THE MANGA GUIDE TO

COMICS INSIDE!

RELATIVITY

HIDEO NITTA MASAFUMI YAMAMOTO KEITA TAKATSU TREND-PRO CO., LTD.

Ohmsha



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"A lot of fun to read. The interactions between the characters are lighthearted, and the whole setting has a sort of quirkiness about it that makes you keep reading just for the joy of it."

-HACK A DAY ON THE MANGA GUIDE TO ELECTRICITY

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"For parents trying to give their kids an edge or just for kids with a curiosity about their electronics, *The Manga Guide to Electricity* should definitely be on their bookshelves."



bookshelves." —SACRAMENTO BOOK REVIEW "This is a solid book and I wish there were more like it in the IT world." —SLASHDOT ON THE MANGA GUIDE TO DATABASES

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"Makes it possible for a 10-year-old to develop a decent working knowledge of a subject that sends most college students running for the hills." —SKEPTICBLOG ON *THE MANGA GUIDE TO MOLECULAR BIOLOGY*

"This book is by far the best book I have read on the subject. I think this book absolutely rocks and recommend it to anyone working with or just interested in databases." —GEEK AT LARGE ON *THE MANGA GUIDE TO DATABASES*

"The book purposefully departs from a traditional physics textbook and it does it very well." —DR. MARINA MILNER-BOLOTIN, RYERSON UNIVERSITY ON *THE MANGA GUIDE TO PHYSICS*

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THE MANGA GUIDE™ TO RELATIVITY



THE MANGA GUIDE" TO RELATIVITY

HIDEO NITTA MASAFUMI YAMAMOTO KEITA TAKATSU TREND-PRO CO., LTD.





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PREFACE

Welcome to the world of relativity!

Everyone wonders what relativity is all about. Because the theory of relativity predicts phenomena that seem unbelievable in our everyday lives (such as the slowing of time and the contraction of the length of an object), it can seem like mysterious magic.

Despite its surprising, counterintuitive predictions, Einstein's theory of relativity has been confirmed many times over with countless experiments by modern physicists. Relativity and the equally unintuitive quantum mechanics are indispensable tools for understanding the physical world.

In Newton's time, when physicists considered velocities much smaller than the speed of light, it was not a problem to think that the measurement of motion, that is, space and time, were independent, permanent, and indestructible absolutes. However, by the end of the 19th century, precise measurements of the speed of light combined with developments in the study of electromagnetism had set the stage for the discovery of relativity. As a result, time and space, which had always been considered to be independent and absolute, had to be reconsidered.

That's when Einstein arrived on the scene. Einstein proposed that time and space were in fact relative. He discarded the idea that space and time were absolute and considered that they vary together, so that the speed of light is always constant.

This radical insight created a controversy just as Galileo's claim that Earth orbited the Sun (and not vice versa) shocked his peers. However, once we ventured into space, it was obvious that Earth was indeed moving.

In a similar way, relativity has given us a more accurate understanding of concepts regarding the space-time in which we are living. In other words, relativity is the result of asking what is *actually* happening in our world rather than saying our world *should be* a particular way.

Although this preface may seem a little difficult, I hope you will enjoy the mysteries of relativity in a manga world together with Minagi and his teacher, Miss Uraga. Finally, I'd like to express my deep gratitude to everyone in the development bureau at Ohmsha; re_akino, who toiled over the scenario; and Mr. Keita Takatsu, who converted it into such an interesting manga.

Well, then. Let's jump into the world of relativity.

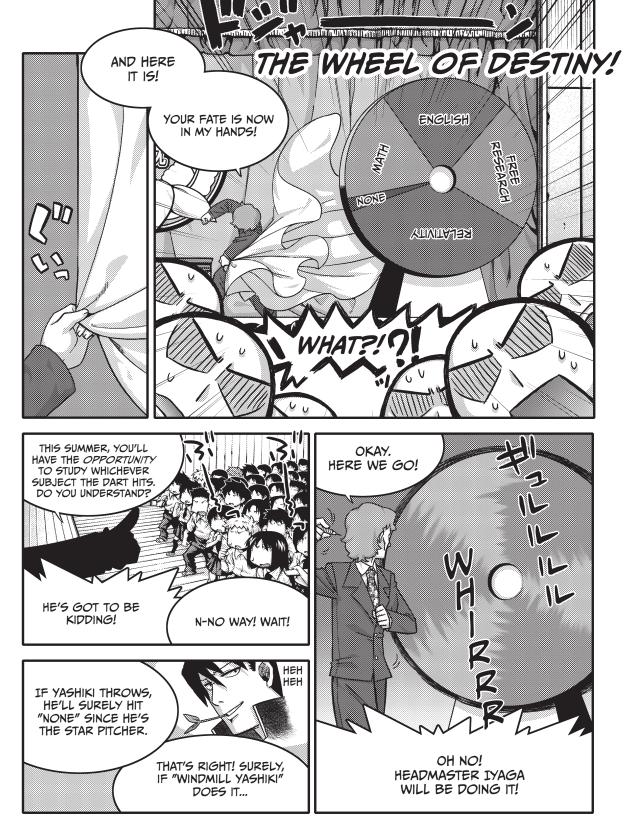
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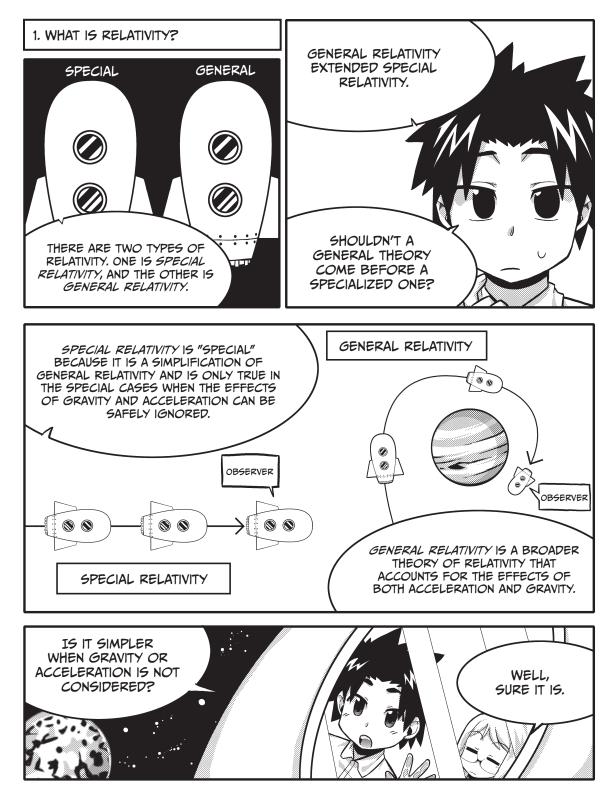
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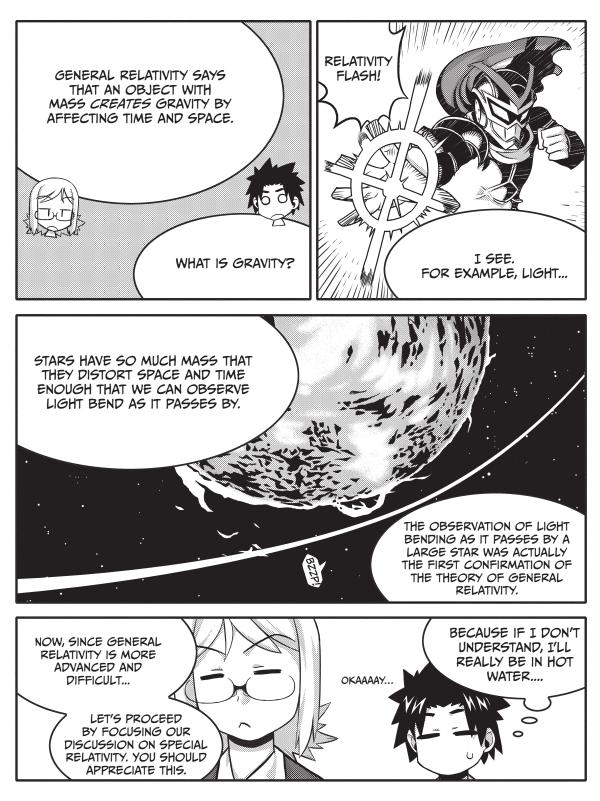
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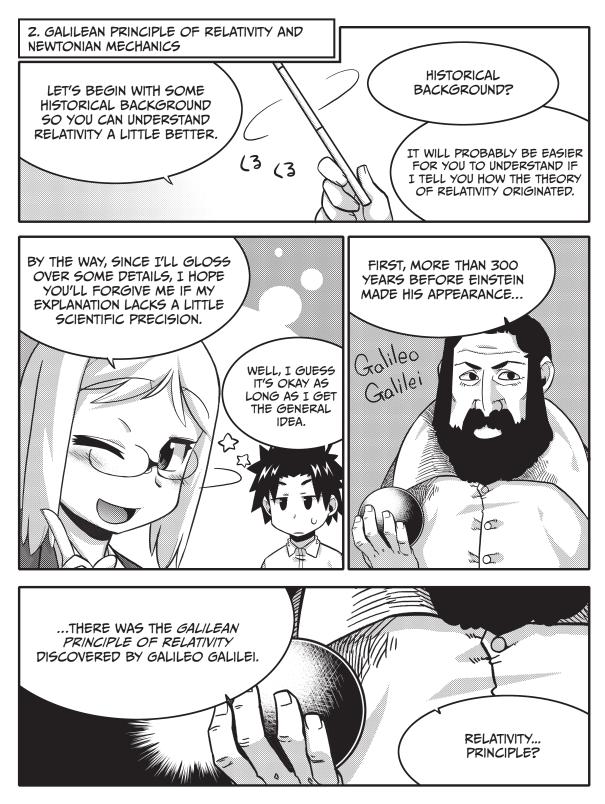
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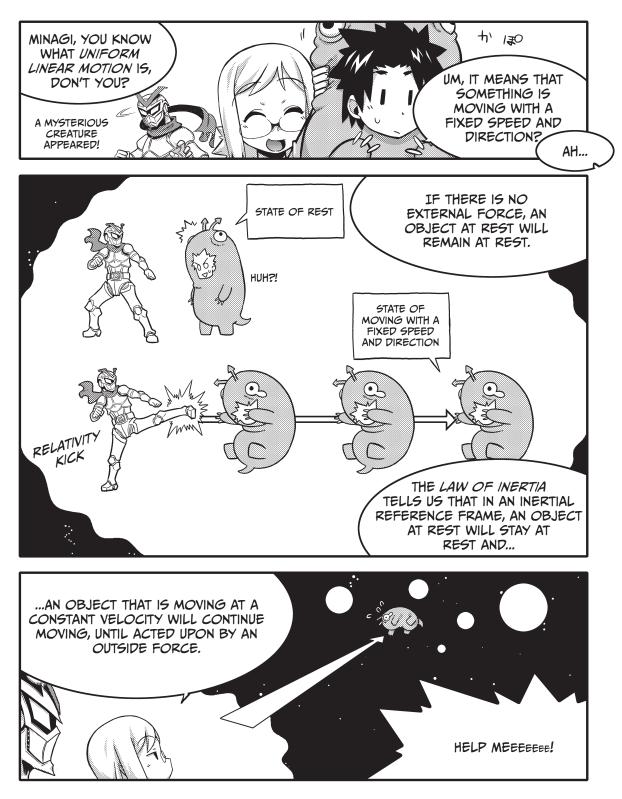


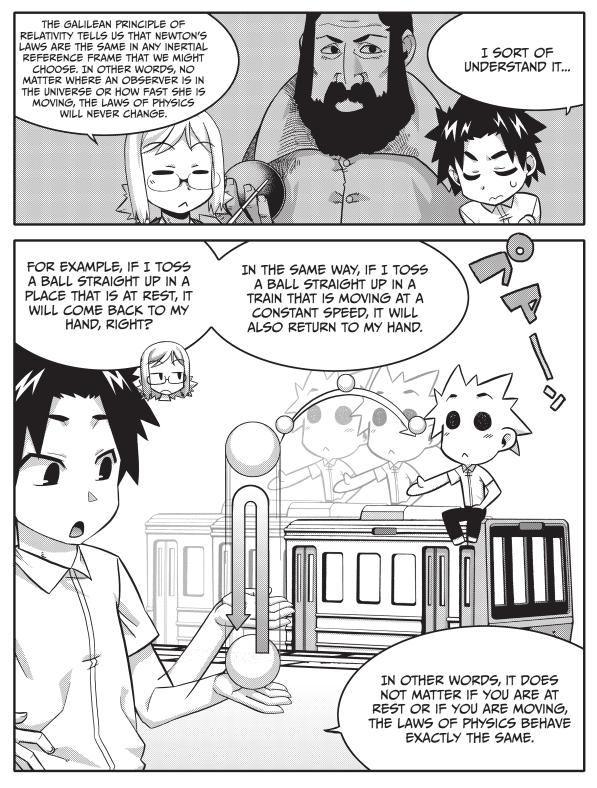






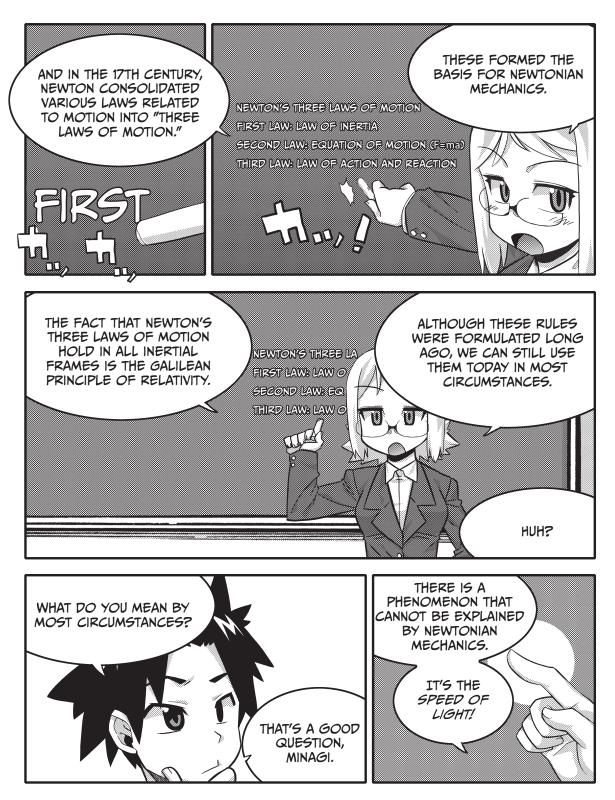


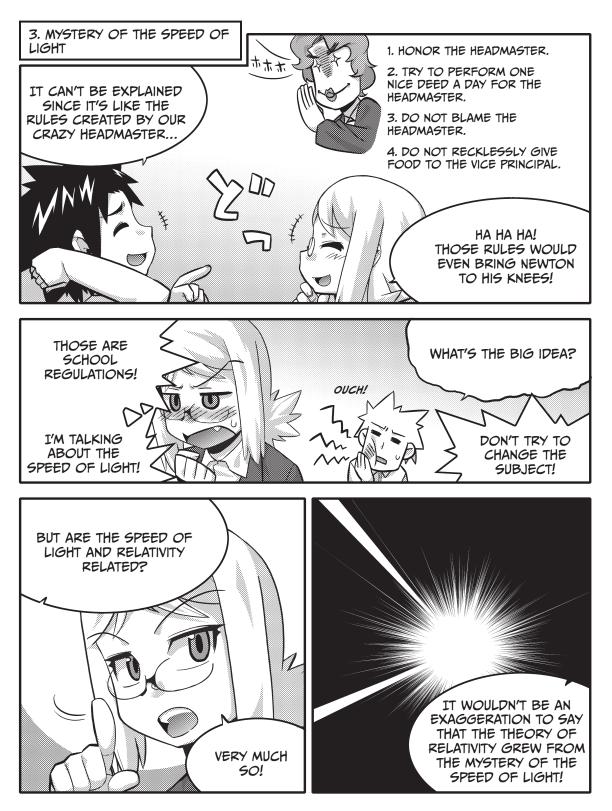




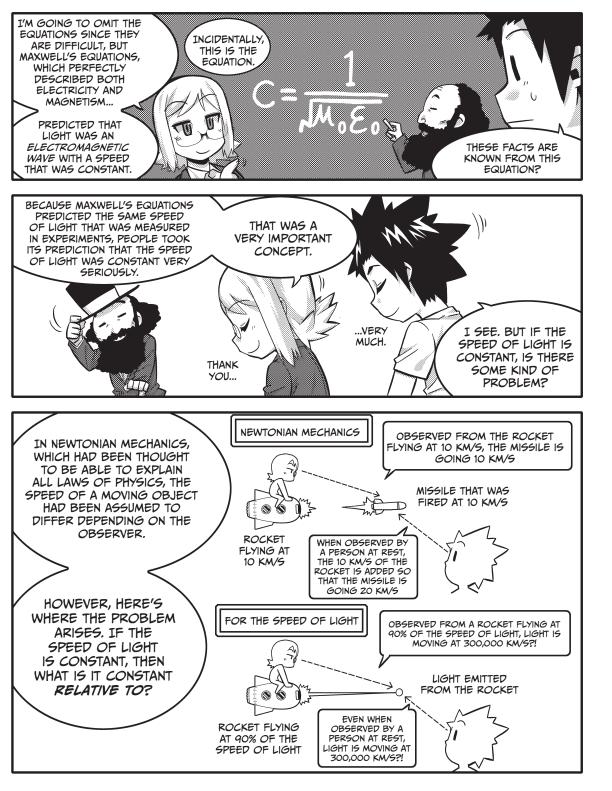


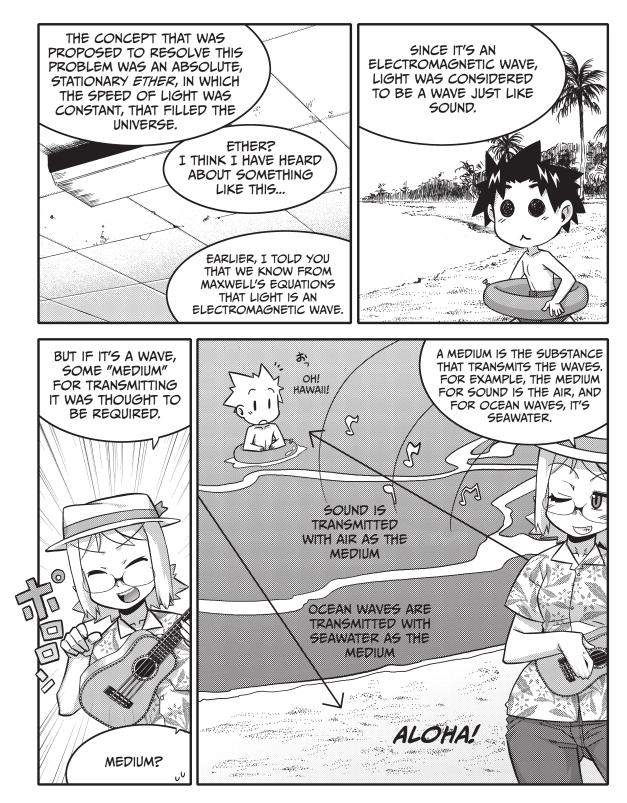


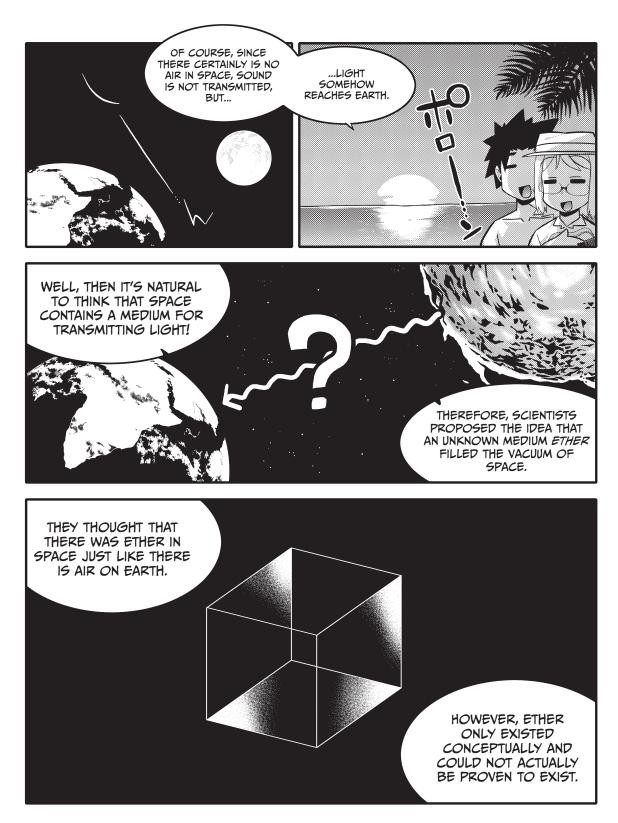


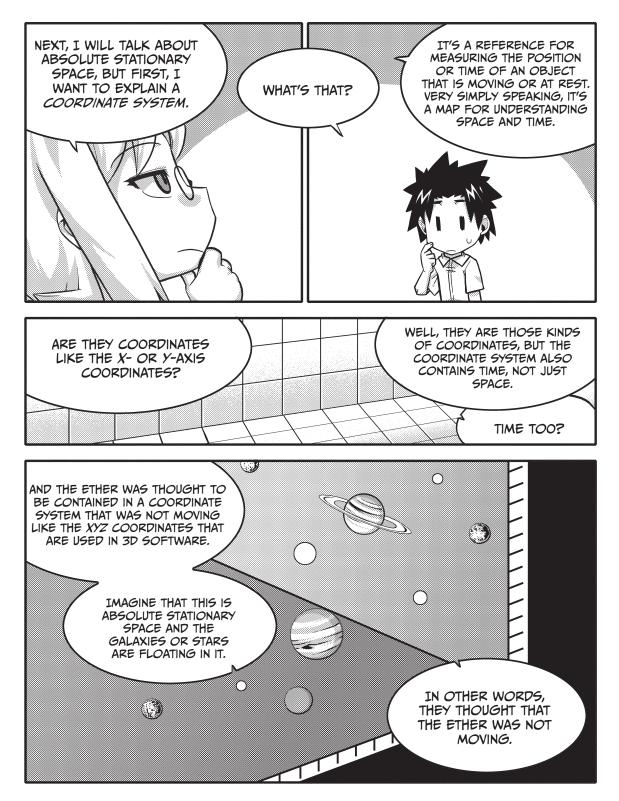




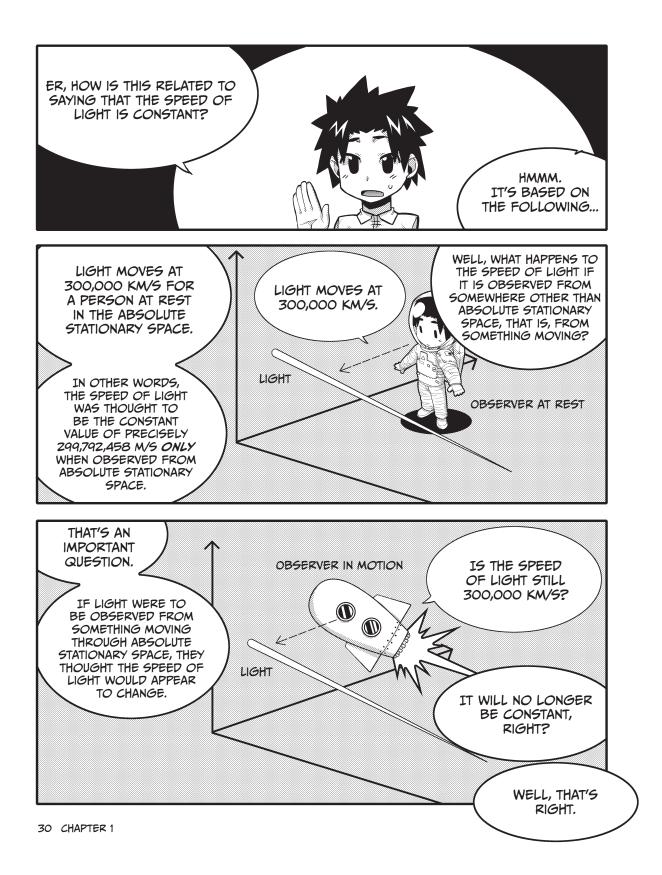


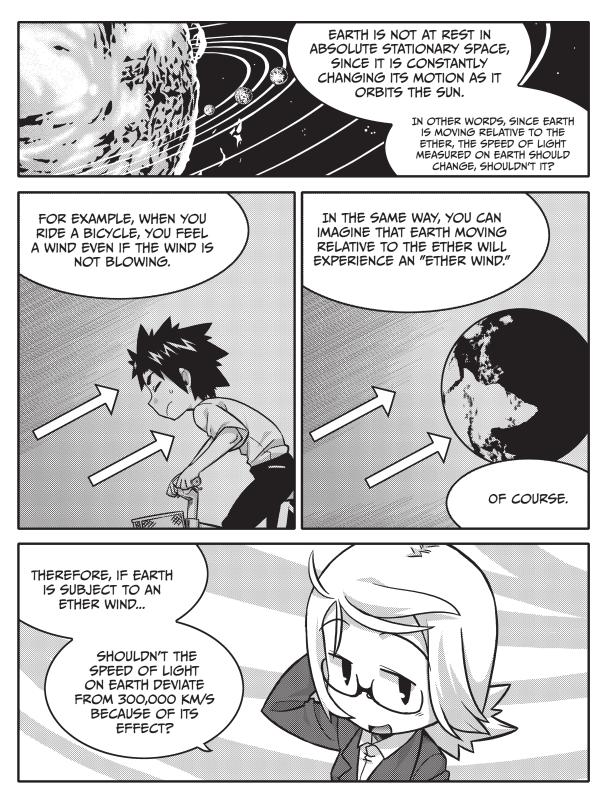


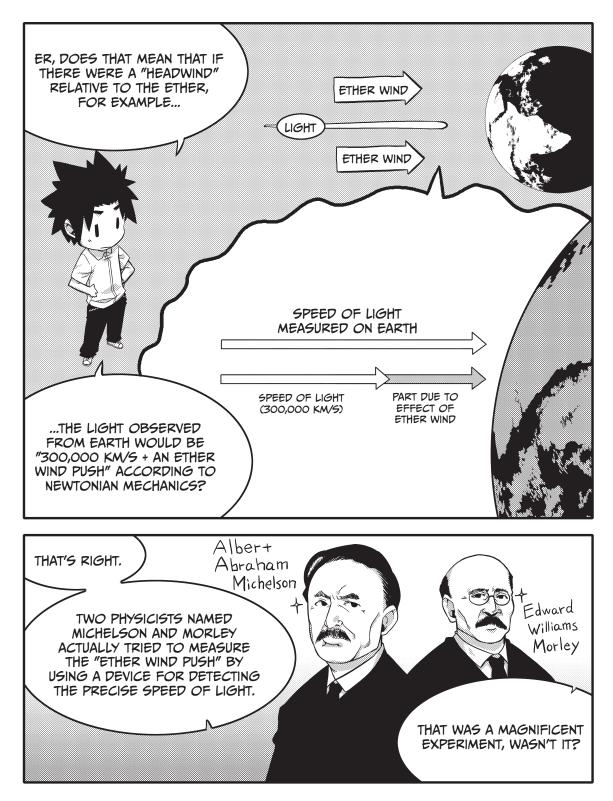


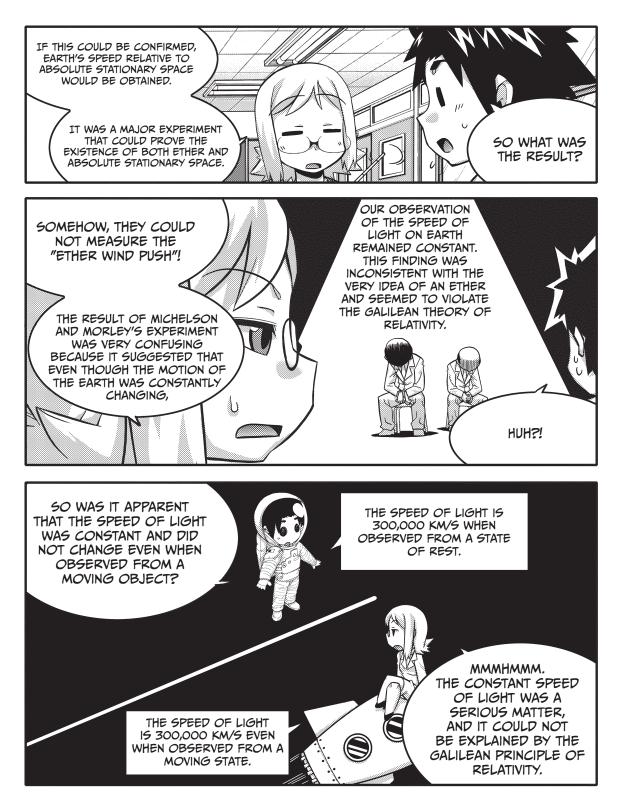


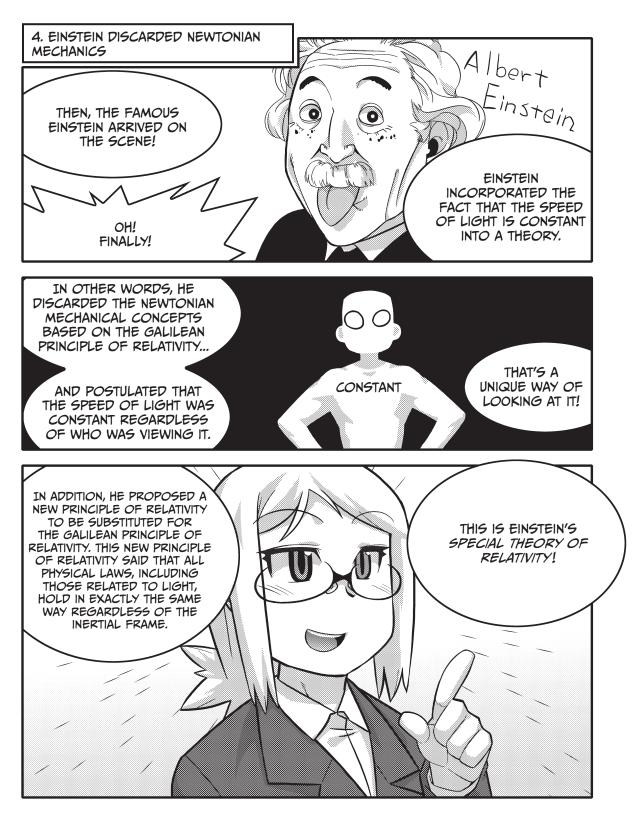


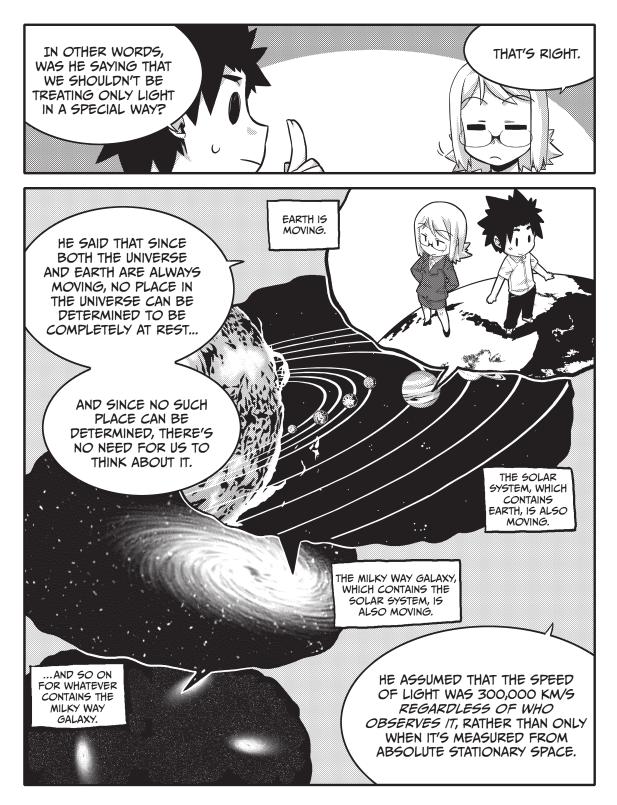


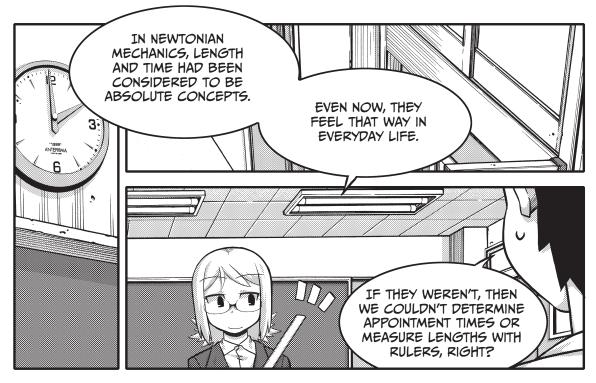




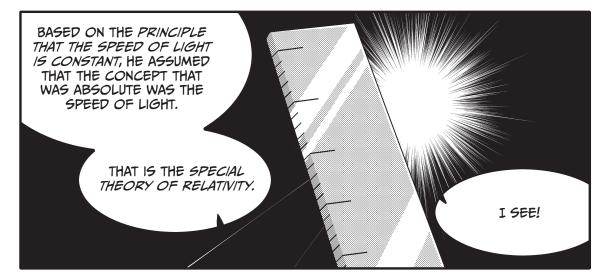


















WHAT IS LIGHT?

Maxwell's equations tell us that light is an electromagnetic wave. The color of light is determined by the wavelength of the electromagnetic wave. Red light has a wavelength of 630 nm, and blue light has a shorter wavelength of approximately 400 nm, where one nanometer (1 nm) = one billionth of a meter (10^{-9} m). Electromagnetic radiation at different wavelengths takes many forms, such as radio waves, X-rays, and gamma (γ) rays (see Figure 1-1).

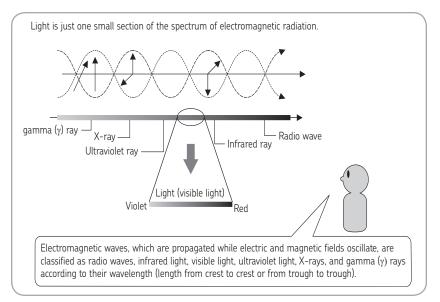


Figure 1-1: Light is an electromagnetic wave.

Although light may seem common enough—it is all around us, after all—it is fundamental to both relativity and quantum theory, the cornerstones of modern physics.

But before we delve into light's true nature, let's introduce the properties of light that have been known for a long time.

First, you know that light is *reflected* by a mirror or the surface of water. You also know about the *refraction* of light—you only need to look at your feet the next time you take a bath or see how your straw "bends" when you put it in a glass of water. Any change in medium changes a wave's direction, due to a change in the wave's speed through that medium.

Some mediums refract light of different wavelengths different amounts. In other words, light of different colors is bent to different degrees, a property known as *dispersion*. This causes white light, which consists of light of all colors, to be spread out into a spectrum of light from red to violet. We can see the seven colors of a rainbow because of dispersion.

These properties of reflection, refraction, and dispersion have been used to create precision camera lenses and telescopes. Figure 1–2 shows what happens to light when it is reflected, refracted, or dispersed.

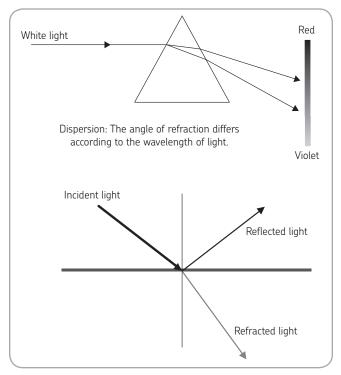


Figure 1-2: Dispersion, reflection, and refraction

Next, more subtle phenomena called *interference* and *diffraction* can be observed. These phenomena stem from the fact that light is a wave. Interference describes what happens when two light waves come together. When the two waves come together, the result is either *constructive interference*, where the waves' amplitudes are added together, or *destructive interference*, where one wave's amplitude is subtracted from the other's. Figure 1-3 shows the different kinds of interference.

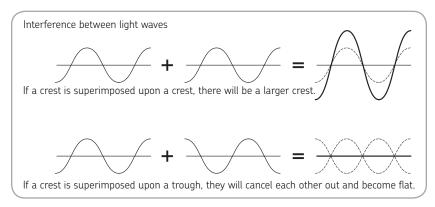


Figure 1-3: Interference can make waves stronger or weaker.

Diffraction can be observed when light passes through a tiny hole about the same size as the wavelength of the light. Due to the constructive and destructive interference of different parts of the light wave with itself, passing through a tiny aperture can cause the light to spread out or bend, as shown in Figure 1-4. Diffraction is often what limits the resolution of microscopes.

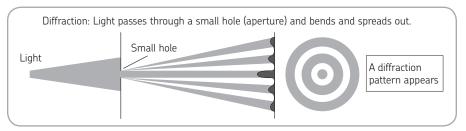


Figure 1-4: Diffraction comes about from interference.

Another property of light is called *polarization*, a property that describes the orientation of the transverse electric and magnetic components of the electromagnetic wave. This property is very useful; it allows special filters to be made (called *polarizing filters*) that allow only light with a specific polarization to pass (see Figure 1–5).

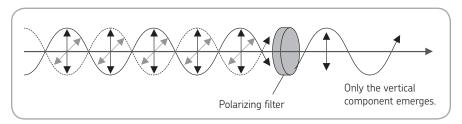


Figure 1-5: Polarization

In *scattering*, light collides with dust and other particles in the air, thereby changing direction (see Figure 1–6). Since blue light (with shorter wavelengths) is scattered by water molecules in the air more than red light (with longer wavelengths), the sky appears blue.

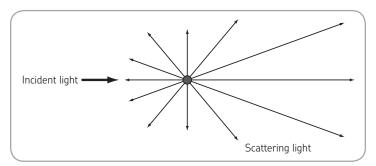


Figure 1-6: Scattering

LIGHT IS CONSTANT (AND THEY PROVE IT EVERY DAY IN A LAB CALLED SPRING-8)

Various tests have been conducted to verify that the speed of light is truly constant. This is important because it is one of the fundamental premises of relativity.

One way that we can test this property is to measure the speed of light coming from an object that is moving very fast. If the speed of light is not constant, the Newtonian notion of "adding" relative velocities predicts that light coming from an object moving towards the observer will be the speed of light plus the speed of the moving object; for example, if the object is moving near the speed of light, then the light from the object should be moving nearly twice the speed of light. If the speed of light is constant, on the other hand, than the light coming from the fast-moving object will just be the speed of light. Measurements confirm that the speed of light is always the same, regardless of the speed of the object from which it comes (see Figure 1–7).

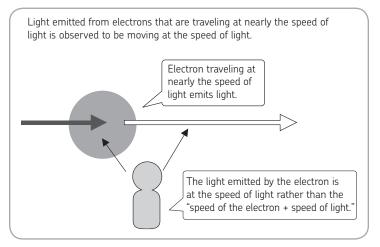


Figure 1-7: Verification that the speed of light is constant at SPring-8

Moving objects near the speed of light for these experiments is extremely difficult, and these experiments are performed at very specialized scientific facilities. SPring-8 is a synchotron radiation facility in Japan's Hyogo Prefecture that performs experiments by smashing together electrons traveling at extremely fast speeds (99.9999998 percent of the speed of light). Besides verifying that the speed of light is constant, these experiments help scientists uncover the basic building blocks of matter.

WHAT'S SIMULTANEOUS DEPENDS ON WHOM YOU ASK! (SIMULTANEITY MISMATCH)

If we consider the principle that "the speed of light is constant," various phenomena appear strange. One of these is the phenomenon called the *simultaneity mismatch*, which means that what is simultaneous for me is not the same as what is simultaneous for you.

I can imagine that you are thinking, "What in the world are you saying?" So let's consider the concept of "simultaneous" again. We will compare the case of Newtonian velocity addition (nonrelativistic addition of velocity) with the case in which the speed of light is constant (relativistic addition of velocity).

Consider Mr. A, who is riding on a rocket flying at a constant velocity, and Mr. B, who is observing Mr. A from a stationary space station. Assume that Mr. A is in the middle of the rocket. Sensors have been placed at the front and back of the rocket. Mr. A throws balls (or emits light) toward the front and back of the rocket. We will observe how those balls (or light beams) hit the sensors at the front and back of the rocket.

CASE OF NEWTONIAN VELOCITY ADDITION (NONRELATIVISTIC ADDITION)

First, we will use the motion of the balls to consider the case in which velocities are added in a Newtonian mechanical manner (before considering relativity).

First, let's look at Mr. A as shown in Figure 1-8. Since from Mr. A's perspective the rocket is not moving, the balls, which are moving at the same velocity from the center toward the sensors at the front and back of the rocket, arrive at the sensors "simultaneously."

Next, when observed by Mr. B from the space station, the rocket advances in the direction of travel. In other words, using the point of departure of the balls (dotted line) as a reference, the front of the ship moves away from the dotted line, and the back of the ship approaches the dotted line. However, since the velocity of the rocket is added to the velocity of the ball in the forward direction, according to normal addition, the ball's velocity increases and it catches up with the front of the ship. On the other hand, the velocity of the ball toward the back of the ship is reduced by the velocity of the rocket (indicated by the short arrow in the figure), and the back of the ship catches up to the ball. Therefore, Mr. B also observes that the balls arrive at the front and back of the ship "simultaneously."

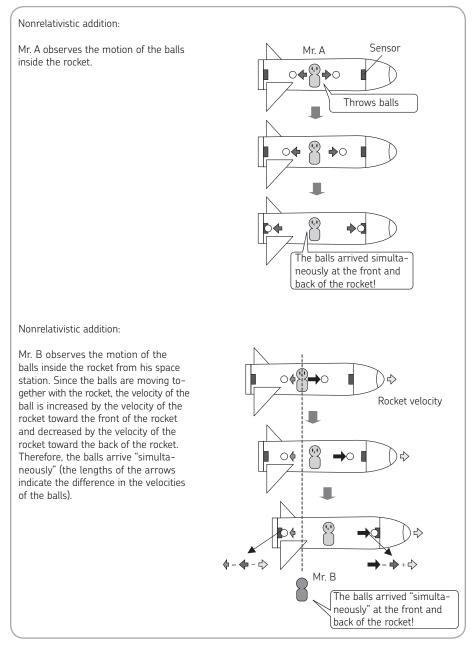


Figure 1-8: Newtonian velocity addition

CASE IN WHICH THE SPEED OF LIGHT IS CONSTANT (RELATIVISTIC ADDITION OF VELOCITY)

Now let's consider the case in which the speed of light is constant. Instead of throwing balls, Mr. A will emit light while traveling at nearly the speed of light (see Figure 1-9).

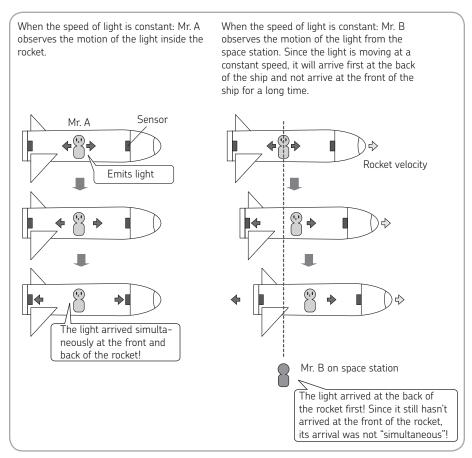


Figure 1-9: Case in which the speed of light is constant (relativistic addition of velocity)

You may have already realized what is at issue: Mr. B's observation will differ from that of Mr. A.

For Mr. A, even when the speed of light is constant, the light will arrive "simultaneously" at the front and back of the rocket.

However, when observed by Mr. B, the light moving towards the front of the ship does not arrive for a long time. It has to overtake the ship, which is moving away at nearly the speed of the light. Therefore, the light arrives at the back of the ship before it reaches the front of the ship.

That's right; when observed by Mr. B, the light does not arrive "simultaneously" at the front and back of the ship.

The simultaneity property of light differs in this way depending on the standpoint of the observer. This is called *simultaneity mismatch*.

GALILEAN PRINCIPLE OF RELATIVITY AND GALILEAN TRANSFORMATION

The Galilean principle of relativity says that "the laws of physics are the same regardless of whether the coordinate system from which the observation is made is at rest or moving at a constant velocity." In other words, Newtonian mechanics (the physical laws that govern motion) are always the same, regardless of whether observations are made in a reference frame that is at rest or one that is moving at a constant velocity. This principle was derived from an experiment in which an iron ball was dropped from the mast of a ship, as shown in Figure 1-10. The iron ball fell directly under the mast whether the ship was moving or at rest.

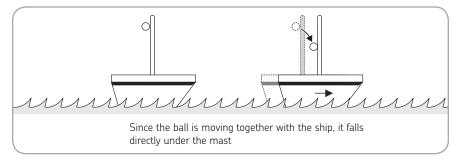


Figure 1-10: Galilean principle of relativity

Since the laws of physics are the same in any reference frame, Galileo arrived at a straightforward way to describe how observations look different depending on which reference frame you are in. Today we use algebraic equations called the *Galilean transformation* to help understand the notion of "adding" relative velocities.

Let's take two coordinate systems, one with the coordinates (x, t) and the other with coordinates (x', t'), where x and x' describe position and t and t' describe time. One can go from one coordinate system to the other, by considering the relative velocity between the two coordinate systems v.

$$x' = x - vt$$
$$t' = t$$

The above equations show the relationship between coordinates from a coordinate system at rest and a coordinate system moving at a constant velocity *v* relative to the coordinate system at rest. Inertial frames are mutually linked in this way by the Galilean transformation. If we compare them using Newton's equation of motion, we can prove that Newton's equation of motion takes the same form in each inertial frame. In other words, when the Galilean principle of relativity holds, Newtonian mechanics will hold.

DIFFERENCES BETWEEN THE GALILEAN PRINCIPLE OF RELATIVITY AND EINSTEIN'S SPECIAL PRINCIPLE OF RELATIVITY

As just described, the Galilean principle of relativity indicates that Newtonian mechanics apply across inertial frames when linked with the Galilean transformation.

On the other hand, the assumption that the speed of light is constant in any reference frame forced scientists to reformulate the Galilean transformation to be consistent with relativity. This new transformation is called the *Lorentz transformation*.

The Lorentz transformation is shown by the equations below, which show the relationship between coordinates from a coordinate system at rest and a coordinate system moving at a constant velocity v relative to the coordinate system at rest. The variables with the prime symbol (') attached represent coordinates observed from the coordinate system at rest; the variables without the prime symbol represent coordinates observed from the system in motion. Note that the speed of light c appears in the equations here. Another point to notice is that time t is transformed in a manner similar to that of length; time does not exist independently but must be considered to be unified with space.

$$x' = \frac{x - vt}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$
$$t' = \frac{t - \frac{v}{c^2}x}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$

WAIT A SECOND-WHAT HAPPENS WITH THE ADDITION OF VELOCITIES?

When we assume that the speed of light is constant, what happens when velocities are added to the mix?

According to the principle of relativity, when calculated based on the Lorentz transformation, the addition of velocities is indicated by the following equation.

$$w = \frac{u+v}{1+\frac{vu}{c^2}}$$

This equation describes the resulting addition of velocities of a missile w when the velocity of a rocket is v and the velocity (observed from the rocket) of the missile shot from the rocket is u, as shown in Figure 1-11. The difference is apparent when this equation is compared with the normal addition (nonrelativistic) equation w = u + v.

If we enter specific velocities in the above equations, we'll obtain some interesting results.

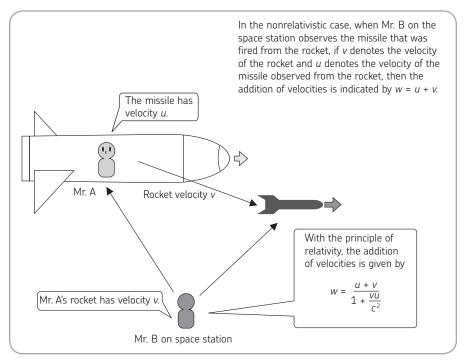


Figure 1-11: Addition of velocities

For example, when the rocket velocity v is 50 percent of the speed of light (0.5c) and the missile velocity u observed from the rocket is 50 percent of the speed of light (also 0.5c), then the missile velocity w observed by Mr. B will be 80 percent of the speed of light (0.8c).

$$w = \frac{(0.5c + 0.5c)}{\left(1 + \frac{(0.5c)^2}{c^2}\right)} = \frac{c}{1.25} = 0.8c$$

This equation also yields an interesting result when v and u are their maximum values. If the rocket velocity v is 100 percent of the speed of light (practically speaking, v = c is impossible for an object with mass, like a rocket) and the missile velocity u observed from the rocket is 100 percent of the speed of light, then the missile velocity w observed by Mr. B will be the speed of light.

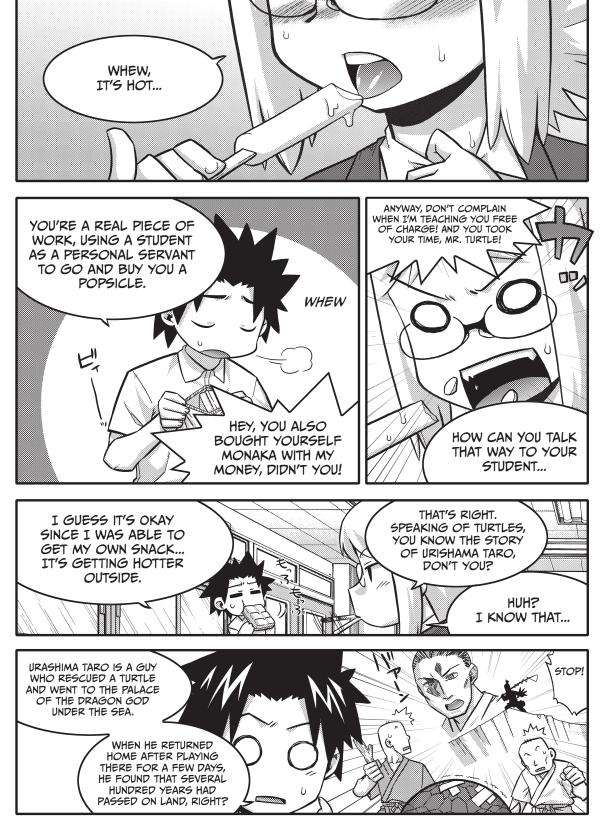
$$w = \frac{\left(c+c\right)}{\left(1+\frac{c^2}{c^2}\right)} = \frac{2c}{2} = c$$

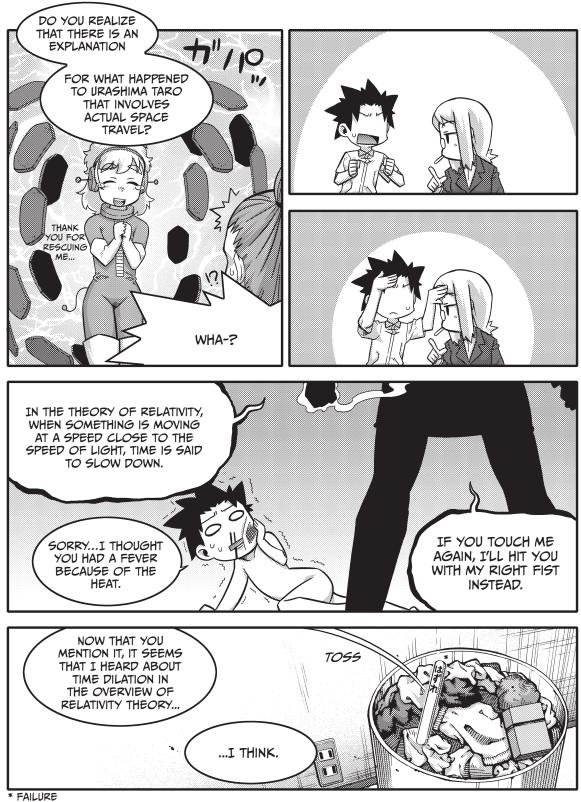
The speed of light cannot be exceeded under any circumstances!



WHAT DO YOU MEAN, TIME SLOWS DOWN?

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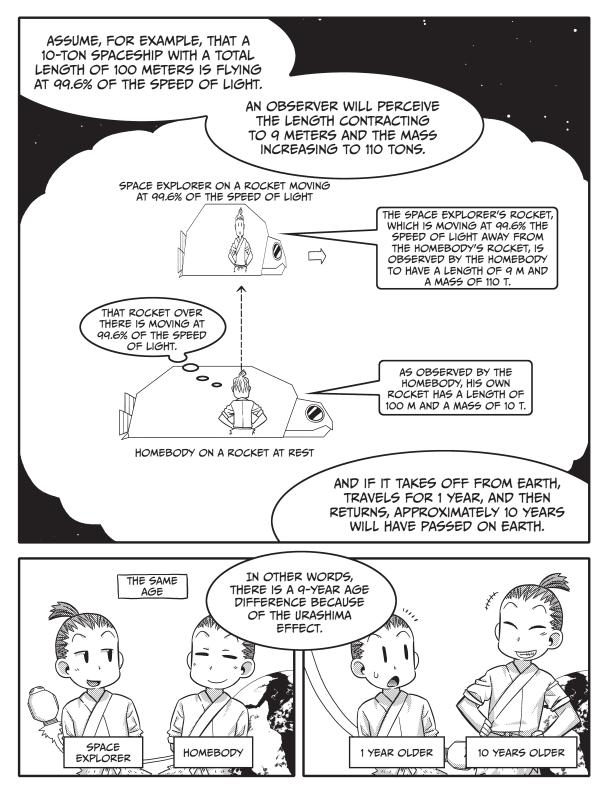




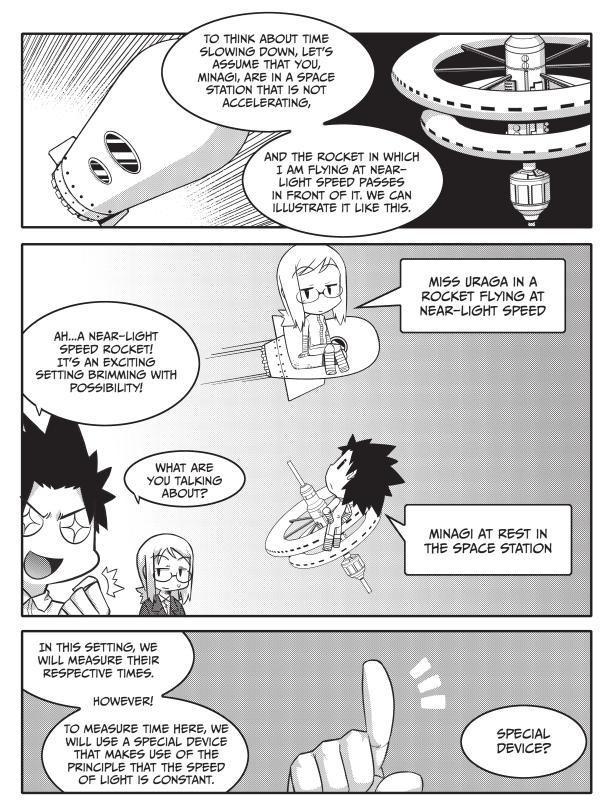
TRANSLATOR'S NOTE: IN JAPAN, POPSICLE STICKS HAVE FORTUNES ON THEM.

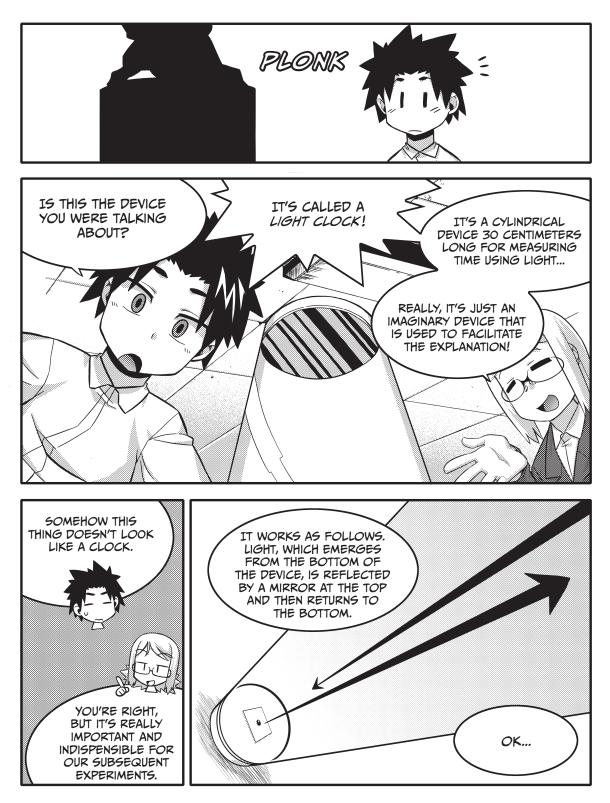
1. URASHIMA EFFECT (TIME DILATION)

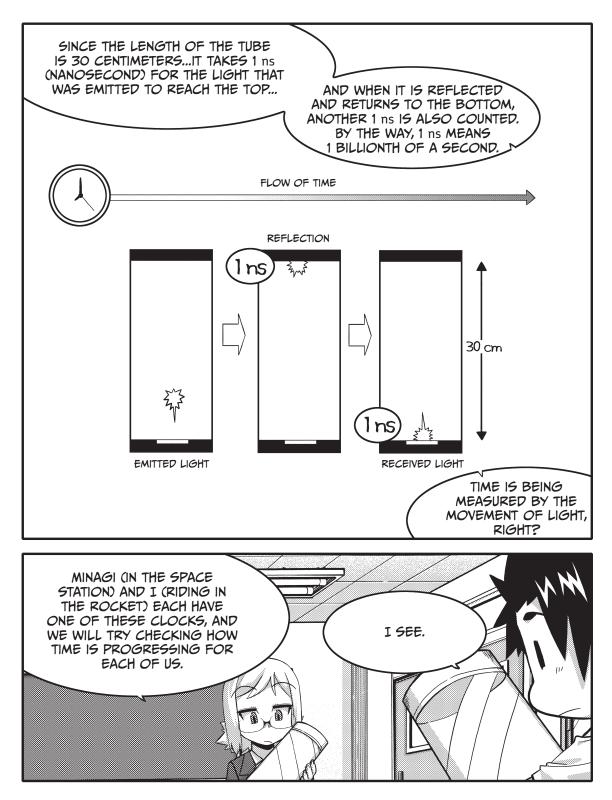


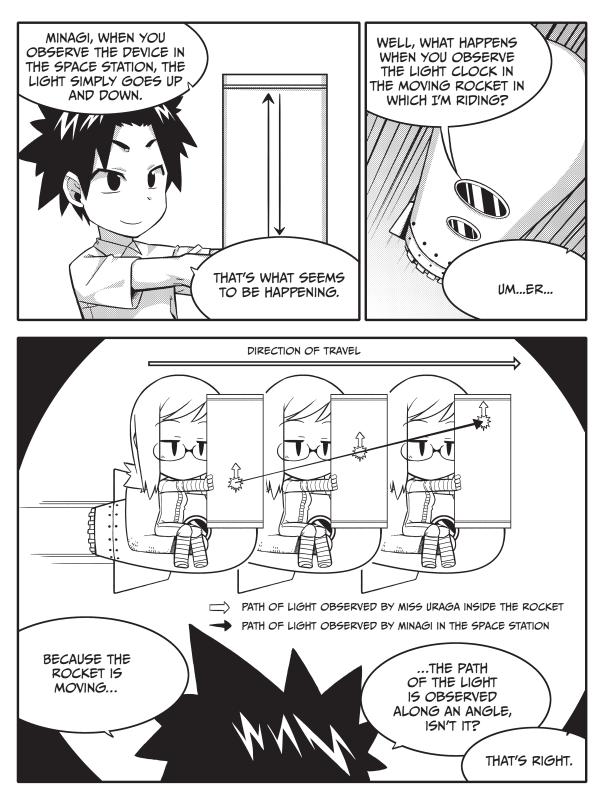


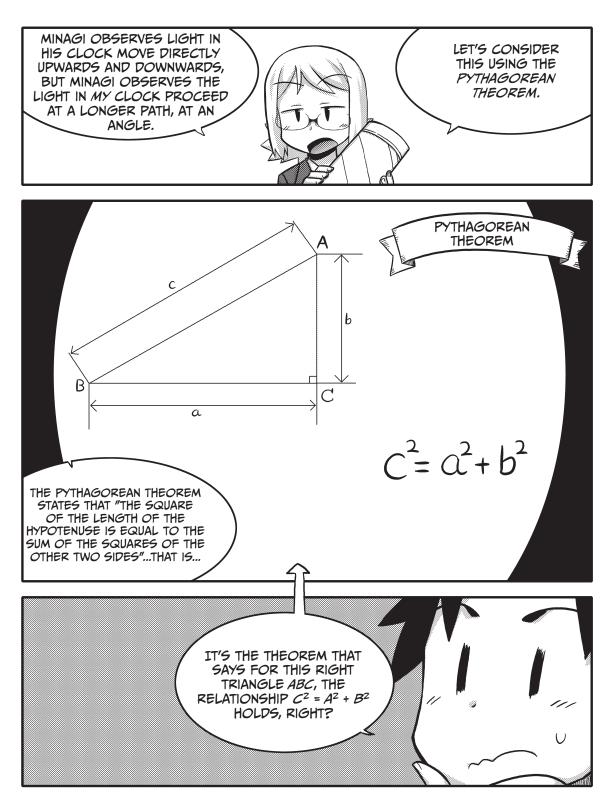


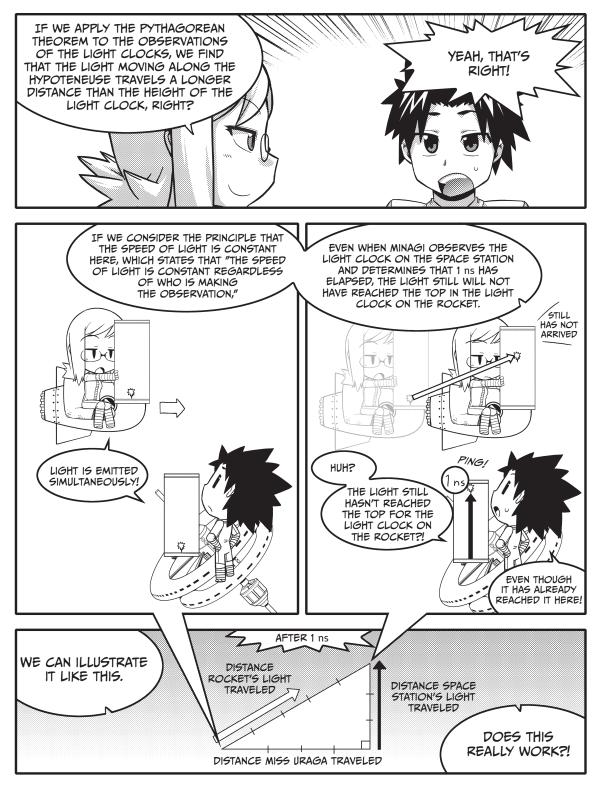


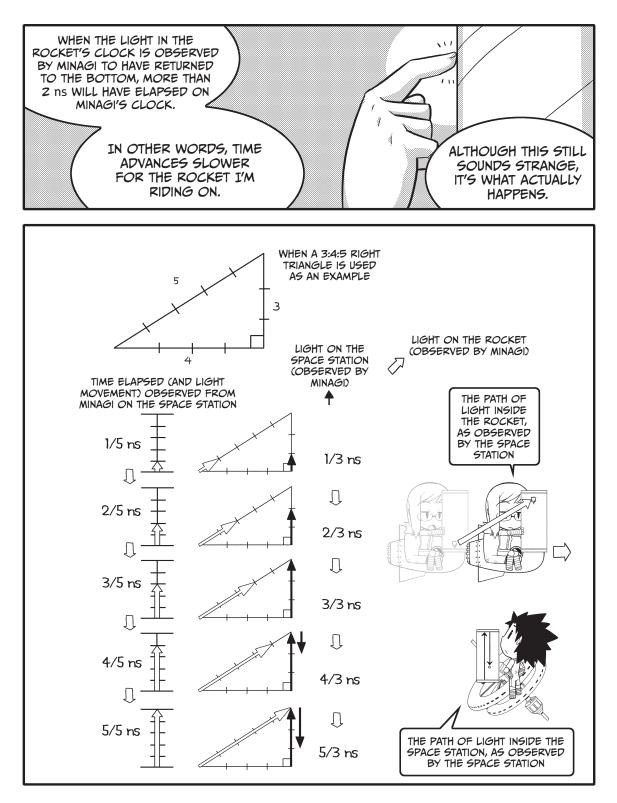


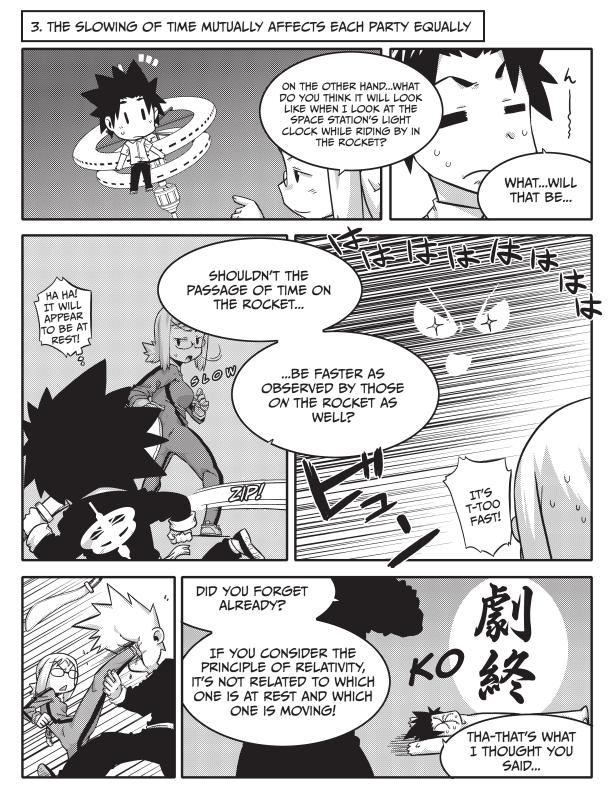


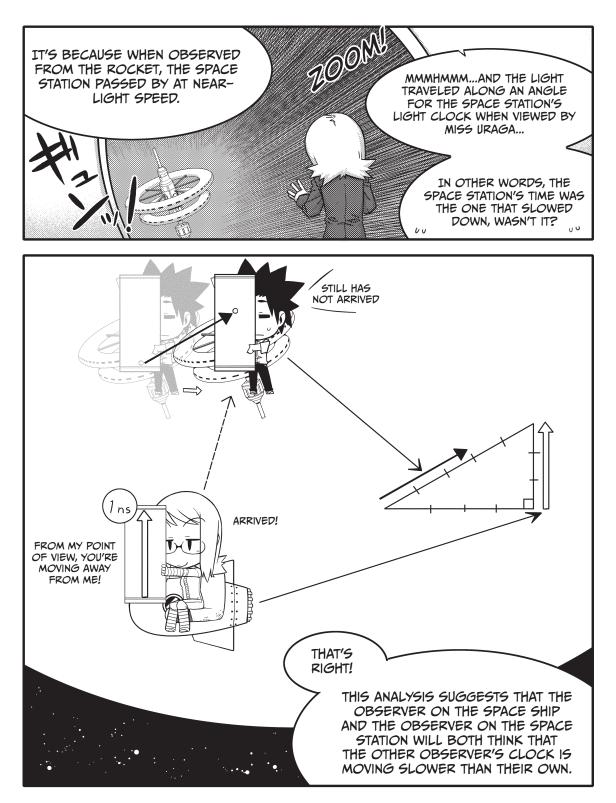




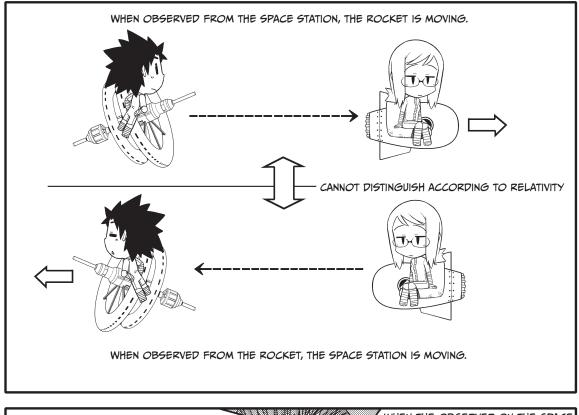


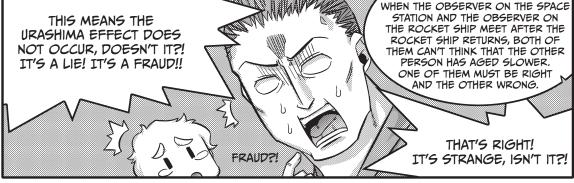


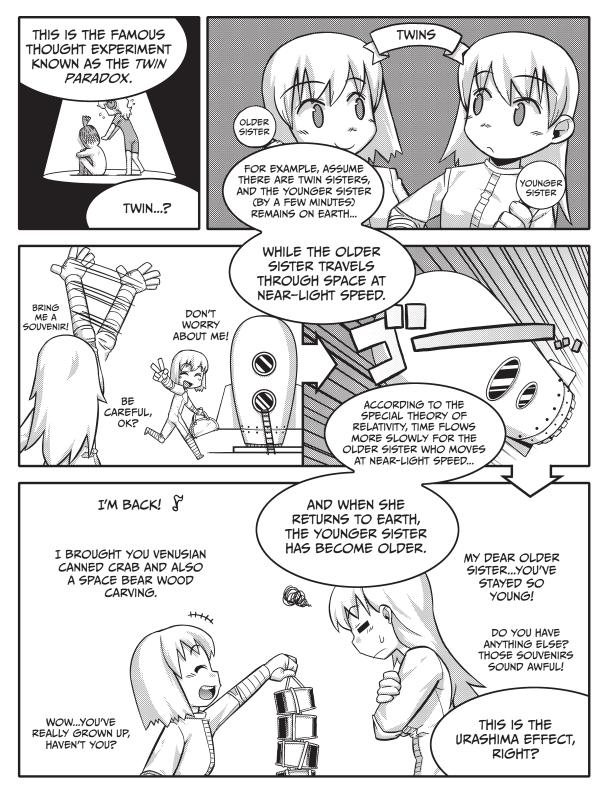


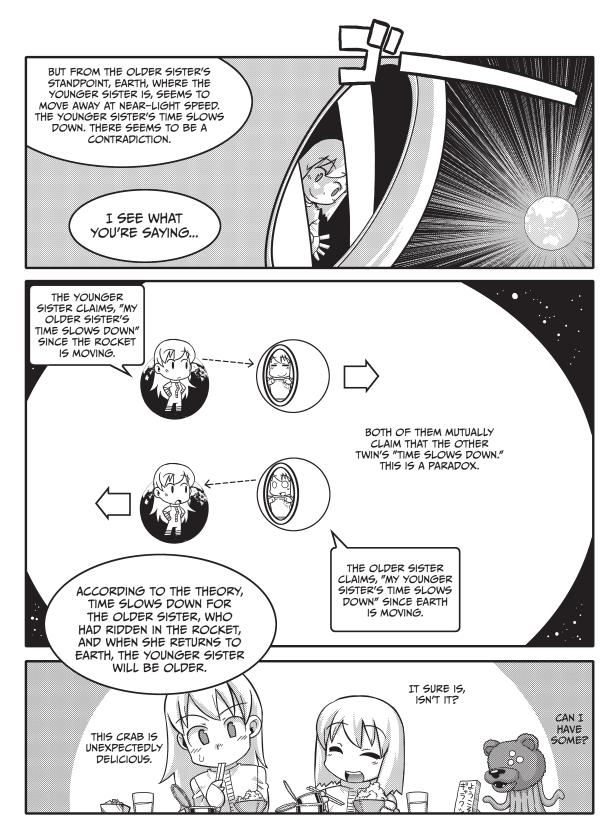


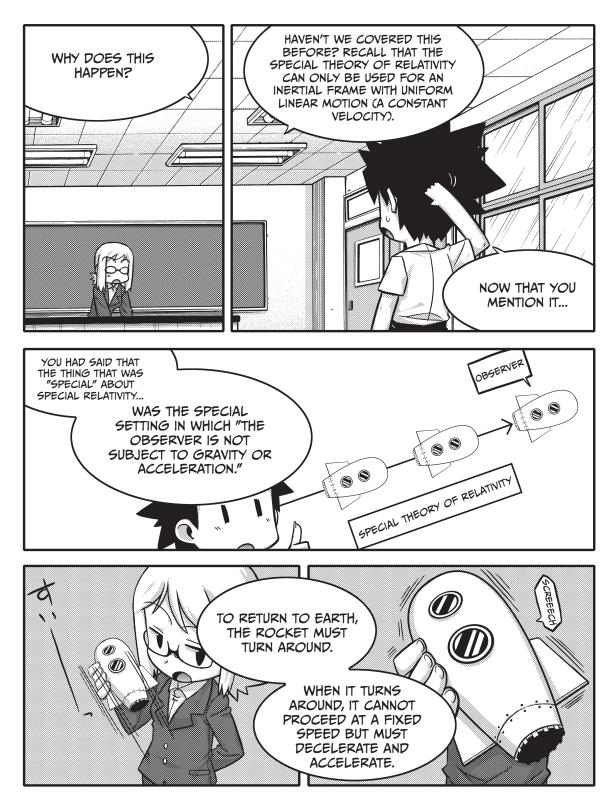


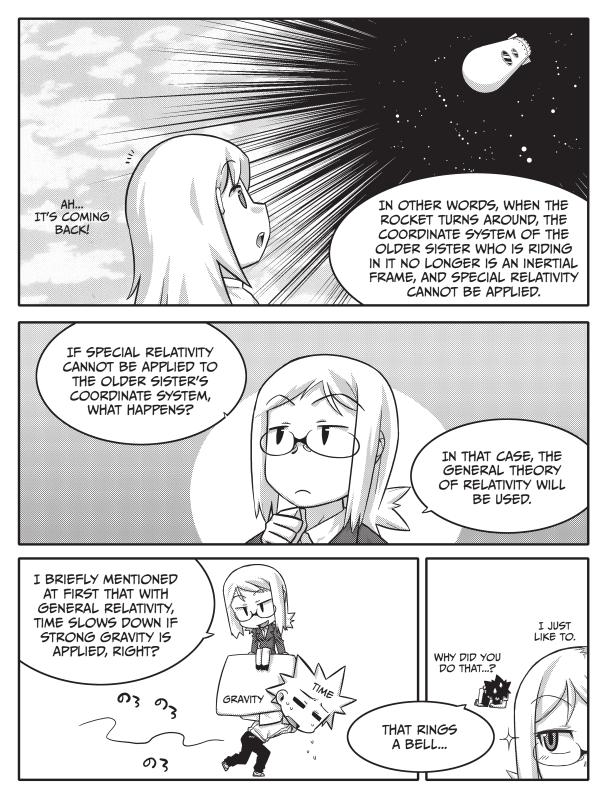


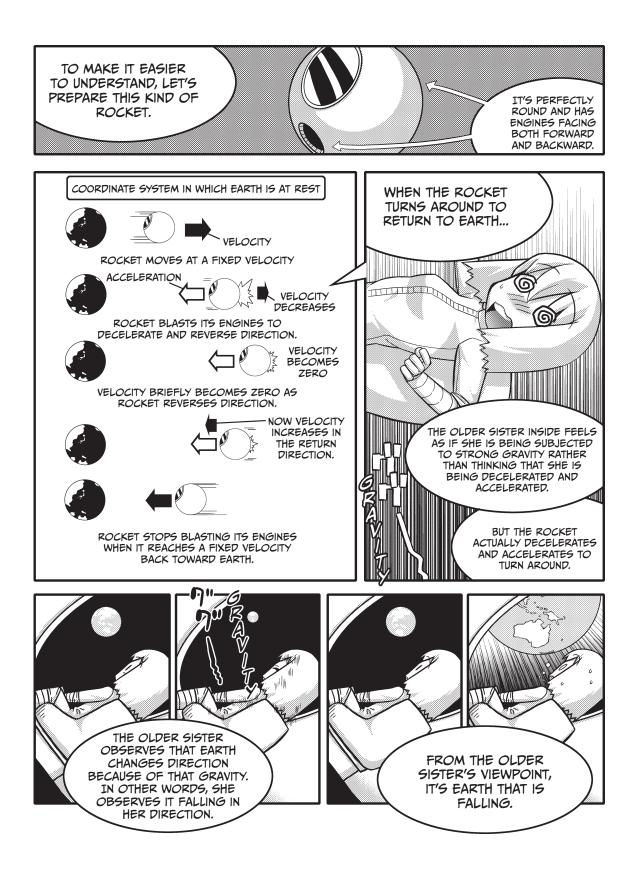
























USING THE PYTHAGOREAN THEOREM TO PROVE TIME DILATION

We learned that according to the theory of relativity, time slows down for an object that is moving at a velocity close to the speed of light. But how much does time slow? When we used the Pythagorean theorem earlier, we considered this question using a triangle. Now we can consider it using a formula.

Let *t* denote the amount of time that has passed according to Space-Station Man looking at the rocket's light clock and *t*' denote the amount of time that has passed according to Rocket Man looking at his own clock (see Figure 2-1).

t: Time observed by Space-Station Man

t': Time observed by Rocket Man

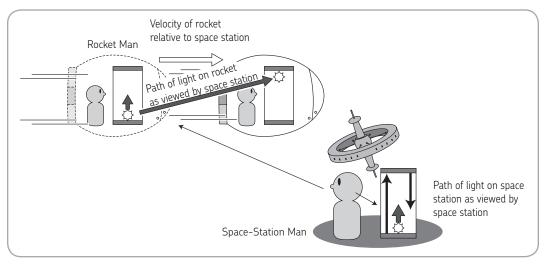


Figure 2-1: Rocket Man and Space-Station Man

When Rocket Man observes his own light clock, the light just goes up and down because the light clock is moving together with Rocket Man. Therefore, if *c* denotes the speed of light, when the light advances by the height of the light clock, it will have moved a distance of *ct'*.

Now if Space-Station Man observes the movement of the light in the rocket's light clock, the light, of course, moves at the speed of light *c* along an upward slanted path accompanying the movement of the rocket. That slanted line points towards the mirror (at the top) of the rocket's light clock. Measured using Space-Station Man's time *t*, that distance is *ct*. Similarly, since Space-Station Man sees the bottom of the rocket's light clock (from where the light was emitted) moving horizontally at the rocket's velocity *v*, the bottom will move by a distance of *vt* to the right in the time *t* that the light takes to reach the top.

This determines the three sides of a triangle (see Figure 2-2).

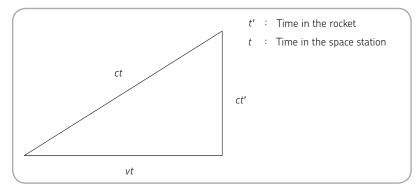


Figure 2-2: Distances moved by light as viewed on rocket and space station, expressed as sides of a right triangle

Therefore, from the Pythagorean theorem, we have $c^2t^2 = c^2t'^2 + v^2t^2$. Move the v^2 term to the left side of the equation:

$$(c^2 - v^2)t^2 = c^2t'^2$$

And switch the left and right sides:

$$c^{2}t'^{2} = (c^{2} - v^{2})t^{2}$$

Dividing by c^2 , we now have this:

$$t'^2 = \left(1 - \frac{v^2}{c^2}\right)t^2$$

Now take the square root of both sides and use the positive solution:

$$t' = \sqrt{1 - \frac{v^2}{c^2}} \times t$$

This is the relationship between Rocket Man's time *t* and Space-Station Man's time *t*.

Note that
$$t' < t$$
 since $\sqrt{1 - \frac{v^2}{c^2}} < 1$

A second (1 s) measured on the Rocket Man's light clock thus corresponds to a longer time measured by the Space-Station Man. So Space-Station Man sees the Rocket Man's clock tick off seconds at a slower rate than his own. In other words, time advances more slowly for Rocket Man than for Space-Station Man. It is also apparent by considering the term

$$\sqrt{1-\frac{v^2}{c^2}}$$

that this time-slowing effect is greater the closer v is to c.

With this formula, we can calculate the time dilation effect for an object moving at any relative speed.

HOW MUCH DOES TIME SLOW DOWN?

We have learned that time slows down for a moving object. Let's now determine exactly how much it slows down. We will use space travel as an example for this calculation.

As we saw above, the slowing of time is related to the velocity of the moving object. Recall that the closer the velocity of the moving object is to the speed of light, the greater the time-slowing effect becomes.

As a destination for our space travel, let's pick the star that is closest to our Sun— Alpha Centauri (see Figure 2-3). If we chose a destination inside our solar system, we wouldn't put the theory of relativity to the test because we can go to Mars or Venus in just a few years even with current technology, going nowhere near the speed of light.

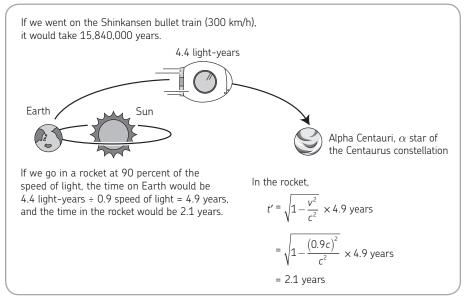


Figure 2-3: Traveling to Alpha Centauri

Alpha Centauri (the α star of the Centaurus constellation) is 4.4 light-years away from Earth. A light-year, which is the distance that light travels in one year, is approximately 9,460,800,000,000 km.^{*} Traveling 4.4 light-years on the Shinkansen bullet train (at 300 km/h) would take approximately 15,840,000 years. But if we fly that distance to Alpha Centauri at 90 percent of the speed of light, the journey will take us only 2.1 years. This is despite the fact that 4.9 years will pass on Earth!

If astronauts were sent out from Earth to Alpha Centauri, even if they returned as soon as they got there, news reports on Earth would follow them for approximately 10 years, but when their families welcomed them home, they would have aged only 4.2 years.

This relativistic time slowing effect is even greater if you move faster.

To understand this relationship, let's consider traveling from the Milky Way galaxy, which contains our Sun, to the Andromeda galaxy (M31), which is near the Milky Way. The Andromeda galaxy is visible in a dark, clear winter sky as a faint smudge in the Andromeda constellation. It is approximately 2,500,000 light-years away. Because light from the galaxy takes 2,500,000 years to reach us, the Andromeda galaxy that we see now is actually the galaxy 2,500,000 years ago. Even if an explosion occurred in the Andromeda Galaxy today, we wouldn't see it until 2,500,000 years later; the light would still take that long to reach us.

If astronauts travel to the Andromeda galaxy at 99.999999999 percent of the speed of light, 11.2 years will pass for the one-way trip in the spaceship, but nearly 2,500,000 years will pass on Earth! Therefore, when the spaceship returns, although the astronauts will have aged 22.4 years, the people on Earth who will greet them will be from a time 5,000,000 years later.

$$t' = \sqrt{1 - \frac{v^2}{c^2}} \times t$$

$$t' = \sqrt{1 - \left(\frac{0.9999999999c}{c}\right)^2} \times t$$

t' = 11.2 years

* Let's do the math: $\frac{300,000 \text{ km}}{1 \text{ second}} \times \frac{60 \text{ seconds}}{1 \text{ minuter}} \times \frac{60 \text{ minutes}}{1 \text{ hourr}} \times \frac{24 \text{ hours}}{1 \text{ day}} \times \frac{365 \text{ days}}{1 \text{ year}}$ = 9,460,800,000 km!

For the astronauts on the spaceship, the round-trip to Andromeda will take 22.4 years, as shown in Figure 2-4. For the people on Earth, the astronauts will return only 22.4 years older than when they left 5 million years earlier!

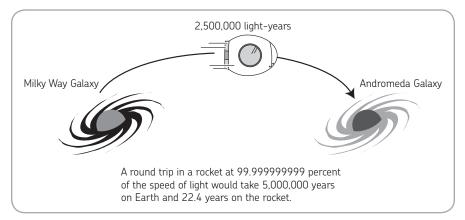


Figure 2-4: Traveling to the Andromeda galaxy



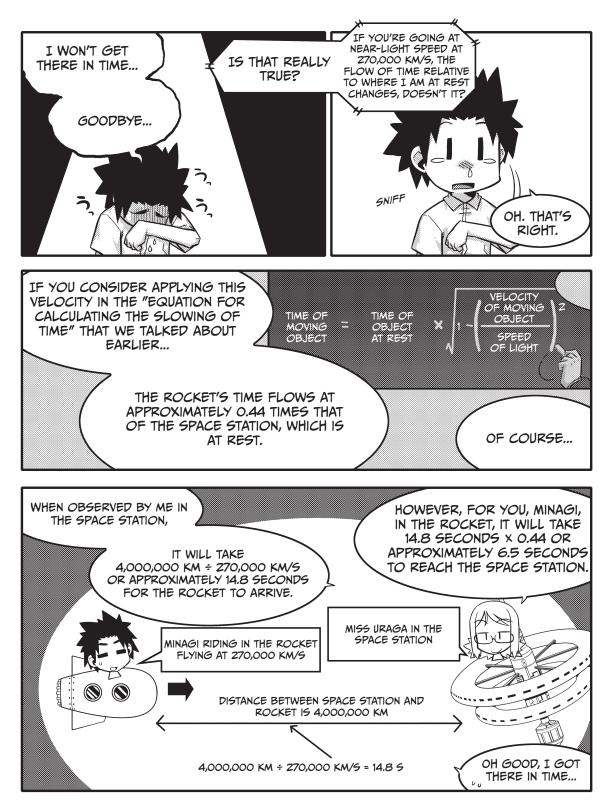
THE FASTER AN OBJECT MOVES, THE SHORTER AND HEAVIER IT BECOMES?

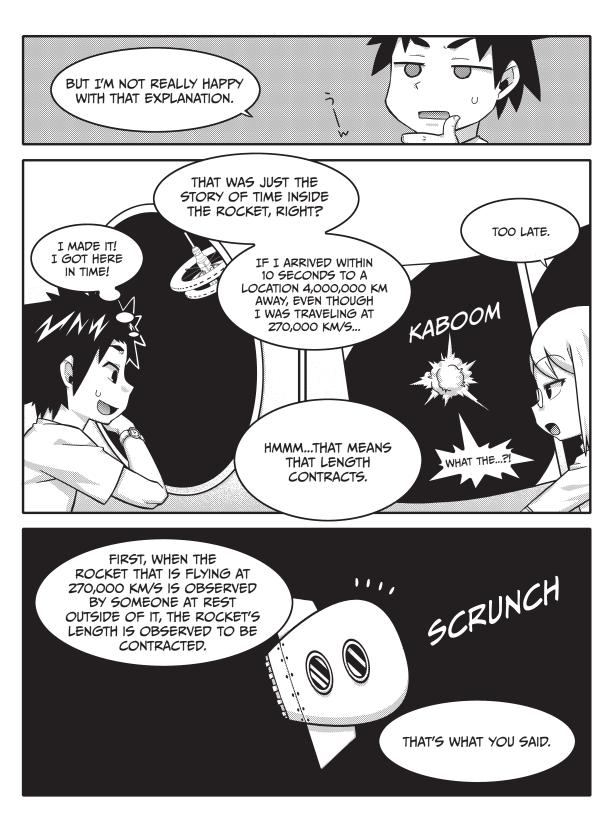




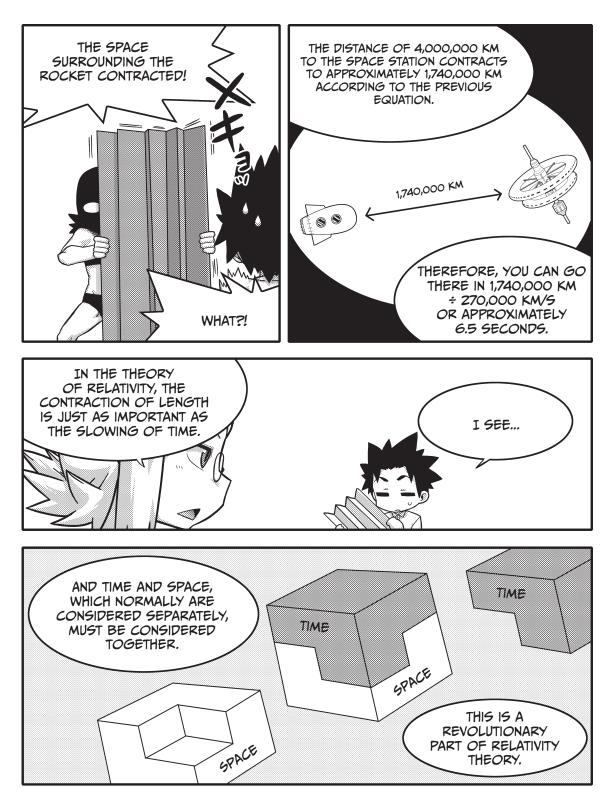




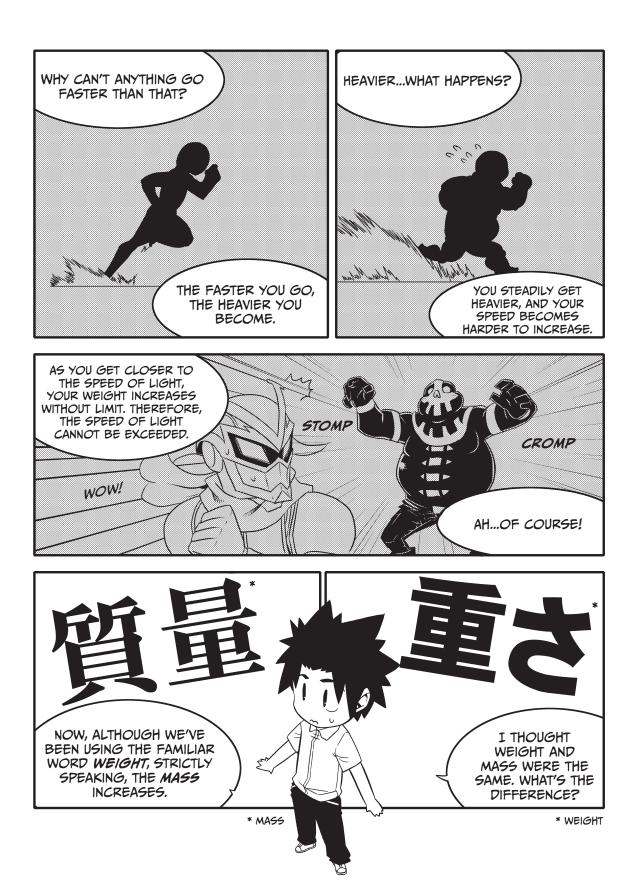


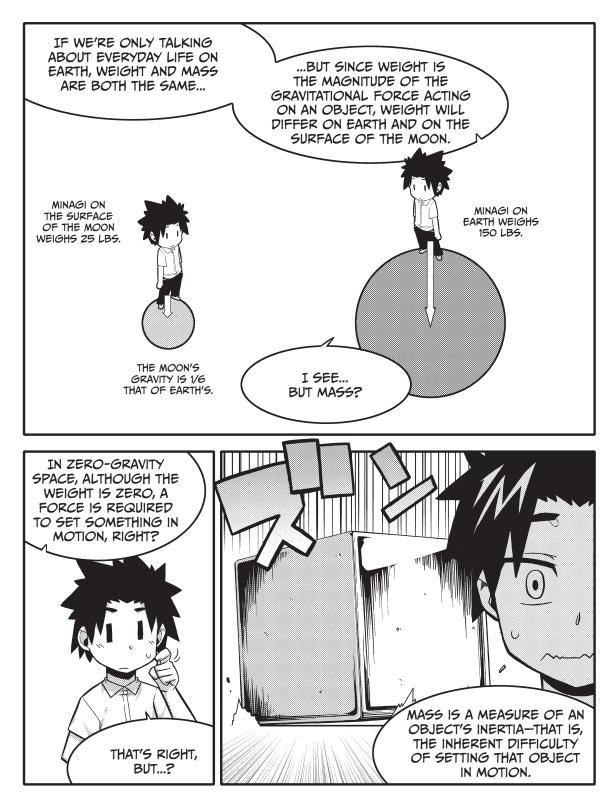






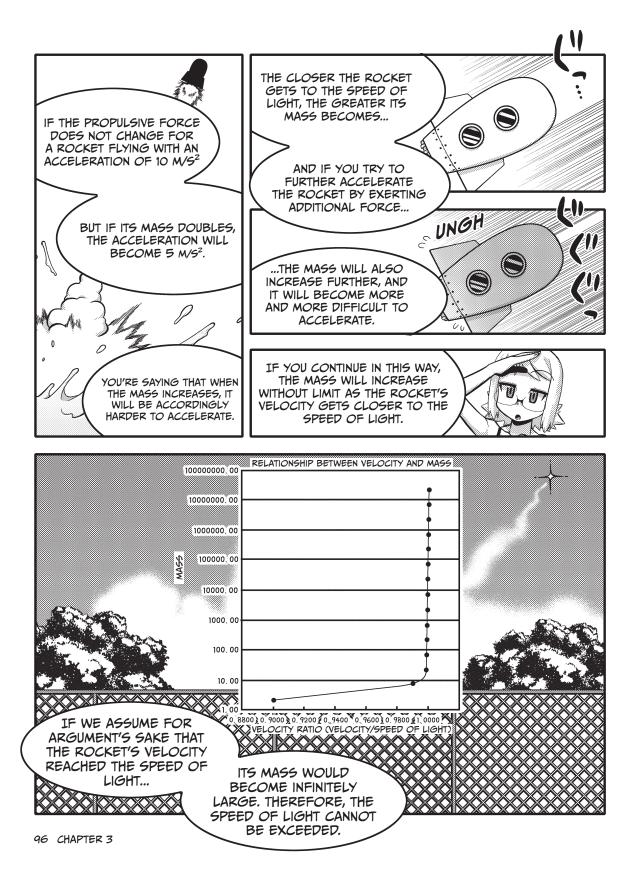






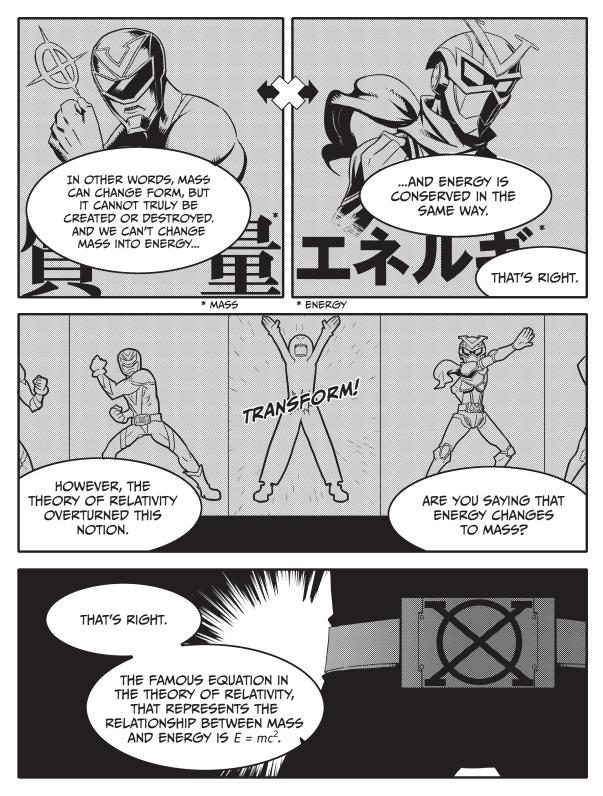


THE FASTER AN OBJECT MOVES, THE SHORTER AND HEAVIER IT BECOMES? 95

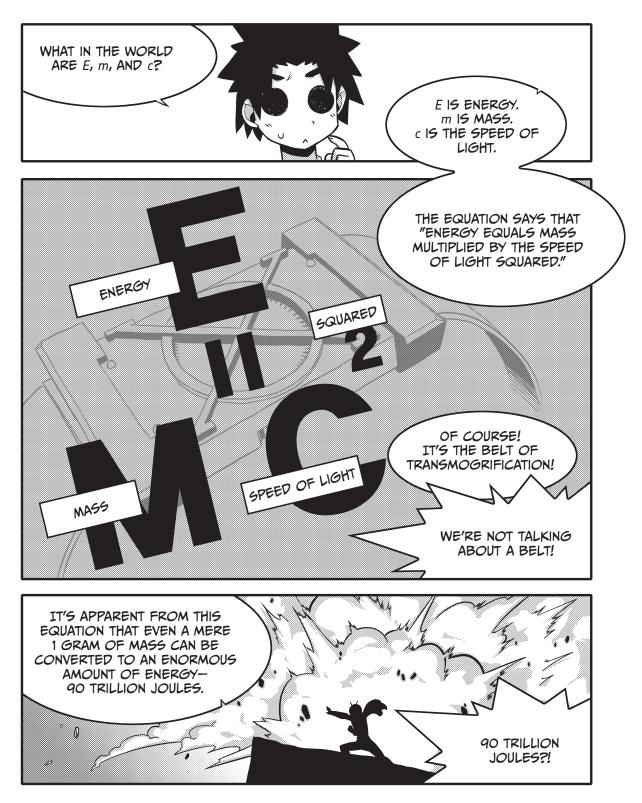


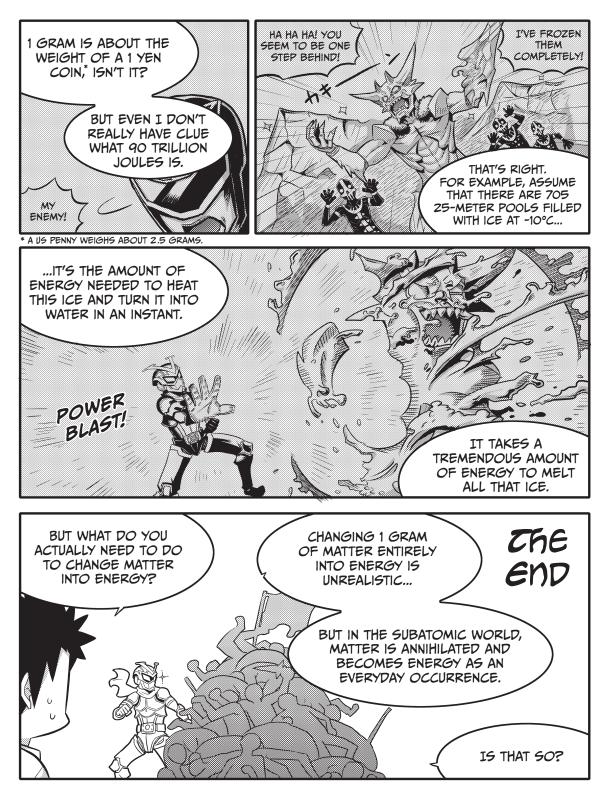


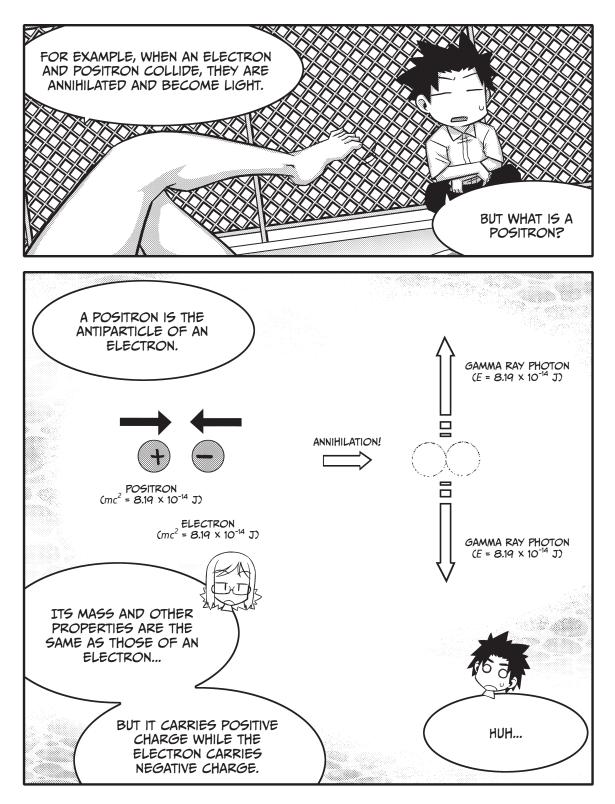


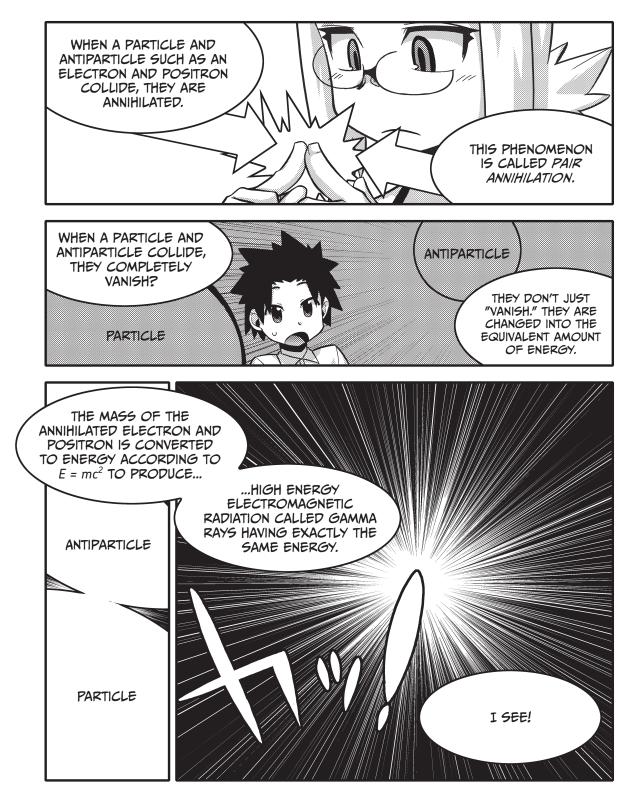


THE FASTER AN OBJECT MOVES, THE SHORTER AND HEAVIER IT BECOMES? 99

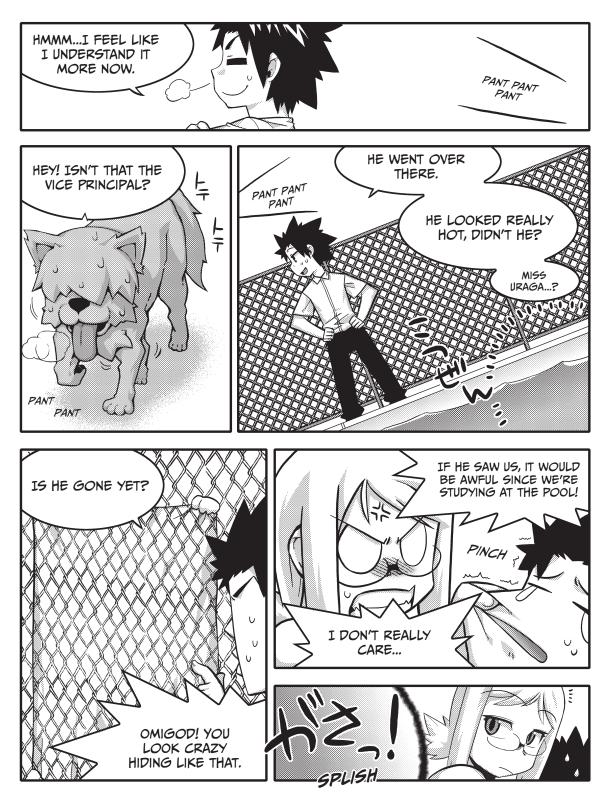








THE FASTER AN OBJECT MOVES, THE SHORTER AND HEAVIER IT BECOMES? 103





THE FASTER AN OBJECT MOVES, THE SHORTER AND HEAVIER IT BECOMES? 105

USING AN EQUATION TO UNDERSTAND LENGTH CONTRACTION (LORENTZ CONTRACTION)

Let's use an equation to see how length contracts.

In this case, let's assume that a rocket is flying at a constant velocity v (see Figure 3-1).

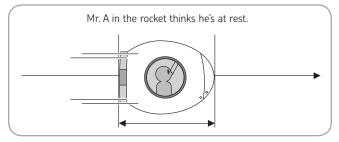


Figure 3-1: A person riding in the rocket measures the positions of the front and back ends of the rocket.

When the person riding in the rocket measures the positions of the front and back ends of the rocket, he finds the front end is at position x'_2 , and the back end is at position x'_1 . Therefore, the rocket's length is $l_0 = x'_2 - x'_1$.

Now what happens if this situation is observed from outside the rocket, for example, from a space station as in Figure 3-2?

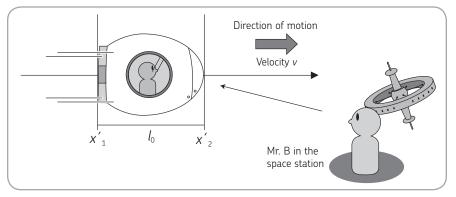


Figure 3-2: The rocket viewed from a space station

To calculate the contraction in the length of a rocket as it moves past an observer at close to the speed of light, let's consider two points in the rocket's frame of reference: x'_1 at the front of the ship and x'_2 at the back of the ship. Using the Lorentz transformation, introduced in "Wait a Second—What Happens with the Addition of Velocities?" on page 48, we can calculate how an observer who watches the ship pass measures the points at the front x_1 and back of the ship x_2 , in his reference frame. The length that the observer on the outside of the ship measures will be shorter than length that the astronaut measures. This effect, *relativistic length contraction*, comes from the contraction of space at speeds close to the speed of light.

Using the Lorentz transformation $x' = \frac{x - vt}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$, we calculate the following positions:

$$x_{1}' = \frac{x_{1} - vt_{1}}{\sqrt{1 - \left(\frac{v}{c}\right)^{2}}}$$
$$x_{2}' = \frac{x_{2} - vt_{2}}{\sqrt{1 - \left(\frac{v}{c}\right)^{2}}}$$

 $x_2 = \frac{1}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$

If we let $l = x_2 - x_1$ represent the rocket's length as observed from outside the rocket, then since

$$l_{0} = x_{2}' - x_{1}' = \frac{x_{2} - vt_{2}}{\sqrt{1 - \left(\frac{v}{c}\right)^{2}}} - \frac{x_{1} - vt_{1}}{\sqrt{1 - \left(\frac{v}{c}\right)^{2}}} = \frac{(x_{2} - x_{1}) - (t_{2} - t_{1})v}{\sqrt{1 - \left(\frac{v}{c}\right)^{2}}}$$

is measured at the same time, $t_1 - t_2 = 0$ because $t_2 = t_1$, so l_0 is calculated as follows:

$$l_{0} = \frac{(x_{2} - x_{1}) - (t_{2} - t_{1})v}{\sqrt{1 - \left(\frac{v}{c}\right)^{2}}} = \frac{(x_{2} - x_{1})}{\sqrt{1 - \left(\frac{v}{c}\right)^{2}}} = \frac{l}{\sqrt{1 - \left(\frac{v}{c}\right)^{2}}}$$

The length of the spaceship measured from outside the spaceship,

$$l = l_0 \sqrt{1 - \left(\frac{v}{c}\right)^2}$$

is therefore less than the length l_0 of the same spaceship measured from inside the ship. We know this to be true because the coefficient

$$\sqrt{1-\left(\frac{v}{c}\right)^2} < 1$$

due to the fact that the speed of the spaceship must be slower than the speed of light (v < c).

MUONS WITH EXTENDED LIFE SPANS

Our discussion of time slowing down and length contracting is not just a theoretical proposition. The slowing of time is observed every day.

High-energy elementary particles called *cosmic rays* are raining down on Earth all day, every day. When those cosmic rays collide with molecules in Earth's upper atmosphere, muons are generated with a certain probability. A *muon* is a type of elementary particle that is similar to an electron. The life span of a muon is approximately 2 millionths of a second in a laboratory on the ground at rest. Therefore, when a muon is produced in the upper atmosphere, several tens to several hundreds of kilometers from the ground, it would fly only 300,000 km/s $\times 2/1,000,000$ s = 0.6 km, even if it were flying at a velocity extremely close to the speed of light. Based on these calculations, it should not reach the surface of Earth. But muons are observed on the surface of Earth! This seemingly impossible event occurs because the muon's life span is extended according to the special theory of relativity (see Figure 3-3). The extension of the lifetime of muons by time dilation has been verified in the laboratory by accelerating muons to near the speed of light.

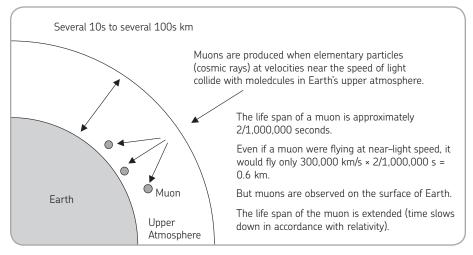


Figure 3-3: Muon life span

Let's now apply the concept of length contraction to the example of the muon. In the reference frame of the muon, the muon's lifetime is not extended—it is still 2 millionths of a second. From the reference frame of the muon, Earth is rushing toward the particle at close to the speed of light. The length between Earth and the muon contracts, however, as shown in Figure 3-4. And because the distance between Earth and the muon contracts, the muon reaches the planet's surface within its lifetime

Both the dilation of time from Earth's perspective and the contraction of length from the muon's perspective are consistent. In this way, the dilation of time and the contraction of length change together according to the theory of relativity.

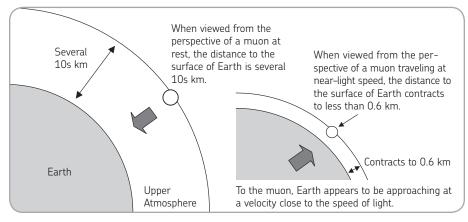


Figure 3-4: Distance contracts as well as time.

MASS WHEN MOVING

Now let's consider how the mass of an object is related to its velocity according to relativity. Let's start by reviewing the laws of motion. Before relativity was understood, the Galilean transformation and Newton's law of motion were used to describe motion.

GALILEAN TRANSFORMATION

The Galilean transformation describes the relationship between coordinate systems moving at velocity v:

$$x' = x - vt$$
 and $t' = t$

where x' and t' represent position and time, respectively, in one system and x, v, and t represent position, velocity, and time, respectively, in the other system.

NEWTON'S SECOND LAW OF MOTION

Newton's second law of motion is represented as follows:

$$f = ma = m \frac{d^2 x}{dt^2}$$
 ,

where *f* represents force, *m* represents mass, *a* represents acceleration, and acceleration can be considered the second derivative of displacement with respect to time:

$$a = \frac{d^2 x}{dt^2}$$

According to Galileo's principle of relativity, the laws of physics operate exactly the same way, whether measured while at rest or while moving. In other words, whether you toss a ball in the air inside an elevator that is at rest or inside an elevator that is moving at a constant velocity, the ball will move up and down and return to your hand in the same way.

Now, let's look at the laws of motion in two different reference frames and verify that the laws of physics do not change when we don't take into account relativity. We'll look at the laws of motion in a reference frame that is at rest, in which the position of the ball is measured x, and in the reference frame of the moving elevator, in which the position of the ball is measured x' (see Figure 3–5).

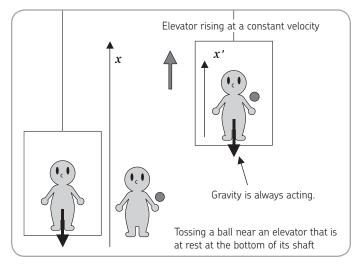


Figure 3-5: Elevator moving at a constant velocity

In this case, the velocity of the ball, which is moving in the x' direction inside the elevator, is given by

$$\frac{dx'}{dt'}$$

If we substitute the Galilean transformation x' = x - vt here, we obtain the following:

$$\frac{dx'}{dt'} = \frac{d}{dt'} (x - vt) = \frac{dx}{dt'} - v \frac{dt}{dt'} = \frac{dx}{dt} - v$$

We used the relationship $\frac{dt'}{dt} = 1$ here, because dt' = dt.

If we differentiate again, we obtain the following:

$$\frac{d^2x'}{dt'^2} = \frac{d}{dt'} \left(\frac{dx}{dt} - v \right) = \frac{d^2x}{dt^2}$$

Since the only force acting on the ball here is gravity, if we let *g* denote gravity, we have this:

$$g = f = ma = m\frac{d^2x}{dt^2}$$

Note that g is a force, not accleration due to gravity in this equation. Now, if we let a' denote the acceleration inside the elevator, which is moving at a constant velocity, and let f' denote force, we have

$$m\frac{d^2x}{dt^2} = m\frac{d^2x'}{dt'^2} = ma' = f' = g$$

and the form of the equation of motion is unchanged. Since the form of the equation of motion does not change, the laws of physics remain the same.

Let's consider how the situation described above changes when we consider relativity and replace the Galilean transformation with the Lorentz transformation.

LORENTZ TRANSFORMATION

$$x' = \frac{x - vt}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$
$$t' = \frac{t - \frac{v}{c^2}x}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$

As we have shown in this chapter, time and space become intermixed within the framework of relativity. Therefore, when describing the coordinates of an object, it is not sufficient to give its position in three-dimensional space (x, y, z)—we must also consider its time (t). Because the units of position and time are different (meters and seconds, respectively), we multiply time by the speed of light so that we can describe the coordinate position of an object with four dimensions that all have the same unit: length. The following shows that two reference frames are mutually transformed; note that time and space are transformed together.

$$(ct, x, y, z) \leftrightarrow (ct', x', y', z')$$

If we use this thinking to extend the equation of motion so that its form does not change even when a Lorentz transformation is used, it is apparent that mass, which had been considered constant in Newtonian mechanics, is represented in a form similar to the Lorentz transformation:

$$m = \frac{m_0}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$

Here, m_0 , which is called the *rest mass* or *invariant mass*, is the mass measured in a coordinate system at rest (v = 0).

RELATIONSHIP BETWEEN ENERGY AND MASS

In the same way, if we consider energy in a form that matches the Lorentz transformation, it is represented as follows:

$$E = \frac{m_0 c^2}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$

If we substitute the earlier relationship,

$$m = \frac{m_0}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$

we derive the famous relationship between energy and mass: $E = mc^2$.

Now when $|x| \ll 1$, if we use the approximation $(1 + x)^{\alpha} \approx 1 + \alpha x$ under the condition

$$\left(\frac{v}{c}\right)^2 \ll 1$$

(velocity *v* is sufficiently small compared with the speed of light), we obtain the following:

$$E = \frac{m_0 c^2}{\sqrt{1 - \left(\frac{v}{c}\right)^2}} = m_0 c^2 \left[1 - \left(\frac{v}{c}\right)^2\right]^{-\frac{1}{2}} \cong m_0 c^2 \left[1 + \frac{1}{2} \left(\frac{v}{c}\right)^2\right] = m_0 c^2 + \frac{1}{2} m_0 v^2$$

The total energy of an object is the sum of its kinetic energy

$$E_{k}=\frac{1}{2}m_{0}v^{2}$$

and its rest energy ($E = mc^2$). This means that even when an object is not moving, it has energy associated with its mass. Rest energy $E = mc^2$ is similar in form to the kinetic energy in Newtonian mechanics, where

$$E_k = \frac{1}{2}m_0 v^2$$

and $m_0 c^2$ is called the *rest energy*.

DOES LIGHT HAVE ZERO MASS?

The equation that we derived above for the mass of an object in motion,

$$m = \frac{m_0}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$

tells us that as the velocity of an object approaches the speed of light, its energy approaches infinity (see Figure 3-6). Therefore, the only way that light can exist (without having infinite energy) is if its mass is 0.

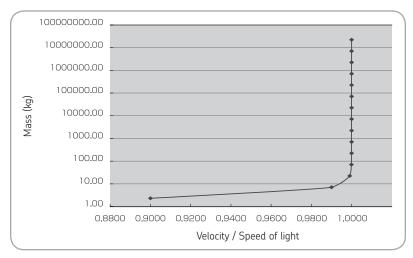


Figure 3-6: Relationship between mass and velocity



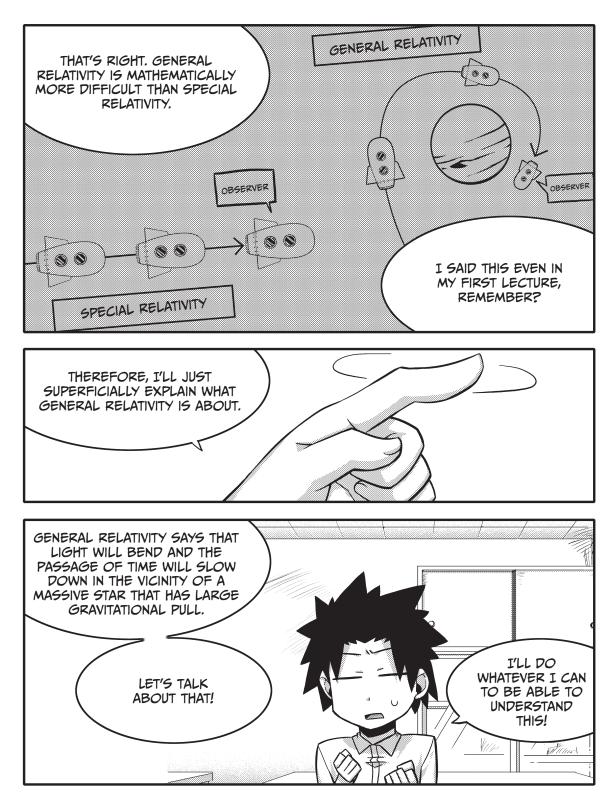
WHAT IS GENERAL RELATIVITY?

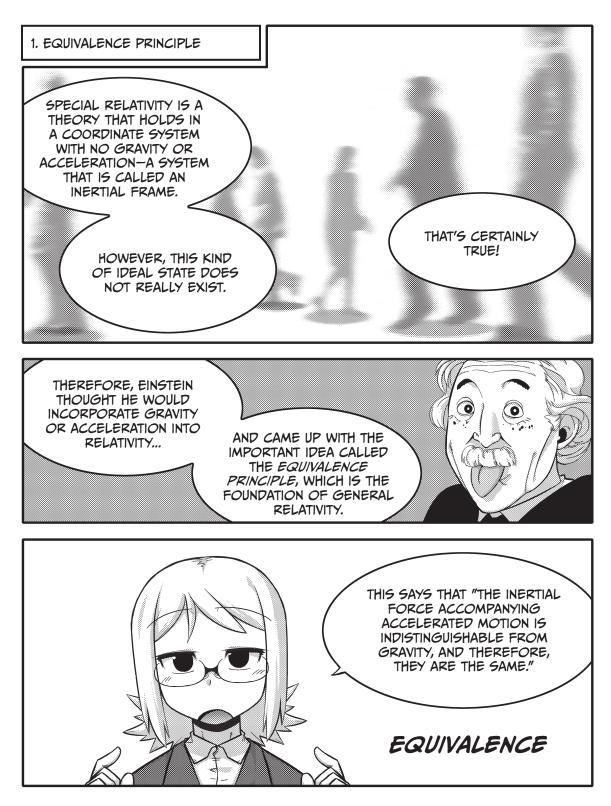
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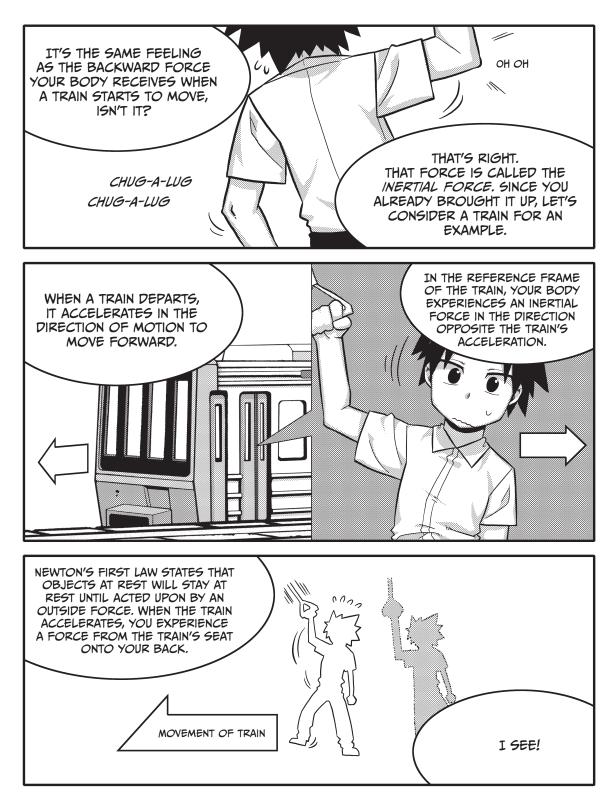


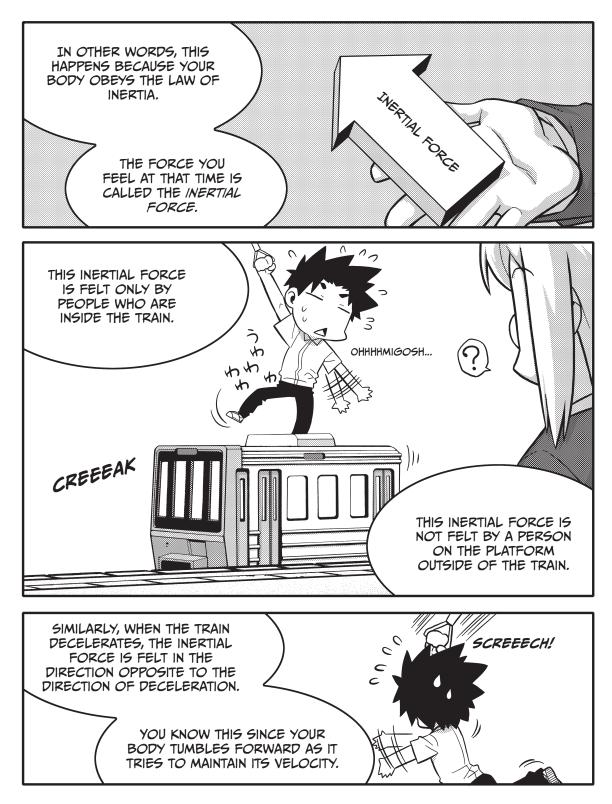


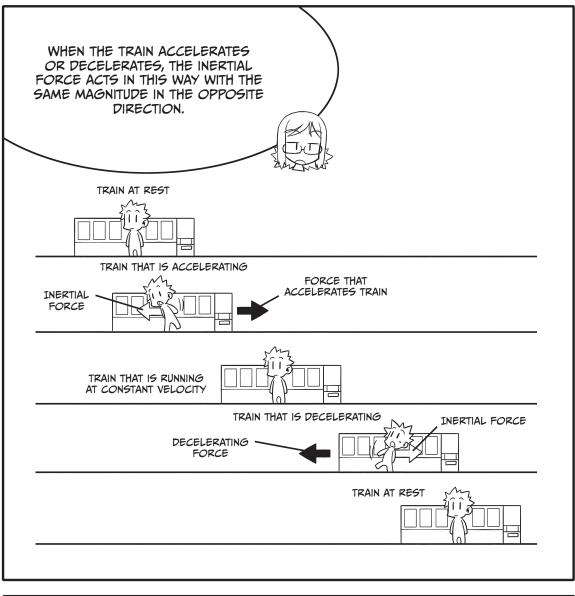




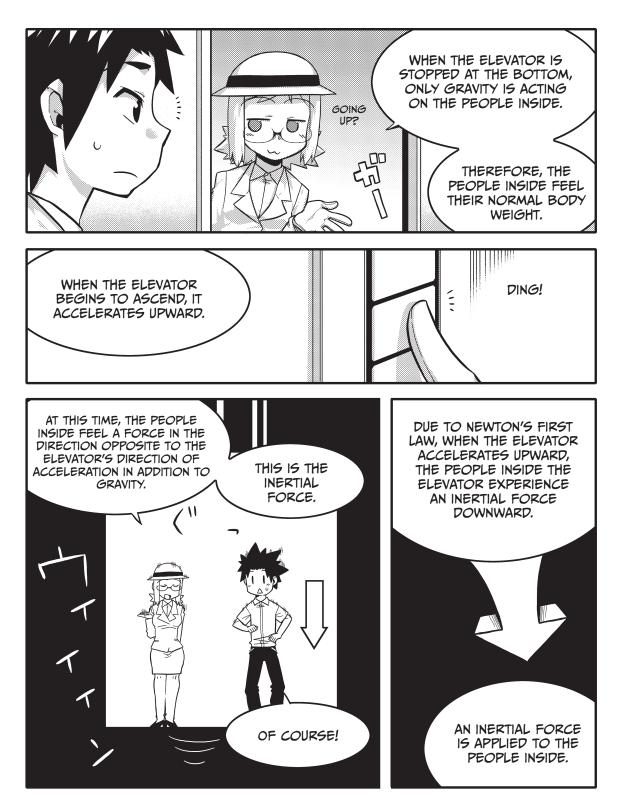


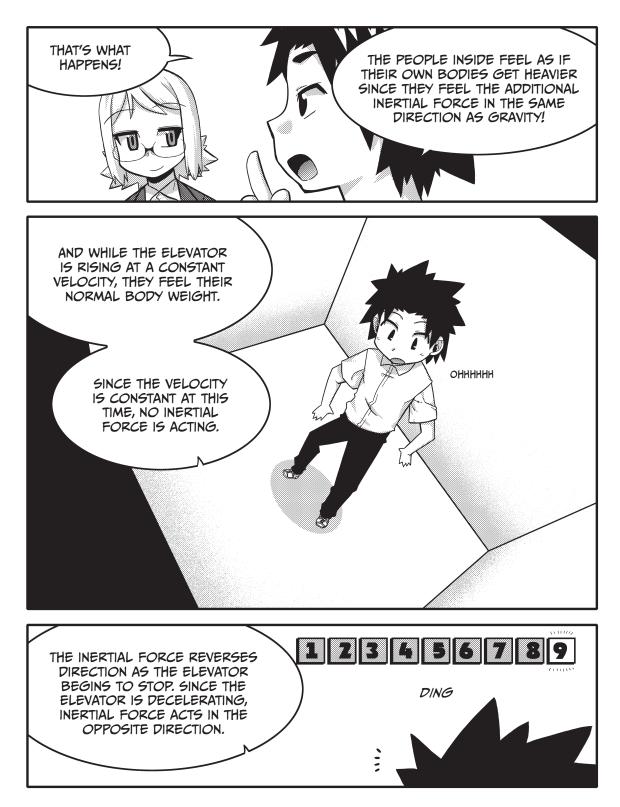


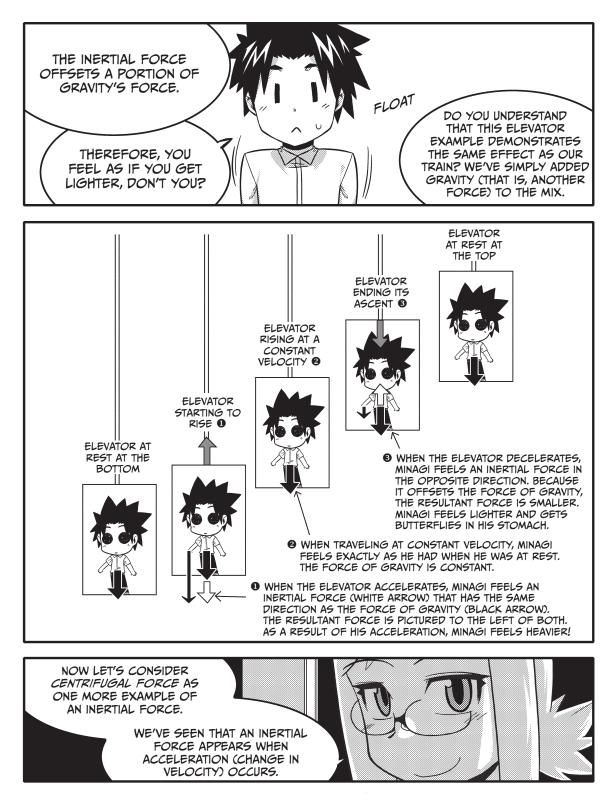


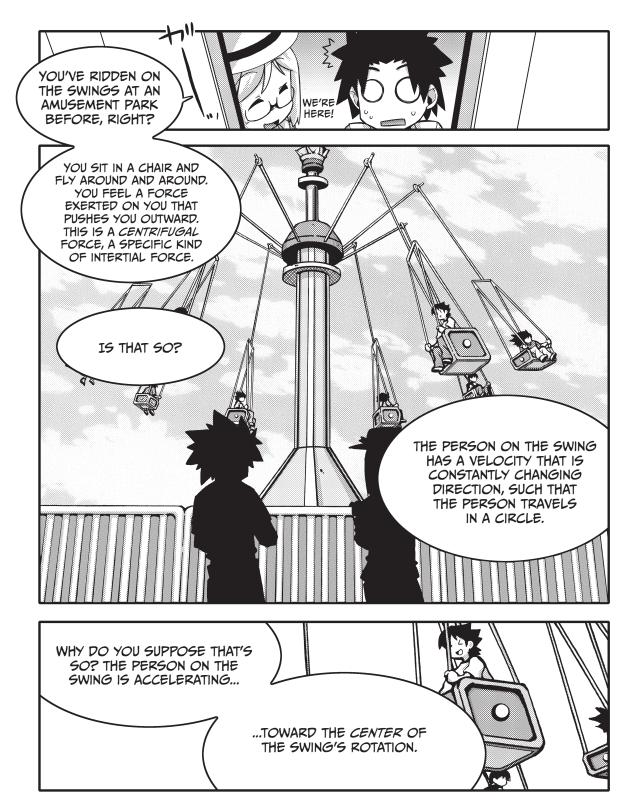


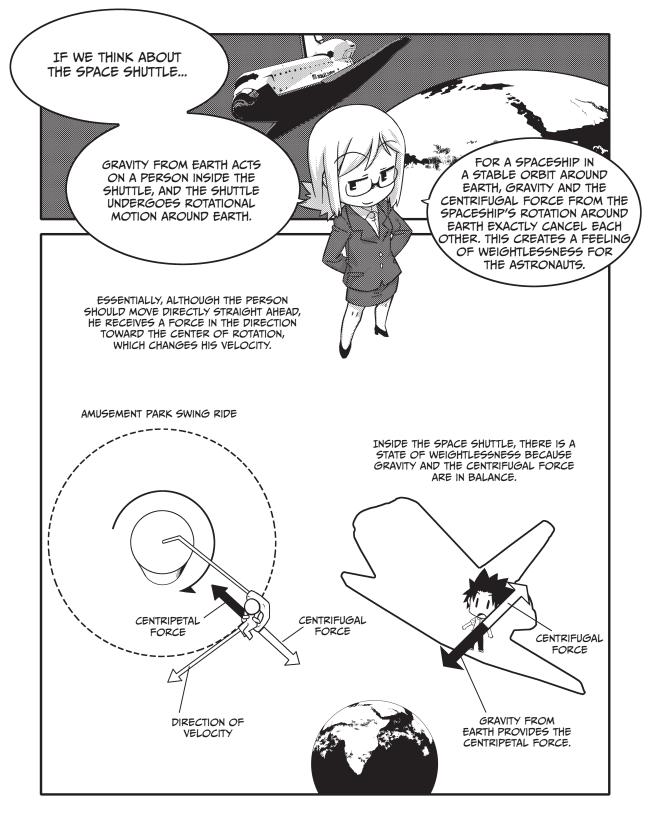


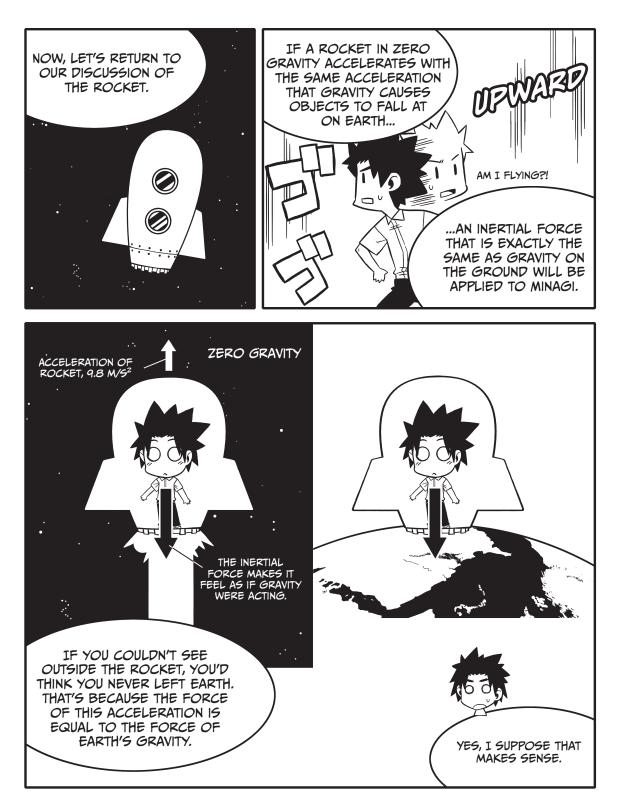


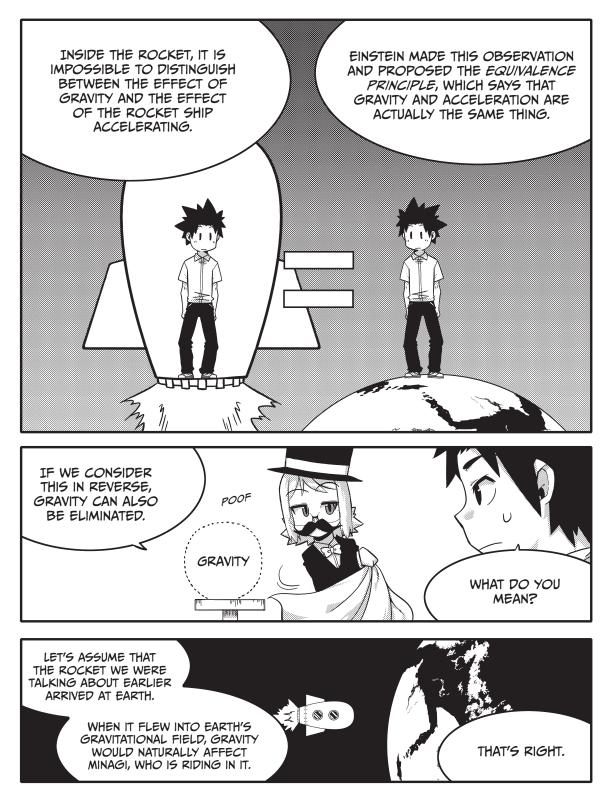


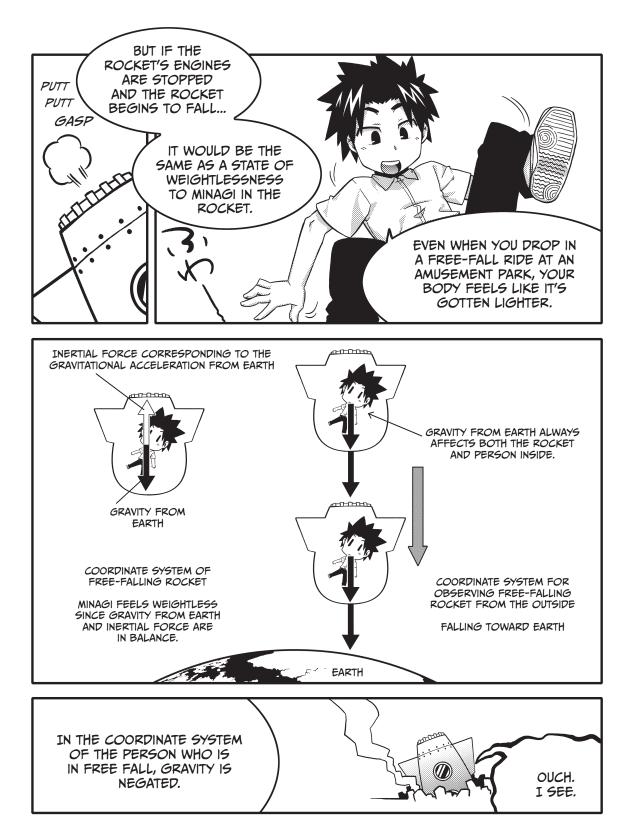


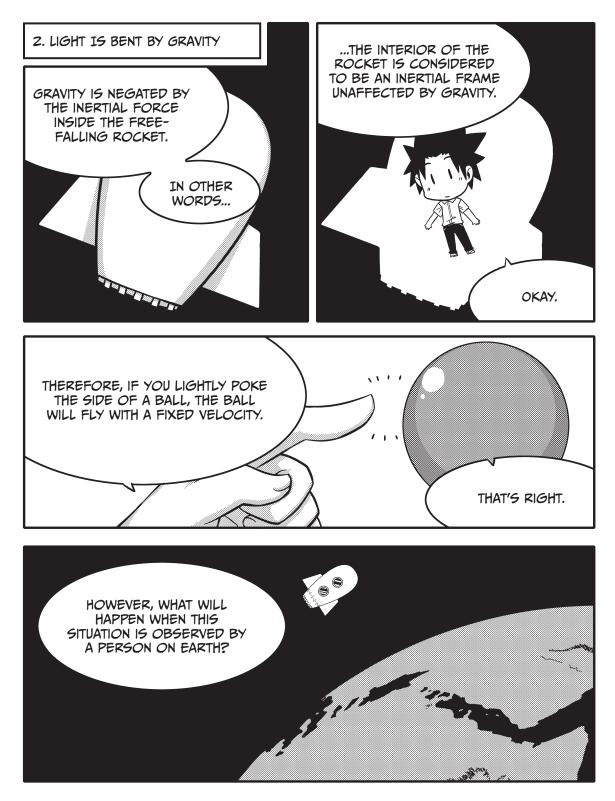


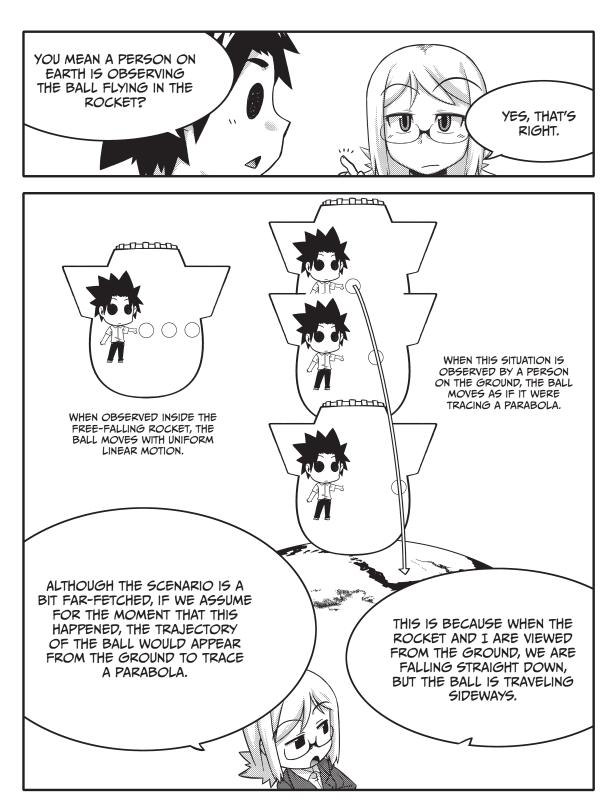


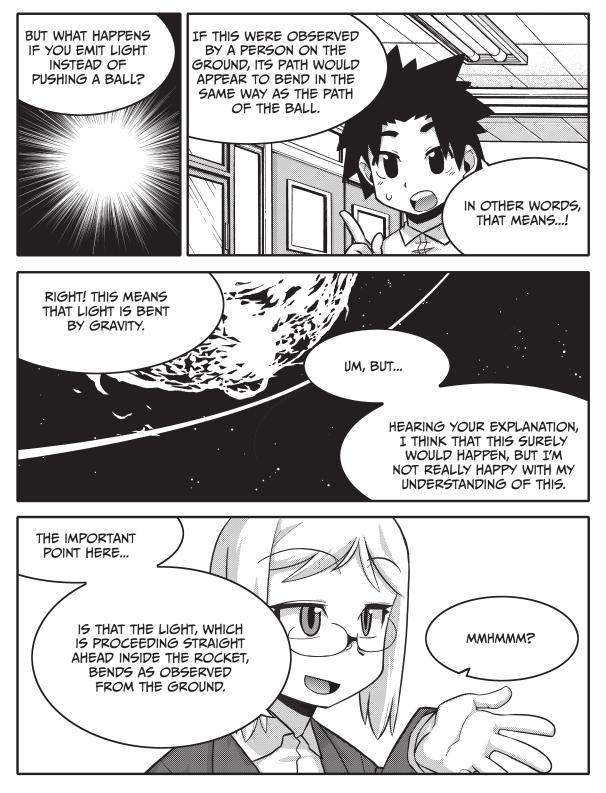


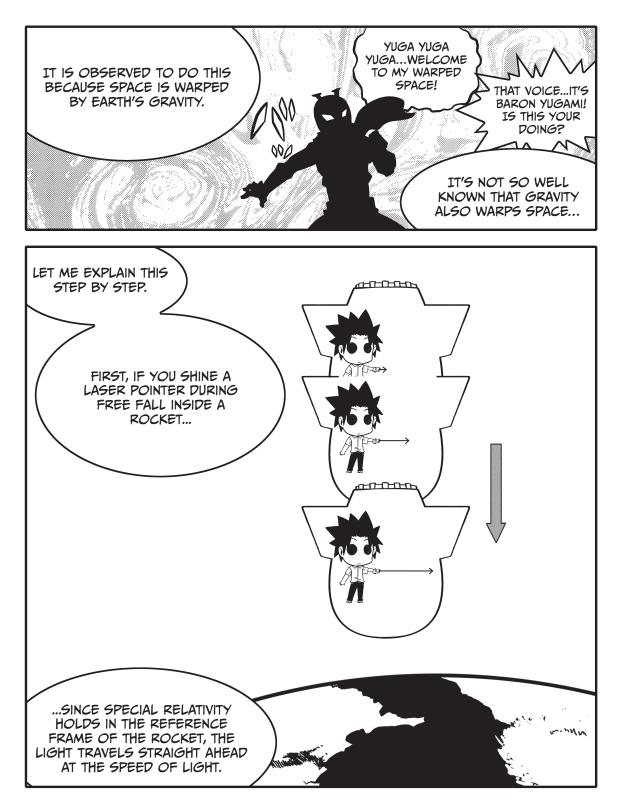


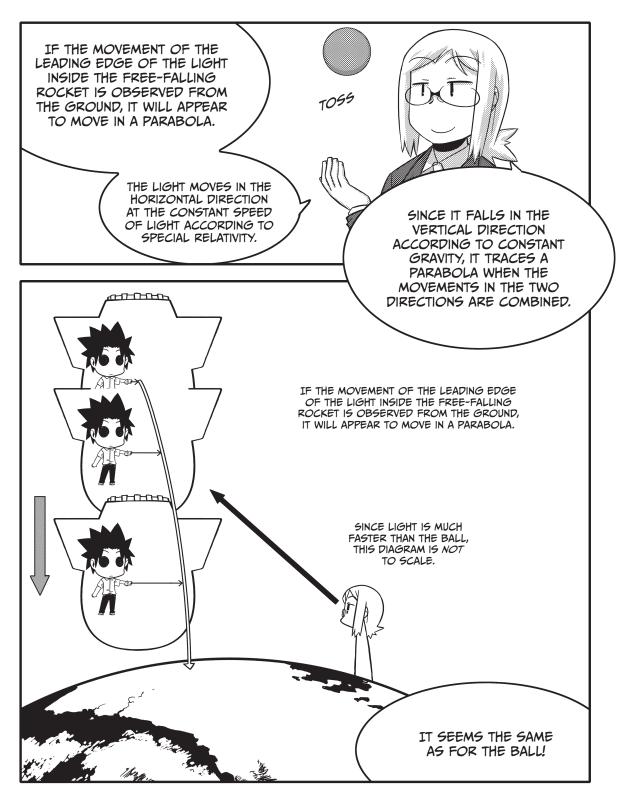












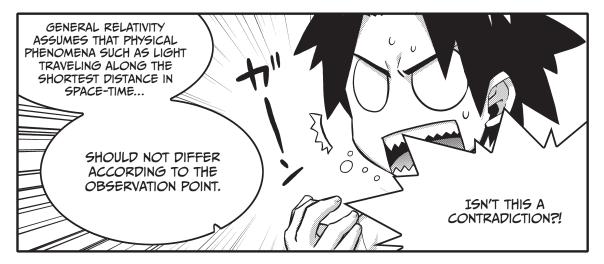
INSIDE THE FREE-FALLING ROCKET, THE LIGHT TRAVELS THE SHORTEST DISTANCE IN SPACE IN THE INERTIAL FRAME...OR, IN OTHER WORDS, IN A STRAIGHT LINE.

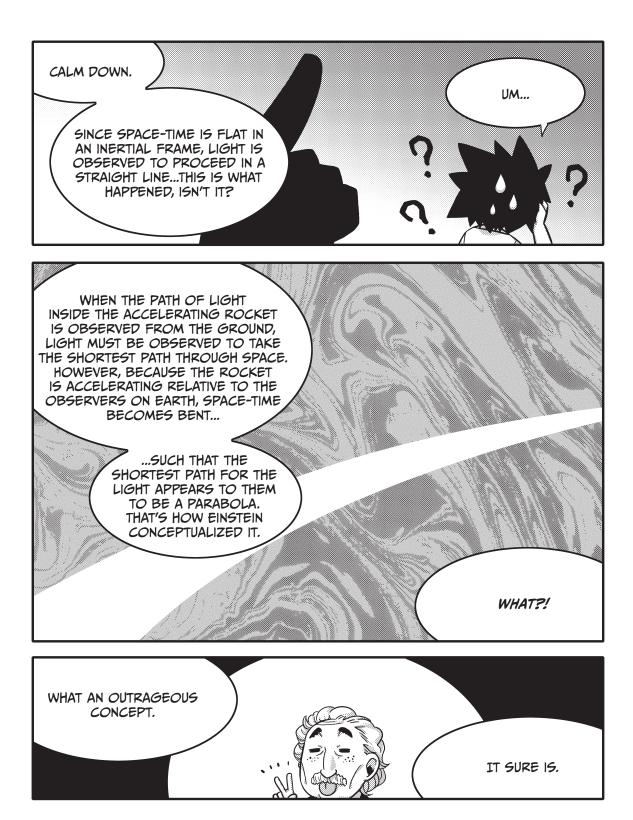
AND WHEN THE INTERIOR OF THE FREE-FALLING ROCKET IS OBSERVED FROM THE GROUND... THE LIGHT BENDS.

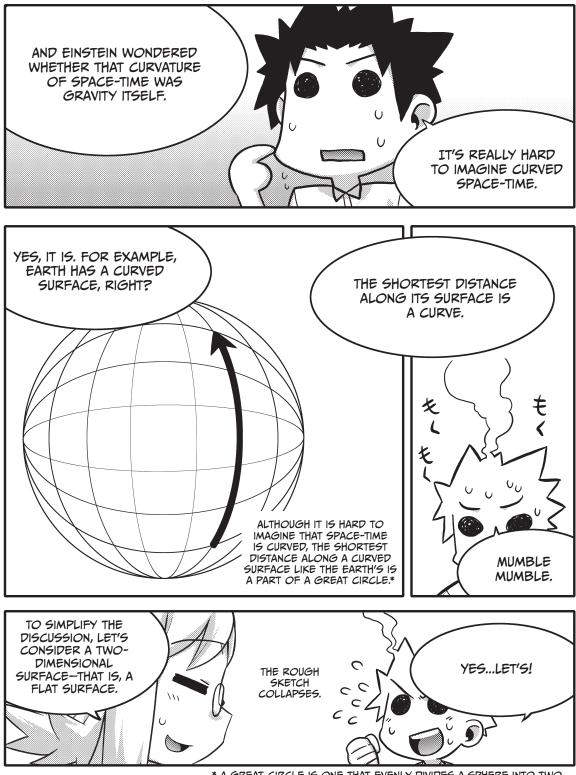
> IT CERTAINLY APPEARS TO BE A ROUNDABOUT PATH THAT IS NOT TRAVELING ALONG THE SHORTEST DISTANCE.

DOESN'T THE WAY IN WHICH THE LIGHT TRAVELS APPEAR TO DIFFER WHEN IT IS OBSERVED FROM INSIDE THE ROCKET AND WHEN IT IS OBSERVED FROM THE GROUND?

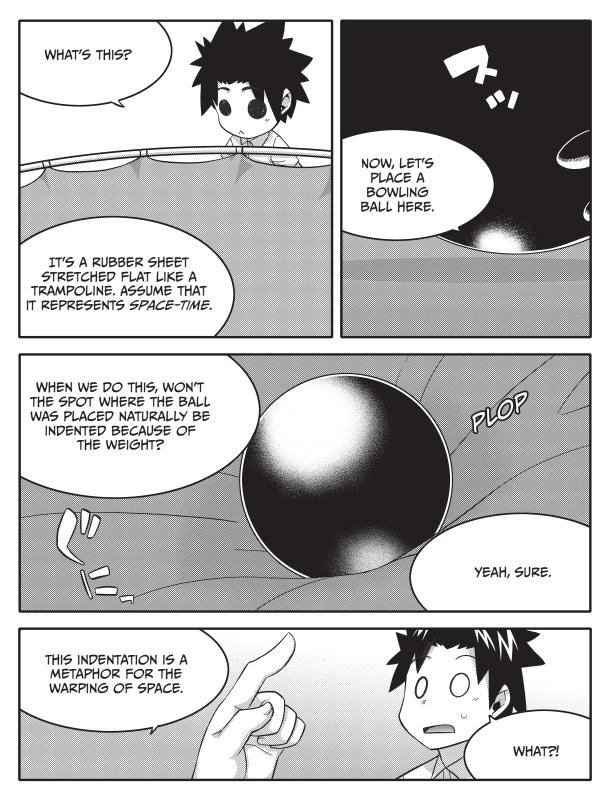


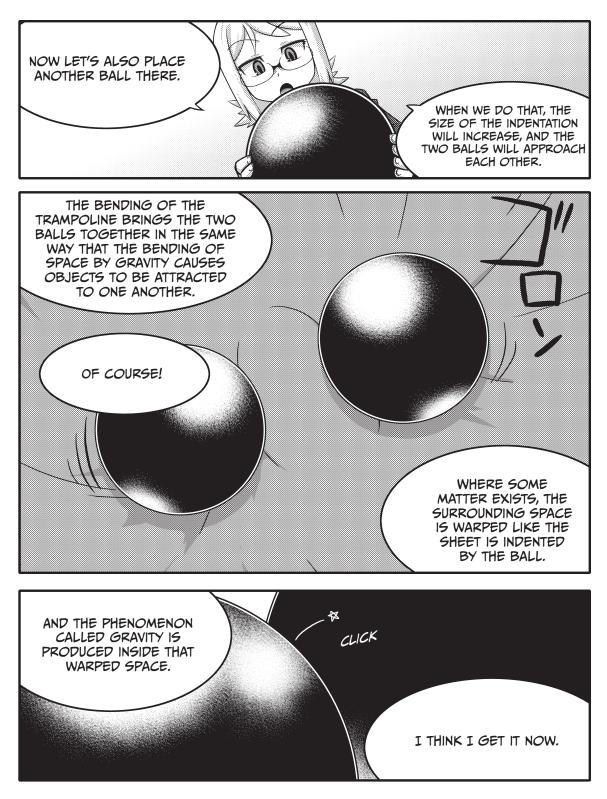




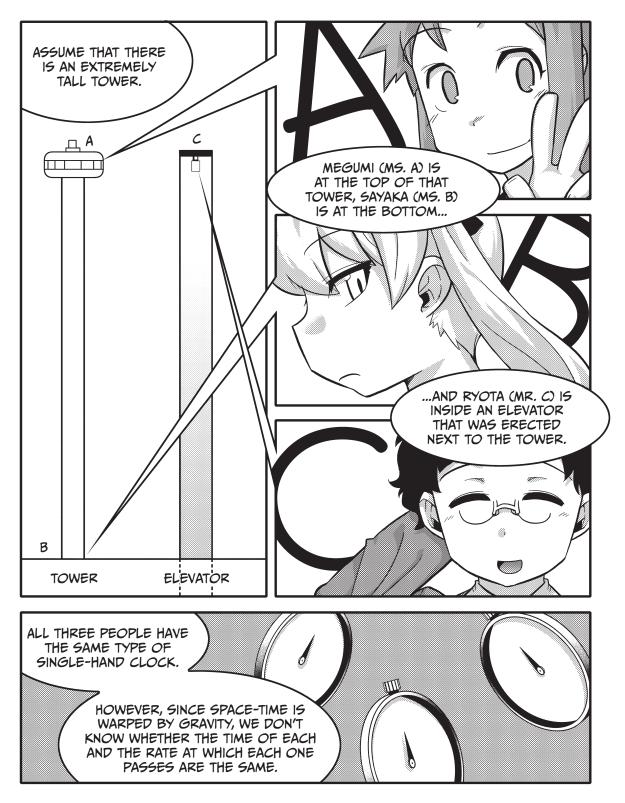


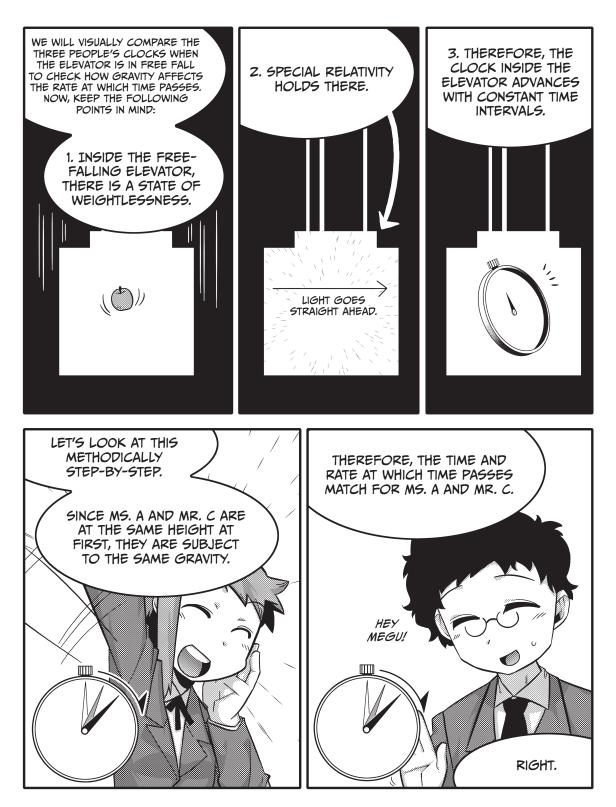
* A GREAT CIRCLE IS ONE THAT EVENLY DIVIDES A SPHERE INTO TWO EQUAL HALVES. THE SHORTEST DISTANCE BETWEEN TWO POINTS ON A SPHERE IS ALWAYS A PORTION (OR A MINOR ARC) OF GREAT CIRCLE.



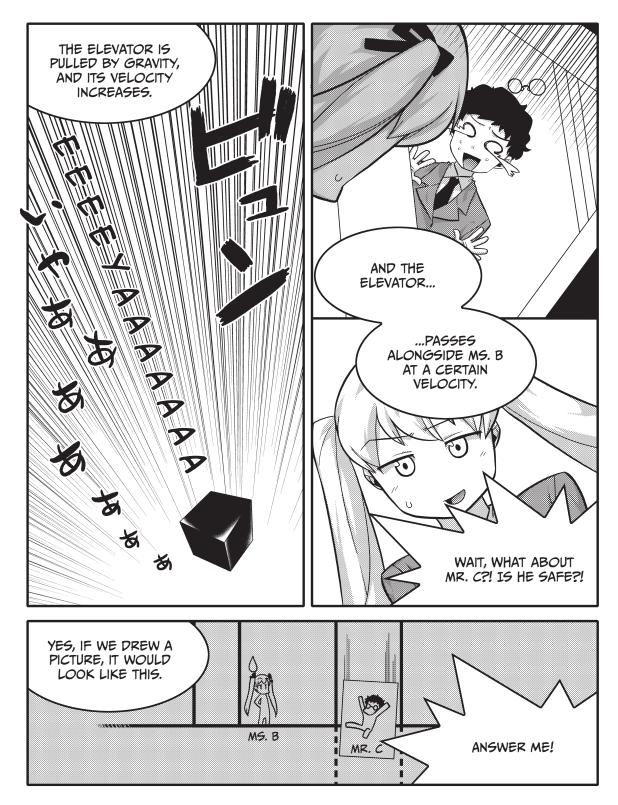


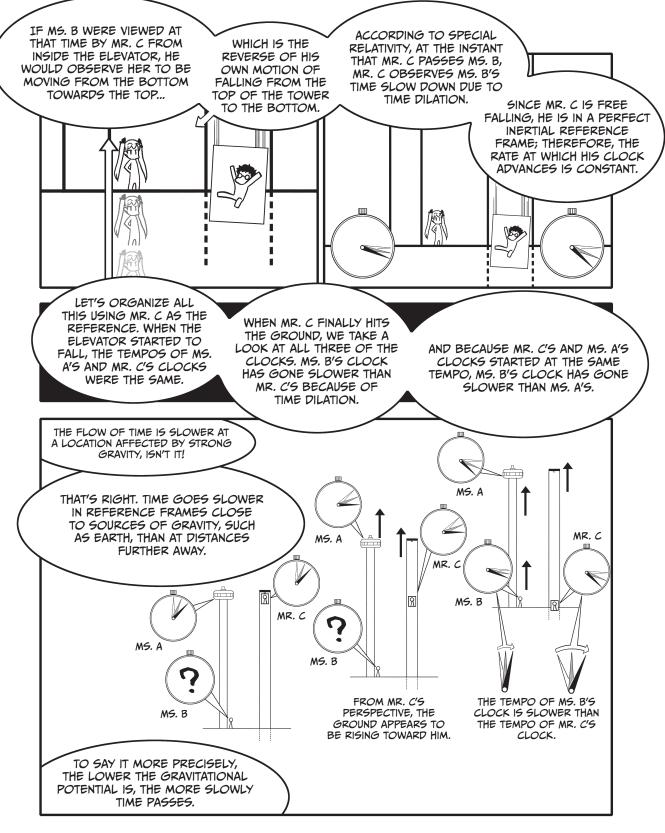




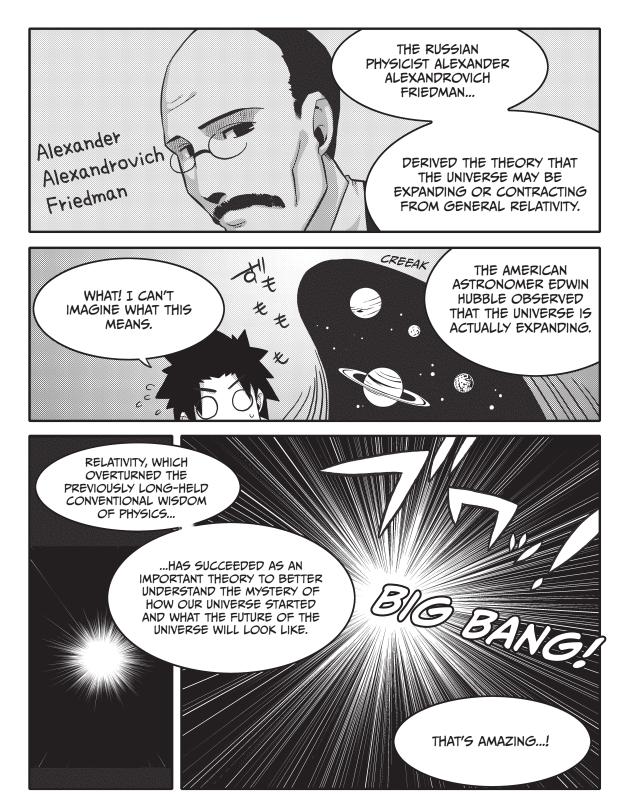


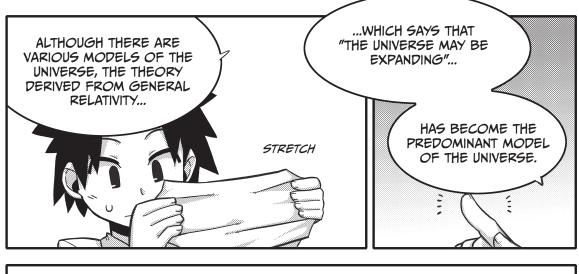




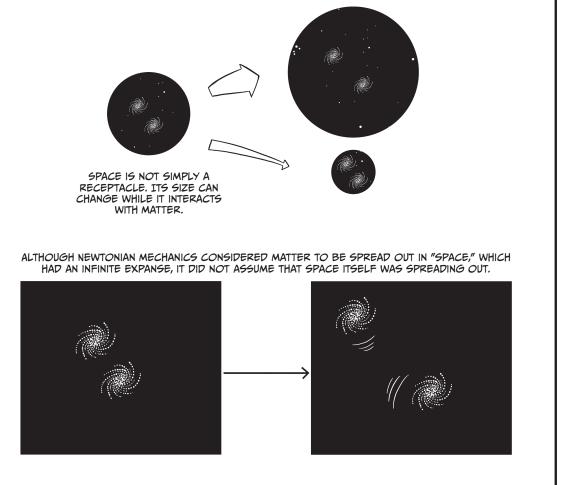


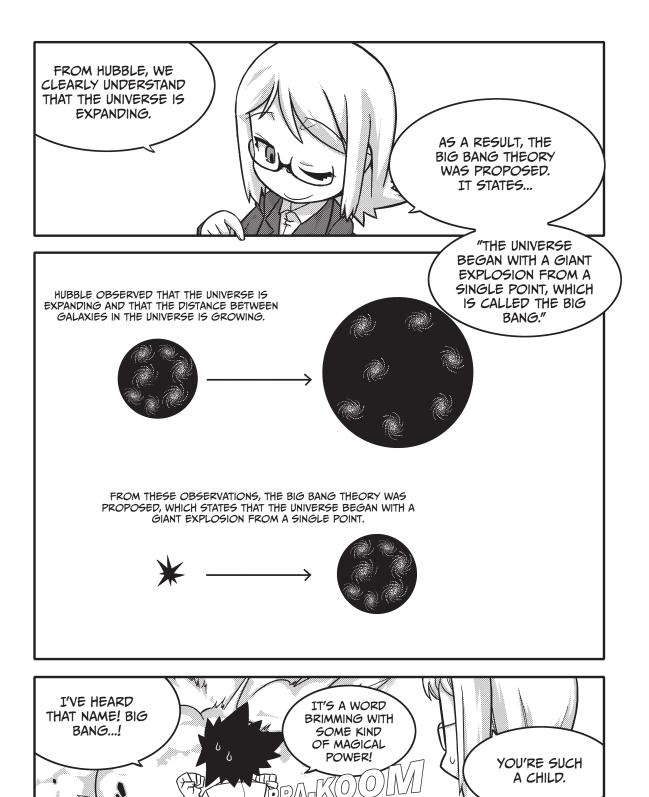






IT IS KNOWN THAT THE UNIVERSE MAY BE EXPANDING OR CONTRACTING AS SPACE-TIME.



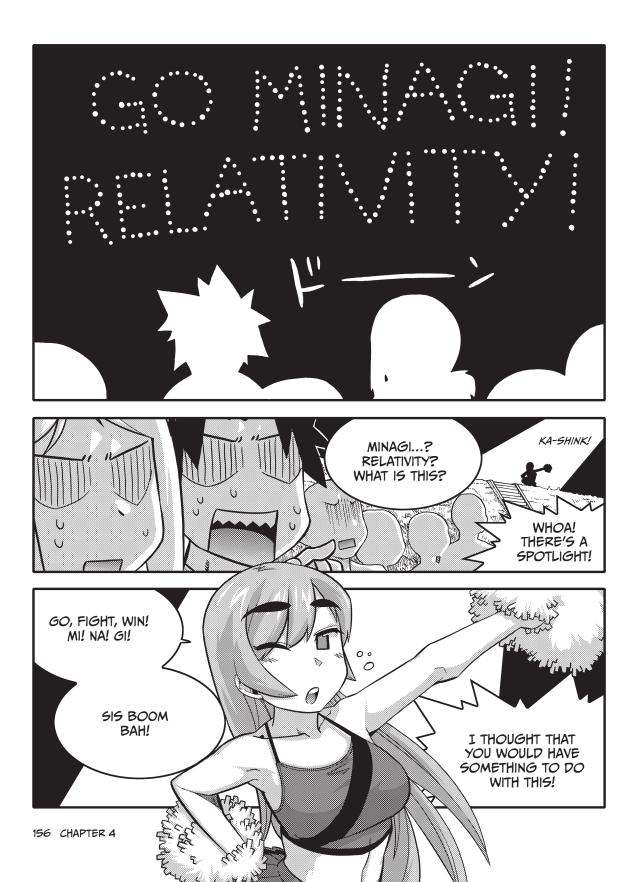








YOU KNOW, MISS URAGA, YOU'RE





THE SLOWING OF TIME IN GENERAL RELATIVITY

Let's use some equations to look at the "slowing of time" in general relativity based on the explanation in the manga.

As in the manga, we assume that Ms. A is at the top of a tall tower, Ms. B is at the bottom of the tower, and Mr. C is inside the elevator next to the tower, as shown in Figure 4-1.

We also assume that each of the three people has the same clock. However, since space-time is warped by gravity, we don't know whether the time of each and the rate at which each one passes (tempo) are the same.

Therefore, we will check the rate at which time passes due to gravity under the following three conditions:

- 1. Inside the free-falling elevator, there is a state of weightlessness.
- Since special relativity holds there, the clock inside the elevator advances with constant time intervals.
- 3. Ms. A's clock at the top of the tower and Ms. B's clock at the bottom of the tower each advance with different constant time intervals.

In addition, we will use the following procedure to check the rate at which time passes due to gravity.

- Align the rates at which time passes for Mr. C's clock inside the elevator and Ms. A's clock at the beginning of the descent.
- 2. Compare the rates at which time passes for Mr. C's clock inside the elevator and Ms. B's clock at the end of the descent.

At first, since Ms. A and Mr. C are at the same height, they are affected by the same gravity.

Let *z* denote the height direction at that location, and let ϕ_1 denote the gravitational potential. The gravitational potential is the quotient of the potential energy divided by the mass of an object. For example, the potential energy of gravity near the surface of the Earth is *mgh*, and the gravitational potential is *gh*.

Therefore, we will align Ms. A's and Mr. C's times and the rates at which time passes.

Let $\Delta \tau_1$ denote the time that passes at Ms. As location, and let $\Delta \tau_2$ denote the time that passes at Ms. B's location.

Now let's assume that the cable that is holding the elevator is cut and the elevator starts to free-fall. Since the falling velocity immediately after the cable is cut (the velocity at which Ms. A is flying upward when viewed from Mr. C's perspective) is v = 0, the tempos of Ms. A's and Mr. C's clocks are the same.

$\mathbf{0} \qquad \Delta \tau_1 = \Delta \tau_3$

The elevator is pulled by gravity, and its velocity gradually increases. The elevator passes alongside Ms. B at a certain velocity (v).

If Ms. B were viewed at that time by Mr. C inside the elevator, he would observe her to be moving upward toward himself, which is the reverse of his own motion (falling from the top of the tower toward the bottom) viewed from his surroundings (see Figure 4-2).

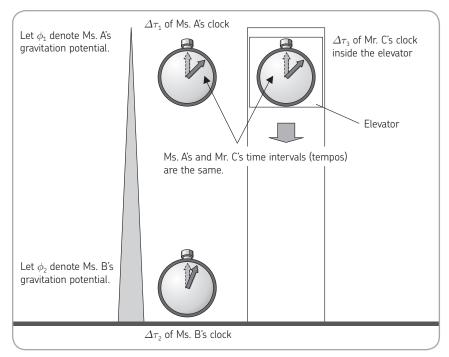


Figure 4-1: Aligning the rates at which time passes for Mr. C's clock inside the elevator and Ms. A's clock at the beginning of the descent

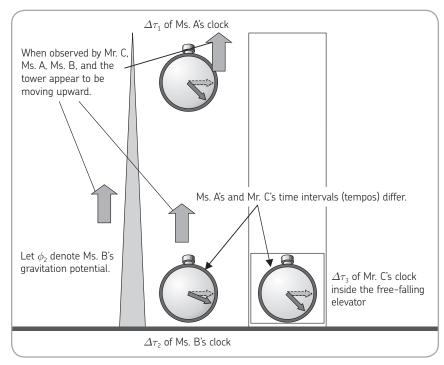


Figure 4-2: Comparing the rates at which time passes for the clocks of Mr. C inside the elevator and Ms. B at the end of the descent

$$\mathbf{2} \qquad \Delta \tau_2 = \Delta \tau_3 \sqrt{1 - \left(\frac{v}{c}\right)^2} \,,$$

and from equations ${f 0}$ and ${f 0}$, we eliminate Δau_3 to obtain

$$\mathbf{6} \quad \frac{\Delta \tau_2}{\Delta \tau_1} = \sqrt{1 - \left(\frac{v}{c}\right)^2} < 1$$

This shows that the tempo of Ms. B's clock becomes slower than the tempo of Mr. C's clock. The rate at which time passes for Ms. B's clock, where the gravitational potential is low (ϕ_2 , close to the source of gravity), is slower than the rate at which time passes for Ms. A's clock, where the gravitational potential is high (ϕ_1 , far from the source of gravity).

In other words, the lower the gravitational potential is, the slower time will pass.

Let's assume that the velocity v is low and that we can use Newtonian mechanics (if we let $x = \frac{v}{c}$, then $x \ll 1$).

Therefore, when ϕ_1 denotes the gravitational potential at Ms. A's location and ϕ_2 denotes the gravitational potential at Ms. B's location, we have $\phi_1 > \phi_2$.

Due to the conservation of energy, at the instant before the elevator hits the ground, all of its potential energy has been converted to kinetic energy, and its velocity v is given by this expression:

$$\left(\phi_1 - \phi_2\right)m = \frac{1}{2}mv^2$$

We obtain the following:

$$\mathbf{\Theta} \quad \phi_1 - \phi_2 = \frac{1}{2}v^2$$

When $x \ll 1$, we can use the following approximation formula:

$$(1+x)^{\alpha} \approx 1 + \alpha x$$

Since $x \ll 1$ because $x = \frac{v}{c}$, we have the following:

$$\sqrt{1 - \left(\frac{v}{c}\right)^2} = \left(1 - x^2\right)^{\frac{1}{2}} \approx 1 - \frac{1}{2}x^2 = 1 - \frac{1}{2}\left(\frac{v}{c}\right)^2$$

If we use this together with equation 6, we obtain the following equation:

If we substitute $\frac{1}{2}v^2 = \phi_1 - \phi_2$ from equation **(9**) into equation **(9**), we obtain the following equation:

$$\mathbf{\bullet} \quad \frac{\Delta \tau_2}{\Delta \tau_1} \approx 1 - \frac{1}{2} \left(\frac{\mathbf{v}}{\mathbf{c}} \right)^2 = 1 - \frac{\phi_1 - \phi_2}{\mathbf{c}^2}$$

Also, if we rearrange the above equation slightly, since $\frac{\phi_1 - \phi_2}{c^2} \approx 1 - \frac{\Delta \tau_2}{\Delta \tau_1} = \frac{\Delta \tau_1 - \Delta \tau_2}{\Delta \tau_1}$ holds, we have the following:

In other words, gravitational potential is related to the slowing of time (time dilation), as shown in equation **1**.

As shown in Figure 4-3, the difference in gravitational potential $\phi_1 - \phi_2$ between the ground and an object at a height *h* is given by the gravitational potential *gh*.

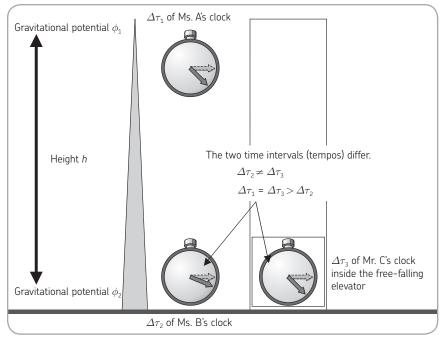


Figure 4-3: State of relatively weak gravity on the ground

If we let $\phi_2 = 0$, *h* denote the height up to ϕ_1 , and *g* denote the gravitational acceleration near the ground and then substitute $\phi_1 = gh$ and $\phi_2 = 0$ into equation O, we obtain the following:

$$\frac{\Delta \tau_1 - \Delta \tau_2}{\Delta \tau_1} \approx \frac{\phi_1 - \phi_2}{c^2} = \frac{gh - 0}{c^2} = \frac{gh}{c^2}$$

The clock at the higher altitude measures a time $\Delta \tau_1$ that is ahead of the clock below, which reads $\Delta \tau_2$. We know this to be true, because in the equation above, gh/c^2 and $\Delta \tau_1$ are both greater than zero, and therefore $\Delta \tau_1 - \Delta \tau_2 > 0$; hence, $\Delta \tau_1 > \Delta \tau_2$.

THE TRUE NATURE OF GRAVITY IN GENERAL RELATIVITY

Space-time surrounding the existence of matter is warped, as explained in the manga. It is apparent that this warping of space-time has the same effect as gravity in attracting surrounding matter.

Einstein unified these effects in a set of equations called the *Einstein field equations*. The Einstein field equations showed that time and space (space-time), which were previously thought to exist as a framework for measuring the motion of matter, were fundamentally connected with matter itself.

PHENOMENA DISCOVERED FROM GENERAL RELATIVITY

This section introduces the following phenomena, which were discovered from general relativity:

- Gravitational lensing
- · Anomalous perihelion precession of Mercury
- Black holes

BENDING OF LIGHT (GRAVITATIONAL LENSING) NEAR A LARGE MASS (SUCH AS THE SUN)

Gravitational lensing is the phenomenon that when light passes in the vicinity of the Sun, the path of that light bends.

Space is bent in the vicinity of the Sun because of the large mass of the Sun, as shown in Figure 4-4. Since light advances along that curvature, the light from a distant star bends, and the direction of the star is observed with a slight shift. This effect was verified during a total solar eclipse. It is noted as the first proof discovered of general relativity.

Also, when light is coming from a distant galaxy, as shown in Figure 4-5, if a massive object (such as a galaxy) lies at an intermediate point, it will bend the light from the distant galaxy as though there were a condensing lens at that intermediate point. This bend may make the distant galaxy seem distorted. Many instances of this effect have been observed. This is another instance of gravitational lensing.

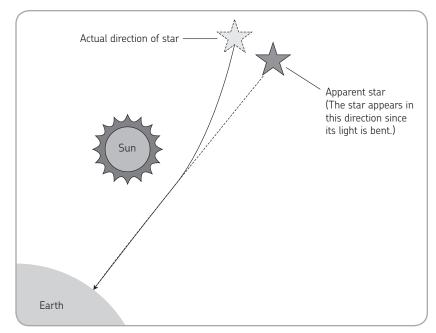


Figure 4-4: Bending of light near a large mass

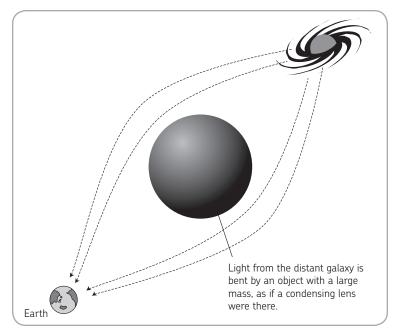


Figure 4-5: Gravitational lensing

ANOMALOUS PERIHELION PRECESSION OF MERCURY

The *perihelion* is the point in the orbit of a planet that is closest to the Sun, as shown in Figure 4–6. We know that the perihelion of Mercury moves by approximately 574 arc seconds per century. Note that the "seconds" mentioned here are angular units rather than units of time. A minute of an arc is 1/60 of 1 degree, and a second is 1/60 of that. In other words, an arc second is 1/3600 of a degree. If the shift revolves by approximately 547 arc seconds per century, it is a shift of only approximately 0.16 degree per century.

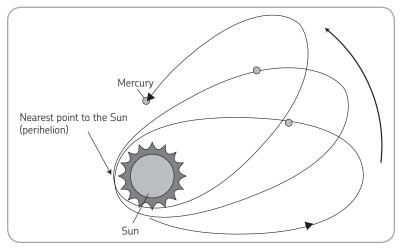


Figure 4-6: The anomalous shift in the perihelion of Mercury

Various causes of this shift, such as the effects of gravity of other planets, were investigated using Newtonian mechanics. However, none of these investigations were able to account for 43 arc seconds of perihelion movement. But when general relativity was used to check the shift in the perihelion of Mercury by calculating the Sun's warping of spacetime, it was found to be shifted by exactly 43 arc seconds.

BLACK HOLES

A *black hole* is a condition in which mass is extremely concentrated and gravity becomes so strong that not even light can escape from it.

A supernova stellar explosion occurs at the end of the life of a star having a mass several times that of the Sun. This event forms a region in space in which mass is extremely concentrated and gravity becomes stronger. Gravity becomes so strong that even light may not be able to escape. That region is a black hole.

Since light cannot escape from it, a black hole cannot be directly observed. However, if other stars exist in the vicinity of the black hole, gas from those stars will stream toward the black hole and form an accretion disk, that is, a cloud of diffuse matter around the black hole. When gas streams into the black hole from that accretion disk, X-rays and gamma rays are emitted.

A black hole candidate was discovered in 1971 in the constellation Cygnus, and currently, supermassive black holes are believed to exist at the center of galactic systems.

GLOBAL POSITIONING SYSTEM AND RELATIVITY

The Global Positioning System (GPS) uses 24 satellites orbiting the Earth to determine position. Each satellite broadcasts a signal toward Earth that includes the radio broadcast time. A receiver on the ground (such as a car navigation system) receives those signals. The radio waves of the signals reach the receiver at the speed of light (approximately 300,000,000 m/s).

When the time the signal was received is compared with the broadcast time, and that time difference is multiplied by the speed of light, the distance to the satellite is known. In other words, if we assume that the distance between the satellite and receiver is 20,000 km, then the radio waves reach the receiver in 20,000,000 m \div 300,000,000 m/s = 0.067 seconds. That calculation is performed using radio waves from three satellites to accurately determine the position on the ground.

However, if there is an error in that time difference, an error will also occur in the calculation of the distance between the satellite and receiver. For example, if the satellite broadcast time is offset by 1 microsecond (10^{-6} second), the distance will be offset by 300 meters (300,000,000 m/s × 0.000001 s = 300 m).

A GPS satellite orbits the Earth at an altitude of 20,000 km and a velocity that causes it to make 1 revolution in approximately 12 hours. At that velocity, the effect of special relativity causes its time to slow by 7.1 microseconds per day. However, since it is located high above the surface of the Earth, the effect of general relativity causes its time to pass faster than time on the surface of the Earth by 46.3 microseconds per day (the effect of general relativity is represented by equation **@** on page 161). As a result, the time broadcast from the GPS is slowed by 39.2 microseconds per day (see Figure 4-7). The design of the GPS system takes into consideration the effects of both special and general relativity with extreme precision.

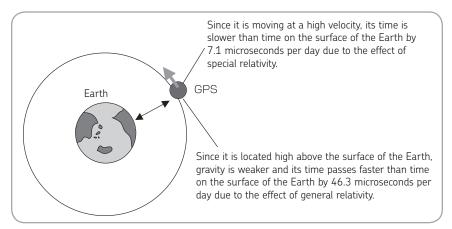


Figure 4-7: The Global Positioning System compensates for the effects of relativity.

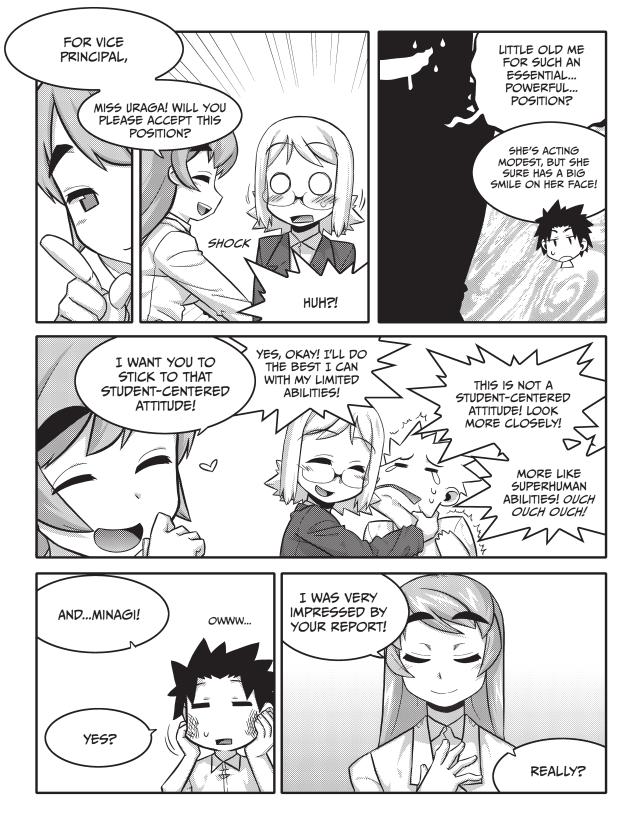














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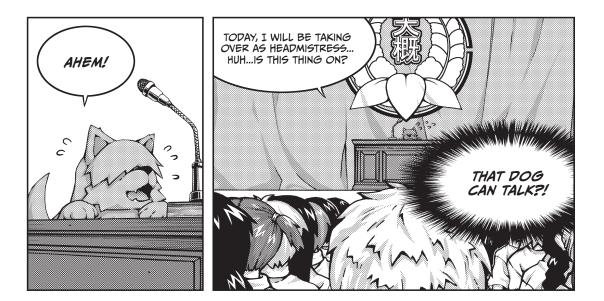
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ABOUT THE SUPERVISING EDITOR

Hideo Nitta completed his doctorate at Waseda University Graduate School of Science and Engineering, 1987, majoring in Theoretical Physics and Physics Education. He is currently a professor in the Department of Education at Tokyo Gakugei University. He is also the author of *The Manga Guide to Physics* (Ohmsha, No Starch Press).

ABOUT THE AUTHOR

Masafumi Yamamoto completed his doctorate at Hokkaido University Graduate School of Engineering Division of Applied Physics, 1984. He is currently Representative Director of Yaaba, Ltd.

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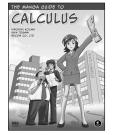
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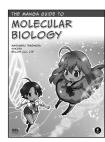
















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