 3702 | 3807 | 3852 | 3897 | 3942 | 3987 | 4032 | 45 |
| 964 | 4077 | 4122 | 4167 | 4212 | 4257 | 4302 | 4347 | 4392 | 4437 | 4482 | 45 |
| 965 | 4527 | 4572 | 4617 | 4662 | 4707 | 4752 | 4797 | 4842 | 4887 | 4932 | 45 |
| 966 | 4977 | 5022 | 5067 | 5112 | 5157 | 5202 | 5247 | ${ }^{5} 292$ | 5337 | 5382 | 45 |
| 967 | 5426 | 5471 | 5516 | 5561 | 5606 | 5651 | 5696 | 5741 | 5786 | 5830 | 45 |
| 968 | 5875 | 5920 | 5965 | 6010 | 6055 | 6100 | 6144 | 6189 | 6234 | 6279 | 45 |
| 969 | 6324 | 6369 | 6413 | 6458 | 6503 | 6548 | 6593 | 6637 | 668; | 6727 | 45 |
| 970 | 6772 | 6817 | 6861 | 6906 | 6951 | 6996 | 7040 | 7085 | 7130 | 7175 | 43 |
| 971 | 7219 | 7264 | 7309 | 7353 | 7398 | 7443 | 7488 | 7532 | 7577 | 7622 | 45 |
| 972 | 7666 | 7711 | 7756 | 7800 | 7845 | 7890 | 7934 | 7979 | 8024 | 8068 | 45 |
| 973 | 8113 | 8157 | 8202 | 8247 | 8291 | 8336 | 8381 | 8425 | 8470 | 8514 | 45 |
| 974 | 85 ¢̇9 | 8604 | 8648 | $869^{3}$ | 8737 | 8782 | 8826 | 8871 | 8916 | 8960 | 45 |
| 975 | 9005 | 9049 | 9004 | 9138 | 9183 | 9227 | 9272 | 9316 | 9361 | 9405 | 45 |
| 976 | 9450 | 9494 | $00^{0} 39$ | 9583 | 9628 | 9672 | 9717 | 9761 | 9806 | 9850 | 44 |
| 977 | - $9^{8} 95$ | 9939 | 9983 | -028 | 0072 | 0117 | 0161 | 0206 | 0250 | 0294 | 44 |
| 978 | 990339 | 0383 | 0428 | 0472 | 0516 | 0561 | 0605 | -650 | 0694 | 0738 | 44 |
| 979 | 0783 | 0827 | 0871 | 0916 | 0,60 | 1004 | 1049 | 1093 | 1137 | 1182 | 44 |
| 980 | 1226 | 1270 | 1315 | 1359 | 1403 | 1448 | 1492 | 1536 | 1580 | 1625 | 44 |
| $99^{1}$ | 1669 | 1713 | 1758 | 1802 | 1846 | 1890 | $10^{35}$ | 1979 | 2023 | 2067 | 44 |
| 982 | 2111 | 2156 | 2200 | 2244 | 2288 | 2333 | 2377 | 2421 | 2465 | 2509 | 44 |
| 983 | 2554 | 2598 | 2642 | 2686 | 2730 | 2774 | 2819 | 2863 | 2907 | 2051 | 44 |
| 984 | 2995 | 3039 | 3083 | 3127 | 3172 | 3216 | 3200 | 3304 | 3348 | 3392 | 44 |
| 985 | 3436 | 3480 | 3524 | 3568 | 3613 | 3657 | 3701 | 3745 | 3789 | 3833 | 44 |
| 986 | 3877 | 3921 | 3965 | 4009 | 4053 | 4097 | 4141 | 4185 | 4229 | 4273 | 44 |
| 987 | 4317 | 4361 | 4405 | 4449 | 4493 | 4537 | 4581 | 4625 | 466 | 4713 | 44 |
| 988 | 4757 | 4801 | 4845 | 4889 | 4933 | 4977 | 5021 | 5065 | 5108 | 5152 | 44 |
| 989 | 5196 | 5240 | 52 | 5328 | 5372 | 5416 | 5460 | 5504 | 554 | 5591 | 44 |
| 990 | 5635 | 5679 | 5723 | 5767 | 5811 | 5854 | 5898 | 5942 | 5986 | 6030 | 44 |
| 991 | 6074 | 6117 | 6161 | 6205 | 6249 | 6293 | 6337 | 6380 | 6424 | 6468 | 44 |
| 992 | 6512 | 6555 | 6509 | 6643 | 6687 | 6731 | 6774 | 6818 | 6862 | 6906 | 44 |
| 993 | 6949 | 6993 | 7037 | 7080 | 7124 | 7168 | 7212 | 7255 | 7299 | 7343 | 44 |
| 994 | 7356 | 7430 | 7474 | 7517 | 7561 | 7605 | 7648 | 7692 | 7736 | 7779 | 44 |
| 995 | 7823 | 7867 |  | 7954 | 7998 | 80.11 | 8085 | 8129 | 8172 | 8216 | 44 |
| 996 | 8259 | 8303 | 8347 | 8390 | 8434 | 8477 | 8521 | 8564 | S608 | 8652 | 44 |
|  | 8695 | 8739 | 8782 | 8826 | 8869 | 8913 | 8956 | 9000 | 9043 | 9087 | 44 |
| 998 | 9131 | 9174 | 9218 | 9261 | 9305 | 9348 | 9392 | 9435 | 9479 | 9522 | 44 |
| 999 | y 365 | 9609 | 9652 | 9696 | 9739 | 9783 | 9826 | 9570 | 9913 | 9957 | 43 |
| N. | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | S | 9 | D. |

## TABLEIL.

## IOGARITHMIC SINES AND TANGENCS.

FOR<br>EVERY DEGREE AND MINUTE OF THE QUADRAN'T.

If the logarithms of the values in Table III. be each increased by 10, the results will be the values of this table.
The logarithmic Secants and Cosecants are not given. They may be readily obtained, as follows:-Subtract the logarithmic Cosine from 20, and the remainder will be the logarithmic Secant; subtract the logarithmic Sire from 20, and the reanainder will be the logarithmic Cosecant.

| 18 |  | LO¢ARITHMIC SINES, |  |  | TANGENTS, ETC. |  | Table II. |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $0^{\circ}$ |  |  |  |  |  |  |  | $19^{\circ}$ |
| , | Sine. | D. | Cosine. | D. | Tang. | D. | Cotang. | 1 |
| 0 | Inf. Neg. |  | 10.000000 |  | Inf. Neg. |  | Infinite. | 60 |
| 1 | $6 \cdot 463726$ | 501717 | 000000 | 00 | 6.463726 | 501717 | 13.536274 | 59 |
| 2 | 764756 | 293485 | 000000 | 00 | 764756 | 293483 | 235244 | 58 |
| 3 | 940847 | 208231 | 000000 | 00 | 940847 | 208231 | $0^{059153}$ | 57 |
| 4 | 7.065786 | 161517 | 000000 | 00 | 7-065786 | 161517 | 12.934214 | 56 |
| 5 | 162696 | 131968 | 000000 | 00 | 162696 | 131969 | 837304 | 55 |
| 6 | 241877 | 111575 | 9-999999 | 01 | 241878 | 111578 | 758122 | 54 |
| 7 | 308824 | 96653 | 999999 | 01 | 308825 | 99653 | 691175 | 53 |
| 8 | 366816 | 85254 | 999999 | 01 | 366817 | 85254 | 633183 | 52 |
| 9 | 417968 | 76263 | 999999 | OI | 417970 | 76263 | 582030 | 51 |
| 10 | 463726 | 68988 | 999993 | OI | 463727 | 68988 | $53627^{3}$ | 50 |
| 11 | 7-505118 | 62981 | 9.999998 | or | 7.505120 | 62981 | 12.494880 | 49 |
| 12 | 542906 | 57936 | 999997 | OI | 542909 | 57933 | 457091 | 48 |
| 13 | 577668 | 53641 | 392997 | OI | $5777^{5} 2$ | 53642 | 422328 | 47 |
| 14 | 609853 | 49938 | 997996 | $0:$ | 609857 | 49939 | 390143 | 46 |
| 15 | 639816 | 46714 | 999996 | OI | 639820 | 46715 | 360180 | 45 |
| 16 | 667845 | 43881 | 999995 | 01 | 667849 | 43882 | 332151 | 44 |
| 17 | 694173 | 41372 | 999995 | OI | 694179 | 41373 | 305821 | 43 |
| 18 | 718997 | 39135 | 999994 | O1 | 719003 | 39136 | 280997 | 42 |
| 19 | $74247^{8}$ | 37127 | 999993 | 01 | 742484 | 37128 | 2505516 | 41 |
| 20 | 764754 | 35315 | $99999^{3}$ | OI | 764761 | 35136 | 235239 | 40 |
| 21 | 7-785943 | 33672 | 9-999992 | 01 | 7-785951 | 33673 | 12.214049 | 39 |
| 22 | 806146 | 32175 | 999991 | OI | 806155 | 32176 | 193845 | 38 |
| 23 | 825451 | 30805 | 999990 | 01 | 825460 | 30806 | 174540 | 37 |
| 24 | 843934 | 29547 | 999989 | 02 | 843944 | 29549 | 156056 | 36 |
| 25 | 861662 | 28388 | 999989 | 02 | 861674 | 28390 | 138326 | 35 |
| 26 | 878695 | 27317 | 999988 | 02 | 878708 | 27318 | 121292 | 34 |
| 27 | 895085 | 26323 | 999987 | 02 | 895099 | 26325 | 104901 | 33 |
| 28 | 910879 | 25399 | 999986 | 02 | 910894 | 25401 | 089106 | 32 |
| 29 | 926119 | 24538 | 999985 | 02 | 926134 | 24540 | 073866 | 31 |
| 30 | 940842 | 23733 | 999983 | 02 | 940858 | 23735 | 059142 | 30 |
| 31 | $7 \cdot 955082$ | 22980 | 9.999982 | 02 | $7 \cdot 955100$ | 22981 | 12.044900 |  |
| 32 | 968870 | 22273 | 999981 | 02 | 968889 | 22275 | 031111 | 28 |
| 33 | 982233 | 21608 | 999980 | 02 | 982253 | 21610 | 017747 | 27 |
| 34 | 995198 | 20981 | 999979 | 02 | 995219 | 20983 | 004781 | 26 |
| 35 | $8 \cdot 007787$ | 20300 | 999977 | 02 | 8-007809 | 20392 | 11.992: 21 | 25 |
| 36 | 020021 | 19831 | 999976 | 02 | 020044 | 19833 | 979956 | 24 |
| 37 | 031919 | 19302 | 999975 | 02 | 031945 | 19305 | 968055 | 23 |
| 38 | 043501 | 18801 | 999973 | 02 | 043527 | 18803 | 956473 | 22 |
| 39 | 054781 | 18325 | 999972 | 02 | 054809 | 18327 | 945191 | 21 |
| 4 C | 065776 | 17872 | 999971 | 02 | 065806 | 17874 | 934194 | 20 |
| 41 | 8.076500 | 17441 | 9.999969 | 02 | 8.076531 | 17444 | 11.923469 | 10 |
| 42 | 086965 | 17031 | 999968 | 02 | 086997 | 17034 | 913003 | 18 |
| 43 | 097183 | 16639 | 999966 | 02 | 097217 | 16642 | 902783 | :7 |
| 44 | 107167 | 16265 | 999964 | o3 | 107203 | 16268 | 892797 | 16 |
| 45 | 116926 | 15908 | 999963 | o3 | 116963 | 15910 | 883037 | 15 |
| 46 | 126471 | 15566 | 999961 | -3 | 126510 | 15568 | 873490 | 14 |
| 47 | 135810 | 15238 | 999959 | o3 | 135851 | 15241 | 864149 | 13 |
| 48 | i44953 | 14924 | 999958 | o3 | 144996 | 14927 | 855004 | 12 |
| 49 | 153907 | 14622 | 999956 | o3 | $153 y 52$ | ${ }^{14627}$ | 846048 | 11 |
| 50 | 162681 | 14333 | 999954 | o3 | 162727 | 14336 | 837273 | 10 |
| 51 | 8.171280 | 14054 | 9.999952 | -3 | 8.171328 | 14057 | 11.828672 | 8 |
| 52 | 179713 | 13786 | 999950 | o3 | 179763 | 13790 | 820237 | 8 |
| 53 | 187985 | 13529 | 999948 | o3 | 185036 | 13532 | 811064 | 7 |
| 54 | 196102 | 13280 | 999946 | o3 | 196156 | 13284 | 803844 | 6 |
| 55 | 204070 | 13041 | 999944 | o3 | 204126 | 13044 | 795874 | 5 |
| 56 | 211885 | 12810 | 999942 | 04 | 211953 | 12814 | 788047 | 4 |
| 57 | 219581 | 12587 | 999940 | 04 | 219641 | 12590 | 750359 | 3 |
| 58 | 227134 | 12372 | 999938 | 04 | 227195 | 12376 | 772805 | 2 |
| 59 | 234557 | 12164 | 999936 | 04 | 23.4621 | 12168 | 765379 | 1 |
| 60 | 241855 | 11963 | 999934 | 04 | 241921 | 11967 | 758079 |  |
| , | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | , |
| $90^{\circ}$ |  |  |  |  |  |  |  | $89^{\circ}$ |


| Table II. |  | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  |  | 18 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 |  |  |  |  |  |  | $178^{\circ}$ |  |
|  | Sine. | D. | Cosine. | D. | Tang. | D. | Cotang. | ' |
| 0 | 8.241855 | 11963 | 9.999934 | 04 | 8-241921 | 11967 | $11 \cdot 758079$ | 60 |
| 1 | 249033 | 11768 | 999932 | 04 | 249102 | 11772 | 750898 | 59 |
| 2 | 256094 | 11580 | 999929 | 04 | 256165 | 11584 | 743835 | 58 |
| 3 | 263042 | 11398 | 999927 | 04 | 263115 | 11402 | 736885 | 57 |
| 4 | 269881 | 11221 | 999925 | 04 | 269956 | 11225 | 730044 | 56 |
| 5 | 276614 | 11050 | 999922 | 04 | 276691 | 11054 | 723309 | 55 |
| 6 | 283243 | 10883 | 999920 | 04 | 283323 | 10887 | 716677 | 54 |
| 7 | 289773 | 10721 | 999918 | 04 | 289856 | 10726 | 710144 | 53 |
| 8 | 296207 | 10565 | 999915 | 04 | 296292 | 10570 | 703708 | 52 |
| 9 | 302546 | 10413 | 999913 | 04 | 302634 | 10418 | 697366 | 51 |
| 10 | 308794 | 10266 | 999910 | 04 | 308884 | 10270 | 691116 | 50 |
| 11 | 8-314954 | 10122 | 9•999907 | 04 | 8.315046 | 10126 | 11.684954. | 49 |
| 12 | 321027 | 9982 | 999905 | 04 | 321122 | 9987 | 678878 | 48 |
| 13 | 327016 | 9847 | 999902 | 04 | 327114 | 9851 | 672886 | 47 |
| 14 | 332924 | 9714 | 999899 | o5 | 333025 | 9719 | 666975 | 46 |
| 15 | 338753 | 9586 | 999897 | -5 | 338856 | 9590 | 661144 | 45 |
| 16 | 344504 | 9460 | 999894 | o5 | 344610 | 9465 | 655390 | 44 |
| 17 | 350181 | 9338 | 999891 | o5 | 350289 | 9343 | 649711 | 43 |
| 18 | 355783 | 9219 | 999888 | -5 | 355805 | 9224 | 644105 | 42 |
| 19 | 361315 | 9103 | 999885 | o5 | 361430 | 9108 | 638570 | 41 |
| 20 | 366777 | 8990 | 999882 | o5 | 366895 | 8995 | 633105 | 40 |
| 21. | $8 \cdot 372171$ | 8880 | 9-999879 | o5 | $8 \cdot 372292$ | 8885 | 11.627708 | 39 |
| 22 | 377499 | 8772 | 999876 | o5 | 377622 | 8777 | 622378 | 38 |
| 23 | 382762 | 8667 | 999873 | o5 | 382889 | 8672 | 617111 | 37 |
| 24 | 387962 | 8564 | 999870 | 05 | 388002 | 8570 | 611908 | 36 |
| 25 | 393101 | 8464 | 999867 | o5 | 393234 | 8470 | 606766 | 35 |
| 26 | 398179 | 8366 | 999864 | o5 | 398315 | 8371 | 601685 | 34 |
| 27 | 403199 | 8271 | 999861 | o5 | 403338 | 8276 | 596662 | 33 |
| 28 | 408161 | 8177 | 999858 | o5 | 408304 | 8182 | 591696 | 32 |
| ${ }_{2} 9$ | 413068 | 8086 | 999854 | o5 | 413213 | 8091 | 586787 | 31 |
| 30 | 417919 | 7996 | 999851 | o6 | 418068 | 8002 | 581932 | 30 |
| 31 | $8 \cdot 422717$ | 7909 | 9.999848 | o6 | 8-4.22869 | 7914 | 11.577131 | 29 |
| 32 | 427462 | 7823 | 999844 | 06 | 427618 | 7830 | 572382 | 28 |
| 33 | 432156 | 7740 | 999841 | o6 | 432315 | 7745 | 567685 | 27 |
| 34 | 436800 | 7657 | 999838 | o6 | 436962 | 7663 | 563038 | 26 |
| 35 | 441394 | 7577 | 999834 | o6 | 441560 | 7583 | 558440 | 25 |
| 36 | 445941 | 7499 | 99983 I | o6 | 446110 | 7505 | 553890 | 24 |
| 37 | 450440 | 7422 | 999827 | o6 | 450613 | 7428 | 549387 | 23 |
| 38 | 454893 | 7346 | 999824 | o6 | 455070 | 7352 | 544930 | 22 |
| 39 | 459301 | 7273 | 999820 | o6 | 459481 | 7279 | 540519 | 21 |
| 40 | 463665 | 7200 | 999816 | 06 | 463849 | 7206 | 536151 | 20 |
| 41 | 8-467985 | 7129 | 9.999813 | o6 | 8.468172 | 7135 | 11.531828 | 19 |
| 42 | 472263 | 7060 | 999809 | o6 | 472454 | 7066 | 527546 | 18 |
| 43 | 476498 | 6991 | 999805 | 06 | 476693 | 6998 | 523307 | 17 |
| 44 | 480693 | 6924 | 999801 | 06 | $48089^{2}$ | $6{ }^{31}$ | 5 I 9108 | 16 |
| 45 | 484848 | 6859 | 999797 | 07 | 485050 | 6865 | 514950 | 15 |
| 40 | 488963 | 6794 | 999794 | 07 | 489170 | 6801 | 510830 | 14 |
| 47 | 493040 | 6731 | 999790 | 07 | 493250 | 6738 | 506750 | 13 |
| 43 | 497078 | 6669 | 999786 | 07 | 497293 | 6676 | 502707 | 12 |
| 49 | 501080 | 6608 | 999782 | 07 | 501278 | 6615 | 498702 | 11 |
| 50 | 505045 | 6548 | 999778 | 07 | 501267 | 6555 | 494733 | 10 |
| 51 | 8.508974 | 6489 | 7-999774 | 07 | 8.509200 | 6496 | 11.490800 |  |
| 52 | 512867 | 6431 | 999769 | 07 | 513098 | 6439 | 486902 | 8 |
| 53 | 516726 | 6375 | 999765 | 07 | 516961 | 6382 | 483039 | 7 |
| 54 | 520551 | 6319 | 999761 | 07 | 520790 | 6326 | 479210 | 6 |
| 55 | 524343 | 6264 | 99975 | 07 | 524586 | 6272 | 475414 | 5 |
| 56 | 528102 | 6211 | 999753 | 07 | 528349 | 6218 | 471651 | 4 |
| 57 | 531828 | 6158 | 999748 | 07 | 532080 | 6165 | 467920 | 3 |
| 58 | 535523 | 6106 | 999744 | 07 | 535779 | 6113 | 464221 | 2 |
| 59 | 539186 | 6055 | 999740 | 07 | 539447 | 6062 | 460553 | 1 |
| 60 | 542819 | 6004 | 999735 | 07 | 543084 | 6012 | 456916 | 0 |
| , | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | , |
| $91^{\circ}$ |  |  |  |  |  |  |  | $88^{\circ}$ |


| 20 | LGđARITHMIC SINES, TANGENTS, ETC. |  |  |  |  |  | TABI | 11. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2^{3}$ |  | D. | Cosine. | D. | Tang. | D. | $177^{\circ}$ |  |
| , | Sine. |  |  |  |  |  | Cotang. | , |
| 0 | 8.542819 | 6004 | 9.999735 | 07 | 8.543084 | 6012 | 11.456916 | 60 |
| 1 | 546422 | 5955 | 999731 | 07 | 546691 | 5962 | 453309 | 59 |
| 2 | 549995 | 5906 | 999726 | 07 | 550268 | 5914 | 449732 | 58 |
| 3 | 550339 | 5858 | 999722 | 08 | 553817 | 5866 | 446183 | 57 |
| 4 | 557054 | 58 II | 999717 | 08 | 557336 | 5819 | 442664 | 56 |
| 5 | 560540 | 5765 | 099713 | 08 | 560828 | 5773 | 439172 | 55 |
| 6 | 563999 | 5719 | 999708 | $\bigcirc 8$ | 564291 | 5727 | 435709 | 54 |
| 7 | 56743 I | 5674 | 999704 | 08 | 567727 | 5682 | 432273 | 53 |
| 8 | 570836 | 5630 | 999699 | 08 | 571137 | 5638 | 428863 | 52 |
| 9 | 574214 | 5587 | 999694 | 08 08 | 574520 | 5595 | 425480 | 51 |
| 10 | 577566 | 5544 | 999689 | 08 | 577877 | 5502 | 422123 | 50 |
| 11 | 8.580892 | 5502 | 9.999685 | 08 | 8.581208 | 5510 | 11.418792 | 49 |
| 12 | 584193 | 5460 | 999680 | 08 | 584514 | 5468 | 415486 | 48 |
| 13 | 587469 | 5419 | 999675 | 08 | 587795 | 5427 | 412205 | 47 |
| 14 | 590721 | 5379 | 999670 | 08 | 591051 | 5.387 | 408949 | 46 |
| 15 | 593948 | 5339 | 999665 | 08 | 594283 | 5347 | 405717 | 45 |
| 16 | 597152 | 5300 | 999660 | 08 | 597492 | 5308 | 402508 | 44 |
| 17 | 600332 | 5261 | 999655 | 08 | 600677 | 5270 | 399323 | 43 |
| 18 | 603489 606623 | 5223 | 999650 | 08 | 603839 | 5232 | 396161 | 42 |
| 19 | 606623 6097 | 5186 | 999645 | 09 | 606978 | 5194 | 3 l 3022 | 41 |
| 20 | 609734 | 5149 | 999640 | 09 | 610094 | 5158 | 389906 | 40 |
| 21 | 8.612823 | 5112 | 9.999635 | 09 | 8.613189 | 5 I 21 | II $\cdot 3868 \mathrm{II}$ | 39 |
| 22 | 615891 | 5076 | 999629 | 09 | 616262 | 5085 | 383738 | 38 |
| 23 | 618937 | 5041 | 999624 | 09 | 619313 | 5050 | 380687 | 37 |
| 24 | 621962 | 5006 | 999619 | 09 | 622343 | 5015 | 377657 | 36 |
| 25 | 624965 | 4972 | 999614 | $\bigcirc$ | 625352 | 4981 | 374648 | 35 |
| 26 | 627948 | 4938 | 999608 | 09 | 628340 | 4947 | 371660 | 34 |
| 27 | 630911 | 4904 | 999603 | 9 | 631308 | 4913 | 368692 | 33 |
| 28 | 633854 | 4871 | 999597 | 09 | 634256 | 4880 | 365744 | 32 |
| 29 | 636776 | 4839 | 999502 | 09 | 637184 | 4848 | 362816 | 31 |
| 30 | 639680 | 4806 | 999586 | 09 | 640093 | 4816 | 359907 | 30 |
| 3 I | 8.642563 | 4775 | 9.999581 | 09 | 8.642982 | 4784 | 11.357018 | 29 |
| 32 | 645428 | 4743 | 999575 | 09 | 645053 | 4753 | 354147 | 28 |
| 33 | 648274 | 4712 | 999570 | 09 | 648704 | 4722 | 351296 | 27 |
| 34 | 651102 | 4682 | 999564 | 09 | 6515037 | 4691 | 348463 | 26 |
| 35 | 653911 | 4652 | 999558 | 10 | 654352 | 4661 | 345648 | 25 |
| 36 | 656702 | 4622 | 999503 | 10 | 657149 | 4631 | 3.42851 | 24 |
| 37 | 659475 | 40.92 | 999547 | 10 | 659928 | 4602 | 3.0072 | 23 |
| 38 | 662230 | 4563 | 999541 | 10 | 662689 | 4573 | 337311 | 22 |
| 39 | 664968 | 4535 | 999535 | 10 | 665433 | 4544 | 334567 | 21 |
| 40 | 667689 | 4506 | 999529 | 10 | 668160 | 4526 | 3318¢0 | 20 |
| 41 | $8 \cdot 670393$ | 4479 | 9.999524 | 10 | 8.670870 | 4488 | 11.329130 | 19 |
| 42 | 673030 | 4451 | 999518 | 10 | 673563 | 4461 | 326437 | 18 |
| 43 | 675751 | 4424 | 999512 | 10 | 676239 | 4434 | 323761 | 17 |
| 44 | 678405 | 4397 | 999506 | 10 | 678900 | 4417 | 321100 | 16 |
| 45 | 681043 | 4370 | 999500 | 10 | 681544 | 4380 | 318456 | 15 |
| 46 | 683665 | 4344 | $99949^{3}$ | 10 | 684172 | 4354 | 315828 | 14 |
| 47 | 686272 | 4318 | 999437 | Io | 686784 | 4328 | 313216 | 13 |
| 48 | 688863 | 4202 | 999481 | 10 | 689381 | 4303 | 310619 | 12 |
| 49 | 691438 | 4267 | 999475 | 10 | 691963 | 4277 | 308037 | 11 |
| 50 | 693999 | $42 ¢ 2$ | 999469 | 10 | 694029 | 4252 | 305471 | 10 |
| 5 I | 8.696543 | 4217 | 9.999463 | II | 8.697081 | 4228 | 11.302919 |  |
| 52 | 699073 | 4192 | 999456 | 11 | 699617 | 4203 | 300383 | 8 |
| 54 | 701509 | 4184 | 999400 | II <br> II | 702139 | 4170 | 297561 | 7 |
| 55 | 706577 | 4121 | 999437 | 11 | 707140 | 4132 | 292860 | 5 |
| 56 | 709049 | 4097 | 999431 | II | 709618 | 4108 | 290382 | 4 |
| 57 | 711507 | 4074 | 999424 | 11 | 712083 | 4085 | 257917 | 3 |
| 58 | 713052 | 4051 | 999418 | II | -14534 | 4062 | 255466 | 2 |
| 59 | 716383 | 4029 | 999411 | 11 | 716972 | 4040 | 283028 | 1 |
| 60 | 718803 | 4006 | 999404 | 11 | 719396 | 4017 | 280604 | 0 |
| 1 | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | , |
| $92^{\circ}$ |  |  |  |  |  |  |  | $87^{\circ}$ |


| Table II. |  | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  |  | 21 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $3^{\circ}$ |  |  |  |  |  |  |  | $176{ }^{\circ}$ |
| , | Sine. | D. | Cosine. | D. | Tang. | D. | Cotang. | , |
| 0 | 8.718800 | 4006 | 9-9994c4 | 11 | 3.719396 | 4017 | 11-280604 | 60 |
| 1 | 721204 | 3984 | 999398 | 11 | 721806 | 3995 | 278194 | 59 |
| 2 | 723595 | 3962 | 999391 | 11 | 724204 | 3974 | 275796 | 58 |
| 3 | 725972 | 3941 | 999384 | 11 | 726588 | 3952 | 273412 | 57 |
| 4 | 728337 | 3919 | 999378 | 11 | 728059 | 3930 | 271041 | 56 |
| 5 | 730688 | 3898 | 999371 | 11 | 731317 | 3909 | 268683 | 55 |
| 6 | 733027 | 3877 | 999364 | 12 | 733663 | 3889 | 266337 | 54 |
| 7 | 735354 | 3857 | 999357 | 12 | 735996 | 3868 | 264004 | 53 |
| 8 | 737667 | 3836 | 999350 | 12 | 738317 | 3848 | 261683 | 52 |
| 9 | 739969 | 3816 | 999343 | 12 | 740626 | 3827 | 259374 | 51 |
| 10 | 742259 | 3796 | 999336 | 12 | 742922 | 3807 | 257078 | 50 |
| 11 | $8 \cdot 744536$ | 3776 | 5.999 229 | 12 | 3.745207 | 3787 | 15.254793 | 49 |
| 12 | 746802 | 3756 | 999322 | 12 | 747479 | 3768 | 252521 | 48 |
| 13 | 749055 | 3737 | 999315 | 12 | 749740 | 3749 | 250260 | 47 |
| 14 | 751297 | 3717 | 999308 | 12 | 751989 | 3729 | 248011 | 46 |
| 15 | 753528 | 3698 | 999301 | 12 | 754227 | 3710 | 245773 | 45 |
| 16 | 755747 | 3679 | 999294 | 12 | 756453 | 3692 | 243547 | 44 |
| 17 | 757955 | 3661 | 999287 | 12 | 758668 | 3673 | 241332 | 43 |
| 18 | 760151 | 3642 | 999279 | 12 | 760872 | 3655 | 239128 | 42 |
| 19 | 762337 | 3624 | 999272 | 12 | 763065 | 3636 | 236935 | 41 |
| 20 | 764511 | 3606 | 99926 | 12 | 765246 | 3618 | 234754 | 40 |
| 21 | $8 \cdot 766675$ | 3588 | 9.999257 | 12 | $8 \cdot 767417$ | 3600 | 11-232583 | 39 |
| 22 | 768828 | 3570 | 999250 | 13 | 769578 | 3583 | 230422 | 38 |
| 23 | 770970 | 3553 | 999242 | 13 | 771727 | 3565 | 228273 | 37 |
| 24 | 773101 | 3535 | 999235 | 13 | 773866 | 3548 | 226134 | 36 |
| 25 | 775223 | 3518 | 999227 | 13 | 775995 | 3531 | 224005 | 35 |
| 26 | 777333 | 3501 | 999220 | 13 | 778114 | 3514 | 221886 | 34 |
| ? 7 | 779434 | 3484 | 999212 | 13 | 780222 | 3447 | 219778 | 33 |
| 28 | 781524 | 3467 | 999205 | 13 | 782320 | 3480 | 217680 | 32 |
| 29 | 783605 | 3451 | 999197 | 13 | 784408 | 3464 | 215592 | 31 |
| 30 | 785675 | 3431 | 999189 | 13 | 786486 | 3447 | 213514 | 30 |
| 31 | $8 \cdot 787736$ | 3418 | 9.999181 | 13 | 8.788554 | 3431 | 11-211446 | 29 |
| 32 | 789787 | 3402 | 999174 | 13 | 790613 | 3414 | 209387 | 28 |
| 33 | 791828 | 3386 | 999166 | 13 | 792662 | 3399 | 207338 | 27 |
| 3.4 | 793859 | 3370 | 999158 | 13 | 794701 | 3383 | 205299 | 26 |
| 35 | 795881 | 3354 | 999150 | 13 | 796731 | 3368 | 203269 | 25 |
| 36 | 797894 | 3339 | 999142 | 13 | 798752 | 3352 | 201248 | 24 |
| 37 | 799897 | 3323 | 999134 | 13 | 800763 | 3337 | 199237 | 23 |
| 38 | 801892 | 3308 | 999126 | 13 | 802765 | 3322 | 197235 | 22 |
| 39 | 803876 | 3293 | 99918 | 13 | 804758 | 3307 | 195242 | 21 |
| 40 | 805852 | 3278 | 999110 | 13 | 806742 | 3292 | 193258 | 20 |
| 41 | 8.807819 | 3263 | 9.999102 | 13 | 8.808717 | 3278 | H-191283 | 19 |
| 42 | 809777 | 3249 | 999094 | 14 | 810683 | 3262 | 189317 | 18 |
| 43 | 811726 | 3234 | 999086 | 14 | 812641 | 3248 | 187359 | 17 |
| 44 | 813667 | 3219 | 999077 | 14 | 814589 | 3233 | 185411 | 16 |
| 45 | 815599 | 3205 | 999069 | 14 | 816529 | 3219 | 183471 | 15 |
| 46 | 817522 | 3191 | 999061 | 14 | 818461 | 3205 | 181539 | 14 |
| 47 | 819436 | 3177 | 999053 | 14 | 820384 | $3 \mathrm{Ig1}$ | 179616 | 13 |
| 48 | 821343 | 3163 | 999044 | 14 | 822298 | 3177 | 177702 | 12 |
| 49 | 823240 | 3142 | 999036 | 14 | 824205 | 3163 | 175795 | 11 |
| 50 | 825130 | 3135 | 999027 | 14 | 826103 | 3150 | 173897 | 10 |
| 51 | 8.827011 | 3122 | 9.999019 | 14 | 8.827992 | 3136 | 11.172008 | 8 |
| 52 | 828884 | 3108 | 999010 | 14 | 829874 | 3123 | 170126 | 8 |
| 53 | 830749 | 3095 | 999002 | 14 | 831748 | 3110 | 168252 | 7 |
| 54 | 832607 | 3082 | 998993 | 14 | 833613 | 3096 | 166387 | 6 |
| 55 | 834456 | 3069 | 998984 | 14 | 835471 | 3083 | 164529 | 5 |
| 56 | 836297 | 3056 | 998976 | 14 | 837321 | 3070 | 162679 | 4 |
| 57 58 | 838130 | 3043 | 998967 | 15 | 839163 | 3057 | 160837 |  |
| 58 | 839956 | 3030 | 998958 | 15 | 840998 | 3045 | 159002 | 2 |
| 59 60 | 841774 843585 | 3017 | 998950 | 15 | 842825 | 3032 | 157175 | 1 |
| 60 | 843585 | 3oco | 998941 | 15 | 844644 | 3019 | 155356 | 0 |
| , | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | , |
| 930 |  |  |  |  |  |  |  | $86^{\circ}$ |


| 22 | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  |  | TABL | II. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $4{ }^{\circ}$ | Sine. | D. | Cosine | D. | Tang. | D. | $175^{\circ}$ |  |
|  |  |  |  |  |  |  | Cotang. | 1 |
| 0 | 8.843585 | 3005 | 9.998941 | 15 | 8.844644 | 3019 | II - 155356 | 60 |
| 1 | 845387 | 2992 | 798932 | 15 | 846450 | 3007 | 153545 | 59 |
| 2 | 847183 | 2980 | 998923 | 15 | 848260 | 2995 | 151740 | 58 |
| 3 | 848971 | 2967 | 998914 | 15 | 850057 | 2982 | 149943 | 57 |
| 4 | 850751 | 2955 | 998005 | 15 | 851846 | 2970 | 148154 | 56 |
| 5 | 852525 | 2943 | 998806 | 15 | 8503628 | 2958 | 146372 | 55 |
| 6 | 854291 | 2931 | 998887 | 15 | 855403 | 2946 | 144597 | 54 |
| 7 | 856049 | 2919 | 998878 | 15 | 857171 | 2935 | 142829 | 53 |
| 8 | 857801 | 2907 | 998869 | 15 | 858932 | 2923 | 141068 | 52 |
| 9 | - 859546 | 2896 | 998860 | $: 5$ | 860686 | 2911 | 139314 | 51 |
| 10 | 861283 | 2884 | 99885 I | 15 | 862433 | 2900 | 137567 | 50 |
| 11 | 8.863014 | 2873 | 9.99884I | 15 | 8.864173 | 2888 | 11.135827 | 49 |
| 12 | 864738 | 2861 | 998832 | 15 | 865006 | 2877 | 134094 | 48 |
| 13 | 866455 | 2850 | 998823 | 16 | 867632 | 2806 | 132368 | 47 |
| 14 | 868165 | 2839 | 998813 | 16 | 869351 | 2854 | 130649 | 46 |
| 15 | 869868 | 2828 | 998804 | 16 | 871064 | 2843 | 128936 | 45 |
| 16 | 871565 | 2817 | 998705 | 16 | 872770 | 2832 | 127230 | 44 |
| 17 | 873255 | 2806 | 998785 | 16 | 874469 | 2821 | 125531 | 43 |
| 18 | 874938 | 2705 | 998776 | 16 | 876162 | 2811 | 123838 | 42 |
| 19 | 876615 | 2786 | 998766 | 16 | 877849 | 2800 | 122151 | 41 |
| 20 | 878285 | 2773 | 998757 | 16 | 879529 | 2789 | 120471 | 40 |
| 21 | 8.879949 | 2763 | 9.998747 | 16 | 8.881202 | 2779 | 11.118798 | $\begin{aligned} & 39 \\ & 38 \end{aligned}$ |
| 22 | $\begin{aligned} & 881607 \\ & 883258 \end{aligned}$ | 2752 | 998738 | 16 | 882869 | 2768 | 117131 |  |
| 23 |  | 2742 | 998728 | 16 | 884530 | 2758 | 115470 | 37 |
| 24 | 884903886542 | 2731 | 998718 | 16 | 886185 | 2747 | 113815 | 36 |
| 25 |  | 2721 | 998708 | 16 | 887833 | 2737 | 112167 | 35 |
| 26 | $\begin{aligned} & 886542 \\ & 888174 \end{aligned}$ | 2711 | $99869^{9}$ | 16 | 889476 | 2727 | 1110524 | 34 |
| 27 | 888174 <br> 889801 | 2700 | 998689 | 16 | 891112 | 2717 | 108888 | 33 |
| 28 | $\begin{aligned} & 891421 \\ & 893035 \end{aligned}$ | 2690 | 998679 | 16 | 892742 | 2707 | 107258 | 32 |
| 29 |  | 2680 | 998669 | 17 | 894366 | 2697 | 105634 | 31 |
| 30 | $\begin{aligned} & 893035 \\ & 894643 \end{aligned}$ | 2670 | 998659 | 17 | 895084 | 2687 | 104016 | 30 |
| 31 | 8.896246 | 2660 | 9.998649 | 17 | 8.897596 | 2677 | 11102404 | 29 |
| 32 | 897842 | 2651 | 998639 | 17 | 899203 | 2667 | 100797 | 28 |
| 33 | 899432 | 2641 | 998629 | 17 | 900803 | 2658 | 099197 | <7 |
| 34 | 901017 | 2631 | 998619 | 17 | 902398 | 2648 | 097602 | 26 |
| 3,5 | 902596 | 2622 | 998609 | 17 | 903987 | 2638 | 096013 | 25 |
| 36 | 904169 | 2612 | 99859 | 17 | 905570 | 2629 | 094430 | 24 |
| 37 | 905736 | 2603 | 99850 | 17 | 907147 | 2620 | 092853 | 23 |
| 38 | 907297 | 2503 | 993578 | 17 | 908719 | 2610 | 0 O 1281 | 22 |
| 39 | 908853 | 2584 | 998506 | 17 | 910285 | 2601 | 089715 | 21 |
| 40 | 910404 | 2575 | 998558 | 17 | 911846 | 2592 | 088154 | 20 |
| 41 | 8.911949 | 2566 | 9.998548 | 17 | 8.913401 | 2583 | 11.086599 | 19 |
| 42 | 913488 | 2556 | 998537 | 17 | 914951 | 2574 | 0850 29 | 18 |
| 43 | 915022 | 2547 | 998527 | 17 | 916405 | 2505 | -83505 | 17 |
| 44 | 916550 | 2538 | 998516 | 18 | 918034 | 2556 | 081966 | 16 |
| 45 | 918073 | 2529 | 998506 | 18 | 919568 | 2547 | 080432 | 15 |
| 46 | 919591 | 2520 | 998405 | 18 | 921096 | 2539 | 078904 | 14 |
| 47 | 921103 | 2512 | 998485 | 18 | 922619 | 2530 | 077381 | 13 |
| 48 | 922610 | 2503 | 998474 | 18 | 924136 | 25.21 | 075864 | 12 |
| 49 | 924112 | 2494 | 99846 | 18 | 925649 | 2512 | 074351 | 11 |
| 50 | 925009 | 2486 | 998453 | 18 | 927156 | 2503 | 072844 | 10 |
| 51 | 8.927100 | 2477 | 9.998442 | 18 | 8.928658 | 2495 | 11.0713 ¢2 | $\begin{aligned} & 3 \\ & 3 \\ & 7 \\ & 6 \\ & 5 \\ & 4 \\ & 2 \\ & 1 \\ & 0 \end{aligned}$ |
| 52 | 928587 | 2469 | 99843 I | 18 | $930 \cdot 55$ | 2486 | 069845 |  |
| 53 | 930068 | 2460 | 998421 | 18 | 931647 | 2478 | 068353 |  |
| 54 | 931544 | 2452 | 998410 | 18 | 933134 | 2470 | 066866 |  |
| 55 | 933015 | 2443 | 998309 | 18 | 934616 | 2461 | 065384 |  |
| 56 | 934481 | 2435 | 998388 | 18 | 936003 | 2453 | 063907 |  |
| 57 | 935942 | 2427 | 998377 | 18 | 9375565 | 2445 | 062435 |  |
| 58 | 937398 | 2419 | 998366 | 18 | 939032 | 2437 | 060968 |  |
| 59 | 938850 | 2411 | 998355 | 18 | 940424 | 2430 | 050506 |  |
| 60 | 940296 | 2403 | 998344 | 18 | 941952 | 2421 | 058048 |  |
| 1 | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | ' |
| $94^{\circ}$ |  |  |  |  |  |  |  | $85^{\circ}$ |


| T'able II. |  | LOGARITHMIC |  | SINES | ANGENTS, ETC |  |  | 23 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $5^{\circ}$ |  | D. | Cosine. | D. | Tang. | D. | $174^{\circ}$ |  |
| , | Sine. |  |  |  |  |  | Cotang. | , |
| 0 | $8 \cdot 940296$ | 2403 | 9.998344 | 19 | $8 \cdot 941952$ | 2421 | 11.058048 | 60 |
| 1 | 941738 | 2394 | 998333 | 19 | 943404 | 2413 | -56596 | 59 |
| 2 | 943174 | 2387 | 998322 | 19 | 944852 | 2405 | 055148 | 58 |
| 3 | 944606 | 2379 | 998311 | 19 | 946295 | 2397 | 053705 | 57 |
| 4 | 946034 | 2371 | 998300 | 19 | 947734 | 2390 | 052266 | 56 |
| 5 | 947456 | 2363 | 998289 | 15, | 949168 | 2382 | -50832 | 55 |
| 6 | 948874 | 2355 | 998277 | 19 | 950597 | 2374 | 049403 | 54 |
| 7 | 950287 | 2348 | 998266 | 19 | 95021 | 2366 | 047979 | 53 |
| 8 | 951696 | 2340 | 998255 | 19 | 953441 | 2360 | 046559 | 52 |
| 9 | 953100 | 2332 | 998243 | 19 | 954856 | 2351 | 045144 | 51 |
| 10 | 954499 | 2325 | 998232 | 19 | 956267 | 2344 | 043733 | 50 |
| 11 | 8.955894 | $2317$ | $9 \cdot 998220$ | 19 | 8.957674 | 2337 | 11.042326 | 49 |
| 12 | 957284 | 2310 | 998209 | 19 | 959075 | 2329 | 040925 | 48 |
| 13 | 958670 | 2302 | 998197 | 19 | 960473 | 2323 | o3q527 | 47 |
| 14 | 960052 | 2295 | 998186 | 19 | 961866 | 2314 | 038134 | 46 |
| 15 | 961429 | 2288 | 998174 | 19 | 963255 | 2307 | 036745 | 45 |
| 16 | 962801 | 2280 | 998163 | 19 | 964639 | 2300 | 035361 | 44 |
| 17 | 964170 | 2273 | 998151 | 19 | 966019 | 2293 | o33981 | 43 |
| 18 | 965534 | 2266 | 998139 | 20 | 967394 | 2286 | 032606 | 42 |
| 19 | 966893 | 2259 | 998128 | 20 | 968766 | 2279 | 031234 | 41 |
| 20 | 968249 | 2252 | 998116 | 20 | 970133 | 2271 | 029867 | 40 |
| 218.969600 |  | 2244 | 9.998104 | 20 | 8-971496 | 2265 | 11.028504 | 39 |
| 22 | 970947 | 2238 | $99^{80}{ }^{2}$ | 20 | 972855 | 2257 | 027145 | 38 |
| 23 | 972289 | 2231 | 998080 | 20 | 974209 | 2251 | 025791 | 37 |
| 24 | 973628 | 2224 | $99^{8068}$ | 20 | 975560 | 2244 | 024440 | 36 |
| 25 | 974962 | 2217 | 998056 | 20 | 976906 | 2237 | 023094 | 35 |
| 26 | 976293 | 2210 | 998044 | 20 | 978248 | 2230 | 021752 | 34 |
| 27 | 977619 | 2203 | $99^{8032}$ | 20 | 979586 | 2223 | 020414 | 33 |
| 28 | 978941 | 2197 | $99^{8020}$ | 20 | 980921 | 2217 | 019079 | 32 |
| 29 | 980259 | 2190 | 998008 | 20 | 982251 | 2210 | -17749 | 3 I |
| 30 | 981573 | 2183 | 997996 | 20 | 983577 | 2204 | 016423 | 30 |
| 3 I | 8.982883 | 2177 | 9.997984 | 20 | 8.984899 | 2197 | 11.015101 | 29 |
| 32 | 984189 | 2170 | 997972 | 20 | 986217 | 2191 | -13783 | 28 |
| 33 | 985491 | 2163 | 997959 | 20 | 987532 | 2184 | OI 2468 | 27 |
| 34 | 986789 | 2157 | 997947 | 20 | 988842 | 2178 | OIII58 | 26 |
| 35 | 988083 | 2150 | 997935 | 21 | 990149 | 2171 | 009851 | 25 |
| 36 | 989374 | 2144 | 997922 | 21 | 991451 | 2165 | 008549 | 24 |
| 37 | 990660 | 2138 | 997910 | 21 | 992750 | 2158 | 007250 | 23 |
| 38 | 991943 | 2131 | 997807 | 21 | 994045 | 2152 | 005955 | 22 |
| 39 | 993222 | 2125 | 997885 | 21 | 995337 | 2146 | 004663 | 21 |
| 40 | 994497 | 2119 | 997872 | 21 | 996624 | 2140 | 003376 | 20 |
| 41 | $8 \cdot 995768$ | 2112 | $9.997^{860}$ | 21 | 8.997908 | 2134 | $11 \cdot 002092$ | 19 |
| 42 | 997036 | 2106 | 997847 | 21 | 999188 | 2127 | 000812 | 18 |
| 43 | 998299 | 2100 | 997835 | 21 | 9-000465 | 2121 | $10 \cdot 999535$ |  |
| 44 | 999560 | 2094 | 997822 | 21 | 001738 | 2115 | 998262 | 15 |
| 45 | 9-000816 | 2087 | 997809 | 21 | 003007 | 2109 | 996993 | 15 |
| 46 | 002069 | 2082 | 997797 | 21 | 004272 | 2103 | 995728 | 14 |
| 47 | 003318 | 2076 | 997784 | 21 | 005534 | 2097 | 994466 | 13 |
| 48 | 004563 | 2070 | 997771 | 21 | 006792 | 2091 | $99^{3208}$ | 12 |
| 49 | oo5805 | 2064 | 997758 | 21 | 008047 | 2085 | 991953 | 11 |
| 50 | 007044 | 2058 | 997745 | 21 | 009298 | 2080 | 990702 | 10 |
| 51. | 9.008278 | 2052 | 9.997732 | 21 | 9.010546 | 2074 | 10.988454 | 8 |
| 52 53 53 | 009510 | 2046 | 997719 | 21 | 011790 | 2068 | 988210 |  |
| 53 54 5 | -10737 | 2040 | 997706 | 21 | 013031 | 2062 | 986969 |  |
| 54555 | 011962 | 2034 | $9976{ }^{3}$ | 22 | 014269 | 2056 | 985732 | 6 |
|  | -13182 | 2029 | 997680 | 22 | 015502 | 2051 | 984498 | 5 |
| 55 <br> 56 | O14400 | 2023 | 997667 | 22 | 016732 | 2045 | 983268 | 4 |
| 56 <br> 5 | $\begin{aligned} & 15613 \\ & 016824 \end{aligned}$ | 2017 | 997654 | 22 | 017959 | 2040 | 982041 | 3 |
| 58 <br> 59 |  | 2012 | 997641 | 22 | 019183 | 2033 | 980817 | 2 |
|  | oi6824 or8o3I 019235 | 2006 2000 | 997628 | 22 | 020403 | 2028 | 979597 | 1 |
| 60 |  | 2000 | 997614 | 22 | 021620 | 2023 | 978380 | 0 |
|  | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | , |
| $95^{\circ}$ |  |  |  |  |  |  |  | $84^{\circ}$ |




| 26 | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  |  | T $\mathrm{T}_{\text {ABL }}$ | II. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $8^{\circ}$ | Sine. | D. | Cosine. | D. | Tang. | D. | $171^{\circ}$ |  |
| , |  |  |  |  |  |  | Cotang. | 1 |
| 0 | 9.143555 | 1496 | 9.995753 | 30 | 9.147803 | 1526 | 10.852197 | 60 |
| 1 | 144453 | 1493 | 995735 | 30 | 148718 | 1523 | 851282 | 59 |
| 2 | 145349 | 1490 | 995717 | 30 | 149632 | 1520 | 850368 | 58 |
| 3 | 146243 | 1487 | 995699 | 30 | 150544 | 1517 | 849456 | 57 |
| 4 | 147136 | 1484 | 995681 | 30 | 151454 | 1514 | 848546 | 56 |
| 5 | 148026 | 1481 | 995664 | 30 | 152363 | 15 II | 847637 | 55 |
| 6 | 148915 | 1478 | 995646 | 30 | 153269 | 1508 | 846731 | 54 |
| 7 | 149802 | 1475 | 995628 | 30 | 154174 | 15 c 5 | 845826 | 53 |
| 8 | 150586 | 1472 | 995610 | 30 | 155077 | 1502 | 844923 | 52 |
| 9 | 151569 | 1469 | 995591 | 30 | 155978 | 1499 | 844022 | 51 |
| 10 | 15245 | 1466 | 995573 | 30 | 156877 | 1496 | 843123 | 50 |
| 11 | 9-153330 | 1463 | 9.995555 | 30 | 9-157775 | 1493 | 10.842225 | 49 |
| 12 | 154208 | 1460 | 995537 | 30 | 158671 | 1490 | 841329 | 48 |
| 13 | 155083 | 1457 | 995519 | 30 | 159565 | 1487 | 840435 | 47 |
| 14 | 155057 | 1454 | 995501 | 3 I | 160457 | 1484 | 839543 | 46 |
| 15 | 156830 | 1451. | 995482 | 3 I | 161347 | 1481 | 838653 | 45 |
| 16 | 157700 | 1448 | 995464 | 31 | 162236 | 1479 | 837764 | 44 |
| 17 | 158569 | 1445 | 995446 | 3 I | 163123 | 1476 | 836877 | 43 |
| 18 | 159435 | 1442 | 995427 | 3 I | 164008 | 1473 | 835992 | 42 |
| 19 | 160301 | 1.439 | 995409 | 3 I | 164892 | 1470 | 835108 | 41 |
| 20 | 161164 | 1436 | 995390 | 31 | 165774 | 1467 | 834226 | 40 |
| 21 | 9-162025 | 1433 | 9.995372 | 3 I | 9.166654 | 1464 | 10.833346 | 39 |
| 22 | 162885 | 1430 | 995353 | 31 | 167532 | 1461 | 832468 | 38 |
| .23 | 163743 | 1427 | 995334 | 31 | 168409 | 1458 | 831591 | 37 |
| 24 | 164600 | 1424 | 995316 | 3 I | 169284 | 1455 | 830716 | 36 |
| 25 | 165454 | 1422 | 995297 | 3 I | 170157 | 1453 | 829843 | 35 |
| 26 | 166307 | 1419 | 995278 | 3 I | 171029 | 1450 | 828971 | 34 |
| 27 | 167159 | 1416 | 995260 | 3 I | 171899 | 1447 | 828101 | 33 |
| 28 | 168008 | 1413 | 995241 | 32 | 172767 | 1444 | 827233 | 32 |
| 29 | 168856 | 1410 | 995222 | 32 | 173634 | 1442 | 826366 | 31 |
| 30 | 169702 | 1407 | 995203 | 32 | 174499 | 1439 | 825501 | 30 |
| 31 | 9.170547 | 1405 | 9.995184 | 32 | 9.175362 | 1436 | 10.824638 |  |
| 32 | 171389 | 1402 | 995165 | 32 | 176224 | 1433 | 823776 | 28 |
| 33 | 172230 | I399 | 995146 | 32 | 177084 | 1431 | 822916 | 27 |
| 34 | 173070 | 1396 | 995127 | 32 | 1779 ¢2 | 1428 | 822058 | 26 |
| 35 | 173908 | 1394 | 995108 | 32 | 178799 | 1425 | 821201 | 25 |
| 36 | 174744 | 1391 | 995089 | 32 | 179655 | 1423 | 820345 | 24 |
| 37 | 175578 | 1388 | 995070 | 32 | 180508 | 1420 | 819492 | 23 |
| 38 | 176411 | 1386 | 995051 | 32 | 181360 | 1417 | 818640 | 22 |
| 39 | 177242 | 1383 | 995032 | 32 | 182211 | 1415 | 817789 | 21 |
| 40 | 178072 | 138 c | 99 จิอ 3 | 32 | 1830 ¢) | 1412 | 816911 | 20 |
| 41 | 9.178900 | 1377 | 9.994993 | 32 | 9.183907 | 1409 |  |  |
| 42 | 179726 | 1374 | 994974 | 32 | $1847^{5} 2$ | 1407 | 815248 | 18 |
| 43 | 180551 | 1372 | 994955 | 32 | 185507 | 1404 | 814403 | 17 |
| 44 | 181374 | 1369 | 994935 | 32 | 186439 | 1402 | 813561 | 16 |
| 45 | 182196 | 1366 | 994910 ́ | 33 | 187280 | 1399 | 812720 | 15 |
| 46 | 183016 | 1364 | 994896 | 33 | 188120 | 1306 | 811880 | 14 |
| 47 | 183834 | 1361 | 994877 | 33 | 18 Sq 58 | 1393 | 811042 | 13 |
| 48 | 18465 I | 1359 | 994857 | 33 | 189794 | 1391 | 810206 | 12 |
| 49 | 185466 | 1356 | 994838 | 33 | 190629 | I 389 | $80937:$ | 11 |
| 50 | 186280 | 1353 | 994818 | 33 | 191462 | 1386 | 808538 | 10 |
| 51 | 9-187092 | 1351 | 9.994798 | 33 | 9.192294 | 1384 | 10.807706 |  |
| 52 | 187903 | 1348 | 994779 | 33 | 193124 | 1381 | 806876 | 8 |
| 53 | 188712 | 1346 | 994759 | 33 | 193953 | 1379 | 806047 |  |
| 54 | 189519 | 1343 | 994739 | 33 | 194780 | 1376 | 805220 | 5 |
| 55 | 190325 | 1341 | 994720 | 33 | 195606 | 1374 | 804394 | 5 |
| 56 | 191130 | 1338 | 994700 | 33 | 196430 | 1371 | 803570 | 4 |
| 57 58 | 191933 | 1336 | 994680 | 33 | 197253 | 1369 | 802747 | 3 |
| 58 | 19274 | 1333 | 994660 | 33 | 198074 | 1366 | 801926 | 2 |
| 59 | 193534 | 1330 | 994640 | 33 | 198894 | 1364 | 801106 | 1 |
| 60 | 194332 | 1328 | 994620 | 33 | 199713 | 1361 | 800287 | - |
| 1 | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | , |
| $98^{\circ}$ |  |  |  |  |  |  |  | $81^{\circ}$ |



| 28 |  | LOGARITHMIC SINES, TANGENTS, ETC |  |  |  |  | Table II. |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $10^{\circ}$ |  |  |  |  |  |  | $169^{\circ}$ |  |
| , | S:ne. | D. | Cosine. | D. | Tang. | D. | Cotang. | , |
| 0 | 9.239670 | 1193 | 9.993351 | 37 | 9.246319 | 1230 | 10.753681 | 60 |
| 1 | 240386 | 1191 | 993329 | 37 | 247057 | 1228 | 752943 | 59 |
| 2 | 241101 | 1189 | 993307 | 37 | 247794 | 1226 | 752206 | 58 |
| 3 | 241814 | 1187 | 993284 | 37 | 248530 | 1224 | 751470 | 57 |
| 4 | 242526 | 1185 | 993262 | 37 | 249264 | 1222 | 750736 | 56 |
| 5 | 243237 | 1183 | 993240 | 37 | 249998 | 1220 | 750002 | 55 |
| 6 | 243947 | 1181 | 993217 | 38 | 250730 | 1218 | 749270 | 54 |
| 7 | 244656 | 1179 | $99^{3195}$ | 38 | 251461 | 1217 | 748539 | 53 |
| 8 | 24 F363 | 1177 | 993172 | 38 | 252191 | 1215 | 747809 | 52 |
| 9 | 246069 | 1175 | 993149 | 38 | 252920 | 1213 | 747080 | 5 I |
| 10 | $24677^{\circ}$ | 1173 | 993127 | 38 | 253648 | 1211 | 746352 | 50 |
| 1 | 9.247478 | 1171 | 9.993104 | 38 | 9.254374 | 1209 | 10.745626 | 49 |
| 12 | 248181 | 1169 | 993081 | 38 | 255100 | 1207 | 744900 | 48 |
| 13 | 248883 | 1167 | 993059 | 38 | 255824 | 1205 | 744176 | 47 |
| 14 | 249583 | 1165 | 993036 | 38 | 256547 | 1203 | 743453 | 46 |
| 15 | 250282 | 1163 | 993013 | 38 | 250269 | 1201 | 742731 | 45 |
| 16 | 250980 | 1161 | 992990 | 38 | 257990 | 1200 | 742010 | 44 |
| 17 | 251677 | 1159 | 992967 | 38 | 258710 | 1198 | 741290 | 43 |
| 18 | 252373 | 1158 | 992944 | 38 | 259429 | 1196 | 940571 | 42 |
| 19 | 253067 | 1156 | 992921 | 38 | 260146 | 1194 | 739854 | 41 |
| 20 | 253761 | 1154 | 992898 | 38 | 260863 | 1192 | 739137 | 40 |
| 21 | 9.254453 | 1152 | 9.992875 | 38 | 9.261578 | 1190 | $10 \cdot 738422$ | 39 |
| 22 | 255144 | 1150 | 992852 | 38 | 262292 | 1189 | 737708 | 38 |
| 23 | 255834 | 1148 | 992829 | 39 | 263005 | 1187 | 736995 | 37 |
| 24 | 256523 | 1146 | 992806 | 39 | 263717 | 1185 | 736283 | 36 |
| 25 | 257211 | 1144 | 992783 | 39 | 264428 | 1183 | 735572 | 35 |
| 26 | 257898 | 1142 | 992759 | 39 | 265138 | 1181 | 734862 | 34 |
| 27 | 258583 | 1141 | 992735 | 39 | 265847 | 1179 | 734153 | 33 |
| 28 | 259268 | 1139 | 992713 | 39 | 266555 | 1178 | 733445 | 32 |
| 29 30 | 259951 | 1137 | 992690 | 39 | 267261 | 1176 | 732739 | 3 I |
| 30 | 260633 | 1135 | 992666 | 39 | 267967 | 1174 | 732033 | 30 |
| 31 | 9.261314 | 1133 | 9.992643 | 39 | 9-268571 | 1172 | 10.731329 | 29 |
| 32 | 261994 | 1131 | 992619 | 39 | 269375 | 1170 | 730625 | 28 |
| 33 | 262673 | 1130 | 992596 | 39 | 270077 | 1189 | 729923 | 27 |
| 34 | 263351 | 1128 | 992572 | 39 | 270779 | 1167 | 720221 | 26 |
| 35 | 264027 | 1126 | 992549 | 39 | 271479 | 1165 | 728521 | 25 |
| 36 | 264703 | 1124 | 992525 | 39 | 272178 | 1164 | 727822 | 24 |
| 37 | 265377 | 1122 | 992501 | 39 | 272876 | 1162 | 727124 | 23 |
| 38 | 266051 | 1120 | 992478 | 40 | 273573 | 1160 | 726427 | 22 |
| 39 | 266723 | 1119 | 992454 | 40 | 274269 | 1158 | 725731 | 21 |
| 40 | 267395 | 1117 | 992430 | 40 | 274964 | 1157 | 725036 | 20 |
| 41 | 9-268065 | 1115 | 9.992406 | 40 | 9.275658 | 1155 |  | 19 |
| 42 | 268734 | 1113 | 992382 | 40 | 276351 | 1153 | 723649 | 18 |
| 43 | 269402 | 1111 | 992359 | 40 | 277043 | 1151 | 722957 | 17 |
| 44 | 270069 | 1110 | 992335 | 40 | 277734 | 1150 | 722266 | 16 |
| 45 | 270735 | 1108 | 992311 | 40 | 278424 | 1148 | 721576 | 15 |
| 46 | 271400 | 1106 | 992287 | 40 | 279113 | 1147 | 720887 | 14 |
| 47 | 272064 | 1105 | 992263 | 40 | 279801 | 1145 | 720199 | 13 |
| 48 | 272726 | 1103 | 992239 | 40 | 280488 | 1143 | 719512 | 12 |
| 49 | 273388 | 1101 | 992214 | 40 | 281174 | 1141 | 718826 | 11 |
| 50 | 274049 | 1099 | 992190 | 40 | 281858 | 1140 | 718142 | 10 |
| 51 | 9.274708 | 1098 | 9.992166 | 40 |  | 1138 |  |  |
| 52 | 275367 | 1096 | 992142 | 40 | 283225 | 1136 | 716775 | 8 |
| 53 | 276025 | 1094 | 992118 | 41 | 283907 | 1135 | 716093 | 6 |
| 54 | 276681 | 1092 | 992093 | 41 | 284588 | 1133 | 715412 | 6 |
| 55 | 277337 | 10 l 1 | 992069 | 41 | 285268 | 1131 | 714732 | 5 |
| 56 | 277991 | 1089 | 992044 | 41 | 285947 | 1130 | 714053 | 4 |
| 57 | 278645 | 1087 | 992020 | 41 | 286624 | 1128 | 713376 | 3 |
| 58 | 279297 | 1086 | 991596 | 41 | 287301 | 1126 | 712699 | 2 |
| 59 | 279948 | 1084 | 991971 | 41 | 287977 | 1125 | 712023 | 1 |
| 60 | 280599 | 1082 | 991947 | 41 | 288652 | 1123 | 711348 | 0 |
|  | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | , |
| $100^{\circ}$ |  |  |  |  |  |  |  | $79^{\circ}$ |

Table II. LOGARITHMIC SINES: TANGENTS, ETC.


| 20 |  | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  | Table II. |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $12^{\circ}$ |  |  |  |  |  |  | $167^{\circ}$ |  |
| , | Sine. | D. | Cosine | D. | Tang. | D. | Cotang. | , |
| 0 | 9.317879 | 990 | 9.990404 | 45 | 9.327475 | 1035 | 10.672525 | 60 |
| 1 | 318473 | 988 | 990378 | 45 | 328095 | 1033 | 671905 | 5 |
| 2 | 319066 | 987 | 990351 | 45 | 328715 | 1032 | 671285 | 58 |
| 3 | 319658 | 986 | 990324 | 45 | 329334 | 1030 | 670666 | 57 |
| 4 | 320249 | 984 | 990297 | 45 | 329953 | 1029 | 670047 | 56 |
| 5 | 320840 | 983 | 990270 | 45 | 330570 | 1028 | 669430 | 55 |
| 6 | 321430 | 982 | 990243 | 45 | 331187 | 1026 | 668813 | 54 |
| 7 | 322019 | 980 | 990215 | 45 | 331803 | 1025 | 668197 | 53 |
| 8 | 322607 | 979 | 990188 | 45 | 332418 | 1024 | 667582 | 52 |
| 9 | 323194 | 977 | 990161 | 45 | 333033 | 1023 | 656067 | 51 |
| 10 | 323780 | 976 | 990134 | 45 | 333646 | 1021 | 666354 | 50 |
| 11 | 9.324366 | 975 | 9-990107 | 46 | 9.334259 | 1020 | 10.665741 | 49 |
| 12 | 324950 | 973 | 990079 | 46 | 334871 | 1019 | 665129 | 48 |
| 13 | 325534 | 972 | 990052 | 46 | 335482 | 1017 | 664518 | 47 |
| 14 | 326117 | 970 | 990025 | 46 | 336093 | 1016 | 663907 | 46 |
| 15 | 326700 | 969 | 989997 | 46 | 336702 | 1015 | 663298 | 45 |
| 16 | 327281 | 968 | 989970 | 46 | 337311 | 1013 | 662689 | 44 |
| 17 | 327862 | 966 | 989942 | 46 | 337919 | 1012 | 662081 | 43 |
| 18 | 328442 | 965 | 989915 | 46 | 338527 | 1011 | 661473 | 42 |
| 19 | 329021 | 964 | 989887 | 46 | 339133 | 1010 | 660867 | 41 |
| 20 | 329599 | 962 | 989860 | 46 | 339739 | 1008 | 660261 | 40 |
| 21 | 9.330176 | 961 | 9.989832 | 46 | 9.340344 | 1007 | $10 \cdot 659656$ | 39 |
| 22 | 330753 | 960 | 989804 | 46 | 340248 | 1006 | $65 \mathrm{go5} 2$ | 38 |
| 23 | 331329 | 958 | 989777 | 46 | 341552 | 1004 | 658448 | 37 |
| 24 | 331903 | 957 | 989749 | 47 | 342155 | 1003 | 657845 | 36 |
| 25 | 332478 | 956 | 989721 | 47 | 342757 | 1002 | 657243 | 35 |
| 26 | 333051 | 954 | 989693 | 47 | 343358 | 1000 | 656642 | 34 |
| 27 | 333624 | 953 | 989665 | 47 | 343958 | 999 | 656042 | 33 |
| 28 | 334195 | 952 | 989637 | 47 | 344558 | $99^{8}$ | 655442 | 32 |
| 29 | 334767 | 950 | 989610 | 47 | 345157 | 997 | 654843 | 31 |
| 30 | 335337 | 949 | 989582 | 47 | 345755 | 996 | 654245 | 30 |
| 31 | 9-335906 | 948 | $9 \cdot 989553$ | 47 | 9.346353 | 994 | 10.653647 | 29 |
| 32 | 336475 | 946 | 989525 | 47 | 346949 | 993 | 653051 | 28 |
| 33 | 337043 | 945 | 989497 | 47 | 347545 | 992 | 652455 | 27 |
| 34 | 337610 | 944 | 989469 | 47 | 348141 | 991 | 651859 | 26 |
| 35 | 338176 | 943 | 989441 | 47 | 348735 | 990 | 651265 | 25 |
| 36 | 338742 | 941 | 989413 | 47 | 349329 | 988 | 650671 | 24 |
| 37 | 339307 | 940 | 989385 | 47 | 349922 | 987 | 650078 | 23 |
| 38 | 339871 | 939 | 989356 | 47 | 350514 | 986 | 642486 | 22 |
| 39 | 340434 | 937 | 989328 | 47 | 351106 | 985 | 648894 | 21 |
| 40 | 340996 | 936 | 989300 | 47 | 351697 | 983 | 648303 | 20 |
| 41 | 9.341558 | 935 | 9.989271 | 47 | 9.352287 | 982 | 10.647713 | 19 |
| 42 | 342119 | 934 | 989243 | 47 | 352876 | 981 | 647124 | 18 |
| 43 | 342679 | 932 | 989214 | 47 | 353465 | 980 | 646535 | 17 |
| 44 | 343239 | 931 | 989186 | 47 | 354053 | 979 | 645947 | 16 |
| 45 | 343797 | 930 | 989157 | 47 | 354640 | 977 | 645360 | 15 |
| 46 | 344355 | 929 | 989128 | 48 | 355227 | 976 | 644773 | 14 |
| 47 | 344912 | 927 | 989100 | 48 | 355813 | 975 | 644187 | 13 |
| 48 | 345469 | 926 | 989071 | 48 | $35639^{8}$ | 974 | 643602 | 12 |
| 49 | 346024 | 925 | 9890.12 | 48 | 356982 | 973 | 643018 | 11 |
| 50 | 346579 | 924 | 989014 | 48 | 357566 | 971 | 642434 | 10 |
| 51 | 9.347134 | 922 | 9.988985 | 48 | 9.358149 | 970 | 10.641851 |  |
| 52 | 347687 | 921 | 988956 | 48 | 358731 | 969 | 641269 | 8 |
| 53 | 348240 | 920 | 988927 | 48 | 359313 | 968 | 640687 | 7 |
| 54 | 348792 | 919 | 988898 | 48 | 359893 | 967 | 6 ¢01c? | 6 |
| 55 | 349343 | 917 | 988869 | 48 | 360474 | 966 | 639526 | 5 |
| 56 | 349893 | 916 | 988840 | 48 | 361053 | 965 | 638947 | 4 |
| 59 58 58 | 350443 350992 | 915 | 988811 | 49 | 361632 | 963 | 638368 | 3 |
| 59 | 351540 | 914 | 988782 | 49 | 362210 | 962 | 637790 637213 | 2 |
| 50 | 352088 | 911 | 988724 | 49 | 363364 | 961 960 | 637213 636636 | 1 |
| , | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | 1 |
| $102^{\circ}$ |  |  |  |  |  |  |  | $77^{\circ}$ |


| Table II. |  | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  |  | 31 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $13^{\circ}$ |  |  |  |  |  |  |  | $166^{\circ}$ |
| , | Sine. | D. | Cosine. | D. | Tang. | D. | Cotang. | , |
| 0 | 9.352088 | 911 | 9.988724 | 49 | 9.363364 | 960 | 10.636636 | 60 |
| 1 | 352635 | 910 | 988695 | 49 | 363940 | 959 | 636060 | 59 |
| 2 | 353181 | 909 | 988666 | 49 | 364515 | 958 | 635485 | 58 |
| 3 | 353726 | 928 | 988636 | 49 | 365090 | 957 | 634910 | 57 |
| 4 | 354271 | 907 | 988607 | 49 | 365664 | 955 | 634336 | 56 |
| 5 | 354815 | 905 | 988578 | 49 | 366237 | 954 | 633763 | 55 |
| 6 | 355358 | 904 | 988543 | 49 | 366810 | 953 | 633190 | 54 |
| 7 | 355901 | 903 | 988519 | 49 | 367382 | 952 | 632618 | 53 |
| 8 | 356443 | co2 | 988489 | 49 | 367953 | 951 | 632047 | 52 |
| 9 | 356984 | 901 | 988460 | 49 | 368524 | 950 | 631476 | 51 |
| 10 | 357524 | 899 | 988430 | 49 | 369094 | 949 | 630906 | 50 |
| 11 | 9.358064 | 898 | 9.988401 | 49 | 9.369663 | 948 | 10.630337 | 49 |
| 12 | -358603 | 897 | 988371 | 49 | 370232 | 946 | 629768 | 48 |
| 13 | 359141 | 896 | 988342 | 49 | 370799 | 945 | 629201 | 47 |
| 14 | 359678 | 895 | 988312 | 50 | 371367 | 944 | 628633 | 46 |
| 15 | 360215 | 893 | 988282 | 50 | 371933 | 943 | 628067 | 45 |
| 16 | 360752 | 892 | 988252 | 50 | 372499 | 942 | 627501 | 44 |
| 17 | 361287 | 891 | 988223 | 50 | 373064 | 941 | 626936 | 43 |
| 18 | 361822 | 890 | 988193 | 50 | 373629 | 940 | 626371 | 42 |
| 19 | 362356 | 889 | 988163 | 50 | 374193 | $9^{3} 9$ | 625807 | 41 |
| 20 | 362889 | 888 | 988 r 33 | 50 | 374756 | 938 | 625244 | 40 |
| 21 | 9.363422 | 887 | 9.988103 | 50 | 9.375319 | 937 | 10.624681 | 39 |
| 22 | 363954 | 885 | 988073 | 50 | 375881 | 935 | 624119 | 38 |
| 23 | 364485 | 884 | 988043 | 50 | 376442 | 934 | 623558 | 37 |
| 24 | 365016 | 883 | 988013 | 50 | 377003 | 933 | 622997 | 36 |
| 25 | 365546 | 882 | 987983 | 50 | 377563 | 932 | 622437 | 35 |
| 26 | 366075 | 881 | 987953 | 50 | 378122 | 931 | 621878 | 34 |
| 27 | 366604 | 880 | 987922 | 50 | 378681 | 930 | 621319 | 33 |
| 28 | 367131 | 879 | 987892 | 50 | 379239 | 929 | 620761 | 32 |
| 29 | 367659 | 877 | 987862 | 50 | 379797 | 928 | 620203 | 31 |
| 30 | 368185 | 876 | 987832 | 51 | 380354 | 927 | 619646 | 30 |
| 3 I | $9 \cdot 368711$ | 875 | $9 \cdot 987801$ | 51 | 9.380910 | 926 | 10.619090 | 29 |
| 32 | 369236 | 874 | 987771 | 51 | 381466 | 925 | 618534 | 28 |
| 33 | 369761 | 873 | 987740 | 51 | 382020 | 924 | 617980 | 27 |
| 34 | 370285 | 872 | 987710 | 51 | 382575 | 923 | 617425 | 26 |
| 35 | 370808 | 871 | 987679 | 51 | 383129 | 922 | 616871 | 25 |
| 36 | 371330 | 870 | 987649 | 51 | 383682 | 921 | 616318 | 24 |
| 37 | 371852 | 869 | 987618 | 51 | 384234 | 920 | 615766 | 23 |
| 38 | 372373 | 867 | 987588 | 51 | 384786 | 919 | 615214 | 22 |
| 39 | 372894 | 866 | 987557 | 51 | 385337 | 918 | 614663 | 2 I |
| 40 | 373414 | 865 | 987526 | 51 | 385888 | 917 | 614112 | 20 |
| 41 | $9 \cdot 373933$ | 864 | 9.987496 | 51 | 9.386438 | 915 | 10.613562 | 19 |
| 42 | 374452 | 863 | 987465 | 51 | - 386987 | 914 | 6ı3013 | 18 |
| 43 | 374970 | 862 | 987434 | 51 | 387536 | 913 | 612464 | 17 |
| 44 | 375487 | 861 | 987403 | 52 | 388084 | 912 | 611916 | 16 |
| 45 | 376003 | 860 | 987372 | 52 | 388631 | 911 | 611369 | 15 |
| 46 | 376519 | 859 | 987341 | 52 | 389178 | 910 | 610822 | 14 |
| 47 | 377035 | 858 | 987310 | 52 | 389724 | 909 | 610276 | 13 |
| 48 | 377549 | 857 | 987279 | 52 | 390270 | 908 | 609730 | 12 |
| 49 | 378063 | 856 | 987248 | 52 | 390815 | 907 | 609185 | 11 |
| 50 | 378577 | 854 | 987217 | 52 | 391360 | 906 | 608640 | 10 |
| 51 | $9 \cdot 379089$ | 853 | Э 9.987186 | 52 | 9.391903 | 905 | 10.608097 | 8 |
| 52 | 379601 | 852 | 987155 | 52 | 392447 | 904 | 607553 | 8 |
| 53 | 380113 | 851 | 987124 | 52 | 392989 | 903 | 607011 | 7 |
| 54 | 380624 | 850 | 987092 | 52 | 393531 | 902 | 636469 | 6 |
| 55 | 381134 | 849 | 987061 | 52 | 394073 | 901 | 605927 | 5 |
| 56 | 381643 | 848 | 987030 | 52 | 394614 | 900 | 605385 | 4 |
| 57 58 | 382152 | 847 | 986998 | 52 | 395154 | 809 | 604846 | 3 |
| 58 | 382661 | 846 | 986967 | 52 | 395694 | 898 | 604306 | 2 |
| 59 60 | 383168 383675 | 845 | 986936 | 52 | 396233 | 897 | 603767 | 1 |
| 60 | 383675 | 844 | 986904 | 52 | 396771 | 896 | 603229 | 0 |
| 1 | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | , |
| $103^{\circ}$ |  |  |  |  |  |  |  | $76^{\circ}$ |


| 32 |  | LOGARITHMIC SLNES, TANGENTS, ETC. |  |  |  |  | Table II. |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $14^{\circ} 165^{\circ}$ |  |  |  |  |  |  |  |  |
| , | Sine. | D. | Cosine. | D. | Tang. | D. | Cutang. | , |
| $\begin{array}{r} 0 \\ 1 \\ 2 \\ 3 \\ 4 \\ 5 \\ 6 \\ 7 \\ 8 \\ 9 \\ 10 \end{array}$ | 9.383675 | 844 | 9.986904 | 52 | 9-396771 | 896 | 10.603229 | 6 c |
|  | 384182 | 843 | 986873 | 53 | 397309 | 896 | 602691 | 59 |
|  | 384687 | 842 | 98684 I | 53 | 397846 | 895 | 602154 | 58 |
|  | 385192 | 84 x | 986809 | 53 | 398383 | 894 | 601617 | 57 |
|  | 385697 | 840 | 986778 | 53 | 398919 | $89^{3}$ | 601081 | 56 |
|  | 386201 | 839 | 986746 | 53 | 399455 | 892 | 600545 | 55 |
|  | 386704 | 838 | 986714 | 53 | 399990 | 891 | 600010 | 54 |
|  | 387207 | 837 | 986683 | 53 | 400524 | 890 | 599476 | 53 |
|  | 387709 | 836 | 986651 | 53 | 401058 | 889 | 598942 | 52 |
|  | 388210 | 835 | 986619 | 53 | 401591 | 888 | 598409 | 51 |
|  | 388711 | 834 | 986587 | 53 | 402124 | 887 | 597876 | 50 |
| 11 | 9.389211 | 833 | $9 \cdot 986555$ | 53 | 9.402656 | 886 | 10. 597344 | 49 |
| 12 | 389711 | 832 | 986523 | 53 | 403187 | 885 | 596813 | 48 |
| 13 | 390210 | 831 | 986491 | 53 | 403718 | 884 | 596282 | 47 |
| 14 | 390708 | 830 | 986459 | 53 | 404249 | 883 | $59575{ }^{1}$ | 46 |
| 15 | 391206 | 828 | 986427 | 53 | 404778 | 882 | 595222 | 45 |
| 16 | 391703 | 827 | 986395 | 53 | 405308 | 881 | 594692 | 44 |
| 17 | 392199 | 826 | 986363 | 54 | 405836 | 880 | 594164 | 43 |
| 18 | 392695 | 825 | 986331 | 54 | 406364 | 879 | 593636 | 42 |
| 19 | 393191 | 824 | 986299 | 54 | 406892 | 878 | 593108 | 41 |
| 20 | 393685 | 823 | 986266 | 54 | 407419 | 877 | 592581 | 40 |
| 21 | 9.394179 | 822 | 9.986234 | 54 | 9.407945 | 876 | 10.592055 | 39 |
| 22 | 394673 | 821 | 986202 | 54 | 408471 | 875 | 591529 | 38 |
| 23 | 395166 | 820 | 986169 | 54 | 408996 | 874 | 591004 | 37 |
| 24 | 395658 | ${ }^{81} 9$ | 986137 | 54 | 409521 | 874 | 590479 | 36 |
| 25 | 396150 | 818 | 986104 | 54 | 410045 | 873 | 589955 | 35 |
| 26 | 396641 | 817 | 986072 | 54 | 410569 | 872 | 589431 | 34 |
| 27 | 397132 | 817 | 986039 | 54 | 411092 | 871 | 588908 | 33 |
| 28 | 397621 | 816 | 986007 | 54 | 411615 | 870 | 588385 | 32 |
| 29 | 398111 | 815 | 985974 | 54 | 412137 | 869 | 587863 | 3 I |
| 30 | 398600 | 814 | 985942 | 54 | 412658 | 868 | 587342 | 30 |
| 31 | 9.399088 | 813 | 9.985909 | 55 | 9.413179 | 867 | 10.586821 | 29 |
| 32 | 399575 | 812 | 985876 | 55 | 413699 | 866 | 586301 | 28 |
| 33 | 400062 | 811 | 985843 | 55 | 414219 | 865 | 585781 | 27 |
| 34 | $40054{ }^{2}$ | 810 | 9858 II | 55 | 414738 | 864 | 585262 | 26 |
| 35 | 401033 | 809 | 985778 | 55 | 415257 | 864 | 584743 | 25 |
| 36 | 401520 | 808 | 985745 | 55 | 415775 | 863 | 584225 | 24 |
| 37 | 402005 | 807 | 985712 | 55 | 416293 | 862 | 583707 | 23 |
| 38 | 402489 | 806 | 985679 | 55 | 416810 | 861 | 583190 | 22 |
| 39 | 402972 | $8 \mathrm{c5}$ | 985646 | 55 | 417326 | 860 | 582674 | 21 |
| 40 | 403455 | 804 | ¢85613 | 55 | 417842 | 859 | 582158 | 20 |
| 41 | 9.403938 | 803 | 9.985580 | 55 | 9.418358 | 858 | 10.581642 | 19 |
| 42 | 404420 | 802 | 985547 | 55 | 418873 | 857 | 581127 | 18 |
| 43 | 404001 | 801 | 985514 | 55 | 419387 | 856 | 580613 | 17 |
| 44 | 405382 | 800 | 985480 | 55 | 419901 | 855 | 580099 | 16 |
| 45 | 405862 | 799 | 985447 | 55 | 420415 | 855 | 579585 | 15 |
| 46 | 406341 | 738 | 985414 | 56 | 420927 | 854 | 579073 | 14 |
| 47 | 406820 | 797 | 985381 | 56 | 421440 | 853 | 578560 | 13 |
| 48 | 407299 | 796 | 985347 | 56 | 421952 | 852 | 578048 | 12 |
| 49 | 407777 | 795 | 985314 | 56 | 422463 | 85 I | 577537 | 11 |
| 50 | 408254 | 794 | 985280 | 56 | 422974 | 850 | 577026 | 10 |
| 51 | $9 \cdot 408731$ | 794 | 9.985247 | 56 | 9-423484 | 849 | 10.576516 |  |
| 52 | 409207 | $79^{3}$ | 985213 | 56 | 423293 | 848 | 576007 | 8 |
| 53 | 409682 | 792 | 985180 | 56 | 424503 | 848 | 575497 | 7 |
| 54 | 410157 | 791 | 985146 | 56 | 425011 | 847 | 574989 | 6 |
| 55 | 410632 | 790 | 985113 | 56 | 425519 | 846 | 574481 | 5 |
| 56 | 411106 | 789 | 985079 | 56 | 426027 | 8.45 | 573973 | 4 |
| 57 | 411579 | 788 | 985045 | 56 | 426534 | 844 | 573466 | 3 |
| 58 | 412052 | 787 | 985011 | 56 | 427041 | 843 | 572959 | 2 |
| 59 | 412524 | 786 | 984978 | 56 | 427547 | 843 | 572453 | 1 |
| 60 | 412996 | 785 | 984944 | 56 | 428052 | 842 | 571948 | 0 |
| 1 | Cosine | D. | Sine. | D. | Cotang. | D. | Tang. | , |
| $104^{\circ}$ |  |  |  |  |  |  |  | $75^{\circ}$ |


| Table II. |  | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  |  | 33 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $15^{\circ}$ |  |  |  |  |  |  |  | $64^{\circ}$ |
| , | Sine. | D. | Cosine. | D. | Tang. | D. | Cotang. | , |
| 0 | 9.412996 | 785 | 9.984944 | 57 | 9.428052 | 842 | 10.571948 | 60 |
| 1 | 413467 | 784 | 984910 | 57 | 428558 | 841 | 571442 | 59 |
| 2 | 413938 | 783 | 984876 | 57 | 429062 | 840 | 570938 | 58 |
| 3 | 414408 | 783 | 984842 | 57 | 429566 | 839 | 570434 | 57 |
| 4 | 414878 | 782 | 984808 | 57 | 430070 | 838 | 569930 | 56 |
| 5 | 415347 | 781 | 984774 | 57 | 430573 | 838 | 569427 | 55 |
| 6 | 415815 | 780 | 984740 | 57 | 431075 | 837 | 568925 | 54 |
| 7 | 416283 | 779 | 984706 | 57 | 431577 | 836 | 568423 | 53 |
| 8 | 416751 | 778 | 984672 | 57 | 432079 | 835 | 567921 | 52 |
| 9 | 417217 | 777 | 984638 | 57 | 432580 | 834 | 567420 | 51 |
| 10 | 417684 | 776 | 984603 | 57 | 433080 | 833 | 566920 | 50 |
| 11 | 9.418150 | 775 | $9 \cdot 984569$ | 57 | 9.433580 | 832 | 10.566420 | 49 |
| 12 | 418615 | 774 | 984535 | 57 | 434080 | 832 | 565920 | 48 |
| 13 | 419079 | 773 | 984500 | 57 | 434579 | 831 | 565421 | 47 |
| 14 | 419544 | 773 | 984466 | 57 58 | 435078 | $83 n$ | $564{ }^{\text {n }}$, 2 | 46 |
| 15 | 420007 | 772 | 984432 | 58 | 435070 | 829 | 564424 | 45 |
| 16 | 420470 | 771 | 984397 | 58 | 436073 | 828 | 563927 | 44 |
| 17 | 420033 | 770 | 984363 | 58 | 436570 | 828 | 563430 | 43 |
| 18 | 421395 | 769 | 984328 | 58 | 437067 | 827 | 562933 | 42 |
| 19 | 421857 | 768 | 984294 | 58 | 437563 | 826 | 562437 | 41 |
| 20 | 422318 | 767 | 984259 | 58 | 438059 | 825 | 561941 | 40 |
| 21 | 9.422778 | 767 | 9.984224 | 58 | 9-438554 | 824 | 10.561446 | 39 |
| 22 | 423238 | 766 | 984190 | 58 | 439048 | 823 | 560952 | 38 |
| 23 | 423697 | 765 | 984155 | 58 | 439543 | 323 | 560457 | 37 |
| 24 | 424156 | 764 | 984120 | 58 | 440036 | 822 | 559964 | 36 |
| 25 | 424615 | 763 | 984085 | 58 | 440529 | 821 | 559471 | 35 |
| 26 | 425073 | 762 | 984050 | 58 | 441022 | 820 | 558978 | 34 |
| 27 | 425530 | 761 | 984015 | 58 | 441514 | 819 | 558486 | 33 |
| 28 | 425987 | 760 | 983981 | 58 | 442006 | 819 | 557994 | 32 |
| 29 | 426443 | 760 | 983946 | 58 | 442497 | 818 | 557503 | 31 |
| 30 | 426899 | 759 | 983911 | 58 | 442988 | 817 | 557012 | 30 |
| 31 | 9.427354 | 758 | $9 \cdot 983875$ | 58 | 9.443479 | 816 | 10.556521 | 29 |
| 32 | 427809 | 757 | 983840 | 59 | 443968 | 816 | 556032 | 28 |
| 33 | 428263 | 756 | 983805 | 59 | 444458 | 815 | 555542 | 27 |
| 34 | 428717 | 755 | 983770 | 59 | 444947 | 814 | 555053 | 26 |
| 35 | 429170 | 754 | 983735 | 59 | 445435 | 813 | 554565 | 25 |
| 36 | 429623 | 753 | 983700 | 59 | 445923 | 812 | 554077 | 24 |
| 37 | 430075 | $7{ }^{5} 2$ | 983664 | 59 | 446411 | 812 | 553589 | 23 |
| 38 | 430527 | 752 | 983629 | 59 | 446898 | ${ }_{8}^{811}$ | 553102 | 2 |
| 39 | 430978 | 751 | 983594 | 59 | 447384 | 810 | 552616 | 21 |
| 40 | 431429 | 750 | 983558 | 59 | 447870 | 809 | 552130 | 20 |
| 41 | 9.431879 | 749 | 9.983523 | 59 | 9.448356 | 809 | 10.551644 | 19 |
| 42 | 432329 | 749 | 983487 | 59 | 448841 | 808 | 551159 | 18 |
| 43 | 432778 | 748 | 983452 | 59 | 449326 | 807 | 550674 | 17 |
| 44 | 433226 | 747 | 983416 | 59 | 449810 | 806 | 550190 | 16 |
| 45 | 433675 | 746 | 983381 | 59 | 450294 | 806 | 549706 | 15 |
| 46 | 434122 | 745 | 983345 | 59 | 450777 | 805 | 549223 | 14 |
| 47 | 434569 | 744 | 983309 | 59 | 451260 | 804 | 548740 | 13 |
| 48 | 435016 | 744 | 983273 | 60 | 451743 | 803 | 548257 | 12 |
| 49 | 435462 | 743 | 983238 | 60 | 452225 | 802 | 547775 | 11 |
| 50 | 435908 | 742 | 983202 | 60 | 452706 | 802 | 547294 | 10 |
| 51 | ¢. 436353 | 741 | 9.983166 | 60 | 9.453187 | 801 | 10.546813 |  |
| 52 | 436798 | 740 | 983130 | 60 | 453668 | 800 | 546332 | 8 |
| 53 | 437242 | 740 | 983094 | 60 | 454148 | 799 | 545852 | 7 |
| 54 | 437686 | 739 | 983058 | 60 | 454628 | 799 | 545372 | 6 |
| 55 | 438129 | 738 | 983022 | 60 | 455107 | $79{ }^{8}$ | 544893 | 5 |
| 56 | 438572 | 737 | 982986 | 60 | 455586 | 797 | 544414 | 4 |
| 57 | 439014 | 736 | 982950 | 60 | 456064 | 796 | 543936 | 3 |
| 58 | 439456 | 736 | 982914 | 60 | 456542 | 796 | 543458 | 2 |
| 59 60 | 439897 440338 | 735 734 | 982878 982842 | 60 60 | 457019 457496 | 795 794 | 542981 542504 | 1 |
| 1 | Cosine | D. | Sine. | D. | Cotang. | D. | Tang. | , |
| $105^{\circ}$ |  |  |  |  |  |  |  | $74^{0}$ |


| 34 |  | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  | Table II. |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $16^{\circ}$ |  |  |  |  |  |  |  | $163^{\circ}$ |
| , | Sine. | D. | Cosine. | D. | Tang. | D. | Cotang. | , |
| 0 | 9.440338 | 734 | 9.982842 | 60 | 9.457496 | 794 | 10.542504 | 60 |
| 1 | 440778 | 733 | 982805 | 60 | 457973 | 793 | 542027 | 59 |
| 2 | 441218 | 732 | 982769 | 61 | 458449 | 793 | 541551 | 58 |
| 3 | 441658 | 731 | $9^{82733}$ | 61 | 458925 | 792 | 541075 | 57 |
| 4 | 442096 | 73 ? | 982606 | 61 | 459400 | 791 | 540600 | 50 |
| 5 | 442535 | 730 | 982660 | 61 | 459875 | 790 | 540125 | 55 |
| 6 | 442973 | 729 | 982624 | 61 | 460349 | 790 | 539651 | 54 |
| 7 | 443410 | 728 | 982587 | 61 | 460823 | 789 | 539177 | 53 |
| 8 | 443847 | 727 | 982551 | 61 | 461297 | 788 | 538703 | 52 |
| 9 | 441284 | 727 | 982514 | 61 | 461770 | 788 | 538230 | 51 |
| 10 | 444720 | 726 | 982477 | 61 | 462242 | 787 | 537758 | 50 |
| 11 | 9.445155 | 725 | 9.982441 | 61 | 9.462715 | 786 | 10.537285 | 49 |
| 12 | 445590 | 724 | 982404 | 61 | 463186 | 785 | 536814 | 48 |
| 13 | 446025 | $7<3$ | 982367 | 61 | 463658 | 785 | 536342 | 47 |
| 14 | 446459 | 723 | 982331 | 61 | 464128 | 784 | 535872 | 46 |
| 15 | 446893 | 722 | 982294 | 61 | 464599 | 783 | 535401 | 45 |
| 16 | 447326 | 721 | 982257 | 61 | 465069 | 783 | 534931 | 44 |
| 17 | 447759 | 720 | 982220 | 62 | 465539 | 782 | 534461 | 43 |
| 18 | 448191 | 720 | 982183 | 62 | 466008 | 781 | 533992 | 42 |
| 19 | 448623 | 719 | 982146 | 62 | 466477 | 780 | 533523 | 41 |
| 20 | 449054 | 718 | 982109 | 62 | 466945 | 780 | 533055 | 40 |
| 21 | 9.449485 | 717 | 9.982072 | 62 | 9.467413 | 779 | 10.532587 | 39 |
| 22 | 449915 | 716 | 982035 | 62 | 467880 | 778 | 532120 | 38 |
| 23 | 450345 | 716 | 981998 | 62 | 468347 | 778 | 531653 | 37 |
| 24 | 450775 | 715 | 981961 | 62 | 468814 | 777 | 531186 | 36 |
| 25 | 451204 | 714 | 981924 | 62 | 469280 | 776 | 530720 | 35 |
| 26 | 451632 | 713 | 981886 | 62 | 469746 | 775 | 530254 | 34 |
| 27 | 452060 | 713 | 981849 | 62 | 470211 | 775 | 529789 | 33 |
| 28 | 452488 | 712 | 981812 | 62 | 470676 | 774 | 529324 | 32 |
| 29 | 452915 | 711 | 981774 | 62 | 471141 | 773 | 528859 | 31 |
| 30 | 453342 | 710 | 981737 | 62 | 471605 | 773 | 528395 | 30 |
| 31 | ¢. 453768 | 710 | 9.981700 | 63 | 9.472069 | 772 |  |  |
| 32 | 454194 | 709 | 981662 | 63 | 472532 | 771 | 527468 | 28 |
| 33 | 454619 | 708 | 981625 | 63 | 472995 | 771 | 527005 | 27 |
| 34 | 455044 | 707 | 981587 | 63 | 473407 | 770 | 526543 | 26 |
| 35 | 455.469 | 707 | 98150 | 63 | 473919 | 769 | 526081 | 25 |
| 36 | 455893 | 706 | 981512 | 63 | 474381 | 769 | 525619 | 24 |
| 37 | 453316 | 705 | 981474 | 63 | 474842 | 768 | 525158 | 23 |
| 38 | 456739 | 704 | 981436 | 63 | 475303 | 767 | 524697 | 22 |
| 39 | 457162 | 704 | 981399 | 63 | 475763 | 767 | 524237 | 21 |
| 40 | 457584 | 703 | 981361 | 63 | 476223 | 766 | 523777 | 20 |
| 41 | 9.458006 | 702 | 9.981323 | 63 | 9.476683 | 765 | 10.523317 | 19 |
| 42 | 458427 | 701 | 981285 | 63 | 477142 | 765 | 522858 | 18 |
| 43 | 458848 | 701 | 981247 | 63 | 477601 | 764 | 522399 | 17 |
| 44 | 459268 | 700 | 981209 | 63 | 478059 | 763 | 521941 | 16 |
| 45 | 459688 | 699 | 981171 | 63 | 478517 | 763 | 521483 | 15 |
| 46 | 460108 | 698 | 981133 | 64 | 478975 | 762 | 521025 | 14 |
| 47 | 460527 | 698 | 981095 | 64 | 479432 | 761 | 520568 | 13 |
| 48 | 460946 | 697 | 981057 | 64 | 479889 | 761 | 520111 | 12 |
| 49 | 461364 | 696 | 981019 | 64 | 480345 | 760 | 519655 | 11 |
| 50 | 461782 | 695 |  | 64 | 480801 | 7509 | 519199 | 10 |
| 51 | 9-462199 | 695 | 9-9809 21 | 64 | 9.481257 | $7{ }^{\text {¢ }}$ 9 | 10.518743 | 8 |
| 52 | 462616 | 694 | 950904 | 64 | 481712 | 758 | 518288 | 8 |
| 53 | 463032 | 693 | 950866 | 64 | $4 \mathrm{~S}_{2167}$ | 757 | 517833 | 7 |
| 54 | 463448 | 693 | 980827 | 64 | 482621 | 757 | 517379 | 6 |
| 55 | 463864 | 692 | 990789 | 64 | 453075 | 756 | 516925 | 5 |
| 56 | 464279 | 691 | 980750 | 64 | 453529 | 755 | 516471 | 4 |
| 57 | 464694 | 690 | 980712 | 64 | 483982 | 755 | 516018 | 3 |
| 58 | 465108 | 690 | 980673 | 6.4 | 484435 | 754 | 515565 | 2 |
| 59 | 465522 | 689 | 990635 | 64 | 484887 | 753 | 515113 | 1 |
| 60 | 465935 | 683 | 980596 | 64 | 485339 | 753 | 514661 | 0 |
| 1 | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | 1 |
| $106^{\circ}$ |  |  |  |  |  |  |  | $73^{\circ}$ |


| Table II. |  | LOGARITHMIC |  | SINES | TANGENTS, ETC. |  |  | 35 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $17^{\circ}$ |  |  |  |  |  |  | $162^{\circ}$ |  |
| , | Sine. | D. | Cosine. | D. | Tang. | D. | Cotang. | , |
| 0 | 9.465935 | 688 | 9.9805 66 | 64 | ¢.485339 | 755 | 10.514661 | 60 |
| 1 | 466348 | 688 | - 980558 | 64 | 485791 | 752 | 514209 | 59 |
| 2 | 466761 | 687 | 980519 | 65 | 486242 | 751 | 513758 | 58 |
| 3 | 467173 | 686 | 980480 | 65 | 486693 | 751 | 513307 | 57 |
| 4 | 467585 | 685 | 980442 | 65 | 487143 | 750 | 512857 | 56 |
| 5 | 467996 | 685 | 980403 | 65 | 487593 | 749 | 512407 | 55 |
| 6 | 468407 | 684 | 980364 | 65 | 488043 | 749 | $5 \mathrm{L1} 9^{5} 7$ | 54 |
|  | 468817 | 683 | 980325 | 65 | 488492 | 748 | 511508 | 53 |
| 8 | 469227 | 683 | 980286 | 65 | 488941 | 747 | 511059 | 52 |
| 9 | 469637 | 682 | 980247 | 65 | 489390 | 747 | $5: 0610$ | 51 |
| 10 | 470046 | 681 | 980208 | 65 | 489838 | 746 | 510162 | 50 |
| 11 | 9.470455 | 680 | 9.980169 | 65 | 9.490286 | 746 | 10.50971/4 | 49 |
| 12 | 470863 | 680 | 980130 | 65 | 490733 | 745 | 509257 | 48 |
| 13 | 471271 | 679 | 980091 | 65 | 491180 | 744 | 508820 | 47 |
| 14 | 471679 | 678 | 980052 | 65 | 491627 | 744 | 508373 | 46 |
| 15 | 472086 | 678 | 980012 | 65 | 492073 | 743 | 507927 | 45 |
| 16 | 472492 | 677 | 979973 | 65 | 492519 | 743 | 507481 | 44 |
| 17 | 472898 | 676 | 979934 | 66 | 492965 | 742 | 507035 | 43 |
| 18 | 473304 | 676 | 979895 | 66 | 493410 | 741 | 506590 | 42 |
| 19 | 473710 | 675 | 979855 | 66 | 493854 | 740 | 506146 | 41 |
| 20 | 474115 | 674 | 979816 | 66 | 474299 | 740 | 505701 | 40 |
| 21 | 9.474519 | 674 | 9.979776 | 66 | 9.494743 | 740 | 10.505257 | 39 |
| 22 | 474923 | 673 | 979737 | 66 | 495186 | 739 | 504814 | 38 |
| 23 | 475327 | 672 | 979697 | 66 | 495630 | 738 | 504370 | 37 |
| 24 | 475730 | 672 | 979658 | 66 | 496073 | 737 | 503927 | 36 |
| 25 | 476133 | 671 | 979618 | 66 | 496515 | 737 | 503485 | 35 |
| 20 | 476536 | 670 | 979579 | 66 | 496957 | 736 | 503043 | 34 |
| 27 | 476938 | 669 | 979539 | 66 | $\therefore 97399$ | 736 | 502601 | 33 |
| 28 | 477340 | 669 | 979499 | 66 | 497841 | 735 | 502159 | 32 |
| 29 | 477741 | 668 | 979459 | 66 | 498282 | 734 | 501718 | 3 I |
| 30 | 478142 | 667 | 979420 | 66 | 498722 | 734 | 501278 | 30 |
| 31 | $9 \cdot 478542$ | 667 | 9.979380 | 66 | 9.499163 | 733 | 10.500837 | 29 |
| 32 | 478042 | 666 | 979340 | 66 | 499603 | 733 | 500397 | 28 |
| 33 | 479342 | 665 | 979300 | 67 | 500042 | 732 | 499958 | 27 |
| 34 | 479741 | 665 | 979260 | 67 | 500481 | 731 | 499519 | 26 |
| 35 | 480140 | 664 | 979220 | 67 | 500920 | 731 | 499080 | 25 |
| 36 | 480539 | 663 | 979180 | 67 | 501359 | 730 | 498641 | 24 |
| 37 38 | 480037 | 663 | 979140 | 67 | 501797 | 730 | 498203 | 23 |
| 38 | 48 I 334 | 662 | 979100 | 67 | 502235 | 729 | 497765 | 22 |
| 39 | 481731 | 661 | 979059 | 67 | 502672 | 728 | 497328 | 21 |
| 40 | 482128 | 661 | 979019 | 67 | 503109 | 728 | 496891 | 20 |
| 41 | 9.482525 | 660 | 9•978979 | 67 | 9.503546 | 727 | 10.496454 | 19 |
| 42 | 482921 | 659 | $978{ }^{3} 9$ | 67 | 503982 | 727 | 496018 | 18 |
| 43 | 483316 | 659 | 978898 | 67 | 504418 | 726 | 495582 | 17 |
| 44 | 483712 | 658 | 978858 | 67 | 504854 | 725 | 495146 | 16 |
| 45 | 484107 | 657 | 978817 | 67 | 505289 | 725 | 494711 | 15 |
| 46 | 484501 | 657 | 978777 | 67 | 505724 | 724 | 494276 | 14 |
| 47 | $48489^{5}$ | 656 | 978737 | 67 | 506159 | 724 | 493841 | 13 |
| 48 | 485289 | 655 | 978696 | 68 | 506593 | 723 | 493407 | 12 |
| 49 | 485682 | 655 | 978655 | 68 | 507027 | 722 | 492973 | 11 |
| 50 | 486075 | 654 | 978615 | 68 | 507460 | 722 | 492540 | 10 |
| 51 | 9.486467 |  | $9 \cdot 978574$ | 68 |  | 721 |  |  |
| 52 53 54 | 486860 | 653 | 978533 | 68 | 508326 | 721 | 491674 | 8 |
| 53 | 487251 | 652 | 978493 | 68 | 508759 | 720 | 491241 | 7 |
| 54 55 | 487643 | 651 | 978452 | 68 | 509191 | 719 | 490809 | 6 |
| 55 | 488034 | 651 | 978411 | 68 | 509622 | 719 | 490378 | 5 |
| 56 | 488424 | 650 | 978370 | 68 | 510054 | 718 | 489946 | 4 |
| 53 | 488814 | 650 649 | 978329 978288 | 68 68 | 510485 510916 | 718 | 489515 | 3 |
| 59 | 48959 | 648 | 978247 | 68 | 511346 | 716 | 488654 | 1 |
| 60 | 489982 | 648 | 978206 | 68 | 511776 | 716 | 488224 | 0 |
| 1 | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | , |
| $107^{\circ}$ |  |  |  |  |  |  |  | $72^{\circ}$ |


| 36 |  | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  | Table II. |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $18^{\circ}$ |  |  |  |  |  |  | $161^{\circ}$ |  |
| , | Sine. | D. | Cosine | D. | Tang. | D. | Cotang. | 1 |
| 0 | 9.489982 | 648 | S.978206 | 68 | 9.511776 | 716 | 10.488224 | 60 |
| 1 | 490371 490750 | 648 | 978165 978124 | 68 | 512206 512635 | 716 | 487794 487365 | 50 58 50 |
| $\stackrel{2}{3}$ | 490759 491147 | 647 646 | 978124 | 68 69 | 512635 513064 | 715 | 486936 | 5 |
| 4 | 491535 | 646 | 978042 | 69 | 513493 | 714 | 486507 | 56 |
| 5 | 491922 | 645 | 978001 | 69 | 5 I 921 | 713 | 486079 | 55 |
| 6 | 492308 | 644 | 977959 | 69 | 514349 | 713 | 485651 | 54 |
| 7 | 492605 | 644 | 977918 | S9 | 514777 | 712 | 485223 | 53 |
| 8 | 493081 | 643 | 977877 | 69 | 515204 | 712 | 484796 | 52 |
| 9 | 493466 | 642 | 977835 | 69 | 515631 | 711 | 484369 | 51 |
| 10 | 493851 | 642 | 977794 | 69 | 516057 | 710 | 483943 | 50 |
| 11 | 9-494236 | 641 | 9-977752 | 69 | 9.516484 | 710 | 10.483516 | 49 |
| 12 | 494621 | 641 | 977711 | 69 | 516910 | 709 | 483090 | 48 |
| 13 | 495003 | 640 | 977669 | 69 | ${ }^{2} 17335$ | 709 | 482665 | 47 |
| 14 | 495388 | 639 | 977628 | 69 | 517761 | 708 | 482239 | 46 |
| 15 | 495772 | 639 | 977586 | 69 | 518186 | 708 | 481814 | 45 |
| 16 | 496154 | 638 | 977544 | 70 | 518610 | 707 | 481390 | 44 |
| 17 | 496537 | 637 | 977503 | 70 | 519034 | 706 | 480066 | 43 |
| 18 | 496919 | 637 | 977461 | 70 | 519458 | 706 | 480542 | 42 |
| 19 | 497301 | 636 | 977419 | 70 | 519882 | 705 | 480118 | 41 |
| 20 | 497682 | 636 | 977377 | 70 | 520305 | 705 | 479695 | 40 |
| 21 | $9 \cdot 498064$ | 635 | 9.977335 | 70 | 9.520728 | 704 | 10.479272 | 39 |
| 22 | 498444 | 634 | 977293 | 70 | 521151 | 703 | 478849 | 38 |
| 23 | 498825 | 634 | 977251 | 70 | 521573 | 703 | 478427 | 37 |
| 24 | 499204 | 633 | 977209 | 70 | 521995 | 703 | 478005 | 36 |
| 25 | 499584 | 632 | 977167 | 70 | 522417 | 702 | 477583 | 35 |
| 26 | 499963 | 632 | 977125 | 70 | 522838 | 702 | 477162 | 34 |
| 27 | 500342 | 631 | 977083 | 70 | 523259 | 701 | 476741 | 33 |
| 28 | 500721 | 631 | 977041 | 70 | 523680 | 701 | 476320 | 32 |
| 29 | 501099 | 630 | 976999 | 70 | 524100 | 700 | 475900 | 3 I |
| 30 | 501476 | 629 | 976957 | 70 | 524520 | 699 | 475480 | 30 |
| 3 I | 9.501854 | 629 | 9.976914 | 70 | 9.524940 | 699 | 10.475060 | 29 |
| 32 | 502231 | 628 | 976872 | 71 | 525350 | 698 | 474641 | 28 |
| 33 | 502607 | 628 | 976830 | 71 | 525778 | 698 | 474222 | 27 |
| 34 | 502984 | 627 | 976787 | 71 | 526197 | 697 | 473803 | 26 |
| -35 | 503360 | 626 | 976745 | 71 | 526615 | 697 | 473385 | 25 |
| 36 | 503735 | 626 | 976702 | 71 | 527033 | 696 | 472967 | 24 |
| 37 | 504110 | 625 | 976660 | 71 | 527451 | 696 | 472549 | 23 |
| 38 | 504485 | 625 | 976617 | 71 | 527868 | 695 | 472132 | 22 |
| 39 | 50,460 | 624 | 976574 | 71 | 528285 | 695 | 471715 | 21 |
| 40 | 505234 | 623 | 976532 | 71 | 528702 | 694 | 471298 | 20 |
| 41 | S. 505608 | 623 | 9.976489 | 71 | 9.529112 | 693 | 10.47088 I | 19 |
| 42 | 505981 | 622 | 976446 | 71 | 529535 | 693 | 470465 | 18 |
| 43 | 506354 | 622 | 976404 | 71 | 529051 | 693 | 470049 | 17 |
| 44 | 506727 | 621 | 976361 | 71 | 530366 | 692 | 469634 | 16 |
| 45 | 507099 | 620 | 976318 | 71 | 530781 | 691 | 469219 | 15 |
| 46 | 507471 | 620 | 976275 | 71 | 531196 | 691 | 468804 | 14 |
| 47 | 507843 | 619 | 976232 | 72 | 531611 | 690 | 468389 | 13 |
| 48 | 508214 | 619 | 976189 | 72 | 532025 | 690 | 467075 | 12 |
| 49 | 508585 | 618 | 976146 | 72 | 532439 | 689 | 467561 | 11 |
| 50 | JoSg5̋6 | 618 | 976103 | 72 | 532853 | 689 | 467147 | 10 |
| 51 | 9.50,326 | 617 | 9.976060 | 72 | 9. 533266 | 688 | 10.466734 |  |
| 52 53 53 | 509696 | 616 | 976017 | 72 | 533679 | 688 | 466321 | 8 |
| 53 | 510065 | 616 | 975974 | 72 | 534092 | 687 | 465908 | 7 |
| 54 55 | 510434 | 615 | 975030 | 72 | 534504 | 687 | 465406 | 6 |
| 55 | 510803 | 615 | 975887 | 72 | 534916 | 686 | 465084 | 5 |
| 56 | 511172 | 614 | 975844 | 72 | 535328 | 686 | 464672 | 4 |
| 57 58 58 | 511540 | 613 | 975800 | 72 | 535739 | 685 | 464261 | 3 |
| 58 | 511907 512275 | 613 | 975757 | 72 | 536150 | 685 | 463850 | 2 |
| 59 60 | 512275 | 612 | 975714 | 72 | 536561 | 684 | 463.439 | 1 |
| 60 | 512642 | 612 | 975070 | 72 | 536972 | 684 | 463028 | 0 |
| , | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | , |
| $108^{\circ}$ |  |  |  |  |  |  |  | $1^{0}$ |


| Table II. |  | LOG.ARITHMIC SINES, TANGENTS, ETC. |  |  |  |  |  | 37 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $19^{\circ}$ |  |  |  |  |  |  |  | $160^{\circ}$ |
| , | Sine. | D. | Cosine. | D. | Tang. | D. | Cotang. | , |
| 0 | 9.512642 | 612 | 9.9755670 | 73 | 9.536972 | 684 | 10.463028 | 60 |
| 1 | 513009 | 611 | 975627 | 73 | 537382 | 683 | 462618 | 59 |
| 2 | 513375 | 611 | 975583 | 73 | 537792 | 683 | 462208 | 58 |
| 3 | 513741 | 610 | 975539 | 73 | 538202 | 682 | 461798 | 57 |
| 4 | 514107 | 609 | 975426 | 73 | 538611 | 682 | 461389 | 56 |
| 5 | 514472 | 609 | 975452 | 73 | 5390:0 | 681 | 460980 | 55 |
| 6 | 514837 | 608 | 975408 | 73 | 539429 | 681 | 460571 | 54 |
| 7 | 5.5202 | 608 | 975365 | 73 | 539837 | 680 | 460163 | 53 |
| 8 | 515566 | 607 | 975321 | 73 | 540245 | 680 | 459755 | 52 |
| 9 | 515930 | 607 | 975277 | 73 | 540653 | 679 | 459347 | 51 |
| 10 | 516294 | 606 | 975233 | 73 | 541061 | 679 | 458939 | 50 |
| 11 | $9 \cdot 516657$ | 605 | \%.975189 | 73 | 9.541468 | 678 | 10.458532 | 49 |
| 12 | 517020 | 605 | 975145 | 73 | 541875 | 678 | 458125 | 48 |
| 13 | 517382 | 604 | 975101 | 73 | 542281 | 677 | 457719 | 47 |
| 14 | 517745 | 604 | 975057 | 73 | 542688 | 677 | 457312 | 46 |
| 15 | 518107 | 603 | 975013 | 73 | 543094 | 676 | 456906 | 45 |
| 16 | 518468 | 603 | 974969 | 74 | 543499 | 676 | 456501 | 44 |
| 17 | 518829 | 602 | 974225 | 74 | 543905 | 675 | 456095 | 43 |
| 18 | 519190 | 601 | 974880 | 74 | 544310 | 675 | 455690 | 42 |
| 19 | 5 c 95ı | 601 | 974836 | 74 | 544715 | 674 | 455285 | 41 |
| 20 | 519911 | 600 | 974792 | 74 | 545119 | 674 | 454881 | 40 |
| 21 | $9 \cdot 520271$ | 600 | 9.974748 | 74 | 9.545524 | 673 | 10.454476 | 39 |
| 22 | 520631 | 599 | 974703 | 74 | 545928 | 673 | 454072 | 38 |
| 23 | 520990 | 599 | 974659 | 74 | 546331 | 672 | 453669 | 37 |
| 24 | 521349 | 598 | 974614 | 74 | 546735 | 672 | 453265 | 36 |
| 25 | 521707 | 598 | 974570 | 74 | 547138 | 671 | 452862 | 35 |
| 26 | 522066 | 597 | 974525 | 74 | 547540 | 671 | 452460 | 34 |
| 27 | 522424 | 596 | 974481 | 74 | 547943 | 670 | 452057 | 33 |
| 28 | 522781 | 596 | 974436 | 74 | 548345 | 670 | 451655 | 32 |
| 29 | 523138 | 595 | 974391 | 74 | 548747 | 669 | 451253 | 31 |
| 30 | $52349^{5}$ | 595 | 974347 | 75 | 549149 | 669 | 450851 | 30 |
| 31 | 9.523852 | 594 | 9-974302 | 75 | 9.549550 | 668 | 10.450450 | 29 |
| 32 | 9 524208 | 594 | 974257 | 75 | 549951 | 668 | 450049 | 28 |
| 33 | 524564 | 593 | 974212 | 75 | 550352 | 667 | 449648 | 27 |
| 34 | 524920 | 503 | 974167 | 75 | 550752 | 667 | 449248 | 26 |
| 35 | 525275 | 592 | 974122 | 75 | 551153 | 666 | 448847 | 25 |
| 36 | 525630 | 591 | 974077 | 75 | 551552 | 666 | 448448 | 24 |
| 37 | 525984 | 591 | 974032 | 75 | 551952 | 665 | 448048 | 23 |
| 38 | 526339 | 590 | 973987 | 75 | 552351 | 665 | 447649 | 22 |
| 39 | 526693 | 500 | 973942 | 75 | 552750 | 665 | 447250 | 21 |
| 40 | 527046 | 589 | 973897 | 75 | 553149 | 664 | 446851 | 20 |
| 41 | G. 527400 | 589 | $9 \cdot 973852$ | 75 | 9.553548 | 664 | 10.446452 | 19 |
| 42 | 527753 | 588 | 973807 | 75 | 553946 | 663 | 446054 | 18 |
| 43 | 528105 | 588 | 973761 | 75 | 554344 | 663 | 445656 | 17 |
| 44 | 528458 | 587 | 973716 | 76 | 554741 | 662 | 445259 | 16 |
| 45 | 528810 | 587 | 973671 | 76 | 555139 | 662 | 444861 | 15 |
| 46 | 529161 | 586 | 973625 | 76 | 555536 | 661 | 444464 | 14 |
| 47 | 529513 | 586 | 973580 | 76 | 555933 | 661 | 444067 | 13 |
| 48 | 529864 | 585 | 973535 | 76 | 556329 | 660 | 443671 | 12 |
| 49 | 530215 | 585 | 973489 | 76 | 556725 | 660 | 443275 | 11 |
| 50 | 530565 | 584 | 973444 | 76 | 557121 | 659 | 442879 | 10 |
| 51 |  |  | 9.973398 | 76 |  |  |  |  |
| 52 | $531265$ | 583 | 973352 | 76 | 557913 | 659 | 442087 | 8 |
| 53 | 531614 | 582 | 973307 | 76 | 558308 | 658 | 441692 | 7 |
| 54 | 531963 | 582 | 973261 | 76 | 558703 | 658 | 441297 | 6 |
| 55 | $532312$ | 581 | 973215 | 76 | 559097 | 657 | 440903 | 5 |
| 56 | 532661 | 581 | 973169 | 76 | 559491 | 657 | 440509 | 4 |
| 57 | 533009 | 580 | 973124 | 76 | 559885 | 656 | 440115 | 3 |
| 58 | 533357 | 580 | 973078 | 76 | 560279 | 656 | 439721 | 2 |
| 59 | 533704 | 579 | 973032 | 77 | 560673 | 655 | 439327 | 1 |
| 60 | 534052 | 578 | 972986 | 77 | 561066 | 655 | 438934 | 0 |
| 1 | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | , |
| $109^{\circ}$ |  |  |  |  |  |  |  | $70^{\circ}$ |



| Table IL. |  | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  |  | 39 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $21^{\circ}$ |  |  |  |  |  |  |  | $158^{\circ}$ |
| , | Sine. | D. | Cosine. | D. | Tang. | D. | Cotang. | , |
| 01234567891010 | $\begin{array}{r} 9.554329 \\ 554658 \\ 554987 \\ 555315 \\ 555643 \\ 555971 \\ 556299 \\ 556626 \\ 556953 \\ 557280 \\ 557606 \end{array}$ | 548 | 9.970152 | 81 | 9.584177 | 629 | 10.415823 | 60 |
|  |  | 548 | 970103 | 81 | 584555 | 629 | 415445 | 59 |
|  |  | 547 | 970055 | 81 | 584932 | 638 | 415068 | 58 |
|  |  | 547 | 970006 | 81 | 585309 | 628 | 414691 | 57 |
|  |  | 546 | 969957 | 81 | 585686 | 627 | 414314 | 56 |
|  |  | 546 | 969909 | 81 | 586062 | 627 | 413938 | 55 |
|  |  | 545 | 969860 | 81 | 584439 | 627 | 413551 | 54 |
|  |  | 545 | 969811 | 81 | 586815 | 626 | 413185 | 53 |
|  |  | 544 | 969762 | 81 | 587190 | 626 | 412810 | 52 |
|  |  | 5.44 | 969714 | 81 | 587566 | 625 | 412434 | 51 |
|  |  | 543 | 969665 | 81 | 587941 | 625 | 412059 | 50 |
| 11 | 9.557932 | 5.43 | 9.969616 | 82 | 9.588316 | 625 | 10.411684 | 49 |
| 12 | $9 \cdot 557932$ 558258 5 | 543 | 969507 | 82 | 588691 | 624 | 411309 | 48 |
| 13 |  | 542 | 969518 | 82 | 589056 | 624 | 410034 | 47 |
| 14 |  | 542 | 969469 | 82 | 589440 | 623 | 410560 | 46 |
| 15 | 558909 559234 | 541 | 969420 | 82 | 589814 | 623 | 410186 | 45 |
| 16 | $559558$ | 541 | 969370 | 82 | 590188 | 023 | 409812 | 44 |
| 17 | 559883 | 540 | 969321 | 82 | 590562 | 622 | 409438 | 43 |
| 18 | 560207 56053 I | 540 | 969272 | 82 | 590935 | 622 | 409065 | 42 |
| 19 |  | 539 | 969223 | 82 | 591308 | 622 | 408692 | 41 |
| 20 | $\begin{aligned} & 56053 \mathrm{I} \\ & 560855 \end{aligned}$ | 539 | 969173 | 82 | 591681 | 621 | 408319 | 40 |
| 21 | 9.561178 | 538 | 9.969124 | 82 | 9.592054 | 621 | 10.407946 | 39 |
| 22 | $\begin{aligned} & 561501 \\ & 561824 \end{aligned}$ | 538 | 969075 | 82 | 592426 | 620 | 407574 | 38 |
| 23 |  | 537 | 969025 | 82 | 592799 | 620 | 407201 | 37 |
| 24 | $\begin{aligned} & 561824 \\ & 562146 \end{aligned}$ | 537 | 968976 | 82 | 593171 | 619 | 406829 | 36 |
| 25 | 562468 | 536 | 968926 | 83 | 593542 | 619 | 406458 | 35 |
| 26 | 562790 | 536 | 968877 | 83 | 593914 | 618 | 406086 | 34 |
| 27 | 563112 | 536 | 968827 | 83 | 594285 | 618 | 405715 | 33 |
| 28 | 56343356375 | 535 | 968777 | 83 | 594656 | 618 | 405344 | 32 |
| 29 |  | 535 | 968728 | 83 | 595027 | 617 | 404973 | 31 |
| 30 | $\begin{aligned} & 563755 \\ & 564075 \end{aligned}$ | 534 | 968678 | 83 | 595398 | 617 | 404602 | 30 |
| 3ı | 9.564396 | 534 | 9.968628 | 83 | 9.595768 | 617 | 10.404232 | 29 |
| 32 | ${ }^{564716}$ | 533 | 968578 | 83 | 596138 | 616 | 403862 | 28 |
| 33 | 565036 | 533 | 968528 | 83 | 596508 | 616 | 403492 | 27 |
| 34 | 565356 | 532 | 968479 | 83 | 596878 | 616 | 403122 | 26 |
| 35 | 565676 | 532 | 968429 | 83 | 597247 | 615 | 402753 | 25 |
| 36 | 565995 | 531 | 968379 | 83 | 597616 | 615 | 402384 | 24 |
| 37 | 566314 | 531 | 968329 | 83 | 597985 | 615 | 402015 | 23 |
| 38 | 566632 | 53 I | 968278 | 83 | 598354 | 614 | 401646 | 22 |
| 39 | $\begin{aligned} & 566951 \\ & 567269 \end{aligned}$ | 530 | 968228 | 84 | 598722 | 614 | 401278 | 21 |
| 40 |  | 530 | 968178 | 84 | 599091 | 613 | 400909 | 20 |
| 41 | 9.567587 | 529 | 9.968128 | 84 | 9.599459 | 613 | 10.400541 | 19 |
| 42 | $\begin{aligned} & 567904 \\ & 568222 \end{aligned}$ | 529 | 968078 | 84 | 599827 | 613 | 400173 | 18 |
| 43 |  | 528 | 968027 | 84 | 600194 | 612 | 399806 | 17 |
| 44 | - 568539 | 528 | 967977 | 84 | 600562 | 612 | 399438 | 16 |
| 45 | 568856 | 528 | 967927 | 84 | 600929 | 611 | 399071 | 15 |
| 46 | 569172 | 527 | 967876 | 84 | 601296 | 611 | 398704 | 14 |
| 47 | 569488 | 527 | 967826 | 84 | 601663 | 611 | 398337 | 13 |
| 48 , | 569804570120 | 526 | 967775 | 84 | 602029 | 610 | 397971 | 12 |
| 49 |  | 526 | 967725 | 84 | 602395 | 610 | 397605 | 11 |
| 50 | $\begin{aligned} & 570120 \\ & 570435 \end{aligned}$ | 525 | 967674 | 84 | 602761 | 61 | 397239 | 10 |
| 51 | $9 \cdot 570751$ | 525 | $9 \cdot 967624$ | 84 | 9.603127 | 609 | 10.396873 |  |
| 52 | 571066 | 524 | 967573 | 84 | $60349^{3}$ | 609 | 396507 | 8 |
| 53 | 571380 | 524 | 967522 | 85 | 603858 | 609 | 396142 | 7 |
| 54 | 5771695 571600 | 523 | 967471 | 85 | 604223 | 608 | 395777 | 6 |
| 55 | 572009572323 | 523 | 967421 | 85 | 604588 | 608 | 395412 | 5 |
| 56 |  | 523 | 967370 | 85 | 604953 | 607 | 395047 | 4 |
| 57 | 572636 | 522 | 967319 | 85 | 605317 | 607 | 394683 | 3 |
| 58 | 572950573263 | 522 | 967268 | 85 | 605682 | 607 | 394318 | 2 |
| 59 |  | 521 | 967217 | 85 | 606046 | 606 | 393954 | 1 |
| 60 | 573575 | 521 | 967166 | 85 | 606410 | 606 | 393590 | 0 |
|  | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | , |
| $111^{\circ}$ |  |  |  |  |  |  |  | $68^{\circ}$ |




| 42 |  | LOGARITHMIC SLNES，TANGENTS．ETC． |  |  |  |  | Tabl |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $24^{\circ}$ |  | D | Cosine． | D． | Tang． | D． | $155^{\circ}$ |  |
| ， | Sine． |  |  |  |  |  | Cotang． | 1 |
|  | 9．609313 | 473 | 9．960730 | 94 | 9.648583 | 566 | 10.351417 | 60 |
| 0123456789 | 609507 | 4,2 | 960674 | 94 | 648923 | 566 | 351077 | 59 |
|  | 609880 | 472 | 960618 | 94 | 649263 | 566 | 350737 | 58 |
|  | 610164 | 472 | 960561 | 94 | 649602 | 566 | 350308 | 57 |
|  | 610447 | 471 | 960505 | 94 | 649942 | 565 | 350058 | 56 |
|  | 610729 | 471 | 960448 | 94 | 650281 | 565 | 349719 | 55 |
|  | 611012 | 470 | 960392 | 94 | 650620 | 565 | 349380 | 54 |
|  | 611294 | 470 | 960335 | 94 | 650959 | 564 | 349041 | 53 |
|  | 611576 | 470 | 960279 | 94 | 651297 | 564 | 348703 | 52 |
|  | 611858 | 469 | 960222 | 94 | 651636 | 564 | 348364 | 51 |
| 10 | 612140 | 469 | 960165 | 94 | 651974 | 563 | 348026 | 50 |
| II | 9．612421 | 469 | $9 \cdot 960109$ | 95 | G．652312 | 563 | 10.347688 | 49 |
| 12 | 612702 | 468 | 960052 | 95 | 652650 | 563 | 347350 | 48 |
| 13 | 612983 | 468 | 959995 | 95 | 652988 | 563 | 347012 | 47 |
| 14 | 613264 | 467 | 950938 | 95 | 653326 | 562 | 346674 | 46 |
| 15 | 613545 | 467 | 959882 | 95 | 653663 | 562 | 346337 | 45 |
| 16 | 613825 | 467 | 959825 | 95 | 654000 | 562 | $3: 6000$ | 44 |
| 17 | 614105 | 466 | 959768 | 95 | 654337 | 561 | 345663 | 43 |
| 18 | 614385 | 466 | 9 937 11 | 95 | 654674 | 561 | 345326 | 42 |
| 19 | 614665 | 466 | 950654 | 95 | 655011 | 561 | 344989 | 41 |
| 20 | 614944 | 465 | 959596 | 95 | 655348 | 561 | 344652 | 40 |
|  | 9．615223 | 465 | $9 \cdot 959539$ | 95 | 9.655684 | 560 | 10.344316 | 39 |
| 21 | 615502 | 465 | 9 9⿹勹⿰丿丿 | 95 | 656020 | 560 | 343980 | 38 |
| 23 | 615781 | 464 | 959425 | 95 | 656356 | 560 | 343644 | 37 |
| 24 | 616060 | 464 | 9 95368 | 95 | 656692 | 559 | 343308 | 36 |
| 25 | 616338 | 464 | $9{ }^{\text {jog }} 310$ | 96 | 6507028 | 50.9 | 342972 | 35 |
| 26 | 616616 | 463 | 959253 | 96 | 657364 | 559 | 342636 | 34 |
| 27 | 616894 | 463 | 959195 | 96 | 657699 | 559 | 342301 | 33 |
| 28 | $61717^{2}$ | 462 | 959138 | 96 | 658034 | 558 | 341966 | 32 |
| 2930 | 617450 | 462 | 959080 | 96 | 658369 | 558 | 341631 | 31 |
|  | 617727 | 462 | 959023 | 96 | 658704 | 558 | 341296 | 30 |
| $\begin{aligned} & 31 \\ & 32 \\ & 33 \\ & 34 \\ & 35 \\ & 36 \\ & 37 \\ & 38 \\ & 39 \\ & 40 \end{aligned}$ | 9．618004 | 461 | 9．958965 | 96 | 9.659039 | 558 | 10.340961 | 29 |
|  | 618281 | 461 | 958908 | 96 | 659373 | 557 | 340627 | 28 |
|  | 618535 | 461 | 958850 | 96 | 659708 | 557 | 340292 | 27 |
|  | 618834 | 460 | 958792 | 96 | 6600.12 | 557 | 339958 | 26 |
|  | 61911 C | 460 | 958734 | 66 | 660376 | 557 | 339624 | 25 |
|  | 619386 | 460 | 958677 | 96 | 660710 | 556 | 339290 | 24 |
|  | 619662 | 459 | 958619 | 96 | 661043 | 550 | 338957 | 23 |
|  | 619938 | 409 | 958561 | 96 | 661377 | 556 | 338623 | 22 |
|  | 620213 | $4{ }^{40}$ | 958503 | 97 | 661710 | 550 | 338290 | 21 |
|  | 620488 | 458 | 958445 | 97 | 662043 | 555 | 337957 | 2 C |
| 41 | 9．620－63 | 458 | $9 \cdot 958387$ | 97 | 9．662376 | 555 | 10.337624 | 19 |
| 42 | 621038 | 457 | 958329 | 97 | 662709 | 554 | 337291 | 18 |
| 43 | 621313 | 457 | 958271 | 97 | 663042 | 554 | 336958 | 17 |
| 44 | 621587 | 457 | 958213 | 97 | 663375 | 554 | 336525 | 16 |
|  | 621861 | 450 | 958154 | 97 | 663707 | 554 | 336293 | 15 |
| 45 | 622135 | 435 | 953096 | 97 | 664039 | 553 | 335061 | 14 |
| 47 | 622409 | $4{ }^{56}$ | 958038 | 97 | 664371 | 553 | 335629 | 13 |
| 48 | 622682 | 455 | 957979 | 97 | 664703 | 553 | 335297 | 12 |
| 49 | 622956 | 455 | 957921 | 97 | 665035 | 553 | 334965 | 11 |
|  | 623229 | 450 | 957863 | 97 | 665366 | 552 | 334634 | 10 |
| $\begin{aligned} & 51 \\ & 52 \\ & 53 \\ & 54 \\ & 55 \\ & 56 \\ & 57 \\ & 58 \\ & 59 \\ & 60 \end{aligned}$ | $9.6235{ }^{\text {c }}$ | 454 | 9．957804 | 97 | ¢． 665698 | 552 | 10.334302 | $\begin{aligned} & 9 \\ & 7 \\ & 7 \\ & 6 \\ & 5 \\ & 4 \\ & 3 \\ & 2 \\ & 1 \\ & 0 \end{aligned}$ |
|  | 623774 | 45 | 950776 | 98 | 666029 | 552 | 333971 |  |
|  | 624047 | 454 | 957687 | 98 | 666360 | 551 | 333640 |  |
|  | 624319 | 453 | 957628 | 98 | 666691 | 551 | 333309 |  |
|  | 624501 | $4 \mathrm{~S}^{3}$ | 950500 | 98 | 667021 | 551 | 332979 |  |
|  | 624863 | 453 | 950511 | 98 | 667352 | 551 | 332648 |  |
|  | 625135 | $4{ }^{3} 2$ | 957452 | 98 | 667682 | 550 | 332318 |  |
|  | 625406 | 452 | 9 ¢̄7393 | 98 | 668013 | 550 | 331987 |  |
|  | 625677 | 452 | 957335 | 98 | 668343 | 550 | 331657 |  |
|  | 625948 | 451 | 957276 | 98 | 668673 | 550 | 331327 |  |
| 1 | Cosine． | D． | Sine． | D． | Cotang． | D． | Tang． | ＇ |
| $114^{\circ}$ |  |  |  |  |  |  |  | ${ }^{\circ} 5^{\circ}$ |


| Table II. |  | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  |  | 43 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $25^{\circ}$ |  |  |  |  |  |  |  | $154^{\circ}$ |
| , | Sine. | D. | Cosine. | D. | Tang. | D. | Cotang. | , |
| 0 | 9.625948 | 451 | $9 \cdot 957.276$ | 98 | 9.668673 | 550 | 10.331327 | 60 |
| 1 | 626219 | 451 | 957217 | 98 | 669002 | 549 | 330998 | 5 |
| 2 | 626490 | 451 | 957158 | 98 | 669332 | 549 | 330668 | 58 |
| 3 | 626760 | 450 | 957099 | 98 | 669661 | 549 | 330339 | 57 |
| 4 | 627030 | 450 | 957040 | 98 | 669991 | 548 | 330009 | 56 |
| 5 | 627300 | 450 | 956981 | 98 | 670320 | 548 | 329680 | 55 |
| 6 | 627570 | 449 | 956921 | 99 | 670649 | 548 | 329351 | 54 |
| 7 | 627840 | 449 | 956862 | 99 | 670977 | 548 | 329023 | 53 |
| 8 | 628109 | 449 | 956803 | 99 | 671306 | 547 | 328694 | 52 |
| 9 | 628378 | 448 | 956744 | 99 | 671635 | 547 | 328365 | 5 I |
| 10 | 628647 | 448 | 956684 | 99 | 671963 | 547 | 328037 | 50 |
| 11 | 9.628916 | 457 | 9.956625 | 99 | 9.672291 | 547 | 10. 327700 | 49 |
| 12 | 629185 | 447 | 956566 | 99 | 672619 | 546 | 327381 | 48 |
| 13 | 629453 | 447 | 956506 | 99 | 672947 | 546 | 327053 | 47 |
| 14 | 629721 | 446 | 956447 | 99 | 673274 | 546 | 326726 | 46 |
| 15 | 629989 | 446 | 956387 | 99 | 673602 | 546 | 326398 | 45 |
| 16 | 630257 | 446 | 956327 | 99 | 673929 | 545 | 326071 | 44 |
| 17 | 630524 | 446 | 956268 | 99 | 674257 | 545 | 325743 | 43 |
| 18 | 630792 | 445 | 956208 | 100 | 674584 | 545 | 325416 | 42 |
| 19 | 631059 | 445 | 956148 | 100 | 674911 | 544 | 325089 | 41 |
| 20 | 631326 | 445 | 956089 | 100 | 675237 | 544 | 324763 | 40 |
| 21 | ¢.631593 | 444 | 9.956029 | 100 | 9.675564 | 544 | 10.324436 | 39 |
| 22 | 631859 | 444 | 955969 | 100 | 675890 | 544 | 324110 | 38 |
| 23 | 632125 | 444 | 955009 | 100 | 676217 | 543 | 323783 | 37 |
| 24 | 632392 | 443 | 955849 | 100 | 676543 | 543 | 323457 | 36 |
| 25 | 632658 | 443 | 955789 | 100 | 676869 | 543 | 323131 | 35 |
| 26 | 632923 | 443 | 955729 | 100 | 677194 | 543 | 322806 | 34 |
| 27 | 633189 | 442 | 955669 | 100 | 677520 | 542 | 322480 | 33 |
| 28 | 633454 | 442 | 955609 | 100 | 677846 | 542 | 322154 | 32 |
| ${ }_{2} 29$ | 633719 | 442 | 955548 | 100 | 678171 | 542 | 321829 | 31 |
| 30 | 633984 | 441 | 955488 | 100 | 678496 | 542 | 321504 | 30 |
| 3 I | 9.634249 | 441 | 9.955428 | 101 | 9.678821 | 54 I | 10.321179 | 29 |
| 32 | 634514 | 440 | 955368 | 101 | 679146 | 541 | 320854 | 28 |
| 33 | 634778 | 440 | 955307 | 101 | 679471 | 541 | 320529 | 27 |
| 34 | 635042 | 440 | 955247 | IOI | 679795 | 54 I | 320205 | 26 |
| 35 | 635306 | 439 | 955186 | 101 | 680120 | 540 | 319880 | 25 |
| 36 | 635570 | 439 | 955126 | 101 | 680444 | 540. | 319556 | 24 |
| 37 | 635834 | 439 | 955065 | 10 | 680768 | 540 | 319232 | 23 |
| 38 | 636097 | 438 | 955005 | 10 | 681092 | 540 | 318908 | 22 |
| 39 | 636360 | 438 | 954944 | 101 | 681416 | 539 | 318584 | 21 |
| 40 | 636623 | 438 | 954883 | 101 | 681740 | 539 | 318260 | 20 |
| 41 | 9.636886 | 437 | 9.954823 | 101 | 9.682063 | 539 |  | 19 |
| 42 | 637148 | 437 | 954762 | 101 | 682387 | 539 | 317613 | 18 |
| 43 | 637411 | 437 | 954701 | 101 | 682710 | 538 | 317290 | 17 |
| 44 | 637673 | 437 | 954640 | 101 | 683033 | 538 | 316967 | 16 |
| 45 | 637935 | 436 | 954578 | 101 | 683356 | 538 | 316644 | 15 |
| 46 | 638197 | 436 | 954518 | 102 | 683679 | 538 | 316321 | 14 |
| 47 | 638458 | 436 | 954457 | 102 | 684001 | 537 | 315999 | 13 |
| 48 | 638720 | 435 | 954366 | 102 | 684324 | 537 | 315676 | 12 |
| 49 | 6.38981 | 435 | 954335 | 102 | 684646 | 537 | 315354 | 11 |
| 50 | 639242 | 435 | 954274 | 102 | 684968 | 537 | 315032 | 10 |
| 51 | 9.63,503 | 434 | 9.954213 | 102 | 9.685290 | 536 | 10.314710 |  |
| 52 | 639764 | 434 | 954152 | 102 | 685612 | 536 | 314388 | 8 |
| 53 | 640024 | 434 | 954090 | 102 | 685934 | 536 | 314066 | 7 |
| 54 | 640284 | 433 | 954029 | 102 | 686255 | 536 | 3.13745 | 6 |
| 55 | 640544 | 433 | 953968 | 102 | 686577 | 535 | 313423 | 5 |
| 56 | 640804 | 433 | 953006 | 102 | 686898 | 535 | 3 I 3102 | 4 |
| 57 | 641064 | 432 | 953845 | 102 | 687219 | 535 | 312781 | 3 |
| 58 | 641324 | 432 | 953783 | 102 | 687540 | 535 | 312460 | 2 |
| 5 s | 641583 | 432 | 953722 | 103 | 687861 | 534 | 312139 | 1 |
| 60 | 641842 | 431 | 953660 | 103 | 688183 | 534 | 311818 | 0 |
| 1 | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | 1 |
| $115^{\circ}$ 640 |  |  |  |  |  |  |  |  |


| 44 | LOGARITHMIC SINES, TANGENTS, E'T'C. |  |  |  |  |  | . Table |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $25^{\circ}$ |  | D. | Cosine. | D. | Tang. | D. | $153^{\circ}$ |  |
| , | Sine. |  |  |  |  |  | Cotang. | , |
| 0 | 9.641842 | 431 | 9.953660 | 103 | 9.688182 | 534 | 10.311818 | 60 |
| 1 | 642101 | 431 | ¢53599 | 103 | 688502 | 534 | 311498 | 59 |
| 2 | 642360 | 431 | 953537 | 103 | 688823 | 534 | 311177 | 58 |
| 3 | 642618 | 430 | 953475 | 103 | 689143 | 533 | 310857 | 57 |
| 4 | 642877 | 430 | 953413 | 103 | 689463 | 533 | 310537 | 56 |
| 5 | 643135 | 430 | 953352 | 103 | 689783 | 533 | 310217 | 55 |
| 6 | 643393 | 430 | 953290 | 103 | 690103 | 533 | 309897 | 54 |
| 7 | 643650 | 429 | 953228 | 103 | 690423 | 533 | 309577 | 53 |
| 8 | 643908 | 429 | $9{ }^{53166}$ | - 03 | 690742 | 532 | 309258 | 52 |
| 9 | 644105 | 429 | 953104 | 103 | 691062 | 532 | 308938 | 51 |
| 10 | 644423 | 428 | 953042 | 103 | 691381 | 532 | 308619 | 50 |
| 11 | 9.644680 | 428 | 9.952980 | 104 | 9.691700 | 531 | 10.308300 | 49 |
| 12 | 644936 | 428 | 952918 | 104 | 692019 | 531 | 307981 | 48 |
| 13 | $64512^{3}$ | 427 | 952855 | 104 | 692338 | 531 | 307662 | 47 |
| 14 | 645450 | 427 | 952793 | 104 | 692656 | 531 | 307344 | 46 |
| 15 | 645706 | 427 | 952731 | 104 | 692975 | 531 | 307025 | 45 |
| 16 | 645962 | 426 | 952669 | 104 | 693293 | 530 | 306707 | 44 |
| 17 | 646218 | 426 | 952606 | 104 | 693612 | 530 | 306388 | 43 |
| 18 | 646474 | 426 | 952544 | 104 | 693930 | 530 | 306070 | 42 |
| 19 | 646729 | 425 | 952481 | 104 | 694248 | 5.30 | 305752 | 41 |
| 20 | 646984 | 425 | 952419 | 104 | 694566 | 529 | 305434 | 40 |
| 21 | 9.647240 | 425 | 9.952356 | 104 | 9.694883 | 529 | $10 \cdot 305117$ | 39 |
| 22 | 647494 | 424 | 952294 | 104 | 695201 | 529 | 304799 | 38 |
| 23 | 647749 | 424 | 952231 | 104 | 695518 | 529 | 304482 | 37 |
| 24 | 648004 | 424 | 952168 | 105 | 695836 | 529 | 304164 | 36 |
| 25 | 648258 | 424 | 952106 | 105 | 696153 | 528 | 303847 | 35 |
| 26 | 648512 | 423 | 952043 | 105 | 696470 | 528 | 303530 | 34 |
| 27 | 648766 | 423 | 951980 | 105 | 696787 | 528 | 303213 | 33 |
| 28 | 649020 | 423 | 951917 | 105 | 697103 | 528 | 302897 | 32 |
| 29 | 649274 | 422 | 951854 | j05 | 697420 | 527 | 302580 | 3 I |
| 30 | 649527 | 422 | 951791 | 105 | 697736 | 527 | 302264 | 30 |
| 31 | 9.649781 | 422 | 9.951728 | 105 | ¢. 698053 | 527 | 10.301947 | 29 |
| 32 | 650034 | 422 | 951665 | 105 | ¢ 69836ı | 527 | 301631 | 28 |
| 33 | 650287 | 421 | 951602 | 105 | 698685 | 526 | 301315 | 27 |
| 34 | 650539 | 421 | 951539 | 105 | 699001 | 526 | 300999 | 26 |
| 35 | 650792 | 42 I | 951476 | 105 | 699316 | 526 | 300684 | 25 |
| $36^{\circ}$ | 651044 | 420 | 951412 | :05 | 699632 | 526 | 300368 | 24 |
| 37 | 651297 | 420 | 951349 | 106 | 699947 | 526 | 300053 | 23 |
| 38 | 651549 | 420 | 951286 | 106 | 700263 | 525 | 299737 | 22 |
| 39 | 651800 | 419 | 951222 | 106 | 700578 | 525 | 299422 | 21 |
| 40 | 652052 | 419 | 951159 | 106 | 700893 | 525 | 299107 | 20 |
| 41 | 9.652304 | 419 | $9 \cdot 951096$ | 106 | 9.701208 | 524 | 10.298792. | 19 |
| 42 | 652555 | 418 | 951032 | 106 | 701523 | 524 | $298477^{\circ}$ | 18 |
| 43 | 652806 | 418 | 950968 | 106 | 701837 | 524 | 298163 | 17 |
| 44 | 653057 | 418 | 950905 | 106 | 702152 | 524 | 297848 | 16 |
| 45 | 653308 | 418 | 950841 | 106 | 702466 | 524 | 297534 | 15 |
| 46 | 653558 | 417 | 950778 | 106 | 702781 | 523 | 297219 | 14 |
| 47 | 653808 | 417 | 950714 | 106 | 703095 | 523 | 296205 | 13 |
| 48 | 654059 | 417 | 950650 | 106 | 703409 | 523 | 296591 | 12 |
| 49 | 654309 | 416 | 950586 | 106 | 703722 | 523 | 296278 | 11 |
| 50 | 654558 | 416 | 950522 | 107 | 704036 | 522 | 295964 | 10 |
| 51 | 9.654808 | 416 | 9.950458 | 107 | 9.704350 | 522 | 10.295650 |  |
| 52 | 655058 | 416 | 950394 | 107 | 704663 | 522 | 295337 | 8 |
| 53 | 655307 | 415 | 950330 | 107 | 704976 | 522 | 295024 | 7 |
| 54 54 | 655556 | 415 | 950266 | 107 | 705290 | 522 | 294710 | 6 |
| 55 56 | 655805 | 415 | 9 0202 | 107 | 705603 | 521 | 294397 | 5 |
| 56 | 656054 | 414 | 950138 | 107 | 705916 | 521 | 294084 | 4 |
| 57 58 | 656302 | 414 | 950074 | 107 | 706228 | 521 | $29377^{2}$ | 3 |
| 58 | 656551 | 414 | 950010 | 107 | 706541 | 521 | 293459 | 2 |
| 59 | 656799 | 413 | 949945 | 107 | 706854 | 521 | 293146 | 1 |
| 60 | 657047 | 413 | 949881 | 107 | 707166 | 520 | 292834 | 0 |
| , | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | 1 |
| $116^{\circ}$ |  |  |  |  |  |  |  | $63^{\circ}$ |


| Table II. |  | LOGARITHMIC |  | SINES, | TANGENTS, ETC |  |  | 45 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $27^{\circ}$ |  |  |  |  |  |  |  | $52^{\circ}$ |
| , | Sine. | D. | Cosine. | d. | Tang. | D. | Cotang. | , |
| 0 | 9.657047 | 413 | 9.949881 | 107 | $0 \cdot 707166$ | 520 | 10.292834 | 60 |
| 1 | $6572{ }^{5}$ | 413 | 949816 | 107 | 707478 | 520 | 292522 | 59 |
| 2 | 657542 | 412 | 949752 | 107 | 707790 | 520 | 292210 | 58 |
| 3 | 657790 | 412 | 949688 | 108 | 708102 | 520 | 291808 | 57 |
| 4 | 658037 | 412 | 949623 | 108 | 708414 | 519 | 291586 | 56 |
| 5 | 658284 | 412 | 949558 | 108 | 708726 | 519 | 291274 | 55 |
| 6 | 65853 I | 411 | 949494 | 108 | 709037 | 519 | 290963 | 54 |
| 7 | 658778 | 411 | 949429 | 108 | 709349 | 519 | 290651 | 53 |
| 8 | 659025 | 411 | 949364 | 108 | 709660 | 519 | 290340 | 52 |
| 9 | 659271 | 410 | 949300 | 108 | 709971 | 518 | 290029 | 51 |
| 10 | 659517 | 410 | 949235 | 108 | 710282 | 518 | 289718 | 50 |
| 11 | 9.659763 | 410 | 9.949170 | 108 | 9.710593 | 518 | $10 \cdot 289407$ | 49 |
| 12 | 660009 | 409 | 949105 | 108 | 710904 | 518 | 288096 | 48 |
| 13 | 660255 | 409 | 949040 | 108 | 711215 | 518 | 288785 | 47 |
| 14 | 660501 | 409 | 948975 | 108 | 711525 | 517 | 288475 | 46 |
| 15 | 660746 | 409 | 948910 | 108 | 711836 | 517 | 288164 | 45 |
| 16 | 660991 | 408 | 948845 | 108 | 712146 | 517 | 287854 | 44 |
| 17 | 661236 | 408 | 948780 | 109 | 712456 | 517 | 287544 | 43 |
| 18 | 661481 | 408 | 948715 | 109 | 712766 | 516 | 287234 | 42 |
| 19 | 661726 | 407 | 948650 | 109 | 713076 | 516 | 286924 | 41 |
| 20 | 661970 | 407 | 948584 | 109 | 713386 | 516 | 286614 | 40 |
| 21 | 9.662214 | 407 | 9.948519 | 109 | 9.713696 | 516 | 10. 286304 | 39 |
| 22 | 662459 | 407 | 948454 | 109 | 714005 | 516 | 285995 | 38 |
| 23 | 662703 | 406 | 948388 | 109 | 714314 | 515 | 285686 | 37 |
| 24 | 662946 | 406 | 948323 | 109 | 714624 | 515 | 285376 | 36 |
| 25 | $66310^{\circ}$ | 406 | 948257 | 109 | 714933 | 515 | 285067 | 35 |
| 26 | 663433 | 405 | 948192 | 109 | 715242 | 515 | 284758 | 34 |
| 27 | 663677 | 405 | 948126 | 109 | 715551 | 514 | 284449 | 33 |
| 28 | 663920 | 405 | 948060 | 109 | 715860 | 514 | 284140 | 32 |
| 29 | 664163 | 405 | 947995 | 110 | 716168 | 514 | 283832 | 31 |
| 30 | 664406 | 404 | 947929 | 110 | 716477 | 514 | 283523 | 30 |
| 31 | 9.664648 | 404 | 9.947863 | 110 | 9716785 | 514 | 10. 283215 | 29 |
| 32 | 664891 | 404 | 947797 | 110 | $7170{ }^{3}$ | 513 | 282907 | 28 |
| 33 | 665133 | 403 | 947731 | 110 | 717401 | 513 | 282599 | 27 |
| 34 | 665375 | 403 | 947665 | 110 | 717709 | 513 | 282291 | 26 |
| 35 | 665617 | 403 | 947600 | 110 | 718017 | 513 | 281983 | 25 |
| 36 | 665859 | 402 | 947533 | 110 | 718325 | 513 | 281675 | 24 |
| 37 | 666100 | 402 | 947467 | 110 | 718633 | 512 | 281367 | 23 |
| 38 | 666342 | 402 | 947401 | 110 | 718940 | 512 | 281060 | 22 |
| 39 | 666583 | 402 | 947335 | 110 | 719248 | 512 | 280752 | 21 |
| 40 | 666824 | 401 | 947269 | 110 | 719555 | 512 | 280445 | 20 |
| 41 | 9.667065 | 401 | 9.947203 | 110 | $9 \cdot 719862$ | 512 | 10.280138 | 19 |
| 42 | 667305 | 401 | 947136 | 111 | 720169 | 511 | 27983 I | 18 |
| 43 | 667546 | 401 | 947070 | III | 720476 | 511 | 279524 | 17 |
| 44 | 667786 | 400 | 947004 | 111 | 720783 | 511 | 279217 | 16 |
| 45 | 668027 | 400 | 946937 | 111 | 721089 | 511 | 278911 | 15 |
| 46 | 668267 | 400 | 946871 | 111 | 721396 | 511 | 278604 | 14 |
| 47 | 668506 | 399 | 946804 | 111 | 721702 | 510 | 278298 | 13 |
| 48 | 668746 | 399 | 946738 | 11 | 722009 | 510 | 277991 | 12 |
| 49 | 668986 | 399 | 946671 | III | 722315 | 510 | 277685 | 11 |
| 50 | 669225 | 399 | 946604 | 111 | 722621 | 510 | 277379 | 10 |
| 51 |  |  | 9.946538 | 111 | $5 \cdot 722927$ |  |  |  |
| 52 | 669703 | 308 | 946471 | III | 723232 | 509 | 276768 | 8 |
| 53 | 669942 | 398 | 940404 | 111 | 723538 | 509 | 276462 | 7 |
| 54 | 670181 | 397 | 946337 | 111 | 723844 | 509 | 276156 | 6 |
| 55 | 670419 | 397 | 946270 | 112 | 724149 | 509 | 27585 I | 5 |
| 56 | 670658 | 397 | 946203 | 112 | 724454 | 509 | 275546 | 4 |
| 57 58 | 670896 | 397 | 946136 | 112 | 724760 | 508 | 275240 | 3 |
| 58 | 671134 | 396 | 946069 | 112 | 725065 | 508 | 274935 |  |
| 59 60 | 671372 671609 | 396 396 | 946002 945935 | 112 | 725370 725674 | 508 508 | 274630 274326 | 0 |
| 1 | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | 1 |
| $117^{\circ}$ |  |  |  |  |  |  |  | $62^{\circ}$ |


| 46 | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  |  | . Table | II. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $28^{\circ}$ |  |  |  |  |  |  | $151^{\circ}$ |  |
| , | Sine. | D. | Cosine. | D. | Tang. | D. | Cotang. | , |
| 0 | 9.671609 | 396 | 9.945935 | 112 | 9.725674 | 508 | 10.274326 | 60 |
| 1 | 671847 | 395 | 945868 | 112 | 725979 | 508 | 274021. | 59 |
| 2 | 672084 | 395 | 945800 | 12 | 726284 | 507 | 273716 | 58 |
| 3 | 672321 | 395 | 945733 | 11 | 726588 | 507 | 273412 | 57 |
| 4 | 672558 | 395 | 945666 | 112 | 726892 | 507 | 273108 | 56 |
| 5 | 672795 | 394 | 945598 | 112 | 727197 | 507 | 272803 | 55 |
| 6 | 673032 | 394 | 945531 | 112 | 727501 | 507 | 272499 | 54 |
| 7 | 673268 | 394 | 945464 | II3 | 727805 | 506 | $27219^{5}$ | 53 |
| 8 | 673505 | 394 | 945396 | 113 | 728109 | 506 | 271891 | 52 |
| 9 | 673741 | 393 | 945328 | 113 | 728412 | 506 | 271588 | 51 |
| 10 | 673977 | 39.3 | 945261 | 113 | 728716 | 506 | 271284 | 50 |
| 12 | 9.6742 I 3 | 393 | 9.945193 | 113 | 9.729020 | 506 | 10.270980 | 40 |
|  | 674448 | 392 | 945125 | 113 | 729323 | 505 | 270677 | 48 |
|  | 674684 | 392 | 945058 | 113 | 729626 | 505 | 270374 | 47 |
| 13 | 674919 | 392 | 944990 | 113 | 729929 | 505 | 270071 | 46 |
| 1516 | 675155 | 392 | 944922 | 113 | 730233 | 505 | 269767 | 45 |
|  | 675390 | 391 | 944854 | 113 | 730535 | 505 | 269465 | 44 |
| $\begin{aligned} & 17 \\ & 18 \end{aligned}$ | 675624 | 391 | 944786 | 113 | 730838 | 504 | 269162 | 43 |
| $18$ | 675859 | 391 | 944718 | 113 | 731141 | 504 | 268850 | 42 |
| $\begin{aligned} & 19 \\ & 20 \end{aligned}$ | 676094 676328 | 391 390 | 944650 944582 | 113 | 731444 731746 | 504 504 | 268556 268254 | 4 l |
|  |  | 390 |  | 114 | $731746$ |  |  | 40 |
| 21 | $9 \cdot 676562$ | 390 | 9.944514 | 114 | 9.732048 | 504 | $10 \cdot 267952$ | $\begin{aligned} & 39 \\ & 38 \end{aligned}$ |
| 22 | 676796 677030 | 390 390 | 944446 944377 | 114 | 732351 732653 7325 | 503 503 | $\begin{aligned} & 267649 \\ & 267347 \end{aligned}$ | $38$ |
| 23 | 677030 677264 | 390 380 | 944377 944309 | 114 | 732653 | 503 503 | 267347 267045 | 37 36 |
| 24 | 677498 | 389 | 944309 | 114 | 733257 | 503 | 266743 | 35 |
| 25 | 677731 | 389 | 944172 | 114 | 733558 | 503 | 266442 | 34 |
| $27$ | 677964 | 388 | 944104 | 114 | 733860 | 502 | 266140 | 33 |
| $28$ | 678197 | 388 | 944036 | 114 | 734162 | 502 | 265838 | 32 |
| $29$ | 678430 | 388 | 943967 | 114 | 734463 | 502 | 2655337 | 3 I |
| 30 | 678663 | 388 | 943899 | 114 | 734764 | 502 | 265236 | 30 |
| 31 | 9.678895 | 387 | 9.943830 | 114 | 9735066 | 502 | 10.264934 | 29 |
| 32 | 679128 | 387 | 943761 | 114 | 735367 | 502 | 264633 | 28 |
| 33 | 679360 | 387 | 943693 | 115 | 735668 | 501 | 264332 | 27 |
| 34353 | 679592 | 387 | 943624 | 115 | 735969 | 501 | 264031 | 26 |
|  | 679824 | 386 | 943555 | 115 | 736269 | 501 | 263731 | 25 |
| 36 | 680056 | 386 | 943486 | 115 | 736570 | 501 | 263430 | 24 |
| 373838 | 680288 | 386 | 943417 | 115 | 736870 | 501 | 263130 | 23 |
|  | 680519 | 385 | 943348 | 115 | 737171 | 500 | 262829 | 22 |
| 38 39 | 680750 | 385 | 943279 | 115 | 737471 | 500 | 262529 | 21 |
| 40 | 680982 | 385 | 943210 | 115 | 737771 | 500 | 262229 | 20 |
| 41 | 9.681213 | 385 | 9.943141 | 115 | 9.738071 | 500 | 10.261929 | 19 |
| 42 | 681443 | 384 | 943072 | 115 | 738371 | 500 | 261629 | 18 |
| 4 | 681674 | 384 | 943003 | 115 | 738671 | 499 | 261329 | 17 |
| 4 | 681905 | 384 | 942934 | 115 | 738971 | 499 | 261029 | 16 |
| $\begin{aligned} & 44 \\ & 45 \end{aligned}$ | 682135 | 384 | 942864 | 115 | 739271 | 499 | 260729 | 15 |
| 46 | 682365 | 383 | 942795 | 116 | 739570 | 499 | 260430 | 14 |
| 47 | 68.2595 | 383 | 942726 | 116 | 739870 | 499 | 260130 | 13 |
| 4 | 682825 | 383 | 942656 | 116 | 740169 | 499 | 25983 I | 12 |
| 4 4 5 | 683055 | 383 | 942587 | 116 | 740468 | 498 | 250532 | 11 |
|  | 683284 | 382 | 942517 | 116 | 740767 | 498 | 259233 | 10 |
| 5152 | 9.683514 | 382 |  | 116 |  |  |  |  |
|  | 683743 | 382 | $942378$ | 116 | 741365 | 498 | 258635 | 8 |
| 53 | 683972 | 382 | 942308 | 116 | 741664 | 498 | 258336 | 7 |
| 5455 | 684201 | 381 | 942239 | 116 | 741962 | 497 | 258038 | 6 |
|  | 684430 | 381 | 942169 | 116 | 742261 | 497 | 257739 | 5 |
| 56 | 684658 | 381 | 942099 | 116 | 742559 | 497 | 257441 | 4 |
| 5758 | 684887 | 380 | 942029 | 116 | 742858 | 497 | 257142 | 3 |
|  | 685115 | 380 | 941959 | 116 | 743156 | 497 | 256844 | 2 |
| 59 | 685343 | 380 | 941889 | 117 | 743454 | 497 | 250546 | 1 |
| 60 | 685571 | 380 | 941819 | 117 | 743752 | 496 | 256248 | - |
| , | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | 1 |
| $118^{\circ}$ |  |  |  |  |  |  |  | $61^{\circ}$ |




Tabie 1I. LOGARITHMIC SINES, TANGENTS, ETC

| $81^{\circ}$ |  |  |  |  |  |  |  | $43^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | Sine. | D. | Cosine. | D. | Tang. | D. | Cotang. | , |
| 0 | 9.711839 | 350 | 9.933066 | 126 | 9.778774 | 477 | 10.221226 | So |
| 1 | 712050 | 350 | 932990 | 127 | 779060 | 477 | 220940 | 59 |
| 2 | 712260 | 350 | 93214 | 127 | 779346 | 476 | 220654 | 58 |
| 3 | 712469 | 349 | 932838 | 127 | 779632 | 476 | 220368 | 57 |
| 4 | 712679 | 349 | 932762 | 127 | 779918 | 476 | 220082 | 56 |
| 5 | 712889 | 349 | 932685 | 127 | 780203 | 476 | 219797 | 55 |
| 6 | 713098 | 349 | 932609 | 127 | 780489 | 476 | 219511 | 54 |
| 7 | 713308 | 349 | 932533 | 127 | 780775 | 476 | 219225 | 53 |
| 8 | 713517 | 348 | 932457 | 127 | 781060 | 476 | 218940 | 52 |
| 9 | 713726 | 348 | 932380 | 127 | 781346 | 475 | 218654 | 51 |
| 10 | 713935 | 348 | 932304 | 127 | 781631 | 475 | 218369 | 50 |
| 11 | 9.714144 | 348 | 9.932228 | 127 | $9 \cdot 781916$ | 475 | 10.218084 | 49 |
| 12 | 714352 | 347 | 932151 | 127 | 782201 | 475 | 217799 | 48 |
| 13 | 714561 | 347 | 932075 | 128 | 782486 | 475 | 217514 | 47 |
| 14 | 714769 | 347 | 931998 | 128 | 782771 | 475 | 217229 | 46 |
| 15 | 714978 | 347 | 931921 | 128 | 783056 | 475 | 216944 | 45 |
| 16 | 715186 | 347 | 931845 | 128 | 783341 | 475 | 216659 | 44 |
| 17 | 715394 | 346 | 931768 | 128 | 783626 | 474 | 216374 | 43 |
| 18 | 715602 | 346 | 931691 | 128 | 783910 | 474 | 216090 | 42 |
| 19 | 715809 | 346 | 931614 | 128 | 784195 | 474 | 215805 | 41 |
| 20 | 716017 | 346 | 931537 | 128 | 784479 | 474 | 215521 | 40 |
| 21 | 9.716224 | 345 | 9.931460 | 128 | 9.784764 | 474 | 10.215236 | 39 |
| 22 | 716432 | - 345 | 93ı383 | 128 | 785048 | 474 | 214952 | 38 |
| 23 | 716639 | 345 | 931306 | 128 | 785332 | 473 | 214668 | 37 |
| 24 | 716846 | 345 | 931229 | 129 | 785616 | 473 | 214384 | 36 |
| 25 | 717053 | 345 | 931152 | 129 | 785900 | 473 | 214100 | 35 |
| 26 | 717259 | 344 | 931075 | 129 | 786184 | 473 | 213816 | 34 |
| 27 | 717466 | 344 | 930998 | 129 | 786468 | 473 | - 213532 | 33 |
| 28 | 717673 | 344 | 930021 | 129 | 786752 | 473 | 213248 | 32 |
| 29 | 717879 | 344 | 930843 | 129 | 787036 | 473 | 212964 | 31 |
| 30 | 718085 | 343 | 930766 | 129 | 787319 | 472 | 212681 | 30 |
| 31 | $9 \cdot 718291$ | 343 | 9.930688 | 129 | 9.787603 | 472 | 10.212397 |  |
| 32 | 718497 | 343 | 930611 | 129 | 787886 | 472 | 212114 | 28 |
| 33 | 718703 | 343 | 930533 | 129 | 788170 | 472 | 211830 | 27 |
| 34 | 718909 | 343 | 930456 | 129 | 788453 | 472 | 211547 | 26 |
| 35 | 719114 | 342 | 930378 | 129 | 788736 | 472 | 211264 | 25 |
| 36 | 719320 | 342 | 930300 | 130 | 789019 | 472 | 210981 | 24 |
| 37 | 719525 | 342 | 930223 | 130 | 789302 | 471 | 210698 | 23 |
| 38 | 719730 | 342 | 930145 | 130 | 789585 | 471 | 210415 | 22 |
| 39 | 719935 | 341 | 930067 | 130 | 789868 | 471 | 210132 | 21 |
| 40 | 720140 | 341 | 929989 | 130 | 790151 | 471 | 209849 | 20 |
| 41 | 9.720345 | 341 | 9.929911 | 130 | 9.790434 | 471 | 10. 209566 | 19 |
| 42 | 720549 | 341 | 929833 | 130 | 790716 | 471 | 209284 | 18 |
| 43 | 720754 | 340 | 929755 | 130 | 790999 | 471 | 209001 | 17 |
| 44 | 720958 | 340 | 929677 | 130 | 791281 | 471 | 208719 | 16 |
| 45 | 721162 | 340 | 929599 | 130 | 791563 | 470 | 208437 | 15 |
| 46 | 721366 | 340 | 929521 | 130 | 791846 | 470 | 208154 | 14 |
| 47 | 721570 | 340 | 929442 | 130 | 792128 | 470 | 207872 | 13 |
| 48 | 721774 | 339 | 929364 | 131 | 792410 | 470 | 207590 | 12 |
| 49 | 721978 | 339 | 929286 | 131 | $79269^{2}$ | 470 | 207308 | 11 |
| 50 | 722181 | 339 | 929207 | 131 | 792974 | 470 | 207026 | 10 |
| 51 | 9.722385 | 339 | 0.929129 | 131 |  |  | 10.206744 |  |
| 52 | 722588 | 339 | 929050 | 131 | 793538 | 469 | 206462 | 8 |
| 53 | 722791 | 338 | 928972 | 131 | 793819 | 469 | 206181 | 7 |
| 54 | 722994 | 338 | 928893 | 131 | 794101 | 469 | 205899 | 6 |
| 55 | 723197 | 338 | 928815 | 131 | 794383 | 469 | 205617 | 5 |
| 56 | 723400 | 338 | 928736 | 131 | 794064 | 469 | 205336 | 4 |
| 57 | 723603 | 337 | 928657 | 131 | 794946 | 469 | 205054 | 3 |
| 58 | 723805 | 337 | $9285{ }^{\text {¢ }} 7$ | 131 | 795227 | 469 | 204773 | 2 |
| 59 | 724007 | 337 | 928499 | 131 | 795508 | 468 | 204492 | 1 |
| 60 | 7242 IO | 337 | 928420 | 131 | $79^{5} 789$ | 468 | 204211 | - |
| $\cdot$ | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | 1 |
| $121^{\circ}$ |  |  |  |  |  |  |  | $58^{\circ}$ |





| Table II. |  | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  |  | 53 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $85^{\circ}$ |  |  |  |  |  |  |  | $14^{\circ}$ |
| , | Sine. | 1. | Cosine. | D. | Tang. | D. | Cotang. | , |
| 0 | 9.758591 | 301 | 9.913365 | 147 | $9 \cdot 845227$ | 448 | 10.154773 | 60 |
| 3 | 758772 | 300 | 913276 | 147 | 845496 | 448 | 154504 | 50 |
| 2 | $7589{ }^{\text {j2 }}$ | 3 oc | 913187 | 148 | 845764 | 448 | 154236 | 58 |
| 3 | 759132 | 300 | 913099 | 148 | 846033 | 448 | 153967 | 57 |
| 4 | 759312 | 300 | 913010 | 148 | 846302 | 448 | 153608 | 56 |
| 5 | 759492 | 300 | 912922 | 148 | 846570 | 447 | 153430 | 55 |
| 6 | 759672 | 299 | 912833 | 148 | 846839 | 447 | 153161 | 54 |
| 7 | 759852 | 299 | 912744 | 148 | 847108 | 447 | 152892 | 53 |
| 8 | 760031 | 299 | 912655 | 148 | 847376 | 447 | 152624 | 52 |
| 9 | 760211 | 299 | 912566 | 148 | 847644 | 447 | 152356 | 51 |
| 10 | 760390 | 299 | 912477 | 148 | 847913 | 447 | 152087 | 50 |
| 11 | 9.760569 | 298 | 9.912388 | 148 | 9.848181 | 447 | 10.151819 | 49 |
| 12 | 760748 | 298 | 912299 | 149 | 848449 | 447 | 151551 | 48 |
| 13 | 760927 | 298 | 912210 | 149 | 848717 | 447 | 151283 | 47 |
| 14 | 761106 | 298 | 912121 | 149 | 848986 | 447 | 151014 | 46 |
| 15 | 761285 | 298 | 912031 | 149 | 849254 | 447 | 150746 | 45 |
| 16 | 761464 | 298 | 911942 | 149 | 849522 | 447 | 150478 | 44 |
| 17 | 761642 | 297 | 911853 | 149 | $84979{ }^{\circ}$ | 446 | 150210 | 43 |
| 18 | 761821 | 297 | 911763 | 149 | 850057 | 446 | 149943 | 43 |
| 19 | 761999 | 297 | 911674 | 149 | 850325 | 446 | 149675 | 41 |
| 20 | 762177 | 297 | 911584 | 149 | 850593 | 446 | 149407 | 40 |
| 21 | 9.762356 | 297 | 9.911495 | 149 | 9.85086 | 446 | 10.149139 | 39 |
| 22 | 762534 | 296 | 911405 | 149 | 851129 | 446 | 148871 | 38 |
| 23 | 762712 | 296 | 911315 | 150 | 851395 | 446 | 148604 | 37 |
| 24 | 762889 | 296 | 911226 | 150 | 851664 | 446 | 148336 | 36 |
| 25 | 763067 | 296 | 911136 | 150 | 851931 | 446 | 148069 | 35 |
| 26 | 763245 | 296 | 911046 | 150 | 852199 | 446 | 147801 | 34 |
| 27 | 763422 | 296 | 910956 | 150 | 852466 | 446 | 147534 | 33 |
| 28 | 763600 | 295 | 910866 | 150 | 852733 | 445 | 147267 | 32 |
| 29 | 763777 | 295 | 910776 | 150 | 853001 | 445 | 146999 | 31 |
| 30 | 763954 | 295 | 910686 | 150 | 853268 | 445 | 146732 | 30 |
| 31 | $9 \cdot 764131$ | 295 | 9.910596 | 150 | 9.853535 | 445 | 10.146465 | 29 |
| 32 | 764308 | 295 | 910506 | 150 | 853802 | 445 | 146198 | 28 |
| 33 | 764485 | 294 | 910415 | 150 | 854069 | 445 | 145031 | 27 |
| 34 | 764662 | 294 | 910325 | 151 | 854336 | 445 | 145664 | 26 |
| - 35 | 764838 | 294 | 910235 | 151 | 854603 | 445 | 145397 | 25 |
| 36 | 765015 | 294 | 910144 | 151 | 854870 | 445 | 145130 | 24 |
| 37 | 765191 | 294 | 910054 | 151 | 855137 | 445 | 144863 | 23 |
| 38 | 765367 | 294 | 909963 | 151 | 855404 | 440 | 144596 | 22 |
| 39 | 765544 | 293 | 909873 | 151 | 855671 | 444 | 144329 | 21 |
| 40 | 765720 | 293 | 909782 | 151 | 855938 | 444 | 144062 | 20 |
| 41 | $9 \cdot 765896$ | 293 | 9.909691 | 151 | 9.856204 | 444 | 10.143796 | 19 |
| 42 | 766072 | 293 | 909601 | 151 | 856471 | 444 | 143529 | 18 |
| 43 | 766247 | 293 | 909510 | 151 | 856737 | 444 | 143263 | 17 |
| 44 | 766423 | 293 | 909419 | 151 | 857004 | 444 | 142996 | 16 |
| 45 | 766598 | 292 | 909328 | 152 | 857270 | 444 | 142730 | 15 |
| 46 | 766774 | 292 | 909237 | 152 | 8575 ? 7 | 444 | 142463 | 14 |
| 47 | 766949 | 292 | 909146 | 152 | 857803 | 444 | 142197 | 13 |
| 48 | 767124 | 292 | 909055 | 152 | 858069 | 444 | 141931 | 12 |
| 49 | 767300 | 292 | 908964 | 152 | 858336 | 444 | 141664 | 11 |
| 50 | 767475 | 291 | 908873 | 152 | 858602 | 443 | 141398 | 10 |
| 51 |  | 291 |  | 152 | y. 858868 | 443 |  |  |
| 52 | 767824 | 291 | 908690 | 152 | 859134 | 443 | 140866 | 8 |
| 53 | 767999 | 291 | 908599 | 152 | 859400 | 443 | 140600 | 7 |
| 54 55 | 768173 | 291 | 908507 | 152 | 859666 | 443 | 140334 | 6 |
| 55 56 | 768348 | 290 | 908416 | 153 | 859932 | 443 | 140068 | 5 |
| 56 | 768522 | 290 | 908324 | 153 | 860198 | 443 | 139802 | 4 |
| 57 58 58 | 768697 | 290 | 908233 | 153 | 860464 | 443 | 139536 | 3 |
| 58 59 | 768871 | 290 | 908141 | 153 | 860730 | 443 | 139270 | 2 |
| 59 60 | 769045 | 290 290 | 9080 尔 90798 | 153 153 | 860995 861261 | 443 | 139005 138739 | 1 |
| , | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | ' |
| $125^{\circ}$ |  |  |  |  |  |  |  | $54^{\circ}$ |



| Table II. |  | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  |  | 55 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $37^{\circ}$ |  |  |  |  |  |  |  | $142^{\circ}$ |
| , | Sine. | D. | Cosine. | D. | Tang. | D. | Cotang. | , |
| 0 | 9.779463 | 279 | 9.902349 | 159 | 9.877114 | 438 | 10.122886 | 60 |
| 1 | 779631 | 279 | 902253 | 159 | 877377 | 438 | 12262.3 | 59 |
| 2 | 779798 | 279 | 902158 | 159 | 877640 | 438 | 122360 | 58 |
| 3 | 779966 | 279 | 902063 | 159 | 877903 | 438 | 122097 | 57 |
| 4 | 780133 | 279 | 901967 | 159 | 878165 | 438 | 121835 | 56 |
| 5 | 780300 | 278 | 701872 | 159 | 878428 | 438 | 121572 | 55 |
| 6 | 780467 | 278 | 901776 | 159 | 878691 | 438 | 121309 | 54 |
| 7 | 780634 | 278 | 901681 | 159 | 878953 | 437 | 121047 | 53 |
| 8 | 780801 | 278 | 901585 | 159 | 879216 | 437 | 120784 | 52 |
| $\bigcirc$ | 780968 | 278 | 901490 | 159 | 879478 | 437 | 120522 | 51 |
| 10 | 781134 | 278 | 901394 | 160 | 879741 | 437 | 120259 | 50 |
| 11 | $9 \cdot 781301$ | 277 | 9.901298 | 160 | 9.880003 | 437 | 10.119997 | 49 |
| 12 | 781468 | 277 | 901202 | 160 | 880265 | 437 | 119735 | 48 |
| 13 | 781634 | 277 | 901106 | 160 | 880528 | 437 | 119472 | 47 |
| 14 | 781800 | 277 | 901010 | 160 | 880790 | 437 | 119210 | 46 |
| 15 | 781966 | 277 | 900914 | 160 | 881052 | 4.37 | 118948 | 45 |
| 16 | 782132 | 277 | 900818 | 160 | 881314 | 437 | 118686 | 44 |
| 17 | 782298 | 276 | 900722 | 160 | 881577 | 437 | 118423 | 43 |
| 18 | 782464 | 276 | 900626 | 160 | 881839 | 437 | 118161 | 42 |
| 19 | 782630 | 276 | 900529 | 160 | 882101 | 437 | 117899 | 41 |
| 20 | 782796 | 276 | 900433 | 161 | 882363 | 436 | 117637 | 40 |
| 21 | $9 \cdot 782961$ | 276 | 9.900337 | 161 | 9.882625 | 436 | 10.117375 | 39 |
| 22 | 783127 | 276 | 900240 | 161 | 882887 | 436 | 117113 | 38 |
| 23 | 783292 | 275 | 900144 | 161 | 883148 | 436 | 116852 | 37 |
| 24 | 783458 | 275 | 900047 | 161 | 883410 | 436 | 116590 | 36 |
| 25 | 783623 | 275 | 8999 ¹ | 161 | 883672 | 436 | 116328 | 35 |
| 26 | 783788 | 275 | 899854 | 161 | 883934 | 436 | 116066 | 34 |
| 27 | 783953 | 275 | 899757 | 161 | 884196 | 436 | 115804 | 33 |
| 28 | 784118 | 275 | 899660 | 161 | 884457 | 436 | 115543 | 32 |
| 29 | 784282 | 274 | 899564 | 161 | 884719 | 436 | 115281 | 31 |
| 30 | 784447 | 274 | 899467 | 162 | 884980 | 436 | 115020 | 30 |
| 31 | 9.784612 | 274 | 9.899370 | 162 | 9.885242 | 436 | 10.114758 | 29 |
| 32 | 784776 | 274 | 899273 | 162 | 885504 | 436 | 114496 | 28 |
| 33 | 784941 | 274 | 899176 | 162 | 885765 | 436 | 114235 | 27 |
| 34 | 785105 | 274 | 899078 | 162 | 886026 | 436 | 113974 | 26 |
| 35 | 785269 | 273 | 898981 | 162 | 886288 | 436 | 113712 | 25 |
| 36 | 785433 | 273 | 898884 | 162 | 886549 | 435 | 113451 | 24 |
| 37 | 785597 | 273 | 898787 | 162 | 886811 | 435 | 113189 | 23 |
| 38 | 785761 | 273 | 898689 | 162 | 887072 | 435 | 112928 | 22 |
| 39 | 785925 | 273 | 898592 | 162 | 887333 | 435 | 112667 | 21 |
| 40 | 786089 | 273 | 898494 | 163 | 887594 | 435 | 112406 | 20 |
| 41 | 9.786252 | 272 | $9 \cdot 898397$ | 163 | 9.887855 | 435 | 10.112145 |  |
| 42 | 786416 | 272 | 898299 | 163 | 888116 | 435 | 111884 | 18 |
| 43 | 786579 | 272 | 898202 | 163 | 888378 | 435 | 111622 | 17 |
| 44 | 786742 | 272 | 898104 | 163 | 888639 | 435 | 111361 | 16 |
| 45 | 786906 | 272 | 898006 | 163 | 888900 | 435 | 111100 | 15 |
| 46 | 787069 | 272 | 897908 | 163 | 889151 | 435 | 110839 | 14 |
| 47 | 787232 | 271 | 897810 | 163 | 889421 | 435 | 110579 | 13 |
| 48 | 787395 | 271 | 897712 | 163 | 889682 | 435 | 110318 | 12 |
| 49 | 787557 | 271 | 897614 | 163 | 889943 | 435 | 110057 | 11 |
| 50 | 787720 | 271 | 897516 | 163 | 890204 | 434 | 109796 | 10 |
| 51 | 9.787883 | 271 | 9.897418 | 164 | 9.890465 | 434 | 10.109535 |  |
| 52 | 788045 | 271 | 897320 | 164 | 890725 | 434 | 109275 | 8 |
| 53 | 788208 | 271 | 897222 | 164 | 890986 | 434 | 109014 | 7 |
| 54 | 788370 | 270 | 897123 | 164 | 891247 | 434 | 108753 | 6 |
| 55 | 788532 | 270 | 897025 | 164 | 891507 | 434 | 108493 | 5 |
| 56 | 788694 | 270 | 896926 | 164 | 891768 | 434 | 108232 | 4 |
| 57 58 | 788856 | 270 | 896828 | 164 | 892028 | 434 | 107972 | 3 |
| 58 | 789018 | 270 | 896729 | 164 | 892289 | 434 | 107711 | 2 |
| 59 | 789180 | 270 | 896631 | 164 | 892549 | 434 | 107451 | 1 |
| 60 | 789342 | 269 | 896532 | 164 | 892810 | 434 | 107190 | 0 |
| 1 | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | ' |
| $127^{\circ}$ |  |  |  |  |  |  |  | $52^{\circ}$ |


| 56 | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  |  | . Tabl | II. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $38^{\circ}$ |  |  |  |  |  |  |  |  |
| , | Sze. | D. | Cosine. | D. | Tang. | D. | Cotang. | , |
| 012345678910 | $\begin{array}{r} 9 \cdot 789342 \\ 789504 \\ 789665 \\ 789827 \\ 789988 \\ 790149 \\ 790310 \\ 790471 \\ 790632 \\ 790793 \\ 790954 \end{array}$ | 269 | $9 \cdot 895532$ | 164 | 9.892810 | 434 | 10.107190 | 50 |
|  |  | 269 | 896433 | 165 | 893070 | 43.4 | 106,30 | 59 |
|  |  | 269 | 896335 | 165 | 893331 | 434 | 106669 | 58 |
|  |  | 269 | 896236 | 165 | 893591 | 434 | 106409 | 57 |
|  |  | 269 | 806137 | 165 | 893351 | 434 | 106149 | 56 |
|  |  | 269 | 896033 | 165 | 89, 111 | 434 | 105889 | 55 |
|  |  | 268 | 895939 | 165 | 894372 | 434 | 105629 | 54 |
|  |  | 268 | 895840 | 165 | 894632 | 433 | 105368 | 53 |
|  |  | 268 | 895741 | 165 | 894892 | 433 | 105108 | 52 |
|  |  | 268 | 895641 | 165 | 895152 | 433 | 1048.8 | 51 |
|  |  | 268 | 895542 | 165 | 895412 | 433 | 104588 | 50 |
| 11 | 9.791115 | 268 | 9.895443 | 166 | 9.895672 | 433 | 10.104328 | 49 |
| 12 | 791275 | 257 | 895343 | 166 | 895932 | 433 | 104068 | 48 |
| 13 | 791436 | 267 | 895244 | 166 | $8961{ }^{2} 2$ | 433 | 103508 | 47 |
| 14 | 791506 | 267 | 895145 | 166 | 896452 | 433 | 103548 | 46 |
| 15 | 791757 | 267 | 895045 | 165 | 896712 | 433 | 103288 | 45 |
| 16 | 791917 | 267 | 89.485 | 166 | 896971 | 433 | 103029 | 44 |
|  | 792077 | 267 | 89.4846 | 166 | 89723i | 433 | 102-69 | 43 |
| $\begin{aligned} & 17 \\ & 18 \end{aligned}$ | 792237 | 266 | 89.476 | 166 | $89-491$ | 433 | 102509 | 42 |
| $\begin{aligned} & 10 \\ & 19 \end{aligned}$ | 792307 | 266 | 89.4646 | 166 | 89775 | 433 | . 02249 | 41 |
| $\begin{aligned} & 19 \\ & 20 \end{aligned}$ | 792557 | 266 | 894546 | 166 | 89Soro | 433 | 101990 | 40 |
| 21 | 9.792716 | 266 | 9.894446 | 167 | 9.898270 | 433 | 10.101730 | 39 |
| 22 | $7928-6$ | 266 | 89.4346 | 167 | 898530 | 433 | 101470 | 38 |
| 23 | 793035 | 256 | 89.42 .46 | 167 | 898789 | 433 | 101211 | 37 |
| 2 | 793195 | 265 | 89.1146 | 167 | 899049 | 432 | 100951 | 36 |
| 25 | 793354 | 265 | 894046 | 167 | 899303 | 432 | 100692 | 35 |
| 26 | 793514 | 265 | 893946 | 167 | 899503 | 432 | 100432 | 34 |
| 27 | 793673 | 265 | 893846 | 167 | 899927 | 432 | 100173 | 33 |
|  | 793832 | 265 | 893745 | 167 | $9000 \mathrm{~S}_{7}$ | 432 | 099913 | 32 |
| 28 | 793991 | 265 | 893645 | 167 | 900346 | 432 | 099654 | 31 |
| $\begin{aligned} & 29 \\ & 30 \end{aligned}$ | 791150 | 264 | 893544 | 167 | 900605 | 432 | 099395 | 30 |
| 31 | 9.79.4308 | 26.4 | 9.893414 | 168 | 9.900S64 | 432 | 10.099136 | 20 |
| 333 | 79.1467 | 264 | 8933.43 | 169 | 901124 | 432 | 098876 | 28 |
|  | 79.4626 | 264 | 893243 | 168 | 901383 | 432 | 095617 | 27 |
|  | 704784 | 264 | 893142 | 168 | 901642 | 432 | 098358 | 26 |
| $\begin{aligned} & 34 \\ & 35 \end{aligned}$ | $70 ¢ 942$ | 264 | $8930 \leq 1$ | 168 | 901901 | 432 | 095099 | 25 |
| 36 | 795101 | 26.4 | 892910 | 168 | 902160 | 432 | 097540 | 24 |
| 36 <br> 38 | 795259 | 263 | 892839 | 168 | 902420 | 432 | 09-580 | 23 |
|  | 795417 | 263 | 892739 | 168 | 902679 | 432 | 097321 | 22 |
| 3 3 4 | 795575 | 263 | 892638 | 169 | 902938 | 432 | 09-062 | 21 |
| 4 | 795733 | 263 | 892536 | 168 | 903197 | 431 | 096803 | 20 |
| 41 | 9.795891 | 263 | 9.892435 | 169 | 9.903456 | 431 | 10.096544 | 19 |
|  | 7960.49 | 263 | 892334 | 169 | 903714 | 43 I | 096256 | 18 |
| 4 | 796206 | 263 | Sg2233 | 169 | $9039-3$ | 43 I | 096027 | 17 |
|  | 796364 | 262 | $\varepsilon_{92132}$ | 169 | 904232 | 431 | $095-68$ | 16 |
| 4 | 796521 | 262 | 8g2030 | 169 | 904for | 43 I | 095509 | 15 |
|  | 7966 -9 | 262 | 891029 | 169 | 904750 | 43 I | 095250 | 14 |
|  | 796836 | 262 | 891827 | 169 | 905008 | 431 | 094992 | 13 |
| 4 | 796993 | 262 | 891726 | 169 | 905267 | 431 | 095733 | 12 |
|  | 797150 | 261 | 891624 | 169 | 905526 | 43 I | 09454 | 11 |
| 50 | 797307 | 251 | 891523 | 170 | 90ご-S5 | 43i | 09¢215 | 10 |
| 51 | 9.797464 | 251 | 9.891421 | 170 | 9.906043 | 431 | 10.093957 |  |
| 52 | 797621 | 261 | 891319 | 170 | 906302 | 431 | 093698 | 8 |
| 53 | 797777 | 261 | 891217 | 170 | 906560 | 431 | 093450 | 7 |
| 54 | 797934 | 261 | 891115 | 170 | 906819 | 43 I | 093181 | 5 |
| 5555 | 795091 | 261 | 8 fror 3 | 1-0 | 90-077 | 43 I | 092923 | 5 |
|  | 795247 | 261 | 8,ogil | 170 | 90-336 | 431 | 092664 | 4 |
| 5 | 798403 | 260 | 890509 | 1-0 | 907594 | 43 I | 092406 | 3 |
| 59 | 793560 | 260 250 | 890707 | 170 | 90-853 | 431 430 | 092147 | 2 |
|  | 795716 | 250 | Sc90605 | 170 | 908111 | 430 | 091589 | 1 |
|  | 798872 | 260 | S90503 | 170 | 905369 | 430 | 091631 | $\bigcirc$ |
| I | Cosine. | . | Sine. | D. | Cotang. | D. | Tang. | , |
|  |  |  |  |  |  |  |  | $51^{\circ}$ |

Table II. LOGARITHMIC SINES, T ANGENTS, ETC.

| $39^{\circ} \mathrm{l} 140^{\circ}$ |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | Sine. | D. | Cosine. | D. | Tang. | D. | Cotang. | 1 |
| 0 | 9-798872 | 260 | $9 \cdot 890503$ | 170 | $9 \cdot 908369$ | 430 | 16.091631 | 60 |
| 1 | 799028 | 260 | 890400 | 171 | 908628 | 430 | 091372 | 59 |
| 2 | 799184 | 260 | 890298 | 171 | 908886 | 430 | 091114 | 58 |
| 3 | 799339 | 259 | 890195 | 171 | 909144 | 430 | 090856 | 57 |
| 4 | 799495 | 259 | \$90093 | 171 | 909402 | 430 | ogo598 | 56 |
| 5 | 799651 | 259 | 889990 | 171 | 909660 | 430 | 090340 | 55 |
| 6 | 799806 | 259 | 889888 | 171 | 909918 | 430 | 090082 | 54 |
| 7 | 799962 | 259 | 889785 | 171 | 910177 | 430 | 089823 | 53 |
| 8 | 800117 | 259 | 889682 | 171 | 910435 | 430 | -89565 | 52 |
| 9 | 800272 | 258 | 889579 | 171 | 910693 | 430 | 089307 | 51 |
| 10 | 800427 | 258 | 889477 | 171 | 910951 | 430 | 089049 | 50 |
| 11 | c. 800582 | 258 | G. 8890374 | 172 | 9.911209 | 430 | 10.088791 | 49 |
| 12 | 800737 | 258 | 889271 | 172 | 911467 | 430 | 088533 | 48 |
| 13 | 800892 | 258 | 889168 | 172 | 911725 | 430 | 088275 | 47 |
| 14 | 801047 | 258 | 889064 | 172 | 911982 | 430 | 088018 | 46 |
| 15 | 801201 | 258 | 888961 | 172 | 912240 | 430 | 087760 | 45 |
| 16 | 801356 | 257 | 888858 | 172 | 912498 | 430 | 087502 | 44 |
| 17 | 801511 | 257 | 888755 | 172 | 912756 | 430 | 087244 | 43 |
| 18 | 801565 | 257 | 888651 | 172 | 913014 | 429 | 086986 | 42 |
| 19 | 80:819 | 257 | 888548 | 172 | 913271 | 429 | 086729 | 41 |
| 20 | 801973 | 257 | 888444 | 173 | 913529 | 429 | 086471 | 40 |
| 2 I | 9.802128 | 257 | 9.888341 | 173 | $9 \cdot 913787$ | 429 | $10 \cdot 086213$ | 39 |
| 22 | 802282 | 256 | 888237 | 173 | 914044 | 429 | -85956 | 38 |
| 23 | 802436 | 256 | 888134 | 173 | 914302 | 429 | 085698 | 37 |
| 24 | 802589 | 256 | 888030 | $17^{3}$ | 914560 | 429 | 085440 | 36 |
| 25 | 802743 | 256 | 887926 | 173 | 914817 | 429 | 085183 | 35 |
| 26 | 802897 | 256 | 887822 | 173 | 915075 | 429 | 084925 | 34 |
| 27 | 803050 | 256 | 887718 | 173 | 915332 | 429 | 084668 | 33 |
| 28 | 803204 | 256 | 887614 | 173 | 915590 | 429 | 084410 | 32 |
| 29 | 803357 | 255 | 887510 | 173 | 915847 | 429 | 084153 | 3 I |
| 30 | 803511 | 255 | 887406 | 174 | 916104 | 429 | 083896 | 30 |
| 31 | 9.803664 | 255 | 9.887302 | 174 | 9.916362 | 429 | 10.083638 | 29 |
| ${ }_{3}^{3} 2$ | 803817 | 255 | 887198 | 174 | 916619 | 429 | 083381 | 28 |
| 33 | 803970 | 255 | 887093 | 174 | 916877 | 429 | 083123 | 27 |
| 34 | 804123 | 255 | 886989 | 174 | 917134 | 429 | 082866 | 26 |
| 35 | 80.4276 | 254 | 886885 | 174 | 917391 | 429 | 082609 | 25 |
| 36 | 804428 | 254 | 886780 | 174 | 917648 | 429 | 082352 | 24 |
| 37 | 804581 | 254 | 886676 | 174 | 917906 | 429 | 082094 | 23 |
| 38 | 804734 | 254 | 886571 | 174 | 918163 | 428 | 081837 | 22 |
| 39 | 804886 | 254 | 886466 | 174 | 918420 | 428 | 081580 | 21 |
| 40 | 805039 | 254 | 886362 | 175 | 918677 | 428 | 081323 | 20 |
| 41 | 9.805191 | 254 | 9.886257 | 175 | 9.918934 | 428 | 10.081066 |  |
| 42 | 805343 | 253 | 886152 | 175 | 919191 | 428 | 080809 | 18 |
| 43 | 805495 | 253 | 886047 | 175 | 919448 | 428 | -80552 | 17 |
| 44 | 805647 | 253 | 885942 | 175 | 919785 | 428 | 080295 | 16 |
| 45 | 805799 | 253 | 885837 | 175 | 919962 | 428 | 080038 | 15 |
| 46 | 805951 | 253 | 885732 | 175 | 920219 | 428 | 079781 | 14 |
| 47 | 806103 | 253 | 885627 | 175 | 920476 | 428 | 079524 | 13 |
| 48 | 806254 | 253 | 885522 | 175 | 920733 | 428 | 079267 | 12 |
| 49 | 806406 | 252 | 885416 | 175 | 920990 | 428 | 079010 | 11 |
| 50 | 806557 | 252 | 885311 | 176 | 921247 | 428 | 078753 | 10 |
| 51 | 9.806709 | 252 | 9.885205 | 176 | $7 \cdot 921503$ | 428 | $10 \cdot 078497$ | 8 |
| 52 | 806860 | 252 | 885100 | 176 | 921760 | 428 | 078240 | 8 |
| 53 | 807011 | 252 | 884994 | 176 | 922017 | 428 | 077983 | 7 |
| 54 | 807163 | 252 | 884889 | 176 | 922274 | 428 | 077726 | 6 |
| 55 | 807314 | 252 | 884783 | 176 | 922530 | 428 | 077470 | 5 |
| 56 | 807465 | 251 | 884677 | 176 | 922787 | 428 | 077213 | 4 |
| 57 | 807615 | 251 | 884572 | 176 | 923044 | 428 | 076956 | 3 |
| 58 | 807766 | 251 | 884466 | 176 | 923300 | 428 | 076700 | 2 |
| 59 | 807917 | 251 | 884360 | 176 | 923557 | 427 | 076443 | 1 |
| 60 | 808067 | 251 | 884254 | 177 | 923814 | 427 | 076186 | 0 |
| , | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | 1 |
| 12 |  |  |  |  |  |  |  | $50^{\circ}$ |


| 58 | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  |  | . Tabl | 11. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |  |  | $139^{\circ}$ |  |
| , | Sine. | D. | Cosine. | D. | Tang. | D. | Cotang. | , |
| $\begin{array}{r} 0 \\ 1 \\ 2 \\ 3 \\ 4 \\ 5 \\ 6 \\ 7 \\ 8 \\ 9 \\ 10 \end{array}$ | 9.808067 | 251 | 9.884254 | 177 | $9 \cdot 923814$ | 427 | 10.076186 | 60 |
|  | 808218 | 251 | 884148 | 177 | 924070 | 427 | 075930 | 50 |
|  | 808368 | 251 | 884042 | 177 | 924327 | 427 | 075673 | 58 |
|  | 808519 | 250 | 883936 | 177 | 924583 | 427 | 075417 | 57 |
|  | 808669 | 250 | 883829 | 177 | 924840 | 427 | 075160 | 56 |
|  | 808819 | 250 | 883723 | 177 | 925006 | 427 | 074904 | 55 |
|  | 808969 | 250 | 883617 | 177 | 925352 | 427 | 074648 | 54 |
|  | 809119 | 250 | 883510 | 177 | 923609 | 427 | 074391 | 53 |
|  | 809269 | 250 | 883404 | 177 | 925865 | 427 | 074135 | 32 |
|  | 809419 809569 | 249 249 | 883297 883191 | 178 | 926122 926378 | 427 | 073878 073622 | 51 50 |
|  | $\begin{array}{r}809369 \\ \hline 809\end{array}$ | 249 |  | 178 |  | 427 |  | 50 |
| 11 | $9 \cdot 809718$ 809868 | 249 249 | 9.883084 | 178 178 | 9.926634 | 427 | 10.073366 073110 | 49 |
| 12 | 809868 810017 | 249 249 | 882977 882871 | 178 178 178 | 926890 927147 | 427 | 073110 072853 | 48 |
| 13 | 810017 810167 | 249 249 | 882871 882764 | 178 178 178 | 927147 927403 | 427 427 | 072853 072597 | 47 |
| 15 | 810316 | 248 | 882657 | 178 | 927659 | 427 | 072341 | 45 |
| 16 | 810465 | 248 | 882550 | 178 | 927915 | 427 | 072085 | 44 |
| 1718 | 810614 | 248 | 882443 | 178 | 928171 | 427 | 071829 | 43 |
|  | 810763 | 248 | 882336 | 179 | 928427 | 427 | 071573 | 42 |
| 19 | 810912 | 248 | 882229 | 179 | 928684 | 427 | 071316 | 41 |
| 20 | 811061 | 248 | 882121 | 179 | 928940 | 427 | 071060 | 40 |
| 21 | 9.811210 | 248 | 9.882014 | 179 | 9.929196 | 427 | 10.070804 | 39 |
| 22 | 811358 | 247 | 881907 | 179 | 929452 | 427 | 070548 | 38 |
| 2 | 811507 | 247 | 881799 | 179 | 929708 | 427 | 070292 | 37 |
| 2 | 811655 | 247 | 881692 | 179 | 029964 | 426 | 070036 | 36 |
| 2 | 811804 | 247 | 881584 | 179 | 930220 | 426 | 069780 | 35 |
|  | 811952 | 247 | 881477 | 179 | 930475 | 426 | 069525 | 34 |
| 2 | 812100 | 247 | 881309 | 179 | 930731 | 426 | 069269 | 33 |
|  | 812248 | 247 | 881261 | 180 | 930987 | 426 | 069013 | 32 |
| 29 | 812396 | 246 | 881153 | 180 | 931243 | 426 | 068757 | 31 |
|  | 812544 | 246 | 881046 | 180 | 931499 | 426 | 068501 | 30 |
| 31 | 9.812692 | 246 | 9.880 og 38 | 180 | 9.931755 | 426 | 10.068245 | 29 |
| 32 | 812840 | 246 | 880830 | 180 | 9.32010 | 426 | 067990 | 28 |
| 3 | 812988 | 246 | 880722 | 180 | 932266 | 426 | 067734 | 27 |
| 3 | 813135 | 246 | 880613 | 180 | 932522 | 426 | 067478 | 26 |
| $\begin{array}{\|l} 36 \\ 36 \end{array}$ | 813283 | 246 | 880505 | 180 | 932773 | 426 | 067222 | 25 |
|  | 813430 | 245 | 880397 | 180 | 933033 | 426 | 066967 | 24 |
| $\begin{aligned} & 36 \\ & 3 \end{aligned}$ | 813578 | 245 | 880289 | 181 | 933289 | 426 | 066711 | 23 |
|  | 813725 | 245 | 880180 | 181 | 933545 | 426 | 066455 | 22 |
| 4 | 813872 | 245 | 880072 | 181 | 933800 | 426 | 066200 | 21 |
| 40 | 814019 | 245 | 879963 | 181 | 934056 | 426 | 065944 | 20 |
| 41 | 9.814166 | 245 | 9.879855 | 181 | 9.934311 | 426 | 10.065689 | 19 |
|  | 814313 | 245 | 879746 | 181 | 934567 | 426 | 065433 | 18 |
|  | 814460 | 244 | 879637 | 181 | 934822 | 426 | 065178 | 17 |
|  | 814607 | 244 | 879529 | 181 | 935078 | 426 | 064922 | 16 |
|  | 814753 | 244 | 879420 | 181 | 935333 | 426 | 064667 | 15 |
| $\begin{aligned} & 45 \\ & 46 \end{aligned}$ | 814900 | 244 | 879311 | 181 | 935589 | 426 | 064411 | 14 |
| $\begin{aligned} & 46 \\ & 47 \end{aligned}$ | 815046 | 244 | 879202 | 182 | 935844 | 426 | 064156 | 13 |
|  | $81519^{3}$ | 244 | 879093 | 182 | 936100 | 426 | 063900 | 12 |
|  | 815339 | 244 | 878984 | 182 | 936355 | 426 | 063645 | 11 |
|  | 815485 | 243 | 878875 | 18.2 | 936611 | 426 | 063389 | 10 |
| $\begin{aligned} & 51 \\ & 52 \\ & 53 \\ & 54 \\ & 55 \\ & 50 \\ & 57 \\ & 58 \\ & 59 \\ & 60 \end{aligned}$ | 9.815631 | 243 | 9.878766 | 183 | 9.936866 | 425 | -063134 | 9 |
|  | 815778 | 243 | 878656 | 182 | 937121 | 425 | 062879 | 8 |
|  | 815924 | 243 | 878547 | 182 | 937377 | 425 | 062623 | 7 |
|  | 816069 | 243 | 878438 | 182 | 937632 | 425 | 062368 | 6 |
|  | 816215 | 243 | 878328 | 182 | 937887 | 425 | 062113 | 5 |
|  | 816361 | 243 | 878219 | 183 | 938142 | 425 | 061858 | 4 |
|  | 816507 | 242 | 878109 | 183 | 938398 | 425 425 | 061602 061347 | 3 |
|  | 816652 816708 | 242 | 877999 877800 | 183 | 938653 | 425 | 061347 061092 | 2 |
|  | 816798 816943 | 242 242 | $87780^{\circ}$ 877780 | 183 | 938908 939163 | 425 425 | 061092 060837 | 1 |
|  | Cosine. | D. | Sine. | D. | Cotang. | ט. | Tang. | , |
| $131{ }^{\circ}$ |  |  |  |  |  |  |  | $49^{\circ}$ |



| 60 |  | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  | Table II. |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $42^{\circ}$ |  |  |  |  |  |  | $137^{\circ}$ |  |
| , | Sine. | D. | Cosine. | D. | Tang. | D. | Cotang. | , |
| 0 | 9.825511 | 234 | 9.871073 | 190 | 9.954437 | 423 | 10.045563 | 60 |
| 1 | 82565 r | 233 | 870960 | 190 | 954691 | 423 | 045309 | 59 |
| 2 | 825791 | 233 | 870846 | 190 | 954946 | 423 | 045054 | 58 |
| 3 | 825931 | 233 | $870-32$ | 190 | 955200 | 423 | 044800 | 57 |
| 4 | 826071 | 233 | 870618 | 190 | 955454 | 423 | 044546 | 56 |
| 5 | 826211 | 233 | 870504 | 190 | 955708 | 423 | 044292 | 55 |
| 6 | 826351 | 2.33 | 870390 | 190 | 955961 | 423 | 044039 | 54 |
| 7 | 826491 | 233 | 870275 | 190 | 956215 | 423 | 043785 | 53 |
| 8 | 826631 | 233 | 870161 | 190 | -956469 | 423 | 043531 | 52 |
| 9 | 826770 | 232 | 870047 | 191 | 956723 | 423 | 043277 | 51 |
| 10 | 826910 | 232 | 869933 | 191 | 956977 | 423 | 043023 | 50 |
| 11 | 9.827049 | 232 | 9.869818 | 191 | 9.957231 | 423 | 10.042769 | 49 |
| 12 | 827189 | 232 | 869704 | 191 | 957485 | 423 | 042515 | 48 |
| 13 | 827328 | 232 | 869589 | 191 | 957739 | 423 | 042261 | 47 |
| 14 | 827467 | 232 | 869474 | 191 | 957993 | 423 | 042007 | 46 |
| 15 | 827606 | 232 | 869360 | 191 | 958247 | 423 | 041753 | 45 |
| 16 | 827745 | 232 | 869245 | 191 | 958500 | 423 | 041500 | 44 |
| 17 | 827884 | 231 | 869130 | 191 | 958754 | 423 | 041246 | 43 |
| 18 | 828023 | 231 | 869015 | 192 | 959008 | 423 | 040992 | 42 |
| 19 | 828162 | 231 | 868900 | 192 | 9.59262 | 423 | 040738 | 41 |
| 20 | 828301 | 231 | 868785 | 192 | 959516 | 423 | 040484 | 40 |
| 21 | 9.828439 | 231 | 9.868670 | 192 | 9.959769 | 423 | 10.040231 | 39 |
| 22 | 828578 | 231 | 868555 | 192 | 960023 | 423 | 039977 | 38 |
| 23 | 828716 | 231 | 868440 | 192 | 960277 | 423 | 039723 | 37 |
| 24 | 828855 | 230 | 868324 | 192 | 960530 | 423 | 039470 | 36 |
| 25 | 828993 | 230 | 869209 | 192 | 960784 | 423 | 039216 | 35 |
| 26 | 829131 | 230 | 868093 | 192 | 961038 | 423 | -38962 | 34 |
| 27 | 829269 | 230 | 867978 | 193 | 961292 | 423 | 038708 | 33 |
| 28 | 829407 | 230 | 867862 | 193 | 961545 | 423 | 038455 | 32 |
| 29 | 829545 | 230 | 867747 | 193 | 961799 | 423 | 038201 | 31 |
| 30 | 829683 | 230 | 867631 | 193 | 962052 | 423 | 037948 | 30 |
| 31 | 9.829821 | 229 | 9.867515 | 193 | 9.962306 | 423 | 10.037694 |  |
| 32 | 829950 | 229 | 867399 | 193 | 962550 | 423 | -37440 | 28 |
| 33 | 830097 | 229 | 867283 | 193 | 962813 | 423 | 037187 | 27 |
| 34 | 830234 | 229 | 867167 | 193 | 963067 | 423 | 036933 | 26 |
| 35 | 830372 | 229 | 867051 | 193 | 963320 | 423 | 036680 | 25 |
| 36 | 830509 | 229 | 866935 | 194 | 963574 | 423 | 036426 | 24 |
| 37 | 830646 | 229 | 865819 | 194 | 963828 | 423 | 036172 | 2.3 |
| 38 | 830784 | 229 | 866703 | 194 | 964081 | 423 | 035919 | 22 |
| 39 | 830921 | 228 | 866586 | 194 | 964335 | 423 | 035665 | 21 |
| 40 | 831058 | 228 | 866470 | 194 | 964588 | 422 | 035412 | 20 |
| 41 | 9.831195 | 228 | 9.866353 | 194 | 9.964842 | 422 | 10.035158 | 19 |
| 42 | 831332 | 228 | 866237 | 194 | 965095 | 422 | 034905 | 18 |
| 43 | 831469 | 228 | 866120 | 194 | 965349 | 422 | 034651 | 17 |
| 44 | 831606 | 228 | 865004 | 195 | 965602 | 422 | 034398 | 16 |
| 45 | 831742 | 228 | 865887 | 195 | 965855 | 422 | 034145 | 15 |
| 46 | 831879 | 228 | 865770 | 195 | 966109 | 422 | 033891 | 14 |
| 47 | 832015 832152 | 227 | 865653 | 195 | 966362 | 422 | 033638 | 13 |
| 48 | 832152 832288 | 227 | 865536 | 195 | 966616 | 422 | 033384 | 12 |
| 49 50 | 832288 832425 | 227 | 865419 | 195 | 966869 | 422 | 033131 | 11 |
| 50 | 832425 | 227 | 865302 | $19^{5}$ | 967123 | 422 | 032877 | 10 |
| 51 | 9.832561 | 227 | 9.865185 | $19 \stackrel{5}{5}$ | 9.967376 | 422 | 10.0.32624 |  |
| 52 | 832697 | 227 | 865068 | 195 | 967629 | 422 | 032371 | 8 |
| 53 | 832833 | 227 | 864950 | 195 | 967883 | 422 | 032117 | 7 |
| 54 | 832969 | 226 | 864833 | 196 | 968136 | 422 | 0.31864 | 5 |
| 55 | 833105 | 226 | 864716 | 196 | 968389 | 422 | 031611 | 5 |
| 56 | 833241 | 226 | 864508 | 196 | 968643 | 422 | 0313.7 |  |
| 37 | 833377 | 226 | 864481 | 196 | 968896 | 422 | 031144 | 3 |
| 58 | 833512 | 226 | 864363 | 196 | 969149 | 422 | 030851 | 2 |
| 59 | 833648 | 226 | 864245 | 196 | 969403 | 422 | 030597 | 1 |
| 60 | 833783 | 226 | 864127 | 1.6 | 969656 | 422 | 030345 | $\bigcirc$ |
| 1 | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | , |
| $132^{\circ}$ |  |  |  |  |  |  | $47^{\circ}$ |  |


| Table II. |  | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  |  | 61 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $43^{\circ}$ |  |  |  |  |  |  |  | $136^{\circ}$ |
| , | Sine. | D. | Cosine. |  |  | D. | Cotarg. | 1 |
| 0 | 9.833783 | 226 | 9.864127 | 196 | 9.969656 | 422 | 10.030344 | 60 |
| 1 | 833919 | 225 | 864010 | 196 | 969409 | 422 | 030001 | 5 |
| 2 | 834054 | 225 | 863892 | 197 | 970162 | 422 | 029838 | 58 |
| 3 | 834189 | 225 | 863774 | 197 | 970416 | 422 | 029584 | 57 |
| 4 | 834325 | 225 | 863656 | 197 | 970669 | 422 | 029331 | 56 |
| 5 | 834460 | 225 | 863538 | 197 | 970922 | 422 | 029078 | 55 |
| 6 | 834595 | 225 | 853419 | 197 | 971175 | 422 | 028825 | 54 |
| 7 | 834730 | 22.5 | 363301 | 197 | 971429 | 422 | 028571 | 53 |
| 8 | 834865 | 225 | 863183 | 197 | 971682 | 422 | 028318 | 52 |
| 9 | 834999 | 224 | 863064 | 197 | 971935 | 422 | 028065 | 51 |
| 10 | 835134 | 224 | 862946 | 198 | 972188 | 422 | 027812 | 50 |
| 11 | 9.835269 | 224 | 9.862827 | 198 | 9.972441 | 422 | 10.027559 | 49 |
| 12 | 835403 | 224 | 862709 | 198 | 972695 | 422 | 027305 | 48 |
| 13 | 835538 | 224 | 862590 | 198 | 772948 | 422 | 027052 | 47 |
| 14 | 835672 | 224 | 862471 | 198 | 973201 | 422 | 026799 | 46 |
| 15 | 835807 | 224 | 862353 | 198 | 973454 | 422 | 026546 | 45 |
| 16 | 835941 | 224 | 862234 | 198 | 973707 | 422 | 026293 | 44 |
| 17 | 836075 | 223 | 862115 | 198 | 973960 | 422 | 026040 | 43 |
| 18 | 836200 | 223 | 861996 | 198 | 974213 | 422 | 025787 | 42 |
| 19 | 836343 | 223 | 86.877 | 198 | 974466 | 422 | 025534 | 41 |
| 20 | 836477 | 223 | 861758 | 199 | 974720 | 422 | 025280 | 40 |
| 21 | 9.836611 | 223 | 9.861638 | 199 | 9.974973 | 422 | 10.025027 | 39 |
| 22 | 836745 | 223 | 861519 | 199 | 975226 | 422 | 024774 | 38 |
| 23 | 836878 | 223 | 861400 | 199 | 975479 | 422 | 024521 | 37 |
| 24 | 837012 | 222 | 861280 | 199 | 975732 | 422 | 024268 | 36 |
| 25 | 837146 | 222 | 861161 | 199 | 975985 | 422 | 024015 | 35 |
| 26 | 837279 | 222 | 861041 | 199 | 976238 | 422 | 023762 | 34 |
| 27 | 837412 | 222 | 860922 | 199 | 976491 | 422 | 023509 | 33 |
| 28 | 837546 | 222 | 860802 | 199 | 976744 | 422 | 023256 | 32 |
| 29 | $837679{ }^{\circ}$ | 222 | 860682 | 200 | 976997 | 422 | 023003 | 31 |
| 30 | 837812 | 222 | 860562 | 200 | 977250 | 422 | 022750 | 30 |
| 31 | 9.837945 | 222 | 9.860442 | 200 | 9.977503 | 422 | 10.022497 | 29 |
| 32 | 838078 | 221 | 860322 | 200 | 977756 | 422 | 022244 | 28 |
| 33 | 838211 | 221 | 860202 | 200 | 978009 | 422 | 021991 | 27 |
| 34 | 838344 | 221 | 860082 | 200 | 978262 | 422 | 021738 | 26 |
| 35 | 838477 | 221 | 859962 | 200 | 978515 | 422 | 021485 | 25 |
| 36 | 838610 | 221 | 859842 | 200 | 978768 | 422 | 021232 | 24 |
| 37 | 838742 | 221 | 859721 | 201 | 979021 | 422 | 020979 | 23 |
| 38 | 838875 | 221 | 859601 | 201 | 979274 | 422 | 020726 | 22 |
| 39 | 839007 | 221 | 859480 | 201 | 979527 | 422 | 020473 | 21 |
| 40 | 839140 | 220 | 859360 | 201 | 979780 | 422 | 020220 | 20 |
| 41 | 9.839272 | 220 | 9.859239 | 201 | 9.980033 | 422 | 10.019967 | 19 |
| 42 | 839404 | 220 | 859119 | 201 | 980286 | 422 | 019714 | 18 |
| 43 | 839536 | 220 | 858998 | 201 | 980538 | 422 | 019462 | 17 |
| 44 | 839668 | 220 | 858877 | 201 | 980791 | 421 | 019209 | 16 |
| 45 | 839800 | 220 | 858756 | 202 | 981044 | 421 | -18956 | 15 |
| 46 | 839932 | 220 | 858635 | 202 | 981297 | 421 | -18703 | 14 |
| 47 | 840064 | 219 | 858514 | 202 | 981550 | 421 | -18450 | 13 |
| 48 | 840196 | 219 | 858393 | 202 | 981803 | 421 | 018197 | 12 |
| 49 | 840328 | 219 | 858272 | 202 | 982056 | 421 | 017944 | 11 |
| 50 | 840459 | 219 | 858151 | 20 | 982309 | 421 | 017691 | 10 |
| 51 | ¢. 840591 | 219 | 9.858029 | 202 | 9.982562 | 421 | 10.017438 |  |
| 52 | 840722 | 219 | 857908 | 2 C 2 | 982814 | 421 | 017186 | 8 |
| 53 54 | 840854 | 219 | 857786 | 202 | 983067 | 421 | 016933 | 7 |
| 54 55 5 | 840985 | 219 | 857665 | 203 | 983320 | 421 | 016680 | 5 |
| 55 56 | 841116 | 218 | 857543 | 203 | 983573 | 421 | 016427 | 5 |
| 56 | 841247 | 218 | 857422 | 203 | 983826 | 421 | 016174 | 3 |
| 57 58 | 841378 | 218 | 857300 | 203 | 984079 | 421 | 015921 | 3 |
| 58 50 | 841509 841640 | 218 218 | 877178 857056 | 203 | 984332 | 421 | O15668 | 2 |
| 60 | 8841771 | 218 218 | 857093 85694 | 203 | 984584 984837 | 421 | 015416 0.5163 | 1 |
| ' | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | , |
| $133^{\circ}$ |  |  |  |  |  |  |  | $46^{\circ}$ |


| 62 | LOGARITHMIC SINES, TANGENTS, ETC. |  |  |  |  |  | TABL | II. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $44^{\circ}$ |  | D. | Cosine. | D. | Tang. | D. | $135^{\circ}$ |  |
| , | Sine. |  |  |  |  |  | Cotang. | , |
| 0 | 9.841771 | 218 | 9.856934 | 203 | 9.984837 | 421 | $10 \cdot 015163$ | 60 |
| 1 | 841902 | 218 | 856812 | 203 | 985090 | 421 | 014910 | 5 |
| 2 | 842033 | 218 | 856690 | 204 | 985343 | 421 | 014657 | 58 |
| 3 | 842163 | 217 | 856568 | 204 | 985596 | 421 | 014404 | 57 56 |
| 4 | 842294 | 217 | 856446 856323 | 204 | 985848 | 421 | 014152 | 56 |
| 6 | 842555 | 217 217 | 856201 | 204 | 986355 | 421 | -113646 | 54 |
| 7 | 842685 | 217. | 856078 | 204 | 986607 | 421 | 013393 | 53 |
| 8 | 842815 | 217 | 855056 | 204 | 986860 | 421 | 013140 | 52 |
| 9 | 842946 | 217 | 855833 | 204 | 987112 | 421 | 012888 | 51 |
| 10 | 843076 | 217 | 855711 | 205 | 987365 | 421 | O12635 | 50 |
| 11 | 9.843206 | 216 | 9.855588 | 205 | 9.987618 | 421 | 10.012382 | 49 |
| 12 | 843336 | 216 | 855465 | 205 | 987871 | 421 | 012129 | 48 |
| 13 | 843466 | 216 | 855342 | 205 | 988123 | 421 | 011877 | 47 |
| 14 | 843595 | 216 | 855219 | 205 | 988376 | 42 I | 011624 | 46 |
| 15 | 843725 | 216 | 855096 | 205 | 988629 | 421 | ori371 | 45 |
| 16 | 843855 | 216 | 854973 | 205 | 988882 | 421 | 011118 | 44 |
| 17 | 843984 | 216 | 854850 | 205 | 089134 | 421 | 010866 | 43 |
| 18 | 844114 | 215 | 854727 | 206 | 989387 | 421 | -10613 | 42 |
| 19 | 844243 | 215 | 854603 | 206 | 989640 | 421 | -ro360 | 41 |
| 20 | 844372 | 215 | 854480 | 206 | 989893 | 42 I | 010107 | 40 |
| 21 | 9.844502 | 215 | 9. 854356 | 206 | 7.990145 | 421 | $10 \cdot 009855$ | 39 |
| 22 | 84463I | 215 | 854233 | 206 | 990398 | 421 | 009602 | 38 |
| 23 | 844760 | 215 | 854109 | 206 | 990651 | 421 | 009349 | 37 |
| 24 | 844889 | 215 | 853986 | 206 | 990903 | 42 I | 009097 | 36 |
| 25 | 845018 | 215 | 853562 | 206 | 991156 | 421 | 008844 | 35 |
| 26 | 845147 | 215 | 853738 | 206 | 991409 | 421 | 008591 | 34 |
| 27 | 845276 | 214 | 853614 | 207 | 991662 | 421 | 008338 | 33 32 |
| 28 | 845405 | 214 | 853490 | 207 | 991914 | 421 | 008086 | 32 31 1 |
| 30 | 845662 | 214 | 853242 | 207 | 992420 | 421 | 007580 | 30 |
| 31 | 9.845790 | 214 | 9.853118 | 207 | 9.992672 | 42 I | 10.007328 | 29 |
| 32 | 845919 | 214 | 852994 | 207 | 992925 | 421 | 007075 | 28 |
| 33 | 846047 | 214 | 852869 | 207 | 993178 | 421 | 006822 | 27 |
| 34 | 846175 | 214 | 852745 | 207 | 993431 | 42 I | 006569 | 26 |
| 35 | 846304 | 214 | 852620 | 207 | 993683 | 421 | 006317 | 25 |
| 36 | 846432 | 213 | 852496 | 208 | 993936 | 421 | 006064 | 24 |
| 37 | 846560 | 213 | 852371 | 208 | 994189 | 42 I | -o5์811 | 23 |
| 38 | 846688 | 213 | 852247 | 208 | 9.94441 | 421 | 0055̃59 | 22 |
| 39 | 846816 | 213 | 852122 | 208 | 994694 | 421 | -050306 | 21 |
| 40 | 846944 | 213 | 851997 | 208 | 994947 | 421 | -05053 | 20 |
| 41 | $9 \cdot 847071$ | 213 | 9.851872 | 208 | 9.995199 | 421 | $10 \cdot 004801$ | 19 |
| 42 | 847199 | 213 | 851747 | 208 | 995452 | 421 | 004548 | 18 |
| 43 | 847327 | 213 | 851622 | 208 | 995705 | 421 | 004295 | 17 |
| 44 | 847454 | 212 | 85.1497 | 209 | 995957 | 421 | 004043 | 16 |
| 45 | 847582 | 212 | 85.372 | 209 | 996210 | 421 | 003790 | 15 |
| 46 | 847709 847836 | 212 | 851246 | 209 | 996463 | 421 | 003537 | 14 |
| 47 | 847836 | 212 | 851121 | 209 | 996715 | 421 | 003285 | 13 |
| 48 | 847964 | 212 | 850996 | 209 | 996968 | 421 | 003032 | 12 |
| 49 | 848091 | 212 | 850870 | 209 | 997221 | 42 I | 002779 | 11 |
| 50 | 848218 | 212 | 850745 | 209 | 997473 | 42 I | 002527 | 10 |
| 51 | ¢. 848345 | 212 | 9.850619 | 209 | 9.997726 | 42 I | 10.002274 |  |
| 52 | 848472 | 211 | 850493 | 210 | 997979 | 421 | 002021 | 8 |
| 53 | 848599 | 211 | 850368 | 210 | 998231 | 421 | 001769 | 7 |
| 54 | 848726 | 211 | 850242 | 210 | 998484 | 421 | 001516 | 6 |
| 55 | 848852 | 211 | 850116 | 210 | 998737 | 421 | 001263 | 5 |
| 56 | 848979 | 211 | 849990 | 210 | 998989 | 421 | 001011 | 4 |
| 57 | 849106 | 211 | 849864 | 210 | 999242 | 421 | 000758 | 3 |
| 58 | 849232 | 211 | 849738 | 210 | 999.195 | 421 | 000505 | 2 |
| ${ }_{6}^{5}$ | 849359 849485 | 211 | 849611 849485 | 210 210 | 999747 $10 \cdot 000000$ | 421 421 | 000253 10.000000 | 1 |
| , | Cosine. | D. | Sine. | D. | Cotang. | D. | Tang. | , |
| $134^{\circ}$ |  |  |  |  |  |  |  | $45^{\circ}$ |

## TABLE lll.,

# OF <br> NATURAL SINES AND TANGENTS; 

TO

## EVERY DEGREE ANL MINUTE OF THE QUADRANT.

Ir the given angle is less than $45^{\circ}$, look for the degrees and the title of th.e column, at the top of the page; and for the minutes on the left. But if the arigle is between $45^{\circ}$ and $90^{\circ}$, look for the degrees and the title of the column, at the oottom; and for the minutes on the right.

The Secants and Cosecants, which are not inserted in this table, may be easily supplied. If 1 be divided by the cosine of an are, the quotient will be the secant of that arc. And if I be divided by the sine, the quotient will be the cosecant.

The values of the Sines and Cosines are less than a unit, and are given in decimals, although the decimal point is not printed. So also, the tangents of arce lees than $4^{\circ}$, and cotangents of arcs greater than $4^{\circ}$, are less that a unit ani we expressed in dezimals with the decinca. point onitted.

| 64 |  |  | NATURAL SINES A |  |  |  | COSINES. |  | Table III. |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $0^{\circ}$ |  | $1^{\circ}$ |  | $2^{\circ}$ |  | $3^{\circ}$ |  | $4^{\circ}$ |  |  |
|  | Sine. | Cosine | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosine | Sine. | Cosine |  |
| 0 | 00000 | Unit. |  | 99985 | 03490 |  | 05234 |  |  |  | So |
| 1 | 00029 | Unit. | 01774 | 99984 | 03519 |  | -5263 |  | 07005 | 99754 | 59 |
| 2 | 00058 | Unit. | 01803 | 99984 | 03548 |  | 05292 |  | 07034 | 99752 | 58 |
| 3 | 00087 | Unit. | OI832 | 99983 | 03577 |  | 05321 |  | 07063 | 99750 | 57 |
| 4 | 00116 | Unit. | -1862 | 99983 | 03606 | 99935 | -5350 |  | 07092 | 99748 | 56 |
| 5 | 00145 | Unit. | or891 | 99982 | o3635 | 99934 | -5379 |  | 07121 | 99746 | 55 |
| 6 | 00175 | Unit. | -1920 | 99982 | 03664 | 99933 | 05408 | 99854 | 07150 | 99744 | 54 |
| 7 | 00204 | Unit. | 01949 | 99981 | 03693 | 99932 | 05.437 | $9985{ }^{2}$ | 07179 | 99742 | 53 |
| 8 | 00233 | Unit. | 01978 | 99980 | 03723 | 99931 | 05566 | 9985 I | 07208 | 99740 | 52 |
| 9 | 00262 | Unit. | 02007 | 99980 | 03752 | 99930 | -55495 | 99849 | 07237 | 99738 | 51 |
| 10 | 00291 | Unit. | 02036 | 99979 | 03781 | 99929 | 05524 |  | 07266 | 99736 | 50 |
| 11 | 00320 | 99999 | 02065 | 99979 | -03810 | 99927 | 05553 | 99846 | 07295 | 99734 | 49 |
| 12 | 00349 | 99999 | 02094 | 99978 | - 3839 | 99926 | 055j82 | 99844 | 07324 | 99731 | 48 |
| 13 | 00378 | 99999 | 02123 | 99977 | 03868 | 99925 | 05611 | 99842 | 07353 | 99729 | 47 |
| 14 | 00407 | 99999 | 02152 | 99977 | 03897 | 99924 | 05640 | 99841 | 07382 | 99727 | 46 |
| 15 |  | 99999 | 02181 | 99976 | 03926 | 99923 |  | 99839 | 07411 | 99725 | 45 |
| 16 | 00465 | 99999 | 2211 | 99976 | 03955 | 99922 | 05698 | 99838 | 07440 | 99723 | 44 |
| 17 | 00495 | 99999 | 02240 | 99975 | 03984 | 99921 | 05727 | 99836 | 07469 | 99721 | 43 |
| 18 | 00524 | 99999 | 02269 | 99974 | 0.4013 | 99919 | 05756 | 99834 | 07498 | 99719 | 42 |
| 19 | 00553 | 99998 | 02298 | 99974 | 0.4042 | 99918 | 05785 | 99833 | 07527 | 99716 | 41 |
| 20 | 00582 | 99998 | 02327 | 99973 | 04071 | 99917 | 05814 | 99831 | 075056 | 99714 | 40 |
| 21 | 00611 | 99998 | 02356 | 99972 | 04100 | 99916 | 05844 | 99829 | 07585 | 99712 | 30 |
| 22 | 00640 | 99998 | 02385 | 99972 | 04129 | 99915 | -5873 | 99827 | 07614 | 99710 | 38 |
| 23 | 00669 | 99998 | 02414 | 99971 | 04159 | 99913 | 05902 | 99826 | 07643 | 99708 | 37 |
| 24 | 000́98 | 99998 | 02 | 99970 | 04188 | 99912 | 05931 | 99824 | 07672 | 99705 | 36 |
| 25 | 00727 | 99997 | 0.2 | 99969 | 04217 | 99911 | 05960 | 99822 | 07701 | 99703 | 35 |
| 26 | 00756 | 99997 | 02501 | 9996 | 0.4246 | 99910 | 05989 | 99821 | 07730 | 99701 | 34 |
| 27 | 00785 | 99997 | 02530 | 99968 | 04275 | 99909 | 06018 |  | 07759 | 99699 | 33 |
| 28 | 00814 | 99997 | 02560 | 99967 | 04304 | 99907 | 06047 |  | 07788 |  | 32 |
| 29 | 00844 | 99996 | 02589 | 99966 | 04333 | 99906 | 06075 |  | 07817 | 99694 | 31 |
| 30 | 00 | 99996 |  | 99966 | 0.4362 | 9000 |  | 99 |  | 99692 |  |
| 31 | 00902 | 99996 |  | 99965 | 0.4391 | 99904 | 06 | 99812 | 07875 | 99689 | 29 |
| 32 | 00931 | 99996 | 02676 | 99964 | 04420 | 79902 | 06163 | 99810 | 07904 |  | 28 |
| 33 | 00960 | 99995 | 02705 | 999'33 | 04449 | 99901 | 06192 | 99808 | 07933 | 99685 | 27 |
| 34 | $0^{00989}$ | 99995 | 02734 | 99963 | 04478 | 99900 | 06221 | 99806 | 07962 |  | 26 |
| 35 | oior8 | 99995 | 02763 | 99962 | 04507 | 99898 | 06250 | 99804 | 07991 |  | 25 |
| 36 | 01047 | 99995 | 02792 | 99961 | 04536 |  | 06279 | 99803 | 08020 |  | 24 |
| 3 | 01076 | 99994 | 02821 | 99960 | 04565 |  | 06308 | 99801 | 08049 |  | 23 |
| 38 | orios | 99994 | 02350 | 99959 | 04594 | 99 | 06337 | 99799 | 08078 |  | 22 |
| 39 | 01134 | 99994 | 02879 | 999 | 0.4623 | 99893 | 06366 | 99797 | 08107 |  | 21 |
| 40 | 01164 | 9999 | 02908 | 99 | 04653 | 99892 | 06395 | 99795 | 08136 |  | 20 |
| 41 | 01193 | 99953 | 02938 |  | 04682 |  | 06424 | 99793 | 08165 | 99666 | 1 |
| 42 | O1222 | 99993 | 02967 | 99956 | 0.7711 |  | 06453 | 99792 | 08194 | 99664 | 18 |
| 43 | O125I | 99992 | 02996 | 99955 | 0.7740 |  | 064 42 | 99790 | 08223 | 99661 | 17 |
| 44 | O1280 | 99992 | 03025 | 99954 | 0.769 |  | 06511 | 99788 | 08252 | 99659 | 16 |
| 45 | or309 | 99991 | 03054 | 99953 | 04798 | 99885 | 06540 | 99786 | I | 99657 | 15 |
| 46 | or338 | 99991 | 03083 | 99952 | 04827 | 99883 | 06569 | 99784 | 08310 | 99654 | 14 |
| 47 | -1367 | 99991 | 03II2 | 99952 | 04856 | 99882 | 06598 | 99782 | 08339 | 99652 | 13 |
| 48 | -1396 | 99990 | 03141 | 99951 | 04885 | 99881 | 06627 | 99780 | 08368 | 99649 | 12 |
| 49 | 01425 | 99990 | 03170 | 99950 | 04914 |  | 06656 | 99778 | 08397 | 99647 | 11 |
| 50 | 01454 | 99989 | -3199 | 99949 | 04943 |  | 06685 | 99776 | 08426 | 99644 | 10 |
| 51 | 01483 | 99989 | 03228 | 99948 | 04972 |  | 06714 | 99774 | 08455 | 99642 |  |
| 52 | -15́r3 |  | 03257 | 99947 | 05001 |  | 06743 | 99772 | 08 84 | 99639 |  |
| 53 | 01542 | 99988 | 03286 | 99946 | -5030 | 99873 | 06773 | 99770 | 08513 | 99637 |  |
| 54 | 01571 | 99988 | -03316 | 99945 | -5059 | 99872 | 06802 | 99768 | 085. 42 | 99635 |  |
| 55 | 01600 | 99987 | 03345 | 99944 | 05088 | 99870 | 0683I | 99766 | 08571 | 99632 |  |
| 56 | 01629 | 9998 | o3374 | 99943 | 05117 | 99869 | 06860 | 99764 | 08600 | 99630 |  |
| 57 | -1658 | 99986 | 0340.3 | 99942 | -55146 | 9986 | 06889 | 99762 | -8629 | 99627 |  |
| 58 | 01687 | 999 | 03432 | 99941 | 05175 | 99866 | 06918 | 99760 | 08658 | 99625 |  |
| 59 | 01716 | 99 | 03461 | 99940 | OJ5205 | 99 | 069 ¢7 | 99758 | 08687 | 99622 |  |
| 00 | 01745 | 99 | 03.490 | 99939 | 05 | 99 | 06976 | 99756 | 08716 | 99619 |  |
| , | ne. | ine. | ne. | ine. | Cosine. | Sine. | Cosine. | Sine. | sine. | Sine |  |
|  | $89^{\circ}$ |  | $88^{\circ}$ |  | $87^{\circ}$ |  | $86^{\circ}$ |  | $85{ }^{\circ}$ |  |  |


| Table III. |  |  | NATURAL SINES AND COSINES. |  |  |  |  |  |  |  | 65 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $5^{\circ}$ |  | $6^{\circ}$ |  | $7{ }^{\circ}$ |  | $8^{\circ}$ |  | $9^{\circ}$ |  | , |
|  | Sine | Cosin | Sine. | Cosi | e. | Cosin | ne. | Cosin | Sine. | Cosine. |  |
| $\bigcirc$ |  |  |  |  |  |  | 13917 |  | 15643 |  | 6 |
| 1 | 08745 | 99617 | 10482 | 99449 | 12216 | 99251 | 13946 | 99023 | 15672 |  | 59 |
| 2 | 087 | 99614 | 10511 | 99446 | 12 | 99248 | 13975 | 99019 | 15701 | 98760 | 58 |
| 3 | 08803 | 99612 | 1050 | 99443 | 12274 | 99244 | 14004 | 99015 | 15730 | 98755 | 57 |
| 4 | 0883I | 99609 | 10569 | 99440 | 12 | 99240 | 14033 | 99011 | 15758 | 98751 | 56 |
| 5 | 08860 | 9960 | 10597 | 99437 | 12 | 99237 | 14061 | 99006 | 15787 | 98746 | 55 |
| 6 | 08889 | 99604 | 10626 | 99434 | 12 | $99^{233}$ | 14090 | 99002 | 15816 | 98741 | 5 |
| 7 | 08918 | 99602 | 10655 | 99431 | 12 | 99230 | 14119 | 98998 | 15845 |  | 53 |
| 8 | 08947 | 29509 | 1068 | 99428 | 12 | 99226 | 14148 | 98994 | 15873 |  | 52 |
| 9 | 08976 | 99596 | 10713 | 99424 | 12447 | 99222 | 14177 | 98990 | 15902 | 98728 | 51 |
| 10 | 0,005 | 99594 | 10742 | 9942 I | 12476 | $99^{219}$ | 14205 | 98986 | 15931 | 98723 | 50 |
| 11 | 09034 | 99591 | 10771 | 9948 | 12504 | 99215 | 14234 | 98982 | 15959 | 98718 |  |
| 12 | -9063 | 99588 | 10800 | 99415 | 125 | 99211 | 14263 | 98978 | 15988 | 98714 | 4 |
| 13 | 09092 | 99588 | 10829 | 99.12 | 125 | 99208 | 14292 | 98973 | 16017 | 98709 | 47 |
| 14 | 09121 | 99583 | 10858 | 99409 | 12591 | 99204 | 14320 | 98969 | 16046 | 98704 | 46 |
| 15 | 09150 | 99580 | 108 | 99406 | 12620 | 99200 | 14 | 98965 | 16074 | 98700 | 45 |
| 16 | 091 | 995 | 10 | 99.102 |  | 99 | 14378 | 98961 | 16103 | 98695 |  |
| 17 | 09208 | 995 | 10945 | 99399 | 126 | $99^{1} 2^{3}$ | 14407 | 98957 | 16132 | 98690 | 4 |
| 18 | 09237 | 99572 | 10973 |  | 127 | 99189 | 14436 | 98953 | 16160 | 98686 | 42 |
| 19 | 09266 | 99570 | 11002 | 99 | 127 | 9918 | 14 | 98948 | 16189 | 98681 | 41 |
| 20 | 09295 | $99^{507}$ | 11031 | 99 | 12764 | 99182 | 14493 | 98944 | 16218 | 98676 | 40 |
| 21 | 09324 | 99564 | 11 | 993 | 127 | 99178 | 14522 | 98940 | 16246 | 98671 | 30 |
| 22 | 09353 | 99562 | 11089 | 99383 | 12922 | 99175 | 14551 | 98936 | 16275 | 98667 | 38 |
| 23 | 09382 | 99559 | I 1 | 99380 | 1285 I | 99171 | 14580 | 98931 | 16304 | 98662 | 3 |
| 24 | 094II | 9955 | 11 | 99377 | 80 | 99167 | 14608 | 98927 | 16333 | 98657 | 36 |
| 25 | 09440 | 99553 | 11176 | 99374 | 12908 | 99163 | 1463 | 98923 | 16361 | 98652 | 35 |
| 26 | 09469 | 99551 | 11205 | 99370 | 12937 | 99160 | 14666 | 98919 | 16390 | 98648 | 3 |
| 27 | 09498 | 995 | 11 | 99367 | 12966 | 99156 | 1469 | 98914 | 16419 | 98643 | 3 |
| 28 | 09527 | 9954 | 11263 | $99^{36}$ | 129 | 99152 | 14723 | 98910 |  |  | 32 31 1 |
| 29 | 0955 | 99542 | 11 | 99360 | 13 | 99148 | 14752 |  |  | 98633 | 3 I |
| 30 | 09 | 99540 |  | 99357 |  | 99 |  |  |  |  |  |
| 3 I | 09614 | 99537 | 11342 | 99354 | 13081 | 99141 | 14810 |  | 33 |  | 29 |
| 32 | 09642 | 99534 | 11378 | 99351 | 131 | 99137 | 14838 |  | 16562 |  | 28 |
| 33 | 09571 | 9953I | 11407 | 99347 | 131 | 99133 | 14867 |  | 16591 | 98614 | 27 |
| 34 | 09700 | 99528 | 11436 | 99344 | 1316 | 99129 | 14896 | 988 | 16620 |  | 26 |
| 35 | 09729 | 99526 | 11465 | 99341 | 13197 | 99125 | 14925 | 988 | 16648 | 78604 | 25 |
| 36 | 09758 | 99 | 11494 | 99337 | 13226 | 99122 | 14954 |  | 16677 | 98600 | 2 |
| 38 | 09787 | 99520 | 11523 | 99334 | 13254 | 99118 | 14982 |  | 16706 |  | 2 |
| 38 | 09816 | 99517 | 11552 | 99331 | 13283 | 99114 | 15011 | 988 | 16734 |  | 22 |
| 39 | 09845 | 99514 | 11580 | 99.327 | 13312 | 99110 | 15040 | 98863 | 16763 |  | 21 |
| 40 | 09874 | 99 | 11609 | 69 | 13341 | 991 | 15 | 988 | 16792 | 98580 | 20 |
| 41 | 09903 |  | 11638 | 99 | 13370 | 99102 |  | 98854 | 16820 | 98575 | 18 |
| 42 | 09932 | 99 |  |  | 13399 | 99098 | 15126 | 988 | 16849 | 98570 | 1 |
| 43 | 09961 | 99 |  | $99^{314}$ | 13427 | 99094 | 15155 | 98845 | 16878 | 98565 | 17 |
| 44 | 09990 | 99500 | 11725 | $99^{3}$ | 13456 | 9901 | 15184 | 98841 | o6 | 98561 | 16 |
| 45 | 10 | 99497 | 11 | 99307 | 13485 | 99087 | 15212 | 98 | 16935 | 0856 | 15 |
| 46 | 10048 | 99494 |  | 99303 | 135 | 99083 | 15241 | 98832 | 16964 | 98551 | 14 |
| 47 | 10077 | 99491 | 11812 | 99300 | 13543 | 99079 | 15270 | 98827 | 16992 | 98546 | 13 |
| 48 | 10106 | 99488 | 11840 | 99297 | 13572 | 99075 | 15299 | 988.3 | 17021 | 98541 | 12 |
| 49 | 10135 | 99485 | 11869 | $99^{29} 3$ | 13600 | 99071 | 15327 | 98818 | 17050 | 98536 | 11 |
| 50 | 10 | 99482 | 11898 | $99^{290}$ | 13629 | 99067 | 15356 | 98814 | 17078 | 98531 | 10 |
| 51 | 10192 | 99479 | 11927 | 99286 | 1365 | 99063 | 15385 | 98809 | 17107 | 98526 |  |
| 52 | 10221 | 99476 | 11956 | 99283 | 13687 | 99059 | 15414 | 98805 | 17136 | 98521 |  |
| 53 | 102 | 99473 | 11985 | 99279 | 13716 | 99055 | 15442 | 98800 | 17164 | 98516 |  |
| 54 | 10279 | 99470 | 12014 | 99276 | 13744 | 99051 | 15471 | 98796 | $1719^{3}$ | 98511 |  |
| 56 | 10 | 99467 | 12043 | 99272 | 13773 | 99047 | 15500 | 987 | 17222 | 98506 |  |
| 56 | 10366 | 99464 | 12071 12100 | 99269 99265 | 13831 | 9904 |  | 98 | 17250 | 98 |  |
| 58 | 10395 | 99458 | 12120 | 99262 | 1386 | 99035 | 15586 | 98778 | 17308 | 98491 |  |
| 59 | 10 | 99455 | 12158 | 99258 | 13889 | 99031 | 15615 | $9877^{3}$ | 17336 | 98486 |  |
| 60 | 10 | 99452 | 12187 | 99255 | 13917 | 99027 | 1564 | 9876 | 17365 | 984 |  |
|  | ne. | Sine. | ne. | Sine | Cosine | Sine. | Cosine. | Sine. | Cosine | Sine |  |
|  | $84^{\circ}$ |  | $83^{\circ}$ |  | $82^{\circ}$ |  | $81^{\circ}$ |  | $80^{\circ}$ |  |  |


| 66 |  |  | NATURAL SINES AND |  |  |  | COSINES. |  | Tabie III. |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | $10^{\circ}$ |  | $11^{\circ}$ |  | $12^{\circ}$ |  | $13^{\circ}$ |  | $14^{\circ}$ |  |  |
|  | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosine. |  |
| 0 | 17365 | 98481 | 19081 | 98163 | 20791 | 97815 | 22495 | 97437 | 24192 | 97030 | 60 |
| 1 | 17393 | 98476 | 19109 | 9885 | 20820 | 97809 | 22523 | 97430 | 24220 | 97023 | 5 |
| 3 | 17422 | 98471 | 19138 | 98152 | 20848 | 97803 | 22552 | 97424 | 24249 | 97015 | 58 |
| J | 17451 | 98466 | 19167 | 98146 | 20877 | 97797 | 22580 | 97417 | 24277 | 97008 | 57 |
| 4 | 17479 | 98461 | 19195 | 98140 | 20905 | 97791 | 22608 | 97411 | 24305 | 97001 | 56 |
| 5 | 17508 | 98455 | I9224 | 98135 | 20933 | 97784 | 22637 | 97404 | 24333 | 96904 | 55 |
| 6 | 17537 | 98450 | 19252 | 98129 | 20962 | 97778 | 22665 | 97398 | 24362 | 96987 | 54 |
| 7 | 17565 | 98445 | I9281 | 98124 | 20990 | 97772 | 22693 | 9731 | 24390 | 96980 | 53 |
| 8 | 17504 | 98440 | 19309 | 98118 | 21019 | 97766 | 22722 | 97384 | 24418 | 96973 | 52 |
| 9 | 17623 | 98435 | 19338 | 98112 | 21047 | 97760 | 22750 | 97378 | 24446 | 96966 | 51 |
| 10 | 17651 | 98430 | 19366 | 98107 | 21076 | 97754 | 22778 | 97371 | 24474 | 96959 | 50 |
| 11 | 17680 | 98425 | $19^{3} 95$ | 98101 | 21104 | 97748 | 22807 | 97365 | 24503 | 96952 | 49 |
| 12 | 17708 | 98420 | 19423 | 98006 | 21132 | 97742 | 22835 | 97358 | 24531 | 96945 | 48 |
| 13 | 17737 | 98414 | 19452 | 98000 | 21161 | 97735 | 22863 | 97351 | 24559 | 96937 | 47 |
| 14 | 17766 | 98409 | 19481 | 98084 | 21189 | 97729 | 22892 | 97345 | 24587 | 96930 | 46 |
| 15 | 17794 | 98404 | 19509 | 98079 | 21218 | 97723 | 22920 | 97338 | 24615 | 96923 | 45 |
| 16 | 17823 | 98399 | 19538 | 98073 | 21246 | 97717 | 22948 | 97331 | 24644 | 96916 | 44 |
| 17 | 17852 | 98394 | 19566 | 98067 | 21275 | 97711 | 22977 | 97325 | 24672 | 96909 | 43 |
| 18 | 17880 | 98389 | 19.505 | 98061 | 21303 | 97705 | 23005 | 97318 | 24700 | 96902 | 42 |
| 19 | 17909 | 98383 | 19623 | 98056 | 21331 | 97698 | 23033 | 97311 | 24728 | 96804 | 41 |
| 20 | 17937 | 98378 | 19652 | 98050 | 21360 | $976{ }^{2}$ | 23062 | 97304. | 24756 | 96887 | 40 |
| 21 | 17966 | 98373 | 19680 | 98044 | 21388 | 97686 | 23090 | 97298 | 24784 | 96880 | 39 |
| 22 | 17995 | 98368 | 19709 | 98039 | 21417 | 97680 | 23118 | 97291 | 24813 | 96873 | 38 |
| 23 | 18023 | 98362 | 19737 | 98033 | 21445 | 97673 | 23146 | 97284 | 24841 | 96866 | 37 |
| 24 | 18052 | 98357 | 19766 | 98027 | 21474 | 97667 | 23175 | 97278 | 24869 | 96858 | 36 |
| 25 | 18081 | 98352 | 19794 | 98021 | 21502 | 97661 | 23203 | 97271 | 24897 | 96851 | 35 |
| 26 | 18100 | 98347 | 19823 | 98016 | 21530 | 97655 | 23231 | 97264 | 24925 | 96844 | 34 |
| 27 | 18138 | 98341 | 19851 | 98010 | 21559 | 97648 | 23260 | 97257 | 24954 | 96837 | 33 |
| 28 | 18166 | 98336 | 19880 | 98004 | 21587 | 97642 | 23288 | 97251 | 24982 | 96829 | 32 |
| 29 | 18195 | 98331 | 19908 | 97998 | 21616 | 97636 | 23316 | 97244 | 25010 | 96822 | 31 |
| 3.5 | 18224 | 98325 | 19937 | $9799^{2}$ | 21644 | 97630 | 23345 | 97237 | 25038 | 96815 | 3 c |
| 31 | 18252 | 98320 | 19965 | 97987 | 21672 | 97623 | 23373 | 97230 | 25066 | 96807 | 29 |
| 32 | 18281 | 98315 | 19994 | 97981 | 21701 | 97617 | 23401 | 97223 | 25094 | 96800 | 28 |
| 33 | 18309 | 98310 | 20022 | 97975 | 21729 | 97611 | 23429 | 97217 | 25122 | $9679^{3}$ | 27 |
| 34 | 18338 | 98304 | 20051 | 97969 | 21758 | 97604 | 23458 | 97210 | 25151 | 96786 | 26 |
| 35 | 18367 | 98299 | 20079 | 97963 | 21786 | 97598 | 23486 | 97203 | 25179 | 96778 | 25 |
| 36 | 18395 | $9^{82} 24$ | 20108 | 97958 | 21814 | 97502 | 23514 | $9710^{6}$ | 25207 | 96771 | 24 |
| 37 | 18424 | 98288 | 20136 | 97952 | 21843 | 97585 | 23542 | 97189 | 25235 | 96764 | 23 |
| 38 | 18452 | 98283 | 20165 | 97946 | 21871 | 97579 | 23571 | 97182 | 25263 | 96756 | 22 |
| 39 | 18481 | 98277 | 20193 | 97940 | 21899 | 97573 | 23599 | 97176 | 25291 | 96749 | 21 |
| 40 | 18509 | 98272 | 20222 | 97934 | 21928 | 97566 | 23627 | 97169 | 25320 | 96742 | 20 |
| 41 | 18538 | 98267 | 20250 | 97928 | 21956 | 97560 | 23656 | 97162 | 25348 | 96734 | 19 |
| 42 | 18567 | 98261 | 20279 | 97922 | 21985 | 97553 | 23684 | 97155 | 25376 | 96727 | 18 |
| 43 | 18595 | 98255 | 20307 | 97916 | 22013 | 97547 | 23712 | 97148 | 25404 | 96719 | 17 |
| 44 | 18624 | 98250 | 20336 | 97910 | 22041 | 97511 | 23740 | 97141 | 25432 | 96712 | 16 |
| 45 | 18652 | 98245 | 20364 | 9790' | 22070 | 97534 | 23769 | 9713 | 25460 | 96705 | 15 |
| 46 | 1868ı | 98240 | 20393 | $97^{89} 9$ | 22098 | 97528 | 23797 | 97127 | 25488 | 96697 | 14 |
| 47 | 18710 | 98234 | 20421 | $978{ }^{3}$ | 22126 | 97521 | 23825 | 97120 | 25516 | 96690 | 13 |
| 48 | 18738 | 98229 | 20450 | 97887 | 22155 | 97515 | 23853 | 97113 | 25545 | 96682 | 12 |
| 49 | 18767 | 98223 | 20478 | 97881 | 22183 | 97508 | 23882 | 97106 | 25573 | 96675 | 11 |
| 50 | 18795 | 98218 | 20507 | 97875 | 22212 | 97502 | 23910 | 97100 | 25601 | 96667 | 0 |
| 51 | 18824 | 98212 | 20535 | 97869 | 22240 | 97496 | 23938 | 97093 | 25629 | 96660 | 9 |
| 52 | 18852 | $9^{8207}$ | 20563 | 97863 | 22268 | 97489 | 23966 | 97086 | 25657 | 96653 | 8 |
| 53 54 | 18881 | $9^{8201}$ | 20592 | 97857 | 22297 | 97483 | 23995 | 97079 | 25685 | 96645 | 7 |
| 54 55 | 18910 | $9^{81} 196$ | 20620 | 97851 | 22325 | 97476 | 24023 | 97072 | 25713 | 96638 | 5 |
| 56 | 18938 18967 | 98180 | 206 | 97843 | 22353 2238 | 974 | 24051 | 97065 97058 | 25741 25760 | 96630 | 4 |
| 57 | 18995 | $9^{817}$ | 20706 | 97833 | 22410 | 97457 | 24108 | 97051 | 25798 | 96615 | 3 |
| 58 | 19024 | $9^{8174}$ | 20734 | 97827 | 22438 | 97450 | 24136 | 97044 | 25826 | 96605 | 8 |
| 59 | 19052 | 98168 | 20763 | 97821 | 22467 | 97444 | 24164 | 97037 | 25854 | 96600 | 1 |
| 60 | 19081 | 98163 | 20791 | 97815 | $2249^{5}$ | 97437 | 24192 | 97030 | 25882 | 9657.3 | - |
| , | Cosine. | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosine. | Sine. |  |
|  | $79^{\circ}$ |  | $78^{\circ}$ |  | $77^{\circ}$ |  | $76^{\circ}$ |  | $75^{\circ}$ |  |  |


| T'able III. |  |  | NATURAL |  | SINE |  | COSINES. |  |  |  | 67 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $15^{\circ}$ |  | $16^{\circ}$ |  | $17^{\circ}$ |  | $18^{\circ}$ |  | $19^{\circ}$ |  |  |
|  | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosine | Sine. | Cosi | Sine | Cosi |  |
| 0 | 25882 |  | 27564 |  |  | 95630 | 30902 | 95106 | 32557 | 94552 | 60 |
| 1 | 25910 | 96585 | 27592 | 96118 | 29265 | 95622 | 30929 | 95097 | 32584 | 94542 | 59 |
|  | 25938 | 96578 | 27620 | 96110 | 29293 | 95613 | 30957 | 95088 | 32612 | 9.4533 | 58 |
| 2 | 25966 | 96570 | 27648 | 96102 | 29.321 | 95605 | 30985 | 95079 | 3:639 | 94523 | 57 |
| 4 | $\left\|\begin{array}{l} 25994 \\ 26022 \end{array}\right\|$ | 96562 | 27676 | 96004 | 29348 | 95596 | 31012 | 95070 | 32667 | 94514 | 56 |
|  |  | 96555 | 27704 | 96086 | 29376 | 95588 | 31040 | 9 9061 | 32694 | 94504 | 55 |
| 6 | 26050 | 96547 | 27731 | 96078 | 29404 | 95579 | 31068 | 95052 | 32722 | 94.95 | 54 |
| 7 | 2607926107 | 96540 | 27759 | 96070 | 29432 | 95571 | 31095 | 95043 | 32749 | 94485 | 53 |
|  |  | 96532 | 27787 | 96062 | 29460 | 95562 | 3123 | 95033 | 32777 | 94476 | 52 |
| 9 | 26107 26135 | 96524 | 27815 | 96054 | 29487 | 95554 | 31151 | 95024 | 32804 | 94466 | 5 I |
| 10 | 26163 | 96517 | 27843 | 96046 | 29515 | 95545 | 31178 | 95015 | 32832 | 94457 | 50 |
| 11 | 26191 | 96509 | 27871 | 96037 | 29543 | 95536 | 31206 | 95006 | 32859 | 94447 | 49 |
| 12 | 26219 | 96502 | 27899 | 96029 | 29571 | 95528 | 31233 | 94997 | 32887 | 94438 | 48 |
| 13 | 262472627526303 | 96404 | 27927 | 96021 | 29599 | 95519 | 31261 | 94988 | 32914 | 94428 | 47 |
| 14 |  | 96486 | 27955 | 96013 | 29626 | 95511 | 31289 | 94979 | 32942 | 94418 | 46 |
| 15 |  | 96479 | 27983 | 96005 | 29654 | 95502 | 31316 | 94970 | 32969 | 94409 | 45 |
| 16 | 2633I | 964 | 2801 I |  | 29682 | 9 | 31344 |  |  |  | 44 |
| 17 | 26359 | 96463 | 28039 | 95989 | 29710 | 95485 | 31372 | 94952 | 33024 | 94390 | 43 |
| 18 | $\begin{aligned} & 26387 \\ & 26415 \end{aligned}$ | 96456 | 28067 | 95981 | 29737 | 95476 | 31399 | 94943 | 3305ı | 94380 | 42 |
| 19 |  | 96448 | 28095 | 95972 | 29765 | 95467 | 31427 | 94933 | 33079 | 94370 | 41 |
| 20 | $\begin{aligned} & 26415 \\ & 26443 \end{aligned}$ | 96440 | 28123 | 95964 | 29793 | 95459 | 31454 | 94924 | 33106 | 94361 | 40 |
| 21 | $\begin{aligned} & 26443 \\ & 26471 \end{aligned}$ | 96433 | 28150 | 55956 | 29821 | 95450 | 31482 | 94915 | 33134 | 94351 | 39 |
| 22 | 2650026528 | 96425 | 28178 | 95948 | 29849 | 95441 | 31510 | 94906 | 33161 | 94342 | 38 |
| 23 |  | 96417 | 28206 | 95940 | 29876 | 95433 | 31537 | 94897 | 33189 | 94332 | 37 |
| 24 | 26528 | 96410 | 28234 | 95931 | 29904 | 95424 | 31565 | 94888 | 33216 | 94322 | 36 |
| 25 | 26584 | 96402 | 28262 | 95923 | 29932 | 95415 | 31593 | 94878 | 33244 | 94313 | 35 |
| 26 | 26612 | 96394 | 28290 | 95915 | 29960 | 95407 | 31620 | 94869 | 33271 | 94303 | 34 |
| 27 | $\begin{array}{r} 26640 \\ 26668 \end{array}$ | 96386 | 28318 | 95907 | 29987 | 95398 | 31648 | 94860 | 33298 | 94293 | 33 |
| 28 |  | 96379 | 28346 | 95898 | 30015 | 95389 | 31675 | 94851 | 33326 | 94284 | 32 |
| 29 | $\begin{aligned} & 26696 \\ & 26724 \end{aligned}$ | 96371 | 28374 | 95890 | 30043 | 95380 | 31703 | 94842 | 33353 | 94274 | 31 |
| 30 |  | 96363 | 28402 | 95882 | 30071 | 95372 | 3ı730 | 94832 | 33381 | 04264 | 30 |
| 31 | $\begin{aligned} & 26724 \\ & 2772 \end{aligned}$ | 96355 | 28429 | 95874 | 30098 | 95363 | 31758 | 94823 | 33408 | 94254 | 29 |
| 32 | 26780 | 96347 | 28457 | 95865 | 30126 | 95354 | 3ı786 | 94814 | 33436 | 94245 | 28 |
| 33 | 26808 | 96340 | 28485 | 95857 | 30154 | 95345 | 31813 | 94805 | 33463 | 94235 | 27 |
| 34 | $\left.\begin{array}{\|l\|} 26836 \\ 26864 \end{array} \right\rvert\,$ | 96332 | 285. 3 | 95849 | 30182 | 95337 | 31841 | 94795 | 33490 | 94225 | 26 |
| 35 |  | $9^{63324}$ | 2854 I | 95841 | 30209 | 95328 | 31868 | 94786 | 33518 | 94215 | 25 |
| 36 | 26892 | 96316 | 28569 | 95832 | 30237 | 95319 | 31896 | 94777 | 33545 | 94206 | 24 |
| 37 | $\begin{aligned} & 26920 \\ & 26948 \end{aligned}$ | 96308 | 28597 | 95824 | 30265 | 95310 | 31923 | 94768 | 33573 | 94106 | 23 |
| 38 |  | 96301 | 28625 | 95816 | 30292 | 95301 | 3195ı | 94758 | 33600 | 94186 | 22 |
| 39 | $\begin{aligned} & 26948 \\ & 26976 \end{aligned}$ | 96293 | 28652 | 95807 | 30320 | $9529^{3}$ | 31979 | 94749 | 33627 | 94176 | 21 |
| 40 | 27004 | $9^{6285}$ | 28680 | 95799 | 30348 | 95284 | 32006 | 94740 | 33655 | 94167 | 20 |
| 41 | $\begin{aligned} & 27032 \\ & 27060 \end{aligned}$ | 96277 | 28708 | 95791 | 30376 | 95275 | 32034 | 94730 | 33682 | 94157 | 19 |
| 42 |  | 96269 | 28736 | 95 | 30403 | 95266 | 3206I | 94721 | 33710 | 94147 | 18 |
| 43 | $\begin{array}{r} 27060 \\ 27088 \end{array}$ | 96261 | 28764 | 95774 | 30431 | 95257 | 32089 | 94712 | 33737 | 94137 | 17 |
| 44 | $\begin{aligned} & 27116 \\ & 27144 \end{aligned}$ | 96253 | 28792 | 95766 | 30459 | 95248 | 32116 | 94702 | 33764 | 94127 | 16 |
| 45 |  | 96246 | 288 | 9 | 30 | 95240 | 32144 | 94693 | 33792 | 9¢118 | 15 |
| 46 | 27172272002722 | 96238 | 28847 | 95749 | 30514 | 9523 I | 32171 | 94684 | 33819 | 94108 | 14 |
| 47 |  | 96230 | 28875 | 95740 | 30542 | 95222 | 32199 | 94674 | 33846 | 94098 | 13 |
| 43 | 27228 | 96222 | 28903 | 95732 | 30570 | 952 I 3 | 32227 | 94665 | 33874 | 94088 | 12 |
| 49 | $\begin{aligned} & 27256 \\ & 27284 \end{aligned}$ | $9^{6214}$ | 28931 | 95724 | 30597 | 95204 | 32254 | 94656 | 33901 | 94078 | 11 |
| 50 |  | $9^{6206}$ | 28959 | 95715 | 30625 | 95125 | 32282 | 94646 | 33929 | 94068 | 10 |
| 51 | $\begin{aligned} & 2784 \\ & 2712 \\ & 2712 \end{aligned}$ | 96ı98 | 28987 | 95707 | 30653 | 95186 | 32309 | 94637 | 33956 | 94058 | 9 |
| 52 | 27340 27368 | 96190 | 29015 | 95698 | 30680 | 95177 | 32337 | 94627 | 33983 | 94049 | 8 |
| 53 | $\begin{aligned} & 27368 \\ & 27396 \end{aligned}$ | 96182 | 29042 | 95600 | 30708 | 95168 | 32364 | 94618 | 34011 | 94039 | 7 |
| 54 |  | 96174 | 29070 | 95681 | 30736 | 93159 | 32392 | 94609 | 34038 | 94029 | 6 |
| 55 | 27424 | 96166 | 29098 | 95673 | 30763 | 95150 | 32419 | 94599 | 34065 | 94019 | 5 |
| 56 | 27452 | 96158 | 29126 | 95664 | 30791 | 95142 | 32447 | 94590 | 34093 | 94009 | 4 |
| 57 | 2748027508 | 96150 | 29154 | 95056 | 30819 | 95133 | 32474 | 94580 | 34120 | 93999 | 3 |
| 58 |  | 96142 | 29182 | 95647 | 30846 | 95124 | 32502 | 94571 | 34147 | 93984 | 2 |
| 59 | $\begin{aligned} & 27536 \\ & 27564 \end{aligned}$ | 96134 | 29209 | 95639 | 30874 | 95115 | 32529 | 94561 | 34175 | 93979 | 1 |
| 60 |  | 96126 | 29237 | 95630 | 30902 | 95106 | 32557 | 94552 | 34202 | 93969 | 0 |
| , | Cosine. | Sine | Cosine. | Sire. | Cosine. | Sine. | Cosine. | Sine. | Cosine. | Sine. |  |
|  |  |  |  |  |  |  |  |  |  |  |  |


| 68 |  |  | NATURAL SINES AND COSINES. |  |  |  |  |  | Table III. |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $20^{\circ}$ |  | $21^{\circ}$ |  | $22^{\circ}$ |  | $23^{\circ}$ |  | $24^{\circ}$ |  |  |
|  | Sine. | Cosine. | Sine. | Cusine. | Sine. | Cosine | ine. | Cosine | Sine. | Cosine. |  |
| 0 | 34 | 939 | 35837 | 93358 | 37 | 92718 | 39073 | 92050 | 74 | 91355 |  |
| 1 | 34229 | 93959 | 35864 | 93348 | 37488 | 92707 | 39100 | 92039 | 40700 | 91343 | 59 |
| 2 | 34257 | 93949 | 35891 | 93337 | 37515 | 92697 | 39127 | 92028 | 40727 | 91331 | 58 |
| 3 | 34284 | 93939 | 35918 | 93327 | 37542 | 92686 | 39153 | 92016 | 40753 | 91319 | 57 |
| 4 | 34311 | 93929 | 35945 | 93316 | 37569 | 92675 | 39180 | 92005 | 40780 | 91307 |  |
| 5 | 34339 | 93919 | 35973 | 93306 | 37595 | 92664 | 39207 | 91994 | 40806 | 91295 | 55 |
| 6 | 34366 | 93909 | 36000 | $9^{32} 25^{-}$ | 37622 | 92653 | 39234 | 91982 | 40833 | 91283 | 54 |
| 7 | 34393 | 93899 | 36027 | 93285 | 37649 | 92642 | 39260 | 91971 | 40860 | 91272 | 5 |
| 8 | 34421 | 93889 | 36054 | 93274 | 37676 | 92631 | 39287 | 91959 | 40886 | 91260 | 52 |
| 9 | 34448 | 93879 | 36081 | 93264 | 37703 | 92620 | 39314 | 91948 | 40913 | 91248 | 51 |
| 10 | 34475 | 93869 | 36108 | 93253 | 37730 | 92609 | 39341 | 91936 | 40939 | 91236 | 50 |
| 11 | 34503 | 93859 | 36135 | 93243 | 37757 | 92598 | 39367 | 91925 | 40966 | 91224 | 8 |
| 12 | 34530 | 93849 | 36162 | $9^{3232}$ | 37784 | 92587 | 39394 | 91914 | 40992 | 91212 | 48 |
| 13 | 34557 | 93839 | 36190 | 93222 | 37811 | 92576 | 39421 | 91902 | 41019 | 91200 | 47 |
| 14 | 34584 | 93829 | 36217 | 93211 | 37838 | 92565 | 39448 | 91891 | 41045 | 91188 | 46 |
| 15 | 34612 | 38 | 36244 |  |  |  |  | 91879 | 41072 | 91176 | 45 |
| 16 | 34639 | 93809 | 36271 | $931{ }^{\circ} \mathrm{c}$ | 37892 | 92543 | 39501 | 91868 | 41098 | 91164 | 44 |
| 17 | 34666 | 93799 | 36298 | 93180 | 37919 | 92532 | 39528 | 91856 | 41125 | 91152 | 43 |
| 18 | 34694 | 93789 | 36325 | $9^{3169}$ | 37946 | 92521 | 39555 | 91845 | 41151 | 91140 | 42 |
| 19 | 34721 | 93779 | 36352 | 93159 | 37973 | 92510 | 39581 | 91833 | 41178 | 91128 | 41 |
| 20 | 34748 | 93769 | 36379 | 93148 | 37999 | 92499 | 3,608 | 91822 | 41204 | 91116 | 40 |
| 21 | 34775 | 93750 | 36406 | 93137 | 38026 | 92488 | 39635 | 91810 | 4123 I | 91104 | 3 |
| 22 | 34803 | 93748 | 36434 | 93127 | 38053 | 92477 | 3,665 | 91799 | 41257 | 91092 | 38 |
| 23 | 34830 |  | 36461 | 93116 | 38080 | 92466 |  | 9178 | 41284 | 91080 |  |
| 24 | 34857 | 93728 | 36488 | 93106 | 38107 | 92450 | 39715 | 91775 | 41310 | 91068 |  |
| 25 | 34884 | 93718 | 36515 | 93095 | 38134 | 92444 | 39741 | 91764 | 41337 | 91056 | 3 |
| 26 | 34912 |  | 36542 | 93084 | - 38161 | 92432 |  | 91752 | 41363 | 91044 | 3 |
| 27 | 3.4939 | 936 | 36569 | 93074 | 38188 | 92421 |  | 91741 | 41390 | 91032 | 3 |
| 28 |  | 93688 | 36596 | 93063 | 38215 | 92410 | 39822 | 91729 | 41416 | 91020 | 32 |
| 29 | 34993 | 93677 | 36623 | 93052 | 38241 |  | 3984 | 91718 | 41443 |  | 3 |
| 30 |  |  |  | 93642 | 38268 | 92 |  | 91 | 41 | 90996 |  |
| 31 | 35048 | 93657 | 36677 | 9303 I | 38295 | 92 | 39902 | 91694 | 41496 | 90984 |  |
| 32 | 35075 | 93647 | 36704 | 93020 | 38322 | 92366 | 39928 | 91683 | 41522 | 90972 | 28 |
| 33 | 35102 | 93637 | 36731 | 93010 | 38349 | 92355 | 39955 | 91671 | 41549 | 90060 | 27 |
| 34 | 35130 | 93626 | 36758 | 92999 | 38376 | 92343 | 39982 | 91660 | 41575 | 90948 |  |
| 35 | 35157 | 93616 | 36785 | 92938 | 38403 | 92332 | 40008 | 91648 | 41602 | 90936 | 25 |
| 36 | 35184 | 936 | 36812 | 92978 | 38430 | 92321 | 40035 | 91636 | 41628 | 90924 | 24 |
| 37 | 35211 | 935 | 36839 | 92967 | 38450 | 92310 | 40062 | 91625 | 41655 | 90911 | 23 |
| 38 | 35 | 935 | 36867 | 92956 | 38483 | 92299 | 40088 | 91613 | 41681 | 90899 | 22 |
| 39 | 35266 | 93575 | 36894 | 92945 | 38510 | 9228 | 40115 | 91601 | 41707 | 90 | 21 |
| 40 | 35293 | 93565 | 36921 | 92935 | 38537 | 92276 | 40141 | 915 | 41734 | 83 | 20 |
| 41 | 35320 | 93355 | 36948 | 92924 | 38564 | 92265 | 40168 | 91578 | 41760 | 90863 |  |
| 42 | 35 | 93544 | 36975 | 92913 | 38501 | 92254 | 40195 | 91506 | 41787 | 90851 |  |
| 43 |  | 93534 | 37002 | 92902 | 38617 | 92243 | 4022 | 91555 | 41813 | 90839 |  |
| 44 | 35402 | 93524 | 37029 | 92892 | 38644 | 922 | 40248 | 91543 | 41840 | 826 | 16 |
| 45 | 35 | 93 | 3 | 9 | 38671 | 92.220 | 40275 | - | 41866 |  |  |
| 46 | 35456 | 93503 | 37083 | 92870 | 38698 | 92200 | 40301 | 915 | 41892 | 90802 | 14 |
| 47 | 35484 | $9349^{3}$ | 37110 | 92859 | 38725 | 92108 | 4032 | 91508 | 41919 | 90790 | 13 |
| 48 | 35511 | 93483 | 37137 | 92849 | 38752 | 92186 | 40355 | 91496 | 41945 | 90778 | 12 |
| 49 | 35538 | 93472 | 37164 | 92838 | 38778 | 92175 | 40381 | 91484 | 41972 | 90766 | 11 |
| 50 | 35565 | 93462 | 37191 | 92827 | 38805 | 92164 | 40408 | 91472 | 41998 | 90753 | 10 |
| 51 | 35592 | 93452 | 37218 | 92816 | 38832 | 92152 | 40434 | 91461 | 42024 | 90741 |  |
| 52 | 3561 | 93441 | 37245 | 92805 | 38859 | 92141 | 40461 | 91444 | 420 | 90729 |  |
| 53 | 356 | 93431 | 37272 | 92794 | 38886 | 92130 | 40,488 | 91437 | 42077 | 90717 |  |
| 5 | 35674 | 93420 | 37 | 92784 | 38912 | 92119 | 40514 | 91425 | 42104 | 90704 |  |
| 55 |  | 93410 | 37326 | 92773 | 38939 | 92107 | 40541 | 91414 | 42130 | 90692 |  |
| 56 | 3 | 93400 | 37353 | 92762 | 389006 | 92006 | 40567 | 91402 | 42156 | 90650 |  |
| 5 |  | 93 | 37380 | 92751 | 38993 | 92085 | 40594 | 913 | 4218 | 00668 |  |
| 5 | 358 | 93379 93368 |  | 92740 |  | 920 | 406 | 9 |  | 0063 |  |
| 60 | 35 | 93358 | 37461 | 92718 | 39073 | 9:050 | 40674 | 91355 | 42262 | 90631 |  |
|  | Cosine. | Sire. | Co | Sin | C | Sine. |  | Sine. | Cosine. | Sin |  |
|  | $69^{\circ}$ |  | $68^{\circ}$ |  | $67^{\circ}$ |  | $66^{\circ}$ |  | $65^{\circ}$ |  |  |


| Table III. |  |  | NATURAL SINES AN |  |  |  | COSINES. |  | 69 |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | $25^{\circ}$ |  | $26^{\circ}$ |  | $27^{\circ}$ |  | $28^{\circ}$ |  | $23^{\circ}$ |  | , |
|  | Sine. | Cosine. | Sine. | Cosine | Sine. | Cosine. | Sine. | Cosine, | Sine. | Cosine. |  |
| 0 | 42262 | 90631 | 43 | 89 | 45399 | 89101 | 46947 |  | 48481 | 87462 | 60 |
| 1 | 42288 | 90618 | 43863 | 89867 | 45425 | 89087 | $46 \mathrm{c} 7^{3}$ | 88281 | 48506 | 87448 | 59 |
| 2 | 42315 | 90606 | 43889 | 89854 | 45451 | 89074 | 46999 | 88267 | 48532 | 87434 | 58 |
| 3 | 42341 | 90594 | 43916 | 89841 | 45477 | 89061 | 47024 | 88254 | 48557 | 87420 | 57 |
| 4 | 42367 | 90582 | 43942 | 89828 | 45503 | 89048 | 47050 | 88240 | 48583 | 87406 | 56 |
| 5 | 42394 | 90569 | 43968 | 89816 | 45529 | 89035 | 47076 | 88226 | 48608 | 87391 | 55 |
| 6 | 42420 | 90557 | 43994 | 89803 | 45554 | 89021 | 47101 | 88213 | 48634 | $8 \% 377$ | 54 |
| 7 | 42446 | 90545 | 44020 | 89790 | 45580 | 89008 | 47127 | 88199 | 48659 | 87363 | 53 |
| 8 | 42473 | 90532 | 44046 | 89777 | 45606 | 88995 | 47153 | 88.85 | 48684 | 87349 | 52 |
| 9 | 42499 | 90520 | 44072 | 89764 | 45632 | 88981 | 47178 | 88ı72 | 48710 | 87335 | 51 |
| 10 | 42525 | 90507 | 44098 | 89752 | 45658 | 88968 | 47204 | 88.58 | 48735 | 87321 | 50 |
| 11 | 42552 | 90495 | 44124 | $807^{3} 9$ | 45684 | 88955 | 47229 | 88144 | 48761 | 87306 | 49 |
| 12 | 42578 | 90483 | 4415: | 89726 | 45710 | 88942 | 47255 | 88130 | 48786 | 87292 | 48 |
| 13 | 42604 | 90470 | 44177 | 89713 | 45736 | 88928 | 47281 | $8_{8117}$ | 48811 | 87278 | 47 |
| 14 | 42631 | 90458 | 44203 | 89700 | 45762 | 88915 | 47306 | 88103 | 48837 | 87264 | 46 |
| 15 | 42657 | 90446 | 44229 | 89687 | 45787 | 88902 | 47332 | 88089 | 48862 | 87250 | 45 |
| 16 | 42683 | 90433 | 44255 | 89674 | 45813 | 88888 | 47358 | 88075 | 48888 | 87235 | 44 |
| 17 | 42709 | 90421 | 44281 | 89662 | 45839 | 88875 | 47383 | 88062 | 48913 | 87221 | 43 |
| 18 | 42736 | 90408 | 44307 | 89649 | 45865 | 88862 | 47409 | 88048 | 48938 | 87207 | 42 |
| 19) | 42762 | 90396 | 44333 | 89636 | 45891 | 88848 | 47434 | 88034 | 48964 | $871{ }^{3}$ | 41 |
| 20 | 42788 | 90383 | 44359 | 89623 | 45917 | 88835 | 47460 | 88020 | 48989 | 87178 | 40 |
| 21 | 42815 | 90371 | 44385 | 89610 | 45942 | 88822 | 47486 | 88006 | 49014 | 87164 | 39 |
| 22 | 42841 | 90358 | 44411 | 89507 | 45968 | 88808 | 47511 | 87993 | 49040 | 87150 | 38 |
| 23 | 42867 | 90346 | 44437 | 89584 | 45994 | 88795 | 47537 | 87979 | 49065 | 87136 | 37 |
| 24 | 42894 | 90334 | 44464 | 8951 | 46020 | 88782 | 47562 | 87965 | 49090 | 87121 | 36 |
| 25 | 42920 | 90321 | 44490 | 89558 | 46046 | 88768 8875 | 47588 | 87951 | 49116 | 87107 | 35 34 3 |
| 26 | 42946 42972 | 90309 90296 | 44516 | 89545 89532 | 46072 | 88755 88741 | 47614 47639 | 87937 87923 | 49141 49166 | 87093 87079 | 34 |
| 27 28 | 42972 42999 | 90296 | 44542 | 89532 89519 | 46097 | 88741 88728 | 47639 47665 | 87923 87909 | 49166 | 87079 87064 | 33 32 |
| 29 | 43025 | 90271 | 44594 | 89506 | 46149 | 88715 | 47690 | 87896 | 49217 | 87050 | 31 |
| 30 | 43051 | 90259 | 44620 | 89493 | 46175 | 88701 | 47716 | 87882 | 49242 | 87036 | 30 |
| 31 | 43077 | 90246 | 44646 | 89480 | 46201 | 88688 | 47741 | 87868 | 49268 | 87021 | 29 |
| 32 | 43104 | 90233 | 44672 | 89467 | 46226 | 88674 | 47767 | 87854 | 49293 | 87007 | 28 |
| 33 | 43130 | 90221 | 44698 | 89454 | 46252 | 88661 | 47793 | 87840 | 49318 | 86993 | 27 |
| 34 | 43156 | 90208 | 44724 | 89441 | 46278 | 88647 | 47818 | 87826 | 49344 | 86978 | 20 |
| 35 | 43182 | 90186 | 44750 | 89428 | 46304 | 88634 | 47844 | 87812 | 49369 | 86964 | 25 |
| 36 | 43209 | 90183 | 44776 | 89415 | 46330 | 88620 | 47869 | 87798 | 49394 | 86949 | 24 |
| 37 | 43235 | 90171 | 44802 | 89402 | 46355 | 88607 | 47895 | 87784 | 49419 | 86935 | 23 |
| 38 | 43261 | 90158 | 44828 | 89389 | 46381 | $8859^{3}$ | 47920 | 87770 | 49445 | 86921 | 22 |
| 39 | 43287 | 90146 | 44854 | 89376 | 46407 | 88580 | 47946 | 87756 | 49470 | 86906 | 21 |
| 40 | 43313 | 90133 | 44880 | 89363 | 46433 | 88566 | 47971 | 87743 | 49495 | 86892 | 20 |
| 41 | 43340 | 90120 | 44906 | 89350 | 46458 | 88553 | 47997 | 87729 | 49521 | 86878 | 19 |
| 42 | 43366 | 90108 | 44932 | 8 8 337 | 46484 | 88539 | 48022 | 87715 | 49546 | 86863 | 18 |
| 43 | 43392 | 90095 | 44958 | 89324 | 46510 | 88526 | 48048 | 87701 | 49571 | 86849 | 17 |
| 44 | 43418 | 90083 | 44984 | 89311 | 46536 | 88512 | 48073 | 87687 | 49596 | 86834 | 16 |
| 45 | 43445 | 90070 | 45010 | 89298 | 46561 | 88499 | 48099 | 87673 | 49622 | 86820 | 15 |
| 46 | 43471 | 90057 | 45036 | 89285 | 46587 | 88485 | 48124 | 87659 | 49647 | 86805 | 14 |
| 47 | 43497 | 90045 | 45062 | 89272 | 46613 | 88472 | 48150 | 87645 | 49672 | 86791 | 13 |
| 48 | 43523 | 90032 | 45088 | 89259 | 46639 | 88458 | 48175 | 87631 | 49697 | 86777 | 12 |
| 49 | 43549 | 90019 | 45114 | 89245 | 46664 | 88445 | 48201 | 87617 | 49723 | 86762 | 1 |
| 50 | 43575 | 80007 | 45140 | 89232 | 46690 | 88431 | 48226 | 87603 | 49748 | 86748 | 0 |
| 52 | 43628 |  |  | 89219 89206 | 46716 | 88417 88404 | 4825 | 8758 | 49773 | 86733 | 8 |
| 53 | 43654 | 89968 | 45218 | 89193 | 46767 | 88390 | 48303 | 87561 | 49824 | 86704 | 7 |
| 54 | 43680 | 89956 | 45243 | 89180 | 46793 | 88377 | 48328 | 87546 | 49849 | 86690 | 6 |
| 55 | 43706 | 89943 | 45269 | 89167 | 46819 | 88363 | 48354 | 87532 | 49874 | 86675 | 5 |
| 56 | 43733 | 89930 | 45295 | 89153 | 46844 | 88349 | 48379 | 87518 | 49899 | 86661 | 4 |
| 56 58 | 43759 | 89918 | 45321 | 89140 | 46870 | 88.36 | 48405 | 87504 | 49924 | 86646 | 3 |
| 58 | 43785 | 89905 | 45347 | 89127 | 46896 | 88322 | 48430 | 87490 | 49950 | 86632 | 2 |
| 59 | 43811 | 89892 | 45373 | 89114 | 46921 | 88308 | 48456 | 87476 | 49975 | 86617 | 1 |
| 60 | 43837 | 8,879 | 45399 | 89101 | 46947 | 88295 | 48481 | 87462 | 50000 | 86603 | - |
|  | Cosine. | Sine. | Cosire. | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosi | Sine. |  |
|  | $64^{\circ}$ |  | $63^{\circ}$ |  | $62^{\circ}$ |  | $61^{\circ}$ |  | $60^{\circ}$ |  | , |


| 70 |  |  | NATURAL SINES AND COSINES. |  |  |  |  |  | Table III. |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | $30^{\circ}$ |  | $31^{\circ}$ |  | $32^{\circ}$ |  | $33^{\circ}$ |  | $34^{\circ}$ |  | , |
|  | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosine |  |
| 0 | 50000 | 86603 | 51504 | 85717 | 52992 | 84805 | 54464 | 83867 | 55919 | 82904 | 60 |
| 1 | 50025 | 86588 | 51529 | 85702 | 53017 | 84789 | 54488 | 8385i | 55943 | 82887 | 59 |
| 2 | 50050 | 86573 | 51554 | 85687 | 53041 | 84774 | 54513 | 83835 | 55968 | 82871 | 58 |
| 3 | 50076 | 86559 | 51579 | 85672 | 53066 | 84759 | 54537 | 83819 | 55992 | 82855 | 57 |
| 4 | 50101 | 86544 | 51604 | 85657 | 53091 | 84743 | 54561 | 83804 | 56016 | 82839 | 56 |
| 5 | 50126 | 86530 | 51628 | 85642 | 53115 | 84728 | 54586 | 83788 | 56040 | 82822 | 55 |
| 6 | 50151 | 86515 | 51653 | 85627 | 53140 | 84712 | 54610 | 83772 | 56064 | 82806 | 54 |
| 7 | 50176 | 86Јัı | 51678 | 85012 | 53164 | 84697 | 54635 | 83756 | 56088 | 82790 | 53 |
| 8 | 50201 | 86486 | 51703 | 85597 | 53189 | 84681 | 54659 | 83740 | 56112 | $8277^{3}$ | 52 |
| 9 | 50227 | 86471 | 51728 | 85582 | 53214 | 84666 | 54683 | 83724 | 56136 | 82757 | 51 |
| $i \mathrm{C}$ | 50252 | 86457 | 51753 | 85567 | 53238 | 84650 | 54708 | 83708 | 56160 | 82741 | 50 |
| 11 | 50277 | 86442 | 51778 | 85551 | 53263 | 84635 | 54732 | 83692 | 56184 | 82724 | 49 |
| 12 | 50302 | 86427 | 51803 | 85536 | 53288 | 84619 | 54756 | 83676 | 56208 | 82708 | 48 |
| 13 | 50327 | 86413 | 51828 | 85521 | 53312 | 84604 | 54781 | 83660 | 56232 | 82692 | 47 |
| 14 | 50352 | 86398 | 51852 | 85506 | 53337 | 84588 | 54805 | 83645 | 56256 | 82675 | 46 |
| 15 | 50377 | 86384 | 51877 | 85491 | 53361 | 84573 | 54829 | 83629 | 56280 | 82659 | 45 |
| 16 | 50403 | 86369 | 51902 | 85476 | 53386 | 84557 | 54854 | 83613 | 56305 | 82643 | 44 |
| 17 | 50428 | 86354 | 51927 | 85461 | 53411 | 84542 | 54878 | 83597 | 56329 | 82626 | 43 |
| 18 | 50453 | 86340 | 51952 | 85446 | 53435 | 84526 | 54902 | 83581 | 56353 | 82610 | 42 |
| 19 | 50478 | 86325 | 51977 | 85431 | 53460 | 84511 | 54927 | 83565 | 56377 | 82593 | 41 |
| 20 | 50503 | 86310 | 52002 | 85416 | 53484 | $8449^{5}$ | 54951 | 83549 | 56401 | 82577 | 40 |
| 21 | 50528 | 86295 | 52026 | 85401 | 53509 | 84480 | 54975 | 83533 | 56425 | 82.561 | 39 |
| 22 | 50553 | 86281 | 52051 | 85385 | 53534 | 84464 | 54999 | ${ }^{835} 17$ | 56449 | 82054 | 38 |
| 23 | 50578 | 86266 | 52076 | 85370 | 53558 | 84448 | 55024 | 83501 | 56473 | 82528 | 37 |
| 24 | 50603 | 86251 | 52101 | 85355 | 53583 | 84433 | 55048 | 83485 | 56497 | 82511 | 36 |
| 25 | 50628 | 86237 | 52126 | 85340 | 53607 | 84417 | 55072 | 83469 | 56521 | 82495 | 35 |
| 26 | 50654 | 86222 | 52151 | 85325 | 53632 | 84402 | 55097 | 83453 | 56545 | 82478 | 34 |
| 27 | 50679 | 86207 | 52175 | 85310 | 53656 | 84386 | 55121 | 83437 | 56509 | 82462 | 33 |
| 28 | 50704 | 86192 | 52200 | 85294 | 53681 | 84370 | 55145 | 83421 | 56593 | 82446 | 32 |
| 29 | 50729 | 86178 | 52225 | 85279 | 53705 | 84355 | 55169 | ${ }^{83405}$ | 56617 | 82429 | 31 |
| 30 | 50754 | 86163 | 52250 | 85 | 53730 | 84339 | 55194 | 83 | 56641 | 82413 | 30 |
| 3 I | 50 | 86148 | 52275 | 85249 | 53754 | 84324 | 55218 | 83373 | 56665 | 82396 | 29 |
| 32 | 50804 | 86133 | 52299 | 85234 | 53779 | 84308 | 55242 | 83356 | 56689 | 82380 | 23 |
| 33 | 50829 | 86119 | 52324 | 85218 | 53804 | 84292 | 55266 | 83340 | 56713 | 82363 | 27 |
| 34 | 50854 | 86104 | 52349 | 85203 | 53828 | 84277 | 55291 | 83324 | 56736 | 82347 | 26 |
| 35 | 50879 | 86089 | 52374 | 85188 | 53853 | 84261 | 55315 | 83308 | 56760 | 82330 | 25 |
| 36 | 50904 | 86074 | 52399 | $8517^{3}$ | 53877 | 84245 | 55339 | 83292 | 56784 | 82314 | 2.4 |
| 37 | 50929 | 86059 | 52423 | 85157 | 53902 | 84230 | 55363 | 83276 | 56808 | 82297 | 23 |
| 38 | 50954 | 86045 | 52448 | 85142 | 53926 | 84214 | 55388 | 83260 | 56832 | 82281 | 22 |
| 39 | 50979 | 86030 | 52473 | 85127 | 53951 | $8419^{8}$ | 55412 | 83244 | 56856 | 82264 | 21 |
| 40 | 51004 | 86015 | 52498 | 85112 | 53975 | 84182 | $5{ }^{5} 5436$ | 83228 | 56880 | 82248 | 20 |
| 41 | 51029 | 86000 | 52.522 | 85006 | 54000 | 84167 | 55460 | ${ }_{8}^{8312}$ | 56904 | 82231 | 19 |
| 42 | 51054 | 85985 | 52547 | 85081 | 54024 | 84151 | 55484 | 83195 | 56928 | 82214 | 18 |
| 43 | 51079 | 85970 | 52572 | 85066 | 54049 | 84135 | 555009 | 83179 | 56952 | $8219^{8}$ | 17 |
| 44 | 51104 | 835056 | 52597 | 8505ı | 54073 | 84120 | 55533 | 83163 | 56976 | 82181 | 16 |
| 45 | 51129 | 85941 | 52621 | 85035 | 54097 | 841 | 55557 | 83147 | 57000 | 82165 | 15 |
| 46 | 51154 | 85926 | 52646 | 85020 | 54122 | 84088 | 55581 | 83131 | 57024 | 82148 | 14 |
| 47 | 51179 | 85911 | 52671 | 85005 | 54146 | 84072 | 55605 | 83115 | 57047 | 82132 | 13 |
| 48 | 51204 | 85896 | 52696 | 84989 | 54171 | 84057 | 55630 | $830{ }^{8}$ | 57071 | 82115 | 12 |
| 45 | 51229 | 85881 | 52720 | 84974 | 54195 | 84041 | 55654 | 83082 | 57095 | $8200^{8}$ | 11 |
| 50 | 51254 | 85866 | 52745 | 84959 | 54220 | 84025 | 55678 | 83066 | 57119 | 82082 | 10 |
| 51 | 51279 | 8585ı | 52770 | 84943 | 54244 | 84009 | 55.702 | 83050 | 57143 | 82065 | 9 |
| 52 | 51304 | 85836 | 52794 | 84928 | 54269 | 83994 | 55526 | 83034 | 57167 | 82048 | 8 |
| 53 | 51329 | 85821 | 52819 | 84913 | 54293 | 83978 | 55050 | 83017 | 57191 | 82032 | 7 |
| 54 | 51354 | 85806 | 52844 | 84897 | 54317 | 83962 | 55775 | 83001 | 57215 | 82015 | 6 |
| 55 | 51379 | 85792 | 52869 | 84882 | 54342 | 83946 | 55799 | 82985 | 57238 | 81999 | 5 |
| 56 | 51404 | 85777 | 52893 | 84866 | 54366 | 83930 | 55823 | 82969 | 57262 | 81982 | 4 |
| 57 | 51429 | 85762 | 52918 | 84851 | 54391 | 83915 | 55847 | 82953 | 57286 | 81965 | 3 |
| 58 | 51454 | 85747 | 52943 | 84836 | 54415 | 83899 | 55871 | 82936 | 57310 | 81949 | 2 |
| 59 | 51479 | 85732 | 52967 | 84820 | 54440 | 83883 | 55895 | 82920 | 57334 | 81932 | 1 |
| 60 | 5:504 | 85717 | 52952 | 84805 | 54464 | 83867 | 55919 | 82904 | 57358 | 81915 | - |
|  | Cosine. | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosine. | Sine. |  |
|  | $59^{\circ}$ |  | $58^{\circ}$ |  | $67^{\circ}$ |  | $56^{\circ}$ |  | $55^{\circ}$ |  |  |


| Table III. |  |  | NATURAL SINES AND COSINES. |  |  |  |  |  |  |  | 71 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $35^{\circ}$ |  | $36^{\circ}$ |  | $37^{\circ}$ |  | $38^{\circ}$ |  | $39^{\circ}$ |  |  |
|  | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosine. |  |
| 0 | 57358 | 81915 | 58779 | 80902 | 60182 | 79864 | 6ı566 | 78801 | 62932 | 77715 | So |
| 1 | 5738 I | 81899 | 58802 | 80885 | 60205 | 79846 | 61589 | 78783 | 62955 | 77696 | 59 |
| 2 | 57405 | 81882 | 58826 | 80867 | 60228 | 79829 | 61612 | 78765 | 62977 | 77678 | 58 |
| 3 | 57429 | 81865 | 58849 | 80850 | 60251 | $79^{911}$ | 61635 | 78747 | 63000 | 77660 | 57 |
| 4 | 57453 | 81848 | 58873 | 80833 | 60274 | 79793 | 61658 | 78729 | 63022 | 77641 | 56 |
| 5 | 57477 | 8183.2 | 58896 | 80816 | 60298 | 79776 | 6ı681 | 78711 | 63045 | 77623 | 55 |
| 6 | 57501 | 81815 | 58920 | 80799 | 60321 | 79758 | $6^{6} 704$ | 78694 | 63068 | 77605 | 54 |
| 7 | 57524 | ${ }^{81} 798$ | 58943 | 80782 | 60344 | 79741 | 61726 | 78676 | 63090 | 77586 | 53 |
| 8 | 57548 | 81782 | 58967 | 80765 | 60367 | 79723 | 61749 | 78658 | 63113 | 77568 | 52 |
| 9 | 57572 | 81765 | 5899 a | 8 c 748 | 60390 | 79706 | 61772 | 78640 | 63ı35 | 77550 | 51 |
| 10 | 57596 | 81748 | 59014 | 80730 | 50414 | 79688 | 61795 | 78622 | 63158 | 77531 | jo |
| 11 | 57619 | ${ }^{817} 71$ | 59037 | 80713 | 60437 | 79671 | 61818 | 78604 | 63180 | 77513 | 49 |
| 12 | 57643 | 81714 | 59061 | 80696 | 60460 | 79653 | 61841 | 78586 | 63203 | 77494 | 48 |
| 13 | 57667 | 81698 | 59084 | 80679 | 60483 | 79635 | 61864 | 78568 | 63225 | 77476 | 47 |
| 14 | 57691 | 81681 | 59108 | 80662 | 60506 | 79618 | 61887 | 78550 | 63248 | 77458 | 46 |
| 15 | 57715 | 81664 | 5913I | 80644 | 60529 | 79600 | 61909 | 78532 | 63271 | 77439 | 45 |
| 16 | 57738 | 81647 | 59154 | 80627 | 60553 | 79583 | 61932 | 78514 | 63293 | 77421 | 44 |
| 17 | 57762 | 81631 | 59178 | 80610 | 60576 | $79^{565}$ | 61955 | 78496 | 63316 | 77402 | 43 |
| 18 | 57786 | 81614 | 59201 | 80593 | 60599 | 79547 | 61978 | 78478 | 63338 | 77384 | 42 |
| 19 | 57810 | 81597 | 59225 | 80576 | 60622 | 79530 | 62001 | 78460 | 63.361 | 77366 | 41 |
| 20 | 57833 | 81580 | 59248 | 80558 | b0645 | 79512 | 62024 | 78442 | 63383 | 77347 | 40 |
| 21 | 57857 | 81563 | 59272 | 80541 | 60668 | 79494 | 62046 | 78424 | 63406 | 77329 | 39 |
| 22 | 57881 | 81546 | 59295 | 80524 | 60691 | 79477 | 62069 | 78405 | 63428 | 77310 | 38 |
| 23 | 57904 | 81530 | 59318 | 80507 | 60714 | 79459 | 62092 | 78387 | 63451 | 77292 | 37 |
| 24 | 57928 | 81513 | 59342 | 80489 | 60738 | 79441 | 62115 | 78369 | 63473 | $7727^{3}$ | 36 |
| 25 | 57952 | 81496 | 59365 | 80472 | 60761 | 79424 | 62138 | 78351 | 63496 | 77255 | 35 |
| 26 | 57976 | 81479 | 59389 | 80455 | 60784 | 79406 | 62160 | 78333 | 63518 | 77236 | 34 |
| 27 | 57999 | 81462 | 59412 | 80438 | 60807 | $79^{388}$ | 62183 | 78315 | 63540 | 77218 | 33 |
| 28 | 58023 | 81445 | 59436 | 80420 | 60830 | $79^{3} 71$ | 62206 | 78297 | 63563 | 77199 | 32 |
| 29 | 58047 | 81428 | 59459 | 80403 | 60853 | 79353 | 62229 | 78279 | 63585 | 77181 | 31 |
| 30 | 58070 | 81412 | 59482 | 80386 | 60876 | 79335 | 6225 I | 78261 | 63608 | 77162 | 30 |
| 31 | 58094 | 81395 | 59506 | 80368 | 60899 | 79318 | 62274 | 78243 | 63630 | 77144 | 29 |
| 32 | 58118 | 81378 | 59529 | 80351 | 60922 | $79^{300}$ | 62297 | 78225 | 63653 | 77125 | 28 |
| 33 | 58141 | 81361 | 59552 | 80334 | 60945 | 79282 | 62320 | 78206 | 63675 | 77107 | 27 |
| 34 | 58165 | 81344 | 59576 | 80316 | 60968 | 79264 | 62342 | 78188 | 63698 | 77088 | 26 |
| 35 | 58189 | 81327 | 59599 | 80299 | 60991 | 79247 | 62365 | 78170 | 63720 | 77070 | 25 |
| 36 | 58212 | 81310 | 59622 | 80282 | 61015 | 79229 | 62388 | 78152 | 63742 | 77051 | 24 |
| 37 | 58236 | 81293 | 59646 | 80264 | 6ı038 | 79211 | 62411 | 78134 | 63765 | 77033 | 23 |
| 38 | 58260 | 81276 | 59669 | 80247 | 61061 | 79193 | 62433 | 78116 | 63787 | 77014 | 2 |
| 39 | 58283 | 81259 | 59693 | 80230 | 61084 | 79176 | 62456 | 78098 | 63810 | 76996 | 21 |
| 40 | 58307 | 81242 | 59716 | 80212 | 61107 | 79158 | 62479 | 78079 | 63832 | 76977 | 0 |
| 41 | 58330 | 81225 | 59739 | 80195 | 61130 | 79140 | 62502 | 78061 | 63854 | 76959 | 9 |
| 42 | 58354 | 81208 | 59763 | 80178 | 61153 | 79122 | 62524 | 78043 | 63877 | 76940 | 18 |
| 43 | 58378 | 81191 | 59786 | 80160 | 61176 | 79105 | 62547 | 78025 | 63899 | 76921 | 7 |
| 44 | 58401 | 81174 | 59809 | 80143 | 61199 | 79087 | 62570 | 78007 | 63922 | 76903 | 16 |
| 45 | 58425 | 81157 | 59832 | 80125 | 61222 | 79069 | 62592 | 77988 | 63944 | 76884 | 15 |
| 46 | 58449 | 81140 | 59856 | 80108 | 61245 | 7905 I | 62615 | 77970 | 63966 | 76866 | 14 |
| 47 | 58472 | 81123 | 59879 | 80091 | 61268 | 79033 | 62638 | 77952 | 63989 | 76847 | 13 |
| 48 | 58496 | 81106 | 59902 | 80073 | 61291 | 79016 | 62660 | 77934 | 64011 | 76828 | 2 |
| 49 | 58519 | 81089 | 59926 | 80056 | 61314 | 78998 | 62683 | 77916 | 64033 | 76810 | 11 |
| 50 | 58543 | ${ }^{810} 7^{2}$ | 59949 | 80038 | 61337 | 78980 | 62706 | 77897 | 64056 | 76791 | 0 |
| 51 | 58567 | 81055 | 59972 | 80021 | 61360 | 78962 | 62728 | 77879 | 64078 | 76772 | 9 |
| 52 | 58590 | ${ }^{8} \mathrm{8}$ о38 | 59795 | 80003 | 61383 | 78944 | 62751 | 7786 I | 64100 | 76754 | 8 |
| 53 | 58614 | 81021 | 60019 | 79986 | 61406 | 78926 | 62774 | 77843 | 64123 | 76735 |  |
| 55 | 58637 | 81004 | 60042 | 79968 | 61429 | 78908 | 62796 | 77824 | 64145 | 76717 | 6 |
| 55 56 | 58661 | 809.97 | 60065 | 79951 | 61451 | 78891 | 62819 | 77806 | 64167 | 76698 | 5 |
| 56 | 58684 | 80c70 | 60089 | 79934 | 61474 | 78873 | 62842 | 77788 | 64190 | 76679 | 4 |
| 57 58 58 | 58708 | 80953 | 60112 | 79916 | 61497 | 78855 | 62864 | 77769 | 64212 | 76661 | 3 |
| 59 |  | 80ç | 60135 60158 |  | 61520 | 78837 | 62887 | 77751 | 64234 | 76642 | 2 |
| 60 | 53779 | 80go2 | 60182 | 79864 | 61566 | 7880 I | 62932 | 77715 | 64279 | 76604 | - |
|  | Cosine | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosine. | Sine. |  |
|  | $54^{\circ}$ |  | $53^{\circ}$ |  | $52^{\circ}$ |  | $51^{\circ}$ |  | $50^{\circ}$ |  |  |


| 72 |  |  | NATURAL SINES AND |  |  |  | COSINES. |  | Tabbee III. |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | $40^{\circ}$ |  | $41^{\circ}$ |  | $42^{\circ}$ |  | $43^{\circ}$ |  | $44^{\circ}$ |  |  |
|  | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosine. | Sine. | Cosin | Sine. | Cosine. |  |
| 0 | 64279 | 76604 | 65 | 75 | 66 | 74314 | 68200 | 73135 |  |  | 60 |
| 1 | 64301 | 76586 | 65628 | 75502 | 66935 | 74295 | 68221 | 73116 |  | 71914 | 59 |
| 2 | 64323 | 76567 | 65650 | 75.433 | 66956 | 74276 | 68242 | 73096 |  | 71894 | 58 |
| 3 | 64346 | 76548 | 65672 | 75414 | 66978 | 74256 | 68264 | 73075 | 69529 | 71873 | 57 |
| 4 | 64368 | 76530 | 65694 | 75395 | 66999 | 74237 | 68285 | 730556 | 69549 | 71853 | 56 |
| 5 | 64390 | 76511 | 65716 | 75375 | 67021 | 74217 | 68306 | 73036 | 69570 | 71833 | 55 |
| 6 | 64412 | 76492 | 65738 | 75356 | 67043 | 74198 | 68327 | 73016 | 69591 | 71813 | 54 |
| 7 | 64435 | $7647^{3}$ | 65759 | 75337 | 67064 | 74178 | 68349 | 72996 | 69612 | 71792 | 53 |
| 8 | 64457 | 76455 | 65781 | $7{ }^{5318}$ | 67086 | 74159 | 68370 | 72976 | 69633 | 71772 | 52 |
| 9 | 64479 | 76436 | 65803 | 75299 | 67107 | 74139 | 68391 | 729 9'7 | 69654 | 71752 | 51 |
| 10 | 64501 | 76417 | 65825 | $7{ }^{\text {²20 }}$ | 67129 | 74120 | 68412 | 72937 | 69675 | 71732 | 50 |
| 11 | 645124 | 76398 | 65847 | 75021 | 67151 | 74100 | 68434 | 72917 | 69696 | 71711 | 49 |
| 12 | 64546 | 76380 | 65869 | 75241 | 67172 | 74080 | 68455 | 72897 | 69717 | 71691 | 48 |
| 13 | 640568 | 76361 | 65891 | 75222 | 67194 | 74061 | 68476 | 72877 | 69737 | 71671 | 47 |
| 14 | 643090 | 76342 | 65913 | $7{ }^{\text {J203 }}$ | 67215 | 74041 | 68497 | 72857 | 69758 | 71650 | 46 |
| 15 | 64612 | 76323 |  |  | 67237 | 74022 |  |  | 69779 |  | 45 |
| 16 | 64635 | 76304 | 65956 | 75165 | 67258 | 74002 | 68539 | 72817 | 69800 | 71610 | 44 |
| 17 | 64657 | 76286 | 65978 | 75146 | 67280 | 73983 | 68561 | 72797 | 69821 | 71590 | 43 |
| 18 | 64679 | 76267 | 66000 | 75126 | 67301 | 73963 | 68582 | 72777 | 69842 | 71569 | 42 |
| 19 | 64701 | 76248 | 66022 | 75107 | 67323 | 73944 | 68603 | 72757 | 69862 | 71549 | 41 |
| 20 | 64723 | 76229 | 66044 | 75088 | 67344 | 73924 | 68824 | 72737 | 69883 | 71529 | 40 |
| 21 | 64746 | 76210 | 66066 | 75069 | 67366 | 73904 | 68645 | 72717 | 69904 | 71508 | 39 |
| 22 | 64768 | 76192 | 66088 | 75050 | 67387 | 73885 | 68666 | 72697 | 69925 | 71488 | 38 |
| 23 | 64790 | 76173 | 66109 | 75030 | 67409 | 73865 | 68688 | 72677 | 69946 | 71468 | 37 |
| 24 | 64812 | 76154 | 66131 | 75011 | 67430 | 73846 | 68709 | 72657 | 69966 | 71447 | 36 |
| 25 | 64834 | 76135 | 66153 | 74992 | 67452 | 73826 | 68730 | 72637 | 69987 | 71427 | 35 |
| 26 | 64856 | 76116 | 66175 | 74973 | 67473 | 73806 | 68751 | 72617 | 70008 | 71407 | 34 |
| 27 | 64878 | 76097 | 66197 | 74953 | 67495 | 73787 | 68772 | 72597 | 70029 | 71386 | 33 |
| 28 | 64901 | 76078 | 66218 | 74934 | 67516 | 73767 | 68793 | 72577 | 70049 | 71366 | 32 |
| 29 | 64923 | 76059 | 66240 | 74915 | 67533 | 73747 | 68814 | 72557 | 70070 | 71345 | 31 |
| 30 | 64945 | 76041 |  | 74896 | 675059 | 73 |  | 72 | 7009 | 71325 | 30 |
| 31 | 64967 | 76022 | 66284 | 74876 | 67580 | 73708 | 68857 | 72517 | 12 | 71305 | 29 |
| 32 | 64989 | 76003 | 6630 | 74857 | 67602 | 73688 | 68878 | 72497 | 0132 | 71284 | 28 |
| 33 | 65011 | 75984 | 6632 | 74838 | 67623 | 73669 | 68899 | 72477 | 15 | 1264 | 27 |
| 34 | 65033 | 75965 | 66349 | 74818 | 6";645 | 73649 | 68920 | 72457 | 70174 | 71243 | 26 |
| 35 | 65055 | 75946 | 66371 | 74799 | 67666 | 73629 | 68941 | 72437 | 70195 | 71223 | 25 |
| 36 | 65077 | 75927 | 66393 | 74780 | 67688 | 73610 | 68962 | 724:7 | 70215 | 71203 | 24 |
| 37 | 6510 c | $7{ }^{\text {7 }} 9008$ | 66414 | 74760 | 67709 | 7350 | 68983 | 72397 | 70236 | 71182 | 23 |
| 38 | 65122 | 75889 | 66436 | 74741 | 67730 | 73570 | 69004 | 72377 | 70257 | 71162 | 22 |
| 39 | 65144 | 75870 | 66458 | 74722 | 67752 | 73551 | 69025 | 72357 | 70277 | 71141 | 21 |
| 40 | 65166 | 75851 | 66480 | 74703 | 67773 | 73531 | 69046 | 72337 | 70298 | 71121 | 20 |
| 41 | 65188 | 75832 | 66501 | 74683 | 67795 | 73511 | 69067 | 72317 | 70319 | 71100 | 19 |
| 42 | 65210 | 75813 | 66523 | 74664 | 67816 | 73491 | 69088 | 72297 | 70339 | 71080 | 18 |
| 43 | 65232 | 75794 | 66 | 74644 | 67837 | $7347^{2}$ | 69109 | 72277 | 70360 | 71050 | 17 |
| 44 | 65254 |  | 66566 | 74625 | 67859 | ${ }_{7} 73$ ¢5 2 | 69130 | 72257 | 70381 | 71039 | 6 |
| 45 | 65276 |  |  | 74606 |  | 73 | 69151 | 72236 | 70401 | 71019 | 15 |
| 46 | 65298 | 75738 | 66610 | 74586 | 67901 | 73413 | 69172 | 72216 | 70422 | 70098 | 14 |
| 47 | 65320 | 75719 | 66632 | 74567 | 67923 | 73393 | 69193 | $721{ }^{\text {c }} 6$ | 70443 | 70978 | 13 |
| 48 | 65342 | 75700 | 66653 | 74548 | 67944 | $733{ }^{3}$ | 69214 | 72176 | 70463 | 70950 | 12 |
| 49 | 65364 | 75080 | 66675 | 74528 | 67965 | 73353 | 69235 | 72156 | 70484 | 70937 | 11 |
| 50 | 65386 | 75661 | 66697 | 74509 | 67987 | 73333 | 69256 | 72136 | 70505 | 70916 | 10 |
| 5 I | 65408 | 75042 | 66718 | 74489 | 68008 | 73314 | 69277 | 72116 | 70525 | 70896 | 9 |
| 52 | 65430 | 75623 | 66740 | $7447^{\circ}$ | 68029 | 73294 | 69298 | 72095 | 70546 | 70875 |  |
| 53 | 65450 | $7{ }^{\text {7 }} 5004$ | 66-62 | 74451 | 68051 | 73274 | 69319 | 72075 | 70567 | 70855 |  |
| 54 | 65474 | 70585 | 66783 | 74431 | 68072 | 73254 | 69340 | 72055 | 70587 | 70834 | 6 |
| 55 | 65596 | 75506 | 66805 | 74412 | 68093 | 73234 | 69361 | 72035 | 70608 | 70813 | 5 |
| 56 | 65518 | 75547 | 66827 | 74392 | 68115 | 73215 | 69382 | 72015 | 70628 | 70793 | 4 |
| 57 | 65540 | 75528 | 668.4 | 74373 | 68136 | 73195 | 69403 | 71995 | 706 亿9 | 70772 | 3 |
| 58 | $655 \overline{62}$ | 75509 | 66870 | 74353 | $66_{15} 5$ | ${ }_{7}^{7175}$ | 69424 | 71974 | 70670 | 70752 | 2 |
| 59 | 65584 | 75590 | 66891 | 74334 | 68179 | 73155 | 69445 | 71954 | 70690 | 70731 | 1 |
| 60 | 65 | 75471 | 6égl3 | 74314 | 68200 | 73135 | 69466 | 71934 | 70711 | 70711 | 0 |
|  | Cosine. | Sine. | ne. | Sine | ine. | Sine. | Cosine. | Sine. | Cosine. | Sine |  |
|  | $49^{\circ}$ |  | $48^{\circ}$ |  | $47^{\circ}$ |  | $46^{\circ}$ |  | $45^{\circ}$ |  |  |


| Table III. N |  |  | ATURAL TANGENTS AND CO'TANGENTS. |  |  |  |  |  | 73 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | $0^{5}$ |  | $1^{\circ}$ |  | $2^{\circ}$ |  | $3^{\circ}$ |  | , |
|  | Tangent. | Cotang. | Tangent. | Cotang. | Tangent. | Cotang. | Tange.lt. | Cotang. |  |
| 0 | 00000 | Infinite. | 01746 | 57.2900 | 03492 | 28.6363 | 05241 | 19.0811 | 60 |
| 1 | 00029 | $3437 \cdot 75$ | 01775 | 56.3506 | 03521 | 28.3994 | 05270 | 18.9755 | 59 |
| 2 | 00058 | 1718.87 | 01804 | $55 \cdot 4415$ | 03550 | $28 \cdot 1664$ | 05299 | 18.8711 | 58 |
| $\stackrel{3}{3}$ | 00087 | 1145.92 | 0ı833 | 54.5613 | o3579 | $27 \cdot 9372$ | 05328 | $18.767^{8}$ | 57 |
| 4 | 00116 | $859 \cdot 436$ | 01862 | 53.708 t | -3609 | $27 \cdot 7117$ | 05357 | 18.6656 | 55 |
| 5 | 00145 | 687.549 | or891 | 52.8821 | o3638 | 27-4899 | 05387 | 18.5645 | 55 |
| 6 | 00175 | 572.957 | 01920 | 52.0807 | 03667 | $27 \cdot 2715$ | 05416 | 18.4645 | 54 |
| 7 | c0204 | $491 \cdot 106$ | -1949 | $5 \mathrm{~L} \cdot 3032$ | 03696 | 27.0566 | 05445 | 18.3655 | 53 |
| 8 | 00233 | $429 \cdot 718$ | 01978 | 50.5485 | c3725 | 26.8450 | 05474 | 18.2677 | 52 |
| 10 | 00262 | 381.971 | 02007 | $49 \cdot 8157$ | 03754 | 26.6367 | o5503 | 18.1708 | 51 |
| 10 | 00291 | $343 \cdot 774$ | 02036 | 49.1039 | o3783 | $26.431^{5}$ | 05533 | $18.075=$ | 50 |
| 11 | 00320 | 312.521 | 02066 | 48.4121 | 03812 | $26 \cdot 22,6$ | 05562 | 17.9802 | 49 |
| 12 | oo349 | 286.478 | 02095 | $47 \cdot 739^{5}$ | o3842 | $26 \times 307$ | -5591 | 17.8863 | 48 |
| 13 | 00378 | $264 \cdot 441$ | 02124 | 47.0853 | 03871 | 258348 | 05620 | $17 \cdot 7934$ | 47 |
| 14 | 00407 | 245.552 | 02153 | $46 \cdot 4489$ | 03900 | 25.6418 | 05649 | $17 \cdot 7015$ | 46 |
| 15 | 00436 | $229 \cdot 182$ | 02182 | $45 \cdot 8294$ | -3929 | $25 \cdot 4517$ | 05678 | 17.6106 | 45 |
| 16 | 00465 | 214.858 | 02211 | 45.2261 | o3, 58 | 25.2644 | 05708 | 17.5205 | 44 |
| 17 | 00495 | $202 \cdot 219$ | 02240 | 44.6386 | 03987 | 25.0798 | -5737 | 17.4314 | 43 |
| 18 | 00524 | $190 \cdot 984$ | 02269 | 44.0661 | 04016 | 24.8978 | 05766 | 17.3432 | 42 |
| 15 | oo553 | 180.032 | 02298 | 43.5081 | 04046 | 24.7185 | -5795 | 17.2558 | 41 |
| 2 c | 00582 | 171.885 | 02328 | 42.9641 | 04075 | 24.5418 | 05824 | 17.1693 | 40 |
| 21 | 00611 | $163 \cdot 700$ | 02357 | 42.4335 | 04104 | 24.3675 | o5854 | $17 \cdot 0837$ | 39 |
| 22 | 00640 | $156 \cdot 259$ | 02386 | $41 \cdot 9158$ | 04133 | $24 \cdot 1957$ | o5883 | 16.9990 | 38 |
| 23 | 00669 | 149.465 | 02415 | 41.4106 | 04162 | 24.0263 | -5912 | 16.9150 | 37 |
| 24 | 00698 | 143.237 | 02444 | $40 \cdot 9174$ | 04191 | 23.8593 | 05941 | 16.8319 | 36 |
| 25 | 00727 | 137.507 | $0247^{3}$ | $40 \cdot 4358$ | 04220 | 23.6945 | 05970 | 16.7496 | 35 |
| 26 | 00756 | 132.219 | 02502 | $39 \cdot 9655$ | 04250 | 23.5321 | 05999 | 16.6681 | 34 |
| 27 | 00785 | $127 \cdot 321$ | 0253ı | 39.5059 | 04279 | 23.3718 | 06029 | 16.5874 | 33 |
| 28 | 00814 | 122.774 | 02560 | 39.0568 | 04308 | $23 \cdot 2137$ | o6o58 | 16.5075 | 32 |
| 29 | 00844 | 118.540 | 02589 | 38.6177 38.1885 | 04337 04366 | 23.0577 | 06087 | 16.4283 16.3499 | 31 30 |
| 31 | 00902 | $110 \cdot 892$ | 02648 | $37 \cdot 7686$ | 04395 | 22.7519 | 06145 | 16.2722 | 29 |
| 32 | 00931 | $107 \cdot 426$ | 02677 | 37.3579 | 04424 | 22.6020 | 06175 | $16 \cdot 19^{5} 2$ | 28 |
| 33 34 | 00960 | $104 \cdot 171$ | 02736 | 36.9560 | 04454 | 22.454 l | 06204 | 16.1190 | 27 |
| 34 35 | 00989 01018 | 101.107 | 02735 | 36.5627 | 04483 | 22.3081 | 06233 | 16.0435 | 26 |
| 35 36 | 01018 | 98.2179 | 02764 | 36.1776 | 04512 | $22 \cdot 1640$ | 05262 | 15.8687 | 25 |
| 36 37 | 01047 01076 | $95 \cdot 4895$ 92.9085 | 02793 02822 | 35.8006 35.43 I 3 | 04541 04570 | 22.0217 21.8813 | 06291 06321 | 15.8945 15.8211 | 24 23 |
| 38 | 01105 | 90.4633 | 0285I | $35 \cdot 0695$ | 04599 | 21.7426 | 06350 | 15.7483 | 22 |
| 39 | 01135 | 88.1436 | 0288ı | 34.7151 | 0.4628 | 21.6056 | 06379 | 15.6762 | 21 |
| 40 | 01164 | $85 \cdot 0^{3} 98$ | 02910 | 34.3678 | 04658 | 21.4704 | 06408 | 15.6048 | 20 |
| 41 | 01193 | 83.8435 | 02939 | 34.0273 | 04687 | 21. 3369 | 06437 | 15.5340 | 19 |
| 42 | O1222 | 81.8470 | 02968 | 33.6935 | 04716 | 21-2049 | 06467 | 15.4638 | 18 |
| 43 | O1251 | 79.9434 | 02997 | 33.3662 | 04745 | 21.0747 | 06496 | 15.3943 | 17 |
| 44 | -1280 | $78 \cdot 1263$ | 03026 | 33.0452 | 04774 | 20.9460 | 06525 | 15.3254 | 16 |
| 45 | 01309 | $76 \cdot 3900$ | 03055 | $32 \cdot 73 \mathrm{o} 3$ | 04803 | 20.8188 | 06554 | 15.2571 | 15 |
| 46 | O1338 | $74 \cdot 7292$ | 03084 | 32.4213 | 04832 | 20.6932 | 06584 | 15.1893 |  |
| 47 | 01367 | $73 \cdot 1390$ | 03114 | 32-118I | 04862 | 20.5691 | 066.13 | 15.1222 | 13 |
| 48 | 01396 | 71.6151 | 03143 | $3 \mathrm{I} \cdot 8205$ | 04871 | $20 \cdot 4465$ | 06642 | 15.0557 | 12 |
| 49 | 01425 | $70 \cdot 15.33$ | 03172 | 31.5284 | 04920 | 203253 | 06671 | 14.9898 | 11 |
| 50 | 01455 | $68 \cdot 7501$ | 03201 | 31.2416 | 04949 | 20.2056 | 06700 | 14.8244 | 10 |
| 51 52 | 01484 | 67.4019 | 03230 | 30.9599 | 04978 | 20.0872 | 06730 | $14.859^{6}$ | 8 |
| 52 | 01513 | $66 \cdot 1055$ 64.8580 | 03259 03288 033 | 30.6833 30.4116 | 05007 | 19.9702 19.8546 | 06759 06788 | 14.7954 14.7317 | 8 |
| 54 | 0.571 | 63.6567 | o3317 | 30.1446 | -5066 | 19.7403 | 06817 | 14.6685 | 6 |
| 55 | 01600 | $62 \cdot 4992$ | 03346 | 29.8823 | 05095 | 19.6273 | 06847 | 14.6059 | 5 |
| 56 | 01629 | 61.3829 | 03376 | 29.6245 | 05124 | 19.5156 | 06876 | 1.4.5438 | 4 |
| 57 | 01658 | 60.3058 | 03405 | 29.3711 | 05153 | $19 \cdot 4051$ | 06905 | 14.4823 |  |
| 5 | 01687 | $59 \cdot 2659$ | 03434 | $29 \cdot 1220$ | -5182 | $19 \cdot 20^{59}$ | 06934 | 14.4212 | 2 |
| 54 60 | 01716 | 58.2612 | 03463 | 28.8771 | ${ }^{0} 5212$ | $19 \cdot 1879$ | 06963 | 14.3607 | 1 |
| 60 | 01746 | 57.2900 | 03492 | 28.6363 | 05241 | 19.0811 | 06993 | 14.3007 | 0 |
|  | Cotang. | Tangent. | Cotang. Tangent. |  | $\overline{\text { Cotang. }}$ Tangent. |  | Cotang. Tangent. |  |  |
|  | $89^{\circ}$ |  | $88^{\circ}$ |  | $87^{\circ}$ |  | $86^{\circ}$ |  |  |


| 74 | NATURA |  | TANGENTS |  | AND COTANGENV'S. |  |  | Table III. |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $4^{\circ}$ |  | $5^{\circ}$ |  | $6^{\circ}$ |  | 70 |  |  |
|  | Tangent. | Cotang. | Tangent. | Cotang. | Tangent. | Cotang. | Tangent. | Cotang. |  |
| 0 | 06993 | 14.3007 | 08749 | 11.4301 | 10510 | 9.51436 | 12278 | 8.14435 | 60 |
| 1 | 07022 | 14.2411 | 08778 | 11.3919 | 10540 | 9.48781 | 12308 | 8.12481 | 5 |
| 2 | 07051 | 14.1821 | 08807 | 11.3540 | 1 c 569 | 9.46141 | 12338 | $8 \cdot 10536$ | 53 |
| 3 | 07080 | 14.1235 | 08837 | 11.3i63 | 10599 | 9-43515 | 12367 | 8.08600 | 57 |
| 4 | 07110 | 14.0655 | 08866 | 11.2789 | 10628 | 9.40904 | 12397 | 8.06674 | 55 |
| 5 | 07139 | 14.0079 | 08895 | 11.2417 | 10657 | 9.38307 | 12426 | 8.04756 | 55 |
| 6 | c7168 | 13.9507 | 08925 | II-2048 | 10687 | 9.35724 | 12456 | 8.02848 | 54 |
| 7 | 07197 | 13.8940 | 08954 | II.168I | 10716 | $9 \cdot 33154$ | 12485 | $8 \cdot 00948$ | 53 |
| 8 | 07227 | 13.8378 | 08983 | 11.1316 | 10746 | $9 \cdot 30509$ | 12515 | 7.99058 | 5.2 |
| 9 | 07256 | 13.7821 | هوо13 | 11.0054 | 10775 | 9.28058 | 12544 | $7 \cdot 97176$ | 51 |
| 10 | 07285 | 13.7267 | 09042 | 11.0094 | 10805 | 9.2553o | 12574 | $7 \cdot 95302$ | 50 |
| 11 | 07314 | 13.6719 | 09071 | 11.0237 | 10834 | 9.23016 | 12603 | 7.93438 | 49 |
| 12 | 07344 | 13.6174 | 09101 | $10 \cdot 9882$ | 10863 | 9.20516 | 12633 | $7 \cdot 91582$ | 48 |
| 13 | 07373 | 13.5634 | -9130 | $10 \cdot 9^{5} 29$ | 10893 | 9.18028 | 12662 | 7.89734 | 47 |
| 14 | 07402 | $13.50{ }^{8}$ | -9159 | 10.9178 | 10922 | 9.15554 | 12692 | 7.87895 | 46 |
| 15 | 07431 | 13.4566 | 09189 | 10.8829 | 10952 | 9.13093 | 12722 | 86064 | 45 |
| 16 | 07461 | 13.4039 | 09218 | 10.8483 | 10981 | 9.10646 | 12751 | 7.84242 | 44 |
| 17 | 07490 | 13.3515 | 09247 | 10.8139 | 11011 | 9.08211 | 12781 | 7.82428 | 43 |
| 18 | 07519 | 13.2996 | 09277 | 10.7797 | 11040 | 9.05789 | 12810 | $7 \cdot 80622$ | 42 |
| 19 | 07548 | 13.2480 | 09306 | 10.7457 | 11070 | 9.03379 | 12840 | $7 \cdot 78825$ | 41 |
| 20 | 07578 | 13.1969 | 09335 | 10.7119 | 11099 | 9-00983 | 12869 | $7 \cdot 77035$ | 40 |
| 21 | 07607 | 13.1461 | 09365 | 10.6783 | 11128 | $8 \cdot 98598$ | 12899 | $7 \cdot 75254$ | 39 |
| 22 | 07636 | 13.0958 | $00^{3} 94$ | 10.6450 | 11158 | $8 \cdot 96227$ | 12929 | $7 \cdot 73480$ | 38 |
| 23 | 07665 | 13.0458 | 09423 | 10.6118 | 11187 | $8 \cdot 93867$ | 12958 | 7.71715 | 37 |
| 24 | 07695 | 12.9962 | 09453 | 10.5789 | 11217 | $8 \cdot 91520$ | 12988 | 7.69957 | 36 |
| 25 | 07724 | 12.9469 | 09482 | 10.5462 | 11246 | 8.89185 | 13017 | 7.68208 | 35 |
| 26 | 07753 | 12.8981 | -9511 | 10.5136 | 11276 | 8.86862 | 13047 | 7.66466 | 34 |
| 27 | 07782 | 12.8496 | 09541 | $10 \cdot 4813$ | 11305 | 8.84551 | 13076 | 7.64732 | 33 |
| 28 | 07812 | 12.8014 | 09570 | 10.4491 | 11335 | $8 \cdot 82252$ | 13106 | 7.63005 | 32 |
| 29 | 07841 | 12.7536 | 09600 | 10.4172 | 11364 | $8 \cdot 79964$ | 13136 | 7.61287 | 31 |
| 30 | 07870 | 12.7062 | 09629 | 10.3854 | 11394 | $8 \cdot 77689$ | 13165 | 7-5.9575 | 30 |
| 31 | 07899 | 12.6591 | -9658 | 10.3538 | 11423 | $8 \cdot 75425$ | 13195 | 7.57872 | 29 |
| 32 | 07929 | 12.6124 | 09688 | 10.3224 | 11452 | $8 \cdot 73172$ | 13224 | 7.56176 | 28 |
| 33 | 07958 | 12.5660 | 09717 | 10.2913 | 11482 | $8 \cdot 70931$ | 13254 | 7.54487 | 27 |
| 34 | 07987 | 12.5199 | 09746 | 10.2602 | 11511 | 8.687 이 | 13284 | 7.52806 | 26 |
| 35 | 08017 | 12.4742 | 09776 | 10.2294 | 11541 | 8.66482 | 133 r 3 | $7 \cdot 51132$ | 25 |
| 36 | 08046 | 12.4288 | 09805 | 10.1988 | 11570 | 8.64275 | 13343 | 7.49765 | 24 |
| 37 38 | 08075 | 12.3838 | 09834 | 10.1683 | 11600 | 8.62078 | 13372 | 7.47806 | 23 |
| 38 | 08104 | 12.3390 | 0, 864 | 10.1381 | 11629 | $8 \cdot 59893$ | 13402 | 7-46154 | 22 |
| 39 | 08134 | 12.2946 | -9893 | 10.1080 | 11659 | $8.571^{18}$ | 13432 | 7.44509 | 21 |
| 40 | 08163 | 12.2505 | 09923 | 10.0780 | 11688 | $8 \cdot 55555$ | 13461 | 7.42871 | 20 |
| 41 | 08192 | 12.2067 | 09952 | $10 \cdot 0483$ | 11718 | 8.53402 | 13491 | 7.41240 | 19 |
| 42 | 08221 | 12.1632 | -9981 | 10.0187 | 11747 | $8 \cdot 51250$ | 13521 | 7-39616 | 18 |
| 43 | 08251 | 12.1201 | 10011 | 9.9893o | 11777 | 8-49128 | 13550 | 7.37999 | 17 |
| 44 | 08280 | 12.0772 | 10040 | 9.96007 | 11806 | $8 \cdot 47007$ | 13580 | $7 \cdot 36389$ | 10 |
| 45 | 08309 | 12.0346 | 10069 | -3101 | I 1 | 8.44896 | 13609 | $7 \cdot 34786$ | 15 |
| 46 | 08339 | 11.9923 | 10099 | 9.90211 | 11865 | $8 \cdot 42795$ | 13639 | $7 \cdot 33190$ | 1.4 |
| 47 | 08368 | 11.9504 | 10128 | $9 \cdot 87338$ | 11895 | 8.40705 | 13669 | 7.31600 | 13 |
| 48 | 08397 | 11.9087 | 10158 | 9.84482 | 11924 | 8.38625 | 13693 | $7 \cdot 30018$ | 12 |
| 49 | 08427 | 11.8673 | 10187 | 9.81641 | 11954 | 8.36555 | 13728 | 7-28442 | 11 |
| 50 | 08456 | 11.8262 | 10216 | 9.78817 | 11983 | 8.34496 | 13758 | 7.26873 | 10 |
| 51 52 | 08485 | 11.7853 | 10246 | 9•76009 | 12013 | 8.32446 | 13787 | $7 \cdot 25310$ | 9 |
| 52 53 | 08514 | 11.7448 | 10275 | 9.73217 | 12042 | 8.30406 $8.283-6$ | 13817 138 | 7.23754 7.2204 7.2064 | 8 |
| 54 | 08573 | 11.6645 | 10334 | 9.67680 | 12101 | 8.26355 | 13876 | 7-2066 | 6 |
| 55 | 08602 | 11.6248 | 10363 | 9.64935 | 12131 | 8.24345 | 13906 | $7 \cdot 19125$ | 5 |
| 56 | 08632 | 11.5853 | 10393 | $9 \cdot 62205$ | 12160 | 8.22344 | 13935 | 7-17504 | 4 |
| 57 | 08661 | 11.5461 | 10422 | 9.50490 | 12190 | $8 \cdot 20352$ | 13965 | $7 \cdot 16071$ | 3 |
| 38 | 08690 | 11.5072 | 10452 | $9 \cdot 56791$ | 12219 | $8 \cdot 18370$ | 13995 | 7-14553 | 2 |
| 59 | 08;20 | 11.4685 | 10481 | 9.54106 | 12249 | 8.16398 | 14024 | 7-13042 | 1 |
| 60 | 08749 | 11.4301 | 10510 | 9.51436 | 12278 | 8.14435 | 14054 | 7 | c |
|  | Cotang. | Tangent. | Cotang. Tangent. |  | Cotang. ${ }^{\text {Tangent. }}$ |  | Cotang. ${ }^{\text {a }}$ Tangent. |  | , |
|  |  | $5^{\circ}$ | $84^{\circ}$ |  | $83^{\circ}$ |  | $82^{\circ}$ |  |  |


| Table III. |  | NATURAL TANGENTS |  |  |  | AND COTANGENTS. |  |  | 75 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | $8^{\circ}$ |  | $9^{\circ}$ |  | $10^{\circ}$ |  | $11^{\circ}$ |  | 1 |
|  | Tangent. | Cotang. | Tangent. | Cotang. | Tangent. | Cotang. | Tangent. | Cotang. |  |
| 0 | 14054 | 7-11537 | 15838 | 6.31375 | 17633 | 5.67128 | 19438 | 5.14455 | 60 |
| 1 | 14084 | 7-10038 | 15868 | 6.30189 | 17663 | 5.66165 | 19468 | 5.13658 | 59 |
| 2 | 14113 | 7.08546 | 15898 | 6.29007 | 17693 | 5.65205 | 19498 | 5.12862 | 58 |
| 3 | 14143 | 7.07059 | 15928 | 6.27829 | 17723 | 5.64248 | 19529 | 5.12069 | 57 |
| 4 | 14173 | 7.05579 | 15958 | 6.26655 | 17753 | 5.63295 | 1955 | 5.11279 | 55 |
| 5 | 14202 | $7 \cdot 04105$ | 15088 | 6.25486 | 17783 | 5.62344 | 19589 | 5.1049 ${ }^{\text {c }}$ | 55 |
| 6 | 14232 | 7.02637 | 16017 | 6.24321 | 17813 | 5.61397 | 19619 | 5.09704 | 54 |
| 7 | 14262 | 7.01174 | 16047 | 6.23160 | 17843 | 5.60452 | 19649 | 5.08921 | 53 |
| 8 | 14291 | 6.99718 | 16077 | 6. 22003 | 17873 | 5.50511 | 19680 | 5.08139 | 52 |
| 9 | 14321 | $6 \cdot 98268$ | 16107 | 6.20851 | 17903 | 5.58573 | 19710 | 5.07360 | 51 |
| 10 | 14351 | $6 \cdot 96823$ | 16137 | 6. 19703 | 17933 | 5.57638 | 19740 | 5.06584 | 50 |
| 11 | 14381 | $6 \cdot 95385$ | 16167 | 6.18559 | 17963 | 5.56706 | 19770 | 5.05809 | 49 |
| 12 | 14410 | $6 \cdot 93252$ | 16196 | 6.17419 | 17993 | 5.55777 | 19801 | 5.05037 | 48 |
| 13 | 14440 | 6.92525 | 16226 | 6.16283 | 18023 | 5.54851 | 19831 | 5.04267 | 47 |
| 14 | 14470 | $6 \cdot 21104$ | 16256 | $6 \cdot 15151$ | 18053 | 5.53927 | 19861 | 5.03499 | 46 |
| 15 | 14499 | 6.89688 | 16286 | 6.14023 | 18083 | 5.53007 | 19891 | 5.02734 | 45 |
| 16 | 14529 | 6.88278 | 16316 | 6.12899 | 18113 | 5.52090 | 19921 | 5.01971 | 44 |
| 17 | 14559 | 6.86874 | 16346 | 6.11779 | 18143 | 5.51176 | 19952 | 5.01210 | 43 |
| 18 | 14588 | 6.85475 | 16376 | 6.10664 | 18173 | 5.50264 | 19982 | 5.0045I | 42 |
| 19 | 14618 | 6.84082 | 16405 | 6.09552 | 1820.3 | 5.40356 | 20012 | $4 \cdot 99695$ | 41 |
| 20 | 14648 | 6.82694 | 16435 | 6.08444 | 18233 | 5.4840]1 | 20042 | $4 \cdot 98940$ | 40 |
| 21 | 14678 | 6.81312 | 16465 | 6.07340 | 18263 | 5.47548 | 20073 | $4 \cdot 98188$ | 39 |
| 22 | 14707 | $6 \cdot 79236$ | 16495 | 6.06240 | 18293 | 5.46648 | 20103 | $4 \cdot 97438$ | 38 |
| 23 | 14737 | $6 \cdot 78564$ | 16525 | $6 \cdot 05143$ | 18323 | 5.45751 | 20133 | 4.96690 | 37 36 |
| 24 | 14767 | $6 \cdot 77199$ | 16555 | 6.04051 | 18353 | 5.44857 | 20164 | $4 \cdot 95945$ | 36 35 |
| 25 | 14796 | 6.75838 | 16585 | 6.02962 | 18383 | 5.43966 | 20194 | $4 \cdot 95201$ | 35 34 |
| 26 | 14826 | $6 \cdot 74483$ | 16615 | 6.01878 | 18414 | 5.43077 | 20224 | 4-9.1460 | 34 33 |
| 27 | 14856 | $6 \cdot 73133$ | 16645 | 6.00797 | 18444 | 5.42192 | 20254 | $4 \cdot 93721$ | 33 32 |
| 28 | 14886 | 6.71789 | 16674 | 5.99720 | 18474 | 5.41309 | 20285 | $4 \cdot 92984$ | 32 31 31 |
| 29 30 | 14915 | 6.70450 | 16704 | $5 \cdot 98646$ | 18504 | 5.40429 | 20315 | $4 \cdot 92249$ | 31 30 |
| 30 | 14945 | 6.69116 | 16734 | $5 \cdot 97576$ | 18534 | 5.39552 | 20345 | $4 \cdot 91516$ | 30 |
| 31 | 14975 | 6.67787 | 16764 | 5.96510 | 18564 | 5.38677 | 20376 | $4 \cdot 90785$ |  |
| 32 | 15005 | 6.66463 | 16794 | 5.95448 | 18504 | 5.37805 | 20406 | $4 \cdot 00056$ | $28^{*}$ |
| 33 | 15034 | 6.65144 | 16824 | 5.94390 | 18624 | $5 \cdot 36936$ | 20436 | 4.89330 | 27 |
| 34 | 15064 | 6.63831 | 16854 | $5 \cdot 93335$ | 18654 | 5.36070 | 20466 | 4.88605 | 26 |
| 35 | 15094 | 6.62523 | 16884 | $5 \cdot 92283$ | 18684 | 5.35206 | 20497 | 4.87882 | 25 |
| 36 | 15124 | 6.61219 | 16914 | 5.91235 | 18714 | 5.34345 | 20527 | 4.87162 | 24 |
| 37 38 | 15153 15183 | 6.59921 | 16944 | $5 \cdot 00191$ | 18745 | 5.33487 | 20557 | 4.86444 | 23 |
| 38 | 15183 | 6.58627 | 16974 | 5.89151 | 18775 | $5 \cdot 32631$ | 20588 | $4 \cdot 85727$ | 22 |
| 39 | 15213 | 6.57339 | 17004 | 5.88114 | 18805 | $5.3177^{8}$ | 20618 | 4.85013 | 21 |
| 40 | 15243 | 6.56055 | 17033 | $5 \cdot 87080$ | 18835 | ${ }^{5} \cdot 30928$ | 20648 | 4.84300 | 20 |
| 41 | 15272 | 6.54777 | 17063 | $5 \cdot 86051$ | 18865 | 5.30080 | 20679 | $4.835{ }^{\circ}$ | 19 |
| 42 | 15302 | 6.53503 | 17093 | 5.85024 | 18895 | 5.29235 | 20709 | 4.82882 | 18 |
| 43 | 15332 | 6.52234 | 17123 | 5.84001 | 18925 | $5 \cdot 28393$ | 20739 | $4 \cdot 82175$ | 17 |
| 44 | 15362 | 6.50970 | 17153 | $5 \cdot 82982$ | 18955 | $5 \cdot 27553$ | 20770 | 4.81471 | 16 |
| 45 | 15391 | 6.49710 | 17183 | 5.81966 | 18986 | $5 \cdot 26715$ | 20800 | 4.80769 | 15 |
| 46 | 15421 | 6.48456 | 17213 | $5 \cdot 80953$ | 19016 | 5.25880 | 20830 | $4 \cdot 80068$ | 14 |
| 47 | 15451 | $6 \cdot 47206$ | 17243 | $5 \cdot 79944$ | 19046 | $5 \cdot 25048$ | 20861 | $4 \cdot 79370$ | I3 |
| 48 | 15481 | $6 \cdot 45961$ | 17273 | $5 \cdot 78938$ | 19076 | 5.24218 | 20891 | $4 \cdot 78673$ | 12 |
| 49 | 15511 | 6.44720 | 17303 | $5 \cdot 77936$ | 19106 | $5 \cdot 23391$ | 20921 | $4 \cdot 77978$ | 11 |
| 50 | 15540 | $6 \cdot 43484$ | 17333 | $5 \cdot 76937$ | 19136 | 5. 22566 | 20952 | $4 \cdot 77286$ | 10 |
| 51 | 15570 | 6.42253 | 17363 | $5 \cdot 75941$ | 19166 | 5.21744 | 20982 | $4 \cdot 76595$ | 8 |
| 52 | 15600 15630 | 6.41026 | 17393 | $5 \cdot 74949$ | 19197 | $5 \cdot 20925$ | 21013 | $4 \cdot 75906$ | 8 |
| 53 54 | 15630 | 6.38804 | 17423 | $5 \cdot 73960$ | 19227 | $5 \cdot 20107$ | 21043 | $4 \cdot 15219$ |  |
| 54 55 | 15660 | $6 \cdot 38587$ | 17453 | 5.72974 | 19257 | $5 \cdot 19293$ | 21073 | $4 \cdot 74534$ | 6 |
| 56 | 15689 15719 | 6.37374 6.36165 | 17483 | $5 \cdot 71992$ $5 \cdot 71013$ | 19287 | 5-18480 | 21104 | $4 \cdot 73851$ | 5 |
| 57 | 15749 | 6.34961 | 17543 | $5 \cdot 70037$ | 19317 | + $\begin{aligned} & 5 \cdot 17671 \\ & 5 \cdot 16863\end{aligned}$ | 21134 | $4 \cdot 73170$ 4.72490 | $\stackrel{4}{3}$ |
| 59 | 15779 | 6.33761 | 17573 | $5 \cdot 69064$ | 19378 | $5 \cdot 16058$ | 21195 | $4 \cdot 71813$ | 2 |
| 59 | 15809 | 6.32566 | 17603 | 5.68094 | 19408 | 5.15256 | 21225 | $4 \cdot 71137$ | 1 |
| 60 | 15838 | 6.31375 | 17633 | 5.67128 | 19438 | 5.14455 | 21256 | $4 \cdot 70403$ | - |
|  | Cotang. | Tangent. | $\overline{\text { Cotang. , }}$ Tangent. |  | Cotang. | Tangent. | Cotang. \| Tangent. |  | 1 |
|  | $81^{\circ}$ |  | $80^{\circ}$ |  | $79^{\circ}$ |  | $78^{\circ}$ |  |  |


| , | $12^{\circ}$ |  | $13^{\circ}$ |  | $14^{\circ}$ |  | $15^{\circ}$ |  | 1 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | T'angent. | Cotarg. | Tangent. | Cotang. | Tangent. | Cotang. | Tangent. | Cotang. |  |
| 0 | 21256 | $4 \cdot 70463$ | 23087 | ¢ 431488 | 24933 | $4 \cdot 01078$ | 26795 | $3 \cdot 73205$ | 60 |
| 1 | 21286 | $4 \cdot 69791$ | 23117 | $4 \cdot 32573$ | 24964 | 4.00582 | 26826 | $3 \cdot 72771$ | 59 |
| 3 | 21316 | 4.69121 | 23148 | $4 \cdot 32001$ | 24995 | $4 \cdot 00086$ | 26857 | $3 \cdot 72338$ | 58 |
| 3 | 21347 | 4.68452 | 23179 | 4.31430 | 25026 | $3 \cdot 9950{ }^{2}$ | 26888 | $3 \cdot 71907$ | 57 |
| 5 | 21377 | 1.67786 | 23209 | $4 \cdot 30860$ | 25056 | $3 \cdot 99099$ | 26920 | 3.71476 | 56 |
| 5 | 21408 | 4.67121 | 23240 | 4.30291 | 25087 | $3 \cdot 98607$ | 2695 I | 3.71046 | 55 |
| 6 | 21438 | 4.66458 | 23271 | 4. 29724 | 25118 | 3.98117 | 26982 | $3 \cdot 70616$ | 54 |
| 7 | 21469 | 4.65797 | 23301 | 4.29159 | 25149 | $3 \cdot 97627$ | 27013 | $3 \cdot 70188$ | 53 |
| 今 | 21499 | $4 \cdot 65138$ | 23332 | 4.28505 | 25180 | $3 \cdot 97139$ | 27044 | $3 \cdot 69761$ | 52 |
| 9 | 21529 | 4.64480 | 23363 | $4 \cdot 28032$ | 25211 | $3 \cdot 96651$ | 27076 | $3 \cdot 69335$ | 51 |
| צо | 21560 | 4.63825 | 23393 | $4 \cdot 27471$ | 25242 | 3.96165 | 27107 | 3.68909 | 50 |
| 11 | 21590 | 4.63171 | 23424 | $4 \cdot 26911$ | 25273 | 3.95680 | 27138 | 3.68485 | 49 |
| 12 | 21621 | 4.62518 | 23455 | $4 \cdot 26352$ | 25304 | $3 \cdot 95196$ | 27169 | $3 \cdot 68061$ | 48 |
| 13 | 21651 | 4.61868 | 23485 | $4 \cdot 25795$ | 25335 | $3 \cdot 94713$ | 27201 | $3 \cdot 67638$ | 47 |
| 14 | 21682 | 4.61219 | 23516 | $4 \cdot 25239$ | 25366 | 3.94232 | 27232 | $3 \cdot 67217$ | 46 |
| 15 | 21712 | $4 \cdot 60572$ | 23547 | $4 \cdot 24685$ | 25397 | 3.93̄51 | 27263 | 3.66796 | 45 |
| 16 | 21743 | 4.59927 | 23578 | 4.24132 | 25428 | $3 \cdot 93271$ | 27294 | 3.66376 | 44 |
| 17 | 21773 | $4 \cdot 59283$ | 23608 | $4 \cdot 23580$ | 25459 | $3 \cdot 92793$ | 27326 | $3 \cdot 65957$ | 43 |
| 18 | 21804 | $4 \cdot 58641$ | 23639 | $4 \cdot 23030$ | 25490 | $3 \cdot 92316$ | 27357 | 3.650338 | 42 |
| 19 | 21834 | $4 \cdot 58001$ | 23670 | 4.22481 | 255521 | $3 \cdot 91839$ | 27388 | $3 \cdot 65121$ | 41 |
| 20 | 21864 | $4 \cdot 57363$ | 23700 | $4 \cdot 21933$ | 25552 | $3 \cdot 91364$ | 27419 | 3.64705 | 40 |
| 21 | 21895 | $4 \cdot 56726$ | 23731 | 4.21387 | 25583 | $3 \cdot 90890$ | 27451 | 3.64289 | 39 |
| 22 | 21925 | $4 \cdot 56091$ | 23762 | $4 \cdot 20842$ | 25014 | 3.90417 | 27482 | 3.63874 | 38 |
| 23 | 21956 | $4 \cdot 55458$ | 23793 | 4-20298 | 25645 | $3 \cdot 89945$ | 27513 | 3.63461 | 37 |
| 24 | 21986 | $4 \cdot 54826$ | 23823 | 4.19756 | 25676 | 3.89474 | 27545 | 3.63048 | 36 |
| 25 | 22017 | 4.54196 | 23854 | 4.19215 | 25707 | $3 \cdot 89004$ | 27576 | 3.62636 | 35 |
| 26 | 22047 | $4 \cdot 53568$ | 23885 | 4.18675 | 25738 | 3.88536 | 27607 | 3.62224 | 3.4 |
| 27 | 22078 | $4 \cdot 52941$ | 23916 | 4.18137 | 25769 | 3.88068 | 27638 | 3.61814 | 33 |
| 28 | 22108 | $4 \cdot 52316$ | 23946 | 4.17600 | 25800 | 3.87601 | 27670 | 3.61405 | 32 |
| 29 | 22139 | $4 \cdot 51693$ | 23977 | $4 \cdot 17064$ | 2583 I | 3.87136 | 27701 | 3.60996 | 31 |
| 30 | 22169 | $4 \cdot 51071$ | 24008 | 4.16530 | 25862 | 3.86671 | $2 \cdot 1732$ | 3.60588 | 30 |
| 31 | 22200 | $4 \cdot 5045 \mathrm{I}$ | 24039 | 4. 15997 | 25893 | 3.86208 | 27764 |  |  |
| 32 | 22231 | $4 \cdot 49832$ | 24069 | 4.15465 | 25924 | 3.85745 | 27795 | $3 \cdot 59775$ | 28 |
| 33 | 22261 | 4.42215 | 24100 | $4 \cdot 14934$ | 25955 | 3.85284 | 27826 | 3.59370 | 27 |
| 34 | 22292 | $4 \cdot 48600$ | 24131 | 4.14405 | 25986 | 3.84824 | 27858 | 3.58966 | 26 |
| 35 | 22322 | $4 \cdot 47986$ | 24162 | $4 \cdot 13877$ | 26017 | 3.84364 | 27889 | $3 \cdot 58562$ | 25 |
| 36 | 22353 | $4 \cdot 47374$ | 24193 | $4 \cdot 13350$ | 26048 | 3.83go6 | 27920 | 3.58160 | 24 |
| 37 | 22383 | $4 \cdot 46764$ | 24223 | 4-12825 | 26079 | 3.83449 | 27952 | $3 \cdot 57758$ | 23 |
| 38 | 22414 | $4 \cdot 46155$ | 24254 | 4-12301 | 26110 | 3.82992 | 27983 | $3 \cdot 57357$ | 22 |
| 39 | 22444 | $4 \cdot 45548$ | 24285 | 4-11778 | 26141 | 3.82537 | 28015 | $3 \cdot 56957$ | 21 |
| 40 | 22475 | 4.44942 | 24316 | 4-11256 | 26172 | $3 \cdot 82083$ | 28046 | $3 \cdot 56557$ | 20 |
| 41 | 22505 | $4 \cdot 44338$ | 24347 | 4-10736 | 26203 | 3.81630 | 28077 | $3 \cdot 56159$ | 19 |
| 42 | 22536 | $4 \cdot 43735$ | 24377 | 4-10216 | 26235 | 3.81177 | 28109 | $3 \cdot 55761$ | 18 |
| 43 | 22567 | $4 \cdot 43134$ | 24408 | $4 \cdot 09699$ | 26266 | 3.80726 | 28140 | $3 \cdot 55364$ | 17 |
| 44 | 22597 | $4 \cdot 42534$ | 24439 | $4 \cdot 09182$ | 26297 | $3 \cdot 80276$ | 28172 | 3.54268 | 16 |
| 45 | 22628 | $4 \cdot 41936$ | 24470 | 4.08666 | 26328 | $3 \cdot 79827$ | 28203 | $3 \cdot 5407^{3}$ | 15 |
| 46 | 22658 | 4.41340 | 24501 | 4-08152 | 26359 | $3 \cdot 79378$ | 28234 | $3 \cdot 54179$ | 14 |
| 47 | 22689 | $4 \cdot 40745$ | 24532 | 4.07639 | 26390 | 3.78931 | 28266 | $3 \cdot 53785$ | 13 |
| 48 | 22719 | $4 \cdot 40152$ | 24562 | 4.07127 | 26421 | $3 \cdot 78485$ | 28297 | $3 \cdot 53393$ | 12 |
| 49 | 22750 | $4 \cdot 30560$ | 24503 | 4.06616 | 26452 | $3 \cdot 78040$ | 28329 | $3 \cdot 53001$ | 11 |
| 50 | 22781 | 4.38069 | 24624 | $4 \cdot 06107$ | 26483 | $3 \cdot 77595$ | 28360 | $3 \cdot 52600$ | 10 |
| 51 | 22811 | $4 \cdot 38381$ | 24655 | $4 \cdot 05509$ | 26515 | 3.77152 | 28391 | 3-52219 | 8 |
| 52 | 22842 | $4 \cdot 37793$ | 24686 | $4 \cdot 05092$ | 26546 | 3.76709 | 28423 | $3 \cdot 51829$ | 8 |
| 53 | 22872 | 4.37207 | 24717 | 4.04586 | 26577 | 3-76268 | 28454 | $3 \cdot 51441$ | 7 |
| 54 | 22903 | 4.36623 | 24747 | 4-0408I | 26608 | $3 \cdot 75828$ | 28486 | $3 \cdot 51053$ | 6 |
| 55 | 22934 | $4 \cdot 36040$ | 24778 | $4 \cdot 03578$ | 26639 | 3-75388 | 28517 | 3-50666 | 5 |
| 56 | 22964 | $4 \cdot 35459$ | 24809 | $4 \cdot 03075$ | 26670 | $3 \cdot 74950$ | 28549 | 3-50279 | 4 |
| 57 | 22995 | $4 \cdot 34879$ | 24840 | $4 \cdot 02574$ | 26701 | 3.74512 | 28580 | 3-49894 | 3 |
| 58 | 23026 | 4.34300 | 24871 | 4.02074 | 26733 | $3 \cdot 74075$ | 28612 | $3 \cdot 49509$ | 2 |
| 59 | 2.3056 | 4.33723 | 24902 | $4 \cdot 01576$ | 26764 | $3 \cdot 73640$ | 28643 | 3.49125 | * |
| 6o | 23087 | 4.33148 | 24933 | $4 \cdot 01078$ | 26795 | 73205 | 28675 | 3.48741 | $\checkmark$ |
| , | Cotang. | Tangent. | Cotang. Tangent. |  | Cotang. Tangent. |  | Cotang. Tangent. |  | 1 |
|  | $77^{\circ}$ |  | $75^{\circ}$ |  | $75^{\circ}$ |  | $74^{\circ}$ |  |  |

\begin{tabular}{|c|c|c|c|c|c|c|c|c|c|}
\hline \multicolumn{2}{|l|}{Table III.} \& \multicolumn{7}{|r|}{NATURAL TANGENTS AND COTANGENTS.} \& 77 \\
\hline \multirow[t]{2}{*}{} \& \multicolumn{2}{|c|}{\(16^{\circ}\)} \& \multicolumn{2}{|r|}{\(17^{\circ}\)} \& \multicolumn{2}{|r|}{\(18^{\circ}\)} \& \multicolumn{2}{|l|}{\(19^{\circ}\)} \& \multirow[t]{2}{*}{,} \\
\hline \& Tangent. \& Cotang. \& Tangent. \& Cotang. \& Tangent. \& Cotang. \& Tangent. \& Cotang. \& \\
\hline 0 \& 28675 \& \(3 \cdot 48741\) \& 30573 \& \(3 \cdot 27085\) \& 32492 \& 3.07768 \& 34433 \& \(2 \cdot 90421\) \& 60 \\
\hline , \& 28706 \& 3-48359 \& 30605 \& \(3 \cdot 26745\) \& 32524 \& 3.07464 \& 34465 \& \(2 \cdot 00147\) \& 59 \\
\hline 2 \& 28738 \& \(3 \cdot 47977\) \& 30637 \& 3-26406 \& 32556 \& 3.07160 \& 34498 \& 2.89873 \& 58 \\
\hline 3 \& 28769 \& \(3 \cdot 47596\) \& 30669 \& 3-26067 \& 32588 \& 3.06857 \& 34530 \& 2.89600 \& 57 \\
\hline 4 \& 28800 \& \(3 \cdot 47216\) \& 30700 \& \(3 \cdot 25729\) \& 32621 \& 3.06554 \& 34563 \& 2.89327 \& 56 \\
\hline 5 \& 28832 \& \(3 \cdot 46837\) \& 30732 \& 3-25392 \& 32653 \& 3.06252 \& 34596 \& 2.89055 \& 55 \\
\hline 6 \& 28864 \& 3-46458 \& 30764 \& \(3 \cdot 25055\) \& 3.2685 \& 3.05950 \& 34628 \& 2.88783 \& 54 \\
\hline 7 \& 28895 \& 3-46080 \& 30796 \& 3.24719 \& 32717 \& 3.05649 \& 34661 \& \(2 \cdot 88511\) \& 53 \\
\hline 8 \& 28927 \& \(3 \cdot 45703\) \& 30828 \& 3.24383 \& 32749 \& 3.053和 \& 34693 \& 2.88240 \& 52 \\
\hline 9 \& 28958 \& \(3 \cdot 45327\) \& 30860 \& \(3 \cdot 24049\) \& 32782 \& \(3 \cdot 05049\) \& 34726 \& 2.87970 \& 51 \\
\hline 10 \& 28990 \& \(3 \cdot 44951\) \& 30891 \& \(3 \cdot 23714\) \& 32814 \& \(3 \cdot 04749\) \& 34758 \& 2.87700 \& 50 \\
\hline 11 \& 29021 \& \(3 \cdot 44576\) \& 30923 \& 3-23381 \& 32846 \& \(3 \cdot 04450\) \& 34791 \& 2.87430 \& 49 \\
\hline 12 \& 20053 \& 3-44202 \& 30955 \& 3.23048 \& 32878 \& \(3 \cdot 04152\) \& 34824 \& 2.87161 \& 48 \\
\hline 13 \& 29084 \& \(3 \cdot 43829\) \& 30987 \& 3.22715 \& 32911 \& 3.03854 \& 34856 \& 2.86892 \& 47 \\
\hline 14 \& 29116 \& \(3 \cdot 43456\) \& 31019 \& \(3 \cdot 22384\) \& 32943 \& 3.03556 \& 34889 \& 2.86624 \& 46 \\
\hline 15 \& 29147 \& 3.43084 \& 31051 \& \(3 \cdot 22053\) \& 32975 \& 3.03260 \& 34922 \& 2.86356 \& 45 \\
\hline 16 \& 29179 \& \(3 \cdot 42713\) \& 31083 \& \(3 \cdot 21722\) \& 33007 \& 3.02963 \& 34954 \& \(2 \cdot 86089\) \& 44 \\
\hline 17 \& 29210 \& \(3 \cdot 42343\) \& 31115 \& \(3 \cdot 21392\) \& 33040 \& 3.02667 \& 34987 \& 2.85822 \& 43 \\
\hline 18 \& 29242 \& \(3 \cdot 41973\) \& 31147 \& 3-21063 \& 33072 \& 3.02372 \& 35019 \& 2.85555 \& 42 \\
\hline 19 \& 29274 \& 3.41604
3.41236 \& 31178 \& \(3 \cdot 20734\) \& 33104
33126 \& 3.02077 \& 35052 \& 2.85289 \& 41 \\
\hline 20 \& 29305 \& 3.41236 \& 31210 \& \(3 \cdot 20406\) \& 33136 \& \(3 \cdot 01783\) \& 35085 \& 2.85023 \& 40 \\
\hline 21 \& 29337 \& \(3 \cdot 40869\) \& 31242 \& \(3 \cdot 20079\) \& 33169 \& \(3 \cdot 01489\) \& 35117 \& \(2 \cdot 84758\) \& 39
3 \\
\hline 22 \& \(29^{368}\) \& 3-40502 \& 31274 \& \(3 \cdot 19752\) \& 33201 \& \(3 \cdot 01196\) \& 35150 \& 2.84494 \& 38 \\
\hline 23 \& 29400 \& 3-40136 \& 31306 \& 3.19426 \& 33233 \& 3-00903 \& 35183 \& 2.84229 \& 37 \\
\hline 24 \& 29432 \& \(3 \cdot 39771\) \& 31338 \& \(3 \cdot 19100\) \& 33266 \& \(3 \cdot 00611\) \& 35216 \& 2.83965 \& 36 \\
\hline 25 \& 29463 \& \(3 \cdot 39406\)
\(3 \cdot 39042\) \& 31370
31402 \& \(3 \cdot 18775\)
\(3 \cdot 18451\) \& 33298
33330 \& \begin{tabular}{l}
\(3 \cdot 00319\) \\
\(3 \cdot 00028\) \\
\hline
\end{tabular} \& 35248
3528 I

l \& 2.83702
2.8343 \& 35
34 <br>
\hline 27 \& 29526 \& 3.38679 \& 31434 \& 3-18127 \& 33363 \& $2 \cdot 99738$ \& 35314 \& 2.83176 \& 33 <br>
\hline 28 \& 29558 \& 3.38317 \& 31466 \& $3 \cdot 17804$ \& 33395 \& $2 \cdot 99447$ \& 35346 \& 2.82914 \& 32 <br>
\hline 29 \& 29590 \& $3 \cdot 37955$ \& 31498 \& 3-17481 \& 33427 \& $2 \cdot 99158$ \& 353379 \& 2.82653 \& 3 I <br>
\hline 30 \& 29621 \& 3.37594 \& 31530 \& 3.17159 \& 33460 \& $2 \cdot 98868$ \& 35412 \& 2.82391 \& 30 <br>
\hline 31 \& 29653 \& $3 \cdot 37234$ \& 31562 \& 3.16838 \& 33492 \& $2 \cdot 98580$ \& 35445 \& 2.82130 \& 29 <br>
\hline 32 \& 29685 \& $3 \cdot 36875$ \& 31594 \& 3-16517 \& 33524 \& $2 \cdot 98292$ \& 33577 \& 2.81870 \& 28 <br>
\hline 33 \& 29716 \& 3.36516 \& 31626 \& 3-16197 \& 33557 \& $2 \cdot 98004$ \& 35510 \& 2.81610 \& 27 <br>
\hline 34 \& 29748 \& 3.36158 \& 31658 \& 3.15877 \& 33589 \& $2 \cdot 97717$ \& 35543 \& 2.8135o \& 26 <br>
\hline 35 \& 29780 \& 3.35800 \& 31690 \& $3 \cdot 15558$ \& 3362 I \& $2 \cdot 97430$ \& 35576 \& 2.81091 \& 25 <br>
\hline 36 \& 29811 \& 3.35443 \& 31722 \& $3 \cdot 15240$ \& 33654 \& $2 \cdot 97144$ \& 35608 \& 2.80833 \& 24 <br>
\hline 37
38 \& 29843 \& 3.35087 \& 31754 \& $3 \cdot 14922$ \& 33686 \& 2.96858 \& 35641 \& $2 \cdot 80534$ \& 23 <br>
\hline 38 \& 29875 \& $3 \cdot 34732$ \& 31786
31818 \& 3.14605 \& 33718 \& $2 \cdot 96573$ \& 35674 \& $2 \cdot 80316$ \& 22 <br>
\hline 39 \& 29906 \& $3 \cdot 34377$ \& 31818 \& 3.14288 \& 33751 \& $2 \cdot 96288$ \& 35707 \& $2 \cdot 80059$ \& 21 <br>
\hline 40 \& 29938 \& $3 \cdot 34023$ \& 31850 \& 3.13972 \& 33783 \& $2 \cdot 96004$ \& 35740 \& $2 \cdot 79802$ \& 20 <br>
\hline 41 \& 29970 \& 3.33670 \& 31882 \& 3.13656 \& 33816 \& $2 \cdot 95721$ \& 35772 \& $2 \cdot 79545$ \& 19 <br>
\hline 42 \& 30001 \& ${ }^{3.3331} 17$ \& 31914 \& 3.13341 \& 33848 \& $2 \cdot 95437$ \& 35805 \& 2.79289 \& 18 <br>
\hline 43 \& 30033 \& 3.32965 \& 31946 \& 3.13027 \& 33881 \& $2 \cdot 95155$ \& 35838 \& $2 \cdot 79033$ \& 17 <br>
\hline 44 \& 30065 \& 3.32614 \& 31978 \& $3 \cdot 12713$ \& 33913 \& $2 \cdot 94872$ \& 35871 \& $2 \cdot 78778$ \& 16 <br>
\hline 40 \& 30097 \& $3 \cdot 32264$ \& 32010 \& 3.12400 \& 33945 \& $2 \cdot 94590$ \& 35904 \& $2 \cdot 78523$ \& 15 <br>
\hline 46 \& 30128 \& 3.31914 \& 32042 \& $3 \cdot 12087$ \& 33978 \& $2 \cdot 94309$ \& 35937 \& $2 \cdot 78269$ \& 14 <br>
\hline 47 \& 30160 \& 3.31565 \& 32074 \& 3.11775 \& 34010 \& $2 \cdot 94028$ \& 35969 \& $2 \cdot 78014$ \& 13 <br>
\hline 48 \& 30192 \& $3 \cdot 31216$ \& 32106 \& 3.11464 \& 34043 \& $2 \cdot 93748$ \& 36002 \& $2 \cdot 77761$ \& 12 <br>
\hline 49 \& 30224 \& $3 \cdot 30868$ \& 32 I 39 \& $3 \cdot 11153$ \& 34075 \& $2 \cdot 93468$ \& 36035 \& $2 \cdot 77507$ \& 11 <br>
\hline 50 \& 30255 \& 3.30521 \& 32171 \& 3.10842 \& 34108 \& $2 \cdot 93189$ \& 36068 \& $2 \cdot 77254$ \& 10 <br>
\hline \& 30287 \& 3.30174 \& 32203 \& 3.10532 \& 34140 \& $2 \cdot 92910$ \& 36101 \& $2 \cdot 77002$ \& 9 <br>
\hline 52
53

5 \& | 30319 |
| :--- |
| 30351 | \& $3 \cdot 29829$

$3 \cdot 29483$
3 \& 32235

32267 \& | $3 \cdot 10223$ |
| :--- |
| $3 \cdot 09914$ | \& 34173

34205 \& 2.92632
2.92354
2.9207 \& 36134
36167 \& $2 \cdot 76750$
$2 \cdot 76498$ \& 8 <br>
\hline 54 \& 30382 \& $3 \cdot 29139$ \& 32299 \& $3 \cdot 09606$ \& 34238 \& $2 \cdot 92076$ \& 36199 \& $2 \cdot 76247$ \& 6 <br>
\hline 55 \& 30414 \& 3.28795 \& 32331 \& $3 \cdot 09298$ \& 34270 \& $2 \cdot 91799$ \& 36232 \& $2 \cdot 75996$ \& 5 <br>
\hline 56 \& 30446 \& $3 \cdot 28452$ \& 32363 \& 3.08991 \& 34303 \& $2 \cdot 91523$ \& 36265 \& $2 \cdot 75746$ \& 4 <br>
\hline 57
58 \& 30478 \& $3 \cdot 28109$ \& 32396 \& 3.08685 \& 34335 \& $2 \cdot 91246$ \& 36298 \& 2.75496 \& 3 <br>
\hline 58 \& 30509 \& $3 \cdot 27767$ \& 32428 \& 3.08379 \& 34368 \& $2 \cdot 90971$ \& 36331 \& $2 \cdot 75246$ \& 2 <br>
\hline 50 \& 30541
30573 \& $3 \cdot 27426$
$3 \cdot 27085$ \& 32460
32492 \& 3.08073
3.07768 \& 34400
34433 \& $2 \cdot 90696$ \& 36364 \& $2 \cdot 74997$ \& 1 <br>
\hline \& \& 3 \& 32492 \& 3.07768 \& 34433 \& $2 \cdot 90421$ \& 36397 \& $2 \cdot 74748$ \& 0 <br>
\hline \multirow[t]{2}{*}{} \& Cotang. \& Tangent. \& Cotang. \& Tangent. \& Cotang. \& Tangent. \& Cotang. \& Tangent. \& <br>
\hline \& \multicolumn{2}{|r|}{$73^{\circ}$} \& \multicolumn{2}{|r|}{$72^{\circ}$} \& \multicolumn{2}{|r|}{$71^{\circ}$} \& \multicolumn{2}{|r|}{$50^{\circ}$} \& <br>
\hline
\end{tabular}

| 78 | NATURAI |  | NG | TS | D C |  |  | Table 1 II. |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $20^{\circ}$ |  | $21^{\circ}$ |  | $22^{\circ}$ |  | $23^{\circ}$ |  | , |
|  | Tangent. | Cotang. | Tangent. | Cotang. | Tangent. | Cotang. | Tangent. | Cotang. |  |
| o | 36397 | $2 \cdot 74748$ | 38386 | 2.60509 | 40403 | 2.47509 | 42447 | $2 \cdot 35585$ | 60 |
| 1 | 36430 | $2 \cdot 74499$ | 38420 | 2.60283 | 40436 | 2.47302 | 42482 | 2.35395 | 59 |
| 2 | 36463 | 2.74251 | 38453 | 2.60057 | 40470 | $2 \cdot 47095$ | 42516 | $2 \cdot 35205$ | 58 |
| 3 | 36496 | $2 \cdot 74004$ | 38487 | $2 \cdot 5983 \mathrm{I}$ | 40504 | $2 \cdot 46888$ | 42551 | $2 \cdot 35015$ | 57 |
| 4 | 36529 | $2 \cdot 73756$ | 38520 | $2 \cdot 59606$ | 40538 | $2 \cdot 46682$ | 42585 | $2 \cdot 34825$ | 56 |
| 5 | 36562 | $2 \cdot 735 \mathrm{co}$ | 38553 | $2 \cdot 59381$ | 40572 | $2 \cdot 46476$ | 42619 | 2.34636 | 55 |
| 6 | 36595 | $2 \cdot 73263$ | 38587 | 2.59156 | 40606 | 2.46270 | 42654 | $2 \cdot 34447$ | 54 |
| 7 | 36628 | $2 \cdot 73017$ | 38620 | 2.58932 | 40640 | $2 \cdot 46065$ | 42688 | $2 \cdot 34258$ | 53 |
| 8 | 36661 | $2 \cdot 72771$ | 38654 | 2.58708 | 40674 | $2 \cdot 45860$ | 42722 | $2 \cdot 34069$ | 52 |
|  | 36694 | $2 \cdot 72526$ | 38687 | 2.58484 | 40707 | $2 \cdot 45555$ | 42757 | $2 \cdot 33881$ | 51 |
| 10 | 36727 | $2 \cdot 72281$ | 38721 | $2 \cdot 58261$ | 40741 | $2 \cdot 45451$ | 42791 | $2 \cdot 33693$ | 50 |
| 11 | 36760 | $2 \cdot 72036$ | 38754 | 2.58038 | 40775 | $2 \cdot 45246$ | 42826 | $2 \cdot 33505$ | 49 |
| 12 | 36793 | $2 \cdot 71792$ | 38787 | 2.57815 | 40809 | $2 \cdot 45043$ | 42860 | $2 \cdot 33317$ | 48 |
| 13 | 36826 | $2 \cdot 71548$ | 38821 | 2.57593 | 40843 | 2.44839 | 42894 | $2 \cdot 33130$ | 47 |
| 14 | 36859 | $2 \cdot 71305$ | 38854 | $2 \cdot 57371$ | 40877 | 2.44636 | 42929 | 2.32943 | 46 |
| 15 | 36892 | $2 \cdot 71062$ | 38888 | 2.57150 | 40911 | $2 \cdot 44433$ | 42963 | $2 \cdot 32756$ | 45 |
| 16 | 36925 | $2 \cdot 70819$ | 38921 | 2.56928 | 40945 | 2.44230 | 42998 | $2 \cdot 32570$ | 44 |
| 17 | 36958 | $2 \cdot 70577$ | 38955 | 2.56707 | 40979 | $2 \cdot 44027$ | 43032 | $2 \cdot 32383$ | 43 |
| 18 | 36991 | $2 \cdot 70335$ | 38988 | 2.56487 | 41013 | $2 \cdot 43825$ | 43067 | $2 \cdot 32197$ | 42 |
| 19 | 37024 | $2 \cdot 70094$ | 39022 | 2.56265 | 41047 | $2 \cdot 43623$ | 43101 | $2 \cdot 32012$ | 41 |
| 20 | 37057 | $2 \cdot 69853$ | 39055 | 2.56046 | 41081 | $2 \cdot 43422$ | 43136 | $2 \cdot 31826$ | 40 |
| 21 | 37090 | 2.69612 | 39089 | 2.55827 | 41115 | 2.43220 | 43170 | $2 \cdot 31641$ | 39 |
| 22 | 37124 | $2 \cdot 69371$ | 39122 | 2.55608 | 41149 | $2 \cdot 43019$ | $\dot{4}^{4} 205$ | $2 \cdot 31456$ | 38 |
| 23 | 37157 | $2 \cdot 69131$ | 39156 | 2.55389 | 41183 | 2.42819 | 43239 | $2 \cdot 31271$ | , |
| 24 | 37190 | 2.68892 | 39190 | 2.55170 | 41217 | $2 \cdot 42618$ | 43274 | $2 \cdot 31086$ | 36 |
| 25 | 37223 | 2.68653 | 39223 | 2.54952 | 41251 | $2 \cdot 42418$ | 43308 | $2 \cdot 30902$ | 35 |
| 26 | 37256 | 2.68414 | 39257 | 2.54734 | 41285 | 2.42218 | 43343 | $2 \cdot 30718$ | 34 |
| 27 | 37289 | $2 \cdot 68175$ | 39290 | 2.54516 | 41319 | $2 \cdot 42019$ | 43378 | $2 \cdot 30534$ | 33 |
| 28 | 37322 | $2.679^{3} 7$ | 39324 | 2.54299 | 41353 | 2.41819 | 43412 | 2.3035ı | 32 |
| 29 | 37355 | 2.67700 | 39357 | 2.54082 | 41387 | 2.41620 | 43447 | $2 \cdot 30167$ | 31 |
| 30 | 37388 | 2.67462 | 39391 | 2.53865 | 41421 | 2.41421 | 43481 | 2. 29984 | 30 |
| 31 | 37422 | $2 \cdot 67225$ | 39425 | 2.53648 | 41455 | $2 \cdot 41223$ | 43516 | 2.29801 | 29 |
| 32 | 37455 | 2.66989 | 39458 | 2.53432 | 41490 | $2 \cdot 41025$ | 43550 | 2.29619 | 28 |
| 33 | 37488 | 2.66752 | 39492 | 2.53217 | 41524 | $2 \cdot 40827$ | 43585 | $2 \cdot 29437$ | 27 |
| 34 | 37521 | $2 \cdot 66516$ | 39526 | 2.53001 | 41558 | 2.40629 | 43620 | $2 \cdot 29254$ | 26 |
| 35 | 37554 | $2 \cdot 6628 \mathrm{I}$ | 39559 | 2.52786 | 41592 | $2 \cdot 40432$ | 43654 | $2 \cdot 29073$ | 25 |
| 36 | 37588 | 2.66046 | 39593 | 2.52571 | 41626 | $2 \cdot 40235$ | 43689 | $2 \cdot 28891$ | 24 |
| 37 | 37621 | $2 \cdot 65811$ | 39626 | 2.52357 | 41660 | $2 \cdot 40038$ | 43724 | $2 \cdot 28710$ | 23 |
| 38 | 37654 | 2.65576 | 39660 | 2.52142 | 41694 | $2 \cdot 39841$ | 43758 | $2 \cdot 28528$ | 22 |
| 39 | 37687 | 2.65342 | 39694 | 2.51929 | 41728 | $2 \cdot 39645$ | 43793 | 2.28348 | 21 |
| 40 | 37720 | 2.65109 | 39727 | 2.51715 | 41763 | $2 \cdot 39449$ | 43828 | $2 \cdot 28107$ | 20 |
| 41 | 37754 | $2 \cdot 64875$ | 39761 | 2.51502 | 41797 | $2 \cdot 39253$ | 43862 | $2 \cdot 27987$ | 19 |
| 42 | 37787 | 2.64642 | 39795 | 2.51289 | 41831 | $2 \cdot 39058$ | 43897 | $2 \cdot 27806$ | 18 |
| 43 | 37820 | 2.64410 | 39829 | 2.51076 | 41865 | 2.38862 | 43932 | $2 \cdot 276.26$ | 17 |
| 44 | 37853 | 2.64177 | 39862 | $2 \cdot 50864$ | 41899 | $2 \cdot 38668$ | 43966 | $2 \cdot 27447$ | 16 |
| 45 | 37887 | 2.63945 | 39896 | $2 \cdot 50652$ | 41933 | $2 \cdot 38473$ | 44001 | $2 \cdot 27267$ | 15 |
| 46 | 37920 | 2.63714 | 39930 | 2.50440 | 41968 | $2 \cdot 38279$ | 44036 | $2 \cdot 27088$ | 14 |
| 47 | 37953 | 2.63483 | 39963 | 2.50229 | 42002 | $2 \cdot 38084$ | 44071 | $2 \cdot 26909$ | 13 |
| 48 | 37986 | 2.63252 | 39997 | $2 \cdot 50018$ | 42036 | $2 \cdot 3789 \mathrm{I}$ | 44105 | $2 \cdot 26730$ | 12 |
| 49 | 38020 | 2.63021 | 40031 | $2 \cdot 49807$ | 42070 | $2 \cdot 37697$ | 44140 | $2 \cdot 26552$ | 11 |
| 50 | 38053 | $2 \cdot 62791$ | 40065 | $2 \cdot 49597$ | 42105 | $2 \cdot 37504$ | 44175 | 2.26374 | 10 |
| 51 | 38086 | 2.62561 | 40098 | 2.49386 | 42139 | $2 \cdot 37311$ | 44210 | $2 \cdot 26196$ | 9 |
| 52 | 38120 | 2.62332 | 40132 | $2 \cdot 49177$ | $4217^{3}$ | 2.37118 | 44244 | $2 \cdot 26018$ | 8 |
| 53 | 38153 | 2.62103 | 40166 | $2 \cdot 48967$ | 42207 | $2 \cdot 36925$ | 44279 | 2.25840 | 7 |
| 54 | 38186 | 2.61874 | 40200 | $2 \cdot 48758$ | 42242 | 2.36733 | 44314 | $2 \cdot 25663$ | 6 |
| 55 | 38220 | 2.61646 | 40234 | $2 \cdot 48549$ | 42276 | 2.36541 | 44349 | $2 \cdot 25486$ | 5 |
| 56 | 38253 38286 | 2.61418 | 40267 | 2.48340 2.48 .32 | 42310 | 2.36349 2.36158 2.359 | 44384 | $2 \cdot 25309$ $2 \cdot 25132$ | 4 3 3 |
| 57 58 58 | 38286 38320 | 2.61190 2.60963 | 40301 | 2.48132 | 42345 | 2.36158 2.35067 2.3576 | 44418 | $2 \cdot 25132$ <br> $2 \cdot 24056$ <br> 2.2450 | 3 2 |
| 58 59 | 38320 38353 | 2.60963 2.60736 | 40335 | 2.47924 2.47716 | 42379 42413 | $2 \cdot 3$ 2.3 | 44453 | 2.24956 | 2 |
| 60 | 38386 | 2.60509 | 4 4 403 | 2.47509 | 4244 | 2.35585 | 44523 | $2 \cdot 2460$ | 0 |
|  | Cotang. | Tangent. | Cotang. | Tangent. | Cotang. | Tangent. | Cotang. |  |  |
|  | $69^{\circ}$ |  | $68^{\circ}$ |  | $67^{\circ}$ |  | $66^{\circ}$ |  |  |


| Table III. |  | NATURAL TANGENTS AND COTANGENTS. |  |  |  |  |  |  | 79 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | $24^{\circ}$ |  | $25^{\circ}$ |  | $26^{\circ}$ |  | $27^{\circ}$ |  | , |
|  | Tangent. | Cotang. | Tangent. | Cotang. | Tangent. | Cotang. | Tangent. | Cotang. |  |
| 0 | 44523 | $2 \cdot 24604$ | 46631 | 2.14451 | 48773 | 2.05030 | 50953 | $1 \cdot 96261$ | 60 |
| 1 | 44558 | $2 \cdot 24428$ | 46666 | 2.14288 | 48809 | $2 \cdot 04879$ | 50989 | 1.96120 | 59 |
| 2 | 44593 | $2 \cdot 24252$ | 46702 | $2 \cdot 14125$ | 48845 | $2 \cdot 04728$ | 51026 | I $\cdot 75979$ | 58 |
| 3 | 44627 | $2 \cdot 24077$ | 46737 | 2.13963 | 48881 | 2.04577 | 51063 | $1 \cdot 95838$ | 57 |
| 4 | 44662 | $2 \cdot 23902$ | 46772 | 2.13801 | 48917 | $2 \cdot 04426$ | 51099 | $1 \cdot 95698$ | 56 |
| 5 | 44697 | $2 \cdot 23727$ | 40808 | 2.13639 | 48953 | $2 \cdot 04276$ | 51136 | $1 \cdot 95557$ | 55 |
| 6 | 44732 | $2 \cdot 23553$ | 46843 | 2.13477 | 48989 | $2 \cdot 04125$ | 51173 | $1 \cdot 95417$ | 54 |
| 7 | 44767 | $2 \cdot 23378$ | 46879 | $2 \cdot 13316$ | 49026 | $2 \cdot 03975$ | 51209 | 1-95277 | 53 |
| 8 | 44802 | $2 \cdot 23204$ | 46914 | 2.13154 | 49062 | 2.03825 | 51246 | 1.95137 | 52 |
| 9 | 44837 | $2 \cdot 23030$ | 46950 | 2.12993 | 49098 | 2.03675 | 51283 | 1-94997 | 51 |
| 10 | 44872 | $2 \cdot 22857$ | 46985 | 2.12832 | 49134 | $2 \cdot 03526$ | 51319 | 1.94858 | 50 |
| 11 | 44907 | $2 \cdot 22683$ | 47021 | $2 \cdot 12671$ | 49170 | $2 \cdot 03376$ | 51356 | 1.94718 | 42 |
| 12 | 44942 | $2 \cdot 22510$ | 47 C 56 | 2.12511 | 49206 | 2.03227 | 51393 | $1 \cdot 94579$ | 48 |
| 13 | 44977 | $2 \cdot 22337$ | 47092 | 2.12350 | $49^{242}$ | 2.03078 | 51430 | $1 \cdot 94440$ | 47 |
| 14 | 45012 | $2 \cdot 22164$ | 47128 | 2.12190 | 49278 | 2.02929 | 51467 | $1 \cdot 94301$ | 46 |
| 15 | 45047 | 2.21992 | 47163 | 2-12030 | 49315 | 2.02780 | 51503 | 1.94162 | 45 |
| 16 | 45082 | 2-21819 | 47199 | $2 \cdot 11871$ | 49351 | $2 \cdot 02631$ | 51540 | 1.94023 | 44 |
| 17 | 45117 | $2 \cdot 21647$ | 47234 | $2 \cdot 11711$ | 49387 | $2 \cdot 02483$ | 51577 | $1 \cdot 93885$ | 43 |
| 18 | 45152 | $2 \cdot 21475$ | 47270 | $2 \cdot 11552$ | 40423 | 2.02335 | 51614 | $1 \cdot 93746$ | 42 |
| 19 | 45187 | $2 \cdot 21304$ | 47305 | 2.11392 | 49459 | 2.02187 | 51651 | $1 \cdot 93608$ | 41 |
| 20 | 45222 | 2.21132 | 47341 | $2 \cdot 11233$ | 49495 | 2.02039 | 51688 | 1-93470 | 40 |
| 21 | 45257 | $2 \cdot 20961$ | 47377 | 2-11075 | 49532 | $2 \cdot 01891$ | 51724 | 1-93332 | 39 |
| 22 | 45292 | $2 \cdot 20790$ | 47412 | $2 \cdot 10916$ | 49568 | $2 \cdot 01743$ | 51761 | $1 \cdot 93195$ | 38 |
| 23 | 45327 | $2 \cdot 20619$ | 47443 | $2 \cdot 10758$ | 49604 | $2 \cdot 01596$ | 51798 | $1 \cdot 93057$ | 37 |
| 24 | 45362 | $2 \cdot 20449$ | 47483 | $2 \cdot 10600$ | 49640 | $2 \cdot 01449$ | 51835 | $1 \cdot 92920$ | 36 |
| 25 | 45397 | $2 \cdot 20278$ | 47519 | 2.10442 | 49677 | $2 \cdot 01302$ | 51872 | $1 \cdot 92782$ | 35 |
| 26 | 45432 | $2 \cdot 20108$ | 47555 | 2.10284 | 49713 | $2 \cdot 01155$ | 51909 | $1 \cdot 92645$ | 34 |
| 27 | 45467 | $2 \cdot 19938$ | 47590 | 2.10126 | 49749 | $2 \cdot 01008$ | 51946 | $1 \cdot 92508$ | 33 |
| 28 | 45502 | $2 \cdot 19769$ | 47626 | $2 \cdot 09969$ | 49786 | 2-00862 | 51983 | $1 \cdot 92371$ | 32 |
| 29 | 45537 | $2 \cdot 19599$ | 47662 | $2 \cdot 09811$ | 49822 | $2 \cdot 00715$ | 52020 | $1 \cdot 92235$ | 31 |
| 30 | 45573 | $2 \cdot 19430$ | 47698 | $2 \cdot 09654$ | 49858 | 2.00569 | 52057 | 1-92098 | 30 |
| 31 | 45608 | 2.19261 | 47733 | 2.09498 | 49894 | $2 \cdot 00423$ | 52094 | $1 \cdot 91962$ | 29 |
| 32 | 45643 | $2 \cdot 110092$ | 47769 | $2 \cdot 09341$ | 49931 | $2 \cdot 00277$ | 52131 | $1 \cdot 91826$ | 28 |
| 33 | 45678 | 2.18923 | 47805 | $2 \cdot 09184$ | 49967 | 2-00131 | 52168 | 1.91690 | 27 |
| 34 | 45713 | 2.18755 | 47840 | $2 \cdot 09028$ | 50004 | 1-90986 | 52205 | $1 \cdot 91554$ | 26 |
| 35 | 45748 | 2.18587 | 47876 | $2 \cdot 08872$ | 50040 | 1-99841 | 52242 | $1 \cdot 91418$ | 25 |
| 36 | 45784 | 2.18419 | 47912 | $2 \cdot 08716$ | 50076 | 1.99695 | 52279 | 1-91282 | 24 |
| 37 | 45819 | 2.18251 | 47948 | $2 \cdot 08560$ | 50113 | 1-99550 | 52316 | $1 \cdot 91147$ | 23 |
| 38 | 45854 | 2.18084 | 47984 | $2 \cdot 08405$ | 50149 | 1-99406 | 52353 | $1 \cdot 91012$ | 22 |
| 39 | 45889 | $2 \cdot 17916$ | 48019 | $2 \cdot 08250$ | 50185 | 1-99261 | 52390 | $1 \cdot 90876$ | 21 |
| 40 | 45924 | 2.17749 | 48055 | $2 \cdot 08094$ | 50222 | $1 \cdot 99116$ | 52427 | $1 \cdot 90741$ | 20 |
| 41 | 45960 | $2 \cdot 17582$ | 48091 | $2 \cdot 07939$ | 50258 | 1-98972 | 52464 | $1 \cdot 90607$ | 19 |
| 42 | 45995 | $2 \cdot 17416$ | 48127 | 2.07785 | 50295 | 1-98828 | 52501 | 1.90472 | 18 |
| 43 | 46030 | $2 \cdot 17249$ | 48163 | 2.07630 | 50331 | 1-98684 | 52538 | $1 \cdot 90337$ | 17 |
| 44 | 46065 | $2 \cdot 17083$ | 48198 | 2.07476 | 50368 | 1-98540 | 52575 | 1-90203 | 16 |
| 45 | 46101 | $2 \cdot 16917$ | 43234 | $2 \cdot 07321$ | 50404 | 1.98396 | 52613 | $1 \cdot 90069$ | 15 |
| 46 | 46.36 | $2 \cdot 16751$ | 48270 | $2 \cdot 07167$ | 50441 | $1 \cdot 98253$ | 52650 | 1.89935 | 14 |
| 47 | 46171 | 2. 16585 | 48306 | $2 \cdot 07014$ | 50477 | $1 \cdot 98110$ | 52687 | 1.89801 | 13 |
| 48 | 46306 | 2.16420 | 48342 | $2 \cdot 06860$ | 50514 | $1 \cdot 97956$ | 52724 | 1.89667 | 12 |
| 49 | 46242 | 2.16255 | 48378 | 2-06706 | 505jo | 1-97823 | 52761 | 1-89533 | 11 |
| 50 | 46277 | 2-16090 | 48414 | 2.06553 | 50587 | 1-97680 | 52708 | 1.89400 | 10 |
| 51 | 46312 | 2.15925 | 48450 | $2 \cdot 06400$ | 50623 | 1-97538 | 52836 | 1.89266 | 9 |
| 52 | 46348 | $2 \cdot 15760$ | 48486 | $2 \cdot 06247$ | 50660 | 1-97395 | 52873 | 1.89133 | 8 |
| 53 | 46383 | $2 \cdot 15596$ | 48521 | $2 \cdot 06094$ | 50696 | 1.97253 | 52910 | 1.88000 | 7 |
| 54 | 46418 | $2 \cdot 15432$ | 48557 | 2.05942 | 50733 | 1.97111 | 52947 | 1.88867 | 6 |
| 55 56 | 46454 | 2-15268 | 48593 | $2 \cdot 05790$ | 50769 | 1.96969 | 52984 | 1.88734 | 5 |
| 56 | 46489 46525 | $2 \cdot 15104$ $2 \cdot 14940$ | 48629 48665 | $2 \cdot 05637$ $2 \cdot 05485$ | 50806 50843 | 1.96827 1. 06685 | 53022 53050 | 1.88602 1.88460 | 4 |
| 58 | 46560 | 2.14940 2.14777 | 48701 | 2.05483 | 50879 | 1.96684 | 53059 53096 | 1.88469 1.88337 | 2 |
| 59 | 46595 | $2 \cdot 14614$ | 48737 | $2 \cdot 05182$ | 50916 | 1.96402 | 53134 | 1.8820 | 1 |
| 60 | 46631 | 2.1445 1 | 48773 | $2 \cdot 05030$ | 50093 | 1.96261 | 53171 | 1.88073 | 0 |
|  | Cotang. | Tangent. | Cotang. Tangent. |  | Cotang. Tangent. |  | C $=$ tang. ${ }^{\text {a }}$ Tangent. |  |  |
|  |  | $5^{\circ}$ | $64^{\circ}$ |  | $63^{\circ}$ |  | $62^{\circ}$ |  |  |


| 80 | NATURAI |  | TANGENTS |  | AND COTANGENTS. |  |  | T ABta IIL. |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | $28^{\circ}$ |  | $29^{\circ}$ |  | $30^{\circ}$ |  | \$1 ${ }^{\text {a }}$ |  | , |
|  | Targent. | Cotang. | Tangent. | Cotang. | Tangent. | Cotang. | Tangent. | Cotang. |  |
| 0 | 53171 | 1.88073 | 55431 | 1.80405 | 57735 | 1.73205 | 60086 | 1.66428 | 60 |
| 1 | 53208 | 1.87941 | 55469 | 1.80281 | 57774 | 1.73089 | 60126 | 1.663ı8 | 59 |
| 2 | 53246 | 1.8780) | 55507 | I.80158 | 57813 | $1 \cdot 72973$ | 60165 | 1.66209 | 58 |
| 3 | 53283 | 1.87677 | 55545 | 1.80034 | 57851 | I-72857 | 60205 | 1.65099 | 57 |
| 4 | 53320 | 1.87546 | 55583 | I.79911 | 57890 | I.72741 | 60245 | 1.65990 | 56 |
| 5 | 53358 | 1.87415 | 55521 | 1-79788 | 57929 | 1-72625 | 60284 | 1.65881 | 55 |
| 6 | 53395 | 1.87283 | 55659 | 1-79665 | 57968 | 1.72509 | 60324 | 1.65772 | 54 |
| 7 | 53432 | I. 87152 | 55697 | 1-79542 | 58007 | $1 \cdot 7239^{3}$ | 60364 | 1.65663 | 53 |
| 8 | 53470 | 1.87021 | 55736 | 1-79419 | 58046 | I-72278 | 60403 | 1.65554 | 52 |
| 9 | 53507 | 1.86891 | 55774 | I - 79296 | 58085 | $1 \cdot 72163$ | 60443 | 1.65445 | 51 |
| 10 | 53545 | I. 86760 | 55812 | $1 \cdot 79174$ | 58124 | $1 \cdot 72047$ | 60483 | 1.65337 | 50 |
| 11 | 53582 | 1.86630 | 55850 | 1-7905ı | 58162 | - $\cdot 711^{3} 2$ | 60522 | 1.65228 | 49 |
| 12 | 53320 | I.86499 | 55888 | 1-78929 | 58201 | 1.71817 | 60562 | 1.65120 | 48 |
| 13 | 53657 | 1.86369 | 55926 | $1 \cdot 78807$ | 58240 | 1.71702 | 60602 | 1.65011 | 47 |
| 14 | 53694 | 1.86239 | 55964 | I•78685 | 58279 | 1.71588 | 60642 | 1.64903 | 45 |
| 15 | 53732 | 1.86109 | 56003 | 1.78563 | 58318 | $1 \cdot 71473$ | 60681 | 1.64795 | 45 |
| 16 | 53769 | 1.85979 | 56041 | 1.78441 | 58357 | 1.71358 | 60721 | 1.64687 | 44 |
| 17 | 53807 | 1.85850 | 56079 | $1 \cdot 78319$ | 58396 | 1.71244 | 60761 | 1.64579 | 43 |
| 18 | 53844 | 1.85720 | 56117 | 1-78198 | 58435 | 1.71129 | 60801 | 1.6447 | 42 |
| 19 | 53882 | 1.85591 | 56156 | 1-78077 | 58474 | 1.71015 | 60841 | 1.64363 | 41 |
| 20 | 53920 | 1.85462 | 56194 | 1.77955 | 585 I 3 | $1 \cdot 70901$ | 60881 | 1.64256 | 40 |
| 21 | 53957 | 1.85333 | 56232 | 1-77834 | 58552 | 1.70787 | 60921 | 1.64148 | 39 |
| 22 | 53995 | 1.85204 | 56270 | 1-77713 | 58591 | 1. 70673 | 60960 | 1.64041 | 38 |
| 23 | 54032 | 1.85075 | 56309 | $1 \cdot 77592$ | 58631 | $1 \cdot 70560$ | 61000 | 1.63934 | 37 3 |
| 24 | 54070 | 1.84946 | 56347 | $1 \cdot 77471$ | 58670 | 1.70446 | 61040 | 1.63826 | 36 |
| 25 | $5 / 1107$ 54145 | 1.84818 | 56385 | 1-77351 | 58709 58748 | 1.70332 | 61080 61120 | 1.63719 1.63612 | 35 34 |
| 27 | 54183 | 1.84561 | 56462 | 1-77110 | 58787 | 1.70106 | 61160 | 1.63505 | 33 |
| 28 | 54220 | 1.84433 | 56500 | 1-76990 | 58826 | 1.69992 | 61200 | 1.63398 | 32 |
| 29 | 54258 | 1.84305 | 56539 | 1-76869 | 58865 | 1.69879 | 61240 | 1.632,2 | 31 |
| 30 | 54296 | 1.84177 | 56577 | 1-76749 | 58904 | 1.69766 | 80 | 1.63185 | 30 |
| 31 | 54333 | 1.84049 | 56616 | 1.76630 | 58944 | 1.69653 | 61320 | 1.63079 | 29 |
| 32 | 54371 | 1.83922 | 56654 | 1-765ı0 | 58983 | 1.69541 | 61360 | 1.62972 | 28 |
| 33 | 54409 | 1.83794 | 56693 | 1-76390 | 59022 | 1.69428 | 61400 | 1.62866 | 27 |
| 34 | 54446 | I.83667 | 56731 | 1.76271 | 59061 | 1.69316 | 61440 | 1.62760 | 26 |
| 35 | 54484 | 1.83540 | 55769 | 1-76151 | 59101 | 1.69203 | 61480 | 1.62654 | 25 |
| 36 | 54522 | 1.83413 | 56808 | 1-76032 | 59140 | 1.69091 | 61520 | 1.62548 | 24 |
| 37 | 54560 | I. 83286 | 56846 | 1-75913 | 59179 | 1.68979 | 61561 | 1.62442 | 23 |
| 38 | 54507 | 1.83159 | 56885 | 1-75794 | 59218 | 1.68866 | 61601 | 1.62336 | 22 |
| 39 | 54635 | 1.83033 | 56923 | 1.75675 | 59258 | 1.68754 | 61641 | 1.62230 | 21 |
| 40 | 54673 | I.82906 | 56962 | 1.75556 | 59297 | 1.68643 | 61681 | 1.62125 | 20 |
| 41 | 54711 | 1.82780 | 57000 | 1.75437 | 59336 | 1.68531 | 61721 | 1.62019 | 19 |
| 42 | 54748 | 1.82654 | 57039 | $1 \cdot 75319$ | 59376 | 1.68419 | 61761 | 1.61914 | 18 |
| 43 | 54786 | 1.82528 | 57078 | $1 \cdot 750200$ | 59415 | 1.68308 | 61801 | 1.61808 | 17 |
| 44 | 54824 | 1.82402 | 57116 | 1-75082 | 59454 | 1.68196 | 61842 | 1.61703 | 16 |
| 45 | 54862 | 1.82276 | 57155 | 1-74964 | 59494 | 1.68085 | 61882 | 1.61598 | 15 |
| 46 | 54900 | 1.82150 | $5719^{3}$ | 1.74846 | 59533 | 1.67974 | 61922 | 1.61493 | 14 |
| 47 | 54938 | 1. 82025 | 57232 | 1.74728 | 59573 | 1.67563 | 61962 | 1.61388 | 13 |
| 48 | 54975 | 1.81899 | 57271 | 1.74610 | 59612 | $1.677^{52}$ | 62003 | 1.61283 | 12 |
| 49 | 55013 | 1.81774 | 57309 | 1-74492 | 59651 | 1.57641 | 62043 | 1.61179 | 11 |
| 50 | 55051 | 1.81649 | 57348 | 1.74375 | 59691 | 1.67530 | 62083 | 1.61074 | 10 |
| 51 | 55089 | 1.81524 | 57386 | 1.74257 | 59730 | 1.67419 | 62124 | 1.60970 | 9 |
| 52 | 55127 | 1.81399 | 57425 | 1.74140 | 59770 | 1.67309 | 62164 | 1.60865 | 8 |
| 53 | 55165 | $1 \cdot 81274$ | 57464 | 1-74022 | 59809 | 1.67198 | 62204 | I. 60761 | 7 |
| 54 | 55203 | 1.81150 | 57503 | 1.73905 | 59849 | 1.67088 | 62245 | 1.60657 | 6 |
| 55 | 555241 | 1.81025 | 57541 | 1.73788 | 59888 | 1.66978 | 62285 | 1.60553 | 5 |
| 56 | 55279 | 1.80901 | 57580 | 1.73671 | 59928 | 1.66867 | 62325 | 1.60449 | 4 |
| 57 | 55317 | 1-80777 | 57619 | 1-73555 | 59967 | 1.66757 | 52366 | 1.60345 | 3 |
| 58 | 55355 | 1.80653 | 57657 | 1.73438 | 60007 | 1.66647 | 62406 | 1.60241 | 2 |
| 59 | $5530^{3}$ | 1-80529 | 57606 | 1.73321 | 60046 | 1.66538 | 62446 | 1.60137 | 1 |
| 60 | 55431 | 1-80405 | 57735 | 1.73205 | 60086 | 1.66428 | 62487 | 1.6co33 | 0 |
|  | Cotang, | Tangent. | Cotang. | Tangent. | Cotang. | Tangent. | Cotang | Tangent. |  |
|  | $61^{\circ}$ |  | $60^{\circ}$ |  | $59^{\circ}$ |  | $55^{\circ}$ |  |  |


| Table III. |  | NATURAL TANGENTS |  |  |  | AND COTANGENTS. |  |  | 81 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | $32^{\circ}$ |  | $33^{\circ}$ |  | $34^{\circ}$ |  | $35^{\circ}$ |  | , |
|  | Tangent. | Cotang. | Tangent. | Cotang. | Tangent. | Cotang. | Targent. | Cotang. |  |
| 0 | 62487 | 1.60033 | 64941 | 1.53986 | 67451 | 1-48256 | 7 CO 21 | $1 \cdot 42815$ | 60 |
| 1 | 62527 | 1.59930 | $649^{82}$ | 1.53888 | 67493 | 1-48163 | 70064 | 1.42726 | 59 |
| 2 | 62568 | 1.59826 | 65023 | 1.53791 | 67536 | 1-48070 | 70107 | 1.42638 | 58 |
| 3 | 62608 | I. 59723 | 65065 | 1.53693 | 67578 | 1.47977 | 70151 | $1 \cdot 42550$ | 57 |
| 4 | 62649 | 1.59620 | 65106 | 1.53595 | 67620 | 1.47885 | 70194 | 1.42462 | 56 |
| 5 | 62689 | 1.59517 | 65148 | 1.53497 | 67663 | 1.47792 | 70238 | $1 \cdot 42374$ | 55 |
| 6 | 62730 | 1.59414 | 65189 | 1.53400 | 67705 | 1.47699 | 70281 | 1.42286 | 54 |
| 7 | 62770 | 1.59311 | 65231 | 1.533r,2 | 67748 | 1.47607 | 70325 | $1 \cdot 42198$ | 53 |
| 8 | 52811 | 1-59208 | 65272 | 1.53205 | 67790 | 1.47514 | 70.368 | 1.42110 | 52 |
| 9 | 62852 | $1 \cdot 59105$ | 65314 | 1.53107 | 67832 | 1-47422 | 70412 | 1.42022 | 51 |
| 10 | 62892 | $1 \cdot 50002$ | 65355 | 1.53010 | 57875 | 1.47330 | 70455 | $1.419^{34}$ | 50 |
| 11 | 62933 | 1.58900 | 65397 | 1.52913 | 67917 | 1.47238 | 70499 | 1.41847 | 49 |
| 12 | 62973 | 1.58797 | 65438 | 1.52816 | 57960 | 1.47146 | 70542 | 1.41759 | 48 |
| 13 | 63014 | 1.58695 | 65480 | 1.52719 | 68002 | 1.47053 | 70586 | 1.41672 | 47 |
| 14 | 63055 | 1.58593 | 65521 | 1.52622 | 68045 | 1.46962 | 70629 | 1.41584 | 46 |
| 15 | 63095 | 1.58490 | 65563 | 1.52525 | 68088 | 1-46870 | 70673 | 1-41497 | 45 |
| 16 | 63ı36 | 1. 58388 | 65604 | 1.52429 | 68130 | 1-46778 | 70717 | 1.41409 | 44 |
| 17 | 63177 | 1.58286 | 65646 | 1.52332 | 68173 | 1-46686 | 70760 | 1.41322 | 43 |
| 18 | 63217 | 1.58184 | 65688 | 1.52235 | 68215 | 1-46595 | 70804 | $1 \cdot 41235$ | 42 |
| 19 | 63258 | 1.58083 | 65729 | 1.52139 | 68258 | 1-46503 | 70848 | 1.41148 | 41 |
| 20 | 63299 | 1.57981 | 65771 | 1.52043 | 68301 | I-46411 | 70891 | 1-41061 | 40 |
| 21 | 63340 | 1.57879 | 65813 | 1.51946 | 68343 | 1-46320 | 70935 | $1 \cdot 40974$ | 39 |
| 22 | 63380 | 1.57778 | 65854 | 1.51850 | 68386 | 1-46229 | 70979 | 1-40887 | 38 |
| 23 | 63421 | 1.57676 | 65896 | 1.51754 | 68429 | 1-46137 | 71023 | 1-40800 | 37 |
| 24 | 63462 | $1.5757^{5}$ | 65938 | 1.51658 | 68471 | $1 \cdot 46046$ | 71066 | 1-40714 | 36 |
| 25 | 63503 | 1.57474 | 65980 | 1.51562 | 68514 | 1-45955 | 71110 | 1-40627 | 35 |
| 26 | 63544 | $1.57^{3} 72$ | 66021 | 1.51466 | 98557 | $1 \cdot 45864$ | 71154 | 1-40540 | 34 |
| 27 | 63584 | 1.57271 | 66063 | 1.51370 | 68600 | 1-45773 | 71198 | 1-40454 | 33 |
| 28 | 63625 | 1.57170 | 66105 | 1.51275 | 68642 | 1-45682 | 71242 | 1-40367 | 32 |
| 29 | 63666 | 1.57069 | 66147 | 1.51179 | 68685 | 1-45592 | 71285 | 1-40281 | 31 |
| 30 | 63707 | 1.56969 | 66189 | 1.51084 | 68728 | $1 \cdot 45501$ | 71329 | 1-40195 | 30 |
| 31 | 63748 | 1.56868 | 66230 | 1.50988 | 68771 | 1.45410 | 71373 | 1-40109 | 29 |
| 32 | 63789 | 1.56767 | 66272 | 1.50893 | 68814 | 1-45320 | 71417 | $1 \cdot 40022$ | 28 |
| 33 | 63830 | 1.56667 | 66314 | 1.50797 | 68857 | 1-45229 | 71461 | 1.39936 | 27 |
| 34 | 63871 | 1.56566 | 66356 | 1.50702 | 68900 | 1-45139 | 71.505 | 1.39850 | 26 |
| 35 | 63912 | 1.56466 | 66398 | 1.50607 | 68942 | 1.45049 | 71549 | I.39764 | 25 |
| 36 | 63953 | 1.56366 | 66440 | 1.50512 | 68985 | 1.44058 | $715{ }^{\text {c }} 3$ | 1.39679 | 24 |
| 37 | 63994 | 1.56265 | 66482 | 1.50417 | 69028 | 1-44868 | 71637 | 1.39593 | 23 |
| 38 | 64035 | 1.56165 | 66524 | 1.50322 | 69071 | 1.44778 | 71681 | 1.39507 | 22 |
| 39 | 64076 | 1.56065 | 66566 | 1.50228 | 69114 | 1.44688 | 71725 | 1.39421 | 21 |
| 40 | 64117 | 1.55966 | 66608 | 1.50133 | 69157 | 1.44598 | 71769 | 1.39336 | 20 |
| 41 | 64158 | 1.55866 | 66650 | 1.50038 | 69200 | 1.44508 | 71813 | I. 39250 | 18 |
| 42 | 64199 | 1.55766 | 66692 | 1.49944 | 69243 | 1.44418 | 71857 | 1.39165 | 18 |
| 43 | 64240 | 1.55666 | 66734 | 1.49849 | 69286 | 1.44329 | 71901 | 1.39079 | 17 |
| 44 | 64281 | 1.55567 | 66776 | 1.49755 | 69329 | 1-44239 | 71946 | I.38994 | 16 |
| 45 | 64322 | 1.55467 | 66818 | 1.49661 | 69372 | 1.44149 | 71990 | 1.38909 | 15 |
| 46 | 64363 | 1. 55368 | 66860 | 1-49566 | 69416 | 1-44060 | 72034 | 1.38824 | 14 |
| 47 | 64404 | 1.55269 | 66902 | 1.49472 | 69459 | $1 \cdot 43970$ | 72078 | I. 38738 | 13 |
| 48 | 64446 | 1.55170 | 66944 | 1-49378 | 69502 | 1-43881 | 72122 | I $\cdot 38653$ | 12 |
| 49 | 64487 | 1.55071 | 66986 | 1.49284 | 69545 | 1.43792 | 72166 | I.38568 | 11 |
| 50 | 64528 | 1.54972 | 67028 | 1.49190 | 69588 | $1 \cdot 43703$ | 72211 | 1-38484 | 10 |
| 51 | 64569 | 1.54873 | 67071 | 1.49097 | 69631 | 1.43614 | 72255 | 1.38399 | 8 |
| 52 | 64610 | 1.54774 | 67113 | 1.49003 | 69675 | 1.43525 | 72299 | 1.38314 | 8 |
| 53 | 64652 | 1.54675 | 67155 | 1.48959 | 69718 | 1.43436 | 72344 | 1.38229 | 7 |
| 54 55 | 64693 | 1.540] 76 | 67197 | 1.48816 | 69761 | 1.43347 | 72388 | : 38145 | 6 |
| 55 56 | 64734 | 1.54478 | 67239 | 1.48722 | 69804 | 1.43258 | 72432 | 1.38060 | 5 |
| 56 57 | 64775 | 1.54379 | 67282 | 1.48629 | 69847 | 1.43169 | 72477 | 1.37976 | 4 |
| 57 38 | 64817 | 1.54281 | 67324 | 1.48536 | 69891 | 1.43080 | 72521 | 1.37891 | 3 |
| 59 | 64858 | 1.54183 1.54085 | 67366 | 1.48442 | 69934 | 1.42992 1.42003 | 72505 | 1.37807 | 2 |
| 60 | 64941 | 1.53986 | 67451 | 1.48256 | 70021 | 1.42815 | 72654 | 1.37638 | - |
|  | Cotang. | Tangent. | Cotang. | Tangent. | Cciang. | Tangent. | Cotang. | Tangent. |  |
|  | $57^{\circ}$ |  | $56^{\circ}$ |  | $55^{\circ}$ |  | $54^{\circ}$ |  |  |



| Tablee IIl. N |  |  | ATURAL TANGENTS AND COTANGENTS. |  |  |  |  |  | 83 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | $40^{\circ}$ |  | $41^{\circ}$ |  | $42^{\circ}$ |  | $43^{\circ}$ |  |  |
|  | Tangent. | Cotang. | Tangent. | Cotang. | Tangent. | Cotang. | Tangent. | Cotang. |  |
| 0 | 83910 | 1-19175 | 86929 | 1-15037 | 90040 | 1.11061 | 93252 | 1.07237 | 60 |
| 1 | 83960 | 1-19105 | 86980 | $1 \cdot 14969$ | 90093 | 1-10996 | 93306 | 1.07174 | 50 |
| 2 | 84009 | 1-19035 | 8703 I | $1 \cdot 14902$ | 90146 | 1-10031 | 93360 | 1.07112 | 58 |
| 3 | ${ }^{8405} 9$ | 1-18964 | 87082 | $1 \cdot 14834$ | 90199 | 1. 10867 | 93415 | 1.07049 | 57 |
| 4 | 84108 | 1-18894 | 87133 | 1.14767 | 90251 | 1-10802 | 93469 | 1.06987 | 56 |
| 5 | 84158 | 1-18324 | 87184 | 1-14699 | 90304 | 1-10737 | 93524 | 1.06925 | 55 |
| 6 | 84208 | 1-18754 | 87236 | 1.14632 | 90357 | 1-10672 | 93578 | 1.06862 | 54 |
| 7 | 84258 | 1.18684 | 87287 | 1.14565 | 90410 | 1-10607 | 93633 | 1.06800 | 53 |
| 8 | 84307 | 1-18614 | 87338 | 1-14498 | 90463 | 1-10543 | 93688 | 1.06738 | 52 |
| 9 | 84357 | 1-18544 | 87389 | 1.14430 | 90516 | 1-10478 | 93742 | 1.06676 | 51 |
| 10 | 84407 | 1-18474 | 87441 | 1-14363 | 90569 | I-10414 | 93797 | 1.06613 | 50 |
| 11 | 84457 | 1.18404 | 87492 | 1.14296 | 90621 | 1-10349 | 93852 | 1.06551 | 48 |
| 12 | 84507 | I-18334 | 87543 | 1.14229 | 90674 | 1-10285 | 93906 | 1.06489 | 48 |
| 13 | 84556 | 1-18264 | 87595 | 1-14162 | 90727 | 1-10220 | 93961 | 1.06427 | 47 |
| 14 | 84606 | 1-18194 | 87646 | 1-14095 | 90781 | I - 10156 | 94016 | 1.06365 | 46 |
| 15 | 84656 | 1-18125 | 87698 | 1.14028 | 90834 | 1-10091 | 94071 | 1.063o3 | 45 |
| 16 | 84706 | 1-18055 | 87749 | $1 \cdot 13961$ | 90887 | 1-10027 | 94125 | 1.06241 | 44 |
| 17 | 84756 | 1-17986 | 87801 | 1-13894 | 90040 | 1-09963 | 94180 | 1.06179 | 43 |
| 18 | 84806 | 1-17916 | 87852 | I-13828 | 90993 | 1.09899 | 94235 | 1.06117 | 42 |
| 19 | 84856 | 1.17846 | 87904 | $1 \cdot 13761$ | 91046 | 1-09834 | 94290 | I. 06056 | 41 |
| 20 | 84906 | 1-17777 | 87955 | 1-13694 | 91009 | 1-09770 | 94345 | 1.05994 | 40 |
| 21 | 84956 | 1-17708 | 88007 | I-13627 | 91153 | 1.09706 | 94400 | 1.05932 | 39 |
| 22 | 85006 | 1-17638 | 88059 | 1-13561 | 91206 | 1-09642 | 94455 | 1.05870 | 38 |
| 23 | 85057 | 1-17569 | 88110 | I-13494 | 91259 | 1.09578 | 94510 | I.05809 | 37 |
| 24 | 85107 | 1-17500 | 88162 | I.13428 | 91313 | 1.09514 | 94565 | 1.05747 | 36 |
| 25 | 85157 | 1-17430 | 88214 | I-1336I | 91366 | 1.09450 | 94620 | 1.05685 | 35 |
| 26 | 85207 | 1-17361 | 88265 | 1-13295 | 91419 | 1-09386 | 94676 | 1.05624 | 34 |
| 27 | 85257 | 1-17292 | 88317 | I-13228 | 91473 | 1.00322 | 94731 | 1.05502 | 33 |
| 28 | 85307 | 1-17223 | 88369 | I-13162 | 91526 | 1.09258 | 94786 | 1.05501 | 32 |
| 29 | 85358 | 1-17154 | 88421 | I-13096 | 91580 | 1.09195 | 9484 I | 1.05439 | 31 |
| 30 | 85408 | 1-17085 | 88473 | 1-13029 | 91633 | 1.09131 | 94896 | 1.05378 | 30 |
| 31 | 85458 | 1.17016 | 88524 | 1-12963 | 91687 |  | 94952 |  |  |
| 32 | 85509 | 1-16947 | 88576 | 1-12897 | 91740 | 1.00003 | 95007 | $1 \cdot 05255$ | 28 |
| 33 | 85559 | 1-16878 | 88628 | 1-1283I | 91794 | 1.08040 | 95062 | 1.05194 | 27 |
| 34 | 85609 | 1-16809 | 88680 | 1-12765 | 91847 | 1.08876 | 95118 | 1.05133 | 26 |
| 35 | 85660 | 1-16741 | 88732 | 1-12699 | 91901 | 1-08813 | 95173 | 1-05072 | 25 |
| 36 | 85710 | 1-16672 | 88784 | 1.12633 | 91955 | 1.08749 | 95229 | $1 \cdot 05010$ | 24 |
| 37 | ${ }^{85761}$ | 1-16603 | 88836 | 1-12567 | 92008 | 1.08686 | 95284 | 1.04949 | 23 |
| 38 | 85811 | 1-16535 | 88888 | 1-12501 | 92062 | 1-08622 | 95340 | 1.04888 | 22 |
| 39 | 85862 | 1-16406 | 88940 | 1-12435 | 92116 | 1.08559 | $953 q^{5}$ | 1.04827 | 21 |
| 40 | 85912 | 1-16398 | 88992 | 1-12369 | 92170 | 1.08496 | 95451 | 1.04766 | 20 |
| 41 | 85963 | 1.16329 | 890.5 | 1-12303 | 92223 | 1.08432 | 95506 | I. 04705 | 19 |
| 42 | 86014 | 1-16261 | 89097 | 1-12238 | 92277 | 1.08369 | 95502 | I.04644 | 18 |
| 43 | 86064 | 1-16192 | 89149 | 1-12172 | 92331 | 1.08306 | 95618 | 1.04583 | 17 |
| 44 | 86115 | 1-16124 | 89201 | 1-12106 | 92385 | 1.08243 | 95673 | $1 \cdot 04522$ | 16 |
| 45 | 86166 | 1-16056 | 89253 | 1-12041 | 92439 | I.08179 | 95729 | 1.04461 | 15 |
| 46 | 86216 | 1-15987 | 89306 | 1-11975 | 92493 | 1.08116 | 95785 | 1.04401 | 14 |
| 47 | 86267 | 1-15919 | 89358 | 1-11909 | 92547 | I - 08053 | 9584 I | 1.04340 | 13 |
| 48 | 86318 | 1-1585ı | 89410 | 1-11844 | 92601 | 1.07990 | 95897 | 1.04279 | 12 |
| 49 | 86368 | 1-15783 | 89463 | 1.11778 | 92655 | I. 07927 | 95952 | 1.04218 | 11 |
| 50 | 86419 | 1-15715 | 89515 | 1-11713 | 92709 | I $\cdot 07864$ | 96008 | 1.04158 | 10 |
| 51 | 86470 | 1-15647 | 89567 | 1-11648 | 92763 | 1.07801 | 96064 | 1.04097 | 8 |
| 52 | 86521 | 1-15579 | 89620 | 1-11582 | 92817 | 1.07738 | 96120 | 1-04036 | 8 |
| 53 | 86572 | 1-15511 | 89672 | 1.11517 | 92872 | 1.07676 | 96176 | 10.3976 | 7 |
| 54 | 86623 | $1 \cdot 15443$ | 89725 | 1-11452 | 92926 | 1.07613 | 96232 | 103915 | 6 |
| 55 | 86674 | 1.15375 | 89777 | 1-11387 | 92980 | 1.07550 | 96288 | 1.03855 | 5 |
| 56 | 86725 | 1-15308 | 89830 | 1-11321 | 93034 | I. 07487 | 96344 | 1.03794 | 4 |
| 57 | 86776 | I-15240 | 89883 | 1-11256 | 93088 | 1.07425 | 96400 | 1.03734 | 3 |
| 58 | 86827 | 1-15172 | 89935 | 1-11191 | 93143 | 1.07362 | 96457 | I.03074 | 2 |
| 59 | 86878 | 1.15104 | 89988 | 1-11126 | 93197 | I. 07299 | 96513 | I.03613 | 1 |
| 60 | 86929 | $1 \cdot 15037$ | 90040 | 1-11061 | 93252 | 1.07237 | 96569 | I $\cdot 03553$ | - |
|  | Cotang. ${ }^{\text {Tangent. }}$ |  | Cotang. Tangent. |  | Cotang | Tangent. | Cotang. Tangent. |  |  |
|  |  |  | $48^{\circ}$ |  | $47^{\circ}$ |  | $46^{\circ}$ |  |  |


| 84 | NATU | L TANG | TS | D C | ANGEN | TA | III |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |  |  |  |
|  | Tangent. | Cotang. | , | $\gamma$ | Tangent. | Cotang. | , |
| 0 | 96569 | 1.03553 | 60 | 3 I | 98327 | 1.01702 | 29 |
| 1 | 96625 | 1.03493 | 59 | 32 | 98384 | 1.01642 | 28 |
| 2 | 96681 | 1.03433 | 58 | 33 | 98441 | 1.01583 | 27 |
| 3 | 96738 | 1.03372 | 57 | 34 | 98499 | 1.01524 | 26 |
| 4 | 96794 | 1.03312 | 56 | 35 | 98556 | 1.01465 | 25 |
| 5. | 96850 | I. 03252 | 55 | 36 | 98613 | 1.01406 | 24 |
| 6 | 96907 | 1.03192 | 54 | 37 | 98671 | 1.01347 | 23 |
| 7 | 96963 | 1.03132 | 53 | 38 | 98728 | 1.01288 | 22 |
| 8 | 97020 | $1 \cdot 03072$ | 52 | 39 | 98786 | 1.01229 | 21 |
| 9 | 97076 | 1.03012 | 51 | 40 | 98843 | 1.01170 | 20 |
| 10 | 97133 | 1.02952 | 50 | 41 | 98901 | I.01112 | 19 |
| 11 | 97189 | 1.02892 | 49 | 42 | 98958 | I. 01053 | 18 |
| 12 | 97246 | 1.02832 | 48 | 43 | 99016 | I $\cdot 00994$ | 17 |
| 13 | 97302 | 1.02772 | 47 | 44 | 99073 | 1.00935 | 16 |
| 14 | 97359 | I. 02713 | 46 | 45 | 99131 | 1.00876 | 15 |
| 15 | 97416 | I.02653 | 45 | 46 | -99189 | I. 00818 | 14 |
| 16 | 97472 | 1.02593 | 44 | 47 | 99247 | $1 \cdot 00759$ | 13 |
| 17 | 97529 | 1.02533 | 43 | 48 | 99304 | 1.00701 | 12 |
| 18 | 97586 | 1.02474 | 42 | 49 | 99362 | 1.00642 | 11 |
| 19 | 97643 | 1.02414 | 41 | 50 | 99420 | I $\cdot 00583$ | 10 |
| 20 | 97700 | 1.02355 | 40 | 51 | 99478 | I - 00525 | 9 |
| 21 | 97756 | 1.02295 | 39 | 52 | 99536 | I. 00467 | 8 |
| 22 | 97813 | 1.02236 | 38 | 53 | 99594 | 1-00408 | 7 |
| 23 | 97870 | 1.02176 | 37 | 54 | 99652 | 1.00350 | 6 |
| 24 | 97927 | 1.02117 | 36 | 55 | 99710 | 1.00291 | 5 |
| 25 | 97984 | 1.02057 | 35 | 56 | 99768 | $1 \cdot 00233$ | 4 |
| 26 | 98041 | 1.01998 | 34 | 57 | 99826 | 1.00175 | 3 |
| 27 | 98098 | 1 01939 | 33 | 58 | 99884 | 1.00110 | 2 |
| 28 | 98155 | 1.01879 | 32 | 59 | 99942 | $1 \cdot 00058$ | 1 |
| 29 | 98213 | 1.01820 | 3 I | 60 | Unit. | Unit. | - |
| 30 | 98270 | 1.01761 | 30 |  |  |  |  |
| 1 | Cotang. | Tangent. | 1 | , | Cotang. | Tangent. | , |
|  | $45^{\circ}$ |  |  |  | $45^{\circ}$ |  |  |

## TABLE OF CONSTANTS.

Base of Napier's system of logarithms $=$ $\qquad$ $\varepsilon=2 \cdot 718281828459$
Mod. of common syst. of logarithms $=\ldots$ com. log. $\varepsilon=\mathrm{M}=0.43429448: 903$
Ratio of circumference to diameter of a circle $=\ldots \ldots \ldots \ldots \pi=3 \cdot 141592653590$ $\log . \pi=0.497149872694$
$\pi^{2}=9.869604401089 \ldots \ldots \ldots . \sqrt{ } \pi=1 \cdot 772453850906$
Are of same length as radius $=\ldots \ldots \ldots . .180^{\circ} \div \pi=10800^{\prime} \div \pi=648000^{\prime \prime} \div \pi$
$180^{\circ} \div \pi=57^{\circ} \cdot 2957795130$ log. $=1 \cdot 758_{122632409}$
$10800^{\prime} \div \pi=3437^{\prime} \cdot 7467707849, \ldots \ldots \ldots . . . . . . . . . . . . . . . . . . \log .=3 \cdot 536273882793$
$648000 " \div \pi=206264 " \cdot 8062470964, \ldots \ldots \ldots . . . . . . . . . . . . . \log .=5 \cdot 314425133176$
Tropical year $=365 \mathrm{~d} .5 \mathrm{~h} .4 \mathrm{Sm} .47 \mathrm{~s}, \cdot 588=365 \mathrm{~d} . \cdot 242217456, \log .=2.5625810$
Sidereal year $=365 \mathrm{~d}$. 6h. qm. 10s. $\cdot 742=365 \mathrm{~d} . \cdot 256374332, \log .=2.5625978$
24 h . sol. t. $=24 \mathrm{~h} .3 \mathrm{~m} .56 \mathrm{~s} .055335 \mathrm{sid}$. t. $=24 \mathrm{~h} . \times \mathrm{I} \cdot 0027379 \mathrm{l}, \log . \mathrm{I} \cdot 002=0.0011874$
24 h. sid. $\mathrm{t} .=2$ 亿h. $-(3 \mathrm{~m} .55 \mathrm{~s} . \cdot 90944)$ sol. t. $=24 \mathrm{~h} . \times 0 \cdot 9972696, \log .0 \cdot 997=9 \cdot 9988126$
British imperial gallon $=277 \cdot 274$ cubic inches, $\qquad$ log. $=2 \cdot 4429091$
Length of sec. pend., in inches, at London, $39 \cdot 13929$; Paris, $39 \cdot 1285$; New
York, 39 - 1285.
French metre $=3 \cdot 2808992$ English feet $=39.3707904$ inches.
I cubic inch of water (bar. 30 inches, Fahr. therm. $62^{\circ}$ ) $=\mathbf{2 5 2 \cdot 4 5 8}$ Troy grains.

\begin{tabular}{|c|c|c|c|c|c|c|c|c|c|}
\hline \multirow[t]{3}{*}{} \& \multicolumn{9}{|c|}{Refraction in Declination.} \\
\hline \& \multicolumn{9}{|c|}{For Latitude 150.} \\
\hline \& + \(20^{\circ}\) \& \(+15^{\circ}\) \& \(+10^{\circ}\) \& \(+5^{\circ}\) \& \(0{ }^{\circ}\) \& \(-5^{\circ}\) \& \(-10^{\circ}\) \& \(-15^{\circ}\) \& \(-20^{\circ}\) \\
\hline -h. \& -05" \& - \({ }^{\prime \prime}\) \& +o5" \& \(\mathrm{Io}^{\prime \prime}\) \& \({ }^{15}{ }^{\prime \prime}\) \& \(2 \mathrm{I}^{\prime \prime}\) \& \(27^{\prime \prime}\) \& \(33^{\prime \prime}\) \& \(40^{\prime \prime}\) \\
\hline 2 \& -03 \& +02 \& \(\bigcirc\) \& 12 \& 18 \& \begin{tabular}{l}
23 \\
28 \\
\hline
\end{tabular} \& 29 \& 36 \& 43 \\
\hline 3 \& +oI \& 05 \& 11 \& 16 \& 22 \& 28 \& 34 \& 4 I \& , 49 \\
\hline 4 \& -8 \& 12
34 \& 19
41 \& 24 \& 30
59 \& ( \(\begin{array}{r}37 \\ \text { r }\end{array}\) \& 44
\(\mathrm{r}^{4} 24\) \& r

$\mathrm{I}^{53} 43$ \& [ <br>
\hline 5 \& 29 \& 34 \& 4 I \& 49 \& 59 \& I 10 \& $\mathrm{I}^{\prime 24}$ \& 143 \& 208 <br>
\hline
\end{tabular}

For Latitude $17^{\circ} 30^{\prime}$.

| - h. | -02" | +o2' | -8 ${ }^{\prime \prime}$ | $13^{\prime \prime}$ | $18{ }^{\prime \prime}$ | $24^{\prime \prime}$ | $30^{\prime \prime}$ | $36^{\prime \prime}$ | $44^{\prime \prime}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 2 | $\bigcirc$ | 05 | то | 15 | 21 | 27 | 33 | 40 | 48 |
| 3 | +o2 | 10 | 15 | 2 I | 27 | 33 | 40 | 48 | 57 |
| 4 | 13 | 18 | 23 | 29 58 | + 35 | +43 | ${ }^{1}{ }^{515}$ | ${ }^{1}$ о ${ }^{\text {a }}$ | $\mathrm{I}^{\text {I } 13}$ |
| 5 | 34 | 4 I | 49 | 58 | $\mathrm{I}^{\prime}$ 'о | $\mathrm{I}^{\prime} 23$ | $\mathrm{I}^{\prime}{ }^{51}$ | 206 | 242 |

For Latitude $20^{\circ}$.

| oh. | $0^{\prime \prime}$ | 05" | ro' ${ }^{\prime \prime}$ | $15^{\prime \prime}$ | $2 \mathrm{I}^{\prime \prime}$ | $27^{\prime \prime}$ | $33^{\prime \prime}$ | $40^{\prime \prime}$ | $48^{\prime \prime}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 2 | $\bigcirc 3$ | $\bigcirc 7$ | 13 | 18 | 24 | 30 | 36 | 44 | ${ }^{52}$ |
| 3 | $\bigcirc 6$ | 13 | 18 | 24 | 30 | 36 | 44 | 52 | $\mathrm{I}^{\prime} \mathrm{O}$ |
| 4 | 17 | 22 | 28 | 35 | 42 | . 50 | I'00 | I'ti | I 26 |
| 5 | 39 | 47 | 57 | r'07 | I'20 | r'37 | 200 | 232 | 325 |

For Latitude $22^{\circ} 3^{\circ}$ 。

| - h. | -2' ${ }^{\prime \prime}$ | 08" | $13^{\prime \prime}$ | $18^{\prime \prime}$ | $24^{\prime \prime}$ | $30^{\prime \prime}$ | $3^{6 \prime}$ | $44^{\prime \prime}$ | $52^{\prime \prime}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 2 | 06 | 11 | 15 | 21 | 27 | 33 | 40 | 48 | 57 |
| 3 | 11 | 15 | 2 T | 27 | 33 | 40 | 48 | 57 | '08 |
| 4 | 20 | 26 |  | 39 | , 46 | , 56 | r'07 | I'19 | I 37 |
| 5 | 45 | 53 | r'03 | I'т6 | $\mathrm{I}^{\prime} \mathrm{I}^{\text {I }}$ | I'52 | 221 | 307 | 428 |

For Latitude $25^{\circ}$.

| oh. | $05^{\prime \prime}$ | $10^{\prime \prime}$ | $15^{\prime \prime}$ | $2 \mathrm{I}^{\prime \prime}$ | $27^{\prime \prime}$ | $33^{\prime \prime}$ | $40^{\prime \prime}$ | $48^{\prime \prime}$ | 57 ${ }^{\prime \prime}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 2 | $\bigcirc 8$ | 14 | 19 | 25 | $3{ }^{1}$ | 38 | 46 | ,54 | $\mathrm{I}^{\prime} 05$ |
| 3 | 12 | 18 | 24 | 30 | 37 | 44 | 53 | I'04 | I 18 |
| 4 | 23 | 29 | $\xrightarrow{35}$ | + ${ }^{45}$ | +53 | ${ }^{1} \mathrm{O}$ | $\mathrm{I}^{\prime}$ 5 6 | 131 | I 52 |
| 5 | 49 | 59 | I'ı0 | $\mathrm{I}^{\prime} 24$ | $\mathrm{I}^{\prime} 5^{2}$ | 207 | 244 | 346 | 543 |

For Latitude $27^{\circ} 30^{\prime}$.

| - h. | $08^{\prime \prime}$ | $13^{\prime \prime}$ | $18^{\prime \prime}$ | $24^{\prime \prime}$ | $30^{\prime \prime}$ | $3^{6 \prime}$ | $44^{\prime \prime}$ | $52^{\prime \prime}$ | $\mathrm{I}^{\prime} \mathbf{2}^{\prime \prime}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 2 | 11 | 16 | 22 | 28 | 34 | 41 | 49 | $\mathrm{T}^{\prime} \times$ | 110 |
| 3 | 17 | 22 | 28 | 35 | 42 | 50 | I'oo | 111 | I 26 |
| 4 | 28 | 35 | 42 | ,50 | $\mathrm{r}^{\prime} \mathrm{o}$ | I'ti | I 26 | I 43 | 209 |
| 5 | 54 | 1'05 | I'18 | I'34 | I 54 | 224 | 311 | $43^{8}$ | 815 |

For Latitude $30^{\circ}$.

| - h. | 10" | $15^{\prime \prime}$ | $2 \mathrm{I}^{\prime \prime}$ | $27^{\prime \prime}$ | $33^{\prime \prime}$ | $40^{\prime \prime}$ | $4^{\prime \prime}$ | 57 ${ }^{\prime \prime}$ | $\mathrm{I}^{\prime} 8^{\prime \prime}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 2 | 14 | 19 | 25 | 3 I | 38 | 46 | 54 | r'05 | ¢ 18 |
| 3 | 20 | 26 | 32 | 39 | 47 | ,55 | r'o6 | 119 | I 36 |
| 4 | ,32 | +39 | -46 | , 52 | $\mathrm{r}^{\prime} 06$ | I'19 | I 35 | I 57 | 229 |
| 5 | r'oo | I'то | I'24 | I'52 | 207 | 244 | 346 | 543 | 13 о6 |

For Latitude $32^{\circ} 30^{\prime}$.

| oh. | $13^{\prime \prime}$ | $18^{\prime \prime}$ | $24^{\prime \prime}$ | $30^{\prime \prime}$ | $36^{\prime \prime}$ | $44^{\prime \prime}$ | 52 ${ }^{\prime \prime}$ | $\mathrm{I}^{\prime} \mathrm{I}^{\prime \prime}$ | $\mathrm{I}^{\prime} 14^{\prime \prime}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 2 | 17 | 22 | 28 | 35 | 42 | 50 | $\mathrm{r}^{\prime} \mathrm{o}$ | I 11 |  |
| 3 | 23 | 29 | 35 | 43 | ,51 | r'or | $1{ }^{1} 3$ | 128 | 1 47 |
| 4 | 35 $\times 15$ | +43 | +51 | ${ }^{\text {r }}$ - | I'13 | 127 | I 46 | 213 | 254 |
| 5 | ro3 | I'15 | $\mathrm{I}^{1}{ }^{\text {I }}$ | I 53 | 220 | 305 | 425 | 736 |  |

For Latitude $35^{\circ}$.

| o h.23445 | $\begin{array}{r} 15^{\prime \prime} \\ 20 \\ 26 \\ 39 \\ \mathbf{r}^{\prime \prime} 97 \end{array}$ | $21^{\prime \prime}$253347$\mathrm{r}^{2} 20$ | $\begin{array}{r} 27^{\prime \prime} \\ 32 \\ 39 \\ 56 \\ \mathbf{r}^{\prime} 38 \end{array}$ | $\begin{array}{r} 33^{\prime \prime} \\ 3^{\prime \prime} \\ 47 \\ \mathrm{r}^{\prime} 07 \\ 200 \end{array}$ | $40^{\prime \prime}$46$5^{6}$I'20234 | 48155ro71 36329 | $\begin{aligned} & 57^{\prime \prime} \\ & \mathrm{r}^{\prime \prime} 05 \\ & \text { I } 2 \mathrm{I} \\ & \text { I } 59 \\ & 5 \\ & 59 \end{aligned}$ |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |  |  |  |  |  |
|  |  |  |  |  |  |  |  |  |  |
|  |  |  |  |  |  |  |  |  |  |
|  |  |  |  |  |  |  |  |  |  |

\begin{tabular}{|c|c|c|c|c|c|c|c|c|c|}
\hline 86 \& \multicolumn{9}{|c|}{REFRACTION IN DECLINATION.} \\
\hline \multirow[t]{2}{*}{} \& \multicolumn{9}{|c|}{For Latitude \(37^{\circ} 30^{\prime}\).} \\
\hline \& + \(20^{\circ}\) \& \(+15^{\circ}\) \& \(+10^{\circ}\) \& \(+5^{\circ}\) \& \(0^{\circ}\) \& \(-5^{\circ}\) \& \(-10^{\circ}\) \& \(-15^{\circ}\) \& \(-20^{\circ}\) \\
\hline \[
\begin{aligned}
\& \mathrm{oh} . \\
\& 2 \\
\& 3 \\
\& 4 \\
\& 5
\end{aligned}
\] \& \(\begin{array}{r}188^{\prime \prime} \\ 22 \\ 29 \\ 43 \\ \mathrm{I}^{4} 11 \\ \hline\end{array}\) \& \begin{tabular}{c}
\(24^{\prime \prime}\) \\
28 \\
36 \\
5 S \\
\(\mathrm{I}^{\prime} 26\) \\
\hline
\end{tabular} \& \begin{tabular}{l}
\(30^{\prime \prime}\) \\
35 \\
43 \\
1'01 \\
r 54 \\
\hline
\end{tabular} \&  \& \(\begin{array}{r}44^{\prime \prime} \\ 50 \\ \text { I'O2 } \\ \text { 1 } 27 \\ 249 \\ \hline\end{array}\) \&  \& \begin{tabular}{l} 
I'02' \\
I 12 \\
I 29 \\
214 \\
615 \\
\(\mathrm{I}^{\prime}\) \\
\hline
\end{tabular} \& I' \({ }^{\prime} 4^{\prime \prime}\)
I 26
I 49
254
I4 58 \& I'29

I 45
215
405 <br>
\hline \multicolumn{10}{|c|}{For Latitude $40^{\circ}$.} <br>

\hline $$
\begin{aligned}
& \text { oh. } \\
& 2 \\
& 3 \\
& 4 \\
& 5
\end{aligned}
$$ \& $21^{\prime \prime}$

25
33
47
$1 \times 15$ \& ( ${ }^{27^{\prime \prime}} 3^{32}$ \& $33^{\prime \prime}$
39
48
$\mathrm{I}^{48}$
T 5 5 \&  \&  \&  \&  \& $\mathrm{I}^{\prime} 2 \mathrm{I}^{\prime \prime}$
I 35
202
302
32 I
25 I \& I'39 ${ }^{\prime \prime}$
I 57
236
459 <br>
\hline \multicolumn{10}{|c|}{For Latitude $42^{\circ} 3^{\prime}{ }^{\prime}$.} <br>

\hline $$
\begin{aligned}
& \text { oh. } \\
& 2 \\
& 3 \\
& 4 \\
& 5
\end{aligned}
$$ \& $24^{\prime \prime}$

28
36
36
50
1 6 \& $30^{\prime \prime}$
35
43
r
1
1
3 \& ( ${ }^{366^{\prime \prime}}$ \&  \&  \& $\mathrm{I}^{\prime} \mathbf{O}^{\prime \prime}$
I 12
I 29
2 10
500 \& I'It ${ }^{\prime \prime}$ \& I' $29^{\prime \prime}$
I 45
217
355 \& $\mathrm{I}^{\prime} 49^{\prime \prime}$
211
259
689 <br>
\hline \multicolumn{10}{|c|}{For Latitude $45^{\circ}$.} <br>

\hline $$
\begin{aligned}
& \text { o h. } \\
& 2 \\
& 3 \\
& 4 \\
& 5
\end{aligned}
$$ \& $27^{\prime \prime}$

$3^{2}$
$4{ }^{\circ}$
54
$\mathrm{I}^{\prime} 23$ \& $33^{\prime \prime}$
39
47
r
174
141 \& ( ${ }^{40^{\prime \prime}}$ \& [ ${ }^{48^{\prime \prime}}$ \&  \& I'08
I 19
I $3^{8}$
2 24
540
540 \&  \& I' $39^{\prime \prime}$
I 57
234
438 \& $2^{\prime} 02^{\prime \prime}$
229
329
$8^{29} 15$ <br>
\hline \multicolumn{10}{|c|}{For Latitude $47^{\circ} 30^{\prime}$.} <br>

\hline $$
\begin{aligned}
& 0 \mathrm{~h} . \\
& 2 \\
& 3 \\
& 4 \\
& 5
\end{aligned}
$$ \& $30^{\prime \prime}$

35
43
$5^{\prime} 6$

$\mathrm{r}^{\prime} 27$ \& ( ${ }^{36^{\prime \prime}}$ \& ( ${ }^{44^{\prime \prime}}$ \& | 52 |
| :--- |
| $\mathrm{I}^{\prime \prime} 0$ |
| I 13 |
| I |
| I 40 |
| $25^{2}$ | \& Y'O2'

I 12
I 28
205
401 \& $\mathrm{I}^{\prime} \mathrm{I}^{\prime \prime}$
I 26
I 47
2 40
$63^{0}$ \& I'29
I 45
215
339
I6 39 \& I' $49^{\prime \prime}$
201
256
537 \& 2'18 ${ }^{\prime \prime}$
251
408
4118 <br>
\hline \multicolumn{10}{|c|}{For Latitude $50^{\circ}$.} <br>

\hline \& | $33^{\prime \prime}$ |
| :---: |
| 38 |
| 47 |
| $\mathrm{I}^{4} \mathrm{O}$ |
| I 30 | \& ( ${ }^{40^{\prime \prime}}$ \&  \&  \& | I' $8^{\prime \prime}$ |
| :--- |
| I 88 |
| I 36 |
| 1 16 |
| 422 | \& I' $2 \mathbf{I}^{\prime \prime}$

I 35
229
2 288
728
7 \&  \& $2^{\prime} 02^{\prime \prime}$
228
323
$3^{23} 59$ \& $2^{\prime} 36^{\prime \prime}$
3 19
502
1947 <br>
\hline \multicolumn{10}{|c|}{For Latitude $52^{\circ} 30^{\prime}$.} <br>

\hline \[
$$
\begin{aligned}
& o \mathrm{~h} . \\
& 2 \\
& 3 \\
& 4 \\
& 5
\end{aligned}
$$

\] \& | $36^{\prime \prime}$ |
| :---: |
| 43 |
| 50 |
| I05 |
| I 34 | \& $\begin{array}{r}44^{\prime \prime} \\ 50 \\ \text { 5'00 } \\ \text { I } 18 \\ \text { I } 56 \\ \hline\end{array}$ \&  \& | I'O2' |
| :--- |
| I 17 |
| I 26 |
| 2 10 |
| 316 |
| 10 | \& I'r4 ${ }^{\prime \prime}$

I 26
I 45
228
447

4 \& | I $29^{\prime \prime}$ |
| :--- |
| $\mathrm{I} \mathbf{4}^{\prime 2}$ |
| 2 II |
| $3 \mathrm{I9}$ |
| $85^{2}$ | \& $\mathrm{I}^{\prime} 49^{\prime \prime}$

223
251
251
453 \& $2^{\prime} 188^{\prime \prime}$
249
258
848 \& $3^{\prime} 05^{\prime \prime}$
355
622 <br>
\hline \multicolumn{10}{|c|}{For Latitude $55^{\circ}$.} <br>

\hline $$
\begin{aligned}
& \text { o h. } \\
& 2 \\
& 3 \\
& 4 \\
& 5
\end{aligned}
$$ \& $40^{\prime \prime}$

46
55
I'10

I 37 \& $\begin{array}{r}48^{\prime \prime} \\ 55 \\ \mathrm{r}^{\prime} 06 \\ 123 \\ 201 \\ 201 \\ \hline\end{array}$ \& ( ${ }^{\text {57 }}$ \& | I'08 $8^{\prime \prime}$ |
| :--- |
| I 18 |
| I 35 |
| 206 |
| 3 28 | \& $\mathrm{I}^{\prime} 2 \mathrm{I}^{\prime \prime}$

I 34
I 58
243
4 45
5 \& I' $39{ }^{\prime \prime}$
I 56
230
344
10 48
I \& $2^{\prime} 02^{\prime \prime}$
230
32 I
$3^{21}$
549 \&  \& $3^{\prime} 33^{\prime \prime}$
447
919 <br>
\hline \multicolumn{10}{|c|}{For Latitude $57^{\circ} 30^{\prime}$.} <br>

\hline $$
\begin{aligned}
& \hline \mathrm{o} \mathrm{h.} \\
& 2 \\
& 3 \\
& 4 \\
& 5
\end{aligned}
$$ \& $\begin{array}{r}44^{\prime \prime} \\ 50 \\ 58 \\ \mathrm{I}^{8} \mathrm{II} \\ \mathrm{I} 4 \mathrm{I} \\ \hline\end{array}$ \& $\begin{array}{r}52^{\prime \prime} \\ 59 \\ \mathrm{I}^{\prime} 10 \\ \text { 1 } 25 \\ 206 \\ \hline\end{array}$ \& \[

$$
\begin{aligned}
& \text { I'O2' } \\
& \text { I II } \\
& \text { I } 24 \\
& \text { I } 43 \\
& 242
\end{aligned}
$$
\] \&  \& $\mathrm{I}^{\prime} 29^{\prime \prime}$

I 43
207
250
546

54 \& | $\mathrm{r}^{\prime} 49^{\prime \prime}$ |
| :--- |
| 209 |
| 243 |
| 355 |
| 12 26 | \& $2^{\prime} 18^{\prime \prime}$

247
3 45
614 \& 3'0
$3^{\prime \prime}$
355
550
15
4 \& $4^{\prime} 37^{\prime \prime}$
604
1247 <br>
\hline \multicolumn{10}{|c|}{For Latitude $60^{\circ}$.} <br>

\hline $$
\begin{aligned}
& \mathrm{o} \text { h. } \\
& 2 \\
& 3 \\
& 4 \\
& 5
\end{aligned}
$$ \& \[

$$
\begin{array}{r}
48^{\prime \prime \prime} \\
54 \\
\mathrm{I}^{\prime \prime} 03 \\
118 \\
145
\end{array}
$$

\] \& \[

$$
\begin{array}{r}
57^{\prime \prime} \\
\text { I' }^{\prime \prime} \\
\text { I I5 } \\
\text { I } 34 \\
241
\end{array}
$$

\] \& \[

$$
\begin{array}{ll}
\text { I' }^{\prime} 8^{\prime \prime} \\
\text { I } 17 \\
\text { I } 30 \\
\text { I } 56 \\
2550
\end{array}
$$
\] \& $\mathrm{I}^{\prime} 2 \mathrm{I}^{\prime \prime}$

I 33
I 5 I
2 28
357
357 \& I'39"
I 54
220
2 18

6821 \& $$
\begin{aligned}
& 2^{\prime} 02^{\prime \prime} \\
& 224 \\
& 304 \\
& 304 \\
& 450 \\
& 1533^{2}
\end{aligned}
$$ \& $2^{\prime} 36^{\prime \prime}$

3 I2
424
484
853 \& $3^{\prime} 33^{\prime \prime}$
$43^{8}$
$733^{1}$ \& $5^{\prime} 23^{\prime \prime}$
815
2444 <br>
\hline
\end{tabular}

# TABLES 

FOR OBTAINING

## HORIZONTAL DISTANCES <br> AND

DIFFERENCES OF LEVEL,
FROM
STADIA READINGS.

| 88 | DISTANCES. |  |  |  |  |  |  |  |  | $0^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | a |
| $\bigcirc$ | 0.9986 | 1.9972 | 2.95 | 3.99 | 4.9 |  | 6.9 |  |  |  |
| or | o. 99 | 1. 9972 | 2.99 | 3.9944 | 4.99 |  |  |  |  | I. 4000 |
| 02 | 0. 99 | 1. 9972 | 2.99 | 3.9944 | 4.9930 | 5.9916 | 6.9902 |  |  | I. 4000 |
| 03 | - 99 | 1. 9972 | 2.99 | 3.9944 | 4.9930 | 5.9916 |  |  |  | r. 4000 |
| $\bigcirc 4$ | 0.938 | 1. 9972 | 2.9958 | 3.9944 | $4.933{ }^{\circ}$ | 5.9916 | 6.9902 | 7.9888 | 8.9874 | I. 4000 |
| 05 | 0.99 | 1.9972 | 2.9958 | 3.9944 | 4.9930 | 5.9916 | 6.9902 |  |  | 1.4000 |
|  | -. 998 | I. 9972 | ${ }^{2.9958}$ | 3.9944 | $4.933^{\circ}$ | 5.9916 | 6.9902 |  | 8.9874 | I. 4000 |
| $\bigcirc$ | 0.99 | 1.9972 | 2.99 | 3.9944 | $4.993{ }^{\circ}$ | 5.9916 | 6.9932 |  | 8.9873 | I. 4000 |
| -8 | 0.9 | 1. 9972 | 2.95 | 3.9944 | $4.993{ }^{\circ}$ | 5.9916 |  |  | 8.9873 | I.4000 <br> I. 4000 |
| Io | 0.9986 | 1. 9972 | 2.9958 | 3.9944 | $4.993^{\circ}$ | 5.9916 | 6.9901 |  |  | I. 4000 |
| II |  | 72 | 2.9958 |  | 4.9930 |  | 6.9901 |  |  |  |
| 12 |  | I. 9972 |  | 3.99 | 4.99 |  |  |  |  | 00 |
| 13 | 0.99 | 1. 9972 | 2.9958 | 3.9943 | 4.9929 | 5.9915 | 6990 |  |  | 1.4000 |
| 14 | 0.99 | 1. 9972 | 2.9957 | 3.9943 | 4.9929 | 5.9915 | 6.9901 |  |  | 1.4000 |
| 15 | 0.9986 | 1.9972 | 2.9957 | 3.9943 | 4.9929 | 5.9915 | 6.9901 | 7. | 8.9872 | I. 4000 |
| 16 | 0.99 | 1. 9972 | 2.9957 | 3.9943 | 4.9929 | 5.9915 | 6.9900 |  |  | I.4000 |
| 17 | -. 998 | 1.9972 | 2.9957 | 3.9973 | 4.9929 | 5.9915 | 6.9900 | 7.9 | 8.9872 | I. 4000 |
| 18 | 0.9986 | I.997I | 2.9957 | 3.9743 | 4.9929 | 5.9914 | 6.9900 |  |  | I. 4000 |
| 19 | 0.5 | 1.9971 | 2.9957 | 3.9943 | 4.99 | 5.9914 |  |  |  | I. 4000 |
| 20 | 0.9986 | 1.9971 | 2.9957 | 3.9943 | 4.9928 | 5.9914 | 6.9900 | 7.9885 | 8.9871 | 1.4000 |
| 21 |  | I. 9 | 2.9957 | 3.99 | 4.9928 | 5.9914 | 6. 9899 |  |  |  |
| 22 |  | I.9971 | 2.995 | 3.99 | 4.9928 | 5.9913 |  |  |  |  |
| 23 | 0.9 | 1.9971 | 2.9957 | 3.9942 | 4.9928 | 5.9913 |  |  |  | I. 3999 |
| 24 | 0.998 | 1.997I | 299 | 3.9942 | 4.9927 | 5.9913 |  | 7. |  | I. 3999 |
| 25 | 0.99 | 1.9971 | 2.9956 | 3.9942 | 4.9927 | 5.9913 |  |  |  | I. 3999 |
|  | 0.998 | 1.9971 | 2.9956 | 3.9942 | 4.9927 | 5.9912 |  |  |  | I. 3999 |
| 27 |  | I. | 2.95 | 3.99 | 4.9 | 5.99 |  |  |  |  |
| 29 | 0.9 | I.9971 | 2.99956 2.995 | 3.9941 3.9941 | 4.9927 4.9926 | 5.9912 5.9912 | 6.9897 |  |  |  |
| 30 | 0.978 | I. 9970 | 2.9956 | 3.994 | 4026 | 5.99 | 6.9897 | 7.9882 | 8.9867 | I. 3999 |
| 31 |  | I. 9970 | 2.99 | 3.9941 | 4.9 | 5.9 |  |  |  |  |
| 32 | 0.99 | 1.9970 | 2.9955 | 3.9940 | 4.9026 | 5.9911 |  |  |  |  |
| 33 |  |  |  |  |  | 5.9 |  |  |  |  |
| 35 |  | 1.9970 | 2.9955 | 3.9940 | 4.9925 4.9925 | 5.9910 | 6.9895 |  | 8.9865 | I. 3999 |
| 5 | 0.99 | 1.9970 | 2.9955 | 3.9940 | 4.9924 | 5.9909 |  | 7.9879 |  | I. 3999 |
| 37 |  | 1.9970 | 2.9954 | 3.9939 | 4.9924 | 5.9999 |  |  |  | I. 3999 |
| 38 |  | 1.99 | 2.9954 | 3.99 | 4.9924 | 5.9909 |  |  |  | I. 3999 |
| 39 |  | I. 99 | 2.9954 | 3.9939 | 4.9924 |  |  |  |  | I. 3999 |
| 40 | 8 | 1.9969 | 2.9954 | 3.9939 | 4.9923 | 5.990 | 6.9893 | 7.9877 | 8.98 | I. 3 |
| 41 |  | 1. 9969 | 2.9954 |  |  |  |  |  |  |  |
| 42 | 0.99 | 1. 9969 | 2.9953 | 3.9938 | 4.9922 | 5.9907 |  |  |  |  |
| 43 | -0.99 | I. 9996 | 2.9953 | 3.9938 | 4.9922 | 5.9907 |  |  |  |  |
|  |  | I. | 2.9953 2.9953 2 | 3.9 | 4.9 | 5.9906 |  |  |  |  |
| 46 | O. 9 | I. 996 | 2.9953 | 3.9937 | 4.992 I | 5.9905 | 6.9889 | 7.9874 | 8. 9 |  |
| 47 | O. | I. 99 | $2.995^{2}$ | 3.99 | 4.9921 | 5.9905 |  | 7.9873 |  |  |
| 48 |  | I. 99 | 2.9952 | 3.9936 | 4.9920 | 5.9904 |  |  |  |  |
| 49 50 | 0.998 0.998 | I. 9968 1.9968 | 2.9952 2.9952 | 3.9936 3.9936 | 4.9920 4.9919 | 5.9904 5.9903 | $\begin{array}{\|l\|l} 6.9888 \\ 6.9887 \end{array}$ | $\begin{array}{\|l\|l} 7.9872 \\ 7.9877 \end{array}$ | $\begin{aligned} & 8.9856 \\ & 8.9855 \end{aligned}$ | $\begin{array}{\|l\|l\|} \text { I. } 3998 \\ \text { r. } 3998 \end{array}$ |
| 51 |  |  |  |  |  |  |  |  |  |  |
| 52 |  | I. 99 | 2.9951 | 3.9935 | 4.9919 | 5.9902 |  |  |  |  |
| 53 | 0.99 | 1.9967 | 2.9951 | 3.9934 | 4.9918 | 5.9902 | 6.9885 | 7.9869 | 8.9852 | I. 3998 |
|  | 0.99 | I. 996 | 2.9951 | 39934 | 4.9918 | 5.9901 |  |  | 8.9852 |  |
| 55 | 0.9 | I. 9 | 2.9950 | 3.9934 | 4.9917 | 5.9901 | 4 | 7 | 51 | 1.3998 |
| 56 | 0.9 | I. 9 | 2.9950 | 3.9933 | 4.9917 | 5.9900 | 3 | 7.9867 | 8.9850 | I. 3998 |
| 57 | 0.99 | I. 9 | 2.9950 | 3.9933 | 4.99 |  |  |  |  | 8 |
| 58 | 0.9 | 1.9966 | 2.9949 | 3.993 | 4.9916 |  | 6.9882 | 5 | 4 | 98 |
|  | 0.99 | 1.9966 | 2.9949 | 3.9932 | 4.9915 | 5.98 | 6.9881 | 4 | 8.9847 | 98 |
| 60 | 0.9983 | I. 9966 | 2.9949 | 3.9932 | 4.9915 | 5.98 | 6.988 | 7.98 | 8.98 | I. 3998 |


| $0^{\circ}$ | HEIGHTS. |  |  |  |  |  |  |  | 89 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | b |  |
| 0. | 0.0000 | 00 | oo | 00 | 0.0000 | 0.0000 | 0.0000 | . 0000 | 00 | oo |
| 0.0003 | 0.0006 | 0.0009 | 0.0012 | 0.0015 | 0.0017 | 0.0020 | 0.0023 | 0.0026 | 0.0004 | or |
| 0.0006 | o. | 0.0017 | 0.0023 | 0.0029 | 0.0035 | 0.0041 | 0.0046 | 0.0052 | 0.0008 | 02 |
| 0.0009 | 0.0017 | 0.0026 | 0.0035 | 0.0044 | 0.0052 | 0.006I | 0.0070 | 0.0078 | 0.0012 | 03 |
| 0.0012 | 0.0023 | 0.0035 | 0.0046 | 0.0058 | 0.0070 | 0.008I | 0.0c93 | o.0105 | 0.0016 | 04 |
| 0.0015 | 0.0029 | 0.0044 | 0.0058 | 0.0073 | 0.0087 | 0.0102 | 0.0116 | 0.013 ${ }^{\text {r }}$ | 0.0020 | 5 |
| 0.0017 | 0.0035 | 0.0052 | 0.0070 | 0.0087 | -. 0105 | 0.012 | 0.01 39 | -.0157 | 0.0024 | 06 |
| 0.0020 | 0.004I | 0.0061 | 0.0081 | 102 | 9.0122 | 0.0142 | 0.0163 | o.0183 | 0.0029 | 07 |
| 0.0023 | 0.0046 | 0.0070 | 0.0093 | 0.0116 | -.0139 | -.0163 | 0.0186 | 0.0209 | .0033 | 08 |
| 0.0026 | 0.0052 | 0.0078 | 0.0105 | 0.0131 | -.0157 | 0.0183 | 0.0209 | 0.0235 | 0.0037 | 99 |
| 0.0029 | 0.0058 | 0.0087 | 0.0116 | 0.0145 | -.or 74 | 0.0203 | 0.0232 | 0.026 r | 0.0041 | 10 |
| 0.0032 | 0.0064 | 0.0096 | . 128 | 0.0160 | 0.0192 | 0.0224 | 0.0256 | 0.0288 | 0045 | II |
| 0.0035 | 0.0070 | o.0105 | 0.0139 | 0.0174 | 0.0209 | 0.0244 | 0.0279 | 0.0314 | 0.0049 | 12 |
| 0.0038 | 0.0076 | 0.0113 | 0.0151 | 0.0189 | 0.0227 | 0.0264 | 0.0302 | 0.0340 | 0.0053 | 13 |
| 0.004 T | 0.0081 | 0.0122 | 0.0163 | 0.0203 | 0.0244 | 0.0285 | 0.0325 | 0.0366 | 0.0057 | 14 |
| 0.0044 | 0.0087 | 0.0131 | 0.0174 | 0.0218 | 0.0261 | 0.0305 | 0.0349 | 0.0392 | 0.0061 | 15 |
| 0.0046 | 0.0093 | -.0139 | 0.0186 | 0.0232 | 0.0279 | 0.0325 | 0.0372 | 0.0418 | 0.0065 | 16 |
| 0.0049 | 0.0099 | o.0148 | 0.0198 | 0.0247 | 0.0295 | 0.0346 | 0.0395 | 0.0444 | 0.0069 | 17 |
| 0.0052 | 0.0105 | -.or 57 | 0.0209 | 0.0261 | 0.0314 | 0.0366 | 0.0418 | 0.0471 | 0.0073 | 18 |
| 0.0055 | 0.0110 | 0.0166 | 0.022 I | 0.0276 | 0.0331 | 0.0386 | 0.0442 | 0.0497 | 0.0077 | 19 |
| 0.0058 | 0.0116 | 0.0174 | 0.0232 | 0.0290 | 0.0349 | 0.0407 | 0.0465 | 0.0523 | 0.008I | 20 |
| 0.0061 | 0.0122 | 0.0183 | 0.0244 | 0.0305 | 0.0366 | 0.0427 | 0.0488 | 0.0549 | 0.0086 | 21 |
| 0.0064 | 0.0128 | 0.01 | 0.0256 | 0.0320 | 0.0383 | 0.0447 | 0.0511 | 0.0575 | 0.0090 | 22 |
| 0.0067 | 0.013 | 0.0200 | 0.0267 | 0.033 | 0.0401 | 0.0468 | 0.0534 | 0.060I | 0.0094 | 23 |
| 0.00 | 0.013 | 0.0209 | 0.0279 | 0.0349 | 0.0418 | 0.0488 | 0.0558 | 0.0627 | 0.0098 | 24 |
| 0.0 | 0.0145 | 0.0218 | 0.0290 | 0.036 | 0.0436 | 0.0508 | 0.0581 | 0.0654 | , 102 | 25 |
| 0.0076 | 0.015 | 0.0227 | 0.0302 | 0.0378 | -0.0453 | 0.0529 | 0.0604 | 0.0680 | 0.0106 | 26 |
| 0.0078 | 0.0157 | 0.0235 | 0.0314 | 0.0392 | 0.047 I | 0.0549 | 0.0627 | 0.0706 | 0.0110 | 27 |
| 0.0081 | 0.0 | 0.0244 | 0.0325 | 0.0407 | 0.0488 | 0.0569 | 0.0651 | 0.0732 | 0.0114 | 28 |
| 0.0084 |  | 0.0253 | 0.0337 | 0.0421 | 0.0505 | 0.0590 | 0.0674 | 0.0758 | 0.01 | 29 |
| 0.0087 | 0.0174 | 0.026I | 0.0349 | 0.0436 | 0.0523 | 0.0610 | 0.0697 | 0.0784 | 0.0 | 30 |
| 0. | 0.0180 | 0. | 0.0360 | 0. | 0. | 0.0630 | 0.0720 | 0.08 | O12 | 31 |
| 0.0093 | 0.01 | 0.02 | 0.0372 | 0.0465 | 0.0558 | 0.0651 | 0.0744 | 0.0837 | 0.0130 | 32 |
| 0.0096 | 0.0192 | 0.0288 | 0.0383 | 0.0479 | 0.0575 | 0.0671 | 0.0767 | 0.0863 | 0.0134 | 33 |
| 0.0099 | 0.0198 | 0.0296 | 0.0395 | 0.0494 | 0.0593 | 0.0691 | 0.0790 | 0.0889 | 0.0138 | 34 |
| 0.0102 | 0.0203 | 0.0305 | 0.0407 | 0.0508 | 0.0610 | 0.0712 | 0.0813 | 0.0915 | 0.0143 | 35 |
| 0.0105 | 0.0209 | 0.0314 | 0.0418 | 0.0523 | 0.0627 | 0.0732 | 0.0836 | 0.0941 | 0.0147 | 36 |
| 0.0107 | 0.0215 | 0.0322 | 0.0430 | 0.0537 | 0.0645 | 0.0752 | 0.0860 | 0.0967 | 0.0151 | 37 |
| 0.0110 | 0.0221 | 0.0331 | 0.0441 | 0.05 | 0.066 | 0.0773 | 0.0883 | 0.0993 | 0.0155 |  |
| 0.0113 | 0.0227 | 0.0340 | 0.0453 | 0.0566 | 0.0680 | 0.0793 | 0.0906 | -. 1019 | -.0159 | 39 |
| 0.0116 | 0.0232 | 0.0349 | 0.0465 | 0.058 I | 0.0697 | 0.0813 | 0.0929 | 0.1046 | 0.0163 | 40 |
| 0.01I9 | 0.0238 | 0.0357 | 0.0476 | 0.0595 | 0.0715 | 0.08.34 | 0.0953 | 0.1072 | 0.0167 | 41 |
| 0.0122 | 0.0244 | 0.0366 | 0.0488 | 0.0610 | 0.0732 | 0.0854 | 0.0976 | 0.1098 | 0.0171 | 42 |
| 0.0125 | 0.0250 | 0.0375 | 0.0500 | 0.0624 | 0.0749 | 0.0874 | 0.0999 | -. 1124 | 0.0175 | 43 |
| 0.012 | 0.0256 | 0.0383 | 0.0511 | 0.063 | 0.0767 | 0.0895 | 0. 1022 | 0.1150 | 0.0179 | 44 |
| 0.0131 | 0.0261 | 0.0392 | 0.0523 | 0.06 | 0.0784 | 0.0915 | 0.1046 | 0. 1176 | 0.0183 | 45 |
| 0.0134 | 0.0267 | 0.0401 | 0.0534 | 0.0668 | 0.0802 | 0.0935 | 0. 1069 | 0. 1202 | 0.0187 | 46 |
| 0.0137 | 0.0273 | 0.0410 | 0.0546 | 0.0683 | 0.0819 | 0.0956 | 0.1092 | 0. 1229 | 0.0191 | 47 |
| 0.0139 | 0.0279 | 0.0418 | 0.0558 | 0.0697 | 0.0836 | 0.0976 | 0.1115 | 0. 1255 | 0.0195 | 48 |
| 0.0142 | 0.0285 | 0.0427 | 0.0569 | 0.0712 | 0.0854 | 0.0996 | -. 1138 | 0. 1281 | 0.0200 | 49 |
| 0.0145 | 0.0290 | 0.0436 | 0.0581 | 0.0726 | 0.0871 | 0.1017 | 0.1162 | 0.1307 | 0.020 | 50 |
| 0.0148 | 0.0296 | 0.0444 | 0.0592 | 0.0741 | 0.0889 | 0. 1037 | 0.1185 | 0. 1333 | 0.0208 | 51 |
| 0.0151 | 0.0302 | 0.0453 | 0.0604 | 0.0755 | 0.0906 | 0. 1057 | 0.1208 | 0. 1359 | 0.0212 | 52 |
| 0.0154 | 0.0308 | 0.0462 | 0.0616 | 0.0770 | 0.0923 | 0. 1077 | 0.1231 | 0. 1385 | 0.0216 | 53 |
| 0.0157 | 0.0314 | 0.0470 | 0.0627 | 0.0784 | 0.0941 | -. 1098 | 0. 1254 | C. 1411 | 0.0220 | 54 |
| 0.0160 | 0.0319 | 0.0479 | 0.0639 | 0.0799 | 0.0958 | 0.1118 | 0.1278 | 0. 1437 | 0.0224 | 55 |
| 0.0163 | 0.0325 | 0.0488 | 0.0650 | 0.0813 | 0.0976 | o. 1138 | 0.1301 | 0. 1463 | 0.0228 | 57 |
| 0.0166 0.0168 | 0.0331 0.0337 | 0.0497 | 0.0662 0.0674 | 0.0828 0.0842 | 0.0993 0.1011 | o.i159 0.1179 | 0.1324 0.1348 0.151 | 0.1490 0.1516 | $0.0232$ | 57 |
| 0.0168 0.0171 | 0.0337 0.0343 | 0.0505 | 0.0674 | 0.0842 0.0857 | O. 1011 0.1028 | 0.1179 0.1199 | 0.1348 0.1371 | 0.1516 0.1542 | $\left\|\begin{array}{l} 0.0236 \\ 0.0240 \end{array}\right\|$ | 59 |
| 0.0174 | 0.0349 | 0.0523 | 0.0697 | 0.0871 | 0.1046 | 0.1220 | 0.1394 | 0.1568 | 0.0244 | 60 |




|  | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | a |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 00 |  |  |  |  |  |  |  |  |  |  |
| OI | 0.9974 | 1.9947 | 2.59 | 3.9895 |  |  |  |  |  | - |
| 02 | 0.9973 | 1. 9947 | 2.992 | 3.9894 |  | 5.9841 | 6.9814 | 7.9787 |  | . 399 |
| $\bigcirc 3$ | 0.9973 | ז. 9946 | 2.9920 | 3.9893 |  |  | 6.9812 | 7.9786 |  | 399 |
| 04 | 0.9973 | 1. 9946 | 2.9919 | 3.9892 | 4.9865 | 5.9838 | 6.98 II | 7.9784 | 8.9757 | 399 |
| $\bigcirc 5$ | 0.9973 | 1. 9946 | 2.9918 | 3.9891 | 4.9864 | 5.9837 | 6.9810 | 7.9782 | 8.9755 | 399 |
| c6 | 0.9973 | 1. 9945 | 2.9918 | 3.9890 | 4.9863 | 5.9835 | 6.9808 | 7.978I | 8.9753 | 399 |
| 07 | 0.9972 | I. 9945 | 2.971 | 3.9889 | 4.9862 | 5.9834 | 6.9807 | 7.9779 | 8.975 | 399 |
| O8 | 0.9972 | 1. 9944 | 2.9916 | 3.9889 | 4.986 I | 5.9833 | 6.9805 | 7.9777 | 8.9749 | 399 |
| 09 | 0.9972 | 1. 9944 | 2.9916 |  | 4.9860 | 5.9832 | 6.9804 | 7.9776 |  | 399 |
| Io | 0.9972 | 1.9943 | 2.9915 | 3.9887 | 4.9859 | 5.9830 | 6.9802 | 7.9774 |  | I. 399 |
| II |  |  |  |  |  |  |  | 772 |  |  |
| 12 | 0.9971 | 1. 9343 | 2.9914 | 3.9885 | 4.9856 | 5.9828 | 6.9799 | 7.9770 |  | 999 |
| 13 | 0.9971 | I. 9942 | 2.9913 | $3.988{ }_{4}$ | 4.9855 | 5.9826 | 6.9797 | 7.9768 | 8.9739 | 99 |
| 14 | 0.9971 | I. 9942 | 2.9912 | 3.9883 | 4.985 | 5.9825 | 6.9796 | 7.9767 | 8.9737 | 99 |
| 15 | 0.9971 | I. 9941 | 2.9912 | 3.9882 | 4.9853 | 5.9824 | 6.9794 | 7.9765 | 8.9735 | 990 |
| 16 | 0.9970 | I. 9941 | 2.99 II | 3.988I | 4.985 | 5.9822 | 6.9793 | 7.9763 | 8.9733 | I. 3990 |
| 17 | 0.9970 | I. 9940 | 2.9910 | 3.988I | 4.9851 | 5.982I | 6.9791 | 7.976 r | 8.9731 | 1.3989 |
| 18 | 0.9970 | I. 9940 | 2.9910 | 3.9880 | 4.9850 | 5.9819 | 6.97 | 7.9759 | 8.9729 | I. 3989 |
| 19 | 0.9970 | 1. 9939 | 2.9909 | 3.987 | 4.9848 | 5.9818 | 6.9 | 7.9757 | 8.9727 | I. 3989 |
| 20 | 0.9969 | I. 9939 | 2.9908 | 3.937 | 4.9847 | 5.9817 | 6.9786 | 7.9756 | 8.9725 | 988 |
| 2 I |  |  |  |  |  |  |  | 54 |  |  |
| 22 | 0.996 | I. 993 | 2.990 | 3.9876 | 4.9845 | 5.9814 |  | 7.9752 | 8.9721 |  |
| 23 | 0.996 | I. 9937 | 2.9905 | 3.9875 | 4.9844 | 5.9812 | 6.978 | 7.9750 | 8.9718 |  |
| 24 | 0.9968 | I. 9937 | 2.9905 | 3.9874 | 4.9842 | 5.981 | 6.977 | 7.9748 | 8.9716 | 88 |
| 25 | 0.9968 | 1.9936 | 2.9905 | 3.9873 | 4.984 I |  | 6.97 | 7.9746 | 8.9714 | 88 |
| 26 | 0.9968 | 1. 9936 | 2.9904 | 3.9872 | 4.9840 |  | 6.977 | 7.974 | 8.9712 | S8 |
| 27 | 0.9968 | I. 9935 | 2.9903 | 3.9871 | 4.98 | 5.9806 | 6.9774 | 7.9742 | 8.9710 |  |
|  | 0.9967 | I. 9935 | 2.9902 | 3.9870 |  | 5.9805 | 6.9772 | 7.974 | 8.9707 |  |
| 29 | 0.9967 | 1.9934 | 2.9902 |  |  | 5.9803 | 6.977 | 7.9738 | 8.9705 |  |
| 30 | 0.9967 | I. 9934 | 2.9901 | 3.9868 | 4.9835 | 5.9802 |  | 7.9736 | 8.9703 | 1. 3987 |
| 3 I |  | 1.9933 |  |  |  |  |  | 7.9734 |  |  |
| 32 | 0.9956 | I. 9933 | 2.98 | 3.9866 | 4.9832 | 5.9799 |  | 7.9732 |  |  |
| 33 | 0.9966 | 1.9932 |  | 3.9865 | 4.9831 | 5.97 | 6.9764 | 97 | 8.9696 |  |
| 34 | 0.996 | 1.9932 | 2.98 | 3.9864 | 4.9830 | 5.9796 | 6.9762 | 7.972 | 8.9694 |  |
| 35 | 0.996 | I. 9931 | 2.98 | 3.9863 | 4.9828 | 5.9794 | 6.9760 | 7.9726 | 8.9691 | . 3986 |
| 36 | 0.996 | 1.9931 | 2.9896 | 3.9862 | 4.9827 | 5.9793 | 6.9758 | . 9723 | 8.9689 | S086 |
| 37 | 0.9965 | 1.9930 | 2.9896 |  |  | 5.9791 | 6.9756 | 9721 |  |  |
| 3 | 0.9965 | I. 9930 | 2.9895 | 3.986 | 4.9825 |  | 6.9754 | .9719 |  |  |
| 39 | 0.995 | I. 9929 | 2.9894 |  | 4.9823 |  | 6.9753 | 7.9717 |  |  |
| 40 | 0.9954 | 1.9929 | 2.9893 |  |  |  | 6.975 I | 7.9715 | 8.080 |  |
| 4 I |  |  |  |  |  |  |  |  |  |  |
|  | 0.996 | 1.9928 | 2.98 | 3.9855 | 4.98 | 5.9783 | 6.9747 | 7.9711 |  |  |
| 43 | 0.996 | 1.9927 | 2.989 | 3.9854 | 4.981 | 5.978 I | 6.9745 | 7.9708 |  | . 39 |
| 44 | 0.9963 | I. 9927 | 2.989 | 3.9853 | 4.9816 | 5.9780 | 6.9743 | 7.9706 |  | I. 3984 |
| 45 | 0.9963 | I. 9926 |  | 3.9852 | 4.9815 | 5.9778 | 6.9741 | 7.9704 |  |  |
| 46 | 0.9 | I. 9925 |  | 3.9851 | 4.9814 | 5.9776 | 6.9739 | 7.9702 |  | I |
| 47 | 0.9 | I. 9925 |  | 3.9850 | 4.9812 | 5.9775 | 6.9737 | 7.9700 |  | - |
| 48 | 0.9 | I. 9924 |  | 3.98 | 4.9 | 5.9773 | 6.9735 | 7.9697 |  | - |
| 49 | 0.9 | I. 9924 | 2.9886 | 3.9848 |  | 5.9771 | 6.9733 | 7.9695 | 8.965 | I. 3983 |
| 50 | 0.99 | 1.9923 |  | 3.9846 | 4.9808 | 5.9770 | 6.9731 | 7.9693 | 8.9654 |  |
| 51 |  |  |  |  |  |  |  |  |  |  |
| 52 | 0.9 | I. 9922 | 2. | 3.9844 | 4.9505 | 5.9766 | 6.9727 | 7.9688 | 8.9649 | 3 |
| 53 | 0.9 | 1. 9921 | 2.9 | 3.9843 | 4.9804 | 5.9764 | 6.9725 | 7.9685 | 8.9646 | I. 3983 |
|  | -. 9 | I. 992 | 2.9 | 3. | 4.9802 | 5.9763 | 6.9723 | 7.9683 | 8.9644 |  |
|  |  | 1. 9920 | 2.9 | 3.98 | 4.98 or | 5.9761 | 6.9721 | 7.968 I | 8.9641 | I. 3982 |
|  |  | I. 992 |  | 3. | 4.979 | 5.9759 | 6.9719 | 7.9679 |  | 1. 3982 |
|  | 0.996 | I. 9 |  | 3. | 4.979 | 5.9757 | 6.9717 | 7.9676 | 8.9636 |  |
|  | 0.9959 | I. 9 | 2. |  | 4.9796 | 5.9756 | 6.9715 | 7.9674 | 8.9633 |  |
|  | 0.9959 | 1.9918 | 2. | 3. | 4.9795 | 5.9754 | 6.9713 | 7.9672 |  |  |
|  | 0.9 | I. 9 | 2. | 3. |  | 5.975 | 6 | 7.9 | 8.9628 | I.398I |




| $3^{\circ}$ | HEIGHTS. |  |  |  |  |  |  |  | 95 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | b |  |
|  |  | 0.1566 | 0.2088 | 0.2610 | 0.3131 | 0.3653 |  | 0.4697 | 0.0733 |  |
| 0.052 | 0.1050 | 0. 1574 | 0.2099 | 0.2624 | 0.3149 | 0.3674 | 0.4198 | 0.4723 | 0.0737 | or |
| 0.0528 | 0. 1055 | 0.1583 | 0.2111 | 0.2638 | 0.3166 | 0.3694 | 0.4222 | 0.4749 | 0.074 x | 02 |
| 0.0531 | 0.1061 | 0.1592 | 0.2122 | 0.2653 | 0.3184 | -. 3714 | 0. 4245 | 0.4775 | 0.0745 | 03 |
| 0.0533 | 0.1067 | 0.1600 | . 2134 | 0.2667 | 0.3201 | 0.3734 | 0.4268 | 0.4801 | 0.0749 | 04 |
| 0.0536 | 0. 1073 | 0.1609 | 0.2145 | 0.2682 | 0.3218 | 0.3754 | 0.4291 | 0.4827 | 0.0753 | 5 |
| 0.0539 | -. 1078 | -.1618 | 0.2157 | 0.2696 | 0.3235 | 0.3774 | 0.4314 | 0.4853 | 0.0757 | 06 |
| 0.0542 | -. 1084 | 0.1626 | 0.2168 | 0.2711 | 0.3253 | 0.3795 | 0.4337 | 0.4879 | 0.0761 | 07 |
| 0.0545 | 0.1090 | 0. 1635 | 0.2180 | 0.2725 | 0.3270 | 0.3815 | 0.4360 | -. 4905 | 0.0765 | 8 |
| 0.0548 | 0.1096 | 0.1644 | 0.2192 | 0.2739 | 0.3287 | 0.3835 | 0.4383 | 0.493I | 0.0769 | 9 |
| 0.0551 | 0.1102 | 0.1652 | 0.2293 | 0.2754 | 0.3305 | 0.3856 | 0.4406 | 0.4957 | 0.0773 | ro |
| 0.0554 | 0.1107 | 0.1661 | 0.2215 | 0.2768 | 0.3322 | 0.3876 | 0.4430 | 0.4983 | 0.0777 | 11 |
| 0.0557 | 0.1113 | 0.1670 | 0.2226 | 0.2783 | 0.3340 | 0.3896 | 0.4453 | 0.5009 | 0.0781 | 12 |
| 0.0559 | -.1119 | 0.1678 | 0.2238 | 0.2797 | 0.3356 | 0.3916 | 0.4475 | 0.5035 | 0.0786 | 13 |
| 0.0562 | 0.1125 | 0.1687 | 0.2249 | 0.2812 | 0.3374 | 0.3936 | 0.4498 | 0.5061 | 0.0790 | 14 |
| 0.0565 | 0.1130 | 0.1696 | 0.2261 | 0.2826 | 0.3391 | 0.3956 | 0.4522 | 0.5087 | 0.0794 | 15 |
| 0.0568 | 0.1136 | 0.1704 | 0.2272 | 0.2841 | 0.3409 | 0.3977 | 0.4545 | 0.5113 | 0.0798 | 16 |
| 0.0571 | 0.1142 | 0.1713 | 0.2284 | 0.2855 | 0.3426 | 0.3997 | 0.4568 | 0.5139 | 0.0802 | 17 |
| 0.0574 | 0.1148 | 0.1722 | 0.2296 | 0.2869 | 0.3443 | 0.4017 | 0.4591 | 0.5165 | 0.0806 | 18 |
| 0.0577 | 0.1154 | 0. 1730 | 0.2307 | 0.2884 | 0.3461 | 0.4038 | 0.4614 | 0.5191 | 0.0810 | 19 |
| 0.0580 | 0.1159 | 0. 1739 | 0.2319 | 0.2898 | 0.3478 | 0.4058 | 0.4638 | 0.5217 | 0.08I4 | 20 |
| 0.0583 | 0.1165 | 0.1748 | 0.2330 | 13 | 0.3495 | 0.4078 | 0.4660 | 0.5243 | 0.0818 | 21 |
| 0.0585 | 0.1171 | 0.1756 | 0.2342 | 0.2927 | 0.3512 | 0.4098 | 0.4683 | 0.5269 | 0.0822 | 22 |
| 0.0588 | -. 1177 | 0.1765 | 0.2353 | 0.29 | 0.3530 | 0.4118 | 0.4706 | 0.5295 | 0.0826 | 23 |
| 0.0591 | 0.1182 | 0.1774 | 0.2365 | 0.2956 | 0.3547 | 0.4138 | 0.4730 | 0.5321 | 0.0830 | 24 |
| 0.0594 | 0.1188 | 0.1782 | 0.2376 | 0.2971 | 0.3565 | 0.4159 | 0.4753 | 0.5347 | 0.0834 | 25 |
| 0.0597 | -.1194 | 0.1791 | 0.2388 | 0.2985 | 0.3582 | 0.4179 | 0.4776 | 0.5373 | 0.0838 | 26 |
| 0.0600 | 0.1200 | 0.1799 | 0.2399 | 0.2999 | 0.3599 | 0.4199 | 0.4799 | 0. 5399 | 0.0842 | 27 |
| 0.0603 | 0. 1205 | 0.1808 | 0.241 I | 0.3014 | 0.3616 | .0.4219 | 0.4822 | 0.5425 | $0.0847$ | 28 |
| 0.0606 | 0.12 | 0.1817 | 0.2422 | 0.3028 | 0.3634 | 0.4239 | 0.4845 | 0.545 | $0.0851$ | 29 |
| 0.06c8 | 0.1217 | 0.1825 | 0.2434 | 0.3042 | 0.3651 | 0.4259 | 0.4868 | 0.5477 | 0.0855 | 30 |
| 0.0611 | 0. 1223 | 0. 1834 |  | 0.3057 |  | 0.4280 |  | 0.5503 |  | 3 I |
| 0.0614 | 0.1229 | 0.1843 | 0.2457 | 0.3071 | 0.3686 | 0.4300 | 0.4914 | $0.5529$ | $0.0863$ | 32 |
| 0.0617 | 0. 1234 | 0.1851 | 0.2468 | 0.3086 | 0.3703 | 0.4320 | 0.4937 | 0.5554 | 0.0867 | 33 |
| 0.0620 | $0.1240$ | 0.1860 | 0.2480 | 0.3100 | 0.3720 | 0.4340 | 0.4960 | 0.5580 | 0.087 I | 34 |
| $0.0623$ | 0.1246 | 0.1869 | 0.2492 | 0.3115 | -. 3737 | 0.4360 | 0.4983 | 0.5606 | $0.0875$ | 35 |
| $0.0626$ | 0. 1252 | 0. 1877 | 0.2503 | 0.3129 | 0.3755 | 0.438 I | 0.5006 | 0.5632 | 0.0879 | 36 |
| $0.0629$ | $0.1257$ | 0. 1886 | 0.2515 | 0.3143 | 0.3772 | 0.4401 | 0.5030 | 0.5658 | 0.0883 | 37 |
| $0.0632$ | 0.1263 | 0. 1895 | 0.2526 | 0.3158 | 0.3789 | 0.4421 | 0.5053 | 0.5684 | $0.0887$ | 38 |
| $0.0634$ | -. 1269 | 0.1903 | 0.2538 | 0.3172 | 0.3806 | 0.444I | $0.5075$ | 0.5710 | $0.0891$ | 39 |
| 0.0537 | 0.1275 | 0.1912 | 0.2549 | 0.3187 | 0.3824 | 0.446I | 0.5098 | 0.5736 | 0.0895 | 40 |
| 0.0640 | 0. 1280 | 0.1921 | 0.2561 | 0.3201 |  | 0.448I | 0.5122 | 0.5762 | 0.0899 | 1 |
| 0.0643 | 0.1286 | 0. 1929 | 0.2572 | 0.3215 | 0.3859 | 0.4502 | 0.5145 | $\begin{aligned} & 0.5788 \end{aligned}$ | 0.0903 | 42 |
| $0.0646$ | 0.1 | -. 1938 | 0.2584 | 0.3230 | 0.3876 | 0.4522 | 0.5168 | 0.5814 | 0.0908 | 43 |
| $0.0649$ | 0.1298 | 0. 1946 | 0.2595 | 0.3244 | 0.3893 | 0. 4542 | 0.5190 | 0.5839 | 0.0912 | 44 |
| $0.0652$ | $0.1303$ | $0.1955$ | 0.2607 | 0.3259 | 0.3910 | 0.4562 | 0.5214 | 0.5865 | 0.0916 | 45 |
| $0.0655$ | o. 1309 | o. 1964 | 0.2618 | 0.3273 | 0. 3928 | 0.4582 | 0.5237 | 0.5891 | 0.0920 | 46 |
| $0.0657$ | 0.1315 | 0.1972 | 0.2630 | 0.3287 | 0. 3945 | 0.4602 | 0.5260 | 0.5917 | 0.0924 | 47 |
| 0.0660 | 0.1321 | -. 1981 | 0.2642 | 0.3302 | 0.3962 | 0.4622 | 0.5283 | 0.5943 | 0.0928 | 48 |
| 0.0663 | 0.1326 | o. 1990 | 0.2653 | 0.3316 | -. 3979 | 0.4642 | $0.53{ }^{0} 6$ | 0.5069 | 0.0932 | 49 |
| 0.0666 | 0.1332 | 0. 1998 | 0.2664 | 0.333 I | 0.3997 | 0.4663 | 0.5329 | 0.5995 | 0.0936 | 50 |
| $0.0669$ | 0.1338 | $0.20$ |  | 0.3345 | 0.4014 | 0.4683 | 0.5352 | 0.6021 | 0.0940 | 51 |
| 0.0672 | o. 1344 | 0.2016 | 0.2688 | 0.3359 | 0.4031 | 0.4703 | 0.5375 | 0.6047 | 0.0944 | 52 |
| 0.067 | 0.1349 | 0.2024 | 0.2699 | 0.3374 | 0.4048 | 0.4723 | 0. 5398 | 0.6073 | 0.0948 | 53 |
| 0.0678 | 0.1355 | 0.2033 | 0.2710 | 0.3388 | 0.4066 | 0.4743 | 0.5421 | 0.6099 | 0.0952 | 54 |
| 0.0681 | 0.1361 | 0.2042 | 0.2722 | 0.3403 | 0.4083 | 0.4764 | 0.5444 | 0.6125 | 0.0956 | 55 |
| 0.0683 | 0. 1367 | 0.2050 | 0.2734 | 0.3417 | 0.4100 | 0.4784 | 0.5467 | 0.6151 | 0.0961 | 56 |
| 0.0686 | 0.1373 | 0.2059 | 0.2745 | 0.3431 | 0.4118 | 0.4804 | 0. 5490 | 0.6177 | 0.0965 | 57 |
| 0.0689 | 0. 1378 | 0.2067 | 0.2756 | $0.344^{6}$ | 0.4135 | 0.4824 | 0.5513 | 0.6202 | 0.0969 | 58 |
| 0.0592 | 0. 1384 | 0.2076 | 0.2768 | 0.3460 | 0.4152 | 0.4844 | 0.5536 | 0.6228 | 0.0973 | 59 |
| 0.0695 | 0.1390 | 0.2085 | 0.2780 | 0.3474 | 0.4169 | 0.4864 | 0. 5559 | 0.6254 | 0.0977 | - |


| 96 | DISTANCES. |  |  |  |  |  |  |  |  | $4^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | a |
| - | 0.9937 | 1.9875 | 2.9812 | 3.9750 | 4.9687 | 5.9624 | 6.9562 | 7.9499 | 8.9437 | I. 3966 |
| OI | 0.9937 | r. 9874 | 2.9811 | 3.9748 | 4.9685 | 5.9522 | 6.9559 | 7.9495 | 8.9433 | I. 3966 |
| 02 | 0.9937 | 1. 9873 | 2.9810 | 3.9746 | 4.9683 | 5.9619 | 6.9556 | 7.9493 | 8.9429 | I. 3965 |
| $\bigcirc 3$ | 0.9936 | 1. 9872 | 2.9809 | 3.9745 | 4.968I | 5.9617 | 6.9553 | 7.9489 | 8.9426 | I. 3965 |
| 04 | 0.9936 | 1. 9872 | 2.9807 | 3.9743 | 4.9679 | 5.9615 | 6.9550 | 7.9486 | 8.9422 | I. 3965 |
| 05 | 0.9935 | 1.9871 | 2.9806 | 3.9741 | 4.9677 | 5.9612 | 6.9547 | 7.9483 | 8.9418 | I. 3965 |
| -6 | 0.9935 | 1. 9870 | 2.9805 | 3.9740 | 4.9675 | 5.9610 | 6.9545 | 7.9479 | 8.9414 | I. 3964 |
| 07 | 0.9935 | 1. 9869 | 2.9804 | 3.9738 | 4.9673 | 5.9607 | 6.9542 | 7.9476 | 8.941 I | I. 3964 |
| 08 | 0. 9934 | 1.9868 | 2.9802 | 3.9736 | 4.9671 | 5.9605 | 6.9539 | 7.9473 | 8.9407 | I. 3964 |
| 09 | 0.9934 | 1.9867 | 2.9801 | 3.9735 | 4.9668 | 5.9602 | 6.9536 | 7.9470 | 8.9403 | I. 3953 |
| 10 | 0.9933 | 1.9867 | 2.9800 | 3.9734 | 4.9666 | 5.9600 | 6.9533 | 7.9466 | 8.9400 | I. 3963 |
| II | 0.9933 |  | 2.9799 | 3.973 I |  | 5.9597 | 6.9530 | 7.9463 | 8.9396 | 1. 3963 |
| 12 | 0.9932 | 1. 9865 | 2.9797 | 3.9730 | 4.9662 | 5.9595 | 6.9527 | 7.9459 | 8.9392 | I. 3563 |
| 13 | 0.9932 | 1.9864 | 2.9756 | 3.9728 | 4.9660 | 5.9592 | 6.9524 | 7.9456 | 8.9388 | I. 3962 |
| 14 | 0.9932 | 1.9863 | 2.9795 | 3.9726 | 4.9658 | 5.9589 | 6.9521 | 7.9452 | 8.9,384 | I. 3962 |
| 15 | 0.9931 | 1.9862 | 2.9793 | 3.9725 | 4.9556 | 5.9587 | 6.9518 | 7.9449 | 8.9380 | 1. 3962 |
| 16 | 0.9931 | 1.9861 | 2.9792 | 3.9723 | 4.9654 | 5.9584 | 6.9515 | 7.9446 | 8.9376 | 1. 3962 |
| 17 | 0.9930 | 1.9861 | 2.9791 | 3.9721 | 4.9651 | 5.9582 | 6.9512 | 7.9442 | 8.9373 | I.3961 |
| 18 | 0.9930 | 1.9860 | 2.9790 | 3.9719 | 4.9649 | 5.9579 | 6.9509 | 7.9439 | 8.9369 | I.3961 |
| 19 | 0.9929 | 1.9859 | 2.9788 | 3.9718 | 4.9647 | 5.9577 | 6.9506 | 7.9435 | 8.9365 | I.396I |
| 20 | 0.9929 | 1.9858 | 2.9787 | 3.9716 | 4.9645 | 5.9574 | 6.9503 | 7.9432 | 8.936I | I. 3960 |
| 21 | 0.9929 | 1.9857 | 2.9786 | 3.9714 | 4.9643 | 5.957 | 6.9500 | 7.9428 | 8.9357 |  |
| 22 | 0.9928 | 1.9856 | 2.9784 | 3.9712 | 4.964 I | 5.9569 | 6.9497 | 7.9425 | 8.9353 | I. 3960 |
| 23 | 0.9928 | 1.9855 | 2.9783 | 3.9711 | 4.9638 | 5.9566 | 6.9494 | 7.9421 | 8.9349 | 1. 3959 |
| 24 | 0.9927 | 1.9854 | 2.9782 | 3.9709 | 4.9636 | 5.9563 | 6.9450 | 7.9418 | 8.9345 | I. 3959 |
| 25 | 0.9927 | 1.9854 | 2.9780 | 3.9707 | 4.9634 | 5.956 | 6.9487 | 7.9414 | 8.9341 | I. 3959 |
| 26 | 0.9926 | 1.9853 | 2.9779 | 3.9705 | 4.9632 | 5.9558 | 6.9484 | 7.9410 | 8.9337 | 1.3958 |
| 27 | 0.9926 | 1.9852 | 2.9778 | 3.9703 | 4.9629 | 5.9555 | 6.948 I | 7.9407 | 8.9333 | I. 3958 |
| 28 | 0.9925 | 1.9851 | 2.9776 | 3.9702 | 4.9527 | 5.9553 | 6.9478 | 7.9403 | 8.9329 | I. 3958 |
| 29 | 0.9925 | 1.9850 | 2.9775 | 3.9700 | 4.9625 | 5.9550 | 6.9475 | 7.9400 | 8.9325 | I. 3958 |
| 30 | 0.9925 | 1.9849 | 2.9774 | 3.9698 | 4.9623 | 5.9547 | 6.9472 | 7.9396 | 8.932 | I. 3957 |
| 31 | 0.9924 | 1.9848 | 2.9772 | 3.9696 | 4.9620 | 5.9544 | 6.9468 | 7.9393 | 8.9317 | I. 3957 |
| 32 | 0.9924 | 1.9847 | 2.9771 | 3.9694 | 4.9618 | 5.9542 | 6.9465 | 7.9389 | 8.9312 | I. 3957 |
| 33 | 0.9723 | 1.9846 | 2.9769 | 3.9693 | 4.9616 | 5.9539 | 6.9462 | 7.9385 | 8.9308 | 1. 3956 |
| 34 | 0.9923 | 1. 9845 | 2.9768 | 3.9691 | 4.9613 | 5.9536 | 6.9459 | 7.9381 | 8.9304 | I. 3956 |
| 35 | 0.9922 | 1.9844 | 2.9767 | 3.9689 | 4.9611 | 5.9533 | 6.9456 | 7.9378 | 8.9300 | 1.3956 |
| 36 | 0.9922 | 1. 9844 | 2.9765 | 3.9687 | 4.9609 | 5.953 | 6.9452 | 7.9374 | 8.9296 | I. 3955 |
| 37 | 0.9921 | 1.9843 | 2.976 | 3.9685 | 4.9606 | 5.9528 | 6.9449 | 7.9370 | 8.9292 | I. 3955 |
| 38 | 0.9921 | 1.9842 | 2.9762 | 3.9683 | 4.9604 | 5.9525 | 6.9446 | 7.9367 | 8.9287 | I. 3955 |
| 39 | 0.9920 | 1.9841 | 2.9761 | 3.9681 | 4.9602 | 5.9522 | 6.9443 | 7.9363 | 8.9283 | I. 3954 |
| 40 | 0.9920 | 1.9840 | 2.9760 | 3.9680 | 4.9600 | 5.9519 | 6.9439 | 7.9359 | 8.9279 | I. 3954 |
| 41 | 0.9919 | 1.9839 | 2.9758 | 3.9678 | 4.9597 | 5.9517 | 6.9436 | 7.9355 | 8.9275 | I. 3954 |
| 42 | 0.9919 | 1.9838 | 2.9757 | 3.9676 | 4.9595 | 5.9514 | 6.9433 | 7.9352 | 8.9270 | I. 3953 |
| 43 | 0.9918 | 1.9837 | 2.9755 | 3.9674 | 4.9592 | 5.9511 | 6.9429 | 7.9348 | 8.9266 | I. 3953 |
| 44 | 0.9918 | 1.9536 | 2.9754 | 3.9672 | 4.9590 | 5.9508 | 6.9426 | 7.9344 | 8.9262 | I. 3953 |
| 45 | 0.9918 | 1.9835 | 2.9753 | 3.9679 | 4.9588 | 5.9505 | 6.9423 | 7.9340 | 8.9258 | 1. 3952 |
| 46 | 0.9917 | 1.9834 | 2.9751 | 3.9668 | 4.9585 | 5.9502 | 6.9419 | 7.9336 | 8.9253 | 1.3952 |
| 47 | 0.9917 | 1.9833 | 2.9750 | 3.9566 | 4.9583 | 5.9499 | 6.9416 | 7.9332 | 8.9249 | 1.395 ${ }^{2}$ |
| 48 | 0.9916 | 1.9832 | 2.9748 | 3.9664 | 4.9580 | 5.9496 | 6.9412 | 7.9329 | 8.9245 | 1.3951 |
| 49 | 0.9916 | 1.983 I | 2.9747 | 3.9662 | 4.9578 | 5.9494 | 6.9409 | 7.9325 | 8.9240 | 1.3951 |
| 50 | 0.9715 | 1.9830 | 2.9745 | 3.9660 | 4.9576 | 5.9491 | 6.9406 | 7.9321 | 8.9236 | 1.3951 |
| 51 | 0.9915 | 1.9829 | 2.9744 | 3.9658 | 4.9573 | 5.9488 | 6.9402 | 7.9317 | 8.9231 | 1. 3950 |
| 52 | 0.9914 | 1.9828 | 2.9742 | 3.9556 | 4.9571 | 5.9485 | 6.9399 | 7.9313 | 8.9227 | I. 3950 |
| 53 | 0.9914 | 1.9827 | 2.9741 | 3.9654 | 4.9568 | 5.9482 | 6.9395 | 7.9309 | 8.9223 | 1. 3950 |
| 54 | 0.9913 | 1.9826 | 2.9739 | 3.9653 | 4.9566 | 5.9479 | 6.9392 | 7.9305 | 8.9218 | 1. 3949 |
| 55 | 0.9913 | 1.9825 | 2.9738 | 3.9651 | 4.9563 | 5.9476 | 6.9388 | 7.9301 | 8.9214 | I. $39+9$ |
| 56 | 0.9912 | 1.9824 | 2.9736 | 3.9649 | 4.9561 | 5.9473 | 6.9385 | 7.9297 | 8.9209 | 1. 3949 |
| 57 | 0.9912 | 1.9823 | 2.9735 | 3.9647 | 4.9558 | 5.9470 | 6.938 I | 7.9293 | 8.9205 | I. 3948 |
| 58 | 0.9911 | 1.9822 | 2.9733 | 3.9645 | 4.9556 | 5.9467 | 6.9378 | 7.9289 | 8.9100 | 1.3948 |
| 59 | 0.9911 | 1.9821 | 2.9732 | 3.9643 | 4.9553 | 5.9464 | 6.9375 | 7.9285 | 8.9196 | 1. 3948 |
| 60 | 0.9910 | 1.9820 | 2.9730 | 3.9641 | 4.9551 | 5.9461 | 6.9371 | 7.9281 | 8.9191 | 1.3947 |


| $4^{\circ}$ | HEIGH'TS. |  |  |  |  |  |  |  |  | 97 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | O | b |  |
| 0.0695 | 0.1390 | 0.2085 | 0.2780 | 0.3474 | 0.4169 | 0.4864 | O. 5559 | 0.6254 | 0.0977 | 00 |
| $0.0699$ | 0.1396 | 0.2093 | 0.2791 | 0.3489 | 0.4187 | 0.4884 | 0.5582 | 0.6280 | $0.0981$ | OI |
| 0.0701 | 0.1401 | 0.2102 | 0.2802 | 0.3503 | 0.4204 | 0.4904 | 0.5605 | 0.6306 | 0.0985 | C2 |
| 0.0704 | 0.1407 | 0.2111 | 0.2814 | 0.3518 | 0.4221 | 0.4925 | 0.5628 | 0.6332 | c.0989 | O3 |
| 0.0706 | 0.1413 | 0.2119 | 0.2826 | 0.3532 | 0.4238 | 0.4945 | 0.5651 | 0.6358 | 0.0993 | 04 |
| 0.0709 | 0.1419 | 0.2128 | 0.2837 | 0.3546 | 0.4256 | 0.4965 | 0.5674 | 0.6384 | 0.0997 | O5 |
| 0.0712 | 0.1424 | 0.2136 | 0.2848 | 0.3561 | 0.4273 | 0.4985 | 0.5697 | 0.6409 | 0.1001 | 06 |
| 0.0715 | 0.1430 | 0.2145 | 0.2860 | 0.3575 | 0.4290 | 0.5005 | 0.5720 | 0.6435 | 0.1005 | 07 |
| 0.0718 | 0.1436 | 0.2154 | 0.2872 | 0.3589 | 0.4307 | 0.5025 | 0.5743 | $0.6461$ | 0.1009 | 08 |
| 0.0721 | 0.1442 | 0.2162 | 0.2883 | 0.3604 | 0.4325 | 0.5045 | 0.5766 | 0.6487 | -.1013 | 09 |
| 0.0724 | 0.1447 | 0.2171 | 0.2894 | 0.3618 | 0.4342 | 0.5065 | 0.5789 | 0.6513 | O.1017 | 10 |
| 0.0727 | 0.1453 | 0.2180 | 0.2906 | 0.3633 | 0.4359 | 0.5086 | 0.5812 | 0.6539 | 0.1021 | II |
| 0.0729 | -. 1459 | 0.2188 | 0.2918 | 0.3647 | 0.4376 | 0.51c6 | 0.5835 | 0.6565 | 0.1025 | 2 |
| 0.0732 | 0.1465 | 0.2197 | 0.2929 | 0.366 I | 0.4394 | 0.5126 | 0.5858 | .0.6591 | 0.1029 | 13 |
| 0.0735 | 0.1470 | 0.2205 | 0.2940 | 0.3676 | 0.4411 | 0.5146 | 0.588 I | 0.6616 | 0.1033 | 14 |
| 0.0738 | 0.1476 | 0.2214 | 0.2952 | 0.3690 | 0.4428 | 0.5166 | 0.5904 | $0.6642$ | 0.1037 | 15 |
| 0.0741 | 0.1482 | 0.2223 | 0.2964 | 0.3704 | 0.4445 | 0.5186 | 0.5927 | 0.6668 | 0.1041 | 16 |
| 0.0744 | 0.1488 | 0.2231 | 0.2975 | 0.3719 | 0.4463 | 0.5206 | 0. 5950 | 0.6694 | 0.1046 | 17 |
| 0.0747 | 0.1493 | 0.2240 | 0.2986 | 0.3733 | 0.4480 | 0.5226 | 0.5973 | 0.6720 | 0.1050 | 18 |
| $0.0749$ | 0.1499 | 0.2248 | 0.2998 | 0.3747 | 0.4497 | 0.5246 | 0.5596 | $0.6746$ | $0.1054$ | 19 |
| 0.0752 | 0.1505 | 0.2257 | 0.3010 | 0.3762 | 0.4514 | 0.5266 | 0.6019 | 0.6772 | 0.1058 | 20 |
| 0.0755 | 0.1510 | 0.2266 | 0.3021 | 0.3776 | 0.4531 | 0.5286 | 0.6042 | 0.6797 | $0.1052$ | I |
| $0.0758$ | 0.1516 | 0.2274 | 0.3032 | $0.3791$ | 0.4549 | 0.5307 | $0.6065$ | $06823$ | $0.1066$ | 22 |
| 0.0761 | 0.1522 | 0.2283 | 0.3044 | 0.3805 | c. 4566 | 0.5327 | 0.6088 | 0.6849 | 0.1070 | 23 |
| 0.0764 | 0.1528 | 0.2292 | 0.3056 | 0.3819 | 0.4583 | 0.5347 | 0.6111 | 0.6875 | 0.1074 | 24 |
| 0.0767 | O.1533 | 0.2300 | 0.3067 | 0.3834 | 0.4600 | 0.5367 | 0.6134 | 0.6900 | $0.1078$ | 25 |
| $0.0770$ | 0.1539 | 0.2309 | 0.3078 | 0.3848 | 0.4618 | 0.5387 | $0.6157$ | $0.6926$ | $0.1082$ | 26 |
| 0.0772 | 0.1545 | 0.2317 | 0.3090 | 0.3862 | 0.4635 | 0.5407 | 0.6180 | 0.6952 | 0.1086 | 27 |
| 0.0775 | O.1551 | 0.2326 | 0.3101 | 0.3877 | 0.4652 | 0.5427 | 0.6203 | . 0.6978 | 0.1090 | 28 |
| 0.0778 | 0.1556 | 0.2335 | 0.3113 | 0.3891 | 0.4669 | 0.5447 | $0.6226$ | 0.7004 | $0.1094$ | 29 |
| 0.0781 | -. 1562 | 0.2343 | 0.3124 | 0.3905 | 0.4687 | 0.5467 | 0.6249 | 0.7030 | 0.1c98 | 30 |
| 0.0784 | 0.1568 | 0.2352 | 0.3136 | 0.3920 | 0.4703 | 0.5487 | 0.6271 | 0.7055 | 0.1102 | 31 |
| 0.0787 | O.1574 | 0.2360 | 0.3147 | 0.3934 | 0.4721 | 0.5508 | $0.6294$ | $0.708 \mathrm{I}$ | 0.1107 | 32 |
| $0.0790$ | 0.1579 | 0.2369 | 0.3159 | 0.3948 | 0.4738 | 0.5528 | $0.6318$ | 0.7107 | 0.1111 | 33 |
| $0.0793$ | 0.1585 | 0.2378 | 0.3170 | 0.3963 | 0.4755 | 0.5548 | $0.6340$ | $0.7133$ | C.III5 | 34 |
| 0.0795 | O.1591 | 0.2386 | 0.3182 | 0.3977 | 0.4772 | 0.5568 | $0.6363$ | $0.7159$ | O.III9 | 35 |
| $0.0798$ | 0.1597 | 0.2395 | 0.3193 | 0.3991 | 0.4790 | - 0.5588 | $0.6386$ | $0.7185$ | 0.1123 | 36 |
| $0.0801$ | $0.16 c 2$ | 0.2403 | $0.3204$ | 0.4007 | $0.4807$ | $0.5608$ | $0.6409$ | $0.7210$ | $0.1127$ | 37 |
| 0.0804 | 0.1608 | 0.2412 | 0.3216 | $0.4020$ | 0.4824 | 0.5628 | 0.6432 | 0.7236 | 0.1131 | 38 |
| 0.0807 | O.1614 | 0.2421 | 0.3228 | 0.4034 | 0.484 I | 0.5648 | 0.6455 | 0.7262 | c.II35 | 39 |
| 0.0810 | 0.1620 | 0.2429 | 0.3239 | 0.4049 | 0.4859 | 0.5668 | 0.6478 | 0.7288 | 0.1139 | 40 |
| 0.0813 | 0.1625 | 0.2438 | 0.3250 | 0.4063 | 0.4876 | 0.5688 | 0.6501 | 0.7313 | O.II43 | 41 |
| $0.0815$ | 0.1631 | 0.2446 | 0.3262 | 0.4077 | 0.4893 | 0.5708 | $0.6524$ | $0.7339$ | 0.1147 | 42 |
| $0.0818$ | 0.1637 | 0.2455 | 0.3273 | 0.4092 | 0.4910 | 0.5728 | 0.6546 | 0.7365 | 0.1151 | 43 |
| 0.0821 | 0.1642 | 0.2464 | 0.3285 | 0.4106 | 0.4927 | 0.5748 | 0.6570 | 0.7391 | O.1155 | 44 |
| 0.0824 | 0.1648 | 0.2472 | 0.3296 | 0.4120 | 0.4945 | 0.5768 | 0.6593 | 0.7417 | O.1159 | 45 |
| 0.0827 | 0.1654 | 0.248 I | 0.3308 | 0.4135 | 0.4961 | 0.5788 | $0.6615$ | 0.7442 | $0.1163$ | 46 |
| $0.0830$ | 0.1660 | 0.2489 | 0.3319 | 0.4149 | 0.4979 | 0.5809 | $0.6638$ | 0.7468 | 0.1167 | 47 |
| $0.0833$ | 0.1665 | 0.2498 | 0.3331 | $0.4163$ | 0.4996 | 0.5829 | 0.6662 | 0.7494 | O.II7I | 48 |
| 0.0836 | 0.1671 | 0.2507 | 0.3342 | 0.4178 | 0.5013 | 0.5849 | 0.6684 | 0.7520 | 0.1176 | 49 |
| 0.0838 | 0.1677 | 0.2515 | 0.3354 | 0.4192 | 0.5030 | 0.5869 | 0.6707 | 0.7546 | 0.1180 | 50 |
| 0.0841 | 0.1683 | 0.2524 | 0.3365 | 0.4206 | 0.5048 | 0.5889 | 0.6730 | 0.7572 | 0.1184 | 5I |
| 0.0844 | 0.1688 | 0.2532 | 0.3376 | 0.4221 | 0.5065 | 0.5909 | 0.6753 | 0.7597 | 0.1188 | 52 |
| 0.0847 | 0.1694 | 0.2541 | 0.3388 | 0.4235 | 0.5082 | 0.5929 | 0.6776 | 0.7623 | O.1192 | 53 |
| 0.0850 | 0.1700 | $0.2549$ | 0.3399 | 0.4249 | 0.5099 | 0.5949 | 0.6798 | 0.7648 | 0.1196 | 54 |
| 0.0853 | 0.1705 | 0.2558 | 0.34II | 0.4264 | 0.5116 | 0.5969 | 0.6822 | 0.7674 | 0.1200 | 55 |
| 0.0856 | 0.1711 | 0.2567 | 0.3422 | 0.4278 | 0.5134 | 0.5989 | 0.6845 | 0.7700 | 0.1204 | 56 |
| $0.0858$ | 0.1717 | $0.2575$ | 0.3434 | 0.4293 | 0.5150 | 0.6009 | 0.6867 | 0.7726 | 0.1208 | 57 |
| $0.086 \mathrm{I}$ | $0.1723$ | 0.2584 | 0.3445 | 0.4306 | 0.5168 | 0.6029 | 0.6890 | c. 7752 | $0.1212$ | 58 |
| 0.0864 | 0.1728 | 0.2593 | 0.3457 | 0.4321 | 0.5185 | 0.6049 | 0.6914 | 0.7778 | 0.1216 | 59 |
| 0.0867 | 0.1734 | 0.2601 | 0.3468 | 0.4335 | 0.5202 | 0.6069 | 0.6936 | 0.7803 | C. 1220 | 60 |


| 98 | DISTANCES. |  |  |  |  |  |  |  |  | $5{ }^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | a |
| -o | 0.99 | 1.9820 | 2.9730 | 3.9641 | 4.955I | 5.946I | 6.937 I | 7.928I | 8.9191 | 1. 3947 |
| 01 | c.9910 | 1.9819 | 2.9729 | 3.9639 | 4.9548 | 5.9458 | 6.9367 | 7.9277 | 8.9187 | I. 3947 |
| 02 | 0.9909 | 1.9818 | 2.9727 | 3.9636 | 4.9546 | 5.9455 | 6.9364 | 7.9273 | 8.9182 | I. 3946 |
| $\bigcirc 3$ | 0.9909 | 1.9817 | 2.9726 | 3.9634 | 4.9543 | 5.9452 | 6.9360 | 7.9269 | 8.9177 | I. 3946 |
| 04 | 0.9908 | 1.9816 | 2.9724 | 3.9632 | 4.9541 | 5.9449 | 6.9357 | 7.9265 | 8.9173 | I. 3946 |
| 05 | 0.9908 | 1.9815 | 2.9723 | 3.9630 | 4.9538 | 5.9446 | 6.9353 | 7.9261 | 8.9168 | I. 3945 |
| 06 | 0. 9907 | 1.9814 | 2.9721 | 3.9628 | 4.9535 | 5.9442 | 6.9349 | 7.9257 | 8.9164 | I. 3945 |
| 07 | 0.9907 | 1.9813 | 2.9720 | 3.9626 | 4.9533 | 5.9439 | 6.9346 | 7.9252 | 8.9159 | I. 3944 |
| 08 | 0.9906 | 1.9812 | 2.9718 | 3.9624 | 4.9530 | 5.9436 | 6.9342 | 7.9248 | 8.9154 | I. 3944 |
| 99 | 0.9906 | 1.98II | 2.9717 | 3.9622 | 4.9528 | 5.9433 | 6.9339 | 7.9244 | 8.9150 | I. 3944 |
| 10 | 0.9905 | 1.9810 | 2.9715 | 3.9620 | 4.9525 | 5.9430 | 6.9335 | 7.9240 | 8.9145 | I. 3943 |
| 11 | 0.9904 | 1.980 | 2.9713 | 3.9618 | 4.9522 | 5.9427 | 6.9331 | 7.9236 | 8.9140 | I. 3943 |
| 12 | 0.9904 | 1. 9808 | 2.9712 | 3.9616 | 4.9520 | 5.9424 | 6.9328 | 7.9232 | 8.9136 | I. 3942 |
| ${ }^{1} 3$ | 0.9903 | 1. 9807 | 2.9710 | 3.9614 | 4.9517 | 5.942I | 6.9324 | 7.9227 | 8.9131 | I. 3942 |
| 14 | 0.9903 | I. 9806 | 2.9709 | 3.9612 | 4.9515 | 5.9417 | 6.9320 | 7.9223 | 8.9126 | I. 3941 |
| 15 | 0.9902 | 1.9805 | 2.9707 | 3.9610 | 4.9512 | 5.9414 | 6.9317 | 7.9219 | 8.9121 | I. 3941 |
| 16 | 0.9902 | 1.9804 | 2.9706 | 3.9607 | 4.9509 | 5.9411 | 6.9313 | 7.9215 | 8.9117 | I. 3941 |
| 17 | 0.9901 | 1.9803 | 2.9704 | 3.9605 | 4.9507 | 5.9408 | 6.9309 | 7.9211 | 8.9112 | I. 3940 |
| 18 | 0.9901 | 1.9802 | 2.9702 | 3.9603 | 4.9504 | 5.9405 | 6.9306 | 7.9206 | 8.9107 | I. 3940 |
| 19 | 0.9900 | 1.9801 | 2.9701 | 3.9601 | 4.9501 | 5.9402 | 6.9302 | 7.9202 | 8.9102 | 1. 3940 |
| 20 | 0.99 co | 1.9799 | 2.9639 | 3.9599 | 4.9499 | $5 \cdot 9398$ | 6.9298 | 7.9198 | 8.9098 | I. 3939 |
| 21 | 0.9899 | I. 9798 | 2.9698 | 3.9597 | 4.9496 | 5.9395 | 6.9294 | 7.9193 | 8.9093 | I. 3939 |
| 22 | 0.9899 | 1. 9797 | 2.9695 | 3.9595 | 4.9493 | 5.9392 | 6.9290 | 7.9189 | 8.9088 | I. 3938 |
| 23 | 0.9898 | 1.9796 | 2.9694 | 3.9592 | 4.9490 | 5.9389 | 6.9287 | 7.9185 | 8.9083 | I. 3938 |
| 24 | 0.9898 | I. 9795 | 2.9693 | 3.9590 | 4.9488 | 5.9385 | 6.9283 | 7.9180 | 8.9078 | I. 3938 |
| 25 | 0.9897 | I. 9794 | 2.9691 | 3.9588 | 4.9485 | 5.9382 | 6.9279 | 7.9176 | 8.9073 | I. 3937 |
| 26 | 0.9896 | I. 9793 | 2.9689 | 3.9586 | 4.9482 | 5.9379 | 6.9275 | 7.9172 | 8.9068 | I. 3937 |
| 27 | 0.9896 | 1. 9792 | 2.9688 | 3.9584 | 4.9480 | 5.9375 | 6.9271 | 7.9167 | 8.9063 | I. 3936 |
| 28 | 0.9895 | 1.9791 | 2.9686 | 3.958I | 4.9477 | 5.9372 | 6.9268 | 7.9163 | 8.9058 | I. 3936 |
| 29 | 0.9895 | 1. 9790 | 2.9684 | 3.9579 | 4.9474 | 5.9369 | 6.9264 | 7.9159 | 8.9053 | I. 3936 |
| 30 | 0.9894 | 1.9789 | 2.9683 | 3.9577 | 4.947 I | 5.9366 | 6.0260 | 7.9154 | 8.9048 | I. 3935 |
| 31 | 0.9894 | 1. 9787 | 2.9681 | 3.9575 | 4.9469 | 5.9362 | 6.9256 | 7.9150 | 8.9043 | I. 3935 |
| 32 | 0.9893 | 1. 9786 | 2.9679 | 3.9573 | 4.9466 | 5.9359 | 6.9252 | 7.9145 | 8.9038 | I. 3934 |
| 33 | 0.9893 | I. 9783 | 2.9678 | 3.9570 | 4.9463 | 5.9355 | 6.9248 | 7.9141 | 8.9033 | I. 3934 |
| 34 | 0.9892 | 1.9784 | 2.9676 | 3.9568 | 4.9460 | 5.9352 | 6.9244 | 7.9136 | 8.9028 | I. 3934 |
| 35 | 0.9891 | 1.9783 | 2.9674 | 3.9566 | 4.9457 | 5.9349 | 6.9240 | 7.9132 | 8.9023 | 1.3933 |
| 36 | 0.9891 | 1.9782 | 2.9673 | 3.9564 | 4.9454 | 5.9345 | 6.9236 | 7.9127 | 8.9018 | I. 3933 |
| 37 | 0.9890 | 1.9781 | 2.9671 | 3.9561 | $4.945^{2}$ | 5.9342 | 6.9232 | 7.9123 | 8.9013 | I. 3932 |
| 38 | 0.9880 | 1.9780 | 2.9669 | 3.9559 | 4.9449 | 5.9339 | 6.9228 | 7.9118 | 8.9008 | 1. 3932 |
| 39 | 0.9889 | 1.9778 | 2.9668 | 3.9557 | 4.9446 | 5.9335 | 6.9224 | 7.9114 | 8.9003 | 1.3932 |
| 40 | 0.9889 | 1.9777 | 2.9666 | 3.9555 | 4.9443 | 5.9332 | 6.9220 | 7.9109 | 8.8998 | 1.3931 |
| 41 | 0.9888 | 1. 9776 | 2.9664 | $3.955^{2}$ | 4.9440 | 5.9328 | 6.9216 | 7.9104 | 8.8993 | I. 393 I |
| 42 | 0.9887 | 1. 9775 | 2.9662 | 3.9550 | 4.9437 | 5.9325 | 6.9212 | 7.9100 | 8.8987 | I. 3930 |
| 43 | 0.9887 | 1. 9774 | 2.9661 | 3.9548 | 4.9435 | 5.9321 | 6.920 S | 7.9095 | S. 8982 | I. 3930 |
| 44 | 0.9886 | 1. 9773 | 2.9659 | 3.9545 | 4.9432 | 5.9318 | 6.9204 | 7.9091 | 8.8977 | I. 3930 |
| 45 | 0. 9886 | 1. 9772 | 2.9657 | 3.9543 | 4.9429 | 5.9315 | 6.9200 | 7.9086 | 8.8972 | 1. 3929 |
| 46 | 0.9885 | 1. 9770 | 2.9656 | 3.954 | 4.9426 | 5.9311 | 6.9196 | 7.908I | 8.8967 | I. 3929 |
| 47 | 0.9885 | I. 9769 | 2.9654 | 3.9538 | 4.9423 | 5.9308 | 6.9192 | 7.9077 | 8.8961 | I. 3928 |
| 48 | 0.9884 | 1. 9768 | 2.9652 | 3.9536 | 4.9420 | $5.93{ }^{\text {c }}$ | 6.9188 | 7.9072 | 8. 8956 | I. 3928 |
| 49 | 0.9883 | I. 9767 | 2.9650 | 3.9534 | 4.9417 | 5.9300 | 6.9184 | 7.9067 | 8.8951 | I. 3928 |
| 50 | 0.9883 | 1.9766 | 2.9649 | 3.9531 | 4.9414 | 5.9297 | 6.9180 | 7.9063 | 8.8946 | I. 3927 |
| 51 | 0.9882 | 1. 9765 | 2.9647 | 3.9529 | 4.9411 | 5.9294 | 6.9176 | 7.9058 | 8. 8940 | 1. 3927 |
| 52 | 0.9882 | 1.9763 | 2.9645 | 3.9527 | 4.9408 | 5.9290 | 6.9172 | 7.9053 | 8. 8035 | I. 3926 |
| 53 | 0.988I | 1.9762 | 2.9643 | 3.9524 | 4.9405 | 5.9286 | 6.9167 | 7.9048 | 8. 8930 | 1. 3926 |
| 54 | 0.9880 | 1.9761 | 2.9641 | 3.9522 | 4.9402 | 5.9283 | 6.9163 | 7.9044 | 8. 8924 | I. 3926 |
| 55 | 0.9880 | 1.9760 | 2.9640 | 3.9519 | 4.9399 | 5.9279 | 6.9159 | 7.9039 | S. S919 | I. 3925 |
| 56 | 0.9879 | I. 9759 | 2.9638 | 3.9517 | 4.9396 | 5.9276 | 6.9155 | 7.9034 | S. 8913 | I. 3925 |
| 57 | 0.9879 | I. 9757 | 2.9636 | 3.9515 | 4.9393 | 5.9272 | 6.9151 | 7.9029 | 8. 8908 | I. 3924 |
| 58 | 0.9878 | 1.9756 | 2.9634 | 3.9512 | 4.9390 | 5.9268 | 6.9147 | 7.9025 | 8. 8903 | 1. 3924 |
| 59 | c. 9877 | I. 9755 | 2.9632 | 3.9510 | 4.9387 | 5.9265 | 6.9142 | 7.9020 | 8. 8897 | 1. 3924 |
| 60 | 0.9877 | 1. 9754 | 2.9631 | 3.9508 | 4.9384 | 5.9261 | 6.9138 | 7.9015 | 8. 5892 | 1.3923 |


| $5^{\circ}$ | HEIGHTS. |  |  |  |  |  |  |  |  | 99 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | b |  |
| 0.0867 | 0.1734 | 0.2 | 0.3468 | 0.4335 | 0.5202 | 0.6069 | 0.6936 | 0.7803 | 20 | oo |
| 0.0870 | 0.1740 | 0.2610 | o. 3480 | 0. 4349 | 0. 5219 | 0.6089 | 0.6959 | 0.7829 | 0. 1224 | or |
| 0.0873 | -.r 745 | 0.2618 | 0.3491 | 0.4364 | 0.5236 | -.6109 | 0.6982 | 0.7854 | 0. 1228 | 02 |
| 0.0876 | 0.1751 | 0.2627 | 0.3502 | c. 4378 | 0. 5254 | 0.6129 | 0.7c05 | 0.7880 | 0. 1232 | 03 |
| 0.0878 | 0.1757 | 0.2635 | 0.3514 | 0.4392 | 0.527 I | 0.6149 | 0.7028 | 0. 7906 | 0.1236 | 04 |
| 0.0881 | -.1763 | 0. 2644 | 0.3525 | 0.4407 | 0. 5288 | 0.6169 | 0.7050 | 0.7932 | 0.1240 | 05 |
| 0.0884 | 0.1768 | 0.2653 | 0.3537 | 0.4421 | 0. 5305 | 0.6189 | 0. 7074 | 0.7958 | 0.1244 | 6 |
| 0.0887 | -. 1774 | 0.2661 | 0.3548 | 0.4435 | 0. 5322 | 0.6209 | 0. 7096 | 0. 7983 | 0. 1248 | 7 |
| 0.08g0 | 0. 1780 | 0.2670 | 0.3560 | 0.4450 | 0. 5539 | 0.6229 | 0.7119 | 0.8009 | 0.1253 | 8 |
| 0.0893 | -. 1786 | 0.2678 | 0.3571 | 0. 4464 | 0. 5357 | 0.6249 | 0.7142 | 0.8035 | 0.1257 | 09 |
| 0.0896 | -.1791 | 0.2687 | 0.3582 | 0. 4478 | 0. 5374 | 0.6269 | 0.7165 | 0.8060 | 0.126I | - |
| 0.0898 | 0.1 |  | 0.3594 | O. | 0.5391 | 0.6289 | 0.7188 | 36 |  | 11 |
| 0.0901 | 0.1803 | 0.2704 | 0.3605 | 0.4507 | 0. 5408 | 0.6309 | 0.7211 | c.8112 | 0.1269 | 12 |
| 0.0904 | 0.1808 | 0.2713 | 0.3617 | 0.4521 | 0. 5425 | 0.6329 | 0. 7234 | 0.8138 | 0.1273 | I3 |
| 0.0907 | -. 1814 | 0.2721 | 0.3628 | 0.4535 | 0.5442 | 0.6349 | 0.7256 | 0.8163 | 0.1277 | 14 |
| 0.0910 | 0.1820 | 0.2730 | 0.3640 | 0.4550 | 0. 5459 | 0.6369 | 0. 7279 | 0.8189 | 0.1281 | 15 |
| 0.0913 | 0.1826 | 0.2738 | 0.3651 | 0.4564 | 0.5477 | 0.6389 | 0.7302 | 0.8215 | 0. 1285 | 16 |
| 0.0916 | 0.1831 | 0.2747 | 0.3662 | 0.4578 | 0. 5494 | 0.6409 | 0.7325 | 0.8240 | 0. 1289 | 7 |
| 0.0918 | 0.1837 | 0.2755 | 0.3674 | 0.4592 | 0.5511 | 0.6429 | 0. 7348 | 0.8266 | 0.1293 | 8 |
| 0.0921 | 0.1843 | 0.2764 | 0.3685 | 0.4607 | 0. 5528 | 0.6449 | 0.7371 | 0.8292 | 0. 1297 | 19 |
| 0.0924 | 0.1848 | 0.2773 | 0.3697 | 0.4621 | 0. 5545 | 0.6469 | 0.7394 | 0.8318 | 0.1301 | 2 |
| 0.0927 | 0. 1854 | 0.278 I | 0.3708 | 0.4635 | 0. 5562 | 0.6489 | 0.7416 |  | 0.1305 | 1 |
| 0.0930 | 0.1860 | 0.2790 | 0.3720 | 0. 4649 | 0. 5579 | 0.6509 | 0.7439 | 0.8369 | -.1309 | 22 |
| 0.0933 | 0. 1865 | 0.2798 | 0.3731 | 0.4664 | 0. 5596 | 0.6529 | 0. 7462 | 0.8394 | -.13I3 | 23 |
| 0.0936 | 0.1871 | 0.2807 | 0.3742 | 0.4678 | 0.5614 | 0.6549 | 0. 7485 | 0.8420 | -.1317 | 24 |
| 0.0938 | 0. 1877 | 0.2815 | 0. 3754 | 0.4692 | 0.5631 | 0.6569 | 0.7507 | 0.8446 | -. 1321 | 25 |
| 0.0941 | 0. 1883 | 0.2824 | 0.3765 | 0.4706 | 0.5648 | 0.6589 | 0.7530 | 0.8472 | -.1326 | 26 |
| 0.0944 | 0. 1888 | 0.2833 | 0. 3777 | 0.4721 | 0.5665 | 0.6609 | 0.7553 | 0.8498 | -. 1330 | 27 |
| 0.0947 | 0. 1894 | 0.284 I | 0.3788 | 0.4735 | 0.5682 | 0.6629 | 0.7576 | 0.8523 | 0.1334 | 28 |
| 0.0950 | 0. 1900 | 0.2850 | 0.3800 | 0.4749 | 0.5699 | 0.6649 | 0.7599 | 0.8549 | -. 1338 | 29 |
| 0.0953 | -. 1905 | 0.2858 | 0.3811 | 0.4764 | 0.5716 | 0.6669 | 0.7622 | 0.8574 | -. 1342 | 30 |
| 0.0956 | -. I9II | 0.2867 | 0.3822 | 0.4778 | C. 5734 | 0.6689 | 0. 7645 | 0.8600 | 0. 1346 | 31 |
| 0.0958 | -. 1917 | 0.2875 | 0.3834 | 0.4792 | 0.5751 | 0.6709 | 0. 7667 | 0.8626 | -.1350 | 32 |
| 0.0961 | -. 1923 | 0.2884 | 0.3845 | 0.4806 | 0.5768 | 0.6729 | 0.7690 | 0.8652 | -. 1354 | 33 |
| 0.0964 | -. 1928 | 0.2892 | 0.3856 | 0.4820 | 0.5785 | 0.6749 | 0.7713 | 0.8677 | -. 1358 | 34 |
| 0.0967 | o. 1934 | 0.2901 | 0.3868 | 0.4835 | 0.5802 | 0.6769 | 0.7736 | 0.8703 | -. 1362 | 35 |
| 0.097 | o. 1940 | 0.2909 | 0.3879 | 0.4849 | 0.5819 | 0.6789 | 0.7759 | 0.8728 | -. I366 | 36 |
| 0.09 | -. 1945 | 0.2918 | 0.389I | 0.4863 | 0.5836 | 0.6809 | 0.7782 | 0. 8754 | -. 1370 | 37 |
| 0.0976 | -.1951 | 0.2927 | 0.3902 | 0.4878 | 0. 5853 | 0.6829 | 0.7804 | 0.8780 | -. 1374 | 38 |
| 0.0978 | 0. 1957 | 0.2935 | -. 3914 | 0.4892 | 0.5870 | 0.6849 | 0.7827 | 0.8806 | -. 1378 | 39 |
| 0.0981 | -. 1962 | 0.2944 | 0.3925 | 0.4906 | 0.5887 | 0.6869 | 0.7850 | 0.883 I | 0.1382 |  |
| 0.0984 | c. 1968 | 0.2952 | . 39 | 0.4920 | 0.5905 | 0.6889 | 0.7873 | 0.8857 | 0.1389 | 1 |
| 0.0987 | o. 1974 | 0.2961 | 0.3948 | 0.4935 | 0.5921 | 0.6908 | 0.7895 | 0.8882 | 0.1390 | 42 |
| 0.0990 | -. 1979 | 0.2969 | 0.3959 | 0.4948 | 0.5938 | 0.6928 | 0.7918 | 0.8907 | -. 1395 | 43 |
| 0.0993 | 0. 1985 | 0.2978 | 0.3970 | 0.4963 | 0.5956 | 0.6948 | 0.794I | 0.8933 | 0.1399 | 44 |
| 0.0995 | 0.1991 | 0.2986 | 0.3982 | 0.4977 | 0.5973 | 0.6968 | 0.7963 | 0.8959 | 0.1403 | 45 |
| 0.0998 | 0. 1997 | 0.2995 | 0.3993 | 0.4991 | 0.5990 | 0.6988 | 0.7986 | 0.8985 | 0. 1407 | 46 |
| 0.1001 | 0.2002 | 0.3003 | 0.4004 | 0. 5006 | 0.6007 | 0.7008 | 0.8009 | 0.9010 | 0.14II | 47 |
| 0. 1004 | 0.2008 | 0.3012 | 0.4016 | 0. 5020 | 0.6024 | 0.7028 | 0.8032 | 0.9036 | -.1415 | 48 |
| 0. 1007 | 0.2014 | 0.3020 | 0.4027 | 0.5034 | 0.6041 | 0.7048 | 0.8054 | 0.9061 | 0.1419 | 49 |
| 0.1010 | 0.2019 | 0.3029 | 0.4039 | 0. 5049 | 0.6058 | 0.7068 | 0.8078 | 0.9087 | 0.1423 | 50 |
| 0.1013 | 0.2025 | 0.3038 | 0.4050 | 0.5063 | 0.6075 | 0.7088 | 0.8İo | 0.9113 | 0.1427 | 51 |
| 0.1015 | 0.2031 | 0.3046 | 0.4062 | 0.5077 | 0.6092 | 0.7108 | 0.8123 | 0.9139 | 0.I43I | 52 |
| -. 1018 | 0.2036 | 0. 3055 | 0.4073 | 0.5091 | 0.6109 | 0.7127 | 0.8146 | 0.9164 | O. I435 | 53 |
| 0.1021 | 0.2042 | 0. 3063 | 0.4084 | 0.5105 | 0.6126 | 0.7147 | 0.8i68 | 0.9189 | O. 1439 | 54 |
| 0.1024 | 0.2048 | 0.3072 | 0.4096 | 0.5119 | 0.6143 | 0.7167 | 0.8191 | 0.9215 | 0. 1443 | 55 |
| 0. 1027 | 0.2053 | 0.3080 | 0.4107 | 0.5134 | 0.6160 | 0.7187 | 0.8214 | 0.9240 | 0. 1447 | 56 |
| -.1030 | 0.2059 | 0. 3089 | 0.4118 | 0.5148 | 0.6177 | 0. 7207 | 0.8237 | 0.9266 | 0.1451 | 57 |
| o.1032 | 0.2065 | 0.3097 | 0.4130 | 0.5162 | 0.6194 | 0.7227 | o. 8259 | 0.9292 | 0.1455 | 58 |
| 0.1035 | 0.2071 | 0.3106 | 0.4141 | 0.5176 | 0.6212 | 0.7247 | 0.8282 | 0.9318 | 0.1459 | 59 |
| o. 103 | 0.2076 | 0.3114 | 0.415 | 0.5191 | c. 622 | 0.726 | 0.830 | 0.9343 | 0.14 | 60 |



| $6^{\circ}$ | HEIGHTS. |  |  |  |  |  |  |  | 101 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | b |  |
| 0.1038 | 0.2076 |  |  |  | 0.6229 |  | 0.8305 |  |  |  |
| 0.104I | 0.2082 | 0.3123 | 0.4164 | 0. 5205 | 0.6246 | 0. 7287 | 0.8327 | 0.9368 |  | OI |
| 0. 1044 | 0.2088 | 0.313I | 0.4175 | -. 5219 | 0.6263 | 0.7307 | -. 8350 | 0.9394 | -.147I | 02 |
| 0.1047 | 0.2093 | -.3140 | 0.4186 | 0. 5233 | 0.6280 | 0.7326 | 0. 8373 | 0.9419 | -.1476 | 03 |
| -. 1049 | 0.2099 | $0.314^{8}$ | 0.4198 | 0. 5247 | 0.6297 | -. 7346 | 0.8396 | 0.9445 | -. 1480 | 04 |
| -. 1052 | 0.2105 | 0.3157 | 0.4209 | 0. 5262 | 0.6314 | 0.7366 | 0.8418 | 0.947 I | o. 1484 | 05 |
| -.1055 | 0.2110 | 0.3165 | 0.4220 | 0.5276 | 0.6331 | 0.7386 | 0.844I | 0.9496 | o. 1488 | 06 |
| -. 1058 | 16 | -. 3174 | $0.423^{2}$ | 0.5290 | 0.6348 | 0.7406 | 0.8464 | 0.9522 | -.1492 | 07 |
| 0.1061 | 0.2122 | 0.3182 | 0.4243 | 0.5304 | 0.6365 | 0.7426 | 0.8486 | 0.9547 | o. 1496 | 08 |
| 0.1064 | 0.2127 | o.3191 | 0. 4255 | 0.5318 | 0.6382 | 0. 7446 | -. 8509 | 0.9573 | 0.1500 | 09 |
| -. 1067 | 0.2133 | 0.3200 | 0.4266 | -. 5333 | 0.6399 | 0.7466 | 0.8532 | 0.9599 | 0.1504 | Io |
| 0.1069 | 0.2139 | 0.3208 | 0.4277 | 0.5347 | 0.6416 | 0.7485 | o. 8554 | 0.9624 | -. 1508 | 1 |
| 0. 1072 | 0.2144 | 0.3217 | 0.4389 | 0.5361 | 0.6433 | 0.7505 | 0.8577 | 0.9650 | o. 1512 | 12 |
| 0.1075 | 0.2150 | 0.3225 | 0.43C0 | 0. 5375 | 0.6450 | 0.7525 | 0.8600 | 0.9675 | o.1516 | 13 |
| -. 1078 | 0.2156 | 0.3233 | 0.4311 | 0.5389 | 0.6467 | -. 7545 | 0.8622 | c.97co | -. 1520 | 14 |
| 0.1081 | 0.2161 | 0. 3242 | 0.4323 | c. 5403 | 0.6484 | -. 7565 | 0.8645 | 0.9726 | -. 1524 | 15 |
| 0.1084 | 0.2167 | 0.3251 | 0.4334 | 0.5418 | 0.6501 | 0.7585 | 0.8668 | 0.9752 | o. 1528 | 16 |
| 0.1086 | 0.2173 | 0.3259 | 0.4346 | 0. 5432 | 0.6518 | 0.7605 | 0.8691 | 0.9778 | o. 1532 | 17 |
| -. 1089 | 0.2178 | 0.3268 | 0.4357 | 0. 5446 | 0.6535 | 0.7624 | 0.8714 | 0.9803 | -.1536 | 18 |
| c. 1092 | 0.2184 | 0.3276 | 0.4368 | 0.5460 | 0.6552 | 0. 7644 | 0.8736 | 0.9828 | 0. 1540 | 19 |
| 0.1095 | 0.2190 | 0.3285 | 0.4380 | $\bigcirc .5474$ | 0.6569 | 0.7664 | 0.8759 | 0.9854 | 0.1544 | 20 |
| 0.1098 | 0.2195 | 0.3293 |  | 0. 5488 | 0.6586 | 0.7684 | 0.8782 | 79 | 0.1548 | 21 |
| 0.1101 | o. | 0.3302 | 0.4 | 0.5503 | 0.6603 | 0.7704 | 0.8804 | 0.9905 | -. 1552 | 22 |
| 0.1103 | 0.2207 | 0.33 | 0.4414 | 0.5517 | 0.6620 | 0.7724 | 0.8827 | 0.9931 | 0.1556 | 23 |
| 0.1106 | 0.2 | 0.3319 | 0.442 | 0.553 I | 0.6637 | 0.7743 | 0.8850 | 0.9956 | 0.156I | 24 |
| 0.1109 | 0. | 0.332 | 0.4436 | 0.5545 | 0.6654 | 0.7763 | 0.8872 | 0.9981 | 0.1565 | 25 |
| 0.1112 | 0. | 0.3336 | $0.444^{8}$ | -. 555 | 0.6671 | 0.7783 | 0.8889 | 1.0007 | -. 1569 | 6 |
| O. | 0.22 | 0.3344 | 0.4459 | 0.5573 | 0.6688 | 0.7803 | c. 8918 | 1.0032 | -. 1573 | 7 |
| O.II | 0.2235 | 0.3353 | 0.4470 | 0.558 | 0.6705 | 0.7823 | 0.8940 | I. $\cos 8$ | -. 1577 | 28 |
| O. 1 | 0.224 I | 0.3361 | 0.448I | 0.5602 | 0.6722 | 0.7842 | 0.8c63 | 1.0083 | -.158I | 29 |
| 0.1123 | 0.2246 | 0.3370 | 0.4493 | 0.5616 | 0.6739 | 0.7862 | - 08986 | I.oIC9 | -. 1585 | 30 |
| 0.1126 |  | 0.3 | 0.4504 | 0.5630 | 0.6756 | 0.7882 |  | I. 0134 |  | 3 I |
| 0.1129 | 0.2258 | 0.3386 | 0.4515 | 0. 5644 | 0.6773 | 0.7902 | 0.9031 | 1.0159 | -. 1593 | 32 |
| 0.1132 | 0.2263 | 0.3395 | 0.4527 | -. 5659 | 0.6790 | c. 7922 | 0.9054 | I.0185 | -. 1597 | 33 |
| 0.1134 | 0.2269 | 0.3403 | 0.4538 | 0.5673 | 0.6807 | 0.794 | 0.9076 | 1.0210 | 0.1601 | 34 |
| 0.1137 | 0.2275 | 0.3412 | 0.4549 | 0.5687 | 0.6824 | 0.7961 | 0.9098 | 1.0236 | 0.1605 | 35 |
| 0.1140 | 0.2280 | 0.3421 | 0.4561 | 0.5701 | 0.684 I | 0.8081 | 0.9121 | 1.0262 | -.16c9 | 36 |
| 0.1143 | 0.228 | 0.3429 | 0.4572 | 0.5715 | c. 6858 | 0.8001 | 0.9144 | 1.0287 | -.16I3 | 37 |
| 0.1146 | 0.2292 | 0. 3437 | 0.4583 | 0.5729 | 0.6875 | 0.8021 | 0.9166 | 1.0312 | 0.1617 | 38 |
| 0. 1149 | 0.2297 | 0.3446 | 0.4594 | 0.5743 | 0.6892 | 0.8040 | 0.9189 | 1.0337 | 0.1621 | 39 |
| 0.115I | 0.2303 | 0.3454 | 0.4606 | 0.5757 | 0.6909 | 0.8 | 0.9212 | 1.0363 | 0.1625 | 40 |
| 0.1154 | 0.2309 | 0.3463 | 0.46 | 0.577 I | 0.6926 | 0.80 | 0.9234 | 1.0389 |  | 4 |
| 0.1157 | 0.2314 | 0.347 I | 0. 4628 | 0.5786 | 0.6943 | 0.8100 | 0.9257 | 1.0414 | o. 1633 | 42 |
| 0.1160 | 0.2320 | -.3480 | 0.4640 | 0.5800 | 0.6960 | 0.8119 | 0.9279 | I. 0439 | 0. 1637 | 43 |
| 0.1163 | 0.2326 | 0.3488 | 0.4651 | 0.5814 | 0.6977 | 0.8139 | 0.9302 | I. 0465 | 0.1641 | 44 |
| 0.1166 | 0.2331 | 0.3497 | 0.4662 | 0.5828 | 0.6994 | 0.8159 | 0.9325 | 1.0490 | 0.1645 | 45 |
| o. | 0.2337 | 0.3505 | 0.4674 | 0.5842 | 0.7010 | 0.8179 | 0.9347 | 1.0516 | 0.1650 | 46 |
| 0.1171 | 0.2342 | 0.3514 | 0.4685 | 0.5856 | 0.7027 | 0.8199 | 0.9370 | 1.054I | -. 1654 | 47 |
| 0.1174 | 0.2348 | 0.3522 | 0.4696 | 0.5870 | 0.7045 | 0.8219 | 0.9393 | 1.0567 | -. 1658 | 48 |
| 0.1177 | 0.2354 | -0.3531 | 0.4708 | 0.5884 | 0.706 I | $0.823^{8}$ | $0.94{ }^{15}$ | 1.0592 | 0. 1662 | 49 |
| 0.1180 | 0.2359 | 0.3539 | 0.4719 | 0.5899 | 0.7078 | 0.8258 | $0.943^{8}$ | 1.0617 | o. 166 | 50 |
| 0.1183 | 0.2365 | 0.3548 | 0.4730 | 0. 5913 | 0.7095 | 0.8278 | 0.9460 | 1. 0643 | 0.1670 | 51 |
| 0.1185 | 0.2371 | 0.3556 | 0.4742 | 0.5927 | 0.7112 | 0.8298 | 0.9483 | 1.0669 | 0.1674 | 52 |
| O.1188 | 0.2376 | 0.3565 | 0.4753 | 0.5941 | 0.7129 | 0.8317 | 0.9506 | 1.0694 | 0.1678 | 53 |
| O.II91 | 0.2382 0.2388 | 0.3573 0.3581 | 0.4764 | 0.5955 | 0.7146 | 0.8337 | 0.9528 | 1.0719 | 0.1682 0.1686 | 54 |
| 0.1194 | 0.2388 | 0.3581 0.3500 | 0.4775 | 0.5969 | 0.7163 | 0.8357 | 0.9550 | 1.0744 | 0.1686 | 55 |
| 0.1197 | 0.2393 0 | 0.3590 | 0.4786 | 0.5983 | 0.7180 | 0.8376 | 0.9573 | 1.0769 | 0.1690 | 56 |
| 0.1199 | 0.2399 | 0.3598 0.3607 | 0.4798 | 0. 5997 | 0.7197 | 0.8306 | 0.9596 | 1.0795 | $0.1694$ |  |
| 0.1202 | 0.2405 | 0.3607 | 0.4809 | 0.6011 | 0.7214 | 0.8416 | 0.9618 | I. 0821 | 0.1698 0.1702 | 58 |
| 0.1205 0.1208 | 0.2410 0.2416 | 0.3615 0.3624 | 0.4820 0.4832 | 0.6025 0.6040 | 0.7231 0.7247 | 0.8436 0.8455 | 0.9641 0.9663 | 1.0846 1.0871 | 0.1702 0.1706 | 59 |
| 0.1208 | 0.2416 | 0.3624 | 0.4832 | 0.604 | 0.7247 | 0.8455 | 0.9603 | 1.0871 | 0.1706 |  |


| 102 |  | DISTANCES. |  |  |  |  |  |  |  | $7{ }^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | a |
| - | 0.9838 | 1. 9675 | 2.9513 | 3.9351 | 4.9188 | 5.9026 | 6.8864 | 7.8702 | 8.8539 | 1.3896 |
| OI | 0.9837 | 1. 9674 | 2.9511 | 3.9348 | 4.9185 | 5.9022 | 6.8859 | 7.8696 | 8.8533 | 1. 3896 |
| 02 | 0.9836 | 1. 9673 | 2.9509 | 3.9345 | 4.918 I | 5.9018 | 6.8854 | 7.8690 | 8.8526 | 1.3895 |
| 03 | 0.9836 | I. 9671 | 2.9507 | 3.9342 | 4.9178 | 5.9013 | 6.8849 | 7.8684 | 8.8520 | I. 3895 |
| 04 | 0.9835 | I. 9670 | 2.9505 | 3.9339 | 4.9174 | 5.9009 | 6.8844 | 7.8679 | 8.8514 | I. $3^{89} 94$ |
| 05 | 0.9834 | I. 9668 | 2.9502 | 3.9337 | 4.9171 | 5.9005 | 6.8839 | 7.8673 | 8.8507 | I. 3894 |
| 06 | 0.9833 | I. 9667 | 2.9500 | 3.9334 | 4.9167 | 5.9001 | 6.8834 | 7.8667 | 8.8501 | I. 3893 |
| 07 | 0.9833 | 1. 9665 | 2.9498 | 3.9331 | 4.916 | 5.8996 | 6.8829 | 7.8662 | 8.8494 | I. 3893 |
| 08 | 0.9832 | 1. 9664 | 2.9496 | 3.9328 | 4.9160 | 5.8992 | 6.8824 | 7.8656 | 8.8488 | I. 3892 |
| 09 | 0.9831 | 1. 9663 | 2.9494 | 3.9325 | 4.9156 | 5.8988 | 6.8819 | 7.8650 | 8.8482 | I. 3892 |
| 10 | 0.9831 | 1.9661 | 2.9492 | 3.9322 | 4.9153 | 5.8983 | 6.8814 | 7.8645 | 8.8475 | 1.3891 |
| II | 0.9830 | 1.9660 | 2.9490 | 3.9319 | 4.9149 | 5.8979 | 6.8809 | 7.8639 | 8.8469 | I. 3891 |
| 12 | 0.9829 | I. 9658 | 2.9487 | 3.9316 | 4.9146 | 5.8975 | 6.8804 | 7.8633 | 8.8462 | 1.3890 |
| 13 | 0.9828 | I. 9657 | 2.9485 | 3.9314 | 4.9142 | 5.8970 | 6.8799 | 7.8627 | 8.8456 | I. 3890 |
| 14 | 0.9828 | 1. 9655 | 2.9483 | 3.9311 | 4.9138 | 5.8966 | 6.8794 | 7.8621 | 8.8449 | I. 3889 |
| 15 | 0.9827 | I. 9654 | 2.948 I | 3.9308 | 4.9135 | 5.8962 | 6.8789 | 7.8616 | ${ }_{8} 8.8442$ | 1. 3889 |
| 16 | 0.9826 | 1.9652 | 2.9479 | 3.9305 | 4.9131 | 5.8957 | 6.8783 | 7.8610 | 8.8436 | I. 3888 |
| 17 | 0.9825 | 1.9651 | 2.9476 | 3.9302 | 4.9127 | 5.8953 | 6.8778 | 7.8604 | 8.8429 | I. 3888 |
| 18 | 0.9825 | 1.9650 | 2.9474 | 3.9299 | 4.9124 | 5.8949 | 6.8773 | 7.8598 | 8.8423 | 1. 3887 |
| 19 | 0.9824 | I. 9648 | 2.9472 | 3.9296 | 4.9120 | 5.8944 | 6.8768 | 7.8592 | 8.8416 | I. 3887 |
| 20 | 0.9823 | 1.9647 | 2.9470 | 3.9293 | 4.9117 | 5.8940 | 6.8763 | 7.8586 | 8.8410 | I. 3886 |
| 21 | 0.9 | 1. 9645 | 2.9468 | 3.9290 | 4.9113 | 5.8935 | 6.8758 | 7.8580 | 8.8403 | 1. 3886 |
| 22 | 0.9822 | I. 9644 | 2.9465 | 3.9287 | 4.9109 | 5.8931 | 6.8753 | 7.8574 | 8.8396 | I. 3885 |
| 23 | 0.9821 | 1.9642 | 2.9463 | 3.9284 | 4.9105 | 5.8926 | 6.8747 | 7.8569 | 8.8390 | 1. 3885 |
| 24 | 0.9820 | 1.9641 | 2.9461 | 3.928I | 4.9102 | 5.8922 | 6.8742 | 7.8563 | 8.8383 | 1. 3884 |
| 25 | 0.9820 | 1.9639 | 2.9459 | 3.9278 | 4.9098 | 5.8918 | 6.8737 | 7.8557 | 8.8376 | I. 3884 |
| 26 | 0.9819 | 1.9638 | 2.9457 | 3.9275 | 4.9094 | 5.8913 | 6.8732 | 7.8551 | 8.8370 | I. 3883 |
| 2 | 0.9818 | 1.9636 | 2.9454 | 3.9272 | 4.9091 | 5.8909 | 6.8727 | 7.8545 | 8.8363 | 1. 3883 |
| 28 | 0.9817 | 1. 9635 | 2.9452 | 3.9269 | 4.9087 | 5.8904 | 6.8722 | 7.8539 | 8.8356 | 1. 3882 |
| 29 | 0.9817 | 1.9633 | 2.9450 | 3.9266 | 4.9083 | 5.8005 | 6.8716 | 7.8533 | 8.8349 | 1. 3882 |
| 30 | 0.9816 | 1.9632 | 2.9448 | 3.9263 | 4.9079 | 5.8895 | 6.8711 | 7.8527 | 8.8343 | I.388I |
| 3 I | 0.9815 | 1.9630 | 2.9445 | 3.9260 | 4.9076 | 5.8891 | 6.8706 | 7.8521 | 8.8336 | 1.388I |
| 32 | 0.9814 | I. 9629 | 2.9443 | 3.9257 | 4.9072 | 5.8886 | 6.8700 | 7.8515 | 8.8329 | I. 3880 |
| 33 | 0.9814 | I. 9627 | 2.944 I | 3.9254 | 4.9068 | 5.8882 | 6.8695 | 7.8509 | 8.8322 | 1. 3880 |
| 34 | 0.9813 | 1.9626 | 2.9438 | 3.9251 | 4.9064 | 5.8877 | 6.8690 | 7.8503 | 8.8315 | 1. 3879 |
| 35 | 0.9812 | 1.9624 | 2.9436 | 3.9248 | 4.9060 | 5.8872 | 6.8684 | 7.8497 | 8.8309 | 1. $3^{8} 79$ |
| 36 | 0.98 II | 1.9623 | 2.9434 | 3.9245 | 4.9057 | 5.8868 | 6.8679 | 7.8490 | 8.8302 | 1. 3878 |
| 37 | 0.9811 | 1.9621 | 2.9432 | 3.9242 | 4.9053 | 5.8863 | 6.8674 | 7.8484 | 8.8295 | 1.3878 |
| 38 | 0.9810 | 1.9620 | 2.9429 | 3.9239 | 4.9049 | 5.8859 | 6.8669 | 7.8478 | 8.8288 | 1. 3877 |
| 39 | 0.9809 | 1.9618 | 2.9427 | 3.9236 | 4.9045 | 5.8854 | 6.8663 | 7.8472 | 8.828I | 1.3876 |
| 40 | 0.9808 | 1.9617 | 2.9425 | 3.9233 | 4.9041 | 5.8850 | 6.8658 | 7.8466 | 8.8274 | 1. 3876 |
| 4 I | 0.9807 | 1. 9615 | 2.9422 | 3.9230 | 4.9037 | 5.8845 | 6.8652 | 7.8460 | 8.8267 | 1. $3^{8} 75$ |
| 42 | 0.9807 | 1.9613 | $2.94{ }^{20}$ | 3.9227 | 4.9034 | 5.8840 | 6.8647 | 7.8454 | 8. 8260 | I. 3875 |
| 43 | 0.9806 | 1.9612 | 2.9418 | 3.9224 | 4.9030 | 5.8836 | 6.8642 | 7.8.44 | 8. 8253 | 1. 3874 |
| 44 | 0.9805 | 1.9610 | 2.9416 | 3.9221 | 4.902 | 5.8831 | 6.8636 | 7.844 I | 8.8247 | 1. 3874 |
| 45 | 0.9804 | 1.96c9 | 2.9413 | 3.9218 | 4.9022 | 5.8826 | 6.8631 | 7.8435 | 8.8240 | 1.3873 |
| 46 | 0.9804 | 1.9607 | 2.9411 | 3.9214 | 4.9018 | 5.8822 | 6.8625 | 7.8429 | 8.823,3 | 1.3872 |
| 47 | 0.9803 | I. 9606 | 2.9409 | 3.9211 | 4.9014 | 5.8817 | 6.8620 | 7.8423 | 8.8226 | 1.3872 |
| 48 | 0.9802 | 1.9604 | 2.9406 | 3.9208 | 4.9010 | 5.8812 | 6.8614 | 7.8416 | 8.8219 | 1.3871 |
| 49 | 0.9801 | 1.9603 | 2.9404 | 3.9205 | 4.9006 | 5.8808 | 6.8609 | 7.8410 | 8.8212 | 1. 3871 |
| 50 | 0.980I | 1.9601 | 2.9402 | 3.9202 | 4.9003 | 5.8803 | 6.8604 | 7.8404 | 8.8205 | 1. 3870 |
| 51 | 0.9800 | 1.9599 | 2.9399 | 3.9199 | 4.8999 | 5.8798 | 6.8598 | 7.8398 | 8.8197 | 1.3870 |
| 52 | 0.9799 | 1. 9598 | 2.9397 | 3.9196 | 4.8995 | 5. 5794 | 6.8592 | 7.8391 | 8.8190 | I. 3869 |
| 53 | 0.9798 | I. 9596 | 2.9394 | 3.9193 | 4.8991 | 5.8789 | 6.8587 | 7.8385 | 8.8183 | I. 3868 |
| 54 | 0.9797 | 1.9595 | 2.9392 | 3.9189 | 4.8987 | 5.8784 | 6.858 I | 7.8379 | 8.8176 | 1. 3868 |
| 55 | 0.9797 | 1.9593 | 2.9390 | 3.9186 | 4.8983 | 5.8779 | 6.8576 | 7.8372 | 8. 8169 | I. 3867 |
| 56 | 0.9796 | I. 9592 | 2.9387 | 3.9183 | 4.8979 | 5.8775 | 6.8570 | 7.8366 | 8. S162 | I. 3866 |
| 57 | 0.9795 | 1.9590 | 2.9385 | 3.9180 | 4.8975 | 5.8770 | 6.8565 | 7.8360 | 8. SI55 | I. 3866 |
| 58 | 0.9794 | 1. 9588 | 2.9383 | 3.9177 | 4.8971 | 5.8765 | 6.8559 | 7.8353 | 8. 8148 | I. 3865 |
| 59 | 0.9793 | I. 9587 | 2.9380 | 3.9173 | 4.8967 | 5. 8760 | 6.8554 | 7.8347 | S. SI 40 | 1. 3865 |
| 60 | 0.9793 | 1.9585 | 2.9378 | 3.9170 | 4.8963 | 5.8755 | 6.8548 | 7.834 I | 8.SI33 | 1. 3864 |


| $7^{\circ}$ | HEIGHTS. |  |  |  |  |  |  |  | 103 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | b |  |
| 0. 1208 | 0. | 0.3624 | 0.4832 | 0.6040 |  | 0.8455 | 0.966 | 1.0871 |  |  |
| o. | 0.2421 | 0.3632 | 0. 4843 | 0.6054 | 0. 7264 | 0. 8475 | 0. 9686 | 1.0896 | 0.1710 | II |
| o. | 0.2427 | 0.364I | -. 4854 | 0.6c68 | 0.728I | 0.8495 | 0.9709 | 1.0922 | 0.1714 | 02 |
| 0. | 0.2433 | 0. 3649 | 0.4866 | 0.6082 | 0. 7298 | 0.8515 | 0.973 I | 1.0948 | 0.1718 | 03 |
| 0. | 0.2438 | 0. 3658 | 0. 4877 | 0.6196 | -. 7315 | 0. 8534 | 0.9754 | 1. 0973 | 0.1722 | 04 |
| o. | c. 2444 | 0.3666 | 0.4888 | 0.6ıio | 0.7332 | 0.8554 | 0.9776 | I. C 998 | 0.1726 | 05 |
| - | 0.2450 | 0.3674 | 0.4899 | 0.6124 | 0.7349 | 0. 8574 | 0.9798 | I. 1023 | o. 1730 | 6 |
| o. | 0.2455 | 0. 3683 | -.4910 | 0.6138 | 0.7366 | 0. 8593 | 0.9821 | I. 1048 | -. 1734 | 7 |
| O. | 0.246 I | 0.3691 | 0.4922 | 0.6152 | -. 7382 | 0.8613 | 0.9843 | I. 1074 | 0.1738 | 8 |
| 0.1233 | 0. 2467 | 0.3700 | 0.4933 | 0.6166 | 0.7400 | 0.8633 | 0.9866 | 1.1100 | o. 1743 | 9 |
| 0.1236 | 0.2472 | 0.3708 | -. 4944 | 0.6180 | -. 7417 | 0.8653 | 0.9889 | I. 1125 | -. 1747 | 10 |
| 0.1239 | 0.2478 | 0.3717 | 0. 4956 | 0.6194 | 0.7433 | 0.8672 | 0.9911 | 1.1150 | 0.1751 |  |
| 0.1242 | 0.2483 | 0.3725 | 0.4967 | 0.6209 | 0. 7450 | 0.8692 | 0.9934 | I.1175 | 0. 1755 | 12 |
| 0.1245 | 0.2489 | 0.3734 | 0.4978 | 0.6223 | 0. 7467 | 0.8712 | 0.9956 | I. 1201 | -. 1759 | I3 |
| 0.1247 | 0.2495 | 0.3742 | 0.4989 | 0.6237 | 0. 7484 | 0.8731 | 0.9978 | 1.1226 | 0.1763 | 14 |
| 0.1250 | 0.2500 | 0.3750 | 0.5000 | 0.625 I | 0.7501 | 0.8751 | 1.0001 | I. 1251 | -. 1767 | 15 |
| 0.1253 | 0.2506 | 0.3759 | 0.5012 | 0.6265 | 0.7518 | 0.877 I | 1.0024 | I. 1277 | 0.1771 | 16 |
| 0.1256 | 0.2512 | 0.3767 | 0.5023 | 0.6279 | 0. 7535 | 0.8791 | 1.0046 | I. 1302 | 0.1775 | 17 |
| 0.1259 | 0.251 | 0.3776 | 0. 5034 | 0. 6293 | 0.7552 | 0.88ı0 | 1.0069 | I. 1327 | -. 1779 | 18 |
| 0.1261 | 0.2523 | 0.3784 | 0.5046 | 0.6307 | 0.7568 | 0.8830 | 1.0091 | I. 1353 | 0.1783 | 19 |
| 0.1264 | 0.2528 | 0.3793 | 0.5057 | 0.6321 | -. 7585 | 0.8849 | I.oII4 | I. $137^{8}$ | 0.1787 | 20 |
| 0.1267 | 0.25 | 0.3801 | 0.5068 | 0.6335 | 0.7602 | 0.8869 | I.0136 | 1.1403 | I | 21 |
| 0.1270 | 0.25 | 0.3809 | 0. 5079 | 0.6349 | 0.7619 | o. 8889 | 1.crs8 | I. 1428 | -.1795 | 22 |
| 0.1273 | 0. 2545 | 0.3818 | 0.5090 | 0.6363 | 0.7636 | 0.8908 | 1.0181 | I. 1453 | 0.1799 | 23 |
| 0.1275 | 0.255I | 0.3826 | 0.5102 | 0.6377 | 0.7652 | -. 8928 | 1.0203 | I. 1479 | -. 1803 | 24 |
| 0.1278 | 0. 2556 | 0.3835 | -. 5113 | 0.6391 | 0.7669 | 0.8947 | 1.022 | I. 1504 | 0.1807 | 25 |
| 0.1281 | 0.2562 | -0.3843 | 0.5124 | 0.6405 | 0.7680 | 0.8967 | 1.0248 | I. 1529 | 0.1811 | 26 |
| 0.1284 | 0.2568 | -0.3852 | 0.5136 | 0.6419 | c. 7703 | 0.8987 | 1.0271 | I. 1555 | 0.1815 | 27 |
| 0. 1287 | 0.2573 | C. 3860 | 0.5147 | 0.6433 | 0.7720 | 0.9007 | 1.0294 | I. 1580 | 0.1819 | 28 |
| 0.1289 | 0.2579 | -0.3868 | 0.5158 | 0.6447 | 0.7737 | 0.9026 | 1.0316 | I.1605 | 0.1823 | 29 |
| 0.1292 | 0.2585 | 0.3877 | 0. 5169 | 0.6461 | 0.7754 | 0.9046 | 1.0338 | 1.163I | 0.1827 | 30 |
| 0.1295 | 0. 2590 | 0.3885 | 0.5180 | 0.6475 | 0.7771 | 0.9066 | 1.0361 | I. 1656 |  | 3 I |
| 0.1298 | 0.2596 | 0.3894 | 0. 5192 | 0.6489 | 0.7787 | 0.9085 | 1.0383 | I. 1681 | 0.1835 | 32 |
| 0.1301 | 0. 2601 | 0.3902 | 0. 5203 | 0.6503 | 0. 7804 | 0.9105 | 1.0406 | I. 1706 | o. 1839 | 33 |
| 0.1303 | 0. 2607 | 0.3910 | 0. 5214 | 0.6517 | 0. 7821 | 0.9124 | I. 0428 | I. 1731 | o. 1843 | 34 |
| o. 1306 | 0. 2613 | 0.3919 | 0. 5225 | 0.6532 | 0. 7838 | 0.9144 | I. 0450 | I. 1757 | 0. 1847 | 35 |
| -. 1309 | 0. 2618 | 0.3927 | 0.5236 | 0.6546 | 0. 7855 | 0.9164 | 1.0473 | 1. 1782 | o. 1852 | 36 |
| 0.1312 | 0. 2624 | 0. 3936 | 0. 5248 | 0.6560 | 0. 7871 | 0.9183 | I. 0495 | I. 1807 | o. 1856 | 37 |
| o. 1315 | 0. 2629 | 0.3944 | 0. 5259 | 0.6574 | 0. 7888 | 0.9203 | 1.0518 | I. 1832 | -. 1860 | 38 |
| 0.1318 | 0.2635 | 0.3953 | 0. 5270 | 0.6588 | 0.7905 | 0.9223 | 1.0540 | I. 1858 | 0.1864 | 39 |
| 0.1320 | 0.264 I | 0.3961 | 0.5281 | 0.6602 | 0.7922 | 0.9242 | I. 0562 | I. 1883 | 0.1868 | 40 |
| 0.1323 | 0.2646 | 0.3969 | 0.5292 | 0.6616 | 0.7939 | 0.9262 | 1.0585 | 1. 1908 |  | 4 |
| 0.1326 | 0. 2652 | 0. 3978 | 0. 5304 | 0.6630 | 0.7955 | 0.928 I | 1.0607 | I. 1933 | o. 1876 | 42 |
| -. 1329 | 0. 2657 | 0.3986 | 0.5315 | 0.6644 | 0.7972 | 0.9301 | 1.0630 | I. 1958 | 0.1880 | 43 |
| -. 1332 | 0.2663 | 0. 3995 | 0.5326 | 0.6658 | 0.7989 | 0.9321 | 1.0652 | I. 1984 | o. 1884 | 44 |
| 0.1334 | 0. 2669 | 0.4003 | 0.5337 | 0.6672 | 0.80c6 | 0.9340 | 1.0674 | I. 2009 | o. 1888 | 45 |
| -. I337 | 0. 2674 | 0.4011 | 0.5348 | 0.6686 | 0.8023 | 0.9360 | 1.0697 | I. 2034 | 0.1892 | 46 |
| 0.1340 | 0. 2688 | 0.4020 | 0.5360 | 0.6700 | 0.8039 | 0.9379 | 1.0719 | I. 2059 | 0. 1896 | 47 |
| 0.1343 0.1346 | 0. 2685 | 0. 4028 | 0.5371 | 0.6714 | 0.8056 | 0.9399 | 1.0742 | 1. 2084 | 0.1900 | 48 |
| -. 1346 | 0.2691 | 0.4037 | 0.5382 | 0.6728 | 0. 8073 | 0.9419 | 1.0764 | I. 2 | 0. 1904 | 49 |
| -. 1348 | 0. 2697 | 0.4045 | -0.5393 | 0.6742 | 0.8090 | 0.9438 | 1.0786 | I. 2135 | 0. 1908 | 50 |
| 0.1351 | 0.2702 | 0.4053 | c. 5404 | 0.6756 | 0.8107 | 0.9458 | 1.0809 | 1.2160 | 0. 1912 | 51 |
| o. 1354 | 0.2708 | 0. 4062 | 0.5416 | 0.6770 | 0.8123 | 0.9477 | 1.083I | 1.2185 | o. 1916 | 52 |
| O. I357 | 0.2713 | 0. 4070 | 0.5427 | 0.6783 | 0.8140 | 0.9497 | I.c854 | I. 2210 | 0.1920 | 53 |
| -. 1359 | 0.2719 | 0. 4078 | 0. 5438 | 0.6797 | o.8157 | 0.9516 | 1.0876 | I. 2235 | 0.1924 | 54 |
| 0.1362 | 0.2725 | 0.4087 | 0.5449 | 0.681 I | 0.8174 | 0.9536 | 1.0898 | 1. 2261 | 0. 1928 | 55 |
| 0. 1365 | 0.2730 | 0.4095 | 0.5460 | 0.6825 | 0.8191 | 0.9556 | 1.0921 | 1.2286 | o. 1932 | 56 |
| 0.1368 | 0.2736 | 0.4104 | 0. 5472 | 0.6839 | 0. 8207 | 0.9575 | 1.0943 | I. 2311 | o. 1936 | 57 |
| O. 1371 | 0.2741 | 0.4112 | 0.5483 | 0.6853 | 0. 8224 | 0.9595 | I.0966 | I. 2336 | o. 1940 | 58 |
| O. 1374 | 0.2747 | 0.4121 | 0. 5494 | 0.6867 | 0.8241 | 0.9615 | 1.0988 | I. 2362 | o. 1944 | 59 |
| 0.1376 | 0.2753 | 0.4129 | 0.5505 | 0.6881 | 0.8258 | 0.9634 | I. 1010 | 1.2387 | 0.1948 | 60 |


| 104 |  |  |  |  | DISTANCES. |  |  |  |  | $8^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | a |
| $\infty$ | 0.9793 | 1.9585 | 2.9378 | 3.9170 | 4.8963 | 5.8755 | 6.8548 | 7.834 I | 8.8133 | 64 |
| OI | 0.9792 | 1. 9584 | 2.9375 | 3.9167 | 4.8959 | 5.8751 | 6.8542 | 7.8334 | 8.8126 | I. 3863 |
| 02 | 0.9791 | 1.9582 | 2.9373 | $3.91{ }_{4}$ | 4.8955 | 5.8746 | 6.8537 | 7.8328 | 8.8119 | I. 3862 |
| 03 | 0.9790 | 1.9580 | 2.9370 | 3.9161 | 4.895 I | 5.8741 | 6.8531 | 7.8321 | 8.81 II | 1. 3862 |
| 04 | 0.9787 | 1. 9579 | 2.9368 | 3.9157 | 4.8947 | 5.8736 | 6.8525 | 7.8315 | 8.8104 | 1.386r |
| 05 | 0.9789 | 1. 9577 | 2.9366 | 3.9154 | 4.8943 | 5.8731 | 6.8520 | 7.8308 | 8.8097 | 1. 386 I |
| $\infty$ | 0.9788 | r. 9575 | 2.9353 | 3.9151 | 4.8939 | 5.8726 | 6.8514 | 7.8302 | 8.8090 | I. 3860 |
| 07 | 0.9787 | 1.9574 | 2.9361 | 3.9148 | 4.8935 | 5.8722 | 6.8508 | 7.8295 | 8.8082 | I. 3860 |
| 08 | 0.9786 | 1.9572 | 2.9358 | 3.9144 | 4.8931 | 5.8717 | 6.8503 | 7.8289 | 8.8075 | I. 3859 |
| 09 | 0.9785 | 1.9571 | 2.9356 | 3.9141 | 4.8927 | 5.8712 | 6.8497 | 7.8282 | 8.8068 | I. 3859 |
| 10 | 0.9785 | 1.9567 | 2.9354 | 3.9138 | 4.8923 | 5.8707 | 6.8492 | 7.8276 | 8.806 I | 1. 3858 |
| II | 0.9784 | 1. 9567 | 2.9351 | 3.9135 | 4.8918 | $5.87 c 2$ | 6.8485 | 7.8269 | 8.8053 | 1. 3858 |
| 12 | 0.9783 | 1.9565 | 2.9349 | 3.9131 | 4.8914 | 5.8697 | 6.8480 | 7.8263 | 8.8046 | I. 3857 |
| 13 | 0.9782 | 1.9564 | 2.9346 | 3.9128 | 4. S910 | 5.8692 | 6.8474 | 7.8256 | 8.8038 | I. 3856 |
| 14 | 0.9781 | 1. 9562 | 2.9344 | 3.9125 | 4.8936 | ${ }_{5} 5.8687$ | 6.8468 | 7.8250 | 8.8031 | I. 3856 |
| 15 | 0.9780 | 1.956I | 2.9341 | 3.9122 | 4.8902 | 5.8682 | 6.8463 | 7.8243 | 8.8023 | I. 3855 |
| 16 | 0.9780 | 1. 9559 | 2.9339 | 3.9118 | 4.8898 | 5.8677 | 6.8457 | 7.8236 | 8.8016 | 1.3854 |
| 17 | 0.9779 | 1.9557 | 2.9336 | 3.9115 | 4.8894 | 5.8672 | 6.845 I | 7.8230 | 8.8009 | I. 3854 |
| 18 | 0.9778 | 1.9556 | 2.9334 | 3.9112 | 4.8890 | 5.8667 | 6.8445 | 7.8223 | 8.8001 | I. 3853 |
| 19 | 0.9777 | 1.9554 | 2.9331 | 3.9108 | 4.8885 | 5.8662 | 6.8410 | 7.8217 | 8.7994 | 1. 3853 |
| 20 | 0.9775 | I. 9553 | 2.9329 | 3.9105 | 4.888I | 5.8657 | 6.8434 | 7.8210 | 8.7986 | 1.3852 |
| 21 | 0.9775 | 1.9551 | 2.9326 | 3.9102 | 4.8877 | 5.8652 | 6.8428 | 7.8203 | 8.7979 | I. 3852 |
| 22 | 0.9775 | 1.9549 | 2.9324 | 3.9078 | 4.8873 | 5.8647 | 6.8422 | 7.8197 | 8.7971 | 1.3851 |
| 23 | 0.9774 | 1.9547 | 2.9321 | 3.9095 | 4.8869 | $5.86{ }_{42}$ | 6.8416 | 7.8190 | 8.7964 | I. 3850 |
| 24 | 0.9773 | I. 9546 | 2.9319 | 3.9092 | 4.8864 | 5.8637 | 6.8410 | 7.8183 | 8. 7956 | I. 3850 |
| 25 | 0.9772 | 1.9544 | $2.93{ }^{56}$ | 3.9038 | 4.8860 | 5.8632 | 6.8404 | 7.8176 | 8.7948 | I. 3849 |
| 26 | 0.9771 | 1.9542 | 2.9314 | 3.9035 | 4.8856 | 5.8627 | 6.8398 | 7.8170 | 8.7941 | I. 3849 |
| 27 | 0.9770 | 1.9541 | 2.9311 | 3.909 I | 4.8552 | 5.8622 | 6.839 .3 | 7.8163 | 8.7933 | I. 3848 |
| 28 | 0.9770 | 1.9539 | 2.9309 | 3.9078 | 4.8848 | $5.86 \mathrm{I}_{7}$ | 6.8387 | 7.8156 | 8.7926 | I. 3847 |
| 29 | 0.9769 | 1.9537 | 2.9306 | 3.9075 | 4.8843 | 5.8612 | 6.838 I | 7.8I49 | 8.7918 | I. 3847 |
| 30 | 0.9768 | 1.9535 | 2.9304 | 3.907I | 4.8839 | 5.8607 | 6.8375 | 7.8143 | 8.79II | I. 3846 |
| 31 | 0.9767 | 1. 9534 | 2.9301 | 3.9068 | 4.8835 | 5. 8602 | 6.8369 | $7.813^{6}$ |  | I. 38.46 |
| 32 | 0.9766 | 1.9532 | 2.9298 | 3.9064 | 4.8831 | 5.8597 | 6.8353 | 7.8129 | 8.7895 | I. 3845 |
| 33 | 0.9765 | 1.9531 | 2.9296 | 3.9361 | 4.8826 | 5.8592 | 6.8357 | 7.8122 | 8.7887 | I. 3844 |
| 34 | 0.9764 | 1.9529 | 2.9293 | 3.9058 | 4.8322 | 5.8586 | 6.835 I | 7.8115 | 8.7880 | I. $38+4$ |
| 35 | 0.9764 | 1.9527 | 2.9291 | 3.9054 | 4.8818 | 5.858 I | $6.83+5$ | 7.81 | 8.7872 | I. $3^{8} 843$ |
| 36 | 0.9763 | 1.9525 | 2.9288 | 3.9551 | $4 . \mathrm{SSI}_{3}$ | 5.8576 | 6.8339 | 7.8102 | 8.7864 | I. 3843 |
| 37 | 0.9762 | 1.9524 | 2.9285 | 3.9047 | 4.8809 | 5.8571 | 6.8333 | 7.8095 | 8.7856 | I. 3842 |
| 38 | 0.9761 | 1.9522 | 2.9283 | 3.9044 | 4.8805 | 5.8566 | 6.8327 | 7.8 | 8.7849 | I. 384 I |
| 39 | 0.9760 | 1.9520 | 2.9280 | 3.9040 | 4.8801 4.8796 | 5. | 6.8321 6.8315 | 7.8081 | 8.7841 8.7833 | I. 384 II I. 3840 |
| 40 | 0.9759 | 1.9519 | 2.9 | 3.9037 |  | 5. | 6.8315 | 7.8074 | 8.7833 |  |
| 41 | 0.9758 | 1.9517 | 2.9275 | 3.9034 | 4.8792 | 5.8550 | 6.8309 | 7.8067 | 8.7826 | 1.3840 |
| 42 | 0.9758 | 1.9515 | 2.9273 | 3.9030 | 4.8788 | 5.8545 | 6.8303 | 7.8060 | 8.7818 | 1. 3839 |
| 43 | 0.9757 | 1.9513 | 2.9270 | 3.9027 | 4.8783 | 5.8540 | 6.8296 | 7.8053 | 8.78io | I. 3838 |
| 44 | 0.9756 | 1.9512 | 2.9267 | 3.9023 | 4.8779 | 5.5535 | 6.8290 | 7.8046 | 8.7802 | I. 3838 |
| 45 | 0.9755 | 1.9510 | 2.9265 | 3.9020 | 4.8774 | 5.8529 | $6.828_{4}$ | 7.8039 | $8.779+$ | I. 3837 |
| 46 | 0.9754 | 1.9508 | 2.9262 | 3.9016 | 4.8770 | 5.8524 | 6.8278 | 7.8032 | 8.7786 | I. 3837 |
| 47 | 0.9753 | 1.9506 | 2.9259 | 3.9013 | 4.8765 | 5.8519 | 6.8272 | 7.8025 | 8. 7778 | I. 3886 |
| 48 | 0.9752 | 1.9505 | 2.9257 | 3.9009 | 4.8761 | 5.8514 | 6.8266 | 7.8018 | 8.7770 | I. 3835 |
| 49 | 0.9751 | 1.9503 | 2.9254 | 3.9006 | 4.8757 | 5.8508 | 6.8260 | 7.8011 | 8. 7763 | I. 3835 |
| 50 | 0.9751 | 1.9501 | 2.9252 | 3.9002 | 4.8753 | 5.8503 | 6.8254 | 7.8007 | 8.7755 | I. 3834 |
| 51 | 0.9750 | 1. 9499 | 2.9249 | 3. 8999 | 4.8748 | 5. 4.498 | 6.8247 | 7.7997 | 8.7747 | I. 3834 |
| 52 | 0.9749 | I. 9497 | 2.9246 | 3.8995 | 4.8744 | 5. S 492 | $6.82+1$ | 7.7990 | 8.7739 | I. $3^{833}$ |
| 53 | 0.9748 | 1.9496 | 2.924 | 3.8991 | 4.8739 | 5. 8487 | 6.8235 | 7.7983 | 8.7731 | I. 3832 |
| 54 | 0.9747 | I. 9494 | 2.9241 | 3. SgSS | 4.8735 | 5.8482 | 6.8229 | 7.7976 | 8. 1723 | I. 3832 |
| 55 | 0.9746 | 1.9492 | 2.9238 | 3.SgS4 | 4.8730 | 5.8476 | 6.8222 | 7.7969 | 8.7715 | I. 3831 |
| 56 | 0.9745 | 1.9490 | 2.9236 | 3.898 I | 4.8726 | 5. 8471 | 6.8216 | 7.7961 | 8.7707 | I. $3^{3} 31$ |
| 57 | 0.9744 | $1.9+89$ | 2.9233 | 3.8977 | 4.8721 | 5.8466 | 6.8210 | 7.7954 | 8. 7699 | I. 3830 |
| 58 | c. 9743 | 1.9487 | 2.9230 | 3. 8974 | 4.8717 | 5. 8.460 | 6.8204 | 7.7947 | 8. 7691 | I. 3 S29 |
| 59 60 | 0.9743 | I. 9485 I. $9+83$ | 2.9228 2.9225 | 3.8970 3.8966 | 4.8713 4.8708 | 5.8455 5.8450 | 6.8198 6.8191 | 7.7940 7.7933 | 8.7683 8.7675 | I. 3829 I. 3828 |


| $8^{\circ}$ | HEIGHTS. |  |  |  |  |  |  |  |  | 105 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | b |  |
|  | 0. | 0.4129 | 0.5 | 0.6 |  |  | I. 1010 | 7 | 8 | $\bigcirc$ |
| -.1380 | $0.275^{8}$ | 0.4137 | 0.5516 | 0.6895 | 0.8275 | 0.9654 | I. 1033 | 1.2412 | 0. 1952 | 1 |
| c. $133^{82}$ | 0.2764 | 0.4145 | 0.5527 | 0.6909 | c. 8291 | 0.9673 | I. 1055 | 1. 2437 | -. 1956 | 02 |
| 0.1385 | 0.2769 | 0.4154 | 0. 5538 | 0.6923 | 0.8308 | 0.9692 | 1.1077 | I. 2462 | 0.1960 | 03 |
| 0.1387 | 0.2775 | 0.4162 | 0.5550 | 0.6937 | 0.8324 | 0.9712 | I. 1099 | I. 2487 | -. 1965 | 04 |
| 0.1390 | 0.2780 | 0.4171 | 0.556I | 0.6951 | 0.834 | 0.9731 | I. 1122 | I. 2512 | 0.1969 | 05 |
| 0.1393 | 0.2786 | 0.4179 | 0.5572 | 0.6965 | 0.8358 | 0.9751 | I. 11 | I. 2537 | -. 1973 | 6 |
| 0.1396 | 0.2792 | 0.4187 | 0.5583 | 0.6979 | 0. 8375 | 0.9771 | I. 11 | I. 2562 | -. 1977 | 7 |
| 0.1399 | 0.2797 | 0.4196 | 0.5594 | 0.6993 | 0.8392 | 0.9790 | I. 1189 | 1.2587 | -.1981 | 8 |
| 0.1401 | 0.2803 | 0.4204 | 0.5606 | 0. 7007 | 0.8408 | 0.9810 | I. 1211 | I. 2613 | 0. 1985 | 9 |
| 0.1404 | 0.2808 | 0.4213 | 0.5617 | 0.702I | 0.8425 | 0.9829 | I. 1234 | 1. 2638 | -. 1989 | Io |
|  |  |  |  |  |  |  | 1. 1256 |  |  | 1 |
| 0.1410 | 0.2819 | 0.4229 | 0.5639 | 0.7049 | 0.8458 | 0.9868 | I. 1278 | 1. 2688 | 0.1597 | 12 |
| 0.1413 | 0.2825 | 0.4238 | 0.5650 | 0. 7063 | 0.8475 | 0.9888 | 1.1300 | 1.2713 | OOI | I3 |
| 0.1415 | 0.2831 | 0.4246 | 0.5661 | 0.7077 | 0. 8492 | 0.9907 | 1.1322 | 1.2738 | 0.2005 | 14 |
| 0.1418 | 0.2836 | 0.4254 | 0.5672 | 0.7091 | 0.8509 | 0.9927 | I. 1345 | 1.2763 | 0.2609 | 5 |
| 0.1421 | 0.2842 | 0.4263 | 0.5684 | 0.7104 | 0.8525 | 0.9946 | I. 1367 | 1.2788 | 0.2013 | 6 |
| 0.1424 | 0.2847 | 0.4271 | 0.5695 | 0.7118 | 0.8542 | 0.9966 | I. 1390 | 1.2813 | 0.2017 | 17 |
| 0.1426 | 0.2853 | 0.4279 | 0.5706 | 0.7132 | 0. 8558 | c. 9985 | I. 1412 | 1.2838 | 0.202 | 8 |
| 0.1429 | 0.2858 | 0.4288 | 0.5717 | 0.7146 | 0.8575 | 1.0005 | I. 1434 | 1.2863 | 0.2025 | 19 |
| 0.1432 | 0.2864 | c. 4296 | 0.5728 | 0.7160 | 0.8592 | 1. 0024 | I. 1456 | 1. 2888 | 0.2029 |  |
| 0.1435 | 0.2870 | 0.43 C 4 | 0.5739 | 0.7174 | c. 8609 | 1.0044 | 1. 1478 | 1.2913 | 2033 | 1 |
| 0.14 | 0. 2875 | 0.4313 | 0.5750 | 0.7188 | 0.8626 | 1.co6 3 | I. 1501 | I. 2938 | 0.2037 | 22 |
| 0.1440 | 0.288 I | 0.432 | 0.5762 | 0.7202 | 0.8642 | 1.0083 | I. 1523 | 1.2963 | 0.204I | 23 |
| 0.1443 | 0.2886 | 0.4329 | 0.5773 | 0.7216 | 0.8659 | I. OIO | I. 1545 | I. 2988 | 0.2045 | 24 |
| 0.1 | 0.2892 | $0.433^{8}$ | 0.5784 | 0.7230 | 0.8675 | . 01 | I. 1567 | I.3013 | 0.2049 | 25 |
| 0.1449. | 0.2897 | 0.4346 | 0.5795 | 0.7243 | 0.8692 | I.cI4I | I. 1590 | I. 3038 | 0.2053 | 26 |
| o. | 0.2903 | 0.4354 | 0.5806 | 0.7257 | 0.8709 | 1.0160 | I. 1612 | I. 3063 | 0.2057 | 27 |
| o. | 0.29 | 0.4363 | 0.5817 | 0.7271 | 0.8726 | 1.or80 | 1.1634 | 1. 3088 | 0.206I | 28 |
| 0.1457 | 0.291 | 0.4371 | 0.5828 | 0.7285 | 0.8742 | 1.0199 | I. 1656 | I. 3113 | 0. 2065 | 29 |
| 0.1460 | 0.2920 | 0.4379 | 0.5839 | 0. 7299 | c. 8759 | 1.0219 | I. 1678 | 1.3138 | 0.2569 | 30 |
| 0. 1463 | 0.29 | 0.438 | 0.5850 | 0.7313 | 0.8776 | 1.0238 | 1.1701 |  | 73 | 31 |
| 0. 146 | 0.2931 | 0.4396 | 0.5862 | 0.7327 | 0.8792 | 1.0258 | I. 1723 | 1.3188 | 77 | 32 |
| 0.146 | 0.2936 | 0.4404 | c. 5873 | 0.7341 | 0.8809 | 1.0277 | I. 1745 | I. 3213 | 0.208 I | 33 |
| o. | 0.2942 | 0.44 I 3 |  |  | 0.8825 | 1.0206 |  | 1.3238 | 0.2085 | 34 |
| o. | 0.2947 | 0.442 I |  | 0.7368 | c. 88842 | 1.0316 | 1. 1790 |  | 0. 2089 | 35 |
| 0.1476 | 0.2953 |  |  |  | 0.8859 | 1.0335 | I. 1812 | I. 3288 | 0.2093 | 36 |
| o. 1479 | 0.2958 | 0.4438 | 0. 5 | 0.7396 | 0.8875 | 1.0355 | 1. 1834 | 1.3313 | 0.2097 | 37 |
| 0.1482 | 0.2954 | 0.4446 | 0. 5928 | 0.7410 | 0.8892 | 1.0374 | 1. 1856 | 1.3338 | 0.2101 |  |
| o. 1485 | 0.2970 |  | 0. 5939 |  | 0.8909 | 1.0394 | 1.1878 |  | 0.2105 | 39 |
| 0. 1488 | 0.2975 | 0.4463 | 0.5950 | 0.7438 | 0.8926 | 1.04I3 | 1.1901 | I. 3388 | 0.2110 |  |
|  |  | 0.4471 |  |  |  | 1.0432 | 1. 1923 |  |  | 1 |
| 0.1493 | 0.2986 | 0.4479 | 0.5972 | 0.7466 | 0.8959 | 1.0452 | I. 1945 | I. 3438 | 8 | 42 |
| 0.1496 | 0.2992 | 0.4488 | 0.5984 | 0.7479 | 0. 8975 | 1.0471 | I. 1967 | I. 3463 | 0.2122 | 43 |
| 0.1499 | 0.2997 | 0.4496 | 0.5995 | 0.7493 | 0:8992 | 1.0491 | I. 1989 | I. 3488 | c. 2126 | 44 |
| 0.1501 | 0.300 | 0.4504 | 0.6006 | -0.7507 | 0.9008 | 1.0510 | I. 20 | I.3513 | 0.2130 | 5 |
| 0. 1504 | 0.3008 | 0.4513 | 0.6017 | 0.7521 | 0.9025 | 1.0529 | I. 2034 | I. 3538 | 0.2134 | 46 |
| 0.1507 | 0.3014 | $0.45{ }^{21}$ | 0.6028 | 0.7535 | 0.9042 | I. 05 | I. 2056 | I. 35 | $0.213^{8}$ | 47 |
| 0.1510 | 0.3019 | 0.4529 | 0.6039 | 0. 7549 | 0.9058 | 1.0568 | I. 2078 | 1. 3588 | 0.2142 | 88 |
| -.1513 | 0.3025 | 0.4538 | 0.6050 | 0.7563 | 0.9075 | 1.0588 | 1.2100 | 1.3613 | 0.2146 | 49 |
| 0.1515 | 0.3031 | 0.4546 | 0.606I | 0.7576 | 0.9092 | 1.0607 | I. 21 | 1. 3638 | 0.2150 | 50. |
| 0.1518 | 0.3036 | 0.4554 | 0.6072 | 0.7590 | 0.9108 | 1.06 | I. 2 | 1. 3662 | 0.2154 | 51 |
| 0.1521 | 0.3042 | c. 4562 | 0.6083 | 0.7654 | c.9125 | 1.0646 | 1. 2166 | I. 3687 | 0.2158 | 52 |
| O. 1524 | 0.3047 | 0.4571 | 0.6094 | 0.7618 | c.9142 | 1.0665 | 1.2189 | I. 3712 | 0.2162 | 53 |
| 0.1526 | 0.3053 | 0.4579 | 0.6105 | 0. 7632 | 0.9158 | 1.0684 | 1.221I | r. 3737 | 0.2166 | 54 |
| 0.1529 | 0. 3058 | 0.4587 | 0.6ır6 | 0.7646 | 0.9175 | I. 0704 | I. 2233 | I. 3762 | 0.2170 | 55 |
| 0.1532 | 0. 3064 | 0.4596 | 0.6128 | 0. 7660 | 0.9191 | 1.0723 | I. 2255 | I. 3787 | 0.2174 | 56 |
| O.I535 0.1537 | 0. 3069 | 0.4604 0.4612 | 0.6139 | 0.7673 | 0.9208 | 1.0742 | I. 2277 | r. 3812 r. 3837 | 0.2178 0.2182 | 57 |
| O. 1537 0.1540 | 0.3075 0.3080 | $\begin{aligned} & 0.4612 \\ & 0.4621 \end{aligned}$ | 0.6150 <br> 0.616I | 0.7687 0.7701 0.771 | 0.9224 0.9241 | 1.0762 1.0781 | 1.2299 I. 2321 | I. 3837 1. 3862 | 0.2182 0.2186 | 58 |
| 0.1540 0.1543 | 0.3080 0.3086 | 0.4621 0.4629 | 0.6161 0.6172 | 0.7701 0.7715 | 0.9241 0.9257 | 1.0781 1.08co | 1.2321 1.2343 | 1.3862 1.3866 | 0.2190 | 59 60 |


| 106 |  |  |  |  | DISTANCES. |  |  |  |  | $9^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | a |
| 00 | 0.9742 | 1.9483 | 2.9225 | 3.8966 | 4.8708 | 5.8450 | 6.8191 | 7.7933 |  | 1. 3828 |
| OI | 0.9741 | 1.948I | 2.9222 | 3.8963 | 4.8704 | 5.8444 | 6.8185 | 7.7926 | 8.7666 | r. 3827 |
| 02 | 0.9740 | 1.9480 | 2.9219 | 3.8959 | 4.8699 | 5.8439 | 6.8179 | 7.7918 | 8.7658 | r. 3826 |
| 03 | 0.9739 | 1.9478 | 2.9217 | 3.8956 | 4.8695 | 5.8433 | 6.8172 | 7.7911 | 8.7650 | r. 3826 |
| 04 | 0.9738 | 1.9476 | 2.9214 | 3.8952 | 4.8690 | 5.8428 | 6.8166 | 7.79c4 | 8.7642 | 1. 3825 |
| 05 | 0.9737 | 1.9474 | 2.9211 | 3.8948 | 4.8686 | 5.8423 | 6.8160 | 7.7897 | 8.7634 | 1. 3825 |
| 06 | 0.9736 | 1.9472 | 2.9209 | 3.8945 | 4.8681 | 5.8417 | 6.8153 | 7.7890 | 8.7626 | 1. 3824 |
| 07 | 0.9735 | I. 9471 | 2.9206 | 3.8941 | 4.8676 | 5.8412 | 6.8147 | 7.7882 | 8.7618 | I. 3824 |
| 08 | 0.9734 | 1. 9469 | 2.9203 | 3.8938 | 4.8672 | 5.8406 | 6.814 I | 7.7875 | 8.7609 | 1. $3^{823}$ |
| 09 | 0.9733 | 1.9467 | 2.9200 | 3.8934 | 4.8667 | 5.8401 | 6.8134 | 7.7858 | 8.7601 | I. 3822 |
| 10 | 0.9733 | 1.9465 | 2.9198 | 3.8930 | 4.8663 | 5.8395 | 6.8128 | 7.7861 | 8.7593 | 1.3821 |
| 11 | 0.9732 | 1.9463 | 2.9195 | 3.8927 | 4.8658 | 5.8390 | 6.8122 | 7.7853 | 8.7585 | I. 382 I |
| 12 | 0.9731 | 1.9461 | 2.9192 | 3.8923 | 4.8654 | 5.8384 | 6.8115 | 7.7846 | 8.7577 | 1.3820 |
| 13 | $0.973{ }^{\circ}$ | 1.9460 | 2.9189 | 3.8919 | 4.8649 | 5.8379 | 6.8109 | 7.7838 | 8.7568 | r. $3^{819}$ |
| 14 | 0.9729 | I. 9458 | 2.9187 | 3.8916 | 4.8644 | 5.8373 | 6.8102 | 7.7831 | 8.7560 | I. $3^{819}$ |
| 15 | 0.9728 | I. 9456 | 2.9184 | 3.8912 | 4.8640 | 5.8368 | 6.8096 | 7.7824 | 8.7552 | 1.3818 |
| 16 | 0.9727 | I. 9454 | 2.918I | 3.8908 | 4.8635 | 5.8362 | 6.8089 | 7.7816 | 8.7543 | 1.3818 |
| 17 | 0.9726 | I. 9452 | 2.9178 | 3.8904 | 4.8631 | 5.8357 | 6.8083 | 7.7809 | 8.7535 | 1.3817 |
| 18 | 0.9725 | I. 9450 | 2.9176 | 3.8901 | 4.8626 | 5.8351 | 6.8076 | 7.7802 | 8.7527 | I. 3816 |
| 19 | 0.9724 | ז. 9449 | 2.9173 | 3.8897 | 4.86 | 5.8346 | 6.8070 | 7.7794 | 8.7518 | I. 3816 |
| 20 | 0.9723 | 1. 9447 | 2.9170 | 3.8893 | 4.8617 | 5.8340 | 6.8063 | 7.7787 | 8.7510 | r. $3^{815}$ |
| 21 | 0.9722 | 1.9445 | 2.9167 |  | 4.8612 | 5.8334 | 6.8057 | 7.7779 | 8.7502 | r. $3^{814}$ |
| 22 | 0.9721 | 1.9443 | 2.9164 | 3.8886 | 4.8607 | 5.8329 | 6.8050 | 7.7772 | 8.7493 | 1.3814 |
| 23 | 0.9721 | 1.9441 | 2.9162 | 3.8882 | 4.8603 | 5.8323 | 6.8044 | 7.7764 | 8.7485 | 1.3813 |
| 24 | 0.9720 | 1.9439 | 2.9159 | 3.8878 | 4.8598 | 5.8318 | 6.8037 | 7.775: | 8.7476 | r. 3813 |
| 25 | 0.9719 | 1.9437 | 2.9156 | 3.8875 | 4.8593 | 5.8312 | 6.8031 | 7.7749 | 8.7468 | r. $3^{812} 2$ |
| 26 | 0.9718 | 1.9435 | 2.9153 | 3.8871 | 4.8589 | 5.8306 | 6.8024 | 7.7742 | 8.7460 | r. 38 IrI |
| 27 | 0.9717 | 1.9434 | 2.9150 | 3.8867 | 4.8584 | 5.8301 | 6.8018 | 7.7734 | 8.7451 | I. 3811 |
| 28 | 0.9716 | 1.9432 | 2.9148 | 3.8863 | 4.8579 | 5.8295 | 6.8011 | 7.7727 | 8.7443 | I. 3810 |
| 29 | 0.9715 | 1.9430 | 2.9145 | 3.8860 | 4.8575 | 5.8290 | $6.8 \mathrm{co4}$ | 7.7719 | 8.7434 | I. 3810 |
| 30 | 0.9714 | 1.9428 | 2.9142 | 3.8856 | 4.8570 | 5.8284 | 6.7998 | 7.7712 | 8.7426 | 1.3809 |
| 31 | 0.9713 | 1.9426 | 2.9139 | 3.8852 | 4.8565 | 5.8278 | 6.7991 | 7.7704 | 8.7417 | I. 3808 |
| 32 | 0.9712 | 1. $9+24$ | 2.9136 | 3. $884_{4} 8$ | 4.8560 | 5.8272 | 6.7984 | 7,7697 | 8.7409 | I. 3808 |
| 33 | 0.9711 | 1.9422 | 2.9133 | 3.8844 | 4.8556 | 5.8267 | 6.7978 | 7.7689 | 8.7400 | I. 3807 |
| 34 | 0.9710 | 1. 9420 | 2.9130 | 3.8841 | 4.8551 | 5.8261 | 6.7971 | 7.7681 | 8.7391 | I. 3806 |
| 35 | $0.97 c 9$ | I.9418 | 2.9128 | 3.8837 | 4.8546 | 5.8255 | 6. 7964 | 7.7674 | 8.7383 | I. 3806 |
| 36 | 0.9708 | I. 9417 | 2.9125 | 3.8833 | 4.854 I | 5.8250 | 6.7958 | 7.7666 | 8.7374 | I. 3 So5 |
| 37 | 0.9707 | I. 9415 | 2.9122 | 3.8829 | 4.8537 | 5.8244 | 6.7951 | 7.7658 | 8.7366 | I. $3 \mathrm{SO}_{4}$ |
| 38 | 0.9706 | 1.9413 | 2.9119 | 3.8825 | 4.8532 | 5.8238 | 6.7944 | 7.7651 | 8.7357 | I. 3 SO4 |
| 39 | 0.9705 | 1.9411 | 2.9116 | 3.8822 | 4.8527 | 5.8232 | 6.7938 | 7.7643 | 8.7349 | I. 3803 |
| 40 | 0.9704 | 1.9407 | 2.9113 | 3.8818 | 4.8522 | 5.8227 | 6.7931 | 7.7636 | 8. 7340 | 1. 3 SO2 |
| 41 | 0.9703 | 1.9407 | 2.9110 | 3.8814 | 4.8517 | 5.822I | 6.7924 | 7.7628 | 8.7331 | 1. 3802 |
| 42 | 0.9703 | I. 9405 | 2.9108 | 3.8810 | 4.8513 | 5.8215 | 6.7918 | 7.7620 | 8.7323 | r. 3801 |
| 43 | 0.9702 | I. 9403 | 2.9105 | 3.8806 | 4.8508 | 5.8209 | 6.7911 | 7.7612 | 8.7314 | I. 3800 |
| 44 | 0.9701 | 1.9401 | 2.9102 | 3.8802 | 4.8503 | 5.8203 | 6.7904 | 7.7604 | 8.7305 | I. 3799 |
| 45 | 0.9700 | 1.9397 | 2.9099 | 3.8798 | 4.8498 | 5.8198 | 6.7897 | 7.7597 | 8. 7296 | 1. 3799 |
| 46 | 0.9699 | I. 9397 | 2.9096 | 3.8794 | 4.8493 | 5.8192 | 6.7890 | 7.7589 | 8.7288 | 1. 3798 |
| 47 | 0.9698 | 1.9395 | 2.9093 | 3.8791 | 4.8488 | 5.8186 | 6.7884 | 7.7581 | 8.7279 | I. 3797 |
| 48 | 0.9697 | ז. 9393 | 2.9090 | 3.8787 | 4.8483 | 5.8180 | 6.7877 | 7.7573 | 8.7270 | r. 3797 |
| 49 | 0.9696 | 1.9391 | 2.9087 | 3.8783 | 4.8479 | 5.8174 | 6.7870 | 7.7566 | 8.7261 | 1. 3796 |
| 50 | 0.9695 | 1.9389 | 2.9084 | 3.8779 | 4.8474 | 5.8168 | 6.7863 | 7.7558 | 8.7253 | 1. 3795 |
| 51 | 0.9694 | 1. 9388 | 2.9 c 8 I | 3.8775 | 4.8469 | 5.8163 | 6.7856 | 7.7550 | 8.7244 | 1. 3795 |
| 52 | 0.9693 | 1. 9386 | 2.9078 | 3.8771 | 4.8464 | 5.8157 | 6.7849 | 7.7542 | 8. 7235 | 1.3794 |
| 53 | 0.9692 | 1.9384 | 2.9075 | 3.8767 | 4.8459 | 5.8151 | 6.7843 | 7.7534 | 8.7226 | 1.3793 |
| 54 | 0.9691 | 1.9382 | 2.9072 | 3.8763 | 4.8454 | 5.8145 | 6.7836 | 7.7526 | 8. 7217 | 1.3792 |
| 55 | 0.9690 | 1.9380 | 2.9069 | 3.8759 | 4.8449 | 5.8139 | 6. 7829 | 7.7519 | 8. 7208 | 1.3792 |
| 5 | 0.9689 | 1. 9378 | 2.9066 | 3.8755 | 4.8444 | 5.8133 | 6. 7822 | 7.7511 | 8.7199 | I. 3791 |
| 57 | 0.9688 | 1. 9376 | 2.9064 | 3.8751 | 4.8439 | 5.8127 | 6.7815 | 7.7503 | 8.7191 | 1.3790 |
| 58 | 0.9687 | 1.9374 | 2.9061 | 3.8747 | 4.8434 | 5.8121 | 6.7808 | 7.7495 | 8. 7182 | 1.3789 |
| 6 | 0.9686 | 1.9372 | 2.9058 | 3.8744 | 4.8429 | 5. 8115 | 6.7801 | 7.7487 | 8.7173 | 1. 3 - -89 |
|  | 0.9685 | 1.9370 | 2.9055 | 3.8740 | 4.8424 | 5.810 | 6.7794 | 7.7479 | 8.7164 | 1.370 |


| $9^{\circ}$ | HEIGHTS. |  |  |  |  |  |  |  | 107 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | b | , |
| C. I543 | 0.3086 | 0.4629 | 0.6172 | 0.7715 | 0.9257 | 1.0800 | 1.2343 | 1. 3886 | 0.2190 | $\infty$ |
| O. 1546 | 0.3091 | 0.4637 | 0.6183 | 0.7729 | 0.9274 | 1.0820 | I. 2365 | $\text { I. } 39 \text { II }$ | $0.2194$ | OI |
| O. 1548 | 0.3097 | 0.4645 | 0.6194 | 0.7742 | 0.9290 | 1.0839 | 1.2387 | I. 3936 | 0.2198 | 02 |
| 0.155I | 0.3102 | 0.4654 | 0.6205 | 0.7756 | 0.9307 | I. 0858 | 1.2410 | I.3961 | 0.2202 | 03 |
| 0.1554 | 0.3108 | 0.4662 | 0.6216 | 0.7770 | 0.9324 | 1.0878 | I. 2432 | I. 3986 | 0.2206 | 04 |
| O.I557 | 0.3II3 | 0.4670 | 0.6227 | 0.7784 | 0.9340 | 1.0897 | I. 2454 | 1.4010 | 0.2210 | 05 |
| O.1559 | 0.3119 | 0.4678 | 0.6238 | 0.7797 | 0.9357 | 1.0916 | 1.2476 | 1. 4035 | 0.2214 | 06 |
| 0.1562 | 0.3124 | 0.4687 | 0.6249 | 0.78 II | 0.9373 | 1.0936 | I. 2498 | 1.4060 | 0.2218 | 07 |
| -. 1565 | 0.3130 | 0.4695 | 0.6260 | 0.7825 | 0.9390 | 1.0955 | 1.2520 | 1.4085 | 0.2222 | 08 |
| 0.1568 | 0.3136 | 0.4703 | 0.6271 | 0. 7839 | 0.9407 | I. 0975 | 1.2542 | 1.4110 | 0.2226 | 09 |
| O.I57I | 0.314 I | 0.4712 | 0.6282 | 0.7853 | 0.9423 | 1. 0994 | 1.2564 | 1.4135 | 0.2230 | IO |
| 0.1573 | 0.3147 | 0.4720 | 0.6293 | 0.7866 | 0.9440 | I. IOI3 | I. 2586 | 1.4160 | 0.2234 | II |
| 0.1576 | 0.3152 | 0.4728 | 0.6304 | 0. 7880 | 0.9456 | I. 1032 | 1.2608 | 1.4184 | 0.2238 | 12 |
| O.I579 | 0.3158 | 0.4736 | 0.6315 | 0.7894 | 0.9473 | I. 1052 | I. 2630 | 1.4209 | 0.2242 | I3 |
| 0.1582 | 0.3163 | 0.4745 | 0.6326 | 0.7908 | 0.9489 | I. 107I | I. 2652 | 1. 4234 | 0.2246 | 14 |
| 0.1584 | 0.3169 | 0.4753 | 0.6337 | 0.7922 | 0.95c6 | I. 1090 | 1. 2674 | 1. 4259 | 0.2250 | 15 |
| O. 1587 | 0.3174 | 0.4761 | 0.6348 | 0.7935 | 0.9523 | 1.1110 | 1.2697 | 1. 4284 | 0.2254 | I6 |
| 0.1590 | 0.3180 | 0.4769 | 0.6359 | 0. 7949 | 0.9539 | I. 1129 | 1.2719 | 1.4308 | 0.2258 | 17 |
| 0.1593 | 0.3185 | 0.4778 | 0.6370 | 0.7963 | 0.9556 | I. 1148 | 1.274I | 1.4333 | 0.2262 | 18 |
| $0.1595$ | 0.3191 | 0.4786 | 0.6381 | 0.7977 | 0.9572 | 1.1167 | 1.2763 | 1. 4358 | $0.2266$ | 19 |
| 0.1598 | 0.3196 | 0.4794 | 0.6392 | 0.7991 | 0.9589 | I.1187 | 1.2785 | 1.4383 | 0.2270 | 20 |
| 0.1601 | 0.3202 | 0.4802 | 0.6403 | 0.8004 | 0.9505 | I. 1206 | 1.2807 | 1.4407 | 0.2274 | 2 I |
| 0.1604 | 0.3207 | 0.48 II | 0.6414 | 0.8018 | 0.9622 | I. 1225 | 1. 2829 | 1.4432 | $0.2278$ | 22 |
| $0.1606$ | 0.3213 | 0.4819 | 0.6425 | 0.8032 | 0.9638 | I. 1244 | 1.285I | I. 4457 | $0.2282$ | 23 |
| 0.1609 | 0.3218 | 0.4827 | 0.6436 | 0.8046 | 0.9655 | I. 1264 | 1. 2873 | 1.4482 | 0.2287 | 24 |
| $0.1612$ | 0.3224 | 0.4835 | 0.6447 | 0.8059 | 0.9671 | 1.1283 | 1. 2895 | I. 4506 | 0.2291 | 25 |
| $0.16 I 5$ | 0.3229 | 0.4844 | 0.6458 | 0.8073 | 0.9688 | I. 1302 | 1.2917 | I. 4531 | 0.2295 | 26 |
| 0.1617 | 0.3235 | 0.4852 | 0.6469 | 0.8087 | 0.9704 | I. 1321 | 1.2939 | 1.4556 | 0.2299 | 27 |
| 0.1620 | 0.3240 | 0.4860 | 0.6480 | 0.8100 | 0.9721 | I. 1341 | I.2961 | I.4581 | 0.2303 | 28 |
| $0.1623$ | 0.3246 | 0.4868 | 0.6491 | $0.8114$ | 0.9737 | I. 1360 | 1. 2983 | $\text { 1. } 4605$ | $0.2307$ | 29 |
| 0.1626 | 0.3251 | 0.4877 | 0.6502 | 0.8128 | 0.9754 | I. 1379 | 1.3005 | 1.4630 | 0.23 II | 30 |
| 0. 1628 | 0.3257 | 0.4885 | 0.6513 | 0.8142 | 0.9770 | I. 1398 | 1. 3027 | 1. 4655 | 0.2315 | 31 |
| $0.1631$ | 0.3262 | 0.4893 | 0.6524 | 0.8155 | 0.9787 | I. 1418 | I. 3049 | I. 4680 | 0.2319 | 32 |
| $0.1634$ | 0.3268 | $0.4901$ | 0.6535 | 0.8169 | 0.9803 | I. 1437 | I. 307 I | 1.4704 | 0.2323 | 33 |
| 0.1637 | 0.3273 | 0.4910 | 0.6546 | 0.8183 | 0.9819 | I. 1456 | 1. 3092 | I. 4729 | 0.2327 | 34 |
| 0.1639 | 0.3279 | 0.4918 | 0.6557 | 0.8196 | 0.9836 | I. 1475 | I.3II4 | 1.4754 | 0.233 I | 35 |
| $0.1642$ | 0.3284 | 0.4926 | 0.6568 | 0.8210 | 0.9852 | I. 1494 | 1.3136 | 1. 4778 | 0.2335 | 36 |
| $0.1645$ | 0.3290 | 0.4934 | 0.6579 | 0.8224 | 0.9869 | I. 1514 | I. 3158 | 1. 4803 | $0.2339$ | 37 |
| 0.1648 | 0.3295 | 0.4943 | 0.6590 | 0.8238 | 0.9885 | I. 1533 | 1.3180 | 1. 4828 | 0.2343 | 38 |
| 0.1650 | 0.3301 | 0.4951 | 0.6601 | 0.8251 | 0.9902 | I. 1552 | I. 3202 | 1. 4853 | 0.2347 | 39 |
| 0.1653 | 0.3306 | 0.4959 | 0.6612 | 0.8265 | 0.9918 | 1. 157 I | 1.3224 | 1. 4877 | 0.2351 | 40 |
| 0.1656 | 0.3311 | 0.4967 | 0.6623 | 0.8279 | 0.9934 | I. 1590 | I. 3246 | 1.4901 | 0.2355 | 4I |
| $0.1658$ | 0.3317 | 0.4975 | 0.6634 | 0.8292 | 0.9951 | 1.1609 | I. 3268 | 1.4926 | $0.2359$ | 42 |
| $0.166 I$ | 0.3322 | 0.4984 | 0.6645 | 0.8306 | 0.9967 | I. 1629 | I. 3290 | $\text { 1. } 495 \mathrm{I}$ | $0.2363$ | 43 |
| 0.1664 | 0.3328 | 0.4992 | 0.6656 | 0.8320 | 0.9984 | I. 1648 | 1.3312 | 1. 4976 | 0.2367 | 44 |
| 0.1667 | 0.3333 | 0.5000 | 0.6667 | -0.8334 | 1.0000 | I. 1667 | I. 3334 | 1.5000 | 0.2371 | 45 |
| 0.1669 | 0.3339 | 0.5008 | 0.6678 | 0.8347 | 1.0016 | 1.1686 | I. 3355 | 1.5025 | 0.2375 | 46 |
| 0.1672 | 0.3344 | 0.5017 | 0.6689 | 0.8361 | 1.0033 | 1.1705 | I. 3378 | 1.5050 | 0.2379 | 47 |
| 0.1675 | 0.3350 | 0.5025 | 0.6700 | 0.8375 | $\text { I. } 0549$ | I. 1724 | I. 3399 | $1.5074$ | $0.2383$ | 48 |
| 0.1678 0.1680 | 0.3355 0.336 I | 0.5033 0.5041 | 0.671I | 0.8388 | 1.0066 | I. I743 | I. 3421 | I. 5098 | 0.2387 | 49 |
| 0.1680 | 0.3361 | 0.504 I |  |  | I.0082 | 1.1763 | I. 3443 | 1.5123 | 0.2391 | 50 |
| 0.1683 | 0.3366 | 0.5049 | 0.6732 | 0.8416 | 1.0099 | I. 1782 | 1.3465 | 1.5148 | 0.2395 | 51 |
| 0.1686 | 0.3372 | 0.5057 | 0.6743 | 0.8429 | I.OII5 | 1.1801 | I. 3487 | 1.5172 | 0.2399 | 52 |
| 0.1689 | 0.3377 | 0.5066 | 0.6754 | 0.8443 | I.OI32 | I. 1820 | 1.3509 | 1.5197 | 0.2403 | 53 |
| 0.1691 | 0.3383 | $0.5074$ | $0.6765$ | 0.8457 | 1.0148 | 1. 1839 | 1.353I | 1.5222 | 0.2407 | 54 |
| 0.1694 | 0.3388 | 0.5082 | 0.6776 | 0.8470 | 1.0165 | I. 1859 | I. 3553 | I. 5247 | 0.2411 | 55 |
| 0.1697 | 0.3394 | 0.5090 | 0.6787 | 0.8484 | I.OI8I | I. 1878 | I. 3574 | I. 5271 | 0.2415 | 56 |
| 0.1700 | 0.3399 | 0.5099 | 0.6798 | 0.8498 | I.0197 | I. 1897 | r. 3596 | I. 5296 | 0.2419 | 57 |
| $0.1702$ | 0.3404 | 0.5107 | $0.6809$ | 0.8511 | 1.0213 | I.1916 | $\text { I. } 3618$ | I. 5320 | $0.2423$ | 58 |
| 0.1705 | 0.3410 | 0.5115 | 0.6820 | 0.8525 | I. 0230 | I. 1935 | I. 3640 | I. 5345 | 0.2427 | 59 |
| 0.1708 | 0.3415 | 0.5123 | 0.6831 | 0.8539 | 1.0246 | I. 1954 | 1.3662 | I. 5369 | $0.243^{\text {I }}$ | 60 |


| 108 |  |  |  |  | DISTANCES. |  |  |  |  | $10^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | a |
| $\infty$ | 0.9685 | 1.93 | 2.9055 | 3.8740 | 4.8424 | 5.8109 | 6.7794 | 7.7479 | 8.7164 | 1.3788 |
| OI | 0.9684 | 1.9368 | 2.9052 | 3.8736 | 4.8419 | 5.8103 | 6.7787 | 7.747 I | 8.7155 | $\begin{aligned} & \text { I. } 3787 \end{aligned}$ |
| 02 | 0.9683 | 1. 9366 | 2.9049 | 3.8732 | 4.8414 | 5.8097 | 6.7780 | 7.7463 | 8.7146 | 1. 3786 |
| 03 | 0.9682 | 1. 9364 | 2.9046 | 3.8728 | 4.8409 | 5.8091 |  | 7.7455 | 8.7137 | 1. 3786 |
| 04 | 0.9681 | 1.9362 | 2.9043 | 3.8724 | 4.8404 | 5.8085 | 6.7766 | 7.7447 | 8.7128 | 1. 3785 |
| c5 | 0.9680 | 1.9360 | 2.9040 | 3.8720 | 4.8399 | 5.8079 | 6.7759 | 7.7439 | 8.7119 | 1. 3784 |
| 06 | 0.9679 | 1.9358 | 2.9037 | 3.8716 | 4.8394 | 5.8073 | 6.7752 | 7.7431 | 8.7110 | 1. 3783 |
| 07 | 0.9678 | 1.9356 | 2.9034 | 3.8712 | 4.8389 | 5.8067 | 6.7745 | 7.7423 | 8.7101 | 1.3783 |
| 08 | 0.9577 | 1. 9354 | 2.9031 | 3.8707 | 4.8384 | 5.806I | 6.7738 | 7.7415 | 8.7092 | 1. 3782 |
| 09 | 0.9676 | 1.9352 | 2.9028 | 3.8703 | 4.8379 | 5.8055 | 6.7731 | 7.7407 | 8.7083 | 1. 3781 |
| 10 | 0.9675 | 1.9350 | 2.9025 | 3.8699 | 4.8374 | 5.8049 | 6.7724 | 7.7399 | 8.7074 | 1.3780 |
| II | 0.9674 | 1. 9348 | 2.9022 | 3.8695 | 4.8369 | 5.8043 | 6.7717 | 7.7391 | 8.7065 | 1.3780 |
| 12 | 0.9673 | 1. 9346 | 2.9019 | 3.8691 | 4.8364 | 5.8037 | 6.7710 | 7.7383 | 8.7056 | 1. 3779 |
| 13 | 0.9672 | I. 9344 | 2.9015 | 3.8687 | 4.8359 | 5.8031 | 6.7703 | 7.7375 | 8.7046 | 1. 3778 |
| 14 | 0.9671 | I. 9342 | 2.9012 | 3.8683 | 4.8354 | 5.8025 | 6.7606 | 7.7366 | 8.7037 | I. 3777 |
| 15 | 0.9670 | I. 9340 | 2.9009 | 3.8679 | 4.8349 | 5.8019 | 6.7689 | 7.7358 | 8.7028 | 1. 3776 |
| 16 | 0.9669 | 1.9338 | 2.9006 | 3.8675 | 4.8344 | 5.8011 | 6.768I | 7.7350 | 8.7019 | 1. 3776 |
| 17 | 0.9668 | 1. 9336 | 2.9003 | 3.8671 | 4.8339 | 5.8007 | 6.7674 | 7.7342 | 8.7010 | I. 3775 |
| 18 | 0.9667 | I. 9333 | 2.9000 | 3.8667 | 4.8334 | 5.8000 | 6.7667 | 7.7334 | 8.7001 | I. 3774 |
| 19 | 0.9666 | I. 9331 | 2.8997 | 3.8663 | 4.8329 | 5.7994 | 6. 7660 | 7.7326 | 8.6991 | 1. 3773 |
| 20 | 0.9665 | 1.9329 | 2.8994 | 3.8659 | 4.8324 | 5.7988 | 6.7653 | 7.7318 | 8.6982 | 1.3773 |
| 21 | 0.9664 | 1.9327 | 2.8991 | 3.8655 | 4.8318 | 5.7982 | 6. 7646 | 7.7309 | 8.6973 | 1.3772 |
| 22 | 0.9663 | I. 9325 | 2.8988 | 3.8651 | 4.8313 | 5.7976 | 6.7638 | 7.7301 | 8.6964 | 1.3771 |
| 23 | 0.9662 | 1.9323 | 2.8985 | 3.8646 | 4.8308 | 5.7970 | 6.7631 | 7.7293 | 8.6954 | 1. 3770 |
| 24 | 0.9661 | 1.9321 | 2.8992 | 3.8642 | 4.8303 | 5.7963 | 6.7624 | 7.7285 | 8.6945 | I. 3769 |
| 25 | 0.9660 | 1.9319 | 2.8979 | 3.8638 | 4.8298 | 5.7957 | 6.7617 | 7.7276 | 8.6936 | 1. 3769 |
| 26 | 0.9659 | 1.9317 | 2.8976 | 3.8634 | 4.8293 | $5 \cdot 7951$ | 6.7610 | 7.7268 | 8.6927 | I. 3768 |
| 27 | 0.9657 | 1.9315 | 2.8972 | 3.8630 | 4.8287 | 5.7945 | 6.7602 | 7.7260 | 8.6917 | 1. 3767 |
| 28 | 0.9656 | 1.9313 | 2.8969 | 3.8626 | 4.8282 | 5.7939 | 6.7595 | 7.7252 | 8.6508 | 1. 3766 |
| 29 | 0.9655 | 1.93II | 2.8966 | 3.8622 | 4.8277 | 5.7932 | 6.7588 | 7.7243 | 8.6859 | I. 3765 |
| 30 | 0.9654 | 1.9309 | 2.8963 | 3.8617 | 4.8272 | 5.7926 | 6.758 I | 7.7235 | 8.6889 | 1. 3765 |
| 31 | 0.9653 | 1.9307 | 2.8960 | 3.8613 | 4.8267 | 5.7920 | 6.7573 | 7.7227 | 8.6880 | 1. 3764 |
| 32 | 0.9652 | 1.9305 | 2.8957 | 3.8609 | 4.8261 | 5.7914 | 6.7566 | 7.7218 | 8.6870 | I. 3763 |
| 33 | 0.9651 | 1.9302 | 2.8954 | 3.8605 | 4.8256 | 5.7907 | 6.7559 | 7.7210 | 8.6861 | 1. 3762 |
| 34 | 0.9650 | 1.9300 | 2.8951 | 3.8601 | 4.8251 | 5.7901 | 6.7551 | 7.7201 | 8.6852 | 1.3761 |
| 35 | 0.9649 | I. 9298 | 2.8947 | 3.8597 | 4.8246 | 5.7895 | 6.7544 | 7.7193 | 8.6842 | I. 3761 |
| 36 | 0.9648 | I. 9296 | 2.8944 | 3.8592 | 4.8240 | 5.7888 | 6.7537 | 7.7185 | 8.6833 | 1. 3760 |
| 37 | 0.9647 | 1.9294 | 2.8941 | 3.5588 | 4.8235 | 5.7882 | 6.7529 | 7.7176 | 8.6823 | I. 3759 |
| 38 | 0.9646 | 1.9292 | 2.8938 | 3.8584 | 4.8230 | 5.7876 | 6.7522 | 7.7168 | 8.6814 | 1. 3759 |
| 39 | 0.9645 | 1.9290 | 2.8935 | 3.8580 | 4.8225 | 5.7870 | 6.7515 | 7.7159 | 8.6804 | 1. 3758 |
| 40 | 0.9644 | 1.9288 | $2.893{ }^{2}$ | 3.8576 | 4.8219 | 5.7863 | 6.7507 | 7.7151 | 8.6795 | 1. 3757 |
| 41 | 0.9643 | 1. 9286 | 2.8928 | 3.8571 | 4.8214 | 5.7857 | 6.7500 | 7.7143 | 8.6785 | 1. 3756 |
| 42 | 0.9642 | 1. 9284 | 2.8925 | 3.8567 | 4.8209 | 5.7851 | 6.7492 | 7.7134 | 8.6776 | 1. 3755 |
| 43 | 0.9641 | 1. 9281 | 2.8922 | 3.8563 | 4.8203 | 5.7844 | 6.7485 | 7.7126 | 8.6766 | I. 3755 |
| 44 | 0.9640 | 1.9279 | 2.8919 | 3.8558 | 4.8198 | 5.7838 | 6.7477 | 7.7117 | 8.6757 | 1. 3754 |
| 45 | 0.9639 | 1.9277 | 2.8916 | 3.8554 | 4.8193 | 5.7831 | 6.7470 | 7.7108 | 8.6747 | 1. 3753 |
| 46 | 0.9638 | 1. 9275 | 2.8912 | 3.8550 | 4.8187 | 5.7825 | 6.7462 | 7.7100 | 8.6737 | 1. 3752 |
| 47 | 0.9636 | I. 9273 | 2.8909 | 3.8546 | 4.8182 | 5.7819 | 6.7455 | 7.7091 | 8.6728 | 1. 3752 |
| 48 | 0.9635 | 1. 9271 | 2.8506 | 3.8541 | 4.8177 | 5.7812 | 6. 7448 | 7.7083 | 8.6718 | 1. 3751 |
| 49 | 0.9634 | 1. 9269 | 2.8903 | 3.8537 | 4.8172 | 5.7806 | 6.7440 | 7.7074 | 8.6709 | I. 3750 |
| 50 | 0.9633 | 1. 9266 | 2.8900 | 3.8533 | 4.8166 | 5.7799 | 6.7433 | 7.7066 | 8.6699 | I. 3749 |
| 51 | 0.9632 | 1.9264 | 2.8896 | 3.8529 | 4.8161 | 5.7793 | 6.7425 | 7.7057 | 8.6689 | I. 3748 |
| 52 | 0.9631 | 1. 9262 | 2.8893 | 3.8524 | 4.8155 | 5.7786 | 6.7417 | 7.7049 | 8.6680 | I. 3748 |
| 53 | 0.9630 | 1.9260 | 2.8890 | 3.8520 | 4.8150 | 5.7780 | 6.7410 | 7.7040 | 8.6670 | 1. 3747 |
| 54 | 0.9629 | 1. 9258 | 2.8887 | 3.8516 | 4.8145 | 5.7773 | 6.7402 | 7.7031 | 8.6660 | 1. 3746 |
| 55 | 0.9628 | 1.9256 | 2.8883 | 3.8511 | 4.8139 | 5.7767 | 6.7395 | 7.7023 | 8.6650 | 1.3745 |
| 56 | 0.9627 | I. 9254 | 2.8880 | 3.8507 | 4.8134 | 5.7761 | 6.7387 | 7.7014 | 8.6641 | 1.3744 |
| 57 | 0.9626 | I. 9251 | 2.8877 | 3.8503 | 4.8128 | 5.7754 | 6.7380 | 7.7005 | 8.6631 | I. 3744 |
| 58 | 0.9625 | I. 9249 | 2.8874 | 3.8498 | 4.8123 | 5.7748 | 6. 7372 | 7.6997 | 8.6621 | 1.3743 |
| 5 | 0.9624 | 1.9247 | 2.8871 | 3.8494 | 4.8118 | 5.7741 | 6.7365 | 7.6988 | 8.6612 | 1.3742 |
| 60 | 0.9622 | 1.9245 | 2.8867 | 3.8490 | 4.8112 | 5.7735 | 6.7357 | 7.6979 | 8.660 | 1.3742 |


| $10^{\circ}$ | HEIGHTS. |  |  |  |  |  |  |  |  | 109 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | b | , |
| 0.1708 | 0.3415 | 0.5123 | 0.6831 | 0.8539 | 1.0246 | 1. 1954 | 1.3662 | 1.5369 . | 0.2431 | 00 |
| 0.1710 | 0.3421 | 0.5131 | 0.6842 | 0.8552 | 1.0262 | 1. 1973 | 1.3683 | 1. 5394 | 0.2435 | OI |
| 0.1713 | 0.3426 | 0.5140 | 0.6853 | 0.8566 | 1.0279 | 1. 1992 | I. 3705 | I. 5419 | 0.2439 | 02 |
| -. 1716 | 0.3432 | 0.5148 | 0.6864 | 0.8580 | 1.0295 | 1.2 CII | I. 3727 | 1.5443 | 0.2443 | 03 |
| O.1719 | 0.3437 | 0.5156 | 0.6875 | 0.8593 | 1.0312 | 1.2030 | I. 3749 | 1.5468 | 0.2447 | 04 |
| 0.1721 | 0.3443 | 0.5164 | 0.6886 | 0.8607 | 1.0328 | 1.2050 | I. 3771 | 1.5493 | 0.245 I | 05 |
| O. 1724 | 0.3448 | 0.5173 | 0.6896 | 0.8620 | 1.0345 | 1.2069 | I. 3793 | 1.5517 | 0.2455 | c6 |
| c. 1727 | 0. 3454 | 0.5180 | 0.6907 | 0.8634 | 1.0361 | 1. 2088 | I. 3814 | I.5541 | 0.2459 | 07 |
| 0.1730 | 0.3459 | 0.5189 | 0.6918 | 0.8648 | 1.0377 | 1.2107 | 1.3836 | 1.5566 | 0.2463 | 08 |
| O. 1732 | 0.3464 | 0.5197 | 0.6929 | 0.866 I | 1.0393 | 1.2126 | I. 3858 | I. 5590 | 0.2467 | 09 |
| O. 1735 | 0.3470 | 0.5205 | 0.6940 | 0.8675 | 1.0410 | 1.2145 | I. 3880 | I.5615 | 0.247 I | 10 |
| 0.1738 | 0.3475 | C. 5213 | 0.6951 | 0.8688 | 1.0426 | 1.2164 | 1.3902 | 1.5639 | 0.2475 | II |
| 0.1740 | 0.348 I | 0.5221 | 0.6962 | 0.8702 | 1.0442 | 1.2183 | I. 3923 | I. 5664 | 0.2479 | 12 |
| 0.1743 | 0.3486 | 0.5229 | 0.6973 | 0.8716 | 1.0459 | 1.2202 | 1. 3945 | I. 5688 | 0.2483 | I3 |
| 0.1746 | 0.3492 | 0.5238 | 0.6984 | 0.8729 | I. 0475 | 1.222I | I. 3967 | 1.5713 | 0.2487 | 14 |
| 0.1749 | 0.3497 | 0.5246 | 0.6994 | 0.8743 | 1.0492 | 1.2240 | 1.3989 | 1.5737 | 0.2491 | 15 |
| 0.1751 | 0.3503 | 0.5254 | 0.7005 | 0.8757 | 1.0508 | 1.2259 | 1.4010 | 1.5762 | c. 2495 | 16 |
| 0.1754 | 0.3508 | 0.5262 | 0.7016 | c. 8770 | 1.0524 | 1.2278 | 1.4032 | 1.5786 | 0.2499 | 17 |
| 0.1757 | 0.3513 | 0.5270 | 0.7027 | c. 8784 | 1.0540 | 1.2297 | 1. 4054 | 1.5810 | 0.2503 | 18 |
| $\text { ०. } 1759$ | 0.3519 | 0.5278 | 0.7038 | 0.8797 | 1.0557 | 1.2316 | 1.4076 | I. 5835 | $0.2507$ | 19 |
| 0.1762 | 0.3524 | 0.5287 | 0.7049 | 0.88 II | 1.0573 | 1.2335 | 1.4098 | 1.5860 | 0.2511 | 20 |
| 0.1765 | 0.3530 | 0.5295 | 0.7060 | 0.8824 | 1.0589 | 1.2354 | 1.4119 | 1. 5884 | 0.2515 | 21 |
| $0.1768$ | 0.3535 | 0.5303 | 0.7070 | 0.8838 | 1.0606 | 1.2373 | 1.4141 | 1.5908 | 0.2519 | 22 |
| 0.1770 | 0.354I | 0.5311 | 0.7081 | 0.8852 | 1.0622 | 1.2392 | 1.4162 | 1.5933 | 0.2523 | 23 |
| 0.1773 | 0.3546 | 0.5319 | 0.7092 | 0.8865 | 1.0638 | 1.2411 | I. 4184 | 1.5957 | 0.2527 | 24 |
| 0.1776 | 0.3552 | 0.5327 | 0.7103 | 0.8879 | 1.0655 | 1.2430 | 1.4206 | $1.5932$ | $0.253 \mathrm{I}$ | 25 |
| 0.1778 | 0.3557 | 0.5335 | 0.7114 | 0.8892 | 1.0671 | 1.2449 | $\text { I. } 4228$ | $1.6006$ | $0.2535$ | 26 |
| $0.178 \mathrm{I}$ | 0.3562 | 0.5344 | 0.7125 | 0.8906 | 1.0687 | 1.2468 | I. 4250 | 1.6031 | 0.2539 | 27 |
| 0.1784 | 0.3568 | 0.5352 | 0.7136 | 0.8920 | 1.0703 | 1.2487 | 1.4271 | 1. 6055 | 0.2543 | 28 |
| $0.1787$ | 0.3573 | 0.5360 | 0.7146 | 0.8933 | r.0720 | 1.2506 | I. 4293 | $1.6079$ | $0.2547$ | 29 |
| 0.1789 | 0.3579 | 0.5368 | 0.7157 | 0.8947 | 1.0736 | 1.2525 | 1.4314 | 1.6104 | 0.2551 | 30 |
| O.I792 | 0.3584 | 0.5376 | 0.7168 | 0.8960 | 1.0752 | 1.2544 | 1. 4336 | 1.6ı28 | 0.2555 | $3{ }^{1}$ |
| 0.1795 | 0.3590 | 0.5384 | 0.7179 | 0.8974 | 1.0769 | 1.2563 | 1. 4358 | 1.6153 | $0.2559$ | 32 |
| 0.1797 | $0.3595$ | 0.5392 | 0.7190 | 0.8987 | 1.0785 | 1.2582 | 1.4380 | $\text { 1. } 6177$ | $0.2563$ | 33 |
| 0.1800 | $0.3600$ | $0.5401$ | 0.7201 | 0.9001 | 1.0801 | 1.2601 | $\text { 1. } 4402$ | $1.6202$ | $0.2567$ | 34 |
| $0.1803$ | 0.3606 | c.54c9 | 0.7212 | 0.9014 | 1.0817 | 1.2620 | 1.4423 | $1.6226$ | $0.2571$ | 35 |
| $0.1806$ | $0.3611$ | 0.5417 | 0.7222 | 0.9028 | 1.0834 | 1.2639 | I. 4445 | $1.6250$ | $0.2575$ | 36 |
| 0.1808 | 0.3617 | 0.5425 | 0.7233 | 0.9041 | 1.0850 | I. 2658 | I. 4466 | 1. 6275 | 0.2579 | 37 |
| 0.18II | 0.3622 | 0.5433 | 0.7244 | 0.9055 | 1.0866 | 1.2677 | I. 4488 | 1. 6299 | 0.2583 | $3^{8}$ |
| $0.1814$ | 0.3627 | 0.5441 | 0.7255 | c. 9069 | I.c882 | 1.2696 | 1.4510 | $1.6323$ | $0.2587$ | 39 |
| 0.1816 | 0.3633 | 0.5449 | 0.7266 | 0.9082 | I.0898 | 1.2715 | I.453I | 1. 6348 | 0.2591 | 40 |
| O.1819 | 0.3638 | 0.5457 | 0.7276 | 0.9096 | 1.C915 | 1.2734 | 1. 4553 | 1.6372 | 0.2595 | 4 I |
| 0.1822 | 0.3644 | 0.5465 | 0.7287 | 0.9109 | 1.0931 | 1.2753 | I. 4574 | 1. 6396 | $0.2599$ | 42 |
| $0.1825$ | 0.3649 | $0.5474$ | 0.7298 | 0.9123 | 1.0947 | 1.2772 | $\text { 1. } 4596$ | $1.6421$ | $0.2603$ | 43 |
| 0.1827 | 0.3654 | 0.5482 | 0.7309 | 0.9136 | 1.0963 | 1.2790 | 1.4618 | 1.6445 | 0.2607 | 44 |
| 0.1830 | 0.3660 | 0.5490 | 0.7320 | 0.9150 | 1.0979 | 1.2809 | 1. 4639 | 1.6469 | 0.2611 | 45 |
| 0.1833 | 0.3665 | 0.5498 | 0.7330 | 0.9163 | 1.c996 | 1.2828 | 1.4661 | $1.6493$ | $0.2615$ | 46 |
| 0.1835 | 0.3671 | 0.5506 | 0.7341 | 0.9177 | 1.1012 | 1.2847 | 1.4682 | $1.6518$ | $0.26 \mathrm{I} 9$ | 47 |
| 0.1838 | 0.3676 | 0.5514 | 0.7352 | 0.9190 | 1. 1028 | 1.2866 | 1.4704 | 1.6542 | 0.2623 | 48 |
| 0.1841 | 0.3681 | 0.5522 | 0.7363 | 0.9204 | I. 1044 | 1.2885 | I. 4726 | 1. 6566 | 0.2627 | 49 |
| 0.1843 | 0.3687 | 0.5530 | 0. 7374 | 0.9217 | 1.106I | 1.2904 | 1.4748 | 1.6591 | 0.2631 | 50 |
| 0.1846 | 0.3692 | 0.5538 | 0.7384 | 0.9231 | 1.1077 | 1.2923 | 1.4769 | 1.66I5 | 0.2635 | 51 |
| $\text { 0. } 1849$ | 0.3598 | 0.5546 | 0.7395 | 0.9244 | I. I593 | 1.2942 | 1.4790 | 1. 6639 | $0.2639$ | 52 |
| 0.1852 | 0.3703 | 0.5555 | 0.7406 | 0.9258 | I. 1109 | 1.2961 | 1.4812 | I $\cdot 6664$ | 0.2643 | 53 |
| 0.1854 | 0.3708 | 0.5563 | 0.7417 | 0.9271 | I.1125 | $1.2979$ | I. 4834 | 1.6688 | 0.2647 | 54 |
| 0.1857 | 0.3714 | 0.5571 | 0.7428 | 0.9285 | I.II4I | 1.2998 | .1. 4855 | 1.6712 | 0.2651 | 55 |
| 0.1860 0.1862 | 0.3719 | 0.5579 | 0.7438 | 0.9298 | I.II58 | 1.3017 | 1.4877 | 1.6736 | 0.2655 | 56 |
| O.1862 | 0.3725 | 0.5587 | 0.7449 | 0.9312 | 1.1174 | 1. 3036 | 1.4898 | 1. 6761 | 0.2659 | 57 |
| O.I865 0.1868 | 0.3730 0.3735 | 0.5595 | 0.7460 | 0.9325 | I.II90 | I. 3055 | 1. 4920 | $\text { 1. } 6785$ | $0.2663$ | 58 |
| 0. 1868 | 0.3735 | 0.5603 | 0.7471 | 0.9339 | 1.1206 | I. 3074 | 1. 4942 | 1.6809 | $0.2667$ | 59 |
| 0.1870 | 0.3741 | 0.5611 | 0.7482 | 0.9352 | 1.1222 | 1.3093 | 1.4963 | 1.6834 | 0.2671 | 60 |


| 110 |  | DISTANCES. |  |  |  |  |  |  |  | $11^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | 1 | 2 | 3 | 4 | 5 | 6 | \% | 8 | 9 | a |
| 00 | 0.9622 | 1.9245 | 2.8867 | 3.8490 | 4.8112 | 5.7735 | 6.7357 | 7.6979 | 8.6602 | 742 |
| OI | 0.9621 | 1.9243 | 2.8864 | 3.8485 | 4.8107 | 5.7728 | 6.7349 | 7.6971 | 8.6592 | I. 374 I |
| 02 | 0.9620 | 1. 9240 | 2.8861 | 3.8481 | 4.8 IOI | 5.7721 | 6.7342 | 7.6962 | 8.6582 | r. 3740 |
| 03 | 0.96I9 | r. 9238 | 2.8857 | 3.8477 | 4.8096 | 5.7715 | 6.7334 | 7.6953 | 8.6572 | r. 3739 |
| 04 | 0.9618 | 1.9236 | 2.8854 | 3.8472 | 4.8090 | 5.7708 | 6.7326 | 7.6944 | 8.6562 | I. 3738 |
| 05 | 0.9617 | 1.9234 | 2.8851 | 3.8468 | 4.8085 | 5.7702 | 6.7319 | 7.6936 | 8.6553 | r. 3738 |
| 06 | 0.9616 | 1.9232 | 2.8848 | 3.8463 | 4.8079 | 5.7695 | 6.7311 | 7.6927 | 8.6543 | 1. 3737 |
| 07 | 0.9615 | 1.9230 | 2.8844 | 3.8459 | 4.8074 | 5.7689 | 6.7303 | 7.6918 | 8.6533 | I. 3736 |
| 08 | 0.9614 | 1.9227 | 2.8841 | 3.8455 | 4.8068 | 5.7682 | 6.7296 | 7.6909 | 8.6523 | 1. 3735 |
| 09 | 0.9613 | 1.9225 | 2.8838 | 3.8450 | 4.8063 | 5.7675 | 6.7288 | 7.6 gor | 8.6513 | 1. 3734 |
| 10 | 0.96 II | 1.9223 | 2.8834 | 3.8446 | 4.8057 | 5.7669 | 6.7280 | 7.6892 | 8.6503 | I. 3734 |
| II | 0.961 | 1.9221 | 2.8831 | 3.844 I | 4.8052 | 5.7662 | 6.7273 | 7.6883 | 8.6493 | 1. 3733 |
| 12 | 0.9609 | 1.9218 | 2.8828 | 3.8437 | 4.8046 | 5.7655 | 6.7265 | 7.6874 | 8.6483 | 1. 3732 |
| 13 | 0.9608 | 1.9216 | 2.8824 | 3.8433 | 4.8041 | 5.7649 | 6.7257 | 7.6865 | 8.6473 | I. 3731 |
| 14 | 0.9607 | 1.9214 | 2.8821 | 3.8428 | 4.8035 | 5.7642 | 6.7249 | 7.6856 | 8.6463 | I. 3730 |
| 15 | 0.9606 | 1.9212 | 2.8818 | 3.8424 | 4.8030 | 5.7635 | 6.7241 | 7.6847 | 8.6453 | 1. 3730 |
| 16 | 0.9605 | 1.9210 | 2.8814 | 3.8419 | 4.8024 | 5.7629 | 6.7234 | 7.6838 | 8.6443 | 1.3729 |
| 17 | 0.9604 | 1.9207 | 2.88 II | 3.84 I5 | 4.8018 | 5.7622 | 6.7226 | 7.6830 | 8.6433 | 1.3728 |
| 18 | 0.9603 | 1.9205 | 2.8808 | 3.8410 | 4.8013 | 5.76I5 | 6.7218 | 7.6821 | 8.6423 | 1. 3727 |
| 19 | 0.9601 | 1.9203 | 2.8804 | 3.8406 | 4.8007 | 5.7609 | 6.7210 | 7.6812 | 8.6413 | 1.3726 |
| 20 | 0.96 co | 1.9201 | 2.88 כI | 3.8401 | 4.8 coz | 5.7602 | 6.7202 | 7.6803 | 8.6403 | 1. 3726 |
| 21 | 0.959 | 1.9198 | 2.8798 | 3.8397 | 4.7996 | 5.7595 | 6.7195 | 7.6794 | 8.6393 | 1. 3725 |
| 22 | 0.9598 | 1.9196 | 2.8794 | 3.8392 | 4.7990 | 5.7589 | 6.7187 | 7.6785 | 8.6383 | 1.3724 |
| 23 | 0.9597 | 1.9194 | 2.8791 | 3.8388 | 4.7985 | 5.7582 | 6.7179 | 7.6776 | 8.6373 | 1.3723 |
| 24 | 0.9596 | 1.9192 | 2.8788 | 3.8383 | 4.7979 | 5.7575 | 6.7171 | 7.6767 | 8.6363 | 1. 3722 |
| 25 | 0.9595 | 1.9189 | 2.8784 | 3.8379 | 4.7974 | 5.7568 | 6.7163 | 7.6758 | 8.6352 | 1. 3722 |
| 26 | 0.9594 | 1.9187 | 2.8781 | 3.8374 | 4.7958 | 5.7562 | 6.7155 | 7.6749 | 8.6342 | I. 3721 |
| 27 | 0.9592 | I.9185 | 2.8777 | 3.8370 | 4.7962 | 5.755 | 6.7147 | 7.6740 | 8.6332 | 1. 3720 |
| 28 | 0.9591 | 1.9183 | 2.8774 | 3.8365 | 4.7957 | 5.7548 | 6.7139 | 7.673I | 8.6322 | 1.3719 |
| 29 | 0.9590 | 1.9180 | 2.8771 | 3.8361 | 4.7951 | 5.7541 | 6.7131 | 7.6722 | 8.6312 | 1.3718 |
| 30 | 0.9589 | 1.9178 | 2.8767 | 3.8356 | 4.7945 | 5.7534 | 6.7124 | 7.67I3 | 8.6302 | 1.3718 |
| 3 I | 0.9588 | 1.9 | 2.8764 | 3.8352 | 4.7940 | 5.7528 | 6.7116 | 7.6704 | 8.6291 | r.3717 |
| 32 | 0.9587 | 1.9174 | 2.8760 | 3.8347 | 4.7934 | 5.7521 | 6.7108 | 7.6694 | 8.6281 | r. 3716 |
| 33 | 0.9586 | 1.9171 | 2.8757 | 3.8343 | 4.7928 | 5.7514 | 6.7100 | 7.6685 | 8.627 I | I. 3715 |
| 34 | 0.9585 | 1.9169 | 2.8754 | 3.8338 | 4.7923 | 5.7507 | 6.7092 | 7.6676 | 8.6261 | I. 3714 |
| 35 | 0.9583 | 1.9167 | 2.8750 | 3.8333 | 4.7917 | 5.7500 | 6.7084 | 7.6667 | 8.6250 | I. 3714 |
| 36 | 0.9582 | 1.9164 | 2.8747 | 3.8329 | 4.79 II | 5.7493 | 6.7076 | 7.6658 | 8.6240 | I. 3713 |
| 37 | 0.958 I | 1.9162 | 2.8743 | 3.8324 | 4.7905 | 5.7487 | 6.7068 | 7.6649 | 8.6230 | I. 3712 |
| 38 | 0.9580 | 1.9160 | 2.8740 | 3.8320 | 4.7900 | 5.7480 | 6.7060 | 7.6640 | 8.6219 | I. 3711 |
| 39 | 0.9579 | 1.9158 | 2.8736 | 3.8315 | 4.7894 | 5.7473 | 6.7052 | 7.6630 | 8.6209 | I. 3710 |
| 40 | $0.957^{8}$ | 1.9155 | 2.8733 | 3.8311 | 4.7888 | 5.7466 | 6.7044 | 7.6621 | 8.6199 | 1.3710 |
| 4I | 0.9577 | 1.9153 | 2.8730 | 3.8306 | 4.7883 | 5.7459 | 6. 7036 | 7.6612 | 8.6189 | 1.3709 |
| 42 | 0.9575 | 1.9151 | 2.8726 | 3.8301 | 4.7877 | 5.7452 | 6.7027 | 7.6603 | 8.6178 | I. 3708 |
| 43 | 0.9574 | I.9148 | 2.8723 | 3.8297 | 4.7871 | 5.7445 | 6.7019 | 7.6593 | 8.6168 | 1.3707 |
| 44 | 0.9573 | 1.9146 | 2.8719 | 3.8292 | 4.7865 | 5.7438 | 6.7011 | 7.6584 | 8.6157 | 1.3706 |
| 45 | 0.9572 | 1.9144 | 2.8716 | 3.8287 | 4.7859 | 5.7431 | 6.7003 | 7.6575 | 8.6147 | 1.3706 |
| 46 | 0.9571 | 1.9141 | 2.8712 | 3.8283 | 4.7854 | 5.7424 | 6.6995 | 7.6566 | 8.6136 | I. 3705 |
| 47 | 0.9570 | 1.9139 | 2.8709 | 3.8278 | 4.7848 | 5.7417 | 6.6987 | 7.6556 | 8.6126 | 1.3704 |
| 48 | 0.9568 | 1.9137 | 2.8705 | 3.8274 | 4.7842 | 5.7410 | 6.6979 | 7.6547 | 8.6116 | 1.3703 |
| 49 | 0.9567 | 1.9134 | 2.8702 | 3.8269 | 4.7836 | 5.7403 | 6.6971 | 7.6533 | 8.6105 | 1.3702 |
| 50 | 0.9566 | 1.9132 | 2.8698 | 3.8264 | 4.7830 | 5.7396 | 6.6962 | 7.6529 | 8.6095 | 1.3702 |
| 51 | 0.9565 | 1.9130 | 2.8695 | 3.8260 | $4 \cdot 7824$ | 5.7389 | 6.6954 | 7.6519 | 8.6084 | 1.3701 |
| 52 | 0.9564 | 1.9127 | 2.8691 | 3.8255 | 4.7819 | 5.7382 | 6.6946 | 7.6510 | 8.6073 | 1.3700 |
| 53 | 0.9563 | 1.9125 | 2.8688 | 3.8250 | 4.7813 | 5.7375 | 6.6938 | 7.6500 | 8.6063 | I. 3699 |
| 54 | 0.9561 | 1.9123 | 2.8684 | 3.8245 | 4.7807 | 5.7368 | $6 \cdot 6930$ | 7.6491 | 8.6052 | 1.3698 |
| 55 | 0.9560 | 1.9120 | 2.8681 | 3.8241 | 4.7801 | 5.7361 | 6.6921 | 7.6482 | $8.60+2$ | 1. 369 S |
| 56 | 0.9559 | 1.9118 | 2.8677 | 3.8236 | 4.7795 | 5.7354 | 6.6913 | 7.6472 | 8.6031 | 1. 3697 |
| 57 | 0.9558 | 1.9116 | 2.8674 | 3.8231 | 4.7789 | $5.73+7$ | 6.6905 | 7.6463 | 8.6021 | I. 3696 |
| 58 | 0.9557 | 1.9113 | 2.8670 | 3.8227 | 4.7783 | $5.73+0$ | 6.6897 | 7.6453 | S.6010 | 1. 3695 |
| 59 | 0.9556 | 1.9111 | 2.8667 | 3.8222 | 4.7778 | 5.7333 | 6.6889 | 7.6414 | 8.6000 | 1. 3694 |
| 60 | 0.9554 | 1.9109 | 2.8663 | 3.8217 | 4.7772 | 5.7326 | 6.6850 | 7.6435 | 8.5989 | 1. 3694 |


| $11^{\circ}$ | HEIGHTS. |  |  |  |  |  |  |  | 111 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | b |  |
| 0.1870 | 0.3741 | 0.5611 | 0.7482 |  | 1. 1222 | 1.3 | 1.4963 | 1. 6834 | 7 | ¢ |
| 0.1873 | 0.3746 | 0.5619 | 0.7492 | 0.9366 | 1.1239 | 1.3112 | 1.4985 | 1. 6858 | 0.2675 | ${ }^{\circ}$ |
| 0.1876 | 0. 3752 | 0.5627 | 0.7503 | 0.9379 | 1. 1255 | 1.3131 | 1. 5006 | 1. 6882 | 0.2679 | 02 |
| 0.1878 | 0. 3757 | 0.5635 | 0.7514 | 0.9392 | 1.1271 | I. 3149 | I. 5028 | 1. 6906 | 0.2683 | 3 |
| 0.1881 | 0.3762 | 0.5644 | 0.7525 | 0.9406 | 1.1287 | 1. 3168 | 1.5050 | 1.6931 | 0. 2687 | 04 |
| 0. 1884 | 0.3768 | 0. 5652 | 0.7536 | 0.9420 | 1.1303 | 1. 3187 | 1.5071 | 1.6955 | 0.2691 | 05 |
| 0. 1887 | 0.3773 | 0. 5660 | 0.7546 | 0.9433 | 1.1319 | 1. 3206 | 1.5093 | 1. 6979 | 0.2695 | 06 |
| 0. 1889 | 0.3778 | 0. 5668 | 0. 7557 | 0.9446 | I.1335 | I. 3224 | 1.5114 | 1.7003 | 0.2699 | 07 |
| 0.1892 | 0. 3784 | 0. 5676 | 0.7568 | 0.9460 | 1.135 | 1. 3243 | I.5135 | I. 7027 | 0.2703 | 08 |
| 0.1895 | 0.3789 | 0.5684 | 0.7578 | 0.9473 | 1.1368 | 1. 3262 | 1.5157 | 1.7051 | 0.2707 | 09 |
| 0.1897 | 0.3795 | 0.5692 | 0.7589 | 0.9487 | 1.1384 | 1.328I | 1.5178 | 1. 7076 | 0.27 II | o |
| 0.1900 | 0.3800 | 0. 5700 | c. 7600 | 0.9500 | 1.1400 | 1.3300 | 1.5200 | 1.7100 | 0.2715 | II |
| 0.1903 | 0.3805 | 0.5708 | 0.7611 | 0.9513 | 1.1416 | 1.3319 | 1.5222 | 1.7124 | 0.2719 | 12 |
| 0. 1905 | 0.3811 | 0.5716 | 0.7622 | 0.9527 | 1.1432 | 1. 3338 | 1. 5243 | 1. 7149 | 0.2723 | 13 |
| 0.1908 | 0.3816 | 0.5724 | 0.7632 | 0.9540 | 1.1448 | 1. 3357 | 1.5265 | 1.7173 | 0.2727 | 14 |
| 0.1911 | 0.3821 | 0. 5732 | 0. 7643 | 0.9554 | I. 1464 | 1. 3375 | 1. 5286 | 1.7197 | 0.2731 | 15 |
| -. 1913 | 0.3827 | 0. 5740 | 0.7654 | 0.9567 | 1.1480 | I. 3394 | 1. 5307 | 1.7221 | 0.2735 | 16 |
| -. 1916 | 0.3832 | 0. 5748 | 0. 7664 | 0.9580 | I. 1497 | I.3413 | I. 5329 | r. 7245 | 0.2739 | 17 |
| -.1919 | 0.3838 | 0.5756 | 0.7675 | 0.9594 | 1.1513 | I. 3432 | 1.5350 | 1. 7269 | 0.2743 | 18 |
| 0.192I | 0.3843 | 0.5764 | 0.7686 | 0.9607 | 1.1529 | I. 3450 | 1.5372 | 1. 7293 | 0.2747 | 19 |
| 0. 1924 | -. 3848 | 0.5773 | 0.7697 | 0.9621 | 1. 1545 | 1.3469 | 1.5394 | 1.7317 | 0.2751 | 20 |
| 0.1927 | 0. 3854 | 0.5781 | 0.7707 | 0.9634 | 1.156 | 1. 3488 | 1. 5415 | 1.7341 | 0.2755 | 21 |
| 0.1930 | 0.3859 | 0.5789 | 0.7718 | 0.9648 | 1.1577 | 1.3507 | 1.5436 | 1.7366 | 0.2759 | 22 |
| 0. 1932 | 0.3864 | 0.5797 | 0.7729 | 0.9661 | 1.1593 | 1.3525 | 1.5458 | 1. 7390 | 0.2763 | 23 |
| 0. 1935 | 0.3870 | 0. 5805 | 0.7740 | 0.9674 | 1.1609 | I. 3544 | I. 5479 | 1.7414 | 0.2767 | 24 |
| 0.1938 | 0.3875 | 0.5813 | 0.7750 | 0.9688 | 1.1625 | I.3563 | 1.5500 | 1. 7438 | 0.2771 | 25 |
| -. 1940 | 0.3880 | 0.5821 | 0.776I | 0.9701 | 1.164I | I.3581 | 1. 5522 | I. 7462 | 0.2775 | 26 |
| 0. 1943 | 0. 3886 | 0.5829 | 0.7772 | 0.9714 | 1.1657 | I.3600 | 1. 5543 | I. 7486 | 0.2779 | 27 |
| 0. 1946 | 0.3891 | 0.5837 | 0.7782 | 0.9728 | 1.1674 | 1.3619 | 1.5565 | 1.7510 | 0.2783 | 28 |
| -. 1948 | 0.3896 | 0. 5845 | 0. 7793 | 0.974 | 1.1689 | 1. 3637 | 1.5586 | 1. 7534 | 0.2787 | 29 |
| 0.1951 | 0.3902 | 0.5853 | 0.78c4 | 0.9755 | 1.1705 | I. 3656 | 1. 5607 | 1.7558 | 0.2791 | 30 |
| 0. 195 | 0.3907 | 0.5861 | 0. |  | 1. | 1. 3675 | 1.5629 |  | 95 |  |
| 0.1956 | 0.3913 | 0.5869 | 0.7825 | 0.9781 | 1.1738 | 1. 3694 | I. 5650 | 1. 7607 | 0.2799 | 32 |
| 0. 1959 | -0.3918 | 0. 5877 | 0. 7836 | 0.9795 | 1. 1753 | 1.3712 | 1.5671 | 1.7630 | 0.28 c 3 | 33 |
| c. 1962 | 0.3923 | 0. 5885 | c. 7846 | 0.9808 | 1.1770 | 1.3731 | 1.5693 | 1. 7654 | 0.2807 | 34 |
| 0. 1964 | 0.3929 | 0.5893 | 0.7857 | 0.9821 | 1. 1786 | 1.3750 | 1.5714 | 1.7679 | 0.28 II | 35 |
| -. 1957 | -0.3934 | 0.5901 | 0. 7868 | 0.9835 | 1.1802 | 1.3769 | I. 5736 | 1. 7703 | 0.2815 | 36 |
| 0. 1970 | 0.3939 | 0.5909 | 0. 7878 | 0.9848 | 1.1818 | 1.3787 | 1.5757 | 1.7727 | 0.2819 | 37 |
| 0.1972 | 0.3945 | 0. 5917 | 0.7889 | 0.9861 | I. 1834 | 1. 3806 | I. 5778 | 1.775I | 0.2823 |  |
| 0. 1975 | 0.3950 | 0.5925 | 0.7900 | 0.9875. | I. 1850 | 1. 3825 | I. 5800 | I. 7775 | 0.2827 | 39 |
| 0. 1978 | 0.3955 | 0. 5933 | 0.7910 | 0.9888 | I. 1866 | I. 3843 | 1.582I | 1.7798 | 0.283 I | 40 |
| 0. 1980 | 0.3961 | 0.5941 | 0.7921 | 0.9901 | 1. 1882 | 1. 3862 | 1. 5842 | 1. 7823 | 0.2835 | 4 |
| 0. 1983 | 0.3966 | 0.5949 | 0.7932 | 0.991 | 1.1898 | 1.388I | 1. 5864 | I. 7847 | 0.2839 | 42 |
| 0. 1986 | 0.3971 | 0. 5957 | 0. 7942 | 0.9928 | I. 1914 | 1.3899 | 1. 5885 | 1. 7871 | 0.2843 | 43 |
| -. 1988 | 0.3977 | 0. 5965 | 0. 7953 | 0.9941 | I. 1930 | I. 3918 | I. 5906 | 1.7895 | 0.2847 | 44 |
| o. 1991 | 0.3982 | 0. 5973 | 0. 7964 | 0.9955 | I. 1946 | I. 3937 | I. 5928 | I.7919 | 0.2851 | 45 |
| o. 1994 | 0.3987 | 0.5981 | 0.7974 | 0.9968 | I. 1962 | I. 3955 | I. 5949 | I. 7942 | 0.2855 | 46 |
| 0. 1996 | 0. 3993 | 0.5989 | 0.7985 | 0.9981 | 1. 1978 | I. 3974 | 1.5970 | I. 7966 | 0. 2859 | 47 |
| 0. 1999 | 0. 3998 | 0.5997 | 0.7996 | 0.9994 | I. 1993 | I. 3992 | 1.5991 | 1.7990 | 0.2863 | 48 |
| 0.2002 | 0.4003 | 0.6005 | $0.8 \mathrm{co6}$ | 1.0008 | 1.2010 | 1.4011 | 1.6013 | 1.8014 | 0.2867 | 49 |
| 0.2004 | 0.4009 | 0.6013 | 0.8017 | 1.002 1 | 1.2026 | 1.4030 | 1.6034 | 1.8038 | 0.2871 | 50 |
| 0.2007 | 0.4014 | 0.6021 | 0. | 1.0034 | 1.2041 | 1. 4048 | 1. 6055 | 1.8062 | 0. 2875 | 51 |
| 0.2010 | 0.4019 | 0.6029 | 0.8038 | 1.0048 | I. 2057 | 1. 4067 | 1.6077 | r. 8086 | 0. 2879 | 52 |
| 0.2012 | 0.4024 | -.6037 | 0. 8049 | 1.006r | 1. 2073 | 1.4085 | 1.6098 | I.8110 | 0.2883 | 53 |
| 0.2015 | 0.4030 | -.6045 | 0.8060 | 1.0075 | 1.2089 | 1.4104 | 1.6119 | 1.8134 | 0.2887 | 54 |
| 0.2018 | 0.4035 | 0.6053 | 0.8070 | 1.0088 | 1.2105 | I.4123 | 1.6141 | 1.8158 | 0.2891 | 55 |
| 0.2020 | 0.4040 | 0.6061 | 0.8081 | 1.0101 | 1.2121 | 1.4141 | 1.6162 | 1.8182 | 0.2895 | 56 |
| 0.2023 0.2026 | 0.4046 0.4051 | 0.6069 0.6077 | 0.8092 0.8102 | I. l II14 | 1.2137 | I. 4160 | 1. 6183 | 1. 8206 | 0.2899 |  |
| 0.2026 0.2028 | 0.4051 0.4056 | 0.6077 0.6085 | 0.8102 0.8113 | I. 12128 I.OI4I | 1.2153 1.2169 | I. 4179 | 1.6204 I. 6226 | 1.8230 1.8254 | 0.2903 0.2507 | 58 59 |
| 0.2028 0.2031 | 0.4056 0.4062 | 0.6085 0.6092 | 0.8113 0.8123 | I.0141 I.0154 | I. 21169 1.2185 | 1.4197 I.4216 | 1.6226 I. 6247 | 1.8254 1.8278 | 0.2507 0.2911 | 59 |


| 112 |  | DISTANCES. |  |  |  |  |  |  | $12^{\circ}$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | a |
| $\infty$ | 0.9554 |  | 2.8663 | 3.8217 | 4.7772 | 5.7326 | 6.6880 | 7.6435 | .8.5989 | 94 |
| OI | 0.9553 | 1.91c6 | 2.8659 | 3.8213 | 4.7766 | 5.7319 | 6.6872 | 7.6425 | 8.5978 | 1. 3693 |
| 02 | $0.955^{2}$ | 1.9104 | 2.8656 | 3.8208 | 4.7760 | 5.7312 | $6.686_{4}$ | 7.6416 | 8.5958 | 1.3693 |
| 03 | 0.9551 | 1.9102 | 2.8652 | 3.8203 | 4.7754 | 5.7305 | 6.6855 | 7.6406 | 8. 5957 | I. 3692 |
| 04 | 0.9550 | 1. 9099 | 2.8649 | 3.8198 | 4.7748 | 5.7297 | 6.6847 | 7.6397 | 8.5946 | 1.3691 |
| 05 | $0.954^{8}$ | 1.9097 | 2.8645 | 3.8194 | 4.7742 | 5.7290 | 6.6839 | 7.6387 | 8.5936 | I. 3690 |
| 06 | 0.9547 | I. 9094 | 2.8642 | 3.8189 | 4.7736 | 5.7283 | 6.6830 | 7.6378 | 8.5925 | 1. 3689 |
| 07 | 0.9546 | I. 9092 | 2.8638 | 3.8184 | 4.7730 | 5.7276 | 6.6822 | 7.6368 | 8.5914 | I. 3688 |
| 08 | 0.9545 | 1.9090 | 2.8634 | 3.8179 | 4.7724 | 5.7269 | 6.6814 | 7.6359 | 8.5903 | I. 3687 |
| 09 | 0.9544 | 1.9087 | 2.8631 | 3.8175 | 4.7718 | 5. 7262 | 6.6805 | 7.6349 | 8.5893 | I. 3687 |
| 10 | 0.9542 | I.gos5 | 2.8627 | 3.8170 | 4.7712 | 5.7255 | 6.6797 | 7.6340 | 8.5882 | 1.3686 |
| II | 0.9541 | 1.9082 | 2.8 |  | 4.7706 | 5.7247 | 6.6789 | 7.6330 |  |  |
| 12 | 0.9540 | 1.9030 | 2.8620 | 3.8160 | 4.7700 | 5.7240 | 6.6780 | 7.6320 | 8.5860 | 1. 3684 |
| 13 | 0.9539 | 1.9ग78 | 2.8616 | 3.8155 | 4.7694 | 5.7233 | $6.677^{2}$ | 7.6311 | 8.5849 | I. 3683 |
| 14 | 0.9538 | I. $¢ 075$ | 2.8613 | 3.8150 | 4.7688 | 5.7226 | 6.6763 | 7.6301 | 8.5839 | I. 3682 |
| 15 | 0.9536 | I. 9073 | 2.8609 | 3.8146 | 4.7682 | 5.7219 | 6.6755 | 7.6291 | 8.5828 | I. 368 I |
| 16 | 0.9535 | 1.9フ70 | 2.8605 | 3.8141 | 4.7676 | 5.7211 | 6.6747 | 7.6282 | 8.5817 | 1.368I |
| 17 | 0.9534 | 1.9068 | 2.8602 | 3.8136 | 4.7670 | 5.7204 | 6.673 | 7.6272 | 8.5806 | I. 3680 |
| 18 | 0.9533 | I. 9 J66 | 2.8598 | 3.8131 | 4.7664 | 5.7197 | 6.6730 | 7.6262 | 8.5795 | 1.3679 |
| 19 | 0.9532 | $\mathrm{I}_{1.9053}$ | 2.8595 | 3.8126 | 4.7658 | 5.7190 | 6.6721 | 7.6253 | 8.5784 | 1.3678 |
| 20 | 0.9530 | I. 9061 | 2.8591 | 3.8122 | 4.7652 | 5.7182 | 6.6713 | 7.6243 | 8.5774 | 1.3677 |
| 21 | 0.9529 |  | 2.8588 |  | 4.7646 | 5.7175 |  | 7.6233 | 8.5763 | I. 3676 |
| 22 | 0.9528 | 1.9356 | 2.8584 | 3.8112 | 4.7640 | 5.7168 | 6.6696 | 7.6224 | 8.5752 | I. 3675 |
| 23 | 0.9527 | I. 9553 | 2.8580 | 3.8107 | 4.7634 | 5.7160 | 6.6687 | 7.6214 | 8.5741 | I. 3674 |
| 24 | 0.9526 | 1.9531 | 2.8577 | 3.8102 | 4.7628 | 5.7153 | 6.6679 | 7.6204 | 8.5730 | 1. 3673 |
| 25 | 0.9524 | 1.9049 | 2.8573 | 3.8097 | 4.7622 | 5.7146 | 6.6670 | 7.6194 | 8.5719 | I. 3672 |
| 26 | 0.9523 | I. 9046 | 2.8569 | 3.8032 | 4.7615 | 5.7138 | 6.6662 | 7.6185 | 8.5708 | I. 3672 |
| 27 | 0.9522 | 1.9044 | 2.8566 | 3.8087 | 4.7609 | 5.7131 | 6.6653 | 7.6175 | 8.5697 | I. 367 I |
| 28 | 0.9521 | 1.9041 | 2.8562 | 3.8083 | 4.7603 | 5.7124 | 6.6644 | 7.6165 | 8.5686 | I. 3670 |
| 29 | 0.9519 | 1.9039 | 2.8558 | 3.8078 | 4.7597 | 5.7117 | 6.6636 | 7.6155 | 8.5675 | I. 3669 |
| 30 | 0.9518 | 1.9036 | 2.8555 | 3.8073 | $4 \cdot 7591$ | 5.7109 | 6.6627 | 7.6146 | 8.5664 | I. 3668 |
| 31 | 0.9517 | 1.9034 | 2.8551 | 3.8058 | 4.7585 | 5.7102 | $6.6619$ | 7.6136 | 8.5653 | I. 3667 |
| 32 | 0.9516 | I.g031 | 2.8547 | $3.805_{3}$ | 4.7579 | 5.7094 | $6.6610$ | 7.6126 | $8.564^{2}$ | I. 3667 |
| 33 | 0.9514 | 1.9029 | 2.8543 | 3.8058 | 4.7572 | 5.7087 | 6.6601 | 7.6116 | 8.5630 | I. 3666 |
| 34 | 0.9513 | 1.9227 | 2.8540 | 3.8053 | 4.7566 | 5.7030 | 6.6593 | 7.6106 | 8.5619 | I. 3665 |
| 35 | 0.9512 | 1.9024 | 2.8536 | 3.8048 | 4.7560 | 5.7072 | 6.6584 | 7.6096 | 8.5608 | I. 3664 |
| 36 | 0.9511 | 1.9022 | 2.8532 | 3.8043 | 4.7554 | 5.7055 | 6.6575 | 7.6086 | 8.5597 | I. 3663 |
| 37 | 0.9510 | 1.9019 | 2.8529 | 3.8038 | 4.7548 | 5.7057 | 6.6567 | 7.6076 | 8.5586 | I. 3662 |
| 38 | 0.9508 | 1.9017 | 2.8525 | 3.8033 | 4.7542 | 5.7050 | 6.6558 | 7.6066 | 8.5575 | I. 3661 |
| 39 | 0.9507 | 1.9014 | 2.8521 | 3.8028 | 4.7535 | $5.70+2$ | 6.6550 | 7.6057 | 8.5564 | I. 3660 |
| 40 | 0.9506 | 1.9312 | 2.8518 | 3.8023 | 4.7529 | 5.7035 | $6.654^{1}$ | 7.6047 | 8.5553 | 1. 3660 |
| 4 I | 0.9505 | 1.9009 | 2.8514 | 3.8018 | 4.7523 | 5.7028 | 6.6532 | 7.6037 | 8.554I | 1. 3659 |
| 42 | 0.9503 | I. 9007 | 2.8510 | 3.8013 | 4.7517 | 5.7020 | 6.6523 | 7.6027 | 8.5530 | 1.3658 |
| 43 | 0.9502 | 1.9004 | 2.8506 | 3.8008 | 4.7510 | 5.7013 | 6.6515 | 7.0017 | 8.5519 | I. 3657 |
| 44 | 0.9501 | 1.9002 | 2.8503 | 3.8003 | 4.7504 | 5.7005 | 6.6506 | 7.6007 | 8.5508 | I. 3656 |
| 45 | 0.9500 | 1.8999 | 2.8499 | 3.7998 | 4.7495 | 5.6998 | 6.6497 | 7.5997 | 8.5496 | I. 3655 |
| 46 | 0.9498 | 1.8997 | 2.8495 | 3.7993 | 4.7492 | 5.6990 | 6.6488 | 7.5987 | 8.5485 | I. 3654 |
| 47 | 0.9497 | 1.8934 | 2.8491 | 3.7988 | 4.7485 | 5.6983 | 6.6480 | 7.5977 | 8.5474 | 1. 3653 |
| 48 | 0.9495 | 1. 8972 | 2.8488 | 3.7983 | 4.7479 | 5.6975 | 6.647 I | 7.5967 | 8.5463 | I. 3652 |
| 49 | 0.9495 | I. 8989 | 2.8484 | 3.7978 | 4.7473 | 5.6968 | 6.6462 | 7.5957 | 8.5451 | I. 3651 |
| 50 | 0.9493 | 1. 8987 | 2.8480 | 3.7973 | 4.7467 | 5.6960 | 6.6453 | 7.5947 | 8.5440 | I. 3651 |
| 51 | 0.9492 | 1. 8984 | 2.8476 | 3.7968 | 4.7460 | 5.6952 | 6.6444 | 7.5937 | 8.5429 | r. 3650 |
| 52 | 0.9491 | 1.8982 | 2.8472 | 3.7963 | 4. 7454 | 5.6945 | 6.6436 | 7.5926 | 8. 5417 | 1. 3649 |
| 53 | 0.9490 | 1. 8979 | 2.8469 | 3.7958 | 4.7448 | 5.6937 | 6.6427 | 7.5916 | 8. 5406 | I. 3648 |
| 54 | 0.9488 | 1. 8977 | 2.8465 | 3.7953 | 4.7441 | 5.6930 | 6.6418 | 7.5906 | 8. 5395 | I. 3647 |
| 55 | 0.9487 | 1. 8974 | 2.846 I | 3.7948 | 4.7435 | 5.6922 | 6.6409 | $7 \cdot 5896$ | 8.5383 | I. 3646 |
| 56 | 0.9486 | 1. 8971 | 2.8457 | 3.7943 | $4.7+29$ | 5.6914 | 6.6400 | 7.5886 | 8.5372 | I. 3645 |
| 57 | 0.9484 | I. 8969 | 2.8453 | 3.7938 | 4.7422 | 5.6907 | 6.6391 | 7.5876 | 8.5360 | I. 3644 |
| 58 | 0.9483 | 1. 8966 | 2.8450 | 3.7933 | 4.7416 | 5.6899 | 6.6382 | 7.5866 | 8.5349 | 1. 3643 |
| 59 | 0.9482 | 1. 8964 | 2.8446 | 3.7928 | 4.7410 | 5.6892 | 6.6374 | 7.5856 | 8.5338 | I. 3642 |
| 60 | 0.948 I | 1.8961 | 2.8442 | 3.7923 | 4.7403 | 5.6884 | 6.6365 | 7.5845 | 8.5326 | 1.3641 |


| $12^{\circ}$ | HEIGHTS. |  |  |  |  |  |  |  |  | 11 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | b |  |
| 0.2031 | 0.4062 | 0.60 |  |  | 1.2185 | 1. 4216 | I. | 1. 8278 | II | -0 |
| 0.2033 | 0.4067 | 0.6100 | 0.8134 | I.0167 | 1.2201 | 1. 4234 | 1. 6268 | 1.8302 | 0.2915 | OI |
| 0.2036 | 0.4072 | 0.6108 | 0.8144 | 1.018I | 1.2217 | 1. 4253 | I. 6289 | 1. 8325 | 0.2919 | 02 |
| 0.2039 | 0.4078 | 0.6116 | 0.8155 | 1.0194 | 1.2233 | 1. 4272 | 1.6310 | 1. 8349 | 0.2923 | 03 |
| 0.2041 | $0.40{ }_{3}$ | 0.6124 | 0.8166 | 1.0207 | 1. 2248 | I. 4290 | I. 6331 | 1. 8373 | 0.2927 | 04 |
| 0.2044 | 0.4088 | 0.6132 | 0.8176 | 1.0221 | 1. 2264 | 1.4309 | 1. 6353 | 1. 8397 | 0.2931 | 05 |
| 0.2047 | 0.4093 | 0.6140 | 0.8187 | 1.0234 | 1.2280 | I. 4327 | 1. 6374 | I. 8420 | 0.2935 | o6 |
| 0.2049 | 0.4099 | 0.6148 | 0.8198 | 1.0247 | 1.2296 | 1. 4346 | 1. 6395 | 1. 8444 | 0.2939 | 07 |
| c. 2052 | 0.4104 | 0.6156 | 0.8208 | 1.0260 | 1.2312 | I. 4364 | 1. 6416 | 1. 8468 | 0.2943 | o8 |
| 0.2055 | 0.4109 | 0.6164 | 0.8219 | 1.0273 | 1. 2328 | 1.4383 | 1. 6438 | 1. 8492 | 0.2947 | $\bigcirc 9$ |
| 0.2057 | 0.4115 | 0.6172 | 0.8229 | 1.0287 | 1.2344 | 1.4401 | 1.6459 | 1.8516 | 0.2951 | O |
| 0.2060 | 0.4120 | 0.6 |  | 1.0300 | 1.2360 | 1.4420 | I. 6480 |  | 0.2955 |  |
| 0.2053 | c. 4125 | 0.6188 | 0.8250 | 1.0313 | I. 2376 | I. 4438 | I. 6501 | 1. 8564 | 0.2959 | 12 |
| 0.2065 | 0.4131 | 0.6196 | 0.8261 | 1.0326 | I. 2392 | I. 4457 | I. 6522 | 1. 8588 | 0.2962 | 13 |
| 0.2068 | 0.4136 | 0.6204 | 0.8272 | I. 3340 | I. 2408 | I. 4475 | I. 6543 | I.86ıI | 0.2966 | 14 |
| 0.2071 | 0.4141 | 0.6212 | 0.8282 | 1.0353 | I. 2424 | I. 4494 | I. 6565 | I. 8635 | 0.2970 | 15 |
| 0.2073 | 0.4146 | 0.6220 | 0.8293 | I. 3636 | I. 2439 | 1.4512 | I. 6586 | 1. 8659 | 0.2974 | 16 |
| 0.2076 | 0.4152 | 0.6227 | 0.8303 | I. 0379 | I. 2455 | I.4531 | 1.6607 | 1. 8682 | 0.2978 | 17 |
| 0.2078 | 0.4157 | 0.6235 | o. 8314 | 1.0392 | I. 2471 | I. 4549 | I. 6628 | 1. 8706 | 0.2582 | 18 |
| 0.2081 | 0.4162 | 0.6243 | 0.8324 | 1.0406 | I. 2487 | I. 4568 | I. 6649 | I. 8730 | 0.2986 | 19 |
| 0.2084 | 0.4168 | 0.6251 | 0.8335 | 1.0419 | I. 2503 | 1.4587 | I. 6670 | I. 8754 | 0. 2990 | - |
| 0.2086 | 0.4173 | 0.625 | 0.8346 | 1.0432 | 1.2518 | 1.4605 | 1.6691 | 1. 8778 | 0.2994 | 21 |
| 0.2089 | 0.4178 | 0.6267 | 0.8356 | 1.0445 | 1. 2534 | 1. 4623 | 1. 6712 | I. 8801 | 0.2598 | 22 |
| 0.2092 | 0.4183 | 0.6275 | 0.8367 | I. 0458 | I. 2550 | 1. 4642 | 1. 6734 | 1. 8825 | 0.3002 | 23 |
| 0.2094 | 0.4189 | 0.6283 | 0.8377 | 1.0472 | 1.2566 | 1. 4660 | 1. 6755 | I. 8849 | 0.3006 | 24 |
| 0.2097 | 0.4194 | 0.6291 | 0.8388 | I. 0485 | I. 2582 | I. 4679 | I. 6776 | 1. 8873 | 0.3010 | 25 |
| 0.2100 | 0.4199 | 0.6299 | 0.8398 | 1.0498 | I. 2598 | 1. 4697 | 1. 6797 | I. 8896 | 0.3014 | 26 |
| -. | 0.4204 | 0.6307 | 0.8409 | 1.0511 | I. 2613 | 1.4715 | I. 6818 | 1. 8920 | 0.3018 | 27 |
| 0.2105 | 0.4210 | 0.6315 | 0.8420 | 1.0524 | I. 2629 | I. 4734 | 1. 6839 | 1. 8944 | 0.3022 | 28 |
| 0.2107 | 0.4215 | 0.6322 | 0. 8430 | I. 0537 | 1. 2645 | 1.4752 | 1.6860 | I. 8967 | 0.3026 | 29 |
| 0.2110 | 0.4220 | 0.6330 | 0.8440 | 1.0551 | 1. 2661 | 1.477 | 1.688ı | 1.8991 | 0.3030 | $3^{\circ}$ |
| 13 | 0.4226 | 0.6338 | 0.8451 | 1.0564 | I. 2677 | 1. 4790 | 1.6,02 | 1.9015 | 0.3034 | 3 I |
| 0.2115 | 0.4231 | 0.6346 | 0.8462 | I. 0577 | 1.2692 | 1. 4808 | т. $¢ 923$ | I. 9039 | 0.3038 | $3^{2}$ |
| 0.2118 | 0.4236 | 0.6354 | 0.8472 | I.0550 | 1.2708 | 1. 4826 | I. 6944 | 1. 9062 | 0.3042 | 33 |
| 0.2121 | 0.4241 | 0.6362 | 0. 8483 | I.o¢03 | 1.2724 | 1. 4845 | 1. 6965 | I. 9086 | 0.3046 | 34 |
| 0.2123 | 0.4247 | 0.6370 | 0.8493 | 1.0616 | 1.2740 | I. 4863 | I. 6986 | I.9110 | 0. 3050 | 35 |
| 0.212 | 0.4252 | 0.6378 | 0.8504 | 1.0630 | I. 2755 | 1.488I | 1. 7007 | 1.9133 | c. 3054 | 36 |
| 0.2129 | 0.4257 | 0.6386 | 0.8514 | 1.0643 | 1.2771 | I. 4900 | 1. 7028 | 1.9157 | 0. 3058 | 37 |
| 0.2 | 0.4262 | 0.6394 | 0.8525 | 1.0656 | 1. 2787 | 1. 4918 | 1. 7049 | I.9181 | 0.3062 | $3^{8}$ |
| 0.2134 | 0.4268 | 0.6401 | 0.8535 | 1.0669 | 1.2803 | I. 4937 | I. 7070 | 1.9204 | 0.3066 | 39 |
| 0.2136 | 0.4273 | 0.6409 | 0.8546 | 1.0682 | I. 2818 | 1. 4955 | 1. 7091 | 1. 9228 | 0. 3070 | $4^{\circ}$ |
| 0.2139 | 0.4278 | 0.6417 | 0.8556 | 1.0695 | 1.2834 | I. 4973 | 1.7112 | 1.925I | 0.3074 | $4{ }^{1}$ |
| 0.2142 | 0.4283 | 0.6425 | 0.8567 | 1.0708 | 1.2850 | 1. 4992 | 1. 7133 | I. 9275 | 0.3078 | 42 |
| 0.2144 | 0.4289 | 0.6433 | 0.8577 | 1.0721 | 1. 2866 | 1.5010 | 1.7154 | I. 9299 | 0.3082 | 43 |
| 0.2147 | 0.4294 | 0.644 I | 0.8588 | I. 073 | 1.288I | 1.5028 | 1. 7175 | 1.9322 | 0.3086 | 44 |
| 0.2150 | 0.4299 | 0.6449 | 0. 8598 | 1.0748 | 1. 2897 | I. 5047 | 1. 7196 | 1. 9346 | 0.3090 | 45 |
| 0.2152 | 0.4304 | 0.6457 | 0.8609 | 1.0761 | I. 2913 | I. 5065 | 1. 7217 | 1.9370 | 0.3094 | 46 |
| 0.2155 | 0.4310 | 0.6464 | 0.8619 | 1.0774 | 1.2929 | I. 5084 | 1. 7238 | 1. 9393 | 0.3098 | 47 |
| 0.2157 | 0.4315 | 0.6472 | 0.8630 | 1.0787 | I. 2944 | 1.5102 | 1.7259 | 1.9417 | 0.3102 | 4 |
| 0.2160 | 0.4320 | 0.6480 | 0.8640 | 1.0800 | 1.2960 | 1.5120 | 1. 7280 | 1. 9440 | 0.3106 | 49 |
| 0.2163 | 0.4325 | 0.6488 | 0.8651 | 1.0813 | I. 2976 | I.5138 | 1.7301 | 1.9464 | 0.3110 | 50 |
| 0.2165 | 0.433 I | 0.6496 | 0.866r | 1.0826 | 1.2992 | I. 5157 | 1.7322 | 1. 9487 | 0.3114 | 51 |
| 0.2168 | 0.4336 | 0.6504 | 0.8671 | 1.083 | 1.3007 | 1.5175 | 1. 7343 | 1.9511 | 0.3118 | 52 |
| 0.2170 | 0.4341 | 0.6511 | 0.8682 | 1.085 | 1. 3023 | I. 5193 | 1.7364 | ז. 9534 | 0.3121 | 53 |
| 0.2173 | 0.4346 | 0.6519 | 0.8692 | 1.086 | 1. 3039 | 1.5212 | I. 7385 | 1. 9558 | 0.3125 | 54 |
| 0.2176 | 0.4351 | 0.6527 | 0.8703 | 1.0879 | I. 3054 | I. 5230 | I. 7406 | 1.958I | 0.3129 | 55 |
| 0.2178 | 0. 4357 | 0.6535 | 0.8713 | 1.0892 | 1. 3070 | 1. 5248 | 1.7427 | I. 9605 | 0.3133 | 56 |
| 0.2181 | 0.4362 | 0.6543 | 0.8724 | 1.0505 | 1.3186 | 1. 5267 | 1. 7448 | I. 9629 | 0.3137 | 57 |
| 0.2184 | 0.4367 | 0.655 I | 0.8734 | 1.0918 | I.3101 | I. 5285 | 1. 7468 | 1. 9652 | 0.3141 | 58 |
| 0.2186 | 0.4372 | 0.6559 | 0.8745 | 1.0931 | 1.3117 | 1.5303 | I. 7489 | 1.9676 | 0.3145 | 59 |
| 0.2189 | 0.4378 | 0.65 | 0.8755 | 1.0944 | 1.3133 | 1.5322 | 1.7510 | 1.9699 | 0.3149 | 60 |


| 114 |  | DISTANCES. |  |  |  |  |  |  |  | $13^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | a |
| 00 | 0.948 I | 1.896I | 2.84 | 3.7923 | 4.7403 | 5.6884 | 6.6365 | 7.5845 | 8.5320 |  |
| cr | 0.9479 | 1. 8959 | 2.8438 | 3.7918 | 4.7397 | 5.6876 | 6.6356 | 7.5835 | 8.5315 | I. 3040 |
| 02 | 0.9478 | т. 8956 | 2.8434 | 3.7912 | 4.7391 | 5.6869 | 6.6347 | 7.5825 | 8.5303 | I. 3639 |
| 03 | 0.9477 | 1. 8954 | 2.8431 | 3.7907 | 4.7384 | 5.6861 | 6.6338 | 7.5815 | 8.5292 | I. 3638 |
| 04 | 0.9476 | 1.8951 | 2.8427 | 3.7902 | 4.7378 | 5.6853 | 6.6329 | 7.5804 | 8.5280 | I. 3637 |
| 05 | 0.9474 | 1. 8949 | 2.8423 | 3.7897 | 4.7371 | 5.6846 | 6.6320 | 7.5794 | 8.5269 | I. 3636 |
| 06 | 0.9473 | 1. 8946 | 2.8419 | 3.7892 | 4.7365 | 5.6838 | 6.6311 | 7.5784 | 8.5257 | I. 3635 |
| 07 | 0.9472 | 1. 8943 | 2.8415 | 3.7887 | 4.7359 | 5.6830 | 6.6302 | 7.5774 | 8.5245 | I. 3634 |
| 08 | 0.9470 | I. 8941 | 2.841 I | 3.7882 | 4.7352 | 5.6823 | 6.6293 | 7.5763 | 8.5234 | 1. 3634 |
| 99 | 0.9469 | I. 8938 | 2.8407 | 3.7877 | 4.7346 | 5.6815 | 6.6284 | 7.5753 | 8. 5222 | I. 3633 |
| 10 | 0.9468 | 1.8936 | 2.8404 | 3.787 I | 4.7339 | 5.6807 | 6.6275 | 7.5743 | 8.5211 | I. 3632 |
| II | 0.9467 | 1. 8933 | 2.8400 | 3.7866 | 4.7333 | 5.6799 | 6.6266 | 7.5733 | 8.5199 | I. 3631 |
| 12 | 0.9465 | 1.8931 | 2.8396 | 3.7861 | 4.7326 | 5.6792 | 6.6257 | 7.5722 | 8.5188 | I. 3630 |
| 13 | 0.9464 | 1. 8928 | 2.8392 | 3.7856 | 4.7320 | 5.6784 | 6.6248 | 7.5712 | 8.5176 | I. 3629 |
| 14 | 0.9463 | I. 8925 | 2.8388 | 3.7851 | 4.7313 | 5.6776 | 6.6239 | 7.5702 | 8.5164 | I. 3628 |
| 15 | 0.946 r | 1. 8923 | 2.8384 | 3.7846 | 4.7307 | 5.6708 | 6.6230 | 7.5691 | 8.5153 | 1. 3627 |
| 16 | 0.9460 | I. 8920 | 2.8380 | 3.7840 | 4.7300 | 5.6761 | 6.6221 | 7.568I | ${ }^{8.5141}$ | I. 3626 |
| 17 | 0.9459 | I. 8918 | 2.8376 | 3.7835 | 4.7294 | 5.6753 | 6.6212 | 7.5670 | 8.5129 | I. 3625 |
| 18 | 0.9458 | I. 8915 | 2.8373 | 3.7830 | 4.7288 | 5.6745 | 6.6203 | 7.5660 | 8.5118 | I. 3624 |
| 19 | 0.9456 | I. 8912 | 2.8369 | 3.7825 | 4.728 | 5.6737 | 6.6193 | 7.5650 | 8.5106 | I. 3623 |
| 20 | 0.9455 | 1.8910 | 2.8365 | 3.7820 | 4.7275 | 5.6729 | 6.6184 | 7.5639 | 8.5094 | 1. 3622 |
| 21 | 0.9 | 1.8 | 2.8361 | 3.7814 | 4.7268 | 5.6722 |  | 7.5629 | 8.5082 | I. 3621 |
| 22 | 0.9452 | 1. 8905 | 2.8357 | 3.7809 | 4.726 | 5.6714 | 6.6166 | 7.5618 | 8.5071 | I. 3620 |
| 23 | 0.9451 | 1. 8902 | 2.8353 | 3.7804 | 4.7255 | 5.670 | 6.6157 | 7.5608 | 8.5059 | I.3619 |
| 24 | 0.9450 | 1. 8899 | 2.8349 | 3.7799 | 4.7248 | 5.669 | 6.6148 | 7.5597 | 8.5047 | I. 3618 |
| 25 | 0.9448 | 1. 8897 | 2.8345 | 3.7793 | 4.7242 | 5.669 | 6.6139 | 7.5587 | 8.5035 | I. 3618 |
| 26 | 0.9447 | 1. 8894 | 2.8341 | 3.7788 | 4.7235 | 5.6682 | 6.6129 | 7.5576 | 8.5023 | 1. 3617 |
| 27 | 0.9446 | 1.8891 | 2.8337 | 3.7783 | 4.7229 | 5.6674 | 6.6120 | 7.5566 | 8.5012 | 1. 3616 |
| 28 | 0.9444 | 1. 8889 | 2.8333 | 3.7778 | 4.7222 | 5.6667 | 6.61 | 7.5555 | 8.5000 | 1. 3615 |
| 29 | 0.9443 | 1. 8886 | 2.8329 | 3.7772 | 4.7216 | 5.6659 | 6.6 | 7.5545 | 8.4988 | I. 3614 |
| 30 | 0.9442 | 1. 8884 | 2.8325 | 3.7767 | 4.7209 | 5.6651 | 6.6093 | 7.5534 | 8.4976 | 1.3613 |
| 31 | 0.9440 | 1.888I | 2.8321 | 3.7762 | 4.7202 | 5.6643 | 6.6083 | $7 \cdot 5524$ | 8.4964 | 1.3612 |
| 32 | 0.9439 | 1.8878 | 2.8317 | 3.7757 | 4.7196 | 5.6635 | 6.6074 | 7.5513 | $8.495^{2}$ | 1.3611 |
| 33 | 0.9438 | I. 8876 | 2.8313 | 3.775I | 4.7189 | 5.6627 | 6.6065 | 7.5503 | 8.4940 | I. 3610 |
| 34 | 0.9436 | 1. 8873 | 2.8309 | 3.7746 | 4.7182 | 5.6619 | 6.6055 | 7.5492 | 8.4928 | I. 3609 |
| 35 | 0.9435 | I. 8870 | 2.8306 | 3.7741 | 4.7176 | 5.6611 | 6.6046 | 7.548I | 8.4917 | I. 3608 |
| 36 | 0.9434 | I. 8868 | 2.8302 | 3.7735 | 4.7169 | 5.6603 | 6.6037 | 7.5471 | 8.4905 | 1. 3607 |
| 37 | 0.9433 | 1. 8865 | 2.8298 | 3.7730 | 4.7163 | 5.6595 | 6.6028 | 7.5460 | 8.4893 | 1. 3606 |
| 38 | 0.9431 | I. 8862 | 2.8294 | 3.7725 | 4.7156 | 5.6587 | 6.6018 | 7.5449 | 8.488 I | I. 3605 |
| 39 | 0.9430 | I. 8860 | 2.8290 | 3.7719 | 4.7149 | 5.6579 | 6.6009 | 7.5439 | 8.4869 | I. 3604 |
| 40 | 0.9429 | 1. 8857 | 2.8286 | 3.7714 | 4.7143 | 5.657 I | 6.6000 | 7.5428 | 8.4857 | 1. 3603 |
| 41 | 0.9427 | 1. 8854 | 2.8282 | 3.7709 | $4.713^{6}$ | 5.6563 | 6.5990 | 7.5418 | 8.4845 | 1. 3602 |
| 42 | 0.9426 | I. 8852 | 2.8278 | 3.7703 | 4.7129 | 5.6555 | 6.598I | 7.5407 | 8.4833 | 1.3602 |
| 43 | 0.9425 | I. 8849 | 2.8274 | 3.7698 | 4.7123 | 5.6547 | 6.5972 | 7.5396 | 8.4821 | I. 3601 |
| 44 | 0.9423 | I. 8846 | 2.8270 | 3.7693 | 4.7116 | 5.6539 | 6.5962 | 7.5385 | 8.4809 | 1. 3600 |
| 45 | 0.9422 | 1. 8844 | 2.8265 | 3.7687 | 4.7109 | 5.6531 | 6.5953 | 7.5375 | $8.47{ }^{\circ} 6$ | 1. 3599 |
| 46 | 0.9420 | 1. 8841 | 2.8261 | 3.7682 | 4.7102 | 5.6523 | 6.5943 | 7.5364 | 8.4784 | I. 3598 |
| 47 | 0.9419 | 1. 8838 | 2.8257 | 3.7677 | 4.7096 | 5.6515 | 6.5934 | 7.5353 | 8.4772 | 1. 3597 |
| 48 | 0.9418 | I. 8836 | 2.8253 | 3.7671 | 4.7089 | 5.6507 | 6. 5925 | 7.5342 | 8.4760 | I. 3596 |
| 49 | 0.9416 | 1. 8833 | 2.8249 | 3. 7666 | 4.7082 | 5.6499 | 6.5915 | 7.5332 | 8.4748 | I. 3595 |
| 50 | c.9415 | 1.8830 | 2.8245 | 3.7660 | 4.7076 | 5.6491 | 6.5906 | 7.5321 | 8.4736 | 1. 3594 |
| 51 | 0.9414 | 1. 8828 | 2.8241 | 3.7655 | 4.7069 | 5.6483 | 6.5896 | 7.5310 | 8.4724 | 1. 3593 |
| 52 | 0.9412 | 1. 8825 | 2.8237 | 3.7650 | 4.7062 | 5.6474 | 6.5887 | 7.5299 | 8.4712 | 1. 3592 |
| 53 | 0.9411 | 1.8822 | 2.8233 | 3.7644 | 4.7055 | 5.6466 | 6.5877 | 7.5288 | 8.4697 | I.3591 |
| 54 | 0.9410 | 1.8819 | 2.8229 | 3.7639 | 4.7048 | 5.6458 | 6.5868 | 7.5278 | 8.4687 | 1. 3590 |
| 55 | 0.9408 | I. 8817 | 2.8225 | 3.7633 | 4.7042 | 5.6450 | 6.5858 | 7.5267 | 8.4675 | 1. 3559 |
| 56 | 0.9407 | 1.8814 | 2.8221 | 3.7628 | 4.7035 | $5.644^{2}$ | 6.5849 | 7.5256 | 8.4663 | I. 3585 |
| 57 | 0.9406 | I. 8811 | 2.8217 | 3. 7623 | 4.7028 | 5.6434 | 6.5839 | 7.5245 | 8.4651 | 1. 3587 |
| 58 | 0.9404 | I. 8809 | 2.8213 | 3.7617 | 4.7021 | 5.6426 | 6.5830 | 7.523+ | 8.4635 | 1.3586 |
| 59 | 0.9403 | 1. 8806 | 2.8209 | 3.7612 | 4.7015 | 5.6418 | 6.5820 | 7.5223 | 8.4626 | I. 3585 |
| 60 | 0.9402 | 1. 8803 | 2.8205 | 3.7606 | 4.7008 | 5.6499 | 6.58 II | 7.5212 | 8.4614 | 1. 3584 |


| $13^{\circ}$ |  | HEIGHTS. |  |  |  |  |  |  | 115 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | b |  |
|  | 0.43 | 0.6 |  | 1.0 | 1.3 | I. 5 | 1.7510 |  | 9 | - |
| 0.2191 | 0.438 | 0.6574 | 0.876 | 1.0957 | 1.3148 | I. 5340 | I. 753 I | 1.9723 | 0.3153 | or |
| 0.2194 | 0.4388 | 0.6582 | 0. 8776 | 1.0970 | 1.3164 | 1.5358 | I. 7552 | I. 9746 | 0.3157 | 02 |
| 0.2197 | 0.4393 | 0.6550 | c. 8786 | 1.0983 | 1.3180 | r. 5376 | I. 7573 | 1.9769 | 0.3161 | 03 |
| 0.2199 | 0.4398 | 0.6598 | 0.8797 | 1.0996 | 1.3195 | I. 5394 | I. 7594 | r. 9793 | 0.3165 | 04 |
| 0.2202 | 0.4404 | 0.6605 | 0.8807 | I. 1009 | 1.3211 | 1.5413 | 1.7614 | 1.9816 | 0.3169 | 05 |
| 0.2204 | 0.4409 | 0.6613 | 0.8818 | 1.1022 | 1. 322 | I.543I | 1. 7635 | 1.9840 | 0.3173 | 06 |
| 0.2207 | 0.4414 | 0.6621 | 0.8828 | I.1035 | 1. 3242 | 1. 5449 | 1. 7656 | 1.9863 | 0.3177 | 07 |
| 0.2210 | 0.4419 | 0.6629 | 0. 8838 | 1.1048 | 1. 3258 | I. 5467 | 1. 7677 | I. 9886 | 0.3181 | 08 |
| 0.2212 | 0.4425 | 0.6637 | 0.8849 | I. | 1. 3274 | 1. 5486 | 1.7698 | 1.9911 | 0.3185 | 09 |
| 0.2215 | 0.4430 | 0.6645 | 0.8860 | 1.1074 | 1.3289 | 1.5504 | 1.7719 | 1.9934 | 0.3189 | 10 |
| 0.2217 | 0.4435 |  |  | 1.1087 | 1. 3305 | 1.5522 | 1.7740 | 1.9957 | 93 | 1 |
| 0.2220 | 0.4440 | 0.6660 | 0.8880 | I. | 1.3321 | 1.5541 | 1.7761 | 1.9981 | 0.3197 | 12 |
| 0.2223 | 0.4445 | 0.6668 | 0.8891 | 1.1113 | 1. 3336 | 1.5559 | 1.7782 | 2.0004 | 0.3201 | 13 |
| 0.2225 | 0.4451 | 0.6676 | 0.8901 | 1.1 | 1.3352 | 1. 5577 | 1.7802 | 2.002 | 0.3205 | 14 |
| 0.2228 | 0.4456 | 0.6684 | 0.8912 | 1.1139 | 1. 3367 | I. 5595 | 1. 7823 | 2.0051 | 0.32 C 9 | 15 |
| 0.2230 | 0.4461 | 0.6691 | 0.8922 | 1.1152 | 1.3383 | 1.5613 | I. 7844 | 2.0074 | 0.3213 | 16 |
| 0. 2233 | 0.4466 | 0.6699 | 0.8932 | 1.11 | 1. 3399 | 1.5632 | 1. 7865 | 2.0098 | 0.3217 | 17 |
| 0.2236 | 0.4471 | 0.6707 | 0. 8943 | 1.11 | 1.3414 | 1. 56 | I. 788 | 2.0121 | 0.3221 | 18 |
| 0. 2238 | 0.4477 | 0.6715 | c. 8953 | I.1191 | 1.3430 | 1.56 | 1. 7906 | 2.0 | 0.3225 | 19 |
| 0.2241 | 0.4482 | 0.6723 | 0.8964 | 1.1204 | 1. 3445 | I. 5686 | 1. 7927 | 2.0168 | 0.3229 | 20 |
| 0. 2243 | 0.4487 |  |  | I. 1217 | I. 3460 | 1. 5704 |  | 2.0191 | 0.3232 | 21 |
| 0. 2246 | 0.4492 | 0.673 | 0.8984 | I. 1230 | 1. 3476 | 1. 5722 | I. 79 | 2.0214 | 0. 3236 | 22 |
| 0.2249 | 0.4497 | 0.6746 | 0.8994 | I. 1243 | 1. 3492 | 1.5740 | 1.7989 | 2.0237 | 0.3240 | 23 |
| 0.2251 | 0.4502 | 0.6754 | 0.9005 | I. 1256 | I. 3507 | I. 5758 | 1.8010 | 2.0261 | 0.3244 | 24 |
| 0.2254 | 0.4508 | 0.6761 | 0.9015 | I. 1269 | 1.3523 | 1.5777 | 1.8030 | 2.0284 | 0. 3248 | 25 |
| 0.2256 | 0.4513 | 0.6769 | 0.9026 | I. 1282 | I. 3538 | I. 5795 | I. 8051 | 2.0308 | 0. 3252 | 26 |
| 0. 2259 | 0.4518 | 0.6777 | 0.9036 | I. 1295 | I. 3554 | I.5813 | 1. 8072 | 2.0331 | 0.3256 | 27 |
| 0. 2262 | 0.4523 | 0.6785 | 0.9046 | I. 1308 | I. 3570 | I. 5831 | 1. 8093 | 2.0354 | 0.3260 | 28 |
| 0.2264 | 0.4528 | 0.6793 | 0.9057 | I.I32I | I. 3585 | I. 5849 | I.8114 | 2.0378 | 0.3264 | 29 |
| 0.2267 | 0.4534 | 0.6800 | 0.5067 | I. 1334 | 1.3601 | 1.5868 | 1.8134 | 2.0401 | 0.3268 | 30 |
| 0.2269 | 0.4539 | 0.680 |  | 1.I347 | 1.3616 | 1.5886 | 1.8155 | 2.0424 | 0.3272 | 31 |
| 0.2272 | 0.4544 | 0.6816 | 0.9088 | I.1360 | 1.3631 | 1. 5904 | 1.8175 | 2.0447 | 0.3276 | 32 |
| 0.2275 | 0.4549 | 0.6824 | 0.9098 | I. 1373 | 1. 3647 | 1.5922 | 1.8196 | 2.0471 | 0.3280 | 33 |
| 0.2277 | 0.4554 | 0.6831 | 0.91 c 8 | I. 138 | 1. 3663 | I. 5940 | 1.8217 | 2.0494 | 0.3284 | 34 |
| 0.2280 | 0.4559 | 0.6839 | 0.9119 | I. 1398 | 1. 3678 | 1.5958 | 1.8238 | $2.05{ }^{1} 7$ | 0.3288 | 35 |
| 0. 2282 | 0.4565 | 0.6847 | 0.9129 | I.141I | 1. 3694 | I. 5976 | 1. 8258 | 2.0541 | 0.3292 | 36 |
| 0. 2288 | 0. 4570 | 0.6855 | 0.9140 | I. 14 | 1. 3709 | 1.5994 | 1.8279 | 2.0564 | 0.3296 | 37 |
| 0.2287 | 0.4575 | 0.6862 | 0.9150 | 1.1437 | 1. 3725 | 1.6012 | 1.8300 | 2.0587 | 0.3300 | 38 |
| 0.2290 | 0.4580 | 0.6870 | 0.916 | I. 1450 | 1. 3740 | 1.6030 | 1.8320 | 2.0610 | 0.3304 | 39 |
| 0.2293 | 0.4585 | 0.6878 | 0.9170 | 1.1463 | 1.3756 | 1. 6048 | 1.834I | 2.0633 | 0.3308 | 40 |
| 229 | 0.4590 | 0.6886 | 0.918I | 1. 1476 | 1.3771 | 1. 6066 | 1.836I |  | 0.3312 | 41 |
| 0.2298 | 0.4596 | 0.6993 | 0.9191 | I. 1489 | I. 3787 | I. 6085 | 1. 8382 | 2.0680 | 0.3316 | 42 |
| 0.2300 | 0.4601 | 0.6901 | 0.9202 | 1.1502 | 1.3802 | 1.6103 | 1.8403 | 2.0703 | 0.3320 | 43 |
| 0.230 | 0.4606 | 0.6909 | 0.9212 | I. 1515 | I. 3817 | 1.6121 | 1. 8423 | 2.0726 | 0.3324 | 44 |
| 0.2306 | 0.4611 | 0.6917 | 0.9222 | I. 1528 | 1. 3833 | 1.6139 | I. 8444 | 2.0750 | 0.3328 | 45 |
| 0.2308 | 0.4616 | 0.6924 | 0.9232 | 1.1541 | I. 3849 | 1. 6157 | 1. 8465 | 2.0773 | 0.3331 | 4 |
| 0.2311 | 0.4621 | 0.6932 | 0.9243 | I. 1554 | I. 3864 | I. 6175 | 1. 8486 | 2.0796 | 0.3335 | 47 |
| 0.2313 | 0.4626 | 0.6940 | 0.9253 | I. 1566 | 1. 3879 | I.6193 | 1.8506 | 2.0819 | 0.3339 | 4 |
| 0.2316 | 0.4632 | 0.6947 | 0.9263 | I. 1579 | 1. 3895 | I.6211 | I. 8527 | 2.0842 | 0.3343 | 49 |
| 0.2318 | 0.4637 | 0.6955 | 0.9274 | 1. 1592 | 1.3910 | 1. 6229 | 1. 8547 | 2.0866 | 0.33 | 50 |
| 0.23 | 0.4642 | 0.6963 | 0.9284 | 1.1605 | 1. 3926 | 1.6247 | 1. 8568 | 2.0889 | 0.3351 | 51 |
| 0.2324 | 0.4647 | 0.6971 | 0.9294 | 1.16 | I. 394 I | 1.6265 | 1. 8588 | 2.0912 | 0.3355 | 52 |
| 0.2326 | 0. 4652 | 0.6978 | 0.9304 | I. 1630 | I. 3957 | 1.6283 | 1.8609 | 2.0935 | 0.3359 | 53 |
| 0. 2329 | 0.4657 | 0.6986 | 0.9315 | 1.1643 | 1. 3972 | 1.6301 | 1.8630 | 2.0958 | 0.3363 | 54 |
| 0.2331 | 0.4662 | 0. 6994 | 0.9325 | 1.1656 | I. 3987 | 1.6319 | 1. 8650 | 2.0981 | 0.3367 | 55 |
| 0.2334 | 0.4668 | 0.7001 | 0.9335 | I. 1669 | I. 4003 | I. 6337 | I. 8670 | 2.1004 | 0.3371 | 56 |
| 0.2336 | 0.4673 | 0.7009 | 0.9346 | 1.1682 | 1.4018 | 1. 6355 | 1.8691 | 2.1028 | 0.3375 | 57 |
| 0.2339 | 0.4678 | 0.7017 | 0.9356 | 1.1695 | 1.4033 | 1.6373 | 1.8711 | 2.1050 | 0.3379 | 5 |
| 0.2342 | 0.4683 | 0.7025 | 0.9366 | 1.1708 | 1. 4049 | 1.6391 | r. 8732 | 2.1074 | 0.3383 | 59 |
| 0.2344 | 0.468 | 0.7032 | 0.9376 | 1.1720 | 1.4065 | 1.6409 | I. 8753 | 2.1097 | 0.3387 | 60 |


| 116 |  |  |  |  | DISTANCES. |  |  |  |  | $14^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | a |
| 00 | 0.9 | 1.88 | 2.8205 | 3.7606 | 4.7008 |  | 6.58 II | 7.5212 | 8.4614 |  |
| OI | 0.9400 | 1.8800 | 2.8201 | 3.7601 | 4.7001 | 5.6401 | 6.5801 | 7.5202 | 8.4602 | I. 3583 |
| 02 | 0.9399 | 1.8798 | 2.8196 | 3.7595 | 4.6994 | 5.6393 | 6.5792 | 7.5191 | 8.4589 | I. 3582 |
| 03 | 0.9397 | I. 8795 | 2.8192 | 3.7590 | 4.6987 | 5.6385 | 6.5782 | 7.5180 | 8.4577 | I. 3581 |
| 04 | 0.9396 | 1. 8792 | 2.8188 | 3.7584 | 4.6980 | 5.6376 | 6.5773 | 7.5169 | 8.4565 | I. 3580 |
| 05 | 0.9395 | 1.8789 | 2.8184 | 3.7579 | 4.6974 | 5.6368 | 6.5763 | 7.5158 | 8.4552 | I. 3579 |
| - | 0.9393 | 1. 8787 | 2.8180 | 3.7573 | 4.6967 | 5.6360 | 6.5753 | 7.5147 | 8.4540 | I. 3578 |
| 07 | - 0.9392 | I. 8784 | 2.8176 | 3.7568 | 4.6960 | 5.6352 | 6.5744 | 7.5136 | 8.4528 | I. 3577 |
| 08 | 0.9391 | I. 878 I | 2.8172 | 3.7562 | 4.6953 | 5.6344 | 6.5734 | 7.5125 | 8.4515 | I. 3576 |
| 09 | 0.9389 | 1. 8778 | 2.8168 | $3 \cdot 7557$ | 4.6946 | 5.6335 | 6.5725 | 7.5114 | 8.4503 | I. 3575 |
| 10 | 0.0388 | 1. 8776 | 2.8164 | 3.755 I | 4. 6939 | 5.6327 | 6.5715 | 7.5103 | 8.4491 | I. 3574 |
| II | 0.9386 | 1. 8773 | 2.8159 | 3.7546 | 4.6932 | 5.6319 | 6.5705 | 7.5092 | 8.4478 | I. 3573 |
| 12 | 0.9385 | 1.8770 | 2.8155 | 3.7540 | 4.6925 | 5.6310 |  | 7.508I | 8.4466 | I. 3572 |
| I3 | 0.9384 | I. 8767 | 2.8151 | 3.7535 | 4.6918 | 5.6302 | 6. 5686 | $7 \cdot 5070$ | 8.4453 | I. 3571 |
| 14 | 0.9382 | I. 8765 | 2.8147 | 3.7529 | 4.6912 | 5.6294 | 6.5676 | 7.5058 | 8.4441 | I. 3570 |
| 15 | 0.9381 | 1.8762 | 2.8143 | 3.7524 | 4.6505 | 5.6286 | 6.5666 | 7.5047 | 8.4428 | I. 3569 |
| 16 | 0.9380 | 1. 8759 | 2.8139 | 3.7518 | 4.6898 | 5.6277 | 6.5657 | 7.5036 | 8.4416 | I. 3568 |
| 17 | 0.9378 | I. 8756 | 2.8134 | 3.7513 | 4.6891 | 5.6269 | 6.5647 | 7.5025 | 8.4403 | I. 3567 |
| 18 | 0.9377 | 1.8754 | 2.8130 | 3.7507 | 4.6884 | 5.6261 | 6.5637 | 7.5014 | 8.4391 | I. 3566 |
| 19 | 0.9375 | 1.8751 | 2.8126 | 3.7501 | 4.6877 | 5.6252 | 6.5628 | 7.5003 | 8.4378 | I. 3565 |
| 20 | 0.9374 | I. 8748 | 2.8122 | 3.7496 | 4.6870 | 5.6244 | 6.5618 | $7 \cdot 4992$ | 8.4366 | I. 3564 |
| 21 | 0.9373 | 1. 8745 | 2.8118 | 3.7490 | 4.6863 | 5. 6235 | 6.5608 | 7.4981 | 8.4353 | 1.3563 |
| 22 | 0.937 I | 1.8742 | 2.8114 | 3.7485 | 4.6856 | 5.6227 | 6.5598 | $7 \cdot 4969$ | 8.4341 | 1.3562 |
| 23 | 0.9370 | 1.8740 | 2.8109 | 3.7479 | 4.6849 | 5.6219 | 6.55 | 7.4958 | 8.4328 | 1.3561 |
| 24 | 0.936 | 1. 8737 | 2.8105 | 3.7474 | 4.6842 | 5.6210 | 6.557 | 7.4947 | 8.4315 | 1. 3560 |
| 25 | 0.9367 | 1. 8734 | 2.810 | 3.7468 | 4.6835 | 5.6202 | 6.5569 | 7.4936 | 8.4303 | I. 3559 |
| 26 | 0.936 | 1. 8731 | 2.8097 | 3.7462 | 4.6828 | 5.6193 | 6.5559 | 7.4925 | 8.4290 | I. 3558 |
| 27 | 0.9364 | 1. 8728 | 2.8093 | 3.7457 | 4.68 | 5.6185 | 6.5549 | 7.4913 | 8.4278 | 1.3557 |
| 28 | 0.9363 | 1. 8726 | 2.8088 | 3.7451 | 4.6814 | 5.6177 | 6.5539 | 7.4902 | 8.4265 | 1. 3556 |
| 29 | 0.9361 | 1.8723 | 2.8084 | 3.7445 | 4.6807 | 5.6168 | 6.5530 | 7.4891 | 8.4252 | 1. 3555 |
| 30 | 0.9360 | 1.8720 | 2.8080 | 3.7440 | 4.6800 | 5.6160 | 6.5520 | 7.4880 | 8.4240 | I. 3554 |
| 31 | 0.9359 | 1.8717 | 2.8076 | 3.7434 | 4.6793 | 5.6151 | 6.5510 | 7.4868 | 8.4227 | 1. 3553 |
| 32 | 0.9357 | I. 8714 | 2.8071 | 3.7429 | 4.6786 | 5.6143 | $6.55 c 0$ | 7.4857 | 8.4214 | I. $355^{2}$ |
| 33 | 0.9356 | 1.8711 | 2.8067 | 3.7423 | 4.6779 | 5.6134 | 6.5490 | 7.4846 | 8.4202 | 1.3551 |
| 34 | 0.9354 | 1.8709 | 2.8063 | 3.7417 | 4.6772 | 5.6126 | 6.5480 | 7.4834 | 8.4189 | I. 3550 |
| 35 | 0.9353 | 1.8706 | 2.8059 | 3.7412 | 4.6764 | 5.6117 | 6.5470 | 7.4823 | 8.4176 | 1. 3549 |
| 36 | 0.9351 | 1.8703 | 2.8054 | 3.7405 | 4.6757 | 5.6109 | 6.5460 | 7.4812 | 8.4163 | I. 3548 |
| 37 | 0.9350 | 1.8700 | 2.8050 | 3.7400 | 4.6750 | 5.6100 | 6.5450 | 7.4801 | 8.4151 | I. 3547 |
| 38 | 0.9349 | 1.8697 | 2.8046 | 3.7395 | 4.6743 | $5.609^{2}$ | 6.544 I | 7.4789 | 8.4138 | I. 3546 |
| 39 | 0.9347 | 1.8694 | 2.8042 | 3.7389 | 4.6736 | 5.6083 | 6.5431 | 7.4778 | 8.4125 | I. 3545 |
| 40 | 0.9346 | 1.8692 | 2.8037 | 3.7383 | 4.6729 | 5.6075 | 6.542 I | 7.4767 | 8.4112 | 1.3544 |
| 41 | 0.9344 | 1.8689 | 2.8033 | 3.7378 | 4.6722 | 5.6066 | 6.5411 | 7.4755 | 8.4100 | I. 3543 |
| 42 | 0.9343 | 1.8686 | 2.8029 | 3.7372 | 4.6715 | 5.6058 | 6.5401 | 7.4744 | 8.4087 | I. 3542 |
| 43 | 0.9342 | 1.8683 | 2.8025 | 3.7366 | 4.6708 | 5.6049 | 6.5391 | 7.4732 | 8.4074 | I. 354 I |
| 44 | 0.9340 | 1.8680 | 2.8020 | 3.7360 | 4.6701 | 5.6041 | 6.5381 | 7.4721 | 8.4061 | I. 3540 |
| 45 | 0.9339 | 1. 8677 | 2.8016 | 3.7355 | 4.6693 | 5.6032 | 6.5371 | 7.4709 | 8. 4048 | I. 3539 |
| 46 | 0.9337 | 1. 8674 | 2.8012 | 3.7349 | 4.6686 | 5.6023 | 6.5361 | 7.4698 | 8.4035 | I. 3538 |
| 47 | 0.9336 | 1.8672 | 2.8007 | 3.7343 | 4.6679 | 5.6015 | 6.5351 | 7.4686 | 8.4022 | 1.3537 |
| 48 | 0.9334 | 1. 8669 | 2.8003 | 3.7338 | 4.6672 | 5.6006 | 6.5341 | 7.4675 | 8.4009 | I. 3536 |
| 49 | 0.9333 | 1. 8666 | 2.7999 | 3.7332 | 4.6665 | 5.5998 | 6.5331 | 7.4664 | 8.3997 | I. 3535 |
| 50 | 0.9332 | 1. 8663 | 2.7995 | 3.7326 | 4.6658 | 5.5989 | 6.5321 | 7.4652 | 8.3984 | I. 3534 |
| 51 | 0.9330 | 1. 8660 | 2.7990 | 3.7320 | 4.6650 | 5.5980 | 6.5311 | 7.464 I | 8.3971 | 1. 3533 |
| 52 | 0.9329 | I. 8657 | 2.7986 | 3.7315 | 4.6643 | 5.5972 | 6.5300 | 7.4629 | 8.3958 | I. 3531 |
| 53 | 0.9327 | 1. 8654 | 2. 7982 | 3.7309 | 4.6636 | 5.5963 | 6.5290 | 7.4618 | 8.3945 | 1.3530 |
| 54 | 0.9326 | 1.8651 | 2.7977 | 3.7303 | 4.6629 | $5 \cdot 5954$ | 6.5280 | 7.4606 | 8.3932 | 1.3529 |
| 55 | 0.9324 | 1. 8649 | 2.7973 | 3.7297 | 4.6621 | 5.5946 | 6.5270 | 7.4594 | 8.3919 | 1.3528 |
| 56 | 0.9323 | 1.8646 | 2.7959 | 3.7291 | 4.6614 | $5 \cdot 5937$ | 6.5260 | 7.4583 | 8. 3906 | 1.3527 |
| 57 | 0.9321 | 1.8643 | 2.7964 | 3.7286 | 4.6607 | 5.5928 | 6.5250 | 7.4571 | 8. 3893 | 1.3526 |
| 58 | 0.9320 | 1.8640 | 2.7960 | 3.7280 | 4.6600 | 5.5920 | 6.5240 | 7.4560 | 8.3880 | 1.3525 |
| 59 | 0.9319 | I. 8637 | 2.7956 | 3.7274 | 4.6593 | 5.5911 | 6.5230 | $7.45+8$ | 8.3867 | 1. 3524 |
| 60 | 0.9317 | 1.8634 | 2.795 I | 3.7268 | 4.6585 | $5 \cdot 5902$ | 6.5219 | 7.453 | $8.385+$ | 1.3523 |


| $14^{\circ}$ | HEIGHTS. |  |  |  |  |  |  |  |  | 117 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | 13 |  |
| 0.2 | 0.468 | 0.7032 | 0.93 | I. 1720 |  |  | 1. 8753 | 2.1097 | 0.3387 |  |
| 0.2347 | 0.460 | 0.7040 | 0.9386 | I. 1733 | 1. 4080 | 1.6426 | 1. 8773 | 2.1120 | 0.3391 | I |
| 0.2349 | 0. 4698 | 0.7048 | 0.9397 | 1.1746 | 1. 4095 |  | I. 8794 | 2.1143 | 0. 3395 | 2 |
| . $0.235^{2}$ | 0.4704 | 0. 7055 | 0.9407 | I. 1759 | I.4111 | 1.6462 | r.8814 | 2.1166 | 0. 3399 | 03 |
| 0.2354 | 0.4709 | 0.7063 | 0.9417 | 1.1772 | 1.4126 | 1.6480 | I. 8834 | 2.1189 | 0.3403 | 04 |
| 0.2357 | 0.4714 | 0.7071 | 0.9428 | I. 1785 | 1.4141 | 1. 6498 | I. 8855 | 2.1212 | 0.3407 | 5 |
| 0.2359 | 0.4719 | 0.7078 | 0.9438 | 1.1797 | 1.4156 | 1.6516 | 1.8875 | 2.1235 | 0.34II | c6 |
| 0.2362 | 0.4724 | 0.7086 | 0.9448 | 1.1810 | 1. 4172 | 1.6534 | 1.8896 | 2.1258 | 0.3414 | 7 |
| 0.2365 | 0.4729 | 0.7094 | 0.9458 | 1.1823 | I. 4188 | 1. 6552 | I. 8917 | 2.1 | 0.3418 | 8 |
| 0.2367 | 0.4734 | 0.7101 | 0.9468 | I. 1836 | 1.4203 | 1. 6570 | 1.8937 | 2.1304 | 0.3422 | c9 |
| 0.2370 | 0.4739 | 0.7109 | 0.9479 | 1. 1848 | 1.4218 | 1. 6588 | 1. 8958 | 2.1327 | 0.3426 | 10 |
| 0.2372 | 0.4744 | 0.7117 | 0.9489 | I. | 1.4233 | 1.6606 | 1.8978 | 2.1350 | 0. 3430 | 1 |
| 0.2375 | 0.4750 | 0.7124 | 0.9499 | I. 1874 | 1. 4249 | 1.6624 | I. 8998 | 2.1373 | 0.3434 | 12 |
| 0.23771 | 0.4755 | 0.7132 | 0.9509 | I. 1887 | 1. 4264 | 1.6641 | 1.9018 | 2.1396 | $0.343^{8}$ | 13 |
| 0.2380 , | 0.4760 | 0.7140 | 0.9520 | I. 1899 | I. 4279 | I. 6659 | 1.9039 | 2.1419 | 0.3442 | 14 |
| 0.2382 | 0.4765 | 0. 7147 | 0.9530 | I. 1912 | I. 4295 | 1. 6677 | 1.9060 | 2.1442 | 0.3446 | 15 |
| 0.2385 | 0.4770 | 0.7155 | 0.9540 | 1.1925 | 1.4310 | 1. 6695 | 1.9080 | 2.1465 | 0.3450 | 16 |
| 0.2388 | 0.4775 | 0.7163 | 0.9550 | I. 1938 | 1.4326 | 1.6713 | 1.9101 | 2.1488 | c. 3454 | 17 |
| 0.2390 | 0.4780 | 0.7170 | 0.9560 | I. 1951 | I.434I | 1.673I | 1.9121 | 2.1511 | 0.3458 | 18 |
| 0.2393 | 0.4785 | 0.7178 | 0.9571 | I. 1963 | 1.4356 | 1. 6749 | I.9142 | 2.1534 | 0.3462 | 19 |
| 0.2395 | 0.4790 | 0.7186 | c.958I | I. 1976 | I.4371 | 1. 6767 | 1.9162 | 2. 1557 | 0.3466 | 20 |
| 0.2398 | 0.4796 | 0.7193 | 0.9591 | 1.1989 | 1.4387 |  | 1.9182 |  | 0.3470 | 21 |
| 0.2400 | 0.4801 | 0.7201 | 0.960I | 1.2002 | I. 4402 | 1.6802 | 1.9202 | 2.1603 | 0.3474 | 22 |
| 0.2403 | 0.4806 | 0.7209 | 0.9611 | I. 2014 | I. 4417 | I. 6820 | 1.9223 | 2.1626 | 0.3478 | 23 |
| 0.2405 | 0.4811 | 0.7216 | 0.9622 | 1.2027 | I. 4432 | I. 6838 | 1.9243 | 2.1648 | 0.3482 | 24 |
| 0.2408 | 0.4816 | 0.7224 | 0.9632 | I. 2040 | I. 4448 | 1.6856 | I.9264 | 2.1671 | 0.3485 | 25 |
| 0.2411 | 0.4821 | 0.7232 | 0.9642 | I. 2053 | 1.4463 | 1. 6874 | 1.9284 | 2.1694 | 0.3489 | 26 |
| 0.2413 | 0.4826 | 0. 7239 | 0.9652 | I. 2065 | I. 4478 | 1.6891 | 1.9304 | 2.1717 | 0.3493 | 27 |
| 0.2416 | 0.4831 | 0. 7247 | 0.9662 | 1.2078 | I. 4494 | 1. 6909 | I.9325 | 2.1740 | 0.3497 | 28 |
| 0.2418 | 0.4836 | 0. 7254 | 0.9672 | 1.2091 | 1.4509 | 1.6927 | I. 9345 | 2.1763 | 0.3501 | 29 |
| 0.2421 | 0.4841 | 0.7262 | 0.9683 | 1.2103 | 1. 4524 | 1. 6945 | 1.9366 | 2.1786 | 0.3505 | 30 |
| 0.2423 | 0.4846 | 0.7270 | 0.9693 | . 2116 | 1.4539 | 1. 6962 | 1.9386 | 2.1809 | 0.3509 | 31 |
| 0.2425 | 0.4851 | 0.7277 | 0.9703 | x.2129 | 1. 4554 | 1.6980 | 1.9406 | 2.1831 | 0.3513 | 32 |
| 0.2428 | 0.4857 | 0. 7285 | 0.9713 | 1.214' | I. 4570 | 1. 6998 | I.9426 | 2.1855 | 0.3517 | 33 |
| 0.2431 | 0.4862 | 0. 7292 | 0.9723 | 1. 2154 | I. 4585 | 1.7016 | I. 9446 | 2.1877 | 0.352 I | 34 |
| 0.2433 | 0.4867 | 0. 7299 | 0.9733 | 1. 2166 | 1.4600 | 1.7033 | 1.9466 | 2.1900 | 0.3525 | 35 |
| 0.2436 | 0.4872 | 0.7307 | 0.9743 | I. 2179 | I. 4615 | I.705I | I. 9487 | 2.1923 | 0.3529 | 36 |
| 0.2438 | 0.4877 | 0.7315 | 0.9754 | 1.2192 | I. 4630 | 1. 7069 | 1.9507 | 2. 1946 | 0. 3533 | 37 |
| 0.2441 | 0.4882 | 0.7323 | 0.9764 | 1. 2205 | I. 4646 | 1. 7087 | 1.9528 | 2. 1969 | 0.3537 | 38 |
| 0.2443 | 0. 4847 | 0.7330 | 0.9774 | 1. 2217 | I. 4661 | I. 7104 | 1.9548 | 2.1991 | 0.3541 | 39 |
| 0.2446 | 0.4892 | 0.7338 | 0.9784 | 1.2230 | 1.4676 | 1.7122 | 1.9568 | 2.2014 | 0.3545 | 40 |
| 0.2449 | 0.4897 | 0.7346 | 0.9794 | 1.2243 | 1.4691 | 1. 7140 | 1.9588 | 2.2037 | 0.3549 | 1 |
| 0.2451 | 0.4902 | 0.7353 | 0.98 c 4 | 1. 2255 | 1.4706 | 1.7157 | 1.9608 | 2.2059 | 0.3553 | 42 |
| 0.2454 | 0.4907 | 0.7361 | 0.9814 | 1. 2268 | I. 4722 | I. 7175 | 1.9629 | 2.2082 | 0.3556 | 43 |
| 0.2456 | 0.4912 | 0. 7368 | 0.9824 | 1.2281 | I. 4737 | 1.7193 | 1.9649 | 2.2105 | 0.3560 | 44 |
| 0.2459 | 0.4917 | 0.7376 | 0.9835 | 1. 2294 | 1. 4752 | 1.7211 | 1.9670 | 2.2128 | 0.3564 | 45 |
| 0.246 I | 0.4922 | 0.7384 | 0.9845 | 1. 2305 | 1. 4767 | 1.7228 | 1.9690 | 2.2151 | 0.3568 | 46 |
| 0.2464 | 0.4927 | 0.7391 | 0.9855 | 1.2319 | I. 4782 | I. 7246 | I.9710 | 2.2173 | 0.3572 | 47 |
| 0.2466 | 0.4932 | 0.7399 | 0.9865 | 1.2331 | I. 4797 | 1.7263 | 1.9730 | 2.2196 | 0.3576 | 48 |
| 0.2469 | 0.4938 | 0. 7406 | 0.9875 | 1.2344 | 1.4813 | 1.7281 | 1.9750 | 2.2219 | 0.3580 | 49 |
| 0.2471 | 0.4943 | 0.7414 | c. 9885 | 1.2357 | 1.4828 | 1. 7299 | 1.9770 | 2.2242 | 0.3584 | 50 |
| 0.2474 | 0.4948 | 0.7421 | 0.9895 | 1. 2369 | 1. 4843 | 1.7317 | 1.9790 | 2.2264 | 0.3588 | 51 |
| 0.2476 | 0.4953 | 0.7429 | 0.9905 | 1. 2382 | 1. 4858 | I. 7334 | 1.9810 | 2.2287 | 0.3592 | 52 |
| 0.2479 | 0.4958 | 0.7436 | 0.9915 | 1. 2394 | 1. 4873 | 1. 7353 | 1.983 I | 2.2310 | 0.3596 | 53 |
| 0.2481 | 0.4963 | 0.7444 | 0.9926 | 1.2407 | 1. 4888 | 1. 7370 | 1.9851 | 2.2333 | 0.3600 | 54 |
| 0. 2484 | -. 4968 | $0.745^{2}$ | 0.9936 | 1. 2420 | 1.4903 | 1. 7387 | 1.9871 | 2.2355 | 0.3604 | 55 |
| 0.2486 | 0.4973 | 0.7459 | 0.9946 | I. 2432 | I. 4918 | I. 7405 | 1.9891 | 2.2378 | 0. 3608 | 56 |
| 0.2489 | 0. 4978 | 0. 7467 | 0.9956 | I. 2445 | I. 4933 | 1.7422 | 1.9911 | 2.2400 | 0.3612 | 57 |
| 0.2491 | 0. 4983 | $\text { c. } 7474$ | 0.9966 | 1. 2457 | 1. 4949 | 1.7440 | 1.9932 | $2.2423$ | 0.3616 | 58 |
| 0.2494 0.2497 | 0.4988 0.4993 | 0.7482 0.7490 | 0.9976 0.9986 | I. 2470 I. 2483 | I. 4964 I. 4979 | I.7458 <br> I. 7476 | 1.9952 I.9972 | 2.2446 2.2469 | 0.3620 0.3623 | 59 |
| 0.2497 | 0.4993 | 0.7490 | 0.9986 | 1.2483 | 1.4979 | 1.7476 | 1.9972 | 2.2469 | 0.36 | 60 |


| 118 |  | DISTANCES. |  |  |  |  |  |  |  | $15^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | a |
| $\infty$ | 0.9317 | 1.8634 | 2.7951 | 3.7268 | 4.6585 | 5.5902 | 6.5219 | 7.4537 | 8.3854 | 523 |
| OI | 0.9316 | 1. 8631 | 2.7947 | 3.7262 | 4.6578 | $5 \cdot 5894$ | 6.5209 | 7.4525 | 8.3840 | I. 3522 |
| 02 | 0.9314 | 1. 8628 | 2.7942 | 3.7257 |  | 5.5885 | 6.5199 | 7.4513 | 8.3827 | I. 3521 |
| 03 | 0.9313 | 1.8625 | 2.7938 | 3.7251 | 4. | 5.5876 | 6.5189 | 7.4502 | 8.3814 | I. 3520 |
| 04 | 0.93 II | I. 8622 | 2.7934 | 3.7245 | 4.6556 | $5 \cdot 5867$ | 6.5179 | 7.4490 | 8.3801 | I. 3519 |
| 05 | 0.9310 | 1.8620 | 2.7929 | 3.7239 | 4.6549 | $5 \cdot 5859$ | 6.5168 | 7.4478 | 8.3788 | I. 3518 |
| 06 | -0.9308 | 1.8617 | 2. 7925 | 3.7233 | 4.6542 | $5 \cdot 5850$ | 6.5158 | 7.4466 | 8.3775 | I. 3517 |
| 07 | 0.9307 | 1. 8614 | 2. 7921 | 3.7227 | 4.6534 | 5.5841 | 6.5148 | 7.4455 | 8.3762 | I. 3516 |
| 08 | 0.9305 | I. 8611 | 2.7916 | 3.7222 | 4.6527 | 5.5832 | 6.5138 | 7.4443 | ${ }^{8.3749}$ | I. 3515 |
| 99 | 0.9304 | . 8608 | 2.7912 | 3.7216 | 4.6520 | 5.5824 | 6.5128 | 7.4432 | 8.3736 | I. 3514 |
| Io | 0.9302 | 1. 8605 | 2.7907 | 3.7210 | 4.6512 | 5.5815 | 6.5117 | 7.4420 | 8.3722 | I. 3513 |
| II | 0.9301 | 1.8602 | 2.7903 | 3.7204 | 4.6505 | 5.5806 | 6.5107 | 7.4408 | 8.3709 | I. 3512 |
| 12 | 0.9300 | 1. 8599 | 2.7899 | 3.7198 | 4.6498 | 5.5797 | 6.5097 | 7.4396 |  | I. 3511 |
| 13 | 0.9298 | 1. 8596 | 2.7894 | 3.7192 | 4.6490 | $5 \cdot 5788$ | 6.5086 | 7.4384 | 8. 3682 | I. 3510 |
| 14 | 0.9297 | 1.8593 | 2.7890 | 3.7186 | 4.6483 | 5.5779 | 6.5076 | 7.4373 | 8.3669 | I. 3509 |
| 15 | 0.9295 | 1.8593 | 2.7885 | 3.7180 | 4.6476 | 5.577I | 6.5066 | $7 \cdot 4361$ | 8.3656 | I. 3508 |
| 16 | 0.9294 | I. 8587 | 2.788 I | 3.7175 | 4.6468 | 5.5762 | 6.5055 | 7.4349 | 8. 3643 | r. 3507 |
| 17 | 0.9292 | 1. 8584 | 2.7876 | 3.7169 | 4.646 I | 5.5753 | 6.5045 | 7.4337 | 8. 3629 | I. 3506 |
| 18 | 0.9291 | I. 858 I | 2.7872 | $3.716_{3}$ | 4.6453 | 5.5744 | 6.5035 | 7.4325 | 8.3616 | I. 3504 |
| 19 | 0.9289 | 1. 8578 | 2.7868 | 3.7157 | 4.6446 | 5.5735 | 6.5024 | 7.4314 | 8.3603 | r. 3503 |
| 20 | 0.9288 | 1. 8575 | 2.7863 | 3.7151 | 4.6439 | $5 \cdot 5726$ | 6.5014 | 7.4302 | 8.3590 | 1.3502 |
| 21 | 0.9286 | 1. 8572 | 2.7859 | 3.7145 | 4.643 I | 5.5717 | 6. 5004 | 7.4290 | 8.3576 | 3501 |
| 22 | 0.9285 | I. 8570 | 2.7854 | 3.7139 | 4.6424 | 5.5709 | 6.4993 | 7.4278 | 8.3563 | I. 3500 |
|  | 0.9283 | I. 8567 | 2.7850 | 3.7133 | 4.6416 | 5.5700 | 6.4983 | 7.4266 | 8.3549 | I. 3499 |
| 24 | 0.9282 | I. 8564 | 2.7845 | 3.7127 | 4.6409 | 5.5691 | 6.4972 | 7.4254 | 8.3536 | I. 3498 |
|  | 0.9280 | I. 8561 | 2.7841 | 3.7121 | 4.6401 | $5 \cdot 5682$ | 6.4962 | $7 \cdot 4242$ | 8.3523 | I. 3497 |
| 26 | 0.9279 | I. 8558 | 2.7836 | 3.7115 | 4.6394 | $5 \cdot 5673$ | 6.4952 | 7.4230 | 8.3509 | I. 3496 |
| 27 | 0.9277 | I. 8555 | 2.7832 | 3.7109 | 4.6387 | 5.5664 | 6.4941 | 7.4218 | 8.3496 | I. 3495 |
| 28 | 0.9276 | 1. 8552 | 2.7827 | 3.7103 | 4.6379 | 5.5655 | 6.4931 | 7.4207 | 8.3482 | I. 3494 |
| 29 | 0.9274 | I. 8549 | 2.7823 | 3.7097 | 4.6372 | 5. 5646 | 6.4920 | 7.4195 | 8.3469 | I. 3493 |
| 30 | 0.9273 | 1. 8546 | 2.7819 | 3.7091 | 4.6364 | $5 \cdot 5637$ | 6.4910 | 7.4183 | $8.345^{6}$ | 1.3491 |
| 31 | 0.9 | 1.8543 | 2. 7 SI 4 | 3.7085 | 4.6357 | 5.5628 | 6.4899 | 7.4171 | 8.3442 | 3490 |
| 32 | 0.9270 | I. 8540 | 2.7809 | 3. 7079 | $4.63+9$ | 5.5619 | 6.4889 | $7 \cdot 4159$ | 8.3428 | 3489 |
| 33 | 0.9268 | I. 8537 | 2.7805 | 3.7073 | $4.634^{2}$ | 5.5610 | 6.4878 | $7 \cdot 4147$ | 8.3415 | I. 3488 |
| 3 | 0.0267 | I. 8534 | 2.7800 | 3.7067 | 4.6334 | 5.5601 | 6.4868 | 7.4135 | 8.3401 | I. 3487 |
| 3 | 0.9265 | I. 8531 | 2.7796 | 3.705I | 4.6327 | 5.5592 | 6.4857 | 7.4123 | 8.3388 | I. 3486 |
| 36 | 0.9264 | I. 8528 | 2.7791 | 3.7055 | 4.6319 | 5.5583 | 6.4847 | 7.4111 | 8.3374 | I. 3485 |
| 37 | 0.9262 | I. 8525 | 2.7787 | 3.7049 | 4.6312 | 5.5574 | 6.4836 | 7.4098 | 8.3361 | I. 3484 |
| 38 | 0.9261 | 1. 8522 | 2.7782 | 3. 7043 | 4.6304 | 5.5565 | 6.4826 | 7.4086 | S. 3347 | I. 3483 |
| 39 | 0.9259 | 1. 8519 | 2.7778 | 3.7037 | 4.6297 | 5. 5556 | 6.4815 | 7.4074 | S. 3334 | r. 3482 |
| 40 | 0.9258 | I. 5516 | 2.7773 | 3.703 I | 4.6289 | $5 \cdot 5547$ | 6.4805 | $7 \cdot 4062$ | 8.3320 | I. 3480 |
| 41 | 0.9256 | 1.8513 | 2.7769 | 3.7025 | 4.628 I | 5.553 ${ }^{\text {S }}$ | 6.4794 | 7.4050 | 8.3307 | 1. 3479 |
| 42 | 0.9255 | 1.8510 | 2.7764 | 3.7019 | 4.6274 | 5.5529 | 6.4783 | 7.4038 | 8.3293 | I. 3478 |
| 43 | 0.9253 | I. 8506 | 2.7760 | 3.7013 | 4.6266 | 5.5519 | 6.4773 | 7.4026 | 8. 3279 | I. 3477 |
| 44 | 0.9252 | 1. 8503 | 2.7755 | 3.7007 | 4.6259 | 5.5510 | 6.4762 | 7.4014 | S. 3266 | 1. $3+76$ |
| 45 | 0.9250 | I. 8500 | 2.7751 | 3.7001 | 4.6251 | 5.5501 | 6.4751 | $7 \cdot 4002$ | S. 3252 | 1. 3475 |
| 46 | 0.9249 | I. 8497 | 2.7746 | 3.6995 | 4.6243 | 5.5492 | 6.4741 | 7.3990 | S. 3238 | I. 3474 |
| 47 | 0.9247 | I. 8494 | 2.7742 | 3.6989 | 4.6236 | $5 \cdot 54{ }^{8}$ | 6.4730 | 7.3977 | 8.3225 | I. 3473 |
| 48 | 0.9246 | 1.8491 | 2.7737 | 3.6983 | 4.6228 | 5.5474 | 6.4720 | 7.3965 | S. 3211 | I. $3+72$ |
| 49 | 0.9244 | I. 8488 | 2.7732 | 3.6977 | 4.6221 | 5.5465 | 6.4709 | 7.3953 | 8.3197 | I. 3470 |
| 50 | 0.9243 | 1. 8485 | 2.7728 | 3.6970 | 4.6213 | $5 \cdot 5456$ | 6.4698 | 7.3941 | 8.3284 | I. 3469 |
| 51 | 0.9241 | 1. 8482 | 2.7723 | 3.6964 | 4.6205 | 5.5447 | 6.4688 | 7.3929 | 8.3170 | 1. 3468 |
| 52 | 0.9240 | 1. 8479 | 2.7719 | 3.6958 | 4.6198 | $5 \cdot 5437$ | 6.4677 | 7.3916 | 8.3156 | I. 3467 |
| 53 | 0.9238 | 1. 8476 | 2.7714 | 3.6952 | 4.6190 | 5.5428 | 6.4666 | 7.3904 | 8.3142 | I. 3466 |
| 54 | 0.9236 | 1. 8473 | 2.7709 | 3.6946 | 4.6 IS 2 | 5.5419 | 6.4655 | $7 \cdot 3892$ | 8.3128 | I. 3465 |
| 55 | 0.9235 | I. 8470 | 2.7705 | 3.6940 | 4.6175 | 5.5410 | 6.4645 | 7.3880 | 8.3115 | I. 3464 |
| 56 | 0.9233 | 1. 8467 | 2.7700 2.7606 | 3.6934 | 4.6167 4.6159 | 5.5401 | 6.4634 | 7.3867 | 8.3101 8.3087 | I. $3+63$ |
| 57 | 0.9232 | I. 8464 | 2.7696 2.7601 | 3.6928 | 4.6159 | 5.5391 | 6.4623 | 7.3855 | 8.3087 8.3073 | I. $3+616$ |
| 58 | 0.9230 | I. 8461 | 2.7691 2.7686 | 3.6921 | 4.6152 | 5.5382 | 6.4613 | $7.3 S_{4}$ | 8. 3073 | I. 3460 |
| 59 | 0.9228 | I. 8458 | 2.7686 | 3.6915 | 4.6144 | 5.5373 | 6.4602 | 7.3831 | 8.3059 | 1. 3459 |
| 60 | 0.9227 | I. 8455 | 2.7982 | 3.6909 | 4.6137 | $5 \cdot 5364$ | 6.4591 | $7 \cdot 38 \mathrm{IS}$ | 8. 3946 | 1. 3458 |


| $15^{\circ}$ | HEIGHTS. |  |  |  |  |  |  |  | 119 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | b |  |
|  |  |  |  |  | 1. 4979 | 1. 7476 | 1. | 2.2469 | 3 | 0 |
| 0.2409 | 0. 4998 | 0. 7497 | 0.9996 | I. 2495 | 1. 4994 | 1. 7493 | 1.9992 | 2.2491 | 0.3627 | 1 |
| 0.2502 | 0.5003 | 0.7505 | 1.0006 | 1.2508 | 1.5009 | 1.7511 | 2.0012 | 2.2514 | 0.3631 | 2 |
| 0.250 | 0.5008 | 0.7512 | 1.0016 | 1.2520 | 1. 5024 | I. 7528 | 2.0032 | 2.2536 | -. 3635 | 3 |
| 250 | 0.5013 | 0.7520 | 1.0026 | I. 2533 | 1.5040 | I. 7546 | 2.0053 | 2.2559 | 0.3639 | 04 |
| 0.2509 | 0.5018 | 0.7527 | 1.0036 | 1.2545 | I. 5055 | I. 7564 | 2.0073 | 2.2582 | 0.3643 | 5 |
| 25 | 0.5023 | -. 7535 | 1.0046 | 1.255 ${ }^{\circ}$ | 1.5070 | 1.7581 | 2.0093 | 2.2604 | 0. 3647 | 06 |
| 0.2515 | 0.5028 | 0.7542 | 1.0056 | I. 2570 | 1. 5084 | I. 7599 | 2.0113 | 2.2627 | 0.365I | 7 |
| 0.2517 | 0.5033 | 0.7550 | 1.co66 | I. 2583 | 1.5100 | 1.7616 | 2.0133 | 2.2649 | 0.3655 | 8 |
| 0.2519 | -. 5038 | 0.7557 | 1.0076 | I. 2596 | 1.5115 | 1.7634 | 2.0153 | 2.2672 | 0.3659 | O9 |
| 0.2522 | 0.5043 | 0.7565 | 1.0086 | 1. 2608 | 2.5130 | 1.7651 | 2.0173 | 2.2694 | 0.3663 | Io |
|  |  | 72 |  | 1.2621 |  | 1.7669 |  | 2.2717 |  | II |
| 0.2527 | 0. 5053 | -. 7580 | 1.0106 | 1.2633 | 1.5160 | 1. 7686 | 2.0213 | 2.2739 | 0.3671 | 12 |
| 0.2529 | 0.5058 | -. 7587 | 1.0116 | 1.2646 | 1.5175 | 1. 7704 | 2.0233 | 2.2762 | 0.3674 | I3 |
| 0.2532 | 0.5063 | 0.7595 | 1.0126 | I. 2658 | 1.5190 | I. 7721 | 2.0253 | 2.2784 | 0.3678 | 14 |
| 0.2534 | 0.5068 | 0.7602 | 1.0136 | I. 2671 | 1. 5205 | 1. 7739 | 2.0273 | 2.2807 | 0. 3682 | 15 |
| 0.2537 | 0.5073 | 0.7610 | 1.0146 | I. 2683 | 1. 5220 | I. 7756 | 2.0293 | 2.2829 | 0. 3686 | 16 |
| 0.253 | 0.5078 | 0.7617 | 1.0156 | 1. 2696 | 1. 5235 | 1. 7774 | 2.0313 | $2.285^{2}$ | 0.3690 | 17 |
| 0.254 | 0.5083 | 0.7625 | I.0166 | 1.2708 | I. 5250 | 1.7791 | 2.0333 | 2.2874 | 0.3694 | 8 |
| 0.254 | 0.5088 | 0.7632 | I.0176 | I. 2721 | 1. 5265 | 1. 7809 | 2.0353 | 2.2897 | 0.3698 | 19 |
| 0.2547 | 0. 5093 | 0.7640 | 1.0186 | 1.2733 | 1. 5280 | 1.7826 | 2.0373 | 2.2919 | 0.3702 | 20 |
| 0. | 0.5098 | 0.7647 | 1.0196 | 1.2746 | I. 5295 | 1. 7844 | 2.0393 | 2.2942 | 0.3706 |  |
| 0.25 | 0.5103 | 0.7655 | 1.0206 | I. 2758 | I. 5310 | 1. 7861 | 2.0413 | 2.2964 | 0.3710 | 22 |
| 0.25 | 0.5108 | 0.7662 | 1.0216 | 1.2771 | I. 5325 | 1. 7879 | 2.0433 | 2.2987 | 0.3714 | 23 |
| 0.25 | 0.5113 | 0.7670 | 1.0226 | 1.2783 | I. 5340 | 1. 7896 | 2.0453 | 2.3009 | 0.3718 | 24 |
| 0.255 | 0.5118 | 0.7677 | 1.0236 | 1.2796 | I. 5355 | 1.7914 | 2.0473 | 2.3032 | 0.3722 | 25 |
| 0.2562 | 0.5123 | 0. 7685 | I. 0246 | I. 2808 | 1.5370 | 1.793I | 2.0493 | 2.3054 | 0.3725 | 26 |
| 0.2564 | 0.5128 | 0.7692 | 1.0256 | I. 2821 | 1.538 | 1. 7949 | 2.0513 | 2.3077 | 0.3729 | 27 |
| 0.256 | 0.5133 | 0.7700 | 1.0266 | 1. 2833 | 1.540 | 1. 7966 | 2.0533 | 2.3099 | 0.3733 | 28 |
| 0.2569 | 0.5138 | 0.770 | 1.0276 | 1. 2845 | I. 5415 | I. 7984 | 2.0553 | 2.3122 | 0.3737 | 29 |
| 0.2572 | 0.5143 | 0.7715 | 1.0286 | 1. 2858 | I. 5430 |  | 2.0573 | 2.3144 | 0.3741 | 30 |
| 0. |  | O. |  |  | I. | 1.8019 | 2. |  | 45 | 31 |
| 0. | 0.5153 | 0.773 |  | 1. 2883 | 1.5460 | 1. 8036 | 2.0613 | 2.3189 | 0.3749 | 32 |
| o. | -. 5 | 0. 7737 |  | 1. 2895 | 1. 5474 | I. 8053 | 2.0632 | 2.3211 | 0.3753 | 33 |
|  |  | 0.7745 |  |  | 1.5489 | 1.8071 | 2.0652 | 2.3234 | 0.3757 | 34 |
|  | 0.5168 | 0.7752 | 1.0336 | I. 29 | I. 5504 | 1. 8088 | 2.0672 | 2.3256 | 0.3761 | 35 |
|  | 0.5173 | 0.7760 | 1.0346 | 1.2933 | I. 5519 | I. 8106 | 2.0692 | 2.3279 | 0.3765 | 36 |
| 0.2589 | 0.5178 | 0.7767 | 1.0356 | I. 2945 | I. 5534 | I. 8123 | 2.0712 | 2.3301 | 0.3769 | 37 |
| 0.2591 | 0.5183 | 0.7774 | I. 0366 | I. 2957 | 1. 5549 | I. 8140 | 2.0732 | 2.3323 | 0.3773 | 38 |
| 0.2594 | 0.5188 | 0. 7782 | 1.0376 | 1.2970 | 1.5564 | I. 8158 | 2.0752 | 2.3346 | 0.3776 | 39 |
| 0.2596 | 0.5193 | 0.7789 | 1.0386 | 1.2982 | 1. 5578 | 1.81 75 | 2.077 I | 2.3368 | $0.37^{80}$ |  |
| 0.2599 | 0.519 | 0.7797 | 1. 0396 | 1. 2995 | I. 5 | 1.8192 | 2.0791 | 2.3390 | 0.3784 | 4 I |
| Q. 2601 | 0.520 | 0.7804 | 1.0406 | 1. 3007 | I. 560 | 1.8210 | 2.0811 | 2.3413 | 0.3788 | 42 |
| 0. | 0.5208 | 0. 781 I | 1.0416 | I. 3019 | I. 56 | I. 8227 | 2.0831 | 2.3435 | 0.3792 | 43 |
| 0.2606 | 0.5213 | 0.7819 | I. 0426 | I. 3032 | I. 563 | I. 8245 | 2.0851 | 2.3458 | 0.3796 | 44 |
| 0.2609 | 0. 5218 | 0.7826 | 1. 0435 | I. 3044 | I. 56 | 1.8262 | 2.0870 | 2.3479 | 0.3800 | 45 |
| 0.26 | 0.5223 | 0. 7834 | 1. 0445 | I. 3057 | I. 566 | 1. 8279 | 2.0890 | 2.3502 | 0.3804 | 46 |
| 0.26 | 0.5228 | 0.7841 | 1. 0455 | I. 3069 | I. 568 | I. 8297 | 2.0910 | 2.3524 | 0.3808 | 47 |
| 0.26 | 0. 5233 | 0.7849 | 1.0465 | I. 308 I | I. 5698 | I.8314 | 2.0930 | 2.3547 | 0.3812 | 48 |
| c. 2619 | 0.5237 | 0.7856 | I. 0475 | I. 3094 | I. 5712 | I. 833 I | 2.0950 | 2.3569 | 0.3816 | 49 |
| 0.2621 | 0.5242 | 0. 7864 | 1.0485 | 1.3106 | I. 5727 | 1. 8348 | 2.0970 | 2.3591 | 0.3820 | 50 |
| 0.262 | 0.5247 | 0.7871 | 1. 0495 | 1.3118 | 1.5742 | 1. 8366 | 2.0590 | 2.3613 | 0.3824 | 51 |
| 0.2626 | 0.5252 | 0. 7879 | 1.0505 | I.3131 | I. 5757. | 1. 8383 | 2. 1010 | 2.3636 | 0.3827 | 52 |
| 0.2629 0.2631 | 0.5257 0.5262 | 0.7886 | 1.0514 r.0524 | I.3143 | I. 5772 | I. 8400 | 2.1029 | 2.3658 2.3680 | 0.3831 0.3835 | 53 |
| 0.2631 | 0.5262 0.5267 | 0.7893 | 1.0524 I. 0534 | I. 3155 <br> I. 5168 | 1.5787 I 5802 | 1.8418 | 2.1049 | 2.3680 2.3702 | 0.3835 | 54 |
|  | 0.5267 | 0.7901 | 1.0534 | 1.3168 | $\begin{array}{r}\text { I. } 5802 \\ \text { r } \\ \hline\end{array}$ | I. 8435 | 2.1069 | 2.3702 | 0.3839 0.3843 | 55 |
| c. 2636 0.2638 | 0.5272 0.5277 | 0.7908 | I. 0544 <br> I. 0554 | 1.3180 | 1.5816 | 1. 8452 | 2.1088 | 2.3724 | 0.3843 | 56 |
| c. 2641 | 0.5277 0.5282 | 0.798 0.7915 0.7923 | 1.0554 1.0564 | 1.3182 I. 3205 | 1.5831 1.5846 | 1.8449 | 2.1128 | 2.3746 2.3769 | 0.38847 0.3851 | 58 |
| 0.2643 | 0. 5287 | -. 7930 | 1.0574 | 1.3217 | 1.5860 | 1.8504 | 2.1147 | 2.3791 | 0.3855 | 59 |
| 0.2646 | 0.5292 | 0.7938 | 1.0584 | I. 3230 | I. 5875 | 1.8521 | 2.1167 | 2.3813 | 0.3859 | 60 |


| 120 |  |  |  |  | DISTANCES. |  |  |  |  | $16^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | a |
| -0 | 0.9 | 1. 8455 | 2.7682 | 3.6909 | 4.6r37 | 5.5364 | 6.4591 | 7.3818 |  | I. 3458 |
| OI | 0.9226 | 1. $845{ }^{2}$ | 2.7677 | 3.6903 | 4.6129 | 5.5355 | 6.4580 | 7.3806 | 8.3032 | I. 3457 |
| 02 | 0.9224 | r. 8448 | 2.7673 | 3.6897 | 4.6121 | $5 \cdot 5345$ | 6.4569 | $7 \cdot 3794$ | 8.3018 | I. 3456 |
| 03 | 0.9223 | I. 8445 | 2.7668 | 3.6891 | 4.6113 | 5.5336 | 6.4559 | 7.378 I | 8.3004 | 1. 3455 |
| 04 | 0.922 | I. 8442 | 2.7663 | 3.6884 | 4.6006 | 5.5327 | 6.4548 | 7.3769 | 8.2990 | I. 3454 |
| $\bigcirc 5$ | 0.9220 | 1. 8439 | 2.7659 | 3.6878 | 4.6098 | 5.5317 | 6.4537 | 7.3757 | 8.2976 | I. 3453 |
| o5 | 0.9218 | т. 8436 | 2.7654 | 3.6872 | 4.6090 | 5.5308 | 6.4526 | $7 \cdot 3744$ | 8.2962 | I. 345 I |
| 07 | 0.9216 | r. 8433 | 2.7649 | 3.6866 | 4.6082 | 5.5299 | 6.4515 | 7.3732 | 8.2948 | I. 3450 |
| 08 | 0.9215 | 1. 8430 | 2.7645 | 3.6860 | 4.6075 | 5.5290 | 6.4505 | 7.3719 | 8.2934 | I. 3449 |
| 09 | 0.9213 | 1. 8427 | 2.7640 | 3.6854 | 4.6067 | 5.5280 | 6.4494 | 7.3707 | 8.2921 | I. 3448 |
| 10 | 0.9212 | 1.8424 | 2.7636 | 3.6847 | 4.6059 | 5.5271 | 6.4483 | 7.3695 | 8.2907 | I. 3447 |
| II | 0.9210 | 1.842I | 2.763 I | 3.6841 | 4.6051 | 5.5262 | 6.4472 | 7.3682 | 8.2893 | 1. 3446 |
| 12 | 0.9209 | 1. 8417 | 2.7626 | 3.6835 | 4.6044 | 5.5252 | 6.4461 | 7.3670 | 8.2878 | I. 3445 |
| 13 | 0.9207 | 1.8414 | 2.7621 | 3.6829 | 4.6036 | 5.5243 | 6.4450 | 7.3657 | 8.2864 | I. 3444 |
| 14 | 0.9206 | 1.84iI | 2.7617 | 3.6822 | 4.6028 | 5.5234 | 6.4439 | $7 \cdot 3645$ | 8.2850 | I. 3442 |
| 15 | 0.9204 | I. 8408 | 2.7612 | 3.6816 | 4.6020 | 5.5224 | 6.4428 | 7.3632 | 8.2836 | I. 3441 |
| 16 | 0.9202 | 1. 8405 | 2.7607 | 3.6810 | 4.6012 | 5.5215 | 6.4417 | 7.3620 | 8.2822 | I. 3440 |
| 17 | 0.9201 | 1. 8402 | 2.7603 | 3.6804 | 4.6005 | 5.5205 | 6.4406 | 7.3607 | 8.2808 | '. 3439 |
| 18 | 0.9199 | 1. 8399 | 2.7598 | 3.6797 | 4.5997 | 5.5196 | 6.4395 | $7 \cdot 3595$ | 8.2794 | I. 3438 |
| 19 | 0.9198 | I. 8395 | 2.7593 | 3.6791 | 4.5989 | 5.5187 | 6.4385 | $7 \cdot 3582$ | 8.2780 | I. 3437 |
| 20 | 0.9196 | 1. 8392 | 2.7589 | 3.6785 | 4.598 I | 5.5177 | 6.4374 | 7.3570 | 8.2766 | I. 3436 |
| 21 | 0.9195 | 1. 8389 | 2.7584 | 3.6779 | 4.5973 | 5.5168 | 6.4363 | 7.3557 | 8.2752 | 1. 3435 |
| 22 | 0.9193 | 1. 8386 | 2.7579 | 3.6772 | 4.5965 | 5.5158 | 6.4352 | 7.3545 | 8.2738 | I. 3433 |
| 23 | 0.9192 | 1. 8383 | 2.7575 | 3.6766 | 4.5958 | 5.5149 | 6.434 I | 7.3532 | 8.2724 | r. 3432 |
| 24 | 0.9190 | 1. 8380 | 2.7570 | 3.6760 | 4.5950 | 5.5140 | 6.4330 | 7.3519 | 8.2709 | I. 3431 |
| 25 | 0.9188 | 1. 8377 | 2.756 | 3.6753 | 4.5942 | 5.5130 | 6.4319 | 7.3507 | 8.2695 | I. 3430 |
| 26 | 0.9187 | I. 8374 | 2.7560 | 3.6747 | 4.5934 | 5.5121 | 6.4307 | 7.3494 | 8.2681 | I. $3+29$ |
| 27 | 0.9185 | 1.8370 | 2.7556 | 3.6741 | 4.5926 | 5.5111 | 6.4296 | 7.3482 | 8.2667 | I. 3428 |
| 28 | 0.9184 | I. 8367 | 2.755 I | 3.6735 | 4.5918 | 5.5102 | 6.4285 | 7.3469 | 8.2653 | 1. 3427 |
| 29 | 0.9182 | I. 8364 | 2.7546 | 3.6728 | 4.5910 | 5.5092 | 6.4274 | 7.3456 | 8.2639 | I. 3425 |
| 30 | 0.9180 | т. 8361 | 2.754 I | 3.6722 | 4.5902 | 5.5083 | 6.4263 | 7.3444 | 8.2624 | I. 3424 |
| 31 | 0.9179 | 1. 8358 | 2.7537 | 3.6716 | 4.5894 | 5.5073 | 6.4252 | 7.343I | 8.2610 | 1. 3423 |
| 32 | 0.9177 | 1. 8355 | 2.7532 | 3.6709 | 4.5887 | $5.505_{4}$ | 6.4241 | 7.3418 | 8. 2596 | I. 3422 |
| 33 | c.9176 | 1.8351 | 2.7527 | 3.6703 | 4.5879 | 5.5054 | 6.4230 | 7.3406 | 8.2581 | I. 3420 |
| 34 | 0.9174 | I. 8348 | 2.7522 | 3.6696 | 4.587 I | 5.5045 | 6.4219 | 7.3393 | 8.2567 | I. 3419 |
| 35 | 0.9173 | I. 8345 | 2.7518 | 3.6690 | 4.5863 | 5.5035 | 6.4208 | 7.3380 | 8.2553 | I. $3+18$ |
| 36 | 0.9171 | I. 8342 | 2.7513 | 3.6684 | 4.5855 | 5.5026 | 6.4197 | 7.3368 | 8.2539 | I. 3417 |
| 37. | 0.9169 | I. 8339 | 2.7508 | 3.6677 | 4.5847 | 5.5016 | 6.4186 | 7.3355 | 8.2524 | I. 3416 |
| 38 | 0.9168 | I. 8336 | 2.7503 | 3.6671 | 4.5839 | $5 \cdot 5007$ | 6.4174 | 7.3342 | 8.2510 | I. 3415 |
| 39 | 0.9166 | 1.8332 | 2.7499 | 3.6665 | 4.5831 | 5.4997 | 6.4163 | 7.3329 | 8.2496 | I.34I3 |
| 40 | 0.9165 | 1. 8329 | 2.7494 | 3.6658 | 4.5823 | 5.4988 | $6.415^{2}$ | 7.3317 | 8.248I | 1.3412 |
| 4 I | 0.9163 | 1. 8326 | 2.7489 | 3.6652 | 4.5815 | 5.4978 | 6.4141 | 7.3304 | S.2467 | I. 3411 |
| 42 | 0.9161 | I. 8323 | 2.7484 | 3.6646 | 4.5807 | 5.496S | 6.4130 | 7.3291 | 8. 2452 | I.3410 |
| 43 | 0.9160 | I. 8320 | 2.7479 | 3.6639 | 4.5799 | 5.4959 | 6.4118 | $7 \cdot 3278$ | 8.2438 | I. 3409 |
| 44 | 0.9158 | I. $\mathrm{S}_{316}$ | 2. 7475 | 3.6633 | 4.5791 | 5.4949 | 6.4107 | $7 \cdot 3265$ | 8. 2424 | I. 3407 |
| 45 | 0.9157 | 1. 8313 | 2.7470 | 3.6626 | 4.5783 | 5.4939 | 6.4096 | 7.3253 | 8.2409 | I. 3406 |
| 46 | 0.9155 | 1.8310 | 2.7465 | 3.6620 | 4.5775 | 5.4930 | 6.4085 | 7.3240 | 8.2395 | I. 3405 |
| 47 | 0.9153 | 1. 8307 | 2.7460 | 3.6613 | 4.5767 | 5.4920 | 0.4074 | 7.3227 | 8.2380 | I. 3104 |
| 48 | -.9152 | 1. 8304 | 2.7455 | 3.6607 | 4.5759 | 5.4911 | 6.4062 | 7.3214 | 8.2366 | I. 3403 |
| 49 | 0.9150 | 1. 8300 | 2.7450 | 3.6601 | 4.575 I | 5.4901 | 6.4051 | 7.3201 | 8.2351 | 1.3402 |
| 50 | 0.9149 | 1. 8297 | 2.7446 | 3.6594 | 4.5743 | 5.4891 | 6.4040 | $7 \cdot 3188$ | 8.2337 | 1. 3400 |
| 51 | 0.9147 | I. S 294 | 2.744 I | 3.6588 | 4.5735 | 5.4882 | 6.4029 | 7.3176 | 8.2322 | 1. 3399 |
| 52 | 0.9145 | I. S 291 | 2.7436 | 3.6581 | 4.5727 | 5.4872 | 6.4017 | $7 \cdot 3163$ | 8.2308 | I. 3398 |
| 53 | 0.9144 | 1.8287 | 2.743I | 3.6575 | 4.5719 | 5.4862 | 6.4006 | $7 \cdot 3150$ | 8.2293 | I. 3397 |
| 54 | 0.9142 | I. 8284 | 2.7426 | 3.6568 | 4.5710 | 5.4853 | 6.3995 | 7.3137 | 8.2279 | I. 3395 |
| 55 | 0.9140 | I. S 28 I | 2. 742 I | 3.6562 | 4.5702 | 5.4843 | 6.3983 | 7.3124 | 8.2264 | I. 3394 |
| 56 | 0.9139 | 1. 8278 | 2.7417 | 3.6555 | 4.5694 | 5.4833 | 6.3972 | $7 \cdot 3111$ | 8.2250 | 1. 3393 |
| 57 | 0.9137 | 1.8274 | 2.7412 | 3.6549 | 4.5686 | $5 \cdot 4823$ | 6.3961 | 7.3098 | 8.2235 | I. 3392 |
| 58 | 0.9136 | '1.8271 | 2.7407 | 3.6542 | $4 \cdot 5678$ | 5.4814 | 6.3949 | 7.3055 | 8.2221 | 1.3390 |
| 59 | 0.9134 | I. 8268 | 2.7402 | 3.6536 | 4.5670 | 5.4804 | 6.3938 | $7 \cdot 3072$ | 8.2206 | 1.3389 |
| 60 | 0.9132 | I. 8265 | 2.7397 | $3.653^{\circ}$ | 4.5662 | 5.4794 | 6.3927 | $7 \cdot 3059$ | 8. 2192 | 1.33 |



| 122 |  |  |  | DISTANCES. |  |  |  |  |  | $17^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | a |
| OO | 0.9132 | 1. 8265 | 2.7397 | 3.6530 | 4.5662 | 5.4 | 6.3927 | $7 \cdot 3059$ | 8.2192 | 1.3388 |
| or | 0.9131 | 1. 8262 | 2.7392 | 3.6523 | $4 \cdot 5654$ | 5.4785 | 6.3915 | $7 \cdot 3046$ | 8.2177 | I. 3387 |
| 02 | 0.9129 | 1. 8258 | 2.7387 | 3.6517 | 4.5646 | $5 \cdot 4775$ | 6.3904 | 7.3033 | 8.2162 | 1. 3385 |
| 03 | 0.9127 | 1. 8255 | 2.7382 | 3.6510 | 4.5637 | 5.4765 | 6.3892 | $7 \cdot 3020$ | 8.2147 | 1. 3384 |
| 04 | 0.9126 | 1.8252 | 2.7378 | 3.6503 | $4 \cdot 5629$ | $5 \cdot 4755$ | 6.388 I | $7 \cdot 3007$ | 8.2133 | I. 3383 |
| 05 | 0.9124 | I. 8248 | 2.7373 | 3.6497 | 4.5621 | $5 \cdot 4745$ | $6 \cdot 3870$ | 7.2994 | 8.2118 | I. 3382 |
| 06 | 0.9123 | I. 8245 | 2.7368 | 3.6490 | $4 \cdot 5613$ | $5 \cdot 4736$ | $6 \cdot 3858$ | 7.2981 | 8.2103 | I.338I |
| 07 | 0.9121 | 1. 8242 | 2.7363 | 3.6484 | 4.5605 | 5.4726 | 6.3847 | 7.2968 | 8.2089 | I. 3379 |
| 08 | -.9119 | 1.8239 | $2.735^{8}$ | 3.6477 | 4.5597 | 5.4716 | 6.3835 | 7.2955 | 8.2074 | I. 3378 |
| 09 | 0.9118 | I. 8235 | 2.7353 | 3.6471 | 4.5589 | 5.4706 | 6.3824 | 7.2942 | 8.2059 | r. 3377 |
| 10 | 0.9116 | 1.8232 | 2.7348 | 3.6464 | 4.5580 | $5 \cdot 4696$ | 6.3812 | 7.2929 | 8.2045 | 1.3376 |
| II | 0.9 | 1. 8229 | 2.7343 | $3.645^{8}$ | 4.5572 | 5.4687 | 6.3801 | 7.2915 | 8.2030 | 1. 3375 |
| 12 | 0.9113 | 1. 8226 | 2.7338 | 3.6451 | 4.5564 | 5.4677 | 6.3789 | 7.2502 | 8.2015 | I. 3373 |
| 13 | 0.9111 | 1.8222 | 2.7333 | 3.6445 | 4.5556 | 5.4667 | 6.3778 | 7.2889 | 8.2000 | I. 3372 |
| 14 | 0.9109 | 1.8219 | 2.7328 | 3.6438 | 4.5547 | 5.4657 | 6.3766 | 7.2876 | 8.1985 | 1.3371 |
| 15 | 0.9108 | 1.8216 | 2.7324 | 3.6431 | 4.5539 | 5.4647 | 6.3755 | 7.2863 | 8.1971 | 1.3370 |
| 16 | 0.9106 | 1.8212 | 2.7319 | 3.6425 | 4.5531 | 5.4637 | 6.3743 | 7.2850 | 8.1956 | I. 3368 |
| 17 | 0.9105 | 1.8209 | 2.7314 | 3.6418 | 4.5523 | 5.4627 | 6.3732 | 7.2836 | 8.1941 | x. 3367 |
| 18 | 0.9103 | 1. 8206 | 2.7309 | 3.6412 | 4.5515 | 5.4617 | 6.3720 | 7.2823 | 8.1926 | I. 3366 |
| 19 | 0.9101 | 1.8203 | 2.7304 | 3.6405 | 4.5505 | 5.4608 | 6.3709 | 7.2810 | 8.1911 | I. 3365 |
| 20 | 0.9100 | I. 8199 | 2.7299 | 3.6398 | 4.5498 | 5.4598 | 6.3697 | 7.2797 | 8.1897 | 1. 3364 |
| 21 | 0.9098 | 1.8196 | 2.7294 | 3.6392 | 4.5490 | 5.4588 | 6.3686 | 7.2784 | 8.1882 | 1. 3362 |
| 22 | 0.9096 | 1.8193 | 2.7289 | 3.6385 | 4.5482 | $5 \cdot 4578$ | 6. 3674 | 7.2770 | 8.1867 | 1.3361 |
| 23 | 0.9095 | 1.8189 | 2.7284 | 3.6379 | 4.5473 | 5.4568 | 6.3662 | 7.2757 | 8.1852 | 1. 3360 |
| 24 | 0.9093 | I. 8186 | 2.7279 | 3.6372 | 4.5465 | 5.4558 | 6. 3651 | 7.2744 | 8.1837 | 1. 3359 |
| 25 | 0.g091 | 1.8183 | 2.7274 | 3.6365 | 4.5457 | 5.4548 | 6.3639 | 7.2731 | 8.1822 | 1.3358 |
| 26 | 0.9090 | 1.8179 | 2.7269 | 3.6359 | 4.5448 | 5.4538 | 6. 3628 | 7.2717 | 8.1807 | I. 3357 |
| 27 | 0.9088 | 1.8176 | 2.7264 | $3.635^{2}$ | 4.5440 | 5.4528 | 6.3616 | 7.2704 | 8.1792 | 1.3355 |
| 28 | 0.9086 | I. 8173 | 2.7259 | 3.6345 | 4.5432 | 5-4518 | 6.3604 | 7.2691 | 8.1777 | I. 3354 |
| 29 | 0.9085 | I. 8169 | 2.7254 | 3.6339 | 4.5423 | 5.4508 | 6.3593 | 7.2678 | 8.1762 | 1. 3353 |
| 30 | 0.9083 | 1.8166 | 2.7249 | 3.6332 | 4.5415 | $5 \cdot 4498$ | 6.358 I | 7.2664 | S. 1747 | 1.3352 |
| 31 | 0.908 I | 1.8163 | 2.7244 | 3.6325 | 4.540 | 5.4488 | 6.3570 | 7.2651 | 8.1732 | 1.3350 |
| 32 | 0.9080 | I. 8159 | 2.7239 | 3.6319 | 4.5398 | $5 \cdot 4478$ | 6.3558 | 7.2637 | 8. 1717 | I. 3349 |
| 33 | 0.9078 | 1.8156 | 2.7234 | 3.6312 | 4.5390 | $5 \cdot 4468$ | 6.3546 | 7.2624 | 8. 1702 | 1. 3348 |
| 34 | 0.9076 | 1.8153 | 2.7229 | 3.6305 | 4.5382 | $5 \cdot 4458$ | 6.3534 | 7.2611 | 8.1687 | 1. 3347 |
| 35 | 0.9075 | 1.8149 | 2.7224 | 3.6299 | 4.5373 | 5.4448 | 6.3523 | 7.2597 | 8.1672 | 1. 3346 |
| 36 | 0.9073 | 1.8146 | 2.7219 | 3.6292 | 4.5365 | 5.4438 | 6.351 I | 7.2584 | 8.1657 | I. 3344 |
| 37 | 0.9071 | 1.8143 | 2.7214 | 3.6285 | 4.5357 | 5.4428 | 6.3499 | 7.2571 | 8.1642 | I. 3343 |
| 38 | 0.9070 | 1.8139 | 2.7209 | 3.6279 | 4.5348 | 5-4418 | 6.3488 | 7.2557 | 8. 1627 | 1. 3342 |
| 39 | 0.9068 | 1.8136 | 2.7204 | 3.6272 | 4.5340 | 5.4408 | 6.3476 | 7.2544 | 8.1612 | I. 3341 |
| 40 | 0.go56 | 1.8I33 | 2.7199 | 3.6265 | 4.5332 | 5.4398 | 6.3464 | 7.2530 | 8. 1597 | 1. 3339 |
| 4 I | 0.9065 | 1.8129 | 2.7194 | 3.6258 | 4.5323 | $5 \cdot 4388$ | 6.3452 | 7.2517 | S. 1581 | 1. 3338 |
| 42 | 0.9063 | 1.8126 | 2.7189 | 3.6252 | 4.5315 | $5 \cdot 43{ }^{8}$ | 6.344 I | 7.2503 | 8. 1566 | 1. 3337 |
| 43 | 0.9061 | 1.8122 | 2.7184 | 3.6245 | 4.5306 | 5.4367 | 6.3429 | 7.2450 | 8.1551 | I. 3336 |
| 44 | 0.9060 | 1.8119 | 2.7179 | 3.6238 | 4.5298 | 5.4357 | 6.3417 | 7.2476 | 8. 1536 | I. 3335 |
| 45 | 0.9058 | 1.8ir6 | 2.7174 | 3.6231 | 4.5289 | $5 \cdot 4347$ | 6.3405 | 7.2463 | 8.152I | I. 3333 |
| 46 | 0.9056 | 1.8112 | 2.7169 | 3.6225 | 4.5281 | $5 \cdot 4337$ | 6.3393 | 7.2449 | 8. 1505 | I. 3332 |
| 47 | 0.9054 | 1.8109 | 2.7163 | 3.6218 | 4.5272 | 5.4327 | 6.338 I | 7.2436 | 8.1450 | 1.3331 |
| 48 | 0.9053 | 1.8105 | $2.715^{8}$ | 3.6211 | 4.5264 | $5 \cdot 4317$ | 6.3370 | 7.2422 | 8.1475 | 1.3330 |
| 49 | 0.905 | 1.8102 | 2.7153 | 3.6204 | 4.5256 | 5.4307 | $6.335{ }^{\text {8 }}$ | 7.2409 | 8.1460 | 1.3329 |
| 50 | 0.9049 | 1.8099 | 2.7148 | 3.6198 | 4.5247 | $5 \cdot 4297$ | 6.3346 | 7.2395 | 8.1445 | 1.3327 |
| 51 | 0.9048 | 1.8095 | 2.7143 | 3.6191 | 4.5239 | 5.42S6 | 6.3334 | 7.2382 | 8. 1430 | r. 3326 |
| 52 | 0.9046 | 1.8092 | 2.7138 | 3.6184 | 4.5230 | 5.4276 | 6.3322 | 7.2368 | 8.1414 | 1.3325 |
| 53 | 0.9044 | 1.8089 | 2.7133 | 3.6177 | 4.5222 | 5.4266 | 6.3310 | 7.2355 | 8.1399 | I. 3324 |
| 54 | 0.9043 | 1.8085 | 2.7128 | 3.6171 | 4.5213 | 5.4256 | 6.3298 | 7.234 T | 8.1384 | 1.3323 |
| 55 | 0.9041 | 1. 8082 | 2.7123 | 3.6164 | 4.5205 | $5 \cdot 4246$ | 6.3287 | 7.2327 | 8. 1368 | 1.3321 |
| 56 | 0.9039 | 1. 8078 | 2.7118 | 3.6157 | 4.5196 | 5.4235 | 6.3275 | 7.2314 | 8. 1353 | 1.3320 |
| 57 | 0.9038 | 1. 8075 | 2.7113 | 3.6150 | 4.5188 | 5.4225 | 6.3263 | 7.2300 | 8. 1338 | 1.3319 |
| 58 | 0.9036 | 1. 8072 | 2.7107 | 3.6143 | 4.5179 | 5.4215 | 6.3251 | 7.2287 | S. 1322 | 1.3318 |
| 59 | 0.9034 | I. 8068 | 2.7102 | 3.6137 | 4.5171 | 5.4205 | 6.3239 | 7.2273 | 8. 1307 | 1.3316 |
| 60 | 0.9032 | I. 8065 | 2.7097 | 3.6133 | 4.5162 | 5.4195 | 6.3227 | 7.2259 | S. 1292 | 1.3315 |


| $17^{\circ}$ | HEIGHTS. |  |  |  |  |  |  |  | 123 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | b |  |
|  | 0.5 |  |  | 1.39 |  | 1.9544 | 7 | 9 | 3 | -0 |
| 0.2794 | 0.5589 | 0.8383 | 1.1178 | I. 3972 |  | 1.9561 | 2.2356 | 2.5150 | $0.4 C 97$ | 1 |
| 0.2797 | 0. 5594 | 0.8391 | 1.1187 | 1. 3984 | 1.6781 | 1. 9578 | 2.2375 | 2.5172 | 0.4101 | 02 |
| 0.2799 | 0.5599 | 0. 8398 | r.1197 | 1. 3996 | 1. 6796 | 1. 9595 | 2.2394 | 2.5193 | 0.4105 | 3 |
| 0.2802 | 0.5603 | 0.8405 | I. 1207 | 1. 4008 | 1.6810 | 1.9612 | 2.2414 | 2.5215 | 0.4IC9 | 4 |
| 0.2804 | 0. 5608 | 0. 8412 | 1. 1216 | 1.4020 | I. 6825 | 1. 9629 | 2.2433 | 2.5237 | 0.4113 |  |
| 0.2806 | 0.5613 | 0. 8419 |  | I. 4032 | 1. 6839 | 1.9645 | 2.2452 | 2.5258 | 0.4116 |  |
| 0.2809 | 0.5618 | 0.8427 | I. 1236 | 1. 4044 | 1.6853 | 1. 9662 | 2.2471 | 2.5280 | 0.4120 | 07 |
| 0.2811 | 0. 5623 | 0. 8434 | I. 1245 | I. 4056 | 1.6868 | 1. 9679 | 2.2490 | 2.5302 | 0.4124 | 08 |
| 0.2814 | 0. 5627 | 0.8441 | I. 1255 | I. 4068 | 1.6882 | 1.9696 | 2.2510 | 2.5323 | 0.4128 | 09 |
| 0.2816 | 0.5632 | 0.8448 | 1.1264 | 1.4080 | 1. 6897 | 1.9713 | 2.2529 | 2.5345 | 0.4132 | 10 |
| 0.28 |  |  |  |  |  |  |  | 2.53 |  | 11 |
| 0.2821 | 0.5642 | 0. 8463 | 1.1284 | 1.4104 | r.6925 | 1.9746 | 2.2567 | 2.53 | 0.4140 | 12 |
| 0.2823 | 0.5647 | 0.8470 | I. 1293 | I. 411 | 1. 6940 | 1.9763 | 2.2586 | 2.5409 | 0.4144 | 13 |
| 0.2826 | 0.5651 | 0.8477 | I. 1303 | I.4128 | I. 6954 | 1.9780 | 2.2606 | 2.5431 | 0.4148 | 14 |
| 0.2828 | 0.5656 | 0. 8484 | I. 1312 | I.4140 | 1.6969 | I. 9797 | 2.2625 | 2.5453 | 0.4151 | 5 |
| 0.2830 | 0. 5661 | 0.8491 | I. 1322 | 1.4152 | 1.6983 | 1.9813 | 2.2644 | 2.5474 | 0.4155 | 6 |
| 0.2833 | 0. 5666 | 0.8499 | I. 1332 | I. 4164 | 1. 6997 | 1.9830 | 2.2663 | 2.5496 | 0.4159 | 17 |
| 0.2835 | 0. 5670 | 0.8506 | I. 134 I | I. 4176 | 1.7011 | 1.9847 | 2.2682 | 2.5517 | 0.4163 | 18 |
| 0.2838 | 0. 5675 | 0.8513 | 1.1351 | I. 4188 | I. 7026 | 1.9863 | 2.2701 | 2.5539 | 0.4167 | 19 |
| 0.2840 | 0. 5680 | 0.8520 | I. 1360 | 1. 4200 | 1.7040 | 1.9880 | 2.2720 | 2.5560 | 0.4171 | 20 |
| 0.2842 | 0.5685 | 0.8527 |  | 1.4212 |  | 1.9097 |  |  | . 4175 | 1 |
| 0.2845 | 0.5690 | 0.8534 | 1.1379 | 1. 4224 | 1.7069 | 1.9914 | 2.2758 | 2.56 c 3 | 0.4179 | 22 |
| 0.2847 | 0. 5694 | 0.8542 | 1.1389 | 1. 4236 | r. 7083 | I.9930 | 2.2778 | 2.5625 | 0.4183 | 23 |
| 0.2850 | 0.5699 | 0.8549 | I. 1398 | 1. 4248 | 1.7098 | 1.9947 | 2.2797 | 2.5646 | 0.4186 | 24 |
| 0.2852 | 0. 5704 | 0.8556 | 1.1408 | I. 426 | 1.7112 | 1. 9964 | 2.2816 | 2.5668 | 0.4190 | 25 |
| 0.2854 | 0.5709 | 0. 8563 | I. 1417 | 1.4272 | 1.7126 | 1.9980 | 2.2834 | 2.5689 | 0.4194 | 26 |
| 0.2857 | 0.5713 | 0.8570 | I. 1427 | I. 4284 | I. 7140 | 1.9597 | 2.2854 | 2.5710 | 0.4198 | 27 |
| 0.2859 | 0.5718 | 0.8577 | I. 1436 | I. 4296 | 1. 7155 | 2.0014 | 2.2873 | 2.5732 | . 4202 | 28 |
| 0.2861 | 0.5723 | 0.8584 | I. 1446 | I. 4307 | 1.7169 | 2.0030 | 2.2892 | 2.5753 | 0.4208 | 29 |
| 0.2864 | 0.5728 | 0.8592 | I. 1456 | 1.4319 | 1.7183 | 2.6047 | 2.2911 | 2.5775 | 0.4210 | 30 |
| 0.2866 | 0.5732 |  | 1. 1465 | 31 | 1. | 2.co63 | 2.2930 | 2. | 4 | 31 |
| 0.2869 | 0.5737 | 0.8606 | 1. 1474 | 4343 | 1. 7212 | 2.00 | 2.29 | 2.5818 | 4217 | 32 |
| 0.287 | 0. 5742 | 0.8613 | I. 1484 | 335 | 1. 7226 | 2.00 | 2.2968 | 2.5839 | 4221 | 33 |
| 0.287 | 0.5747 | 0.8620 | I. 1494 | I. 436 | I. | 2.0114 | 2.2987 | 2.5861 | 4225 | 34 |
| 0.2876 | 0.5752 | 0.8627 | I. 1503 | 1. 4379 | I. | 2.01 | 2.3006 | 2.5882 | 4229 | 35 |
| 0.2878 | 0. 5756 | 0.8634 | 1.1512 | I. 4390 | 1.7269 | 2.01 | 2.3025 | 2.5903 | 33 | $3^{6}$ |
| 0.2880 | 0.5761 | 0.8641 | I. 1522 |  | 1. 7283 | 2.0163 |  | 2.5924 | 0.4237 | 37 |
| 0.2883 | 0. 5766 | 0.8649 | I. 1532 |  | 1. 7297 |  | 2.3063 | 2.5946 | 0.4241 | 38 |
| 0.2885 | 0.5770 | 0.8656 | I. 1541 |  |  | 2.0196 | 2.3082 | 2.5967 | 0.4245 | 39 |
| 0.2888 | 0.5775 | 0.8663 | I. 1 | 1.4438 | 1.7326 | , | 2.3101 | 2.5988 | 0.4249 | 40 |
| 0.28 go | 0.5 |  |  |  | 1.7340 |  |  | 2.Єoro |  | 41 |
| 0.2892 | 0.5785 |  | I.1569 | 1. 4462 | 1.7354 | 2.0246 | 2.3139 | 2.6031 | 0.4256 | 42 |
| 0.2895 | 0.5789 | c. 8684 | I. 579 | I. 4474 | 1. 7368 | 2.0263 | 2.3158 | 2.6052 | 0.4260 | 43 |
| 0.2897 | 0.5794 | 0.8691 | 1. 1588 | 1. 4485 | 1. 7383 | 2.02 | 2.3177 | 2.6074 | 0. 4264 | 44 |
| 0.2899 | 0.5799 | 0.8698 | 1.1598 | I. 4497 | I. 7397 | 2.0296 | 2.3196 | 2.6095 | 0.4268 | 45 |
| 0.2902 | 0.5804 | 0.8705 | I. 1607 | 1.4509 | I. 7411 | 2.0313 | 2.3215 | 2.6116 | 0.4272 | 46 |
| 0.2904 | 0.580 | 0.8713 | I.1617 | 1.4521 | 1. 7425 | 2.0329 | 2.3233 | 2.6138 | 0.4276 | 47 |
| 0.2907 | 0.5 | 0.8720 | 1.1 | 1.4533 | 1. 7439 | 2.0345 | 2.3252 | 2.6159 | 0.4280 | 48 |
| 0.2909 | 0.5818 | 0.8727 | 1.1636 | I. 45 | 1.7453 | 2.0362 | 2.3271 | 2.6180 | 0.4283 | 49 |
| 0.2911 | 0.5823 | 0.8734 | 1.1645 | 1.4556 | 1.7468 | 2.0379 | 2.3290 | 2.6202 | 0.4287 | 50 |
| 0.291 | 0.5827 | 0.8741 | 1.1654 | 1. 456 | 1.7482 | 2.0395 | 2.3309 | 2.6223 | 0.4291 | 51 |
| 0.2916 | 0.5832 | 0.8748 | 1.1664 | I. 458 | 1. 7496 | 2.0412 | 2.3328 | 2.6244 | 0.4295 | 52 |
| 0.2918 | 0.5837 | 0.8755 | 1. 1673 | 1.4591 | 1.7510 | 2.0428 | 2.3346 | 2.6265 | 0.4299 | 53 |
| 0.2921 0.2923 | 0.5841 0.5846 | 0.8762 | 1. 1683 | 1. 4603 | 1.7524 | 2.0445 | 2.3365 | 2.6286 | 0.4303 | 54 |
| 0.2923 0.2925 | 0.5846 0.5851 | 0.8769 0.8776 | I. 1692 I. 1702 | I. 4615 I. 4627 | 1.7538 | 2.0461 | 2.3384 | 2.6307 2.6320 | 0.4307 0.4311 | 55 |
| 0.292 | 0.5851 0.5856 | 0.8776 0.8783 | 1.1702 I.1711 I. | 1. 4627 r 4639 | 1.7552 | 2.0478 | 2.3403 | 2.6329 | 0.4311 | 56 |
| 0.2928 0.2930 | 0.5856 0.5860 | 0.8783 0.8790 | 1.1711 1.1720 | 1.4639 1.4651 | 1.7567 1.7581 r | 2.0495 | 2.3422 2.3441 | 2.6350 2.637 I | 0.4315 0.4318 | 57 |
| 0.2932 | 0.5865 | 0.8797 | r.1730 | 1. 4662 | r. 7595 | 2.0527 | 2.3460 | 2.6392 | 0.4322 | 59 |
| 0.2935 | 0.5870 | 0.8804 | 1. 1739 | 1.4674 | 1.7609 | 2.0544 | 2.3478 | 2.6413 | 0.4326 | 60 |


| 124 |  |  |  |  | DISTANCES. |  |  |  |  | $18^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | a |
| 00 | 0.9032 | 1. 8065 | 2.7097 | 3.6130 | 4.5162 | 5.4195 | 6.3227 | 7.2259 | 8.1292 | I. 3315 |
| OI | 0.9031 | 1. 806 I | 2.7092 | 3.6123 | 4.5154 | 5.4184 | 6.3215 | 7.2246 | 8.1276 | I.33I4 |
| 02 | 0.9029 | 1.8058 | 2.7087 | 3.6116 | 4.5145 | 5.4174 | 6.3203 | 7.2232 | 8.126I | I.3312 |
| 03 | 0.9027 | I. 8055 | 2.7082 | 3.6109 | 4.5136 | $5 \cdot 4164$ | 6.3191 | 7.2218 | 8.1246 | I.33II |
| 04 | 0.9026 | 1.8051 | 2.7077 | 3.6102 | 4.5128 | 5.4153 | 6.3179 | 7.2205 | 8.1230 | I.3310 |
| 05 | 0.9024 | I. 8048 | 2. 7072 | 3.6095 | 4.5119 | 5.4143 | 6.3167 | 7.2191 | 8.1215 | I. 3309 |
| 06 | 0.9022 | 1. 8044 | 2.7066 | 3.6089 | 4.5111 | $5 \cdot 4133$ | 6.3155 | 7.2177 | 8.1199 | I. 3308 |
| 07 | 0.9020 | 1.8041 | 2.7061 | 3.6082 | 4.5102 | $5 \cdot 4123$ | 6.3143 | 7.2163 | 8.1184 | 1.3306 |
| 08 | 0.9019 | I. 8037 | 2.7056 | 3.6075 | 4.5094 | 5.4112 | 6.3131 | 7.2150 | 8.1168 | I. 3305 |
| 09 | 0.9017 | I. 8034 | 2.7051 | 3.6068 | 4.5085 | 5.4102 | 6.3119 | 7.2136 | 8.1153 | I. 3304 |
| 10 | 0.9015 | 1.8031 | 2.7046 | 3.695I | 4.5076 | 5.4092 | 6.3107 | 7.2122 | 8.1138 | I. 3302 |
| II | 0.9014 | 1.8027 | 2.7041 | 3.6054 | 4.5068 | $5 \cdot 4081$ | 6.3095 | 7.2108 | 8.1122 | I.3301 |
| 12 | 0.9012 | 1.8024 | 2.7035 | 3.6047 | 4.5059 | 5-4071 | 6.3083 | 7.2095 | 8.1106 | I. 3300 |
| $\mathrm{I}_{3}$ | 0.9010 | 1.8020 | 2.7030 | 3.6040 | 4.5050 | 5.406I | 6.3071 | 7.2081 | 8.1091 | 1.3298 |
| 14 | 0.9008 | I. 8017 | 2.7025 | 3.6033 | 4.5042 | 5.4050 | 6. 3059 | 7.2067 | 8.1075 | I. 3297 |
| ${ }^{1} 5$ | 0.9007 | 1.8013 | 2.7020 | 3.6027 | 4.5033 | 5.4040 | 6.3046 | 7.2053 | 8.1060 | I. 3296 |
| 16 | 0.9005 | 1.8010 | 2.7015 | 3.6020 | 4.5025 | 5.4029 | 6.3034 | 7.2039 | 8.1044 | I. 3294 |
| ${ }^{1} 7$ | 0.9003 | 1. 8006 | 2.7010 | 3.6013 | 4.5016 | 5.4019 | 6.3022 | 7.2025 | 8.1029 | I. 3293 |
| 18 | 0.9001 | 1.8003 | 2.7004 | 3.6006 | $4 \cdot 5007$ | $5 \cdot 4009$ | 6.3010 | 7.2012 | 8.1013 | I. 3292 |
| 19 | 0.9003 | 1. 7999 | 2.6999 | 3.5999 | 4.4999 | 5.3998 | 6.2998 | 7.1998 | 8.0998 | I. 3291 |
| 20 | 0.8998 | 1. 7996 | 2.6994 | 3.5992 | 4.4990 | $5 \cdot 3988$ | 6.2986 | 7.1984 | 8.0982 | I. 3289 |
| 21 | 0.8996 | 1. 7993 | 2.6989 | 3.5985 | 4.4981 | 5.3978 | 6.2974 | 7.1970 | 8.0966 | I. 3288 |
| 22 | 0.8995 | 1.7989 | 2.6984 | 3.5978 | 4.4973 | $5 \cdot 3967$ | 6.2962 | 7.1956 | 8.0951 | I. 3287 |
| 23 | 0.8993 | I. 7986 | 2.6978 | 3.5971 | 4.4964 | 5.3957 | 6.2949 | 7.1942 | 8.0935 | I. 3285 |
| 24 | 0.8991 | I. 7982 | 2.6973 | 3.5964 | 4.4955 | 5.3946 | 6.2937 | 7.1928 | 8.0919 | I. 3284 |
| 25 | 0. 8989 | I. 7979 | 2.6968 | 3.5957 | $4 \cdot 4946$ | 5.3936 | 6.2925 | 7.1914 | 8.0904 | I. 3283 |
| 26 | 0. 8988 | 1. 7975 | 2.6963 | 3.5950 | 4.4938 | 5.3925 | 6.2913 | 7.1900 | 8.0888 | 1.3281 |
| 27 | 0.8986 | I. 7972 | 2.6957 | 3.5943 | 4.4929 | 5.3915 | 6.2901 | 7.1886 | 8.0872 | I. 3280 |
| 28 | 0. 8984 | I. 7968 | 2.6952 | 3.5936 | 4.4920 | 5.3934 | 6.2888 | 7.1873 | 8.0857 | I. 3279 |
| 29 | 0.8982 | I. 7965 | 2.6947 | 3.5929 | 4.4912 | $5 \cdot 3894$ | 6.2876 | 7.1859 | 8.0841 | I. 3278 |
| 30 | 0.898I | 1.796I | 2.6942 | 3.5922 | 4.4903 | $5 \cdot 3884$ | 6.2864 | 7.1845 | 8.0825 | 1. 3276 |
| 31 | 0.8979 | 1. 7958 | 2.6937 | 3.5915 | 4.4894 | $5 \cdot 3873$ | 6.2852 | 7.183I | 8.08ı0 | 275 |
| 32 | 0. 8977 | I. 7954 | 2.6931 | 3.5908 | 4.4885 | $5 \cdot 3862$ | 6.2840 | 7.1817 | 8.0794 | I. 3274 |
| 33 | 0. 8975 | I. 7951 | 2.6926 | $3 \cdot 5901$ | 4.4877 | $5 \cdot 3852$ | 6.2827 | 7.1803 | 8.0778 | I. 3272 |
| 34 | 0. 8974 | I. 7947 | 2.6921 | 3.5894 | 4.4868 | $5 \cdot 384 \mathrm{I}$ | 6.2815 | 7.1789 | 8.0762 | 1. 3271 |
| 35 | 0. 8972 | I. 7944 | 2.6915 | 3.5887 | 4.4859 | $5 \cdot 383 \mathrm{I}$ | 6.2803 | 7.1775 | 8.0746 | I. 3269 |
| 36 | 0. 897 | I. 7940 | 2.6910 | 3.5880 | 4.4850 | $5 \cdot 3820$ | 6.2790 | 7.1760 | 8.0731 | I. 3268 |
| 37 | 0.8968 | 1.7937 | 2.6905 | 3.587 | 4.4842 | 5.3810 | 6.2778 | 7.1746 | 8.0715 | 1. 3267 |
| 38 | 0. 8967 | 1.7933 | 2.6900 | $3 \cdot 58$ | 4.4833 | 5.3799 | 6.2766 | 7.1732 | 8.0699 | I. 3265 |
| 39 | 0. 8965 | 1. 7930 | 2.6894 | 3.5859 | 4.4824 | 5.3789 | 6.2754 | 7.1718 | 8.0683 | I. 3264 |
| 40 | 0.8963 | 1.7926 | 2.6889 | 3.5852 | 4.4815 | $5 \cdot 377^{8}$ | 6.2741 | 7.1704 | 8.0567 | I. 3263 |
| 41 | 0.896i | 1. 7923 | 2.6884 | 3.5845 | $4 \cdot 4806$ | 5.3768 | 6.2729 | 7.1690 | 8.0651 | 1. 3252 |
| 42 | 0.8960 | I. 7919 | 2.6879 | 3.5838 | $4 \cdot 4798$ | $5 \cdot 3757$ | 6.2717 | 7.1676 | 8.0636 | I. 3260 |
| 43 | 0.8958 | I. 7915 | 2.6873 | 3.5835 | 4.4789 | $5 \cdot 3746$ | 6.2704 | 7.1662 | 8.0520 | I. 3259 |
| 44 | 0. 8956 | I. 7912 | 2.6868 | $3 \cdot 5824$ | 4.4780 | 5-373 ${ }^{\text {¢ }}$ | 6.2692 | 7.1648 | 8.0604 | I. 3258 |
| 45 | o. 8954 | I. 7908 | 2.6863 | 3.5817 | 4.4771 | 5.3725 | 6.2679 | 7.1634 | 8.0588 | I. 3257 |
| 46 | 0.8952 | 1. 7905 | 2.6857 | 3.5810 | 4.4762 | $5 \cdot 3715$ | 6.2667 | 7.1619 | 8.0572 | I. 3255 |
| 47 | 0.8951 | I. 7901 | 2.6852 | 3.5803 | 4.4753 | 5.3704 | 6.2655 | 7.1605 | 8.0556 | I. 3254 |
| 48 | 0. 8949 | I. 7898 | 2.6847 | 3.5796 | 4.474 | $5 \cdot 3693$ | 6.2642 | 7.1591 | 8.0540 | I. 3253 |
| 49 | 0.8947 | I. 7894 | 2.6841 | 3.5789 | 4.4736 | $5 \cdot 3683$ | 6.2630 | 7.1577 | 8.0524 | I. 3251 |
| 50 | 0.8945 | 1.7891 | 2.6836 | 3.578I | 4.4727 | $5 \cdot 3672$ | 6.2618 | 7.1563 | S.050S | I. 3250 |
| 51 | 0.8944 | 1. 7887 | 2.683 L | 3.5774 | 4.4718 | 5.366I | 6.2605 | 7.1549 | 8.0492 | I. 3249 |
| 52 | 0.8942 | I. 7884 | 2.6825 | 3.5767 | 4.4709 | $5 \cdot 3651$ | 6.2593 | 7.1534 | S.0476 | 1. 3247 |
| 53 | 0.8940 | I. 7880 | 2.6820 | 3.5760 | 4.4700 | $5 \cdot 3640$ | 6.2580 | 7.1520 | S. 0460 | 1. 3246 |
| 54 | -0.8938 | I. 7876 | 2.6815 | 3.5753 | 4.4691 | $5 \cdot 3629$ | 6.2568 | 7.1506 | 8.0+44 | I. 3245 |
| 55 | -. 8936 | I. 7873 | 2.6809 | 3.5746 | 4.4682 | $5 \cdot 3619$ | 6.2555 | 7.1492 | 8.0428 | I. 3243 |
| 56 | -. 8935 | I. 7869 | 2.6804 | 3.5739 | 4.4673 | $5 \cdot 3608$ | 6.2543 | 7.1477 | 8.0412 | I. 3242 |
| 57 | 0. 8933 | 1. 7866 | 2.6799 2.6793 | 3.5732 | 4.4664 | $5 \cdot 3597$ 5 | 6.2530 | 7.1463 | 8.0396 8.0380 | I. 3241 |
| 58 | 0.8931 | 1. 7862 | 2.6793 | 3.5724 | 4.4656 | $5 \cdot 3587$ | 6.2518 | 7.1449 | $8.0380$ | 1. 3239 |
| 5 | 0. 8929 | 1. 7859 | 2.6788 | 3.5717 | 4.4647 | 5.3576 | 6.2505 | 7.1435 | $8.0364$ | $\text { 1. } 3238$ |
| 60 | 0.8928 | 1.7855 | 2.6783 | 3.5710 | 4.4638 | $5 \cdot 3565$ | 6.2493 | 7.1420 | $8.034^{8}$ | I. 3237 |



| 126 | DISTANCES. |  |  |  |  |  |  |  |  | $19^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | a |
| $\bigcirc$ | 0.8928 | 1. 7855 | 2.6783 | 3.5710 | 4.4638 | 5.3565 | 6.2493 | 7.1420 | 8.0348 | 37 |
| OI | 0.8926 | 1.7851 | 2.6777 | 3.5703 | 4.4629 | 5-3554 | 6.2480 | 7.1406 | 8.0332 | I. 3236 |
| 02 | 0.8924 | 1. 7848 | 2.6772 | 3.5696 | 4.4620 | 5.3544 | 6.2468 | 7.1392 | 8.0316 | 1. 3234 |
| 03 | 0.8922 | r. 7844 | 2.6766 | 3.5689 | 4.46 II | $5 \cdot 3533$ | 6.2455 | 7.1377 | 8.0299 | I. 3233 |
| 04 | 0.8920 | 1.7841 | 2.6761 | 3.568 I | 4.4602 | $5 \cdot 3522$ | 6.2443 | 7.1363 | 8.0283 | I. 3232 |
| 05 | 0.8919 | 1. 7837 | 2.6756 | 3.5674 | 4.4593 | 5.3511 | 6.2430 | 7.1349 | 8.0267 | 1. 3230 |
| 06 | 0.8917 | r. 7834 | 2.6750 | 3.5667 | 4.4584 | 5-3501 | 6.2417 | 7.1334 | 8.0251 | 1. 3229 |
| 07 | 0.8915 | r. 7830 | 2.6745 | 3.5660 | 4.4575 | 5.3450 | 6.2405 | 7.1320 | 8.0235 | 1. 3228 |
| 08 | 0.8913 | 1. 7826 | 2.6740 | 3.5653 | 4.4566 | 5.3479 | 6.2392 | 7.1305 | 8.0219 | I. 3226 |
| $\bigcirc$ | 0.8911 | 1. 7823 | 2.6734 | 3.5646 | 4.4557 | 5.3468 | 6.2380 | 7.1291 | 8.0203 | 1. 3225 |
| 10 | 0.8910 | r. 7819 | 2.6729 | 3.5638 | 4.4548 | $5 \cdot 3458$ | 6.2367 | 7.1277 | 8.0186 | 1. 3224 |
| II | 0.8908 | 1.7816 | 2.6723 | 3.563 I | 4.4539 | $5 \cdot 3447$ | 6.2354 | 7.1262 | 8.0170 | 1. 3222 |
| 12 | 0.8906 | 1. 7812 | 2.6718 | 3.5624 | 4.4530 | $5 \cdot 3436$ | 6.2342 | 7.1248 | 8.or 54 | 1. 3221 |
| 13 | 0.8904 | r. 7808 | 2.6712 | 3.5617 | 4.4521 | 5.3425 | 6.2329 | 7.1233 | 8.0137 | r. 3220 |
| 14 | 0.8902 | r. 7805 | 2.6707 | 3.56 c 9 | 4.4512 | 5.3414 | 6.2316 | 7.1219 | 8.0121 | r. 3218 |
| 15 | 0.8901 | r. 7801 | 2.6702 | 3.5602 | 4.4503 | 5.3403 | 6.2304 | 7.1204 | 8.0105 | r. 3217 |
| 16 | 0.8899 | 1. 7797 | 2.6696 | 3.5595 | 4.4494 | 5.3392 | 6.2291 | 7.1190 | 8.0089 | 1. 3215 |
| 17 | 0.8897 | 1. 7794 | 2.6691 | 3.5588 | 4.4485 | $5 \cdot 3382$ | 6.2279 | 7.1175 | 8.0072 | 1. 3214 |
| 18 | 0.8895 | r. 7790 | 2.6685 | 3.5580 | 4.4476 | $5 \cdot 3371$ | 6.2266 | 7.1161 | 8.0056 | I. 3213 |
| 19 | 0.8893 | 1.7787 | 2.6680 | 3.5573 | 4.4467 | 5.3360 | 6.2253 | 7.1146 | 8.0040 | r. 3211 |
| 20 | $0.889^{2}$ | 1.7783 | 2.6675 | $3 \cdot 5566$ |  | $5 \cdot 3349$ | 6.2241 | 7.1132 | 8.0024 | 1.3210 |
| 21 | 0.8890 | 1.7779 | 2.6669 | 3.5559 | 4.4448 | 5.3338 | 6.2228 | 7.1117 | 8.0007 | 1. 3209 |
| 22 | 0.8888 | 1.7776 | 2.6664 | 3.5551 | 4.4439 | $5 \cdot 3327$ | 6.2215 | 7.1103 | 7.9991 | I. 3207 |
| 23 | 0.8886 | 1. 7772 | 2.6658 | 3.5544 | 4.4430 | 5.3316 | 6.2202 | 7.1088 | 7.9974 | 1. 3206 |
| 24 | 0.8884 | 1. 7768 | 2.6653 | 3.5537 | 4.4421 | 5.3305 | 6.2150 | 7.1074 | 7.9558 | 1. 3205 |
| 25 | 0.8882 | 1. 7765 | 2.6647 | 3.5530 | 4.4412 | 5.3294 | 6.2177 | 7.1059 | 7.9942 | 1. 3203 |
| 26 | 0.8881 | 1.7761 | 2.6642 | 3.5522 | 4.4403 | $5 \cdot 3283$ | 6.2164 | 7.1045 | 7.9925 | 1. 3202 |
| 27 | 0.8879 | 1.7758 | 2.6636 | 3.5515 | 4.4394 | $5 \cdot 3273$ | 6.2151 | 7.1030 | 7.9509 | 1.3201 |
| 28 | 0.8877 | 1. 7754 | 2.6631 | 3.5508 | 4.4385 | $5 \cdot 3262$ | 6.2139 | 7.1015 | 7.9892 | r. 3199 |
| 29 | 0.8875 | I. 7750 | 2.6625 | 3.5500 | 4.4376 | $5 \cdot 3251$ | 6.2126 | 7.1001 | 7.9876 | r. 3198 |
| 30 | 0.8873 | 1.7747 | 2.6620 | 3.5493 | 4.4366 | $5 \cdot 3240$ | 6.2113 | 7.0986 | 7.9860 | 1.3197 |
| 31 | 0.8871 | 1.7743 | 2.6614 | 3.5486 | 4.4357 | $5 \cdot 3229$ | 6.2100 | 7.0972 | 7.9843 | 1.3195 |
| 32 | 0.8870 | 1.7739 | 2.6609 | 3.5478 | 4.4348 | $5 \cdot 3218$ | 6.2087 | 7.0957 | 7.9827 | r. 3194 |
| 33 | 0.8868 | 1.7736 | 2.6603 | 3.5471 | 4.4339 | $5 \cdot 3207$ | 6.2075 | 7.0942 | 7.9810 | 1.3193 |
| 34 | 0.8866 | 1. 7732 | 2.6598 | 3.5464 | 4.4330 | $5 \cdot 3196$ | 6.2062 | 7.0928 | 7.9794 | 1.3191 |
| 35 | 0.8864 | I. 7728 | 2.6592 | 3.5456 | 4.4321 | $5 \cdot 3185$ | 6.2049 | 7.0913 | 7.9777 | I. 3190 |
| 36 | 0. 8862 | I. 7725 | 2.6587 | 3.5449 | 4.4311 | 5.3174 | 6.2036 | 7.0898 | 7.9761 | 1. 3189 |
| 37 | 0.8860 | 1.7721 | 2.6581 | 3.5442 | 4.4302 | $5 \cdot 3163$ | 6.2023 | 7.0884 | 7.9744 | 1.3187 |
| 38 | 0.8859 | I.7717 | 2.6576 | 3.5434 | 4.4293 | $5 \cdot 3152$ | 6.2010 | 7.0869 | 7.9727 | 1.3186 |
| 39 | 0.8857 | r. 7714 | 2.6570 | 3.5427 | 4.4284 | 5.3141 | 6.1997 | 7.0854 | 7.9711 | 1.3185 |
| 40 | 0.8855 | 1.7710 | 2.6565 | 3.5420 | 4.4275 | $5 \cdot 3130$ | 6.1985 | 7.0840 | 7.9694 | 1.3183 |
| 4 I | 0.8853 | 1.7706 | 2.6559 | 3.5412 | 4.4265 | $5 \cdot 3119$ | 6.1972 | 7.0825 |  | 1. 3182 |
| 42 | 0.8851 | 1.7702 | 2.6554 | 3.5405 | 4.4256 | $5 \cdot 3107$ | 6.1959 | 7.0810 | 7.9661 | I. 318 I |
| 43 | 0.8849 | 1.7699 | 2.6548 | 3.5398 | 4.4247 | $5 \cdot 3096$ | 6. 1946 | 7.0795 | 7.9645 | I. 3179 |
| 44 | 0.8848 | 1. 7695 | 2.6543 | 3.5390 | 4.4238 | $5 \cdot 3085$ | 6.1933 | 7.0780 | 7.9628 | I. 3178 |
| 45 | 0.8846 | 1.7691 | 2.6537 | 3.5383 | 4.4229 | 5.3074 | 6.1920 | 7.0766 | 7.9611 | 1.3177 |
| 46 | 0. 8844 | 1. 7688 | 2.6532 | 3.5375 | 4.4219 | $5 \cdot 3063$ | 6.1907 | 7.0751 | 7.9595 | I.3175 |
| 47 | 0.8842 | 1. 7684 | 2.6526 | 3.5368 | 4.4210 | $5 \cdot 3052$ | 6.1894 | 7.0736 | 7.9578 | I. 3174 |
| 48 | - 0.8840 | 1.7680 | 2.6520 | $3 \cdot 5361$ | 4.4201 | $5 \cdot 3041$ | 6.1881 | 7.0721 | 7.9561 | I. 3173 |
| 49 | 0.8838 | I. 7677 | 2.6515 | 3.5353 | 4.4192 | 5.3030 | 6.186S | 7.0706 | $7.95+5$ | ז. 3171 |
| 50 | 0.8836 | 1.7673 | 2.6509 | 3.5346 | 4.4182 | 5.3019 | 6.1855 | 7.0692 | 7.9528 | 1.3170 |
| 51 | 0.8835 | 1.7669 | 2.6504 | 3.5338 | 4.4173 | 5.3008 | 6.1842 | 7.0677 | 7.9511 | 1.3169 |
| 52 | 0.8833 | 1. 7665 | 2.6498 | 3.5331 | 4.4164 | 5.2996 | 6.1829 | 7.0662 | 7.9495 | 1.3167 |
| 53 | 0. 8831 | 1. 7662 | 2.6493 | 3.5324 | 4.4154 | 5.2985 | 6.1816 | 7.0647 | 7.9478 | 1.3166 |
| 54 55 | 0. 8829 | 1.7658 | 2.6487 | 3.5316 | 4.4145 | 5.2974 | 6.1803 | 7.0632 | 7.9461 | 1.3165 |
| 55 | 0.8827 | 1.7654 | 2.648 I | 3.5309 | 4.4136 | 5.2963 | 6.1790 | 7.0617 | 7.9444 | I. 3163 |
| 56 | 0.8825 | 1.7651 | 2.6476 | 3.5301 | $4 \cdot 4127$ | 5.2952 | 6.1777 | 7.0602 | 7.9428 | I. 3162 |
| 57 58 5 | 0.8823 0.8822 | 1. 7647 r. 7643 r | 2.6470 2.6165 | 3.5294 3.5286 3.529 | 4.4117 4.4108 | 5.2941 5.2929 | 6.1764 6.1751 | 7.05 S8 | 7.9411 | r. 3160 r. 3150 r 3150 |
| 58 | 0.8822 0.8820 | 1.7643 r. 7639 | 2.6465 | 3.5286 | 4.4108 | 5.2929 | 6.1751 | 7.0573 | 7.9394 | I. 3159 |
| 59 60 | 0.8820 | 1. 26639 r. 7636 | 2.6459 2.6454 | 3.5279 3.5271 | 4.4099 4.4089 | 5.2918 5.2507 | 6.1738 6.1725 | 7.0558 | 7.9377 | I. 3158 I. 3156 |
| 60 | 0.8818 | 1.7636 | 2.6454 | 3.5271 | $4 \cdot 4089$ | 5.2507 | 6.1725 | 7.0543 | 7.9361 | 1.3156 |


| $19^{\circ}$ | HEIGHTS. |  |  |  |  |  |  |  | 127 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | b |  |
|  |  |  | 1.2296 | I. 5370 | 1. 8444 | 2. 1518 | 2.4592 | 6 | 58 | - |
| c. 3076 | 0.6153 | 0.9229 | 1.2305 | 1.5381 | 1. $845^{8}$ | 2.1534 | 2.4610 | 2.7687 | 0.4562 | or |
| 0. 3079 | 0.6157 | 0.9236 | 1. 2314 | I. 5393 | 1. 8472 | 2.1550 | 2.4629 | 2.7707 | 0.4566 | 02 |
| 0.3081 | 0.6162 | 0.9243 | 1. 2324 | I. 5404 | 1. 8485 | 2. 1566 | 2.4647 | 2.7728 | 0.4570 | 03 |
| 0.3083 | 0.6166 | 0.9249 | I. 2333 | I. 5416 | I. 8499 | 2.1582 | 2.4665 | 2.7748 | 0.4573 | 04 |
| -. 3085 | 0.6171 | 0.9256 | 1. 2342 | 1. 5427 | 1. 8512 | 2.1598 | 2.4683 | 2.7769 | 0.4577 | 5 |
| 0. 3088 | 0.6175 | 0.9263 | 1.235 | I. 5439 | 1. 8526 | 2.1614 | 2.4702 | 2.7789 | 0.458 I | 06 |
| 0.3090 | 0.6180 | 0.9270 | 1.2360 | I. 5450 | 1. 8540 | 2.1630 | 2.4720 | 2.7810 | 0.4585 | 7 |
| 0.3092 | 0.6185 | 0.9277 | 1. 2369 | 1.5461 | 1.8554 | 2. 1646 | 2.4738 | 2.7831 | 0.4589 | 08 |
| 0.3095 | 0.6189 | 0.9284 | 1. 2378 | I. 5473 | 1. 8568 | 2.1662 | 2.4757 | 2.7851 | 0.4593 | 09 |
| 0. 3097 | 0.6194 | 0.9290 | 1.2387 | I. 5484 | 1.8581 | 2.1678 | 2.4775 | 2.7871 | 0.4596 | 10 |
|  |  |  |  | I. 5495 |  |  | 2.4793 | 2.7892 |  | 11 |
| 0.3101 | 0.6203 | 0.9304 | 1.2406 | I. 5507 | 1.860 | 2.1710 | 2.48 II | 2.7913 | 0.4604 | 12 |
| 0.3104 | 0.6207 | 0.9311 | 1.2415 | I. 5518 | 1. 8622 | 2.1726 | 2.4830 | 2.7933 | 0.4608 | 3 |
| 0.3106 | 0.6212 | 0.9318 | 1. 2424 | I. 5530 | 1. 8636 | 2.1742 | 2.4848 | 2.7953 | 0.4612 | 14 |
| 0.3108 | 0.6216 | 0.9325 | I. 2433 | I.5541 | 1. 8649 | 2.1758 | 2.4866 | 2.7974 | 0.4616 | 15 |
| 0.3110 | 0.6221 | 0.9331 | I. 2442 | I. 5552 | 1. 8663 | 2. 1773 | 2.4884 | 2.7994 | 0.4619 | 16 |
| 0.3113 | 0.6226 | 0.9338 | I.2451 | I. 5564 | I. 8677 | 2.1789 | 2.4902 | 2.8015 | 0.4623 | 17 |
| 0.3115 | 0.6230 | 0.9345 | 1.2460 | I. 5575 | 1.8690 | 2.1805 | 2.4920 | 2.8035 | 0.4627 | 18 |
| 0.3117 | 0.6235 | 0.9352 | 1. 2469 | I. 5587 | 1.8704 | 2.1821 | 2.4938 | 2.8056 | 0.4631 | 19 |
| 0.3120 | 0.6239 | 0.9359 | 1.2478 | 1.5598 | 1.8718 | 2.1837 | 2.4957 | 2.8076 | 0.4635 | 20 |
| 0.3122 | 0.6244 | 0.9365 |  | 1.5609 | 1.8731 | 2.1853 | 2.4975 |  |  | 21 |
| 0.312 | 0.6248 | 0.9372 | 1.2496 | 1. 5620 | 1.8745 | 2.1869 | 2.4993 | 2.8117 | 0.4642 | 22 |
| 0.312 | 0.6253 | 0.9379 | 1.2505 | 1.5632 | I. 8758 | 2.1885 | 2.5011 | 2.8137 | 0.4646 | 23 |
| 0.3129 | 0.6257 | 0.9386 | 1.2514 | 1. 5643 | 1. 8772 | 2.1900 | 2.5029 | 2.8157 | 0.4650 | 24 |
| 0.3131 | 0.6262 | 0.9393 | 1.2524 | I. 5654 | 1. 8785 | 2.1916 | 2.5047 | 2.8178 | 0.4654 | 25 |
| 0.3133 | 0.6266 | 0.9399 | I. 2533 | I. 5666 | I. 8799 | 2.1932 | 2.5065 | 2.8198 | 0.4658 | 26 |
| 0.3135 | 0.6271 | 0.9406 | 1.2542 | 1. 5677 | 1.8812 | 2. 1948 | 2.5083 | 2.8219 | 0.4662 | 27 |
| $0.313^{8}$ | 0.6275 | 0.9413 | 1.2551 | 1. 5688 | 1.8826 | 2. 1964 | 2.5101 | 2.8239 | 0.4665 | 28 |
| 0.3140 | 0.6280 | 0.9420 | 1.2560 | I. 5700 | 1.8839 | 2.1979 | 2.5119 | 2.8259 | 0.4669 | 29 |
| 0.3142 | 0.6284 | 0.9427 | I. 2569 | I.5711 | I. 8853 | 2. 1995 | $2.513^{8}$ | 2.8280 | 0.4673 | 30 |
| 44 | 0.6289 | 0.943 |  | 5722 | 1.8867 | 2.2 | 2.5156 | 2.8300 |  | 31 |
| 0.3147 | 0.6293 | 0.9440 | 1.2587 | 1. 5734 | 1.8880 | 2.202 | 2.5174 | 2.8320 | 0.468I | 32 |
| 0.3149 | 0.6298 | 0.9447 | 1. 2596 | I. 5745 | 1.8894 | 2.2043 | 2.5192 | 2.8341 | 0.4685 | 33 |
| 0.3151 | 0.6302 | 0.9454 | 1.2605 | 1. 5756 | 1.8907 | 2.2058 | 2.5210 | 2.8361 | 0.4689 | 34 |
| 0.3153 | 0.6307 | 0.9460 | I. 2614 | 1. 5767 | 1.8921 | 2.2074 | 2.5228 | 2.838 I | 0.4692 | 35 |
| 0.3156 | 0.6311 | 0.9467 | I. 2623 | I. 5779 | I. 8934 | 2.2090 | 2.5246 | 2.8401 | 0.4696 | 36 |
| 0.3158 | 0.6316 | 0.9474 | I. 2632 | I. 5790 | 1.8948 | 2.2106 | 2.5264 | 2.8422 | 0.4700 | 37 |
| 0.3160 | 0.6320 | 0.948 I | I. 2641 | 1. 5801 | 1.8961 | 2.21 | 2.5282 | 2.8442 | 0.4704 | 38 |
| 0.3162 | 0.6325 | 0.9487 | I. 2650 | 1.5812 | 1. 8975 | 2.2137 | 2.5300 | 2.8462 | 0.4708 | 39 |
| 0.3165 | 0.6329 | 0.9494 | 1. 2659 | I. 5824 | 1.8988 | 2.2153 | 2.5318 | 2.8482 | 0.4712 |  |
| 0.3167 | 0. 633 | 0.950 | 1.2668 | 1.5835 | 1.9002 | 2.2169 | 2.5336 | 2.8503 | 0.4715 | 4 |
| 0.3169 | 0.6338 | 0.9508 | I. 2677 | I. 5846 | I.9015 | 2.2184 | 2.5354 | 2.8523 | 0.4719 | 42 |
| 0.3171 | 0.6343 | 0.9514 | 1. 2686 | 1. 5857 | 1.9029 | 2.2200 | 2.5372 | 2.8543 | 0.4723 | 43 |
| 0.3174 | 0.6347 | 0.9521 | 1.2695 | 1.5868 | 1.9042 | 2.22 | 2.5390 | 2.8563 | 0.4727 | 44 |
| 0.3176 | 0.6352 | 0.9528 | I. 2704 | 1.5880 | I. 9055 | 2.2231 | 2.5407 | 2.8583 | 0.4731 | 45 |
| 0.3178 | 0.6356 | 0.9535 | 1.2713 | 1.5891 | I.go69 | 2.2247 | 2.5425 | 2.8604 | c. 4735 | 46 |
| 0.3180 | 0.6361 | 0.954 I | I. 2722 | I. 5902 | 1.9082 | 2.2263 | 2.5443 | 2.8624 | 0.4739 | 47 |
| 0.3183 | 0.6365 | 0.9548 | I. 2731 | 1.5913 | 1. 9096 | 2.2279 | 2.5461 | 2.8644 | 0.4742 | 48 |
| 0.3185 | 0.6370 | 0.95 | 1.2740 | 1. 5924 | 1.9509 | 2.2294 | 2.5479 | 2.8664 | 0.4746 | 49 |
| 0.3187 | 0. 6374 | 0.956 I | 1.2748 | I. 5936 | 1.9123 | 2.2310 | 2.5497 | 2.8684 | 0.4750 | 50 |
| 0.3189 | 0.6 | 0.9568 |  | I. 594 | 1.9136 | 2.2326 | 2.5515 | 2.8704 | 0.4754 | 51 |
| 0.3192 | 0.638 | 0.9575 | 1.2766 | I. 5958 | 1.9150 | 2.2341 | 2.5533 | 2.8724 | 0.4758 | 52 |
| 0.319 | 0.6388 | 0.958 I | I. 2775 | 1. 5969 | I.9163 | 2.2357 | 2.5551 | 2.8744 | 0.4761 | 53 |
| 0.3196 | 0.6392 | 0.9588 | 1.2784 | I. 5980 | 1.9177 | 2.2373 | 2.5569 | 2.8765 | 0.4765 | 54 |
| 0.3198 | 0.6397 | 0.9595 | 1. 2793 | I.5991 | I. 9190 | 2.2388 | 2.5586 | 2.8785 | 0.4769 | 55 |
| 0.3201 | 0.640I | 0.9602 | 1.2802 | 1.6co3 | 1.9203 | 2.2404 | 2.5604 | 2.8805 | 0.4773 | 56 |
| 0.3203 | 0.6405 | 0.9608 | 1.28iI | 1.6014 | 1.9216 | 2.2419 | 2. 5622 | 2.8825 | 0.4777 | 57 |
| 0. 3205 | 0.6410 | 0.9615 | 1.2820 | 1. 6025 | I.9230 | 2.2435 | 2.5640 | 2.8845 | 0.4781 | 58 |
| 0.3207 | 0.6414 | 0.9622 | 1.2829 | I. 6036 | 1.9243 | 2.2450 | 2.5658 | 2.8865 | 0.4784 | 59 |
| c.3209 | 0.6419 | 0.9628 | 1.2838 | ı. 6047 | I.c257 | 2.2465 | 2.5675 | 2.8885 | 0.4788 | 60 |

1. 

[^0]:    Case School of Applied Science, Cleteland, Ohio, January, 1887.

[^1]:    * This length was chosen (by Mr. Edward Gunter) because 10 square chains of 66 feet make one acre, and the computation of areas is thus greatly facilitated. For other surveying purposes, particularly for railroad work, a chain of 100 feet is preferable. On the United States Coast and Geodetic Survey the unit of measurement is the French Metre, equal to 3.281 feet nearly.

[^2]:    * To prevent the rery common mistake of calling fortr, sixtr ; or thirty, serentr; it has been suggested to make the 11th, 21st, 31 st, and 41 st links of brass, which would at once show on which side of the middle of the chain was the doubtful mark. This would be particularly useful in Mining Surveying.
    $\dagger$ This must not be confounded with the pieces of wire which have the same name, since one of them is shorter than the "link" used in calculation br half a ring or more, according to the way in which the chain is made.

[^3]:    * The chain used by the Government surveyors of France, which is ten metres, or about half a Gunter's chain in length, is made from one fifth to two fifths of an inch longer than the standard. An inaccuracy of one five-hundredth of its length ( $=1 \frac{1}{2}$ inch on a Gunter's chain) is the utmost allowed not to vitiate the survey.

[^4]:    * Eleven pins are sometimes used, one being of brass. Nine of iron, with four or eight of brass, may also be employed. Their uses are explained in Articles 18 and 19.

[^5]:    * When a chain's length would end in a ditch, pool of water, etc., and the chainmen are afraid of wetting their feet, they can measure part of a chain, to the edge of the water, then stretch the chain across it, and then measure another portion of a chain, so that, with the former portion, it may make up a full chain.

[^6]:    * A horse, on a walk, averages 330 feet per minute, on a trot 650 , and on a common gallop 1,040. For longer times, the difference in horses is more apparent.

[^7]:    * This question is more than two thousand rears old, for Polybius mrites: "Some even of those who are employed in the administration of states, or placed at the head

[^8]:    * To reduce square yards to acres, instead of diriding by 4,840 , it is easier, and very nearly correct, to multiply by 2 , cut off four figures, and add to this product one third of one tenth of itself.

[^9]:    * The French call the narrow opening ceilleton, and the wide one croisée.

[^10]:    * Many of these methods would seldom be required in practice, but cases sometimes occur, as every surveyor of much experience in field-work has found to his serious inconvenience, in which some peculiarity of the local circumstances forbids any of the usual methods being applied. In such cases the collection here given will be found of great value.

    In all the figures, the given and measured lines are drawn with fine full lines, the visual lines, or lines of sight, with broken lincs, and the lines of the result with heavy full lines. The points which are centers around which the chain is swung are inclosed in circles. The alphabetical order of the letters attached to the points shows in what order they are taken.

[^11]:    * This word, like many others used in enginecring, is derived from a French word, borner, to work out or limit, indicating that the Normans introduced the art of surveying into England.
    $\dagger$ Slightly modified from the French alignement.

[^12]:    * A plane is said to be horizontal or level when it is parallel to the surface of standing water, or perpendicular to a plumb-line. A line is horizontal when it lies in a horizontal plane.

[^13]:    * This is another example of the fruitful principle of reversion.

[^14]:    * The student must not confound these two qualities. To say that the sun appears to rise in the eastern quarter of the heavens and to set in the western is corrcet, but not precise. A watch with a second-hand indicates the time of day precisely, but not always correctly. The statement that two and two make five is precise, but is not usually regarded as correct.

[^15]:    ＊In the＂third method＂the bearings should be written obliquely upward，as directed in Art．194，but are not so printed here，from typographical difficulties．

[^16]:    * "In the description of land conveyed, the rule is that known and fixed monuments control courses and distances. So the certainty of metes and bounds will include and pass all the lands within them, though they vary from the given quantity expressed in the deed. In New York, to remove, deface, or alter landmarks maliciously is an indictable offense."-Kent's Commentaries, IV, 515.

[^17]:    * This was demonstrated by Dr. Bowditch in No. 4 of "The Analyst."

[^18]:    * It should be remembered that the following discussions of the latitudes and longitudes of the points of a surver will not be perfectly applicable to those of distant places, such as the cities just named, in consequence of the surface of the earth not being a plane.

[^19]:    * Whenever sines, cosines, tangents, etc., are here named, ther mean the natural . sines, etc., of an are described with a radias equal to one, or to the unit br which the sines, etc., are measured.

[^20]:    * The first traverse-table for surveyors seems to have been published in 1791, by John Gale. The most extensive table is that of Captain Boileau, of the British army, being calculated for every minute of bearing, and to five decimal places, for distances from 1 to 10 . The table in this volume was calculated for it, and then compared with the one just mentioned.
    $\dagger$ In using this or any similar table, lay a ruler across the page, just above or below the line to be followed out. This is a very valuable mechanical assistance.

[^21]:    * It is frequently doubtful, in many calculations, when the final decimal is 5 , whether to increase the preceding figure by one or not. Thus, 43.5 may be called 43 or 44 with equal correctness. It is better, in such cases, not to increase the whole number, so as to escape the trouble of changing the original figure, and the increased

[^22]:    chance of error. If, however, more than one such case occurs in the same column to be added up, the larger and smaller number should be taken alternately.

    * The/ traverse-table admits of many other minor uses. Thus, it mar be used

[^23]:    * A French writer fixes the allowable difference in chaining at $1-400$ of level lines; $1-200$ of lines on moderate slopes ; 1-100 of lines on steep slopes.

[^24]:    * A demonstration of this principle was giren br Dr. Bowaitch, in No. 4 of "The Analyst."

[^25]:    * This is most easily done with the aid of a right-angled triangle, sliding one of the sides adjacent to the right angle along the blade of the square, to which the other side will then be perpendicular.

[^26]:    * It is, however, substantially the same as Mr. Thomas Burgh's "Method to determine the Areas of Right-lined Figures universally," published nearly a century ago.
    $\dagger$ The phrase "meridian distance" is generally used for what is here called "longitude"; but the analogy of "differences of longitude" with "differences of latitude," usually but anomalously united with the word "departure," borrowed from navigation, seems to put beyond all question the propriety of the innovation here introduced.

[^27]:    * The last course is a "preceding course" to the first course, as will appear on remembering that these tiro courses join each other on the ground.

[^28]:    * The north pole is very nearly at the intersection of the line from Polaris to Alioth, and a perpendicular to this line from the small star seen to the left of it in Fig. 189.

[^29]:    * To calculate the time of the north star passing the meridian at its upper culmination: Find in the "American Ephemeris and Nautical Almanac" the right ascen.

[^30]:    * To calculate the times of the greatest elongation of the north star: Find in the "American Ephemeris and Nautical Almanac" its polar distance at the given time. Add the logarithm of its tangent to the logarithm of the tangent of the latitude of the place, and the sum will be the logarithm of the cosine of the hour angle before or after the culmination. Reduce the space to time; correct for sidereal acceleration ( 3 m .56 s . for 24 hours) and subtract the result from the time of the star's passing the meridian on that day, to get the time of the eastern elongation, or add it to get the western.

[^31]:    * To calculate this azimuth: From the logarithm of the sine of the polar distance of the star, subtract the logarithm of the cosine of the latitude of the place; the remainder will be the logarithm of the sine of the angle required. The polar distance can be obtained as directed in the last note.

[^32]:    * Algebraically, always subtract the bearing from the azimuth, and give the rcmainder its proper resulting algebraic sign. It will be the declination; east if plus, and west if minus. Thus, in the first case above, the declination $=+2^{\circ}-\left(-5^{\circ}\right)$ $=+7^{\circ}=7^{\circ}$ east. In the second case, the declination $=+2^{\circ}-\left(+1 \frac{1}{4}^{\circ}\right)=+\frac{8^{\circ}}{4^{\circ}}=\frac{8^{\circ}}{9^{\circ}}$ east. In the third case, the declination $=+2^{\circ}-\left(+10^{\circ}\right)=-8^{\circ}=8^{\circ}$ west.
    $\dagger$ Copied from "United States Coast and Geodetic Survey Report," 1882.

[^33]:    * For table of hourly variation of the declination, see "Report of United States Coast and Geodetic Survey," 1881, p. 136.

[^34]:    * This remeds seems to hare been first suggested by Rittenhouse. It has since been recommended by T. Sopwith, in 1822 ; by E. F. Johnson, in 1831, and by W. Roberts, of Tror, in 1839. The errors of resurreys, in which the change is neglected, were noticed in the "Philosophical Transactions," as long ago as 1679. On magnetic declination, see the following "Reports of the United States Coast and Geodetic Survey"; Report of 1881, Appendix IN; Report of 1882, Appendix III.

[^35]:    * From the Latin word collimo, or collineo, meaning to direct one thing toward another in a straight line, or to aim at. The line of aim would express the meaning.

[^36]:    * The proper care of instruments must not be orerlooked. If varnished, they should be wiped gently with fine and clean linen. If polished with oil, ther should be rubbed more strongly. The parts neither varnished nor oiled should be cleaned with Spanish-white and alcohol. Varnished wood, when spotted, should be wiped with very soft linen, moistened with a little olive-oil or alcohol. Enpainted wood is

[^37]:    cleaned with sand-paper. Apply olive-oil where steel rubs against brass ; and wax softened by tallow where brass rubs against brass. Clean the glasses with kid or buck skin. Wash them, if dirtied, with alcohol.

    * The vernier is so named from its inventor, in 1631. The name "Nonius," often improperly given to it, belongs to an entirely different contrivance for a similar object.

[^38]:    * The student will do well to draw such a scale and rernier on tro slips of thick paper, and move one beside the other till he can read them in any possible position; and so with the following verniers.

[^39]:    * In algebraic language, let $s$ equal the length of one part on the original line, and $v$ the unknown length of one part on the vernier. Let $m$ of the former $=$ $m+1$ of the latter. Then $m s=(m+1) v . \quad v=\frac{m}{m+1} s . \quad s-v=s$ 。
    $\frac{m}{m+1} s=\frac{s}{m+1} . \quad$ If $m s=(m-1) v$, then $v-s=\frac{s}{m-1}$.

[^40]:    * It has been well said that, "in the present state of science, it may be laid down as a maxim that every instrument should be so contrived that the observer may easily examine and rectify the principal parts; for, however careful the instrumentmaker may be, however perfect the execution thereof, it is not possible that any instrument should long remain accurately fixed in the position in which it came out of the maker's hands." (Adams's " Geometrical and Graphical Essays," 1791.)

[^41]:    * The learner will do well to gauge his own precision and that of the instrument (and he may rest assured that his own will be the one chiefly in fault) by measuring, from any station, the angles between successive points all around him, till he gets back to the first point, beginning at different parts of the circle for each angle. The sum of all these angles should exactly equal $360^{\circ}$. He will probably find quite a difference from that.

[^42]:    * Calculated by Alfred Noble and William T. Casgrain, and used on the United States Lake Surrey.

[^43]:    * From Horace Andrews, C. E., assistant on New York State Surrey.

[^44]:    * The greater part of the work was done by W. B. Landreth, C. E.

[^45]:    * This ingenious contrivance is due to Mr. R. Hood, in wbose practice, while running an air-line for a railroad, the necessity occurred.

[^46]:    * The length of the line AZ can also be at once obtained, since it is equal to the square root of the sum of the squares of $A X$ and $\Sigma Z$, or to the latitude dirided by the cosine of the bearing.

[^47]:    * In this figure and the following ones the angular point inclosed in a circle indicates the place at which the instrument is set.

[^48]:    * The teacher can make any number of examples for his own use by taking a tolerably accurate survey, striking out the bearing and distance of any one course, and calculating it precisely as in Case 1 , given below. He can then omit any two quantities at will, to be supplied by the student by means of the rules now to be given.

[^49]:    * This conception of thus changing the bearings is stated to be due to Professor Robert Patterson, of Philadelphia, by whom it was communicated to Mr. John Gummere, and published by him, in 1814, in his "Treatise on Surreying."

[^50]:    * The given lines will be represented by fine full lines, the lines of construction by broken lines, and the lines of the result by heavy full lines.

[^51]:    * The problem may also be performed by making the side on which the dirision.

[^52]:    * As some lines in the figure are not used in the construction, though needed for the demonstration, the student should draw it himself to a large scale.

[^53]:    * If a line be drawn joining the middle points of the parallel bases of a trape-

[^54]:    * Arts. 455 to 462 of this chapter are mainly taken from "Instructions to the Sur-vevor-General of Oregon, being a Manual for Field Operations," prepared, in March, 1851, by John M. Moore, Principal Clerk of Surveys.

[^55]:    * The marks $\mathrm{O}_{1}+$, and $\wedge$, merely refer to the dates of the survers. They are sometimes used to point out lands offered for sale, or reserred, etc.

[^56]:    * Until 1866 they were either 24 or 30 miles apart.

[^57]:    * The surveyor should prepare a diagram of the townships, with the numbers here referred to, in their proper places, as here indicated.

[^58]:    " 3 . Field-notes of the exterior lines of towaships, showing

[^59]:    * These tables were calculated by Edward W. Arms, C. E., for W. \& L. E. Gurley.

[^60]:    * This attachment, shown in Fig. 342, is manufactured by W. \& L. E. Garley, Troy, New York.

[^61]:    * Invented by G. N. Saegmüller, and manufactured by Fauth \& Co., Washington, D. C., from whose catalogue the description is taken.

[^62]:    * A cylindrical surface is here understood to mean that formed by a line moring parallel to itself along any line, instcad of only a circle, as in elementary geometry.

[^63]:    * Made by Henry J. Green, rī Broadrar, Nerr York.

[^64]:    A Topographical Drawing of Eagle Cliff, by E. Hergesheimer, $\Lambda$ ssistant, United States Coast Geodetic Survey. Scale $\frac{1}{1000 \pi}$

[^65]:    * The Plane-Table is not a Goniometer, or Angle-measurer, like the compass, transit, etc., but a Gonigraph, or Anglediawer.

[^66]:    * Manufactured by Fauth \& Co., Washington, D. C.

[^67]:    * The French phrase, to "orient one's sclf," meaning to determine one's position, usually with respect to the four quarters of the hearens, of which the Orient is the leading one, well deserves naturalization in our language.

[^68]:    * More precisely, A being this angle, and not more than $2^{\circ}$ or $3^{\circ}$, the difference between the inclined and horizontal lengths equals the inclined or real length multiplied by the square of the minutes in A, and that by the decimal 0.00000004231 .

[^69]:    * If the triangles were very large, they would have to be regarded as spherical, and the sum of their angles would be more than $180^{\circ}$; but this "spherical excess" would be only $1^{\prime \prime}$ for a triangle containing 76 square miles, $1^{\prime}$ for 4,500 square miles, etc.; and may therefore be neglected in all ordinary surreying operations.

[^70]:    * See "Report of Coast and Geodetic Surrey," 1868, 1876, 1850, $188 \%$.

[^71]:    * For descriptions of various forms of base apparatus, see "Report of United States Coast and Geodetic Survey," 1854, 1857, 1880, 1882; "Report of Primary Triangulation of the United States Lake Survey"; Wright's "Adjustment of Observations," Chapter VII.

[^72]:    * For the arc $A B$ measures the angle $A 0 B$ at the center, which angle $=180^{\circ}$ $-2\left(90^{\circ}-\mathrm{ASB}\right)=2 \mathrm{~A}$ S B. Therefore, any angle inscribed in the circumference and measured by the same arc is equal to $\mathrm{A} S \mathrm{~B}$.

[^73]:    * For merely solving triangles, only Articles $1,2,3,5,6,10,11$, and 12 are needed.
    $\dagger$ The number of seconds in any are which is given in parts of radius, radius being unity, equals the length of the are so given divided by the length of the arc of one second; or multiplied by the number of seconds in radius.

[^74]:    * For the great value of this indirect mode of comparing the sides and angles of triangles, see Comte's "Philosophy of Mathematics" (Harper's, 185\%), page 225.

[^75]:    * Consequently, the note on page 523 may read thus: The number of seconds in any very small are given in parts of radius, radius being unitr, is equal to the length of the are so given divided by $\sin .1$.

[^76]:    * The square, etc., of the sine, etc., of an arc, is often expressed by placing the exponent between the abbreviation of the name of the trigonometrical line and the

[^77]:    * Three numbers, $m, n, p$, arranged in decreasing order of size, form an harmonic proportion, when the difference of the first and the second is to the difference of the second and the third, as the first is to the third. Such are the numbers 6,4 , and 3 ; or 6,3 , and 2 ; or 15,12 , and 10 ; etc. So, in Fig. $58 \%$, are the lines A D, A B, and AC , which thus give $\mathrm{BD}: \mathrm{CB}:: \mathrm{AD}: \mathrm{AC}$; or $\mathrm{AC}: \mathrm{CB}:: \mathrm{AD}: \mathrm{BD}$. The series of fractions, $\frac{1}{1}, \frac{1}{2}, \frac{1}{3}, \frac{1}{4}, \frac{1}{5}$, etc., is called an harmonic progression, because any consecutive three of its terms form an harmonic proportion.

